

## **Response to the comments to the manuscript “A new parameterization of the UV irradiance altitude dependence for clear-sky conditions and its application in the on-line UV tool over Northern Eurasia”**

*Please, note that in this combined file we included both the required marked-up version with the track changes and the final version where the changes are marked by yellow color and line numeration is the same as in the final manuscript.*

### **Comments of the reviewer 1:**

The paper presents a parametrization of the altitude effect on different types of biologically effective UV related irradiance. It is a useful tool for the scientific community using UVI measurements and it is related with specific health issues.

The analysis and the presentation of the results is adequate for publication in ACP after the authors take into account the following suggestions/comments.

1. *Equations 7-9. What are the units used for the solar elevation here? Since the coefficients are very small compared of the constant factor. What is their physical meaning?*

We used a simple polynomial regression method which provides better accounting the RAF dependence on solar elevation ( see their solar angle dependence in Figure below). The coefficients are small but they are very necessary. For example, for  $RAF_{Qery}$  (see the equation below) at high solar elevations small coefficients at term 1 and term 2 are compensated by large multipliers ( $h^2$  and  $h$ , respectively). For example, at  $h=90$  first negative term gives -0.89, and the second term- 1.41, which are of the same order with the constant 0.665. Solar elevation is given in degree.

$$RAF_{Qery}(h) = -1.10E-04 \pm 1.49E-5 h^2 + 1.57E-02 \pm 1.53E-3 h + 0.665 \pm 0.0333$$

Since this is a statistical polynomial regression approach physical meaning of coefficients is not important as we should take into account all terms at once. The polynomial regression approach provides the most accurate account of the RAF dependence on  $h$ . We have added some clarification in the text about this standard method and some other details.

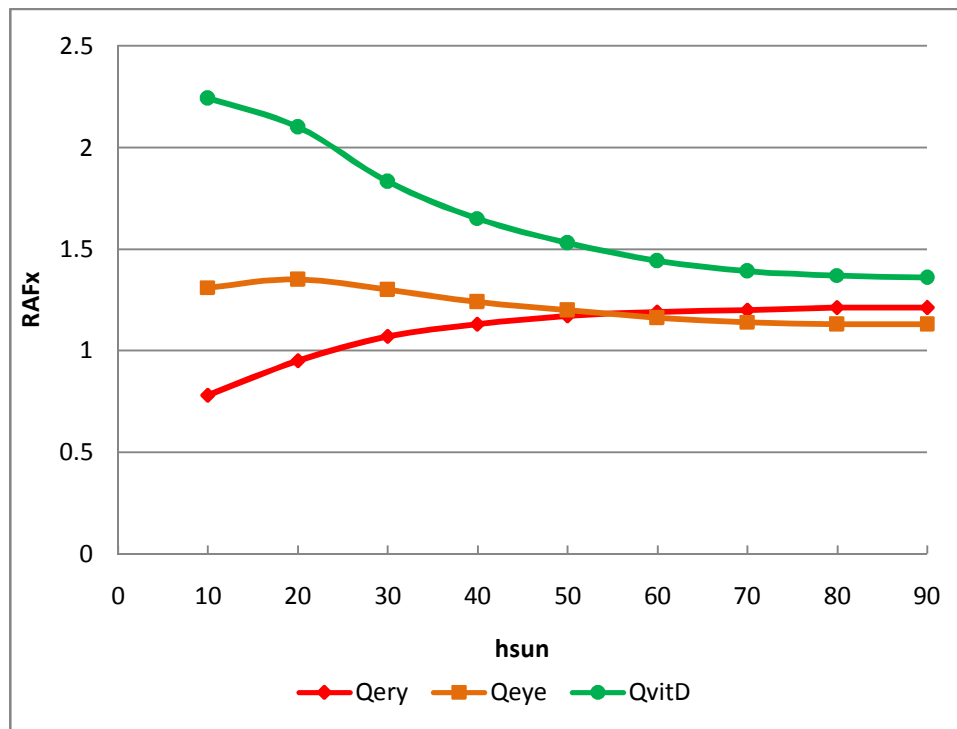


Figure. RAF dependence on solar elevation and polynomial regression for different types of BAUVR.

New variant of the text at line 226-227:

“Using the results of accurate RT modelling and polynomial regression approach we have obtained *RAF* dependencies on solar elevation in degree over  $h=10-90^\circ$  range for different types of BAUVR: “

2 *Line 259. Since we expect the majority of aerosols to be found at lower altitudes, how realistic is to assume that the SSA is non altitude dependent. ?*

We assume this is possible. According to aircraft measurements (Panchenko et al., 2012) SSA at 440nm (the closest wavelength to the UV) changed from 0.87 to 0.93 but this could be a specific feature observed over Western Siberia in visible spectrum for specific conditions. Unfortunately we found no information on SSA altitude dependence in UV spectral region in typical conditions.

However, as mentioned in the text, due to extremely low AOD at high altitudes the UV effects to the SSA changes are negligible. We have added some additional comments to the text:

The updated text is the following at line 281-285:

“In some conditions single scattering albedo and asymmetry factor for visible wavelengths may have the altitude dependence (see, for example, the results of aircraft measurements in Western Siberia (Panchenko et al., 2012)). However, there is no information on the altitude dependence of aerosol properties in UV spectral region from the in-situ measurements over the PEEEX area. Note, that the uncertainty of neglecting the altitude changes in single scattering albedo significantly decreases at small AOD observed at high altitudes and only the altitude changes in aerosol optical depth are usually taken into account in the standard tropospheric aerosol models (WMO, 1986).”

3 SSA: there are publications for SSA at UV wavelengths (e.g. Arola et al., 2009 based Kinne et al simulations)

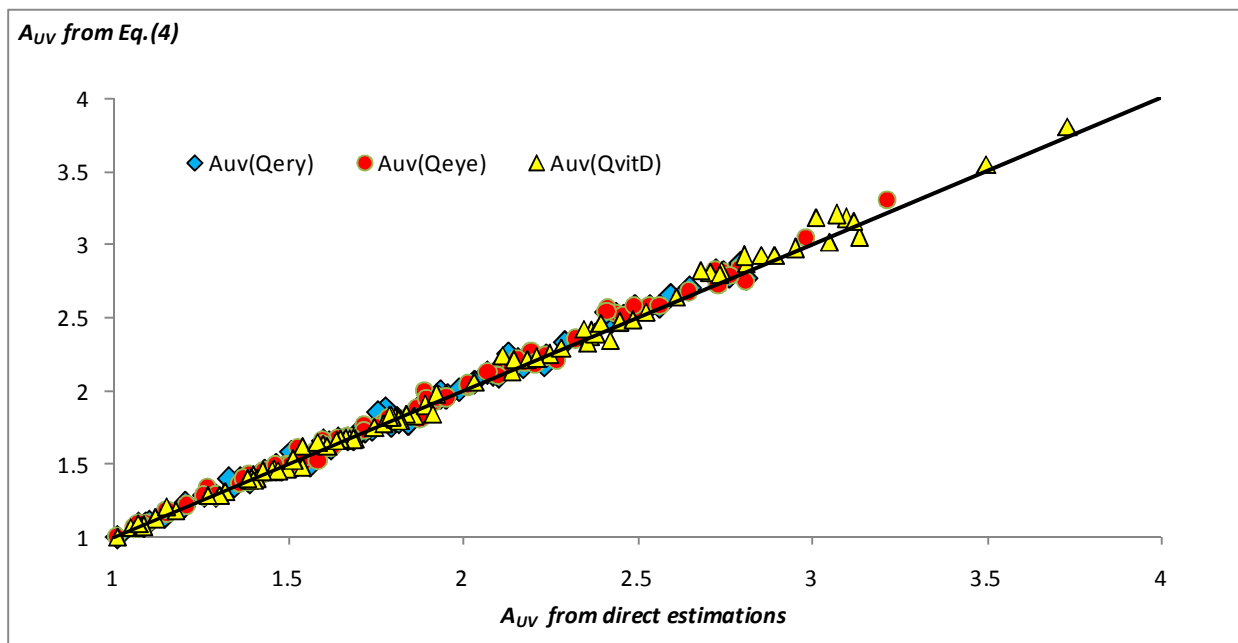
<http://onlinelibrary.wiley.com/doi/10.1029/2009GL041137/abstract>, that report much lower values.

We added the validation of the proposed parameterization over a wider range of the parameters (SSA=0.88, SSA=0.96, Angstrom exponent =0.6, and 1.5) and obtained the same results. We added a few additional references as well.

The text and Figure 2 (new numbering) have been changed.

At line 184-189:

“The model simulations were made for the altitude changes from zero to 5 km with the variations of aerosol optical depth at 340nm within  $AOD_{340} \sim 0.0-0.4$ , variations in total ozone from 350 to 250 DU, and surface albedo changes from zero to  $S=0.9$  at different altitudes. As the input aerosol parameters, within UV spectral region we also used single scattering albedo  $SSA$  varying from 0.88 to 0.96, factor of asymmetry  $g=0.72$ , and Angstrom exponent  $\alpha$  varying from 0.6 to 1.5, which are close to the aerosol background characteristics in Europe (Chubarova, 2009, Arola et al., 2009).”



**Figure 2. The comparison of  $A_{UV}$  amplification factor calculated from Eq.(4) as multiplication of  $A_M A_X A_{AOD} A_S$  with the direct model simulation of UV amplification. All the parameters ( $A_{UV}, A_M A_X A_{AOD} A_S$ ) were obtained from accurate model simulations.**

*Comment. The simulations were performed for different altitudes ( $H=0$  and  $H=5km$ ), aerosol optical depth ( $AOD_{340}= 0, 0.2, 0.4$ ), single scattering albedo ( $SSA=0.88, 0.96$ ), Angstrom exponent ( $\alpha=0.6,1.0,1.5$ ), total ozone ( $X=250,300,350$  DU), surface albedo ( $S=0, S=0.9$ ) and solar elevation ( $h=20^\circ$  and  $50^\circ$ ). For estimating the UV amplification we assume at  $H=0$  km the conditions with 350DU,  $AOD_{340}=0.4$ ,  $S=0\%$  and normalized the  $BAUVR$  at the altitude  $H=5km$  to the value calculated with these parameters.*

In addition, we included the discussion on aerosol properties in UV –B and also added the references there.

At line 262-270:

“The coefficients were obtained according to model simulations for  $0 < AOD_{340} < 0.8$ , single scattering albedo ( $0.8 < SSA < 1$ ), air mass  $m \sim \sinh^{-1}$  ( $m \leq 2$ ), and Angstrom exponent  $\alpha \sim 1$  ( $0.6 < \alpha < 1.5$ ). Note, that these are typical changes in main aerosol properties for European conditions in UV-A spectral range (Chubarova, 2009). However, the Angstrom exponent in UV-B spectral region can differ from this range and be even negative in particular conditions depending on aerosol size distribution and optical properties (Bais et al., 2007). Single scattering albedo in UVB spectral range according to the results of different field campaigns (UNEP, 2015) may vary from 0.7 to 0.97 with low SSA in the presence of black and brown carbon aerosol. Some results demonstrates no existence of SSA spectral dependence in UV (Barnard et al., 2008, Arola et al., 2009) but some results shows its spectral character (UNEP, 2015).”

4 *Is the SSA=0.96 realistic for UV wavelengths?*

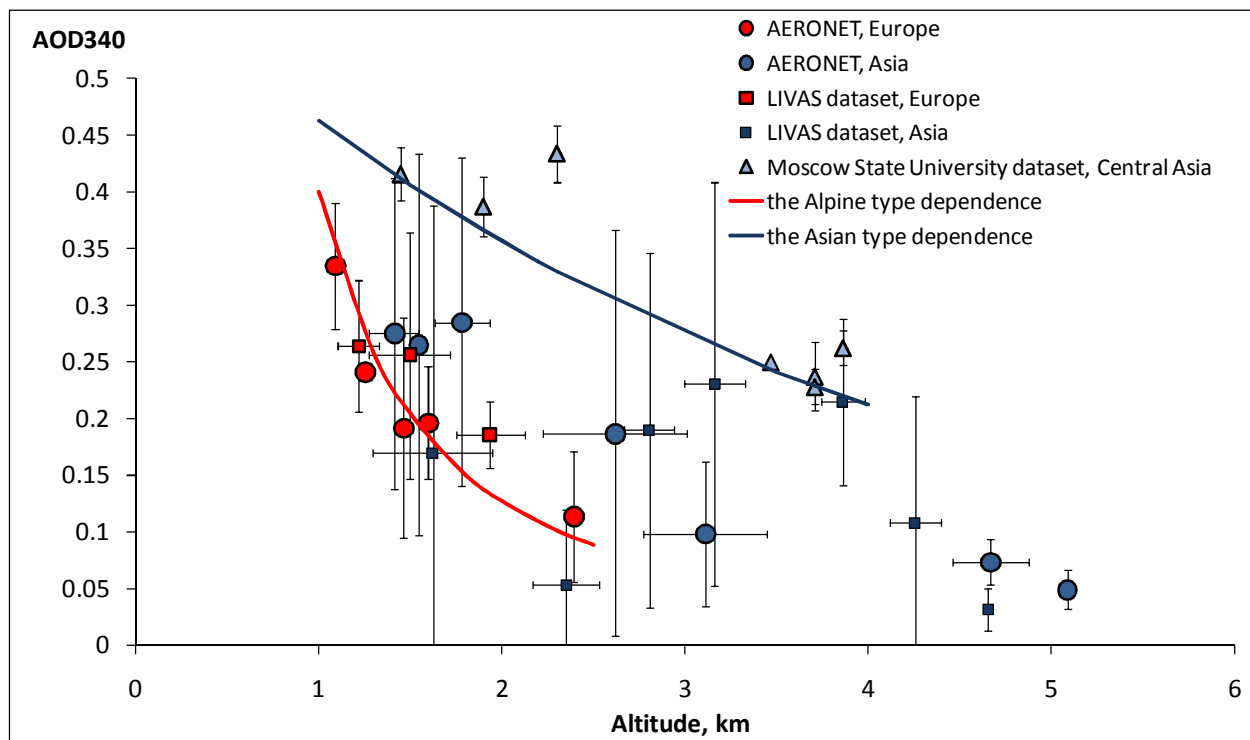
We increased the range of SSA in validation and added the discussion in the text. The SSA=0.96 is the upper boundary of the possible value. The validation of SSA=0.94 have demonstrated a satisfactory agreement with the UV-B measurements in Moscow (see Chubarova, 2009). However, anyone can use different SSA value within 0.8-1.0 range which is an independent parameter in equation (12).

Please, look at the updated text at line 184-189:

“The model simulations were made for the altitude changes from zero to 5 km with the variations of aerosol optical depth at 340nm within  $AOD_{340} \sim 0.0-0.4$ , variations in total ozone from 350 to 250 DU, and surface albedo changes from zero to  $S=0.9$  at different altitudes. As the input aerosol parameters, within UV spectral region we also used single scattering albedo  $SSA$  varying from 0.88 to 0.96, factor of asymmetry  $g=0.72$ , and Angstrom exponent  $\alpha$  varying from 0.6 to 1.5, which are close to the aerosol background characteristics in Europe (Chubarova, 2009, Arola et al., 2009).”

5 *Figure 3 : it would be easier for the reader if more colors could be used as for example aernet Europe and Livas could be mixed now.*

The idea to use only two colors was to show by color the attribution to different regions: Europe and Asia domains. To distinguish between AERONET and LIVAS in Europe we decided to use different size of the markers (circles) to make the difference between these two datasets more clear.



**Figure 4. (old numeration Fig.3) The altitude dependence of aerosol optical depth at 340nm with 1 sigma error bar according to the AERONET, LIVAS and the Moscow State University datasets over European and Asian regions. May-September period. The AOD at 330 nm the Moscow State University dataset and the AOD at 355nm from the LIVAS datasets were recalculated to AOD at  $\lambda=340$  nm using the Angstrom exponent  $\alpha=1.0$ . See further details in the text.**

6 *The provided uncertainty of 1% and 3% has to be clarified more. Here model inputs have errors as they come mostly from measurements. So if the authors would like to provide an uncertainty budget they have to include the propagation of errors coming from the actual measurements and/or fitting procedure they have used. As an example LIVAS 0.1 difference from AERONET is not representative of the actual determination of the AOD at a certain height but as a total column AOD comparison among AERONET and LIVAS.*

We agree, that the uncertainty is mainly due to the errors in input parameters. But our aim was not to show the whole budget of uncertainties but just to compare the exact model simulations with the proposed parameterization ( fitting procedure) that should be made with the same parameters.

We have changed a little the text at line 193-196:

“The correlation for all BAUVR types is higher than 0.99 with the mean relative difference of  $1\pm 3\%$  compared with the exact model simulations if the same input parameters are used. Hence, the proposed approach based on the independent account for the terms, which are affected by different geophysical factors can be applied with high accuracy.”

We had included some assessment of the quality of LIVAS dataset because this is a quite new dataset and some comparison with widely known AERONET dataset may be useful for readers. Yes, we agree

that “LIVAS 0.1 difference from AERONET is not representative of the actual determination of the AOD at a certain height but as a total column AOD comparison among AERONET and LIVAS”. But we used total AOT from LIVAS dataset at the elevated altitudes. That is why we assume that we can leave the text as is.

7        *In addition, the abstract reads: “UV amplification from different factors within a wide range of their changes with mean uncertainty of 1% and standard deviation of 3% compared with the exact model simulations with the same input parameters. ” It is not clear what the authors mean here.*

As it was discussed above we tried to show the uncertainty of the proposed approach - the altitude UV parameterization. To exclude the uncertainty due to the input parameters we have to use the same input parameters in simulations. That is why we have added “*compared with the exact model simulations with the same input parameters*”.

8        *L140-144: More about the threshold concerning vitamin D have to be reported.*

We added the additional information concerning the vitamin D threshold and some other details about the method used for UV resources determination. However, the detailed discussion and the full method description can be found in “Chubarova, N., Zhdanova, Ye.: Ultraviolet resources over Northern Eurasia, *Journal of Photochemistry and Photobiology B: Biology*, Elsevier, 127, 38-51, 2013.”

The text is the following at line 140-155:

“In addition, we estimated UV resources at different altitudes according to the approach given in Chubarova and Zhdanova (2013), which has been developed on the base of international classification of UV index (Vanichetk et al., 2000) and the vitamin D threshold following the recommendations given in CIE (2006). In CIE (2006) there were simple recommendations of choosing the minimum vitamin D dose (*MVitDD*) threshold using one fifth Minimal Erythemal Dose (*MED*) for a one fifth body area. In this study according to the new guidelines a healthy level of vitamin D3 was increased from 400 IU recommended in CIE (2006) to 1000 IU (*Rationalizing nomenclature for UV doses...*, 2014). The possibility to account for the open body fraction as a function of the effective air temperature was also applied in the UV resources estimating method (Chubarova, Zhdanova 2013) as it had been proposed in (McKenzie et al., 2009). According to this approach we defined *noon UV deficiency* when UV dose is smaller than the vitamin D threshold during 11:30-12:30 noon period, and *100% UV deficiency category*, when it is not possible to receive vitamin D throughout the whole day. The *UV optimum* category is determined when the UV dose does not exceed erythemal threshold but it is possible to receive UV dose, necessary for vitamin D at noon hour. Several subclasses of *UV excess* are attributed to the thresholds depending on the standard UV index categories: *moderate UV excess* class, which relates to moderate category of hourly UV index, *high UV excess*, *very high UV excess*, and *extremely high UV excess* category. Further details about this approach can be found in (Chubarova, Zhdanova, 2013).”

## **Response to the reviewer 2:**

General comments:

The authors of the paper propose a parametrization of the altitude effect on three types of biologically effective UV irradiance. In such context this research thoroughly explores the amplification of the effective irradiances as a function of the altitude variation of molecular number density, of ozone and aerosol, and albedo. The implementation of the UV parametrization in the on-line UV tools can be of

potential interest to the researchers involved in studies on the assessment of human UV exposures. The analysis is comprehensive and it will be acceptable for publication after taking into consideration the issues underlined below.

Specific comments:

1. *Introduction L32: The predominant factors which interact with UV radiation determining its variability at the Earth' surface are mentioned, however UV radiation is also controlled by the variation in the cyclic Sun emittance: the 27-day cycle leads to variations less than 1% for  $\lambda > 250\text{nm}$ , 6-8% in the band 245-250nm; the 11-year sunspot cycle determines small changes in irradiance and influences the shortest extra-terrestrial wavelengths. The above factors should be also included (Ref. S., Madronich. The atmosphere and UV-B radiation at ground level. [book auth.] A.R. Young (Eds.) L.O. Bjorn. Environmental UV Photobiology. New York : s.n., 1993, pp. 1-39.)*

We added the solar activity factor in the text and added the reference. However, we decided not to include the details since this is a minor factor for the wavelengths larger than 300nm.

At line 33-35:

“UV radiation is affected by astronomical factors (solar zenith angle, solar-earth distance, solar activity), by different atmospheric characteristics (total ozone content, cloudiness, aerosol, optically-effective gases), and by surface albedo (Madronich, 1993, Bais et al., 2007, Bekki et al., 2011).”

We added an additional reference:

S., Madronich. The atmosphere and UV-B radiation at ground level. [book auth.] A.R. Young (Eds.) L.O. Bjorn. Environmental UV Photobiology. New York : s.n., 1993, pp. 1-39.

2. *L43: How the UV index is calculated should be better specified for readers not familiar with this parameter as well as its reference (COST-713. Action UVB Forecasting. European Communities. Brussels : s.n., 2000).*

We have made the necessary changes in the text. We added the definition in the footnote and included the reference to “K. Vanicek, T. Frei, Z. Litynska, A. Schmalwieser. UV-Index for the Public, COST-713 Action, Brussels, 2000, 27p” has been already included in the reference list.

Footnote 1: “UV index is a widely used characteristic which is equal to erythemally-weighted irradiance expressed in ( $\text{W m}^{-2}$ ) multiplied on 40 (Vanicek et al., 2003)”

*L47-49: the photobiological quantities (erythemally –weighted irradiance, erythemal doses) should be defined.*

We made several corrections in the text and left the terms “erythemally-weighted irradiance” and UV index, which was defined by equation (1) and in the footnote 1 as it was shown above..

The text has been changed in the following way at line 48:

“At the European alpine stations in summer conditions the UV indices are often higher than 11 (Hülßen, 2012). For example, high UV index up to 12 was observed in mountainous areas in Italy (Casale et al. 2015). A significant UV growth with the altitude was also obtained at different sites in Austria and Switzerland (Rieder et al., 2010). In winter, erythemally-weighted UV irradiance is about 60% higher than that at lower-altitude European sites (Gröbner et al., 2010).”

*L46-47: the following references could be acknowledged: Siani, G.R. Casale, H. Diémoz, G. Agnesod, M.G. Kimlin, C.A. Lang and A. Colosimo, Personal UV exposure in high albedo alpine sites, Atmos. Chem. Phys., 2008, 8, 3749–60; Casale G. R., A. M. Siani, H. Diémoz, G. Agnesod, A.V. Parisi, A.*

*Colosimo (2015) Extreme UV Index and Solar Exposures at Plateau Rosà (3500 m a.s.l) in Valle d'Aosta Region, Italy, Science of the Total Environment 512–513 (2015) 622–630;*

Thank you for the useful references. We included one of them concerning the high UV index in the Introduction and another – in the Discussion section..

At line 49:

“For example, high UV index up to 12 was observed in mountainous areas in Italy (Casale et al. 2015).”

At line 474 :

“In this case UV dosimeters which have a spectral response almost identical to that of the UV-induced photobiological effect (Siani et al., 2008) is the most accurate way for evaluating the individual levels of UV exposure.”

*L65. The following reference should be also acknowledged: Seckmeyer, G., Mayer, B., Bernhard, G., Erb, R., Albold, A., Jager, H., Stockwell, W.R.: New maximum UV irradiance levels observed in Central Europe, Atmos. Environ., 31(18), 2971-2976, 1997.*

Thank you. We also included this reference in the list.

At line 67:

“The accurate results of measurements from different field campaigns devoted to the evaluation of altitude UV effects shown in (Bernhard et al., 2008, Blumthaler and Ambach, 1988, Blumthaler et al., 1994, Blumthaler et al., 1997, Dahlback et al., 2007, Lenoble et al., 2004, Piacentini, et al. 2003, Pfeifer et al., 2006, Seckmeyer et al., 1997, Sola et al., 2008, Zaratti et al., 2003) provide precise, however, local character of this phenomenon, which results in various altitude UV gradients.”

*L89: It is worth pointing up that biological action spectra, although helpful to understand the biological reaction, do not express direct information on the possible combined effects of different wavelengths. Additivity for wavelength contributions has been documented for the erythema action spectrum, but not for the vitamin D action spectrum (ref. M.Norval, L. Björn, F. R. de Gruijl, Is the action spectrum for the UV-induced production of pre-vitamin D3 in human skin correct?, Photochem. Photobiol. Sci., 2010, 9, 11–17).*

Yes, we agree and we added the discussion in the text on this point as a footnote 2.

At line 92:

“We used erythemal action spectrum according to CIE (1998), vitamin D spectrum - according to CIE (2006)<sup>2</sup>, and cataract-weighted spectrum according to Oriowo et al. (2001).”

The footnote 2: “Note, that a widely used conception of action spectra, which is based on the additivity of wavelength contribution, still has not be well documented for vitamin D action spectrum (Norval et al., 2010) and needs further studies.”

*L91-98: The weighting function (action spectrum) is generally normalized to unity at the wavelength of maximal sensitivity, in case of erythemal and Vitamin D action spectra are both normalized at 298 nm. Perhaps figures showing the discrepancy among the action spectra could better explain the wavelength-dependent effectiveness of UV radiation in causing the specific reactions.*

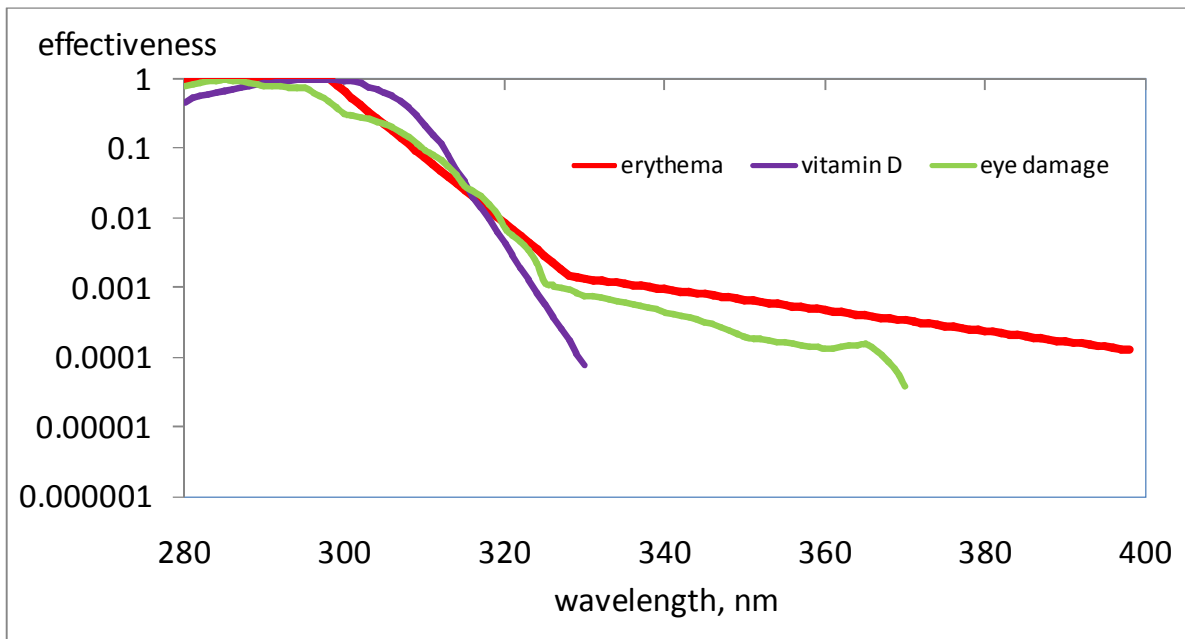
Yes, we agree and we included the recommended Figure in the text.



At line 93:

Various types of BAUVR action spectrum have different efficiency within the UV range (Fig.1).

This Figure 1 is given below:



**Figure 1. Action spectra for erythema (CIE, 1998), vitamin D (CIE, 2006) and for eye damage (cataract) (Oriowo et. al. 2001) effects.**

L137-144:

More clarity is necessary in this part of the text mainly in “noon UV deficiency and UV deficiency category”. How is the vitamin D threshold determined? What are the values of “UV excess”?

We added some description of the method which was described in details in another paper.

The new variant is the following at line 140:

“In addition, we estimated UV resources at different altitudes according to the approach given in Chubarova and Zhdanova (2013), which has been developed on the base of international classification of UV index (Vanichkek et al., 2000) and the vitamin D threshold following the recommendations given in CIE (2006). In CIE (2006) there were simple recommendations of choosing the minimum vitamin D dose (*MVitDD*) threshold using one fifth Minimal Erythemal Dose (*MED*) for a one fifth body area. In this study according to the new guidelines a healthy level of vitamin D3 was increased from 400 IU recommended in CIE (2006) to 1000 IU (Rationalizing nomenclature for UV doses..., 2014). The possibility to account for the open body fraction as a function of the effective air temperature was also applied in the UV resources estimating method (Chubarova, Zhdanova 2013) as it had been proposed in (McKenzie et al., 2009). According to this approach we defined *noon UV deficiency* when UV dose is smaller than the vitamin D threshold during 11:30-12:30 noon period, and *100% UV deficiency category*, when it is not possible to receive vitamin D throughout the whole day. The *UV optimum* category is determined when the UV dose does not exceed erythemal threshold but it is possible to receive UV dose, necessary for vitamin D at noon hour. Several subclasses of *UV excess* are attributed to the thresholds depending on the standard UV index categories: *moderate UV excess* class, which relates to moderate category of hourly UV index, *high UV excess*, *very high UV excess*, and *extremely high UV excess* category. Further details about this approach can be found in (Chubarova, Zhdanova, 2013).”

*” In the context of biomedical radiation effects it should be highlighted that the skin orientation relative to the Sun and the geometry of the human body, could strongly modify the results limited to UV irradiance measurements on horizontal surface.*

We added this information in the “Discussion” section.

At line 472:

The current state of the online interactive tool does not take into account for the skin orientation relative to the Sun and the geometry of the human body which can modify the results limited to UV irradiance simulations on horizontal surface (Hess and Koepke, 2008, Vernez et al., 2014). In this case UV dosimeters which have a spectral response almost identical to that of the UV-induced photobiological effect (Siani et al., 2008) is the most accurate way for evaluating the individual levels of UV exposure. The vitamin D production can be also affected by other factors such as obesity and age (Engelsen, 2010). However, these are the tasks for the future work.

*In addition since the beneficial effect of UV radiation is related to the body area of exposed skin, the length of time to produce sufficient vitamin D decreases with the increase of the exposed body area for all phototypes (See for example McKenzie, R.L., J. B. Liley and L. O. Bjorn (2009) UV Radiation: Balancing Risks and Benefits. Photochem. Photobiol., 85, 88–98.).*

The open body fraction has been included as a parameter in the proposed interactive program. Please, look on the updated version of the text.

At line 142:

“In CIE (2006) there were simple recommendations of choosing the minimum vitamin D dose (*MVitDD*) threshold using one fifth Minimal Erythemal Dose (*MED*) for a one fifth body area. In this study according to the new guidelines a healthy level of vitamin D<sub>3</sub> was increased from 400 IU recommended in CIE (2006) to 1000 IU (Rationalizing nomenclature for UV doses..., 2014). The possibility to account for the open body fraction as a function of the effective air temperature was also applied in the UV resources estimating method (Chubarova, Zhdanova 2013) as it had been proposed in (McKenzie et al., 2009).”

*Yet, obesity and age should also be mentioned as influential factors in vitamin D production (see The Relationship between Ultraviolet Radiation Exposure and Vitamin D Status in Nutrients 2010, 2, 482-495; doi:10.3390/nu2050482).*

We added the proposed items in the Discussion Section at line 476:

“The vitamin D production can be also affected by other factors such as obesity and age (Engelsen, 2010). However, these are the tasks for the future work.”

*In this regard it would be reasonable at least to acknowledge the above issue, whether in the Introduction or in the Discussion. Results*

We included all recommended changes and references in the Discussion Section. See the updated variants above.

*L214: The quadratic and linear terms of eqs 6-7-8 have very small coefficients with respect to the constant term. The authors should provide the physical meaning of these equations, for example: when  $h=0$ , does the *RAF*(erythema or vitamin D) account the diffuse component?*

We used the standard procedure with the standard polynomial regression parameterization. The equations can be used over the  $h=10-90^\circ$  range that is why we did not consider the conditions at  $h=0$ . Physical meaning of this equation is in the fact that RAF has the dependence on solar angle due to changes in spectral distribution of solar irradiance. Similar approach is given in many papers (see, for example, Herman et al., 2010). However, we proposed the equations with smaller number of terms and with high quality.

We showed above (in response to the first reviewer) the example that the small coefficients in the first and second term compared with the constant play an important role at high solar elevation providing the pronounced RAF dependence on solar elevation. Please, look an additional Figure and the text above.

The text has been clarified at line 226:

“Using the results of accurate RT modelling and polynomial regression approach we have obtained *RAF* dependencies on solar elevation in degree over  $h=10-90^\circ$  range for different types of BAUVR:”

*The units of the coefficients in eqs should be specified*

The units of the coefficients are different depending on the degree of the polynomial regression and the number of terms. They have a technical character and we guess that we do not need their unit specification.

We added the necessary information about the type of approximation in the text. Please, look at our response above

*Numbers of the coefficients in Eqs 8 and 9 should be expressed in the same form*

Sorry.

Done:

$$RAF_{Qery}(h) = -1.10E-04 \pm 1.49E-5 h^2 + 1.57E-02 \pm 1.53E-3 h + 0.665 \pm 0.0333 \quad (7)$$

$$R^2 = 0.98$$

$$RAF_{QvitD}(h) = 1.66E-4 \pm 1.0E-5 h^2 - 2.77E-2 \pm 1.1E-3 h + 2.5121 \pm 0.0233 \quad (8)$$

$$R^2 = 0.997$$

$$RAF_{Qeye}(h) = 1.43E-6 \pm 1.0E-6 h^3 - 2.02E-4 \pm 6.6E-5 h^2 + 4.83E-3 \pm 2.9E-3 h + 1.297 \pm 0.035 \quad (9)$$

$$R^2 = 0.98$$

*L286: The authors should give more details about “ the coefficients have been re-affirmed using more statistics”.*

We added some additional information.

New variant at line 311:

“ However, the coefficients have been re-affirmed according to the monthly mean AOD over 1999-2012 period (case number N=137).”

*In discussion or in conclusions: the authors should point out that their analysis is based on irradiance and the question of how well the radiation received by the anatomical area is related to that incident on a horizontal surface should be discussed.*

As we mentioned above we have added a discussion on this point at line 472:

“The current state of the online interactive tool does not take into account for the skin orientation relative to the Sun and the geometry of the human body which can modify the results limited to UV irradiance simulations on horizontal surface (Hess and Koepke, 2008, Vernez et al., 2014). In this case UV dosimeters which have a spectral response almost identical to that of the UV-induced photobiological effect (Siani et al., 2008) is the most accurate way for evaluating the individual levels of UV exposure. The vitamin D production can be also affected by other factors such as obesity and age (Engelsen, 2010). However, these are the tasks for the future work.”

*To determine the individual levels of UV exposure, that is, the real biologically effective doses of sunlight, dosimeters which have a spectral response almost identical to that of the UV-induced photobiological effect, should be mentioned.*

We added the text. Please, look on our response to the previous comment.

*Minor comments: web addresses could be in a web reference list.*

In this variant we did not add it to the reference list since this is not a reference but a part of the text. But if necessary, this could be easily done.

L143: “Ccurrently “should be replaced by “Currently”

Done

Eq 11: the subscript “AOT-0” should be replaced by “AOT0”

Unfortunately I was not able to find this misprint in eq 11.

We used AOD throughout the text after the recommendation before publishing in ACPD.

### **Response to the reviewer 3:**

The manuscript deals with imports subject. i.e., modeling biologically effective UV (BEUV) radiations reaching the ground level. The authors provide useful formula for accurate estimation of the BEUV height dependence. It could improve various presently used UV index forecast models run by national weather services. Thus, the manuscript fits well with the journal scope. The reviewer decision is to accept the manuscript with minor changes listed below.

Minor comments:

1. *Cloudless conditions. It was clearly stated by the authors. However, they recommend using the on line tool based on the proposed formula that could be used also for allsky conditions. Thus they should comment on validity of this formula, especially “As” (albedo) term, for the partially cloudy conditions.*

We decided to remove this part of calculations at high altitudes in cloudy conditions for non zero surface albedo in the interactive tool until the updated version of the algorithm for cloudy conditions is ready. Now, when a user tries to choose the cloudy conditions at high altitudes the following comment appears: “Sorry, this part is under development. You could use the preliminary UV assessment for clear sky conditions at the specified altitude”.

We decided not to include any comments in the text not to confuse a reader.

2 1.179. Angstrom exponent of 1.0 is proposed in the calculations. Authors found that “A” coefficients do not depend significantly on the aerosols characteristics. Probably we can use here any value of Angstrom exponent (the same concern other aerosols characteristics: ssa, asymmetry factor ) because A represents the relative value and the exponent value does not change with the altitude. Authors should comment on the selected aerosols values, which are proper for visible and UVA ranges, but not representative over UV-B range. For example, Angstrom exponent of 1 provides that AOD at 308 nm is about 10% larger than that at 340 nm but many authors suggested that Angstrom exponent in UVB range could be zero and even negative.

We have added the range of single scattering albedo and Angstrom exponent in validation. The updated Figure 2 (new numbering) has been done. We obtained the same statistics (correlation coefficient, averages, standard deviation, etc). Model simulations revealed that the changes within  $\alpha=0.6-1.5$  provide plus- minus 1% difference for various types of BAUVR. Of course, in some conditions it can be abnormal negative change in Angstrom but this would be for specific conditions. In addition we should note that it is very hard to evaluate Angstrom exponent in UV-B region due to different gas absorption which can dramatically influence the evaluation of the Angstrom exponent ( see the discussion, for example, in referenced publication Chubarova et al., 2016. We have added the discussion on this account.

New text at line 184:

“The model simulations were made for the altitude changes from zero to 5 km with the variations of aerosol optical depth at 340nm within  $AOD_{340}\sim 0.0-0.4$ , variations in total ozone from 350 to 250 DU, and surface albedo changes from zero to  $S=0.9$  at different altitudes. As the input aerosol parameters, within UV spectral region we also used single scattering albedo  $SSA$  varying from 0.88 to 0.96, factor of asymmetry  $g=0.72$ , and Angstrom exponent  $\alpha$  varying from 0.6 to 1.5, which are close to the aerosol background characteristics in Europe (Chubarova, 2009, Arola et al., 2009). We compared the  $A_{UV}$  values calculated as a multiplication composite of different separate parameters ( $A_M, A_X, A_{AOD}$ , and  $A_S$ ) according to Eq.(4) with the  $A_{UV}$  values, which were estimated as a ratio of direct simulations of BAUVR at the altitude  $H=5$  km and at zero ground level. The results of the comparisons are shown in Fig.2. One can see a good agreement between the  $A_{UV}$  values obtained using multiplication of  $A_M A_X A_{AOD} A_S$  and the  $A_{UV}$  values from direct estimations of BAUVR. The correlation for all BAUVR types is higher than 0.99 with the mean relative difference of  $-1\pm 3\%$  compared with the exact model simulations if the same input parameters are used. Hence, the proposed approach based on the independent account for the terms, which are affected by different geophysical factors can be applied with high accuracy.

However, we understand that it is important to include the additional discussion on UV aerosol properties in the text. We included the additional discussion on the aerosol properties in UV-B region in 3.4. “Aerosol UV amplification with the altitude” and increase the application of the range of Angstrom exponent from 0.6 to 1.5.

New text at line 262:

“The coefficients were obtained according to model simulations for  $0 < AOD_{340} < 0.8$ , single scattering albedo ( $0.8 < SSA < 1$ ), airmass  $m \sim \sinh^{-1} (m \leq 2)$ , and Angstrom exponent  $\alpha \sim 1$  ( $0.6 < \alpha < 1.5$ ). Note, that these are typical changes in main aerosol properties for European conditions in UV-A spectral range (Chubarova, 2009). However, the Angstrom exponent in UV-

B spectral region can differ from this range and be even negative in particular conditions depending on aerosol size distribution and optical properties (Bais et al., 2007). Single scattering albedo in UVB spectral range according to the results of different field campaigns (UNEP, 2015) may vary from 0.7 to 0.97 with low SSA in the presence of black and brown carbon aerosol. Some results demonstrates no existence of SSA spectral dependence in UV (Barnard et al., 2008, Arola et al., 2009) but some results shows its spectral character (UNEP, 2015). We should also note that direct evaluation of SSA in UV-B spectral region is difficult because UV attenuation due to aerosol can occur together with the absorption in this spectral region by different gases (ozone, sulphur dioxide, formaldehyde, nitrogen dioxide, etc.). As a result we used the aerosol properties at 340nm as input parameters for the BAUVR simulations since we consider typical aerosol conditions without forest fires and heavy industrial smoke. Radiative effects of the existing AOD spectral dependence are relatively small within the UV-B spectral range therefore we consider the same coefficients for different types of BAUVR.”

3 *l.286.” we can see its satisfactory agreement . . .”. I can not see the mentioned agreement. It is better to calculate the exponent value, separately for the AERONET and LIVAS data, and next discuss the agreement with the Pamir/Tien Shan exponent.*

Sorry, we agree that the formulation is not exact and need to be changed. The idea was not to show the quantitative agreement between the different datasets but to reveal the existence of sharp and flat aerosol altitude dependence for so-called Alpine and Asian types and to analyze their effects on UV. As it was already mentioned in the text (old line numbers 300-301):

“We should note that the proposed altitude  $AOD$  dependencies according to (15) and (16) are considered only as a first proxy for the most sharp and flat altitude dependencies.”

The text was changed in the following way at line 303:

“The second one is characterized by a much more gradual altitude  $AOD_{340}$  decrease observed over flat elevated Asian regions according to the AERONET and LIVAS dataset and the data obtained during Moscow State University field campaigns at the high-altitude plateau at Pamir and Tyan’-Shan’ mountainous regions in Central Asia. The main reason of such a character is the existence of the additional aerosol emission sources (i.e. loess, mineral aerosol) from the vast areas of deserts and semi-deserts elevated over sea level of up to 3-4 kilometers.”

Instead of

“The second one is characterized by a much more gradual altitude  $AOD_{340}$  decrease observed over flat elevated Asian regions. The main reason of such a character is the existence of the additional aerosol emission sources (i.e. loess, mineral aerosol) from the vast areas of deserts and semi-deserts elevated over sea level of up to 3-4 kilometers. In addition, Fig.3 demonstrates the  $AOD_{340}$  dependence on altitude according to the data obtained during Moscow State University field campaigns at the high-altitude plateau at Pamir and Tyan’-Shan’ mountainous regions in Central Asia. We can see its satisfactory agreement with the second type of  $f_{AOD}(H)$  obtained from the AERONET and LIVAS dataset. “

*l. 348.” The value  $r_{bio}$  has a relatively small dependence on altitude”. Exactly,  $b$  value is small but  $c$  is also small. Figure 5 shows that  $r_{bio}$  changes significantly (0.4 for  $H=0$*

but 0.2 for  $H=8$  km). Thus, for me it is not so small change.

Sorry, we agree that this part of the text needs editing. This is not exact formulation. We have changed the text in the following way at line 369:

“According to the model estimations the value  $r_{bio}$  in clear sky conditions has the dependence on altitude, which appears due to a decrease mainly in molecular and aerosol loading and can be parameterized by a simple regression as follows:”

l.425-439. The noon dose is mentioned many times but it is not clear how long is the exposition?, 1 hour around noon? Please provide 1 MED value and the vitamin D3 threshold dose for photo-type II and IV used in the calculations.

We added the clarification of the UV resources evaluation in the “Materials and methods” Section at line 142:

“In CIE (2006) there were simple recommendations of choosing the minimum vitamin D dose ( $MVitDD$ ) threshold using one fifth Minimal Erythemal Dose ( $MED$ ) for a one fifth body area. In this study according to the new guidelines a healthy level of vitamin D3 was increased from 400 IU recommended in CIE (2006) to 1000 IU (Rationalizing nomenclature for UV doses..., 2014). The possibility to account for the open body fraction as a function of the effective air temperature was also applied in the UV resources estimating method (Chubarova, Zhdanova 2013) as it had been proposed in (McKenzie et al.,2009). According to this approach we defined *noon UV deficiency* when UV dose is smaller than the vitamin D threshold during 11:30-12:30 noon period, and *100% UV deficiency category*, when it is not possible to receive vitamin D throughout the whole day. The *UV optimum* category is determined when the UV dose does not exceed erythemal threshold but it is possible to receive UV dose, necessary for vitamin D at noon hour. Several subclasses of *UV excess* are attributed to the thresholds depending on the standard UV index categories: *moderate UV excess* class, which relates to moderate category of hourly UV index, *high UV excess*, *very high UV excess*, and *extremely high UV excess* category. Further details about this approach can be found in (Chubarova, Zhdanova, 2013).”

We have also changed the following text ( line 425-439 previous numeration) at line 447:

“Let us analyze the UV resources for skin type 2 and the open body fraction of 0.25 in the Alpine region (approximately 46°N and 7°E) for winter and spring noon conditions. For these conditions the vitamin D threshold is equal to 100 J/m<sup>2</sup> and Minimal Erythemal Dose -250 J/m<sup>2</sup>. According to our estimates on January, 15<sup>th</sup>, at  $H=0$  km for typical (climatological) ozone, aerosol and surface albedo conditions the noon UV deficiency (no vitamin D generation) is observed with noon erythemally UV dose of about 97.2 Jm<sup>-2</sup>, while at the same coordinates at  $H$  higher 0.5 km up to  $H=4.8$  km (the highest point within the Alps, peak Mont Blanc) we obtain the UV optimum conditions with noon erythemal UV dose varying from 100.6 Jm<sup>-2</sup> to 122.9 Jm<sup>-2</sup>.

However, for skin type 4 (Fitspatrick, 1988) with vitamin D threshold of 180 J/m<sup>2</sup> the noon UV deficiency is observed at all altitudes and even at high surface albedo  $S=0.9$  corresponding to the

pure snow with UV dose of  $154.4 \text{ Jm}^{-2}$ . The decrease in open body fraction for this skin type from 0.25 to 0.05, which could take place in frosty weather, provides 100% UV deficiency, when no vitamin D can be generated during the whole day at the  $H=4.8 \text{ km}$  and  $S=0.9$ .

On April, 15<sup>th</sup>, at  $H=0 \text{ km}$  and typical climatological conditions at this geographical point noon UV dose is about  $437.7 \text{ Jm}^{-2}$ . This means that for the open body fraction of 0.4 the moderate UV excess is observed for skin type 2 and the UV optimum – for skin type 4, when vitamin D threshold is  $112.5 \text{ Jm}^{-2}$  and MED threshold -  $450 \text{ Jm}^{-2}$ . At the altitude  $H = 2 \text{ km}$  the conditions are characterized by the moderate UV excess for both skin types 2 and 4 with UV dose of  $463.4 \text{ Jm}^{-2}$ . At the  $H=4 \text{ km}$  a high UV excess is observed for skin type 2 and the moderate UV excess - for skin type 4 with UV dose of  $532.4 \text{ Jm}^{-2}$ .”

”

*l.433. Open body fraction for skin phototype IV of 0.25 and 0.5 on 15th January is highly unreliable during the winter sunbathing. Value of 0.10 here is much more probable.*

Yes. We agree. Since it was interesting to show the change in the class of UV resources, in the new variant we decreased the open body fraction to 0.05.

The new variant of the text at line 456:

“The decrease in open body fraction for this skin type from 0.25 to 0.05, which could take place in frosty weather, provides 100% UV deficiency, when no vitamin D can be generated during the whole day at the  $H=4.8 \text{ km}$  and  $S=0.9$ .”

*l. 718. Fig.3. Here Angstrom exponent=1.3 but 1 was used previously in the text (l.179, l.253).*

Sorry, this is a misprint.

New variant ( the numbering has been changed from 3 to 4 due to including a new Figure 1):

“Fig.4. The altitude dependence of aerosol optical depth at 340nm with 1 sigma error bar according to the AERONET, LIVAS and the Moscow State University datasets over European and Asian regions. May-September period. The AOD at 330 nm the Moscow State University dataset and the AOD at 355nm from the LIVAS datasets were recalculated to AOD at  $\lambda=340 \text{ nm}$  using the Angstrom exponent  $\alpha=1.0$ . See further details in the text. “



# A new parameterization of the UV irradiance altitude dependence for clear-sky conditions and its application in the on-line UV tool over Northern Eurasia

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**Abstract.** A new method for calculating the altitude UV dependence is proposed for different types of biologically active UV radiation (erythemally-weighted, vitamin-D-weighted and cataract-weighted types). We show that for the specified groups of parameters the altitude UV amplification ( $A_{UV}$ ) can be presented as a composite of independent contributions of UV amplification from different factors within a wide range of their changes with mean uncertainty of 1% and standard deviation of 3% compared with the exact model simulations with the same input parameters. The parameterization takes into account for the altitude dependence of molecular number density, ozone content, aerosol and spatial surface albedo. We also provide generalized altitude dependencies of the parameters for evaluating the  $A_{UV}$ . The resulting comparison of the altitude UV effects using the proposed method shows a good agreement with the accurate 8-stream DISORT model simulations with correlation coefficient  $r > 0.996$ . A satisfactory agreement was also obtained with the experimental UV data in mountain regions. Using this parameterization we analyzed the role of different geophysical parameters in UV variations with altitude. The decrease in molecular number density, especially at high altitudes, and the increase in surface albedo play the most significant role in the UV growth. Typical aerosol and ozone altitude UV effects do not exceed 10-20%. Using the proposed parameterization implemented in the on-line UV tool (<http://momsu.ru/uv/>) for Northern Eurasia over the PEEEX domain we analyzed the altitude UV increase and its possible effects on human health considering different skin types and various open body fraction for January and April conditions in the Alpine region.

**Keywords:** UV radiation, altitude dependence, RT modelling, erythemally-weighted irradiance, vitamin D-weighted irradiance, cataract-weighted irradiance, interactive UV-tool.

## 1. Introduction

Biologically active UV radiation (BAUVR) is an important environmental factor, which significantly affect human health and nature (UNEP, 1998; UNEP, 2011). Enhanced levels of UV radiation lead to different types of skin cancer (basal and squamous cell carcinomas, cutaneous melanoma), to various eye diseases (cataract, photokeratitis, squamous cell carcinoma, ocular melanoma, variety of corneal/conjunctival effects), and to immunosuppression. However, small doses of UV radiation have a positive effect on health through the vitamin D generation (UNEP, 2011).

UV radiation is affected by astronomical factors (solar zenith angle, solar-earth [distance](#), [solar activity](#), [distance](#)), by different atmospheric characteristics (total ozone content, cloudiness, aerosol, optically-effective gases), and by surface albedo ([Madronich, 1993](#), (Bais et al., 2007, Bekki et al., 2011)). However, the altitude above sea level has

36 also a significant influence on UV radiation (Bais et al., 2007). There are a lot of studies and special field campaigns  
37 in different geographical regions, which were devoted to the analysis of the altitude UV effect (Bernhard et al.,  
38 2008, Blumthaler and Ambach, 1988, Blumthaler et al., 1994, Blumthaler et al., 1997, Dahlback et al., 2007,  
39 Lenoble et al., 2004, Piacentini, et al. 2003, Pfeifer et al., 2006, Sola et al., 2008, etc.). The UV enhancement at high  
40 altitudes is detected not only due to smaller molecular scattering, but also due to usually observed decreasing in total  
41 ozone content and aerosol, and increasing in surface albedo, which in turn enhances 3D reflection from slopes of  
42 mountains covered by snow (Lenoble et al., 2004). In addition, variation of cloud properties with altitude can also  
43 change the level of UV radiation.

44 The UV records in mountainous areas demonstrate extremely high levels. The highest UV values are observed in  
45 Andes in Bolivia (Pfeifer et al., 2006, Zaratti et al., 2003), where ~~the~~ UV index<sup>1</sup> can be sometimes close to 20. Very  
46 high UV levels were also recorded at high-altitude deserts in Argentina (Piacentini, et al. 2003). In Tibet the UV  
47 index frequently exceeded 15 on clear days and occasionally exceeded 20 on partially cloudy days (Dahlback et al.,  
48 2007). At the European alpine stations in summer conditions the UV indices are often higher than 11 (Hülse,  
49 2012). For example, high UV index up to 12 was observed in mountainous areas in Italy (Casale et al. 2015). A  
50 significant UV growth with the altitude was also obtained at different sites in Austria and Switzerland (Rieder et al.,  
51 2010). In winter, erythemally-weighted UV irradiance is about 60% higher than that at lower-altitude European sites  
52 (Gröbner ~~et al., 2010). The analysis of erythemal UV doses at different sites in Austria and Switzerland also~~  
53 ~~demonstrates a significant growth of UV radiation with the altitude (Rieder et al., 2010).~~ In the Arctic the  
54 comparison of summer UV measurements at Summit (3202m a.s.l.) and Barrow (~0 a.s.l.) stations ~~also~~ shows  
55 significant enhancing of about 30-40% in clear sky conditions at the elevated site (Bernhard et al., 2008).

56 The UV altitude gradients obtained from model calculations vary within the range of 3.5-6%/km in the cloudless  
57 atmosphere if all other parameters (ozone, aerosol, surface albedo) do not change with the altitude (Chubarova and  
58 Zhdanova, 2013). ~~The~~ ~~Even~~ smaller values of the estimated UV altitude gradients (3.5%/km) were obtained in  
59 conditions with high surface albedo at both sea level and high altitude, since the larger diffuse component at sea  
60 level, to some extent, compensates the higher direct flux due to a smaller total optical depth at higher  
61 ~~location altitudes~~. However, the experimental UV altitude gradients are often much higher due to the presence of  
62 additional altitude changes in the atmospheric parameters. According to different field campaigns UV altitude  
63 gradients vary within 5-11%/km (Pfeifer et al., 2006, Zaratti et al., 2003, Schmucki, Philipona, 2002), 11-14%/km  
64 according to (Sola et al., 2008), and in some cases can reach 31%/km (Pfeifer et al., 2006). The existence of spectral  
65 dependence in absorption coefficients of ozone as well as in molecular scattering cross sections provides a  
66 pronounced spectral character of the altitude UV effect, which was obtained in many publications (Blumthaler et al.,  
67 1994, Sola et al., 2008).

68 However, the continuous UV records in mountainous area are still very rare due to the complexity of accurate UV  
69 measurements in severe conditions. The accurate results of measurements from different field campaigns devoted to  
70 the evaluation of altitude UV effects shown in (Bernhard et al., 2008, Blumthaler and Ambach, 1988, Blumthaler et  
71 al., 1994, Blumthaler et al., 1997, Dahlback et al., 2007, Lenoble et al., 2004, Piacentini, et al. 2003, Pfeifer et al.,  
72 2006, Seckmeyer et al., 1997, Sola et al., 2008, Zaratti et al., 2003) provide precise, however, local character of this  
73 phenomenon, which results in various altitude UV gradients.

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<sup>1</sup> UV index is a widely used characteristic, which is equal to erythemally-weighted irradiance expressed in ( $W m^{-2}$ ) multiplied on 40 (Vanicek et al., 2003)

74 At the same time, the accurate RT (Radiative Transfer) model simulations (Liou, 2010) are very time consuming  
75 and can not be used in different on-line tools or other applications. There are also a lot of UV model assessments for  
76 the past and future UV climate scenarios but usually they are given with the coarse spatial resolution, which does  
77 not allow a user to obtain the accurate estimates over the particular mountainous location.

78 In addition, the limiting factor of the UV calculation accuracy is the uncertainty of input geophysical parameters,  
79 which significantly increases at high altitudes. Hence, another task was to obtain some generalized dependencies of  
80 the input parameters using the available data sources.

81 The objective of this paper is to provide the accurate parameterization for different types of biologically active  
82 radiation for the estimation of UV level at different altitudes taking into account the generalized altitude  
83 dependencies of different geophysical parameters. Using the proposed parameterization we will also  
84 discuss the consequence of the enhanced UV level at high altitudes for human health using the classification of UV  
85 resources via a specially developed on-line interactive UV tool.

## 86 2. Materials and methods

87 In order to account for different effects of UV radiation on human health we analyze three types of BAUVR:  
88 erythemally-weighted, vitamin D-weighted, and cataract-weighted irradiances, which are calculated using the  
89 following equation:

$$90 \quad Q_{bio} = \int_{280}^{400} Q_{\lambda} F_{\lambda} d\lambda, \quad (1)$$

93 where  $Q_{\lambda}$  is the spectral flux density,  $F_{\lambda}$  is the respective biological action spectrum.

94 We used erythemal action spectrum according to CIE (1998), vitamin D spectrum - according to CIE  
95 (2006)<sup>2</sup>, and cataract-weighted spectrum according to Oriowo et al. (2001). Various types of BAUVR action  
96 spectrum have different efficiency within the UV range (Fig.1). For their characterization we used the effective  
97 wavelengths, which are calculated as follows:

$$99 \quad \lambda_{eff} = \frac{\int Q_{\lambda} \lambda d\lambda}{\int Q_{\lambda} d\lambda} \quad (2)$$

101 According to our estimates, for example, at high solar elevation ( $h=60^{\circ}$ ) and for the variety of other parameters  
102 (total ozone, aerosol and surface albedo) the effective wavelength for erythemally-weighted irradiance ( $Q_{ery}$ ) is  
103  $\sim 317$  nm, for cataract-weighted irradiance ( $Q_{eye}$ ) -  $\sim 313$  nm, and for vitamin D-weighted irradiance ( $Q_{vitD}$ ) -  
104  $\sim 308$  nm. These changes in effective wavelengths for various BAUVR types indicate their different sensitivity to  
105 the ozone absorption, molecular scattering and aerosol attenuation, which vary dramatically within this spectral  
106 range, and, as a result, explain different BAUVR responses to the changes in these geophysical parameters.

107 All the simulations were fulfilled using one dimensional radiative TUV (Tropospheric Ultraviolet-Visible) model  
108 with 8-stream DISORT RT method (Madronich and Flocke, 1997) and 1 nm spectral resolution. The uncertainty of  
109 the RT method is less than 1% (Liou, 2010). ~~A Badosa et al.(2007) showed a~~ good agreement between the

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<sup>2</sup> Note, that a widely used conception of action spectra, which is based on the additivity of wavelength contribution, still has not be well documented for vitamin D action spectrum (Norval et al., 2010) and needs further studies.

110 | experimental spectral data in different geographical regions and simulated results using this RT method if the input  
111 | atmospheric parameters were known was also shown in (Badosa et al. 2007).-

112 | Several experimental datasets were used. For obtaining the generalized altitude dependence of aerosol optical depth  
113 | (AOD) we used the data of sun/sky CIMEL photometers from different AERONET sites located at different heights  
114 | above sea level (Holben et al., 1998)). These data account for the near-ground emission sources of the aerosol at  
115 | various altitudes in the aerosol column content. The estimated uncertainty for aerosol optical depth in UV spectral  
116 | region is about 0.02. The uncertainty for single scattering albedo is about 0.03 at AOD<sub>440</sub>>0.4 and the uncertainty  
117 | for all other inversion parameters is not higher than 10% (Holben et al., 2006).

118 | In addition, the dataset of historical Moscow State University complex field campaigns over mountainous areas at  
119 | Pamir (38- 40.5° N, 73-74° E H=1.0÷3.9 km), and Tyan'Shan' (43°N, 77°E, H=3.47km) was applied in the analysis  
120 | (Belinski et al., 1968). It includes the data records of total ozone content and aerosol optical depth at 330 nm, which  
121 | had been measured with the help of M-83 filter ozonometer, and UV irradiance less 320 nm – by the UVM-5  
122 | instrument calibrated against the spectroradiometer BSQM (the Boyko's Solar Quartz Monochromator) described in  
123 | (Belinski et al., 1968). The description of the BSQM and the details of the calibration were also discussed in  
124 | Chubarova and Nezval' (2000). The uncertainties of UV measurements less 320nm due to the calibration procedure  
125 | were considered to be about 10% (Belinsky et al., 1968). However, to avoid the calibration errors only relative  
126 | measurements were used in this study. The residual uncertainty due to possible existence of slight variation in  
127 | spectral response of the instrument and their temperature dependence was estimated to be about 5-7%. The obscurity  
128 | of the horizon at all sites was less than 10°. The field campaigns were carried out during summer periods, when no  
129 | snow was detected at the surface. The snow covered mountainous peaks were only observed at Tyan'Shan' at  
130 | relatively large distance of more than 30 km from the site.

131 | We also used the LIVAS database (Lidar Climatology of Vertical Aerosol Structure for Space-Based Lidar  
132 | Simulation Studies, <http://lidar.space.noa.gr:8080/livas/>). This is a 3-dimensional global aerosol climatology based  
133 | on satellite lidar CALIPSO observations at 532 and 1064 nm, EARLINET ground-based measurements and a  
134 | combination of input data from AERONET, aerosol models, etc. The final LIVAS climatology includes 4-year  
135 | (2008 – 2011) time-averaged 1×1° global fields (Amiridis, et al., 2015). We used the annual aerosol extinction  
136 | profiles at 355 nm for calculating aerosol optical depth over various points at different altitudes in the Alpine and  
137 | the Caucasian mountainous regions in Europe and over the high-elevated regions in Asia. It should be mentioned  
138 | that the LIVAS averages all Calipso overpasses over a 1×1° cell and characterizes only the mean altitude within the  
139 | cell. This provides some additional uncertainties in its aerosol extinction altitude dependence evaluation. On  
140 | average, according to (Amiridis et al., 2015) the absolute difference in LIVAS AOD is within 0.1 agreement with  
141 | AERONET AOD values in UV and visible region of spectrum.

142 | In addition, we estimated UV resources at different altitudes according to the approach given in Chubarova and  
143 | Zhdanova (2013), which has been developed on the base of international classification of UV index (Vanichek et  
144 | al., 2000) and the vitamin D threshold following the recommendations given in CIE (2006). In CIE (2006) there  
145 | were simple recommendations of choosing the minimum vitamin D dose (M<sub>VitDD</sub>) threshold using one fifth  
146 | Minimal Erythral Dose (MED) for a one fifth body area. In this study according to the new guidelines a healthy  
147 | level of vitamin D<sub>3</sub> was increased from 400 IU recommended in CIE (2006) to 1000 IU (Rationalizing  
148 | nomenclature for UV doses..., 2014). The possibility to account for the open body fraction as a function of the  
149 | effective air temperature was also applied in the UV resources estimating method (Chubarova, Zhdanova 2013) as it  
150 | had been proposed in (McKenzie et al., 2009). According to this approach we defined *noon UV deficiency* and

151 ~~100% UV deficiency categories,~~ when UV dose is smaller than the vitamin D threshold **during 11:30-12:30 noon**  
 152 **period, and 100% UV deficiency category, when,** and it is not possible to receive vitamin D ~~respectively during~~  
 153 ~~solar noon hour,~~ and throughout the whole day. The *UV optimum* category is determined when the UV dose does  
 154 not exceed erythema threshold but it is possible to receive UV dose, necessary for vitamin D at noon hour. Several  
 155 subclasses of *UV excess* are attributed to the thresholds depending on the standard UV index categories: *moderate*  
 156 *UV excess* class, which relates to moderate category of hourly UV index, *high UV excess*, *very high UV excess*, and  
 157 *extremely high UV excess* category. **Further details about this approach can be found in (Chubarova, Zhdanova,**  
 158 **2013).** Currently, in the assessment of UV resources we do not take into account for the eye damage UV effects,  
 159 since there is no reliable regulation on the UV threshold for this type of BAUVR.

### 160 3. Results

#### 161 3.1. The general description of the approach

162 It is widely known that “the solution of the radiative transfer equation is possible to derive by numerous solution  
 163 methods and techniques” (Liou, 2010). However, the accurate RT methods usually require a lot of computer time  
 164 and can not be used in several applications. The simulated intensity and UV flux density (or irradiance) has a  
 165 complicated non-linear dependence on many geophysical parameters, however, our numerous simulations of UV  
 166 irradiance using the accurate 8-stream discrete ordinate RT method show that within a variety of geophysical  
 167 parameters one can obtain the parameterized altitude correction by taking into account for the quasi-independent  
 168 terms driven by different geophysical factors. Some of them are independent due to different vertical profiles (for  
 169 example, ozone maximum in the stratosphere compared with aerosol and molecular maximum in the troposphere).  
 170 Some of them are dependent (for example, surface albedo UV effects depend on molecular and aerosol loading),  
 171 but, as we show later, this factor can be also considered as a one joint term.

172 Using this assumption, we propose a parameterization, where biologically active UV irradiance at the altitude  $H$   
 173 ( $Q_{bio}(H)$ ) can be estimated from  $Q_{bio}$  at zero altitude ( $H=0$  km a.s.l.) with an independent account for the terms,  
 174 which are affected by different geophysical factors:

$$175 Q_{bio}(M_H, X_H, AOD_H, S_H) = Q_{bio}(M_0, X_0, AOD_0, S_0) \cdot A_M A_X A_{AOD} A_{S(M, AOD, cloud)}, \quad (3)$$

176 where  $A_M$ ,  $A_X$ ,  $A_{AOD}$  are the UV amplifications, respectively, due to the altitude decrease in molecular number  
 177 density ( $M$ ), ozone ( $X$ ), and aerosol optical depth ( $AOD$ ).  $A_{S(M, AOD, cloud)}$  is the UV amplification due to the increase in  
 178 surface albedo  $S$ , which is typically observed with the altitude. This characteristic is also a function of a change in  
 179 molecular number density, aerosol and cloud characteristics with height due to the processes of multiple scattering.  
 180 Further, only the effects in the cloudless atmosphere are considered. The total UV amplification ( $A_{UV}$ ) with altitude  $H$   
 181 can be rewritten from Eq.(3) as:

$$182 A_{UV} = A_M A_X A_{AOD} A_S = \frac{Q_{bio}(M_H, X_H, AOD_H, S_H)}{Q_{bio}(M_0, X_0, AOD_0, S_0)} \quad (4)$$

183 Let us consider separately the effects of different factors on UV irradiance at high altitudes. We specify them by  
 184 using the accurate RT model simulations, different empirical datasets or by applying the important characteristics  
 185 from different publications.

186 The possibility of this approach was tested directly by the accurate modelling for a variety of conditions at different  
 187 solar elevations. The model simulations were made for the altitude changes from zero to 5 km with the variations of

188 aerosol optical depth at 340nm within  $AOD_{340} \sim 0.0-0.4$ , variations in total ozone from 350 to 250 DU, and surface  
 189 albedo changes from zero to  $S=0.9$  at different altitudes. As the input aerosol parameters, within UV spectral region  
 190 we also used single scattering albedo  $SSA$  varying from 0.88 to 0.96, factor of asymmetry  $g=0.72$ , and Angstrom  
 191 exponent  $\alpha$  varying from 0.6 to 1.5, of  $\alpha=1.0$ , which are close to the aerosol background characteristics in Europe  
 192 (Chubarova, 2009, Arola et al., 2009). We compared the  $A_{UV}$  values calculated as a multiplication composite of  
 193 different separate parameters ( $A_M, A_X, A_{AOD}$ , and  $A_S$ ) according to Eq.(4) with the  $A_{UV}$  values, which were estimated as  
 194 a ratio of direct simulations of BAUVR at the altitude  $H=5$  km and at zero ground level. The results of the  
 195 comparisons are shown in Fig.2. Fig.1. One can see a good agreement between the  $A_{UV}$  values obtained using  
 196 multiplication of  $A_M, A_X, A_{AOD}, A_S$  and the  $A_{UV}$  values from direct estimations of BAUVR. The correlation for all  
 197 BAUVR types is higher than 0.99 with the mean relative difference of  $-1\pm 3\%$  compared with the exact model  
 198 simulations if with the same input parameters are used. Hence, The slight variations in aerosol parameters (within  
 199 10%) does not change the proposed approach based on the independent account for the terms, which are affected by  
 200 different geophysical factors can be applied with high accuracy. obtained results.

### 201 3.2. Molecular UV amplification with the altitude

202 A decrease of atmospheric pressure, or molecular number density, with the height is a well-known factor of UV  
 203 amplification. According to the 8-stream DISORT model simulations we found that the BAUVR dependence with  
 204 the altitude has a linear change in the molecular atmosphere, which is clearly seen in Fig.3. Fig.2. Hence, for its  
 205 characterization we can apply a simple gradient approach.

206 For evaluating the UV amplification due to molecular effects the following expression is used:

$$207 \quad A_M = \frac{Q_{bio}(M_H, X_0, AOD_0, S_0)}{Q_{bio}(M_0, X_0, AOD_0, S_0)} = 1 + 0.01 G_{bio, M}(S=0) \Delta H \quad (5)$$

208 where  $G_{bio, M}$  is the relative molecular gradient, in %/km,  $\Delta H$  is the difference in the altitudes, in km. Note, that all  
 209 other parameters do not change with the height.

210 The estimated relative molecular gradients for different types of BAUVR for various conditions are shown in Table  
 211 1. At solar elevation  $h=10^\circ$  there is a decrease in the  $G_{bio, M}$  for different BAUVR and, especially, for vitamin-D  
 212 irradiance due to its smaller effective wavelength and the effects of stronger ozone absorption, which is increased at  
 213 higher ozone content ( $X=500$  DU). However, for solar elevation higher than  $20^\circ$  the sensitivity of the  $G_{bio, M}$  values is  
 214 around 6-7%/km and does not significantly change with variations in  $h$  and  $X$ .

215 As an example, at the altitude of 5 km and at high solar elevation the molecular UV amplification according to Eq.  
 216 (5) lies within ~1.26-1.38 depending on the type of BAUVR (see Table 1), which is in accordance with the accurate  
 217 model simulations. However, at  $h=10^\circ$  the UV amplification for erythemally-weighted and cataract-weighted  
 218 irradiances is about 1.18-1.23, while for vitamin D-weighted irradiance  $A_M$  is only 1.04-1.09 depending on ozone  
 219 content. The maximum UV amplification at the highest peak (m. Everest,  $H=8.848$  km) due to changes only in  
 220 molecular scattering reaches 1.53-1.68 at high solar elevation depending on the type of BAUVR.

### 221 3.3. Ozone UV amplification with the altitude

222 In order to account for the ozone decrease with the altitude we apply the existing linear dependence between UV  
 223 radiation and total ozone  $X$  in log-log scale. This approach was used in the definition of the Radiation Amplification  
 224 Factor ( $RAF$ ) by Booth and Madronich (1994). As a result, the following equation can be written:



225  $\log(Q_{bio}) = RAF(Q_{bio,h}) \log(X_i) + C,$  (6)

226 where  $h$  is the solar elevation,  $C$  is the constant.

227 The  $RAF$  values vary for different types of BAUVR. For example, at high solar elevation Radiation Amplification  
 228 Factor for erythemally-weighted irradiance  $RAF_{Q_{ery}} = 1.2$ , for vitamin D-weighted irradiance -  $RAF_{Q_{vitD}}=1.4$ , for  
 229 cataract-weighted irradiance -  $RAF_{Q_{eye}}=1.1$  (UNEP, 2011). However, we should take into account the  $RAF$   
 230 dependence on solar elevation  $h$  due to the relative changes in solar spectrum with  $h$ . Using the results of accurate  
 231 RT modelling and polynomial regression approach we have obtained  $RAF$  dependencies on solar elevation in degree  
 232 over  $h=10-90^\circ$  range for different types of BAUVR:

233  $RAF_{Q_{ery}}(h) = -1.10E-04 \pm 1.49E-5 h^2 + 1.57E-02 \pm 1.53E-3 h + 0.665 \pm 0.0333$  (7)

234  $R^2 = 0.98$

235  $RAF_{Q_{vitD}}(h) = 1.66E-4 \pm 1.0E-5 h^2 - 0.000166 \pm 0.00001 h - 2.77E-2 \pm 1.1E-3 h + 2.5121 \pm 0.0233$

236 (8)

237  $R^2 = 0.997$

238  $RAF_{Q_{eye}}(h) = 1.43E-6 \pm 1.0E-6 h^3 - 2.02E-4 \pm 6.6E-5 h^2 + 4.83E-3 \pm 2.9E-3 h + 1.297 \pm 0.035$  (9)

240  $R^2 = 0.98$

241 where  $R^2$  – is the determination coefficient. The standard error estimates of the coefficients in the equations are  
 242 given at P=95%.

243 Note, that similar approach for accounting the  $RAF$  solar angle dependence was proposed in Herman (2010) with  
 244 higher power degree.

245 As a result, the BAUVR at the altitude  $H(Q_{bioH})$  with the correction on ozone content can be written as follows:

246  $Q_{bioH} = Q_{bio0}(X_0/X_H)^{RAF(Q_{bio,h})}$  (10)

247 From Eq.(10) we can obtain the altitude UV amplification due to ozone using the altitude ozone gradient  $G_X$   
 248 (DU/km):

249  $A_X = \frac{Q_{bio}(M_0, X_H, AOD_0, A_0)}{Q_{bio}(M_0, X_0, AOD_0, A_0)} = \left( \frac{X_0}{X_0 - G_X * \Delta H} \right)^{RAF(Q_{bio,h})}$  (11)

250 We propose to apply the typical ozone altitude gradient  $G_X$ , which absolute value is about 3.5 DU/km according to  
 251 monthly averaged ozone soundings measurements in Germany and observations in Bolivia (Reuder and Koepke,  
 252 2005; Pfeifer et al., 2006).

253 As an example, if we take into account only for this typical ozone decrease with the altitude, the UV enhancement at  
 254 5 km will be about  $A_X \sim 1.06-1.11$  while at the highest peak (m. Everest, H=8848 m)  $A_X$  will reach 1.11-1.22  
 255 at high solar elevations depending on the BAUVR type and initial ozone content at zero altitude within  $X=250-350$   
 256 DU.

### 257 3.4. Aerosol UV amplification with the altitude

258 Aerosols can significantly change their characteristics with the altitude, affecting the level of BAUVR. Due to  
 259 variations in size distribution and optical properties aerosol may have different radiative properties (aerosol optical

260 depth, single scattering albedo, and phase function). One of the most important aerosol characteristics affecting UV  
 261 radiation is aerosol optical depth.

262 For accounting the aerosol effect on UV attenuation we propose to apply the equation given in (Chubarova, 2009):

$$263 \quad Q_{bio}(AOD_{340}) = Q_{bio(AOD=0)}^* (1 + AOD_{340} B) \quad (12)$$

264 where  $B = (0.42m + 0.93) SSA - (0.49m + 0.97)$ ,  $Q_{bio}^*$  is the BAUVR in aerosol free conditions,  $m$  is the air mass,  $SSA$  is  
 265 the single scattering albedo.

266 The Eq. (12) was obtained from the accurate model simulations for the conditions with low surface albedo ( $S=0.02$ ),  
 267 which is typical for grass. (Here and further we consider AOD at wavelength 340nm ( $AOD_{340}$ ), since this  
 268 wavelength corresponds to the standard UV channel in CIMEL sun/sky photometer, which is used in AERONET).

269 The coefficients were obtained according to model simulations for  $0 < AOD_{340} < 0.8$ , single scattering albedo

270 ( $0.8 < SSA < 1$ ), airmass  $m \sim \sinh^{-1}(m \leq 2)$ , and Angstrom exponent  $\alpha \sim 1$  ( $0.6 < \alpha < 1.5$ ). Note, that these are typical

271 changes in main aerosol properties for European conditions in UV-A spectral range (Chubarova, 2009). However,

272 the Angstrom exponent in UV-B spectral region can differ from this range and be even negative in particular

273 conditions depending on aerosol size distribution and optical properties (Bais et al., 2007). Single scattering albedo

274 in UVB spectral range according to the results of different field campaigns (UNEP, 2015) may vary from 0.7 to

275 0.97 with low SSA in the presence of black and brown carbon aerosol. Some results demonstrates no existence of

276 SSA spectral dependence in UV (Barnard et al., 2008, Arola et al., 2009) but some results shows its spectral

277 character (UNEP, 2015). We should also note that direct evaluation of SSA in UV-B spectral region is difficult

278 because UV attenuation due to aerosol can occur together with the absorption in this spectral region by different

279 gases (ozone, sulphur dioxide, formaldehyde, nitrogen dioxide, etc.). As a result we used the aerosol properties at

280 340nm as input parameters for the BAUVR simulations since we consider typical aerosol conditions without forest

281 fires and heavy industrial smoke. Radiative and airmass  $m \sim \sinh^{-1}(m \leq 2)$ , Angstrom exponent  $\alpha \sim 1$ . Since the

282 radiative effects of the existing AOD spectral dependence are relatively small within the UV-B spectral range

283 therefore we consider the same coefficients for different types of BAUVR.

284 Assuming that single scattering albedo and factor of asymmetry do not change with the altitude, we evaluated the

285 UV amplification with the altitude due to aerosol optical depth. Using Eq.(12) the equation for  $A_{AOD340}$  can be

286 written as follows:

$$287 \quad A_{AOD340} = \frac{Q_{bio}(M_0, X_0, AOD_H, A_0)}{Q_{bio}(M_0, X_0, AOD_0, A_0)} = \frac{1 + AOD_{340, H} B}{1 + AOD_{340, 0} B} \quad (13)$$

288

289 In some conditions single scattering albedo and asymmetry factor for visible wavelengths may have the altitude

290 dependence (see, for example, the results of aircraft measurements in Western Siberia (Panchenko et al., 2012)).

291 However, there is no information on the altitude dependence of aerosol properties in UV spectral region from the in-

292 situ measurements over the PEEEX area. Note, that the uncertainty of neglecting the altitude changes in single

293 scattering albedo significantly decreases at small AOD observed at high altitudes and. ~~We should also note that~~ only

294 the altitude changes in aerosol optical depth are usually taken into account in the standard tropospheric aerosol

295 models (WMO, 1986).

296 The aerosol optical depth at the altitude  $H$  ( $AOD_{340, H}$ ) can be evaluated using the following expression:

$$297 \quad AOD_{340, H} = AOD_{340, 0} f_{AOD}(H), \quad (14)$$

298 where  $f_{AOD}(H)$  is the altitude dependence of aerosol optical depth.



299 There are a lot of model aerosol profiles for the free atmosphere conditions (see, for example, widely used aerosol  
 300 models in WMO, (1986)). However, these profiles can not be applied for high-elevated locations, which are usually  
 301 characterized by a significant emission of primary aerosols or their precursors from nearby surface even in  
 302 background conditions. To account for this kind of altitude AOD dependence we used different ground-based and  
 303 satellite measurements described in the Section 2. Since our objective was to obtain the generalized aerosol altitude  
 304 dependencies we used the monthly mean  $AOD$  data from different archives over different geographical regions in  
 305 Eurasia. The dependence of aerosol optical depth as a function of altitude is shown in [Fig.4.Fig.3](#). Highly variable  
 306  $AOD_{340}$  values at different altitudes may be roughly combined in two groups, which are characterized by different  
 307 rates of aerosol altitude decrease. Hence, in our parameterization we propose to distinguish these two types of the  
 308 altitude aerosol dependence. The first one is characterized by a very strong aerosol optical depth decrease with the  
 309 altitude. It was obtained mostly from the data of European AERONET sites in the Alpine zone as well as from  
 310 several Asian sites in the sharp-peak mountainous areas. This dependence was also confirmed by the LIVAS dataset  
 311 measurements over the same areas.

312 The second one is characterized by a much more gradual altitude  $AOD_{340}$  decrease observed over flat elevated Asian  
 313 regions [according to the AERONET and LIVAS dataset and the data obtained during Moscow State University field](#)  
 314 [campaigns at the high-altitude plateau at Pamir and Tyan'-Shan' mountainous regions in Central Asia](#). The main  
 315 reason of such a character is the existence of the additional aerosol emission sources (i.e. loess, mineral aerosol)  
 316 from the vast areas of deserts and semi-deserts elevated over sea level of up to 3-4 kilometers.

317 ~~In addition, Fig.3 demonstrates the  $AOD_{340}$  dependence on altitude according to the data obtained during Moscow~~  
 318 ~~State University field campaigns at the high altitude plateau at Pamir and Tyan' Shan' mountainous regions in~~  
 319 ~~Central Asia. We can see its satisfactory agreement with the second type of  $f_{AOD}(H)$  obtained from the AERONET~~  
 320 ~~and LIVAS dataset.~~

321 The first, Alpine – like type, can be parameterized as:

$$322 \quad f_{AOD}(H)=H^{-1.65}, R^2=0.4 \quad (15)$$

323 The Alpine type aerosol altitude dependence was firstly obtained for the simulation of the UV climatology over  
 324 Europe –(COST 726, 2010). However, the coefficients have been re-affirmed [according to the monthly mean AOD](#)  
 325 [over 1999-2012 period \(case number N=137\)](#), ~~using more statistics.~~

326 The second, so-called Asian type, can be obtained using the equation according to the Moscow State University  
 327 expedition dataset in Asian region. It is characterized by much flat dependence with the altitude:

$$328 \quad f_{AOD}(H)= \exp(-0.26H), R^2=0.8 \quad (16)$$

329 The proposed dependencies can be considered as the two classes with different altitude aerosol decreasing rates.  
 330 Both dependencies are accounted for the altitudes higher than 1 km, since our analysis of AERONET dataset has  
 331 revealed the absence of the aerosol altitude dependence at the heights below 1 km due to prevailing the effects of  
 332 different aerosol sources or their precursors there. However, over the particular location the altitude  $AOD$   
 333 dependence within the first kilometre can be found, of course.

334 We should note that the proposed altitude  $AOD$  dependencies according to (15) and (16) are considered only as a  
 335 first proxy for the most sharp and flat altitude dependencies. For a particular location and specific geographical  
 336 conditions the AOD altitude dependence can be different. However, a user may easily substitute them in the  
 337 proposed parameterization.

338 Although the  $AOD$  altitude dependence is pronounced, its influence on UV amplification highly depends on initial  
 339 aerosol conditions at  $H=0$  km, the type of the altitude profile, and solar elevation (see Eq. (12)). For example, for the  
 340 Alpine type aerosol altitude profile the UV amplification at  $H=5$  km is about  $A_{AOD}=1.05-1.10$  and does not exceed  
 341  $1.11$  at  $H=8.848$  km for typical aerosol at  $H=0$  km ( $AOD_{340}=0.36$ ). However, for the polluted conditions  
 342 characterized by  $AOD_{340}=0.8$  at  $H=0$  km, the altitude UV amplification at  $H=5$  km is about  $A_{AOD}\sim 1.12-1.27$   
 343 depending mainly on solar elevation. Note, that at  $H=8.848$  km the effect is almost the same ( $A_{AOD}\sim 1.16-1.29$ ). This  
 344 will be further discussed below.

### 345 3.5. UV amplification due to changes in surface albedo with the altitude

346 The increase in surface albedo is one of the important factors, which is necessary to account for the effective  
 347 calculations of BAUVR at high altitudes. Due to significant negative temperature gradients, the snow with high  
 348 surface albedo can be observed even in summer conditions at high altitudes instead of vegetation with low UV  
 349 albedo of about  $S=0.02-0.05$  (Feister and Grewe, 1995). [Fig.5Fig4](#) demonstrates the enhancement in erythemally-  
 350 weighted irradiance due to the increase in surface albedo according to different experimental studies and the results  
 351 of one-dimensional model simulations. One can see the UV increase of around 20% at the effective surface albedo  
 352 close to  $S=0.5$  (Simic et al., 2011, Huber et al., 2004, Smolskaia et al., 2003). On the average, there is an agreement  
 353 between the calculation of UV amplification by 1D model and the measurements at different mountainous regions  
 354 up to effective surface albedo of  $S\sim 0.5$ . However, the accurate comparison of UV measurements with 3D model  
 355 (Diemoz and Mayer, 2007) shows the additional snow effect of about  $\pm 1$ UV index value due to the account of  
 356 overall interactions between radiation and different surfaces. The comparisons of UV spectral actinic flux  
 357 measurements with 1D and 3D model simulations demonstrate the similar range of uncertainties of these models,  
 358 however, 3D model gives, of course, more realistic view of the UV field in mountains since the topography and the  
 359 obstruction of the horizon are taken into account (Wagner et al., 2011). However, currently we do not consider 3D  
 360 effects in our parameterization. Due to small UV albedo over snow free surfaces this factor is negligible in summer,  
 361 while in winter the value of the effective surface albedo in mountainous area can be very high and significantly  
 362 depends on tree line location.

363 To account for surface albedo effects we followed the results obtained in different papers (Green, et al., 1980,  
 364 Chubarova, 1994), where the effects of multiple scattering were accounted using geometric progression approach.  
 365 The same approach with a detailed physical analysis was used in (Lenoble, 1998). Following these publications we  
 366 propose to calculate biologically active UV radiation in conditions with surface albedo  $S$  as follows:

$$367 \quad Q_{bio_S} = Q_{bio_{S=0}} \frac{1}{1-r_{bio}S} \quad (17)$$

368 where  $r_{bio}$  is the coefficient, which characterizes the maximum relative change in  $Q_{bio}$  due to multiple scattering for  
 369 surface albedo variations from 0 to 1. According to (Lenoble, 1998)  $r_{bio}$  is determined as the atmospheric reflectance  
 370 illuminated on its lower boundary. Note, that surface albedo  $S$  characterizes the reflecting properties at ground at the  
 371 considered altitude  $H$ .

372 The application of the equation (17)(46) to  $H=0$  km and to the altitude  $H$  allows us to obtain the following  
 373 expression for  $Q_{bio}$  at  $H$  with surface albedo  $S_H$ : using the known  $Q_{bio}$  at the altitude  $H=0$  km with surface albedo  $S_0$ :

$$374 \quad Q_{bio_{S_H}}(H) = Q_{bio_{S_0}}(H=0) \frac{Q_{bio_{S=0}}(H) \frac{1-r_{bio}(H=0)S_0}{Q_{bio_{S=0}}(H=0) \frac{1-r_{bio}(H)S_H}{1-r_{bio}(H)S_H}}}{1-r_{bio}(H)S_H} \quad (18)$$

375 This equation can be rewritten in the following way:

$$376 \quad \frac{Q_{bioS_H}(H)}{Q_{bioS_0}(H=0)} = \frac{Q_{bioS=0}(H)}{Q_{bioS=0}(H=0)} \frac{1-r_{bio}(H=0)S_0}{1-r_{bio}(H)S_H} \quad (19)$$

377 One can see that in Eq. (19) the left side of the equation  $\frac{Q_{bioS_H}(H)}{Q_{bioS_0}(H=0)}$  is the total UV amplification  $A_{UV}$  defined in  
 378 equation (4); the first term  $(\frac{Q_{bioS=0}(H)}{Q_{bioS=0}(H=0)})$  at the right side of the equation characterizes the total UV amplification in  
 379 conditions with  $S=0$ , while surface albedo effects are accounted only in the last term. Hence, we can write the UV  
 380 amplification due to the effects of surface albedo as follows:

$$381 \quad A_S = \frac{1-r_{bio}(H=0)S_0}{1-r_{bio}(H)S_H} \quad (20)$$

382 According to the model estimations the value  $r_{bio}$  in clear sky conditions has **the a relatively small** dependence on  
 383 altitude, which appears due to a decrease mainly in molecular and aerosol loading and can be **easily** parameterized  
 384 by a simple regression as follows:

$$385 \quad r_{bio}(H) = b H + c, \quad (21)$$

386 where the coefficients  $b$  and  $c$  are given in Table 2 for different types of BAUVR. They were obtained for a variety  
 387 of solar elevation and ozone content taking into account for the altitude changes in molecular scattering as well as  
 388 for altitude dependence of aerosol optical depth  $f_{AOD}(H)$ . The  $r_{bio}(H)$  mainly depends on molecular content and  
 389 aerosol properties, and slightly decreases with the altitude due to reducing in multiple scattering effects with the  
 390 decrease in molecular and aerosol loading.

391 As a result, the UV amplification due to the increase in surface albedo at the altitude  $H$  strongly depends on  
 392 scattering processes and also decreases with the altitude. **Fig.6 Fig-5** shows the maximum  $A_S$  effect due to the  
 393 changes in  $S$  from  $S=0$  at zero level to  $S=1$  at the altitude  $H$  for the different types of BAUVR. The  $A_S$  decreases  
 394 with the altitude from more than 1.6 at  $H=0$  km to about 1.2 at  $H=8.848$  km due to the decrease in  $r_{bio}$ .

### 395 3.6.Validation

396 Using the generalized parameterizations for different geophysical parameters obtained in the previous sections we  
 397 can estimate the total UV amplification  $A_{UV}$  with the altitude from Eq.(4). The results of the validation of the  
 398 proposed method with these altitude parameter dependencies against the accurate RT simulations are shown in  
 399 **Fig.7, Fig-6**. One can see a **high close** correlation ( $r>0.996$ ) between the  $A_{UV}$  values obtained by the proposed method  
 400 and the accurate RT simulations using the 8-stream DISORT method within the changes in altitude from  $H=0$  km to  
 401 8 km, in solar elevation from 20 to 50°, in surface albedo from  $S=0$  to  $S=0.9$ , in ozone from 250 DU to 350 DU at  
 402  $H=0$  km, **and in AOD<sub>340</sub> from 0.2 to 0.4 at  $H=0$  km, and in SSA varying from 0.88 to 0.96 km.** Different altitude  
 403 aerosol profiles were also considered. Validation was made for all three types of BAUVR. Overall, the average bias  
 404 is **0±2% 0±0.2%** for erythemally-weighted irradiance, **0±1% and 1±0.2%** - for **cataract-weighted, and 1±1% - for**  
 405 **vitamin D irradiance other types of BAUVR.** The maximum difference between the  $A_{UV}$  calculated by the proposed  
 406 method and by the accurate model simulations does not exceed 6% at the highest elevation ( $H=8$  km) at low ozone  
 407 content.

408 The comparisons of the total UV amplification according to the proposed method with the total  $A_{UV}$  obtained from  
 409 the experimental dataset as a function of altitude are shown in **Fig.8, Fig-7**. The experimental data were taken from  
 410 the dataset of Moscow State University mountainous field campaigns, which was described in the Section 2. After

411 accounting for the molecular, aerosol, and ozone altitude dependence the simulated UV amplification is in  
412 satisfactory agreement with the obtained experimental results.

#### 413 4. Discussion

414 The total altitude amplification of biologically active UV irradiance  $A_{UV}$  as a function of altitude is shown in  
415 [Fig. 9](#) [Fig. 8](#) for a variety of atmospheric conditions at surface albedo  $S=0$  and  $S=0.9$  and high solar elevation  $h=50^\circ$ .  
416 One can see a distinct altitude difference obtained for different types of BAUVR with larger increase for vitamin D-  
417 weighted irradiance due to its higher sensitivity to ozone content. The difference in  $A_{UV}$  for various BAUVR can  
418 reach 15-20% at high altitudes. The effects of surface albedo on  $A_{UV}$  can be seen if compare the results shown in [Fig.](#)  
419 [9a-7a](#) and [Fig. 9b](#) [Fig. 8b](#). One can see the 2-2.5 times UV increase due to high surface albedo at high altitudes, which  
420 is again more pronounced for vitamin D-weighted radiation with smaller effective wavelength and, hence, more  
421 effective multiple scattering than that for the other types of BAUVR. Larger  $A_{UV}$  values are also observed in  
422 conditions with smaller ozone amount for all three types of BAUVR for both zero and high surface albedo  
423 conditions. High surface albedo  $S=0.9$  provides a significant increase even at zero level, which is similar to the  $A_{UV}$   
424 increase due to altitude change of 6 km. It is clearly seen that typical aerosol and ozone does not play a vital role in  
425  $A_{UV}$ . However, for all types of BAUVR the increase of slightly absorbing aerosol (from  $AOD_{340}=0.2$  to  $AOD_{340}=0.4$ )  
426 provides a noticeable  $A_{UV}$  growth in conditions with relatively small ozone content due to enhancement of multiple  
427 scattering (see [Fig. 9](#) [Fig. 8](#)).

428 The  $A_{UV}$  values are smaller at solar elevation  $h < 20^\circ$  for all types of BAUVR mainly due to decreasing in  $G_{bio\_M}$  (see  
429 the coefficients in Table 1). For example, at  $H=8$  km the UV altitude amplification for vitamin D-weighted radiation  
430 is about  $A_{UV}=1.77$  at  $h=10^\circ$  compared with  $A_{UV}=2.0$  at  $h=50^\circ$  at  $X=250$  DU and  $AOD_{340}=0.4$ .

431 Let us consider the conditions, which are characterized by the most pronounced UV amplification with the altitude -  
432 the conditions with high aerosol loading  $AOD_{340}=0.8$ , low ozone content  $X=250$  DU at  $H=0$  km, and high solar  
433 elevation  $h=50^\circ$ . In addition, we consider the abrupt increase in surface albedo at  $H=2$  km from  $S=0$  to  $S=0.95$ ,  
434 which can be possible due to location of tree line there and pure snow above. The altitude UV amplification due to  
435 these input parameters according to the proposed  $A_{UV}$  parameterization is shown in [Fig. 10](#) [Fig. 9](#). The separate effects  
436 of different factors can be seen in [Fig. 10a](#) [Fig. 9a](#) and their total effects on different BAUVR types are shown in  
437 [Fig. 10b](#) [Fig. 9b](#). One can see a different role of geophysical factors at different altitudes: the prevalence of molecular  
438 scattering especially at high altitudes while the extremely high surface albedo may play the most important role at  
439 the altitudes of its abrupt increase (in our case -  $H=2$  km,  $A_S=1.48$ ). However, in our example the UV amplification  
440 due to surface albedo decreases at the altitude higher than 2 km because of the reduction in multiple scattering. The  
441 UV altitude amplification due to aerosol is more distinct and reaches 1.1-1.2 at high altitudes if there is a strong  
442 aerosol pollution  $AOD_{340}=0.8$  at  $H=0$  km. It is more pronounced for the Alpine-type AOD altitude dependence and  
443 in our example at  $H=2$  km it can be even higher than the  $A_M$  value (see [Fig. 10](#) [Fig. 9](#)). The effects of ozone in UV  
444 amplification do not exceed 1.1-1.2 at high altitudes depending on BAUVR type. We would like to emphasize that  
445 [Fig. 10](#) [Fig. 9](#) is only the illustration of the application of the proposed  $A_{UV}$  parameterization for a given parameters  
446 altitude variations.

447 We implemented the proposed UV altitude parameterization in the developed on-line UV tool <http://momsu.ru/uv/>,  
448 which had been developed for the simulation of erythemally-weighted irradiance and the UV resources over  
449 Northern Eurasia (the PEEEX domain) at  $H=0$  km (Chubarova and Zhdanova, 2013). Using this program it is

450 possible to calculate UV irradiance and UV-resources for different atmospheric conditions at a given geographic  
451 location and specified time. Based on the threshold for vitamin D and erythemally active irradiance the UV  
452 resources are obtained for various skin types and open body fraction. According to the classification we consider  
453 different categories of UV-deficiency, UV-optimum and UV-excess (Chubarova and Zhdanova, 2013). The  
454 interactive on-line UV tool represents a client-server application where the client part of the program is the web-  
455 page with a special form for the input parameters required for erythemally-weighted UV irradiance simulations. The  
456 server part of the program consists of the web-server and the CGI-script, where the different input parameters are set  
457 by a user or taken from the climatological data available at the same site. In addition, in this part of the program  
458 erythemally-weighted irradiance is calculated, visualized and classified according to the proposed UV resources  
459 categories. The proposed UV irradiance altitude parameterization has been also incorporated in the calculation  
460 scheme with additional account for the changes in the atmospheric parameters with the height. This enables a user to  
461 evaluate UV irradiance at any requested elevation above sea level taking into account for a variety of the altitude  
462 dependent parameters.

463 Let us analyze the UV resources for skin type 2 and the open body fraction of 0.25 in the Alpine region  
464 (approximately 46°N and 7°E) for winter and spring noon conditions. For these conditions the vitamin D threshold  
465 is equal to 100 J/m<sup>2</sup> and Minimal Erythemal Dose -250 J/m<sup>2</sup>. According to our estimates on On January, 15<sup>th</sup>, at  
466 H=0 km for typical (climatological) ozone, aerosol and surface albedo conditions the noon UV deficiency (no  
467 vitamin D generation) is observed ~~conditions (with noon erythemally UV dose of about 97.2 Jm<sup>-2</sup> Jm<sup>-2</sup>) are~~  
468 ~~observed at H=0 km for typical (climatological) ozone, aerosol and surface albedo conditions;~~ while at the same  
469 ~~coordinates~~ location the UV optimum is observed at H higher 0.5 km up to H=4.807 km (the highest point within the  
470 Alps, peak Mont Blanc) we obtain the UV optimum conditions with noon erythemally UV dose varying ~~variation~~  
471 from 100.6 Jm<sup>-2</sup> to 122.9 Jm<sup>-2</sup>.

472 However, for skin type 4 ~~according to the skin type classification~~ (Fitspatrick, 1988) with vitamin D threshold of  
473 180 J/m<sup>2</sup> the noon UV deficiency is observed at all altitudes and even at high surface albedo S=0.9 corresponding to  
474 the pure snow with UV dose of 154.4 Jm<sup>-2</sup>. The decrease in ~~If we increase the~~ open body fraction for this skin type  
475 4 from 0.25 to 0.05, which could take place in frosty weather, provides 100% UV deficiency, when no ~~0.5 the~~  
476 vitamin D can be generated during the whole day at the H=4.8 km ~~generation is possible and S=0.9, hence, the~~  
477 ~~conditions are classified as UV optimum.~~

478 On April, 15<sup>th</sup>, at H=0 km ~~for open body fraction 0.25~~ and typical climatological conditions at this geographical  
479 point noon UV dose is about 437.7 Jm<sup>-2</sup>. This means that for the open body fraction of 0.4 the moderate UV excess  
480 is observed for skin type 2, and the UV optimum ~~is~~ for skin type 4 when vitamin D threshold is 112.5 Jm<sup>-2</sup> and  
481 MED threshold - 450 Jm<sup>-2</sup> ~~at H=0 km with noon UV dose of about 437.7 Jm<sup>-2</sup>.~~ At the altitude H = 2 km the  
482 conditions are characterized by the moderate UV excess for both skin types 2 and 2 ~~and, still, by the UV optimum~~  
483 ~~for skin type 4~~ with UV dose of 463.4 Jm<sup>-2</sup>. At the H=4 km a high UV excess is observed for skin type 2 and the  
484 moderate UV excess - for skin type 4 with UV dose of 532.4 Jm<sup>-2</sup>.

485 Thus, the proposed altitude UV parameterization can be effectively used for accurate estimating the BAUVR at  
486 different altitudes with any altitude resolution for a variety of geophysical parameters over the PEEEX domain in  
487 Northern Eurasia. The accurate RT methods like Monte-Carlo, Discrete Ordinate method or others, of course, can be  
488 used instead for UV irradiance simulations, however, their application is very time-consuming and are not possible  
489 in some applications. The proposed approach is especially very useful for the application in different kinds of on-  
490 line UV tools, where it is not possible to use a lot of prescribed calculations for a wide set of different geophysical  
491 parameters or accurate UV modelling.



492 | The current state of the online interactive tool does not take into account for the skin orientation relative to the Sun  
493 | and the geometry of the human body which can modify the results limited to UV irradiance simulations on  
494 | horizontal surface (Hess and Koepke, 2008, Vernez et al., 2014). In this case UV dosimeters which have a spectral  
495 | response almost identical to that of the UV-induced photobiological effect (Siani et al., 2008) is the most accurate  
496 | way for evaluating the individual levels of UV exposure. The vitamin D production can be also affected by other  
497 | factors such as obesity and age (Engelsen, 2010). However, these are the tasks for the future work.

498 | The combination of different altitude dependencies for main geophysical factors in the proposed parameterization  
499 | allows a user to make a reliable altitude UV assessment. ~~It is not, of course, possible to take into account for the~~  
500 | ~~specific altitude dependencies.~~ We should also emphasize that the proposed ozone and aerosol altitude dependencies  
501 | in the troposphere were taken from the experimental data obtained in background conditions and, hence, should be  
502 | applied only for these conditions. However, they can be easily substituted by any other altitude dependence of  
503 | considering factors. ~~using the proposed parameterization.~~

504 | With the application of the proposed method we can also reveal the effects of different geophysical factors on  
505 | various types of BAUVR and estimate their comparative role in the altitude UV effects. And, of course, the  
506 | parameterization ~~can be also used in downscaling the UV results for the regions located at high altitudes~~ from the  
507 | coarse resolution global chemistry-climate- models for the regions located at high altitudes. The proposed method  
508 | can be applied not only over the PEEEX domain but on a global scale over the world. However, more attention should  
509 | be paid in this case to the evaluation of the particular altitude dependence of the different parameters.

## 510 | **5. Conclusions**

511 | The objective of this paper was to develop a flexible parameterization based on rigorous model simulations with  
512 | account for generalized altitude dependencies of molecular density, ozone, and aerosols considering different  
513 | surface albedo conditions. We show that for the specified groups of parameters we can present the altitude UV  
514 | amplification ( $A_{UV}$ ) for different BAUVR as a composite of independent contributions of UV amplification from  
515 | different factors with the mean uncertainty of 1% and standard deviation of 3%. The parameterization takes into  
516 | account for the altitude dependence of molecular number density, ozone content, aerosol loading, and spatial surface  
517 | albedo. We also provide the generalized altitude dependencies of different parameters for evaluating the  $A_{UV}$ . Their  
518 | validation against the accurate RT model (8 stream DISORT RT code) for different types of BAUVR shows a good  
519 | agreement with maximum uncertainty of few percents and correlation coefficient  $r > 0.996$ . It was not possible, of  
520 | course, to cover all the observed variety in the parameters. However, due to the proposed approach the parameter  
521 | altitude dependencies can be easily substituted by a user.

522 | Using this parameterization one can estimate the role of different atmospheric factors in the altitude UV variation.  
523 | The decrease in molecular number density, especially at high altitudes, and the increase in surface albedo play the  
524 | most significant role in  $A_{UV}$  growth. At high solar elevations the UV amplification due to aerosol at  $H=8.848$  km  
525 | does not exceed 1.3 even when  $AOD_{340}=0.8$  at  $H=0$  km. The UV amplification due to aerosol calculated with the  
526 | Alpine-type AOD altitude aerosol dependence is much more pronounced than that calculated with the Asian-type  
527 | AOD altitude dependence, especially at relatively lower altitudes ( $H=2-4$  km). The UV amplification due to ozone  
528 | does not exceed 1.20 and is higher at smaller solar elevations, especially, for vitamin-D-weighted irradiance.

529 | This parameterization was applied to the on-line tool for calculating the UV resources (<http://momsu.ru/uv/>) over the  
530 | PEEEX domain. Using this tool one can easily evaluate the UV conditions (UV deficiency, UV optimum or UV  
531 | excess) at different altitudes for a given skin type and open body fraction. As an example, we analyzed the altitude

532 UV increase and its possible effects on health considering different skin types and various open body fraction for  
533 January and April conditions in the Alpine region. We showed that even in clear sky conditions over the same  
534 geographical point (46°N, 7°E) in mid-January the UV optimum can be observed at altitudes higher than H=0.5 km  
535 for skin type 2, while the noon UV deficiency are still remained at the altitudes up to H=4.8 km for skin type 4 when  
536 open body fraction is 0.25. In mid-April the account for the altitude dependence at 4 km provides the changes from  
537 the UV optimum to UV excess for people with 4 skin type, and from moderate UV excess to a the high UV excess  
538 for skin type 2 and from UV optimum the moderate UV excess conditions — for people with 2 skin type 4 when  
539 open body fraction is 0.4.

540 This approach can be also used in downscaling the results of global chemistry-climate models with the coarse spatial  
541 resolution in mountainous domain and as a simple tool for different types of applications for personal purposes of  
542 users.

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709

710 **FIGURE CAPTIONS**

711 [Fig. 1.](#) Action spectra for erythema (CIE, 1998), vitamin D (CIE, 2006) and for eye damage (cataract) (Oriowo et  
712 [al. 2001](#)) effects.

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714 [Fig.1, Fig.2.](#) The comparison of  $A_{UV}$  amplification factor calculated from Eq.(4) as multiplication of  $A_M A_X A_{AOD} A_S$   
715 with the direct model simulation of UV amplification. All the parameters ( $A_{UV}, A_M A_X A_{AOD} A_S$ ) were obtained from  
716 accurate model simulations.

717 Comment. The simulations were performed for different altitudes ( $H=0$  and  $H=5km$ ), aerosol optical depth ( $AOD_{340}= 0, 0.2,$   
718  $0.4$ ), [single scattering albedo \( \$SSA=0.88, 0.96\$ \), Angstrom exponent \( \$\alpha=0.6, 1.0, 1.5\$ \), total ozone \( \$X=250, 300, 350\$  DU\), surface  
719 albedo \( \$S=0, S=0.9\$ \) and solar elevations \( \$h=20^\circ\$  and  \$50^\circ\$ \). For estimating the UV amplification we assume at  \$H=0\$  km the  
720 conditions with 350DU,  \$AOD\_{340}=0.4\$ ,  \$S=0\%\$  and normalized the BAUVR at the altitude  \$H=5km\$  to the value calculated with  
721 these parameters.](#)

722 [Fig.2.](#)

723 [Fig.3.](#) UV amplification due to decrease in molecular number density with the altitude  $H$  according to accurate  
724 model simulations: TUV, 8- stream DISORT TUV model.  $h=50^\circ$ .  $X=300$  DU.

725 ~~[Fig.3.](#) The altitude dependence of aerosol optical depth at 340nm with 1 sigma error bar according to the AERONET,  
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727 AOD at 330 nm the Moscow State University dataset and the AOD at 355nm from the LIVAS datasets were  
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733 [Fig.5.](#) UV amplification due to the surface albedo increase in mountainous areas according to different experimental  
734 data and model simulations. The error bars of model simulation relates to the different input parameters – altitude of  
735 2 and 3 km, solar elevation of 10, 30 and  $50^\circ$ , total ozone  $X=350DU$ ,  $AOD_{340}=0.17$  at  $P=95\%$ .

736 ~~[Fig. 5.](#) The dependence of  $r_{bio}$  with the altitude for different BAUVR from accurate model simulations for a variety  
737 of geophysical parameters (left axis) and maximum  $A_S$  effects due to changes in surface albedo from  $S=0$  at  $H=0$   
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746 [Fig.7.](#) The comparison between the total altitude UV amplification according to the proposed method and the  $A_{UV}$   
747 values evaluated using the accurate RT model (TUV, 8-stream DISORT method). [See the details in the text.](#)

748 | [Fig.7](#)[Fig.8](#). The comparison between the simulated UV amplification according to the proposed parameterization  
749 | and the UV amplification from the experimental data as a function of altitude. Moscow State University dataset.  
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751 | altitude gradient differed from the typical value. The total ozone was equal to  $X \sim 300$  DU at  $H=0$  km,  $X \sim 240$  DU at  
752 |  $H > 3$  km and  $X \sim 250$  DU at  $H \sim 1-2$  km.

753 | [Fig.8](#)[Fig.9](#). Total UV amplification as a function of the altitude for different types of BAUVR in a variety of  
754 | atmospheric conditions with  $S=0$  (a) and  $S=0.9$  (b). The model parameters at  $H=0$  km:  $X=250-350$  DU,  $AOD_{340}=0.2-$   
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757 | [Fig.9](#)[Fig.10](#). The UV amplification due to molecular  $A_M(Qery)$ ,  $A_M(QvitD)$ , ozone  $A_X(QvitD)$ ,  $A_X(Qery)$ , aerosol  
758 |  $A_{AOD,f1(H)}$ ,  $A_{AOD,f2(H)}$  for the Alpine  $f1(H)$  and Asian  $f2(H)$  types of altitude dependences, and surface albedo  $A_S(Qery)$ ,  
759 |  $A_S(QvitD)$  changes with the altitude (a) and their total altitude effect on  $A_{UV}$  for different types of BAUVR (b). At  
760 |  $H=0$  km:  $AOD_{340}=0.8$ ,  $X=250$  DU. The surface albedo has an abrupt change at 2 km from  $S=0$  to  $S=0.95$ . Solar  
761 | elevation -  $h=50^\circ$ .

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763 **LIST OF THE TABLES:**

764

765 Table 1. Relative molecular gradients  $G_{bio\_M(A=0)}$ (%/km) at different solar elevations and ozone content for different  
 766 types of BAUVR. Accurate model simulations. Zero surface albedo conditions. Determination coefficient  $R^2$  is  
 767 higher than 0.997 in all cases. The standard error of  $G_{bio\_M(A=0)}$  is given in the brackets at P=95% .  
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h, solar elevation, degrees	erythemally-weighted irradiance	cataract-weighted irradiance	vitamin D-weighted irradiance	erythemally-weighted irradiance	cataract-weighted irradiance	vitamin D-weighted irradiance
	X=300 DU			X=500 DU		
10	4.5 (0.04)	3.8 (0.04)	1.8 (0.01)	4.8 (0.05)	3.9 (0.04)	0.8 (0.03)
20	6.4 (0.04)	6.9 (0.05)	7.1 (0.06)	6.0 (0.03)	6.6 (0.04)	6.8 (0.05)
30	6.7 (0.01)	7.2 (0.02)	7.8 (0.02)	6.1 (0.01)	7.0 (0.01)	7.8 (0.02)
40	6.4 (0.02)	6.8 (0.01)	7.3 (0.01)	5.8 (0.02)	6.6 (0.02)	7.4 (0.01)
50	6.0 (0.03)	6.2 (0.03)	6.7 (0.03)	5.5 (0.03)	6.1 (0.03)	6.8 (0.03)
60	5.7 (0.04)	5.8 (0.04)	6.2 (0.04)	5.3 (0.04)	5.7 (0.04)	6.4 (0.04)

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771 Table 2. The coefficients for calculating the  $r_{bio}$  values (Eq. 21) for different types of BAUVR. Model estimations.

772 The standard error of the coefficients is given in the brackets at P=95%.

	erythemally-weighted irradiance	cataract-weighted irradiance	vitamin D-weighted irradiance
$b$	-0.025(0.0002)	-0.025(0.0002)	-0.025(0.0002)
$c$	0.394(0.0009)	0.394(0.0009)	0.405(0.0008)
$R^2$	>0.99	>0.99	>0.99

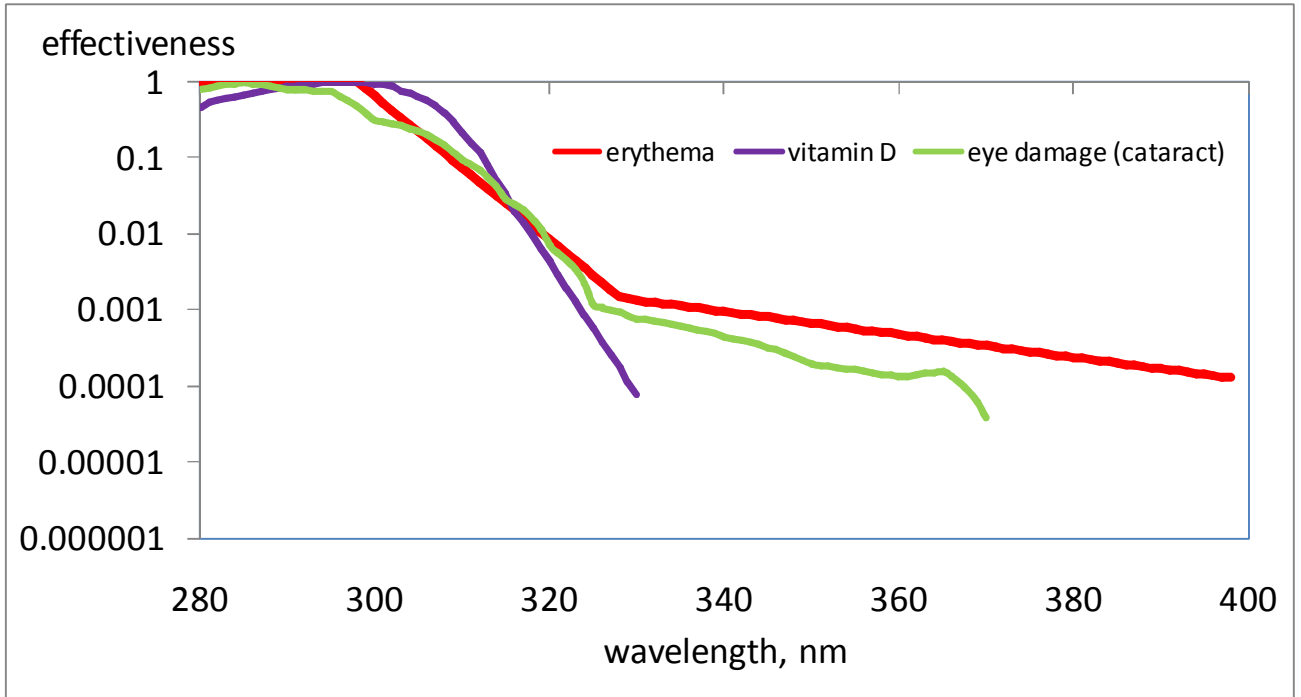
773 Note: the simulations were fulfilled for different combinations of input parameters at  $H=0$ :

774  $X=250-350$  DU,  $AOD_{340}=0.2-0.4$ ,  $S=0-0.9$ ,  $h=20-50^\circ$ .

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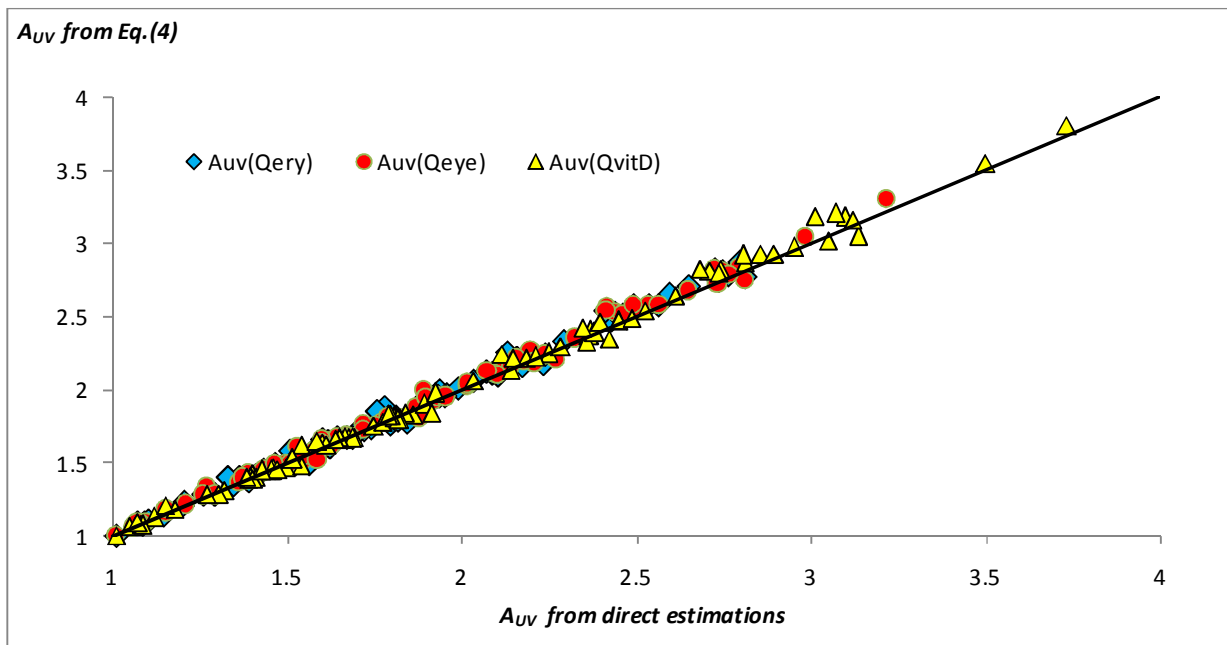


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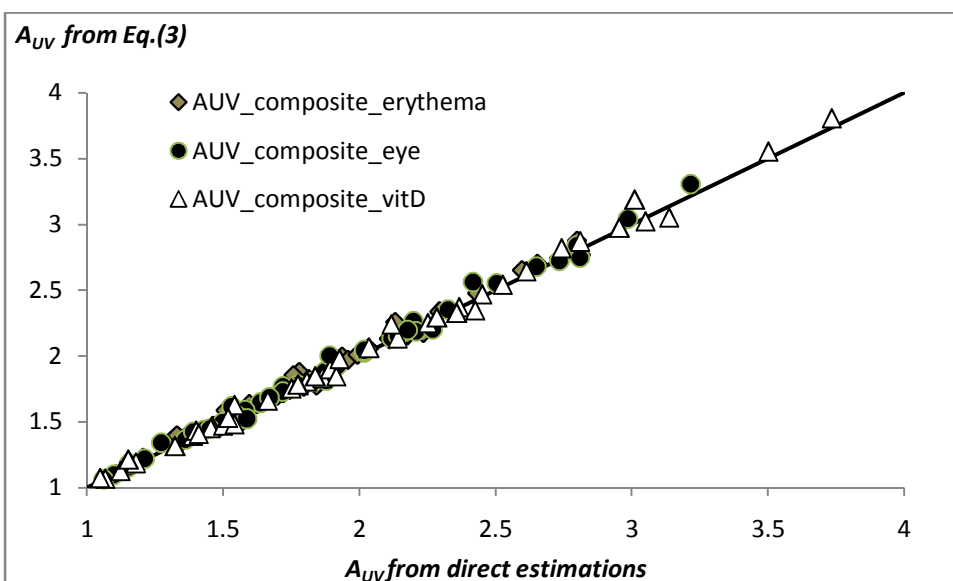
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Figure 2.



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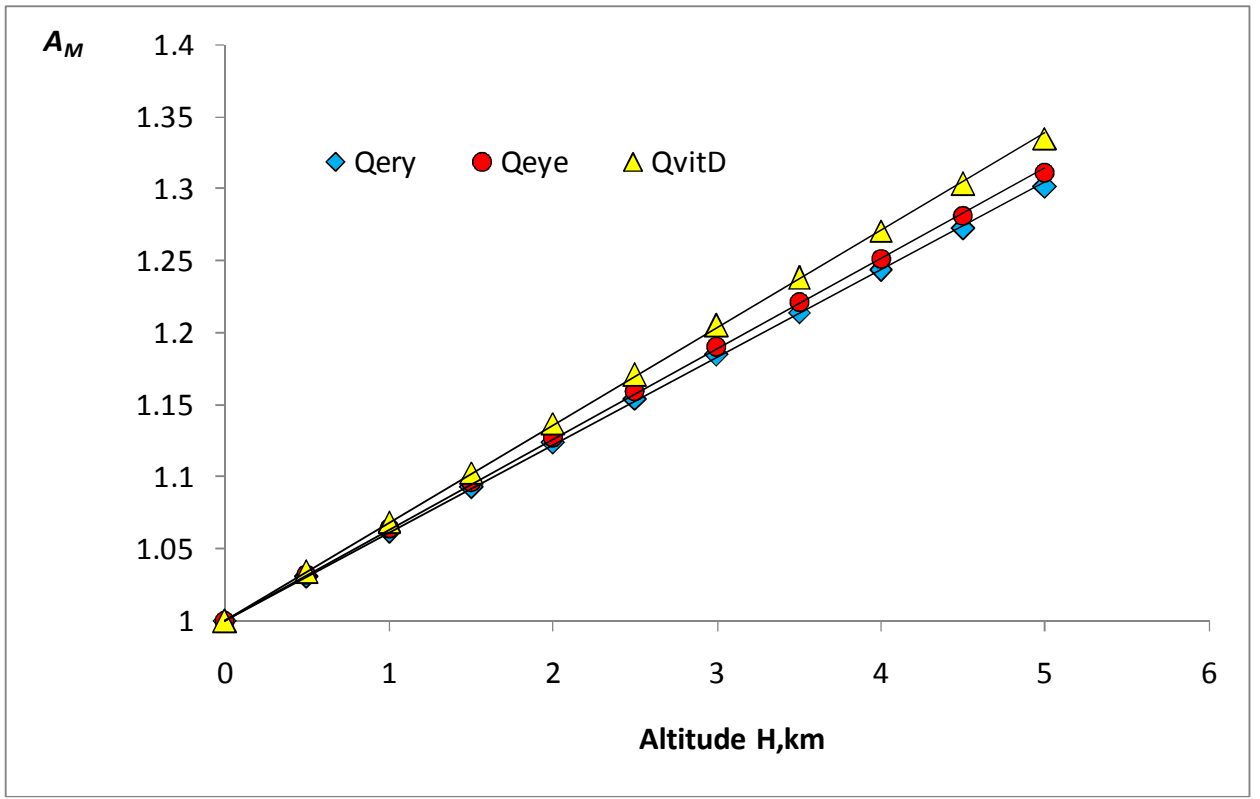
791 **Figure 1.** The comparison of  $A_{UV}$  amplification factor calculated from Eq.(4) as multiplication of  $A_M A_X A_{AOD} A_S$  with the  
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 795  $0.4$ ), **single scattering albedo ( $SSA=0.88, 0.96$ ), Angstrom exponent ( $\alpha=0.6, 1.0, 1.5$ ), total ozone ( $X=250,300,350$  DU), surface  
 796 albedo ( $S=0, S=0.9$ ) and solar elevation\* ( $h=20^\circ$  and  $50^\circ$ ). For estimating the UV amplification we assume at  $H=0$  km the**



797 | *conditions with 350DU, AOD340=0.4, S=0% and normalized the BAUVR at the altitude H=5km to the value calculated with*  
798 | *these parameters.*  
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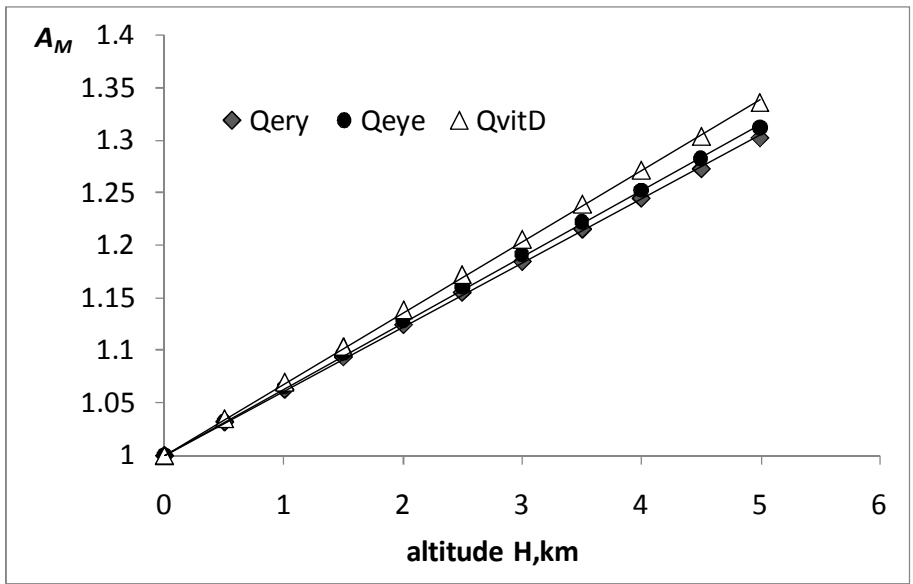
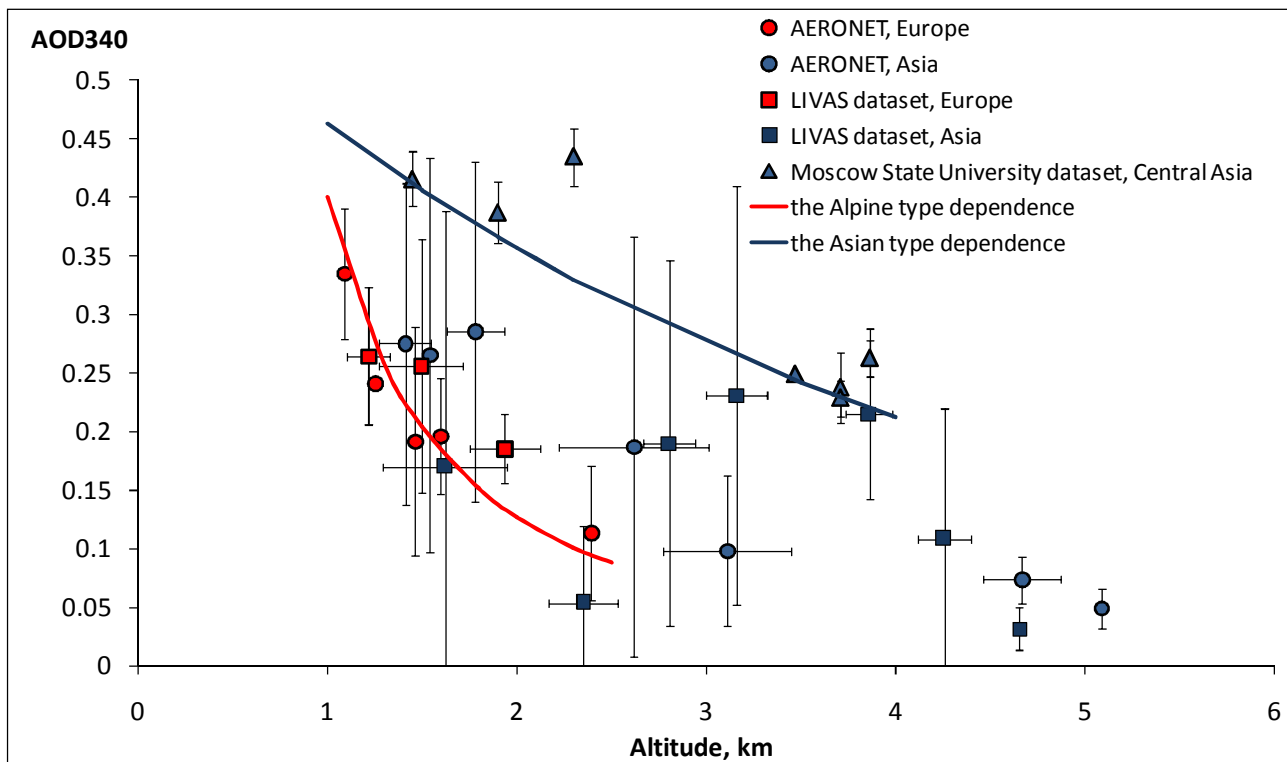
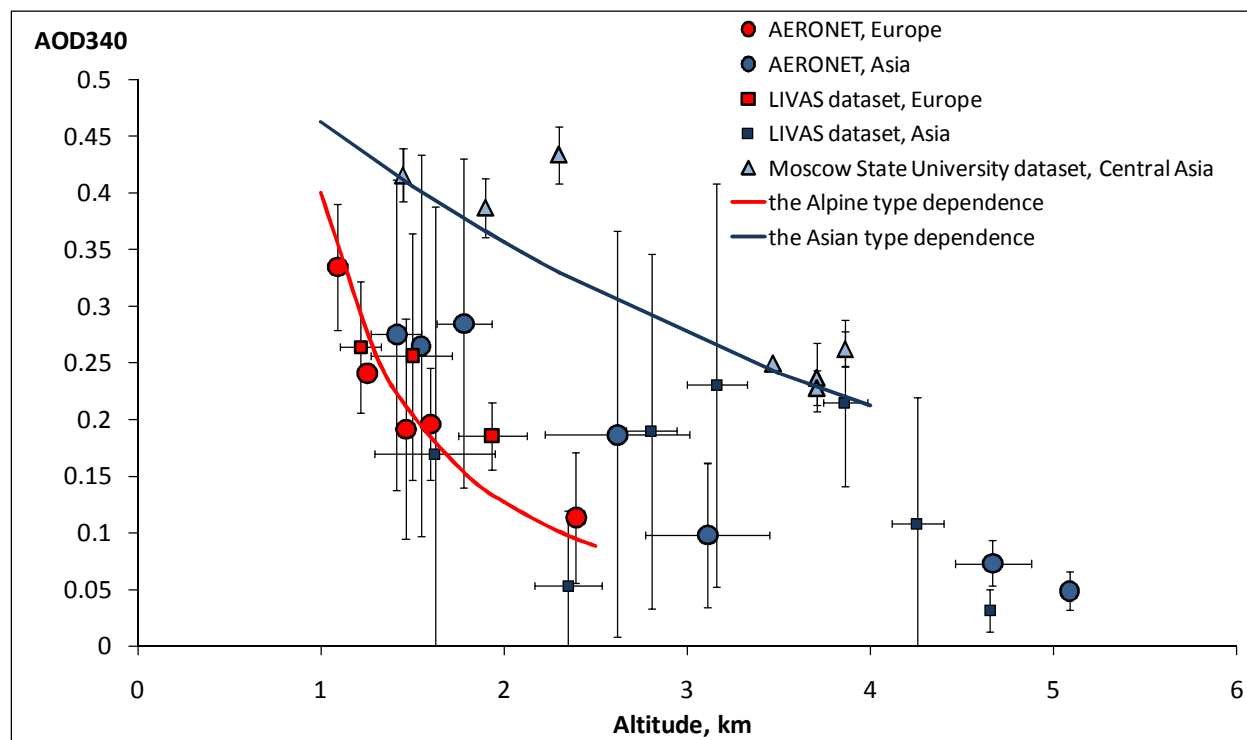


Figure 3.2. UV amplification due to decrease in molecular number density with the altitude  $H$  according to accurate model simulations: 8- stream DISORT TUV model.  $h=50^\circ$ .  $X=300$  DU.  $R^2>0.997$  – for all regression lines.



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835 ~~Figure 3. The altitude dependence of aerosol optical depth at 340nm with 1 sigma error bar according to the~~  
 836 ~~AERONET, LIVAS and the Moscow State University datasets over European and Asian regions. May-September~~  
 837 ~~period. The AOD at 330 nm the Moscow State University dataset and the AOD at 355nm from the LIVAS datasets~~  
 838 ~~were recalculated to AOD at  $\lambda=340$  nm using the Angstrom~~



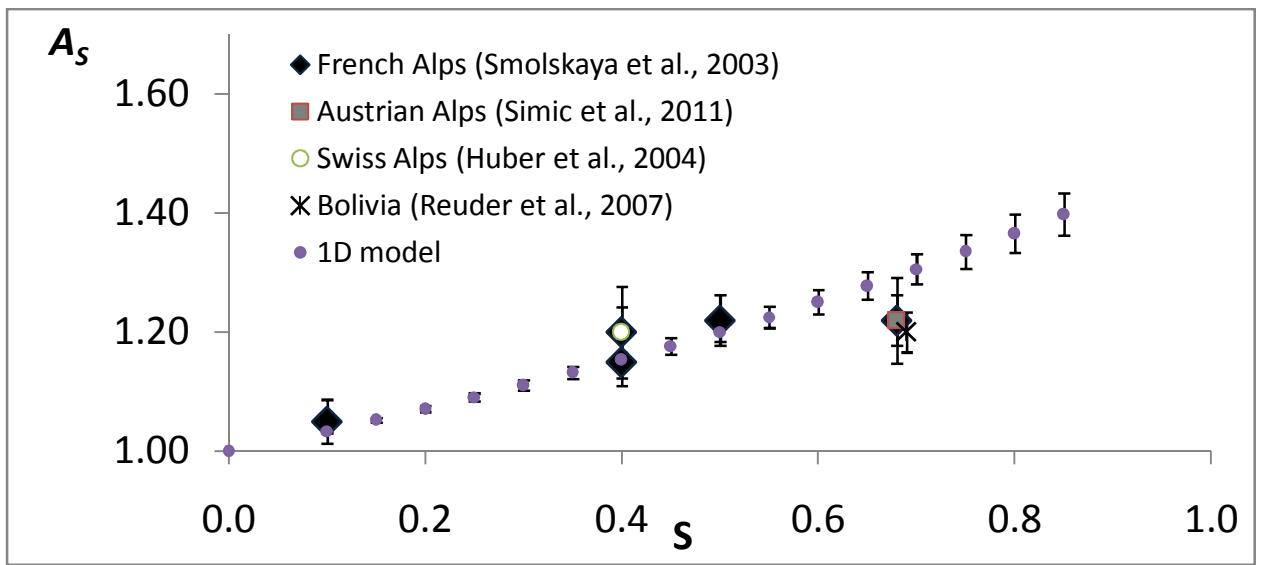
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840 ~~Figure 4. parameter  $\alpha=1.3$ . See further details in the text.~~  
 841 ~~The altitude dependence of aerosol optical depth at 340nm with 1 sigma error bar according to the AERONET, LIVAS~~  
 842 ~~and the Moscow State University datasets over European and Asian regions. May-September period. The AOD at 330 nm~~

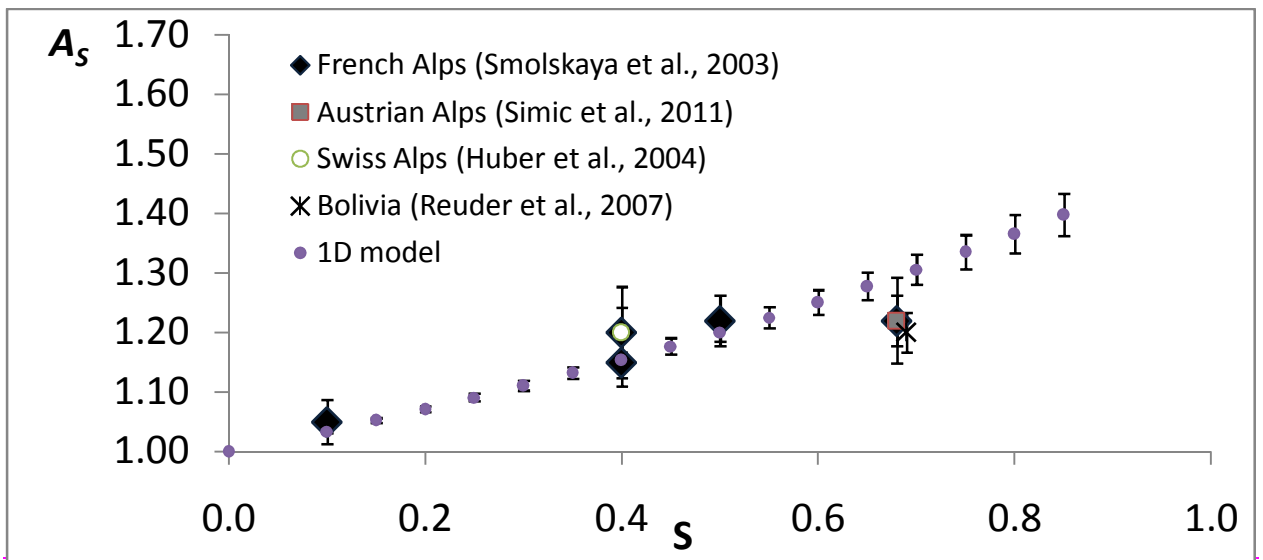
843 | [the Moscow State University dataset and the AOD at 355nm from the LIVAS datasets were recalculated to AOD at](#)  
844 |  [\$\lambda=340\$  nm using the Angstrom exponent  \$\alpha=1.0\$ . See further details in the text.](#)  
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Figure 5.4 UV amplification due to the surface albedo increase in mountainous areas according to different experimental data and model simulations. The error bars of model simulation relates to the different input parameters – altitude of 2 and 3 km, solar elevation of 10, 30 and 50°, total ozone  $X=350\text{DU}$ ,  $\text{AOD}_{340}=0.17$  at  $P=95\%$ .

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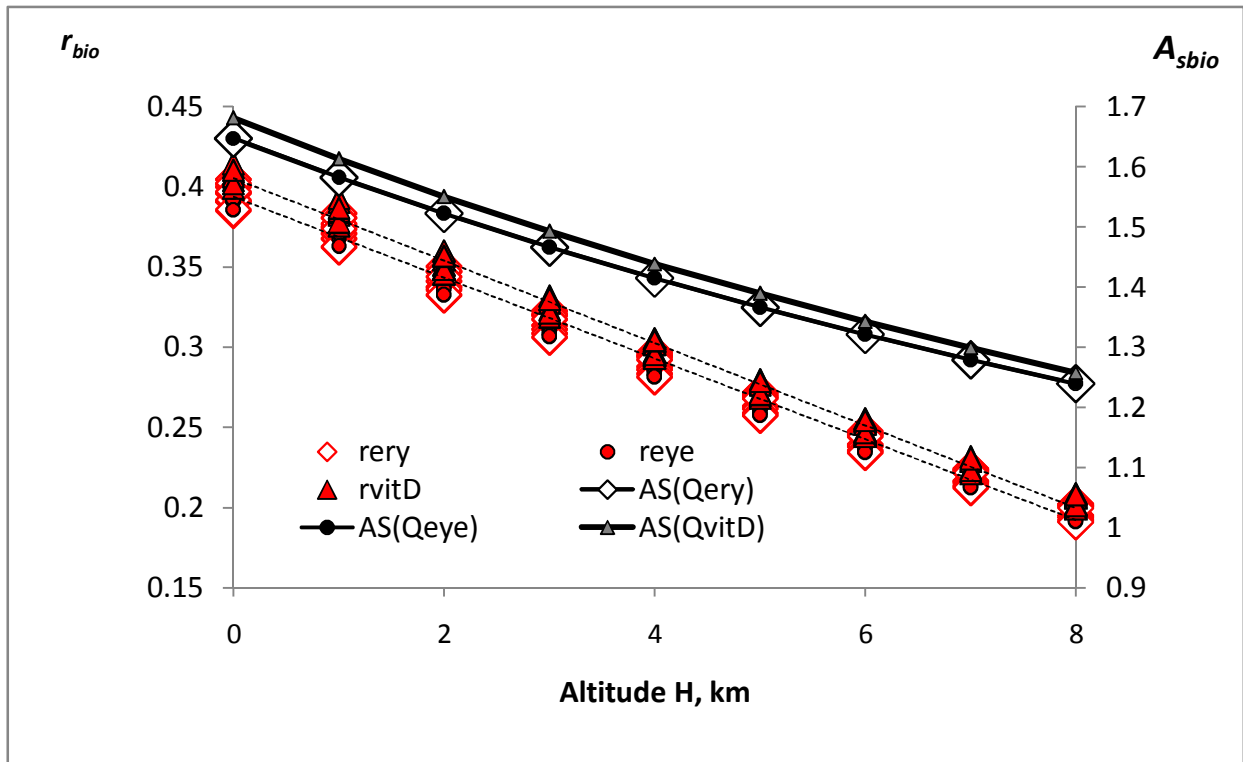
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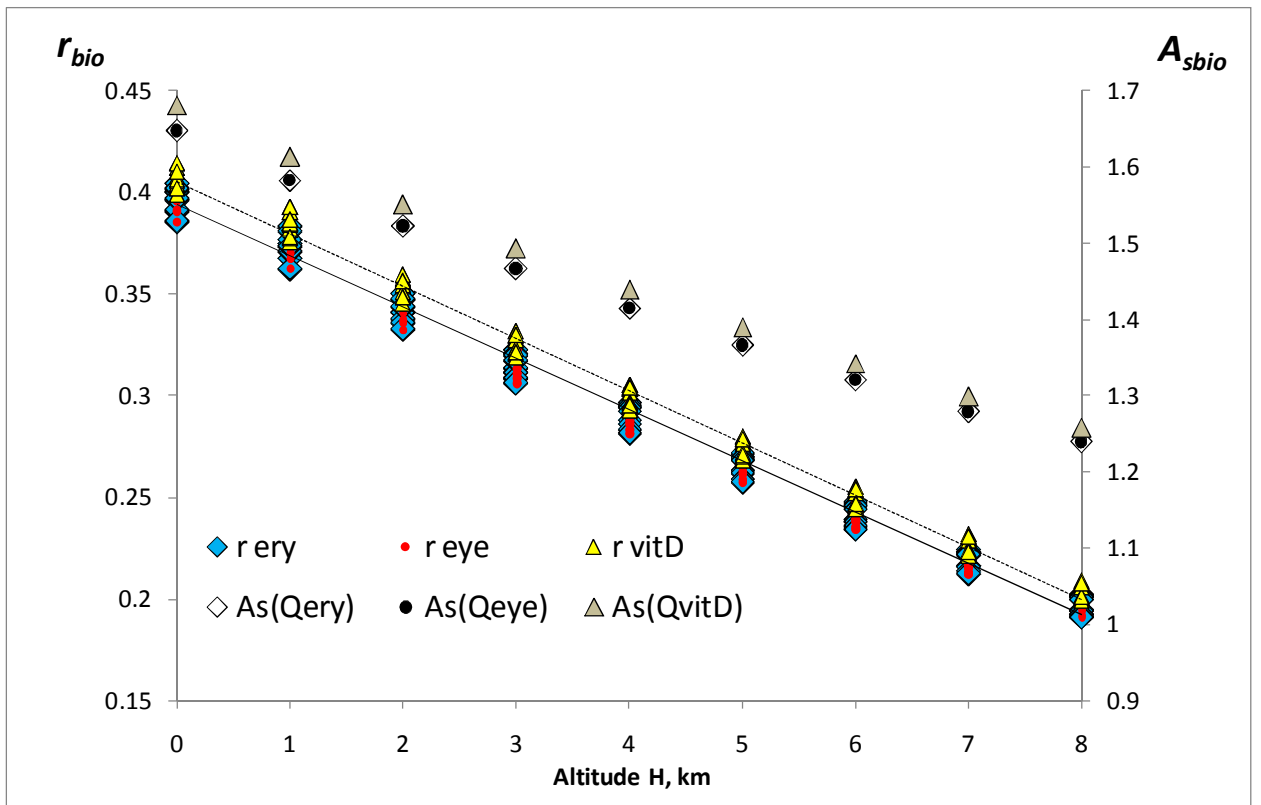
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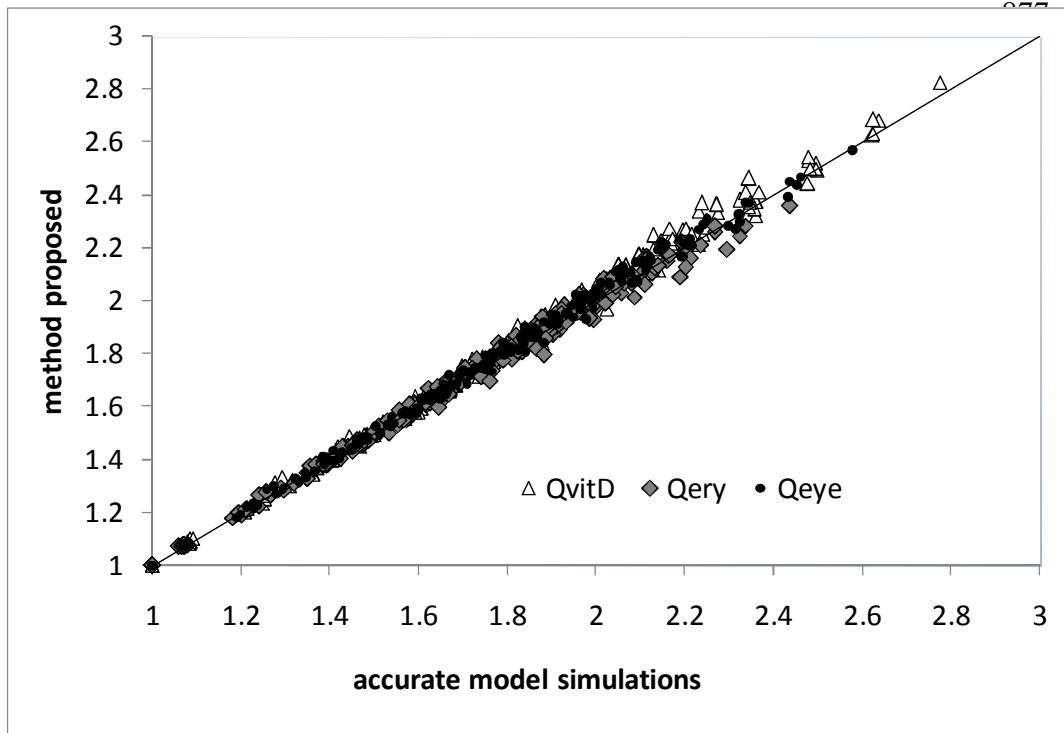
869 **Figure 5.** The dependence of  $r_{bio}$  with the altitude for different BAUVR from accurate model simulations for a  
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 872 for  $r_{Qeye}(H)$  is the same as for  $r_{Qery}(H)$ . The coefficients of the regression equations and the ranges of the input  
 873 parameters at  $H=0$  are given in Table 2.



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The dependence of  $r_{bio}$  with the altitude for different BAUVR from

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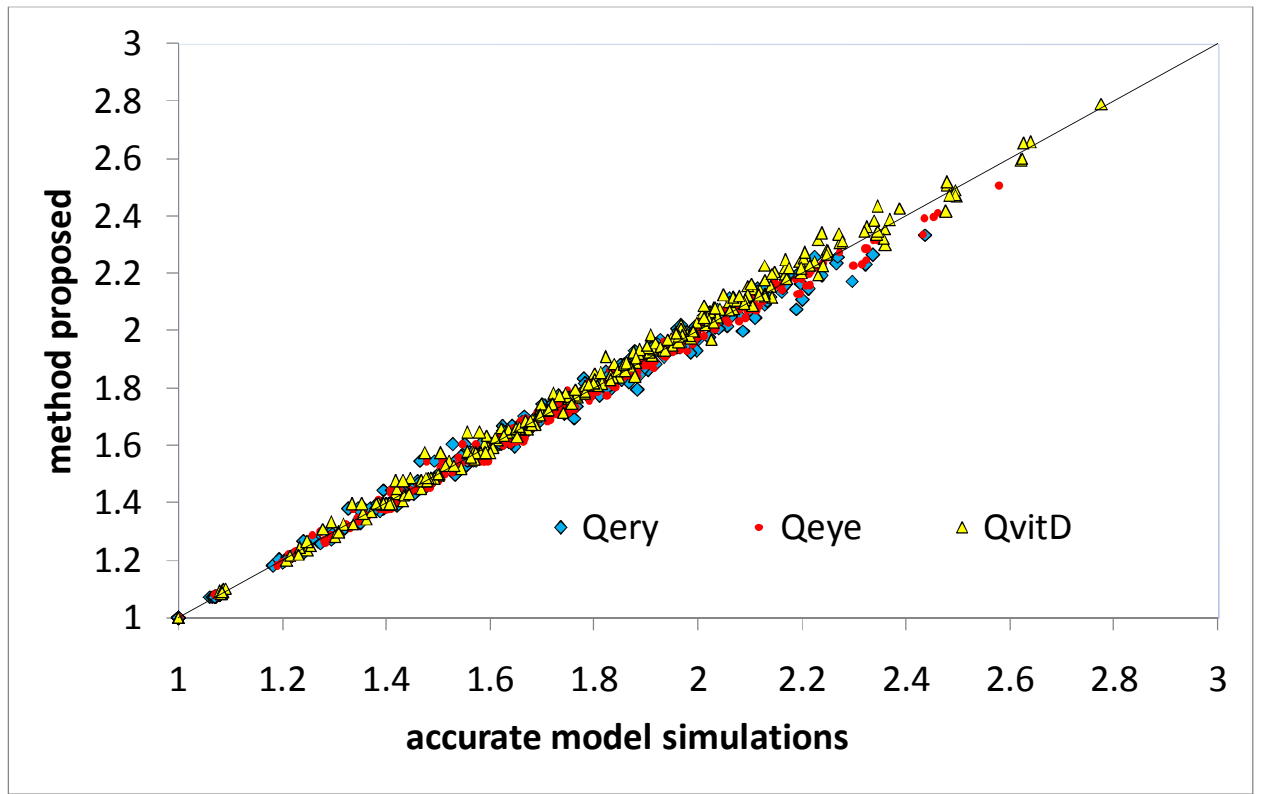
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accurate model simulations for a variety of geophysical parameters (left axis) and maximum  $A_s$  effects due to changes in surface albedo from  $S=0$  at  $H=0$  km to  $S=1$  at level  $H$  (right axis). The  $r_{bio}$  regressions are shown in dashed line. Note, that the regression line for  $r_{Qery}(H)$  is the same as for  $r_{QvitD}(H)$ . The coefficients of the regression equations and the ranges of the input parameters at  $H=0$  are given in Table 2.



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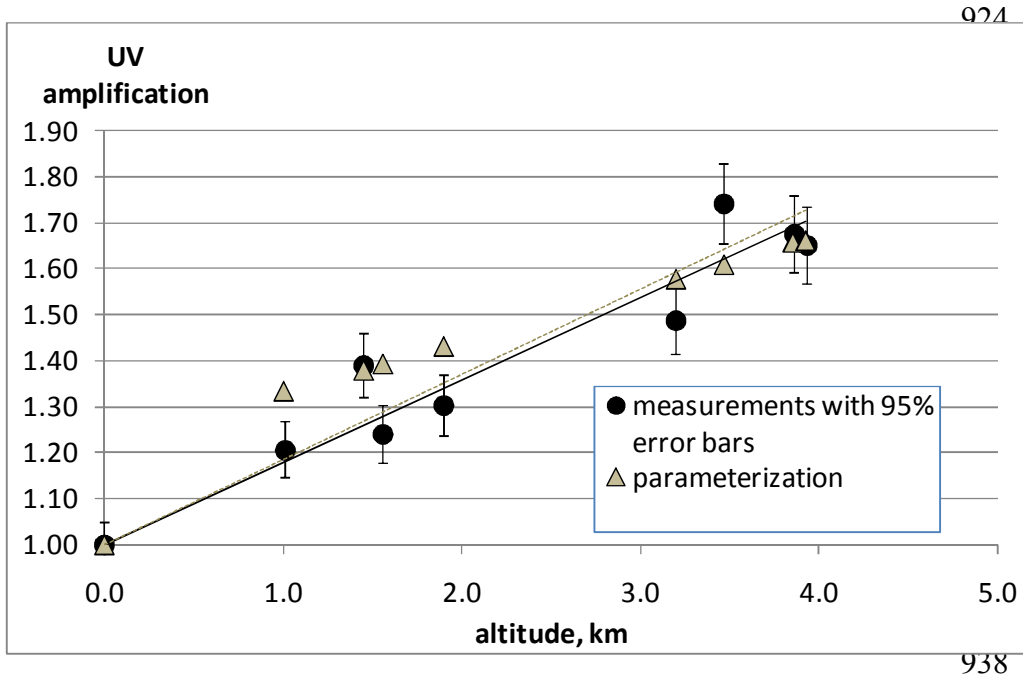
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Figure 7.6: The comparison between the total altitude UV amplification ( $A_{UV}$ ) according to the proposed method and the  $A_{UV}$  values evaluated using the accurate RT model (TUV, 8-stream DISORT method). [See the details in the text.](#)

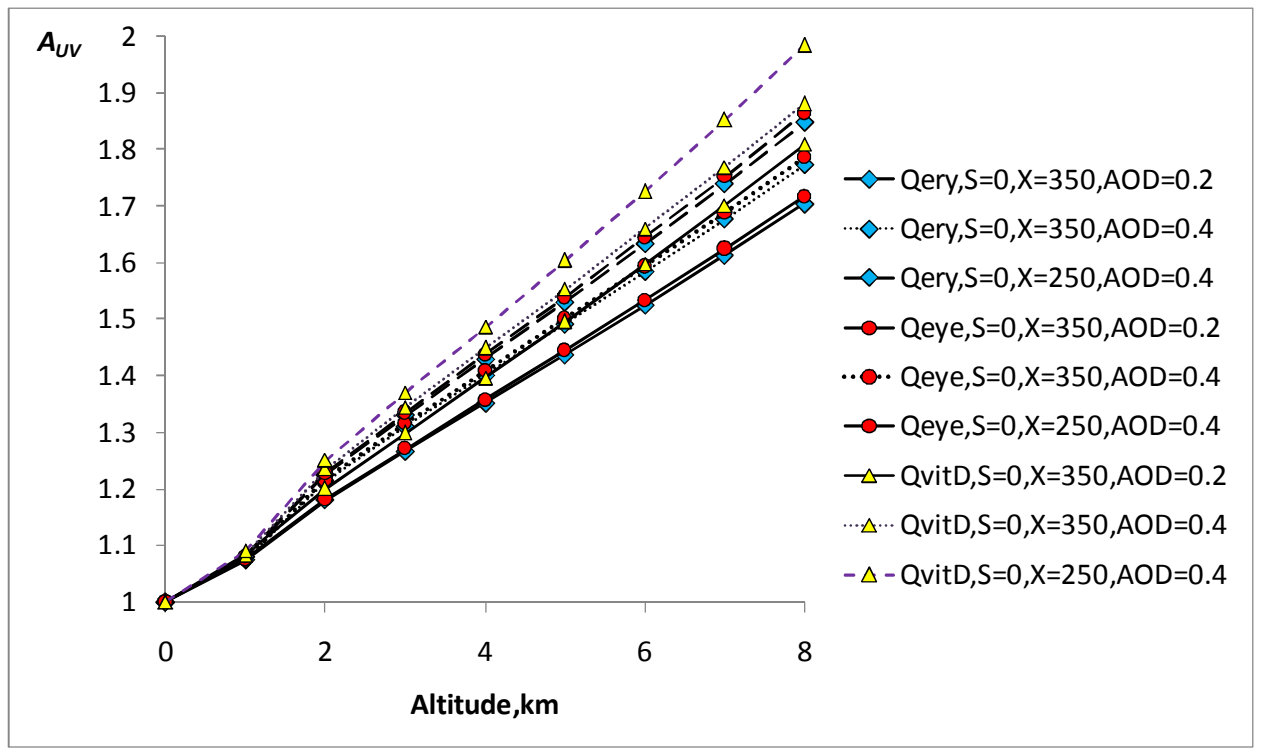
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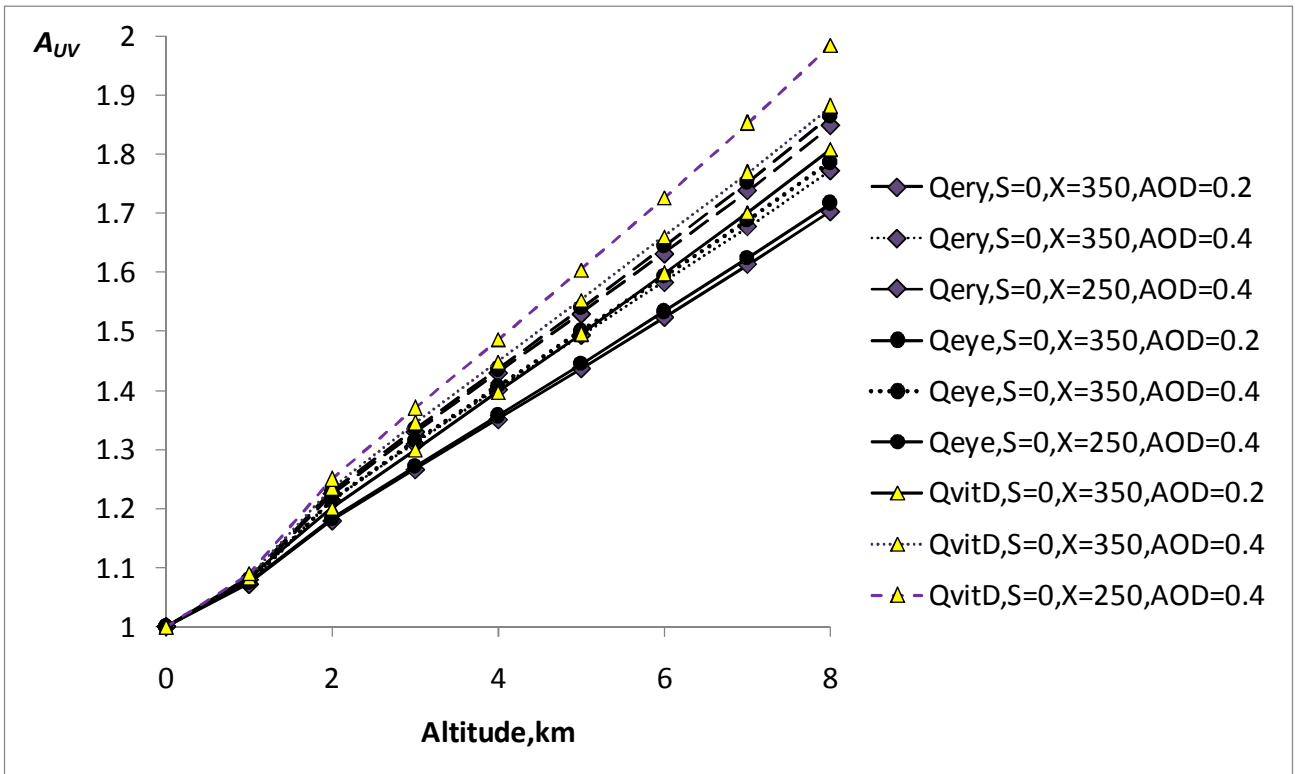
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942 **Figure 8.7.** The comparison between the simulated UV amplification according to the proposed parameterization  
943 and the UV amplification from the experimental data as a function of altitude. Moscow State University dataset.  
944 Solar elevation  $h=50^\circ$ . Clear sky conditions. Note: since we used the data of different field campaigns the ozone  
945 altitude gradient differed from the typical value. The total ozone was equal to  $X \sim 300$  DU at  $H=0$ km,  $X \sim 240$  DU at  
946  $H>3$  km and  $X \sim 250$  DU at  $H \sim 1-2$  km.

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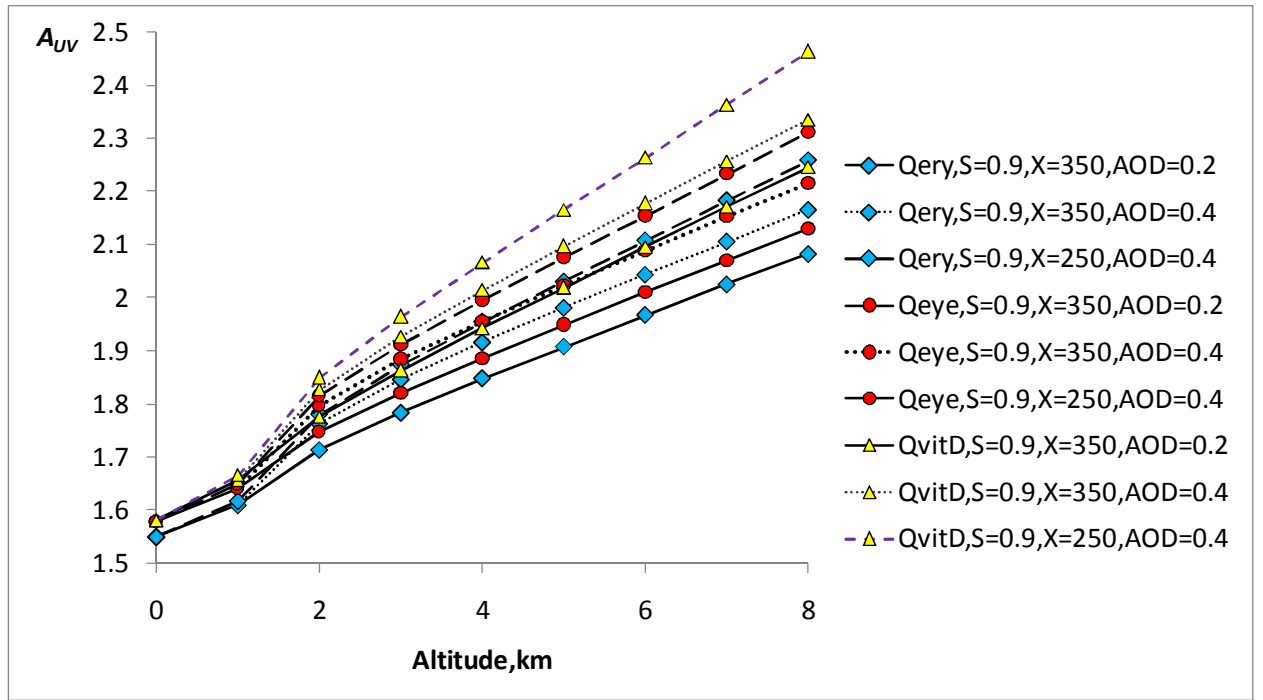


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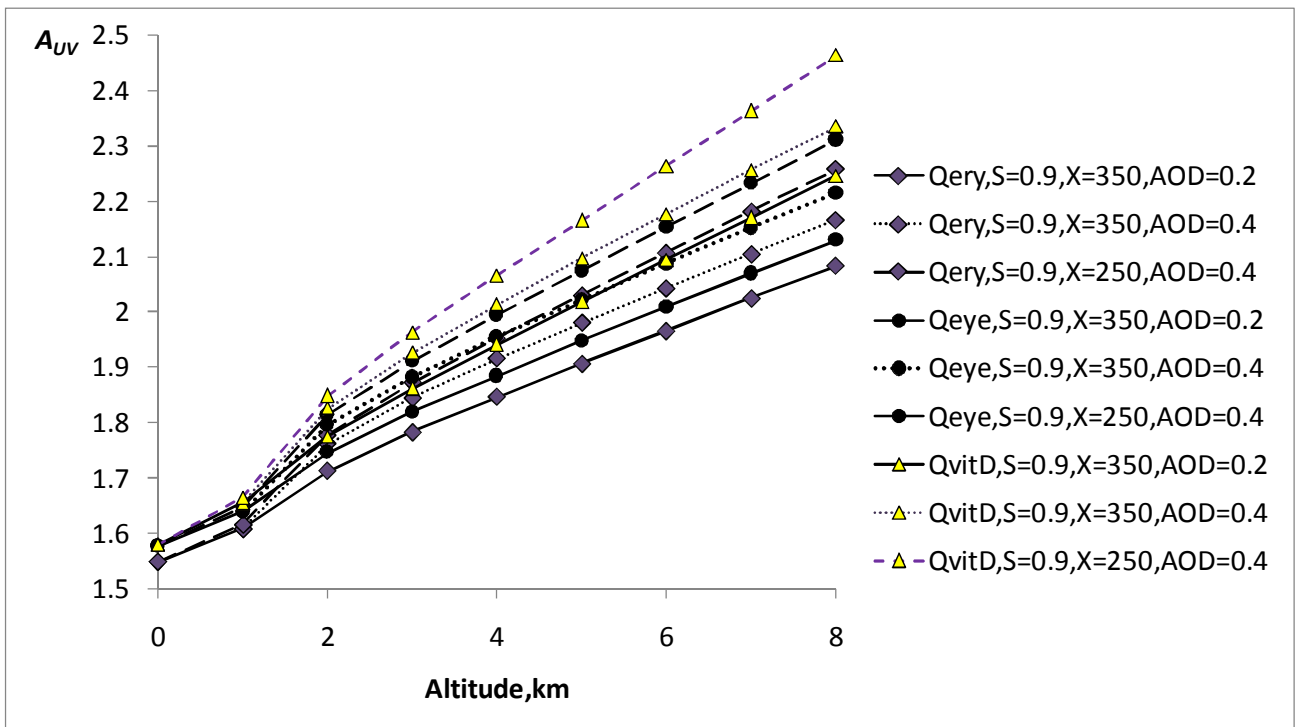
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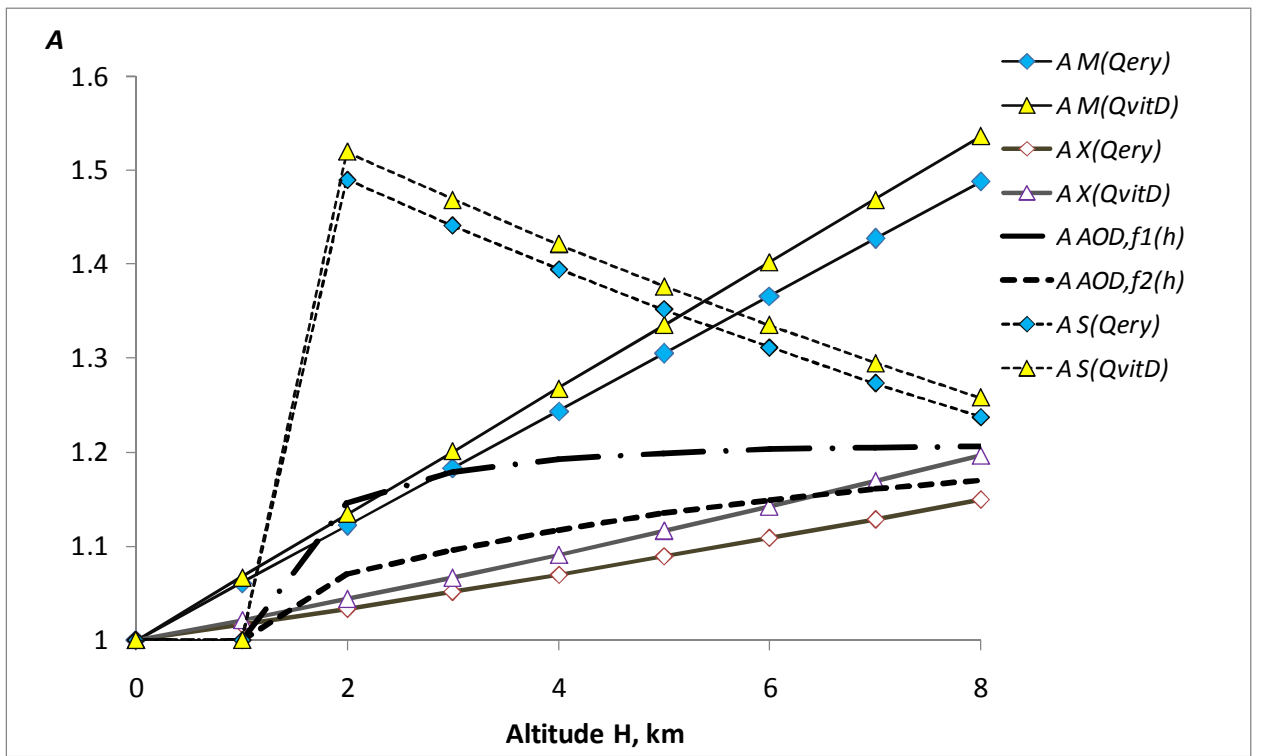
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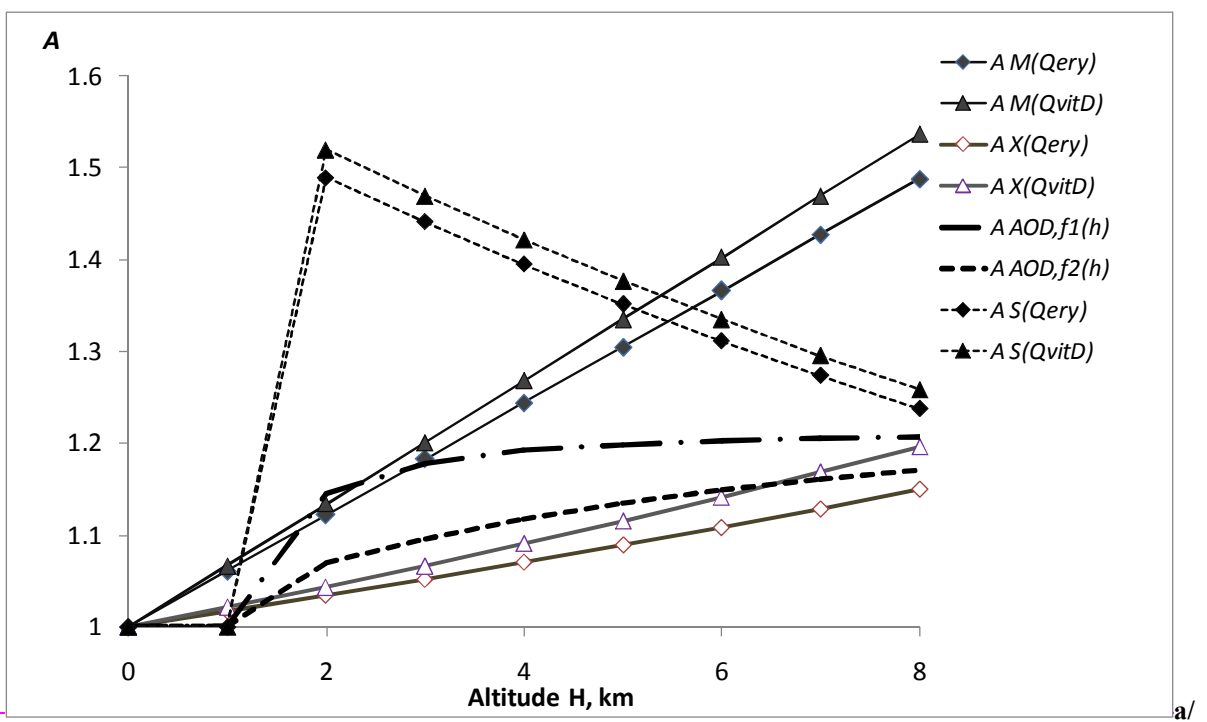
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**Figure 9.8.** Total UV amplification as a function of the altitude for different types of BAUVR in a variety of atmospheric conditions with  $S=0$  (a) and  $S=0.9$  (b). The model parameters at  $H=0$  km:  $X=250-350$  DU,  $AOD_{340}=0.2-0.4$ . The Alpine type of AOD altitude dependence according to the Eq. (15) was taken into account. Solar elevation- $h=50^\circ$ .

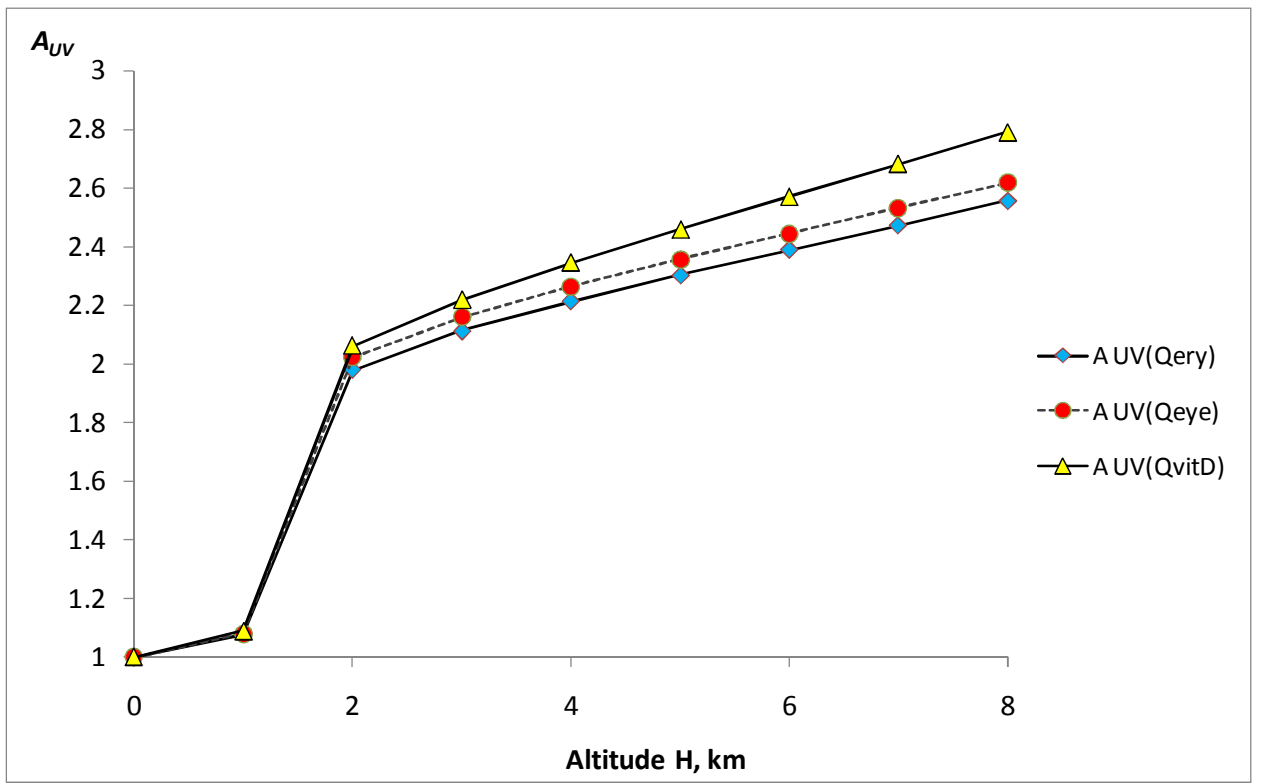




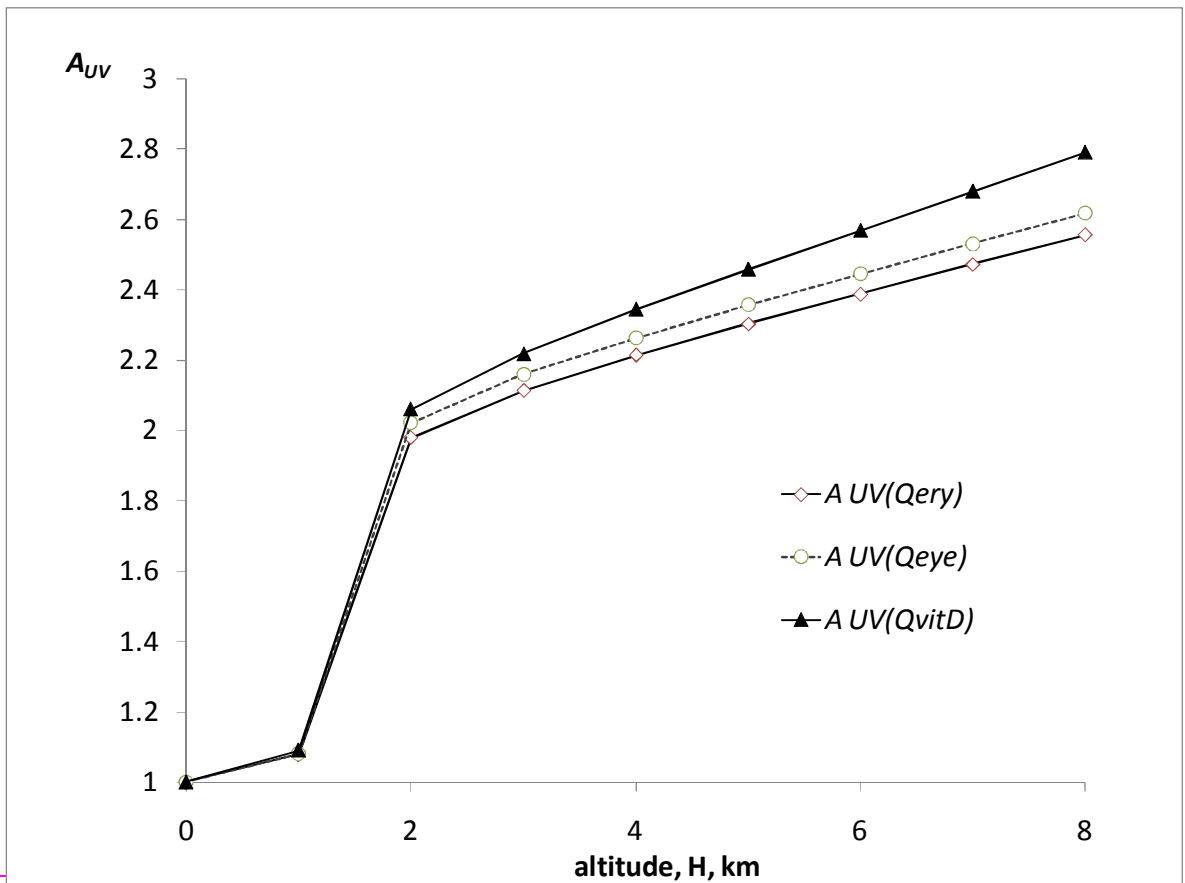
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Figure 10.9- The UV amplification due to molecular  $A_M(Qery)$ ,  $A_M(QvitD)$ , ozone  $A_X(QvitD)$ ,  $A_X(Qery)$ , aerosol  $A_{AOD,f1(H)}$ ,  $A_{AOD,f2(H)}$  for the Alpine  $f1(H)$  and Asian  $f2(H)$  types of altitude dependences, and surface albedo  $A_S(Qery)$ ,  $A_S(QvitD)$

970 changes with the altitude (a) and their total altitude effect on  $A_{UV}$  for different types of BAUVR (b). At  $H=0$  km:  
971  $AOD_{340}=0.8$ ,  $X=250$  DU. The surface albedo has an abrupt change at 2 km from  $S=0$  to  $S=0.95$ . Solar elevation -  $h=50^\circ$ .

# A new parameterization of the UV irradiance altitude dependence for clear-sky conditions and its application in the on-line UV tool over Northern Eurasia

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**Abstract.** A new method for calculating the altitude UV dependence is proposed for different types of biologically active UV radiation (erythemally-weighted, vitamin-D-weighted and cataract-weighted types). We show that for the specified groups of parameters the altitude UV amplification ( $A_{UV}$ ) can be presented as a composite of independent contributions of UV amplification from different factors within a wide range of their changes with mean uncertainty of 1% and standard deviation of 3% compared with the exact model simulations with the same input parameters. The parameterization takes into account for the altitude dependence of molecular number density, ozone content, aerosol and spatial surface albedo. We also provide generalized altitude dependencies of the parameters for evaluating the  $A_{UV}$ . The resulting comparison of the altitude UV effects using the proposed method shows a good agreement with the accurate 8-stream DISORT model simulations with correlation coefficient  $r > 0.996$ . A satisfactory agreement was also obtained with the experimental UV data in mountain regions. Using this parameterization we analyzed the role of different geophysical parameters in UV variations with altitude. The decrease in molecular number density, especially at high altitudes, and the increase in surface albedo play the most significant role in the UV growth. Typical aerosol and ozone altitude UV effects do not exceed 10-20%. Using the proposed parameterization implemented in the on-line UV tool (<http://momsu.ru/uv/>) for Northern Eurasia over the PEEEX domain we analyzed the altitude UV increase and its possible effects on human health considering different skin types and various open body fraction for January and April conditions in the Alpine region.

**Keywords:** UV radiation, altitude dependence, RT modelling, erythemally-weighted irradiance, vitamin D-weighted irradiance, cataract-weighted irradiance, interactive UV-tool.

## 1. Introduction

Biologically active UV radiation (BAUVR) is an important environmental factor, which significantly affect human health and nature (UNEP, 1998; UNEP, 2011). Enhanced levels of UV radiation lead to different types of skin cancer (basal and squamous cell carcinomas, cutaneous melanoma), to various eye diseases (cataract, photokeratitis, squamous cell carcinoma, ocular melanoma, variety of corneal/conjunctival effects), and to immunosuppression. However, small doses of UV radiation have a positive effect on health through the vitamin D generation (UNEP, 2011).

UV radiation is affected by astronomical factors (solar zenith angle, solar-earth distance, solar activity), by different atmospheric characteristics (total ozone content, cloudiness, aerosol, optically-effective gases), and by surface albedo (Madronich, 1993, Bais et al., 2007, Bekki et al., 2011). However, the altitude above sea level has also a

36 significant influence on UV radiation (Bais et al., 2007). There are a lot of studies and special field campaigns in  
37 different geographical regions, which were devoted to the analysis of the altitude UV effect (Bernhard et al., 2008,  
38 Blumthaler and Ambach, 1988, Blumthaler et al., 1994, Blumthaler et al., 1997, Dahlback et al., 2007, Lenoble et  
39 al., 2004, Piacentini, et al. 2003, Pfeifer et al., 2006, Sola et al., 2008, etc.). The UV enhancement at high altitudes is  
40 detected not only due to smaller molecular scattering, but also due to usually observed decreasing in total ozone  
41 content and aerosol, and increasing in surface albedo, which in turn enhances 3D reflection from slopes of  
42 mountains covered by snow (Lenoble et al., 2004). In addition, variation of cloud properties with altitude can also  
43 change the level of UV radiation.

44 The UV records in mountainous areas demonstrate extremely high levels. The highest UV values are observed in  
45 Andes in Bolivia (Pfeifer et al., 2006, Zaratti et al., 2003), where **UV index**<sup>1</sup> can be sometimes close to 20. Very  
46 high UV levels were also recorded at high-altitude deserts in Argentina (Piacentini, et al. 2003). In Tibet the UV  
47 index frequently exceeded 15 on clear days and occasionally exceeded 20 on partially cloudy days (Dahlback et al.,  
48 2007). At the European alpine stations in summer conditions the UV indices are often higher than 11 (Hülse,  
49 2012). **For example, high UV index up to 12 was observed in mountainous areas in Italy (Casale et al. 2015). A**  
50 **significant UV growth with the altitude was also obtained at different sites in Austria and Switzerland (Rieder et al.,**  
51 **2010). In winter, erythemally-weighted UV irradiance is about 60% higher than that at lower-altitude European sites**  
52 **(Gröbner et al., 2010).** In the Arctic the comparison of summer UV measurements at Summit (3202m a.s.l.) and  
53 Barrow (~0 a.s.l.) stations **shows** significant enhancing of about 30-40% in clear sky conditions at the elevated site  
54 (Bernhard et al., 2008).

55 The UV altitude gradients obtained from model calculations vary within the range of 3.5-6%/km in the cloudless  
56 atmosphere if all other parameters (ozone, aerosol, surface albedo) do not change with the altitude (Chubarova and  
57 Zhdanova, 2013). **The** smaller values of the estimated UV altitude gradients (3.5%/km) were obtained in conditions  
58 with high surface albedo at both sea level and high altitude, since the larger diffuse component at sea level, to some  
59 extent, compensates the higher direct flux due to a smaller total optical depth at higher **location**. However, the  
60 experimental UV altitude gradients are often much higher due to the presence of additional altitude changes in the  
61 atmospheric parameters. According to different field campaigns UV altitude gradients vary within 5-11%/km  
62 (Pfeifer et al., 2006, Zaratti et al., 2003, Schmucki, Philipona, 2002), 11-14%/km according to (Sola et al., 2008),  
63 and in some cases can reach 31%/km (Pfeifer et al., 2006). **The existence of spectral dependence in absorption**  
64 **coefficients of ozone as well as in molecular scattering cross sections provides a pronounced spectral character of**  
65 **the altitude UV effect, which was obtained in many publications (Blumthaler et al., 1994, Sola et al., 2008).**

66 However, the continuous UV records in mountainous area are still very rare due to the complexity of accurate UV  
67 measurements in severe conditions. The accurate results of measurements from different field campaigns devoted to  
68 the evaluation of altitude UV effects shown in (Bernhard et al., 2008, Blumthaler and Ambach, 1988, Blumthaler et  
69 al., 1994, Blumthaler et al., 1997, Dahlback et al., 2007, Lenoble et al., 2004, Piacentini, et al. 2003, Pfeifer et al.,  
70 2006, **Seckmeyer et al., 1997**, Sola et al., 2008, Zaratti et al., 2003) provide precise, however, local character of this  
71 phenomenon, which results in various altitude UV gradients.

72 At the same time, the accurate RT (Radiative Transfer) model simulations (Liou, 2010) are very time consuming  
73 and can not be used in different on-line tools or other applications. There are also a lot of UV model assessments for

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<sup>1</sup> **UV index is a widely used characteristic, which is equal to erythemally-weighted irradiance expressed in ( $W m^{-2}$ ) multiplied on 40 (Vanicek et al., 2003)**

74 the past and future UV climate scenarios but usually they are given with the coarse spatial resolution, which does  
 75 not allow a user to obtain the accurate estimates over the particular mountainous location.  
 76 In addition, the limiting factor of the UV calculation accuracy is the uncertainty of input geophysical parameters,  
 77 which significantly increases at high altitudes. Hence, another task was to obtain some generalized dependencies of  
 78 the input parameters using the available data sources.  
 79 The objective of this paper is to provide the accurate parameterization for different types of biologically active  
 80 radiation for the estimation of UV level at different altitudes taking into account the generalized altitude  
 81 dependencies of different geophysical parameters. Using the proposed parameterization we will also discuss the  
 82 consequence of the enhanced UV level at high altitudes for human health using the classification of UV resources  
 83 via a specially developed on-line interactive UV tool.

## 84 2. Materials and methods

85 In order to account for different effects of UV radiation on human health we analyze three types of BAUVR:  
 86 erythemally-weighted, vitamin D-weighted, and cataract-weighted irradiances, which are calculated using the  
 87 following equation:

$$88 \quad Q_{bio} = \int_{280}^{400} Q_{\lambda} F_{\lambda} d\lambda, \quad (1)$$

91 where  $Q_{\lambda}$  is the spectral flux density,  $F_{\lambda}$  is the respective biological action spectrum.

92 We used erythematous action spectrum according to CIE (1998), vitamin D spectrum - according to CIE (2006)<sup>2</sup>, and  
 93 cataract-weighted spectrum according to Oriowo et al. (2001). Various types of BAUVR action spectrum have  
 94 different efficiency within the UV range (Fig.1). For their characterization we used the effective wavelengths, which  
 95 are calculated as follows:

$$96 \quad \lambda_{eff} = \int Q_{\lambda} \lambda d\lambda / \int Q_{\lambda} d\lambda \quad (2)$$

99 According to our estimates, for example, at high solar elevation ( $h=60^{\circ}$ ) and for the variety of other parameters  
 100 (total ozone, aerosol and surface albedo) the effective wavelength for erythemally-weighted irradiance ( $Q_{ery}$ ) is  
 101  $\sim 317$  nm, for cataract-weighted irradiance ( $Q_{eye}$ ) -  $\sim 313$  nm, and for vitamin D-weighted irradiance ( $Q_{vitD}$ ) -  
 102  $\sim 308$  nm. These changes in effective wavelengths for various BAUVR types indicate their different sensitivity to  
 103 the ozone absorption, molecular scattering and aerosol attenuation, which vary dramatically within this spectral  
 104 range, and, as a result, explain different BAUVR responses to the changes in these geophysical parameters.

105 All the simulations were fulfilled using one dimensional radiative TUV (Tropospheric Ultraviolet-Visible) model  
 106 with 8-stream DISORT RT method (Madronich and Flocke, 1997) and 1 nm spectral resolution. The uncertainty of  
 107 the RT method is less than 1% (Liou, 2010). A good agreement between the experimental spectral data in different  
 108 geographical regions and simulated results using this RT method if the input atmospheric parameters were known  
 109 was also shown in (Badosa et al. 2007).

<sup>2</sup> Note, that a widely used conception of action spectra, which is based on the additivity of wavelength contribution, still has not been well documented for vitamin D action spectrum (Norval et al., 2010) and needs further studies.

110 Several experimental datasets were used. For obtaining the generalized altitude dependence of aerosol optical depth  
111 (AOD) we used the data of sun/sky CIMEL photometers from different AERONET sites located at different heights  
112 above sea level (Holben et al., 1998). These data account for the near-ground emission sources of the aerosol at  
113 various altitudes in the aerosol column content. The estimated uncertainty for aerosol optical depth in UV spectral  
114 region is about 0.02. The uncertainty for single scattering albedo is about 0.03 at  $AOD_{440} > 0.4$  and the uncertainty  
115 for all other inversion parameters is not higher than 10% (Holben et al., 2006).

116 In addition, the dataset of historical Moscow State University complex field campaigns over mountainous areas at  
117 Pamir (38- 40.5° N, 73-74° E H=1.0÷3.9 km), and Tyan'Shan' (43°N, 77°E, H=3.47km) was applied in the analysis  
118 (Belinski et al., 1968). It includes the data records of total ozone content and aerosol optical depth at 330 nm, which  
119 had been measured with the help of M-83 filter ozonometer, and UV irradiance less 320 nm – by the UVM-5  
120 instrument calibrated against the spectroradiometer BSQM (the Boyko's Solar Quartz Monochromator) described in  
121 (Belinski et al., 1968). The description of the BSQM and the details of the calibration were also discussed in  
122 Chubarova and Nezval' (2000). The uncertainties of UV measurements less 320nm due to the calibration procedure  
123 were considered to be about 10% (Belinsky et al., 1968). However, to avoid the calibration errors only relative  
124 measurements were used in this study. The residual uncertainty due to possible existence of slight variation in  
125 spectral response of the instrument and their temperature dependence was estimated to be about 5-7%. The obscurity  
126 of the horizon at all sites was less than 10°. The field campaigns were carried out during summer periods, when no  
127 snow was detected at the surface. The snow covered mountainous peaks were only observed at Tyan'Shan' at  
128 relatively large distance of more than 30 km from the site.

129 We also used the LIVAS database (Lidar Climatology of Vertical Aerosol Structure for Space-Based Lidar  
130 Simulation Studies, <http://lidar.space.noa.gr:8080/livas/>). This is a 3-dimensional global aerosol climatology based  
131 on satellite lidar CALIPSO observations at 532 and 1064 nm, EARLINET ground-based measurements and a  
132 combination of input data from AERONET, aerosol models, etc. The final LIVAS climatology includes 4-year  
133 (2008 – 2011) time-averaged  $1 \times 1^\circ$  global fields (Amiridis, et al., 2015). We used the annual aerosol extinction  
134 profiles at 355 nm for calculating aerosol optical depth over various points at different altitudes in the Alpine and  
135 the Caucasian mountainous regions in Europe and over the high-elevated regions in Asia. It should be mentioned  
136 that the LIVAS averages all Calipso overpasses over a  $1 \times 1^\circ$  cell and characterizes only the mean altitude within the  
137 cell. This provides some additional uncertainties in its aerosol extinction altitude dependence evaluation. On  
138 average, according to (Amiridis et al., 2015) the absolute difference in LIVAS AOD is within 0.1 agreement with  
139 AERONET AOD values in UV and visible region of spectrum.

140 In addition, we estimated UV resources at different altitudes according to the approach given in Chubarova and  
141 Zhdanova (2013), which has been developed on the base of international classification of UV index (Vanichek et  
142 al., 2000) and the vitamin D threshold following the recommendations given in CIE (2006). In CIE (2006) there  
143 were simple recommendations of choosing the minimum vitamin D dose (*MVitDD*) threshold using one fifth  
144 Minimal Erythematol Dose (*MED*) for a one fifth body area. In this study according to the new guidelines a healthy  
145 level of vitamin D3 was increased from 400 IU recommended in CIE (2006) to 1000 IU (Rationalizing  
146 nomenclature for UV doses..., 2014). The possibility to account for the open body fraction as a function of the  
147 effective air temperature was also applied in the UV resources estimating method (Chubarova, Zhdanova 2013) as it  
148 had been proposed in (McKenzie et al., 2009). According to this approach we defined *noon UV deficiency* when UV  
149 dose is smaller than the vitamin D threshold during 11:30-12:30 noon period, and *100% UV deficiency category*,  
150 when it is not possible to receive vitamin D throughout the whole day. The *UV optimum* category is determined



151 when the UV dose does not exceed erythema threshold but it is possible to receive UV dose, necessary for vitamin  
 152 D at noon hour. Several subclasses of *UV excess* are attributed to the thresholds depending on the standard UV  
 153 index categories: *moderate UV excess* class, which relates to moderate category of hourly UV index, *high UV*  
 154 *excess*, *very high UV excess*, and *extremely high UV excess* category. Further details about this approach can be  
 155 found in (Chubarova, Zhdanova, 2013). Currently, in the assessment of UV resources we do not take into account  
 156 for the eye damage UV effects, since there is no reliable regulation on the UV threshold for this type of BAUVR.

### 157 3. Results

#### 158 3.1. The general description of the approach

159 It is widely known that “the solution of the radiative transfer equation is possible to derive by numerous solution  
 160 methods and techniques” (Liou, 2010). However, the accurate RT methods usually require a lot of computer time  
 161 and can not be used in several applications. The simulated intensity and UV flux density (or irradiance) has a  
 162 complicated non-linear dependence on many geophysical parameters, however, our numerous simulations of UV  
 163 irradiance using the accurate 8-stream discrete ordinate RT method show that within a variety of geophysical  
 164 parameters one can obtain the parameterized altitude correction by taking into account for the quasi-independent  
 165 terms driven by different geophysical factors. Some of them are independent due to different vertical profiles (for  
 166 example, ozone maximum in the stratosphere compared with aerosol and molecular maximum in the troposphere).  
 167 Some of them are dependent (for example, surface albedo UV effects depend on molecular and aerosol loading),  
 168 but, as we show later, this factor can be also considered as a one joint term.  
 169 Using this assumption, we propose a parameterization, where biologically active UV irradiance at the altitude  $H$   
 170 ( $Q_{bio}(H)$ ) can be estimated from  $Q_{bio}$  at zero altitude ( $H=0$  km a.s.l.) with an independent account for the terms,  
 171 which are affected by different geophysical factors:

$$172 Q_{bio}(M_H, X_H, AOD_H, S_H) = Q_{bio}(M_0, X_0, AOD_0, S_0) \cdot A_M A_X A_{AOD} A_{S(M, AOD, cloud)}, \quad (3)$$

173 where  $A_M$ ,  $A_X$ ,  $A_{AOD}$  are the UV amplifications, respectively, due to the altitude decrease in molecular number  
 174 density ( $M$ ), ozone ( $X$ ), and aerosol optical depth ( $AOD$ ).  $A_{S(M, AOD, cloud)}$  is the UV amplification due to the increase in  
 175 surface albedo  $S$ , which is typically observed with the altitude. This characteristic is also a function of a change in  
 176 molecular number density, aerosol and cloud characteristics with height due to the processes of multiple scattering.  
 177 Further, only the effects in the cloudless atmosphere are considered. The total UV amplification ( $A_{UV}$ ) with altitude  $H$   
 178 can be rewritten from Eq.(3) as:

$$179 A_{UV} = A_M A_X A_{AOD} A_S = \frac{Q_{bio}(M_H, X_H, AOD_H, S_H)}{Q_{bio}(M_0, X_0, AOD_0, S_0)} \quad (4)$$

180 Let us consider separately the effects of different factors on UV irradiance at high altitudes. We specify them by  
 181 using the accurate RT model simulations, different empirical datasets or by applying the important characteristics  
 182 from different publications.

183 The possibility of this approach was tested directly by the accurate modelling for a variety of conditions at different  
 184 solar elevations. The model simulations were made for the altitude changes from zero to 5 km with the variations of  
 185 aerosol optical depth at 340nm within  $AOD_{340} \sim 0.0-0.4$ , variations in total ozone from 350 to 250 DU, and surface  
 186 albedo changes from zero to  $S=0.9$  at different altitudes. As the input aerosol parameters, within UV spectral region  
 187 we also used single scattering albedo  $SSA$  varying from 0.88 to 0.96, factor of asymmetry  $g=0.72$ , and Angstrom

188 exponent  $\alpha$  varying from 0.6 to 1.5, which are close to the aerosol background characteristics in Europe (Chubarova,  
 189 2009, Arola et al., 2009). We compared the  $A_{UV}$  values calculated as a multiplication composite of different separate  
 190 parameters ( $A_M$ ,  $A_X$ ,  $A_{AOD}$ , and  $A_S$ ) according to Eq.(4) with the  $A_{UV}$  values, which were estimated as a ratio of direct  
 191 simulations of BAUVR at the altitude  $H=5$  km and at zero ground level. The results of the comparisons are shown  
 192 in Fig.2. One can see a good agreement between the  $A_{UV}$  values obtained using multiplication of  $A_M A_X A_{AOD} A_S$  and  
 193 the  $A_{UV}$  values from direct estimations of BAUVR. The correlation for all BAUVR types is higher than 0.99 with the  
 194 mean relative difference of  $-1\pm 3\%$  compared with the exact model simulations if the same input parameters are  
 195 used. Hence, the proposed approach based on the independent account for the terms, which are affected by different  
 196 geophysical factors can be applied with high accuracy.

### 197 3.2. Molecular UV amplification with the altitude

198 A decrease of atmospheric pressure, or molecular number density, with the height is a well-known factor of UV  
 199 amplification. According to the 8-stream DISORT model simulations we found that the BAUVR dependence with  
 200 the altitude has a linear change in the molecular atmosphere, which is clearly seen in Fig.3. Hence, for its  
 201 characterization we can apply a simple gradient approach.

202 For evaluating the UV amplification due to molecular effects the following expression is used:

$$203 \quad A_M = \frac{Q_{bio}(M_H, X_0, AOD_0, S_0)}{Q_{bio}(M_0, X_0, AOD_0, S_0)} = 1 + 0.01 G_{bio, M}(S=0) \Delta H \quad (5)$$

204 where  $G_{bio, M}$  is the relative molecular gradient, in %/km,  $\Delta H$  is the difference in the altitudes, in km. Note, that all  
 205 other parameters do not change with the height.

206 The estimated relative molecular gradients for different types of BAUVR for various conditions are shown in Table  
 207 1. At solar elevation  $h=10^\circ$  there is a decrease in the  $G_{bio, M}$  for different BAUVR and, especially, for vitamin-D  
 208 irradiance due to its smaller effective wavelength and the effects of stronger ozone absorption, which is increased at  
 209 higher ozone content ( $X=500$  DU). However, for solar elevation higher than  $20^\circ$  the sensitivity of the  $G_{bio, M}$  values is  
 210 around 6-7%/km and does not significantly change with variations in  $h$  and  $X$ .

211 As an example, at the altitude of 5 km and at high solar elevation the molecular UV amplification according to Eq.  
 212 (5) lies within  $\sim 1.26$ - $1.38$  depending on the type of BAUVR (see Table 1), which is in accordance with the accurate  
 213 model simulations. However, at  $h=10^\circ$  the UV amplification for erythemally-weighted and cataract-weighted  
 214 irradiances is about 1.18-1.23, while for vitamin D-weighted irradiance  $A_M$  is only 1.04-1.09 depending on ozone  
 215 content. The maximum UV amplification at the highest peak (m. Everest,  $H=8.848$  km) due to changes only in  
 216 molecular scattering reaches 1.53-1.68 at high solar elevation depending on the type of BAUVR.

### 217 3.3. Ozone UV amplification with the altitude

218 In order to account for the ozone decrease with the altitude we apply the existing linear dependence between UV  
 219 radiation and total ozone  $X$  in log-log scale. This approach was used in the definition of the Radiation Amplification  
 220 Factor ( $RAF$ ) by Booth and Madronich (1994). As a result, the following equation can be written:

$$221 \quad \log(Q_{bio}) = RAF(Q_{bio, h}) \log(X_i) + C, \quad (6)$$

222 where  $h$  is the solar elevation,  $C$  is the constant.

223 The  $RAF$  values vary for different types of BAUVR. For example, at high solar elevation Radiation Amplification  
 224 Factor for erythemally-weighted irradiance  $RAF_{Q_{ery}} = 1.2$ , for vitamin D-weighted irradiance -  $RAF_{Q_{vitD}} = 1.4$ , for

225 cataract-weighted irradiance -  $RAF_{Q_{eye}}=1.1$  (UNEP, 2011). However, we should take into account the  $RAF$   
 226 dependence on solar elevation  $h$  due to the relative changes in solar spectrum with  $h$ . Using the results of accurate  
 227 RT modelling and polynomial regression approach we have obtained  $RAF$  dependencies on solar elevation in degree  
 228 over  $h=10-90^\circ$  range for different types of BAUVR:

$$229 \quad RAF_{Q_{ery}}(h) = -1.10E-04 \pm 1.49E-5 h^2 + 1.57E-02 \pm 1.53E-3 h + 0.665 \pm 0.0333 \quad (7)$$

$$230 \quad R^2 = 0.98$$

$$231 \quad RAF_{Q_{vitD}}(h) = 1.66E-4 \pm 1.0E-5 h^2 - 2.77E-2 \pm 1.1E-3 h + 2.5121 \pm 0.0233 \quad (8)$$

$$232 \quad R^2 = 0.997$$

$$233 \quad RAF_{Q_{eye}}(h) = 1.43E-6 \pm 1.0E-6 h^3 - 2.02E-4 \pm 6.6E-5 h^2 + 4.83E-3 \pm 2.9E-3 h + 1.297 \pm 0.035 \quad (9)$$

$$234 \quad R^2 = 0.98$$

235 where  $R^2$  – is the determination coefficient. The standard error estimates of the coefficients in the equations are  
 236 given at P=95%.

237 Note, that similar approach for accounting the  $RAF$  solar angle dependence was proposed in Herman (2010) with  
 238 higher power degree.

239 As a result, the BAUVR at the altitude  $H(Q_{bioH})$  with the correction on ozone content can be written as follows:

$$240 \quad Q_{bioH} = Q_{bio0} (X_0/X_H)^{RAF(Q_{bio,h})} \quad (10)$$

241 From Eq.(10) we can obtain the altitude UV amplification due to ozone using the altitude ozone gradient  $G_X$   
 242 (DU/km):

$$243 \quad A_X = \frac{Q_{bio}(M_0, X_H, AOD_0, A_0)}{Q_{bio}(M_0, X_0, AOD_0, A_0)} = \left( \frac{X_0}{X_0 - G_X * \Delta H} \right)^{RAF(Q_{bio,h})} \quad (11)$$

244 We propose to apply the typical ozone altitude gradient  $G_X$ , which absolute value is about 3.5 DU/km according to  
 245 monthly averaged ozone soundings measurements in Germany and observations in Bolivia (Reuder and Koepke,  
 246 2005; Pfeifer et al., 2006).

247 As an example, if we take into account only for this typical ozone decrease with the altitude, the UV enhancement at  
 248 5 km will be about  $A_X \sim 1.06-1.11$  while at the highest peak (m. Everest,  $H=8848$  m)  $A_X$  will reach 1.11-1.22 at high  
 249 solar elevations depending on the BAUVR type and initial ozone content at zero altitude within  $X=250-350$  DU.

### 250 3.4. Aerosol UV amplification with the altitude

251 Aerosols can significantly change their characteristics with the altitude, affecting the level of BAUVR. Due to  
 252 variations in size distribution and optical properties aerosol may have different radiative properties (aerosol optical  
 253 depth, single scattering albedo, and phase function). One of the most important aerosol characteristics affecting UV  
 254 radiation is aerosol optical depth.

255 For accounting the aerosol effect on UV attenuation we propose to apply the equation given in (Chubarova, 2009):

$$256 \quad Q_{bio}(AOD_{340}) = Q_{bio}^*(AOD=0) (1 + AOD_{340} B) \quad (12)$$

257 where  $B = (0.42m + 0.93) SSA - (0.49m + 0.97)$ ,  $Q_{bio}^*$  is the BAUVR in aerosol free conditions,  $m$  is the air mass,  $SSA$  is  
 258 the single scattering albedo.

259 The Eq. (12) was obtained from the accurate model simulations for the conditions with low surface albedo ( $S=0.02$ ),  
 260 which is typical for grass. (Here and further we consider AOD at wavelength 340nm ( $AOD_{340}$ ), since this  
 261 wavelength corresponds to the standard UV channel in CIMEL sun/sky photometer, which is used in AERONET).  
 262 The coefficients were obtained according to model simulations for  $0 < AOD_{340} < 0.8$ , single scattering albedo  
 263 ( $0.8 < SSA < 1$ ), airmass  $m \sim \sinh^{-1}$  ( $m \leq 2$ ), and Angstrom exponent  $\alpha \sim 1$  ( $0.6 < \alpha < 1.5$ ). Note, that these are typical  
 264 changes in main aerosol properties for European conditions in UV-A spectral range (Chubarova, 2009). However,  
 265 the Angstrom exponent in UV-B spectral region can differ from this range and be even negative in particular  
 266 conditions depending on aerosol size distribution and optical properties (Bais et al., 2007). Single scattering albedo  
 267 in UVB spectral range according to the results of different field campaigns (UNEP, 2015) may vary from 0.7 to  
 268 0.97 with low SSA in the presence of black and brown carbon aerosol. Some results demonstrates no existence of  
 269 SSA spectral dependence in UV (Barnard et al., 2008, Arola et al., 2009) but some results shows its spectral  
 270 character (UNEP, 2015). We should also note that direct evaluation of SSA in UV-B spectral region is difficult  
 271 because UV attenuation due to aerosol can occur together with the absorption in this spectral region by different  
 272 gases (ozone, sulphur dioxide, formaldehyde, nitrogen dioxide, etc.). As a result we used the aerosol properties at  
 273 340nm as input parameters for the BAUVR simulations since we consider typical aerosol conditions without forest  
 274 fires and heavy industrial smoke. Radiative effects of the existing AOD spectral dependence are relatively small  
 275 within the UV-B spectral range therefore we consider the same coefficients for different types of BAUVR.

276 Assuming that single scattering albedo and factor of asymmetry do not change with the altitude, we evaluated the  
 277 UV amplification with the altitude due to aerosol optical depth. Using Eq.(12) the equation for  $A_{AOD340}$  can be  
 278 written as follows:

$$279 \quad A_{AOD340} = \frac{Q_{bio}(M_0, X_0, AOD_H, A_0)}{Q_{bio}(M_0, X_0, AOD_0, A_0)} = \frac{1 + AOD_{340, H} B}{1 + AOD_{340, 0} B} \quad (13)$$

280  
 281 In some conditions single scattering albedo and asymmetry factor for visible wavelengths may have the altitude  
 282 dependence (see, for example, the results of aircraft measurements in Western Siberia (Panchenko et al., 2012)).  
 283 However, there is no information on the altitude dependence of aerosol properties in UV spectral region from the in-  
 284 situ measurements over the PEEEX area. Note, that the uncertainty of neglecting the altitude changes in single  
 285 scattering albedo significantly decreases at small AOD observed at high altitudes and only the altitude changes in  
 286 aerosol optical depth are usually taken into account in the standard tropospheric aerosol models (WMO, 1986).

287 The aerosol optical depth at the altitude  $H$  ( $AOD_{340, H}$ ) can be evaluated using the following expression:

$$288 \quad AOD_{340, H} = AOD_{340, 0} f_{AOD}(H), \quad (14)$$

289 where  $f_{AOD}(H)$  is the altitude dependence of aerosol optical depth.

290 There are a lot of model aerosol profiles for the free atmosphere conditions (see, for example, widely used aerosol  
 291 models in WMO, (1986)). However, these profiles can not be applied for high-elevated locations, which are usually  
 292 characterized by a significant emission of primary aerosols or their precursors from nearby surface even in  
 293 background conditions. To account for this kind of altitude AOD dependence we used different ground-based and  
 294 satellite measurements described in the Section 2. Since our objective was to obtain the generalized aerosol altitude  
 295 dependencies we used the monthly mean AOD data from different archives over different geographical regions in  
 296 Eurasia. The dependence of aerosol optical depth as a function of altitude is shown in Fig.4. Highly variable  $AOD_{340}$   
 297 values at different altitudes may be roughly combined in two groups, which are characterized by different rates of  
 298 aerosol altitude decrease. Hence, in our parameterization we propose to distinguish these two types of the altitude

299 aerosol dependence. The first one is characterized by a very strong aerosol optical depth decrease with the altitude.  
300 It was obtained mostly from the data of European AERONET sites in the Alpine zone as well as from several Asian  
301 sites in the sharp-peak mountainous areas. This dependence was also confirmed by the LIVAS dataset  
302 measurements over the same areas.

303 The second one is characterized by a much more gradual altitude  $AOD_{340}$  decrease observed over flat elevated Asian  
304 regions according to the AERONET and LIVAS dataset and the data obtained during Moscow State University field  
305 campaigns at the high-altitude plateau at Pamir and Tyan'-Shan' mountainous regions in Central Asia. The main  
306 reason of such a character is the existence of the additional aerosol emission sources (i.e. loess, mineral aerosol)  
307 from the vast areas of deserts and semi-deserts elevated over sea level of up to 3-4 kilometers.

308 The first, Alpine – like type, can be parameterized as:

$$309 \quad f_{AOD}(H)=H^{1.65}, R^2=0.4 \quad (15)$$

310 The Alpine type aerosol altitude dependence was firstly obtained for the simulation of the UV climatology over  
311 Europe (COST 726, 2010). However, the coefficients have been re-affirmed according to the monthly mean AOD  
312 over 1999-2012 period (case number N=137).

313 The second, so-called Asian type, can be obtained using the equation according to the Moscow State University  
314 expedition dataset in Asian region. It is characterized by much flat dependence with the altitude:

$$315 \quad f_{AOD}(H)= \exp(-0.26H), R^2=0.8 \quad (16)$$

316 The proposed dependencies can be considered as the two classes with different altitude aerosol decreasing rates.  
317 Both dependencies are accounted for the altitudes higher than 1 km, since our analysis of AERONET dataset has  
318 revealed the absence of the aerosol altitude dependence at the heights below 1 km due to prevailing the effects of  
319 different aerosol sources or their precursors there. However, over the particular location the altitude AOD  
320 dependence within the first kilometre can be found, of course.

321 We should note that the proposed altitude AOD dependencies according to (15) and (16) are considered only as a  
322 first proxy for the most sharp and flat altitude dependencies. For a particular location and specific geographical  
323 conditions the AOD altitude dependence can be different. However, a user may easily substitute them in the  
324 proposed parameterization.

325 Although the AOD altitude dependence is pronounced, its influence on UV amplification highly depends on initial  
326 aerosol conditions at  $H=0$  km, the type of the altitude profile, and solar elevation (see Eq. (12)). For example, for the  
327 Alpine type aerosol altitude profile the UV amplification at  $H=5$  km is about  $A_{AOD}=1.05-1.10$  and does not exceed  
328  $1.11$  at  $H=8.848$  km for typical aerosol at  $H=0$  km ( $AOD_{340}=0.36$ ). However, for the polluted conditions  
329 characterized by  $AOD_{340}=0.8$  at  $H=0$  km, the altitude UV amplification at  $H=5$  km is about  $A_{AOD}\sim 1.12-1.27$   
330 depending mainly on solar elevation. Note, that at  $H=8.848$  km the effect is almost the same ( $A_{AOD}\sim 1.16-1.29$ ). This  
331 will be further discussed below.

### 332 3.5. UV amplification due to changes in surface albedo with the altitude

333 The increase in surface albedo is one of the important factors, which is necessary to account for the effective  
334 calculations of BAUVR at high altitudes. Due to significant negative temperature gradients, the snow with high  
335 surface albedo can be observed even in summer conditions at high altitudes instead of vegetation with low UV  
336 albedo of about  $S=0.02-0.05$  (Feister and Grewe, 1995). Fig.5 demonstrates the enhancement in erythemally-  
337 weighted irradiance due to the increase in surface albedo according to different experimental studies and the results

338 of one-dimensional model simulations. One can see the UV increase of around 20% at the effective surface albedo  
 339 close to  $S=0.5$  (Simic et al., 2011, Huber et al., 2004, Smolskaia et al., 2003). On the average, there is an agreement  
 340 between the calculation of UV amplification by 1D model and the measurements at different mountainous regions  
 341 up to effective surface albedo of  $S\sim 0.5$ . However, the accurate comparison of UV measurements with 3D model  
 342 (Diemoz and Mayer, 2007) shows the additional snow effect of about  $\pm 1$ UV index value due to the account of  
 343 overall interactions between radiation and different surfaces. The comparisons of UV spectral actinic flux  
 344 measurements with 1D and 3D model simulations demonstrate the similar range of uncertainties of these models,  
 345 however, 3D model gives, of course, more realistic view of the UV field in mountains since the topography and the  
 346 obstruction of the horizon are taken into account (Wagner et al., 2011). However, currently we do not consider 3D  
 347 effects in our parameterization. Due to small UV albedo over snow free surfaces this factor is negligible in summer,  
 348 while in winter the value of the effective surface albedo in mountainous area can be very high and significantly  
 349 depends on tree line location.

350 To account for surface albedo effects we followed the results obtained in different papers (Green, et al., 1980,  
 351 Chubarova, 1994), where the effects of multiple scattering were accounted using geometric progression approach.  
 352 The same approach with a detailed physical analysis was used in (Lenoble, 1998). Following these publications we  
 353 propose to calculate biologically active UV radiation in conditions with surface albedo  $S$  as follows:

$$354 \quad Q_{bio_S} = Q_{bio_{S=0}} \frac{1}{1-r_{bio}S} \quad (17)$$

355 where  $r_{bio}$  is the coefficient, which characterizes the maximum relative change in  $Q_{bio}$  due to multiple scattering for  
 356 surface albedo variations from 0 to 1. According to (Lenoble, 1998)  $r_{bio}$  is determined as the atmospheric reflectance  
 357 illuminated on its lower boundary. Note, that surface albedo  $S$  characterizes the reflecting properties at ground at the  
 358 considered altitude  $H$ .

359 The application of the equation (17) to  $H=0$  km and to the altitude  $H$  allows us to obtain the following expression  
 360 for  $Q_{bio}$  at  $H$  with surface albedo  $S_H$ : using the known  $Q_{bio}$  at the altitude  $H=0$  km with surface albedo  $S_0$ :

$$361 \quad Q_{bio_{S_H}}(H) = Q_{bio_{S_0}}(H=0) \frac{Q_{bio_{S=0}}(H) \frac{1-r_{bio}(H=0)S_0}{Q_{bio_{S=0}}(H=0) 1-r_{bio}(H)S_H}}{Q_{bio_{S=0}}(H=0) 1-r_{bio}(H)S_H} \quad (18)$$

362 This equation can be rewritten in the following way:

$$363 \quad \frac{Q_{bio_{S_H}}(H)}{Q_{bio_{S_0}}(H=0)} = \frac{Q_{bio_{S=0}}(H) \frac{1-r_{bio}(H=0)S_0}{Q_{bio_{S=0}}(H=0) 1-r_{bio}(H)S_H}}{Q_{bio_{S=0}}(H=0) 1-r_{bio}(H)S_H} \quad (19)$$

364 One can see that in Eq. (19) the left side of the equation  $\frac{Q_{bio_{S_H}}(H)}{Q_{bio_{S_0}}(H=0)}$  is the total UV amplification  $A_{UV}$  defined in  
 365 equation (4); the first term  $\left(\frac{Q_{bio_{S=0}}(H)}{Q_{bio_{S=0}}(H=0)}\right)$  at the right side of the equation characterizes the total UV amplification in  
 366 conditions with  $S=0$ , while surface albedo effects are accounted only in the last term. Hence, we can write the UV  
 367 amplification due to the effects of surface albedo as follows:

$$368 \quad A_S = \frac{1-r_{bio}(H=0)S_0}{1-r_{bio}(H)S_H} \quad (20)$$

369 According to the model estimations the value  $r_{bio}$  in clear sky conditions has the dependence on altitude, which  
 370 appears due to a decrease mainly in molecular and aerosol loading and can be parameterized by a simple regression  
 371 as follows:

$$372 \quad r_{bio}(H) = b H + c, \quad (21)$$



373 where the coefficients  $b$  and  $c$  are given in Table 2 for different types of BAUVR. They were obtained for a variety  
374 of solar elevation and ozone content taking into account for the altitude changes in molecular scattering as well as  
375 for altitude dependence of aerosol optical depth  $f_{AOD}(H)$ . The  $r_{bio(H)}$  mainly depends on molecular content and  
376 aerosol properties, and slightly decreases with the altitude due to reducing in multiple scattering effects with the  
377 decrease in molecular and aerosol loading.

378 As a result, the UV amplification due to the increase in surface albedo at the altitude  $H$  strongly depends on  
379 scattering processes and also decreases with the altitude. Fig.6 shows the maximum  $A_S$  effect due to the changes in  $S$   
380 from  $S=0$  at zero level to  $S=1$  at the altitude  $H$  for the different types of BAUVR. The  $A_S$  decreases with the altitude  
381 from more than 1.6 at  $H=0$  km to about 1.2 at  $H=8.848$  km due to the decrease in  $r_{bio}$ .

### 382 3.6.Validation

383 Using the generalized parameterizations for different geophysical parameters obtained in the previous sections we  
384 can estimate the total UV amplification  $A_{UV}$  with the altitude from Eq.(4). The results of the validation of the  
385 proposed method with these altitude parameter dependencies against the accurate RT simulations are shown in  
386 Fig.7. One can see a high correlation ( $r>0.996$ ) between the  $A_{UV}$  values obtained by the proposed method and the  
387 accurate RT simulations using the 8-stream DISORT method within the changes in altitude from  $H=0$  km to 8 km,  
388 in solar elevation from 20 to 50°, in surface albedo from  $S=0$  to  $S=0.9$ , in ozone from 250 DU to 350 DU at  $H=0$  km,  
389 in  $AOD_{340}$  from 0.2 to 0.4 at  $H=0$  km, and in SSA varying from 0.88 to 0.96. Different altitude aerosol profiles were  
390 also considered. Validation was made for all three types of BAUVR. Overall, the average bias is  $0\pm2\%$  for  
391 erythemally-weighted irradiance,  $0\pm1\%$  - for cataract-weighted, and  $1\pm1\%$  - for vitamin D irradiance. The  
392 maximum difference between the  $A_{UV}$  calculated by the proposed method and by the accurate model simulations  
393 does not exceed 6% at the highest elevation ( $H=8$  km) at low ozone content.

394 The comparisons of the total UV amplification according to the proposed method with the total  $A_{UV}$  obtained from  
395 the experimental dataset as a function of altitude are shown in Fig.8. The experimental data were taken from the  
396 dataset of Moscow State University mountainous field campaigns, which was described in the Section 2. After  
397 accounting for the molecular, aerosol, and ozone altitude dependence the simulated UV amplification is in  
398 satisfactory agreement with the obtained experimental results.

### 399 4.Discussion

400 The total altitude amplification of biologically active UV irradiance  $A_{UV}$  as a function of altitude is shown in Fig.9  
401 for a variety of atmospheric conditions at surface albedo  $S=0$  and  $S=0.9$  and high solar elevation  $h=50^\circ$ . One can see  
402 a distinct altitude difference obtained for different types of BAUVR with larger increase for vitamin D-weighted  
403 irradiance due to its higher sensitivity to ozone content. The difference in  $A_{UV}$  for various BAUVR can reach 15-  
404 20% at high altitudes. The effects of surface albedo on  $A_{UV}$  can be seen if compare the results shown in Fig. 9a and  
405 Fig.9b. One can see the 2-2.5 times UV increase due to high surface albedo at high altitudes, which is again more  
406 pronounced for vitamin D-weighted radiation with smaller effective wavelength and, hence, more effective multiple  
407 scattering than that for the other types of BAUVR. Larger  $A_{UV}$  values are also observed in conditions with smaller  
408 ozone amount for all three types of BAUVR for both zero and high surface albedo conditions. High surface albedo  
409  $S=0.9$  provides a significant increase even at zero level, which is similar to the  $A_{UV}$  increase due to altitude change  
410 of 6 km. It is clearly seen that typical aerosol and ozone does not play a vital role in  $A_{UV}$ . However, for all types of

411 BAUVR the increase of slightly absorbing aerosol (from  $AOD_{340}=0.2$  to  $AOD_{340}=0.4$ ) provides a noticeable  $A_{UV}$   
 412 growth in conditions with relatively small ozone content due to enhancement of multiple scattering (see Fig.9).  
 413 The  $A_{UV}$  values are smaller at solar elevation  $h<20^\circ$  for all types of BAUVR mainly due to decreasing in  $G_{bio\_M}$  (see  
 414 the coefficients in Table 1). For example, at  $H=8$  km the UV altitude amplification for vitamin D-weighted radiation  
 415 is about  $A_{UV}=1.77$  at  $h=10^\circ$  compared with  $A_{UV}=2.0$  at  $h=50^\circ$  at  $X=250$  DU and  $AOD_{340}=0.4$ .  
 416 Let us consider the conditions, which are characterized by the most pronounced UV amplification with the altitude -  
 417 the conditions with high aerosol loading  $AOD_{340}=0.8$ , low ozone content  $X=250$ DU at  $H=0$  km, and high solar  
 418 elevation  $h=50^\circ$ . In addition, we consider the abrupt increase in surface albedo at  $H=2$  km from  $S\sim 0$  to  $S=0.95$ ,  
 419 which can be possible due to location of tree line there and pure snow above. The altitude UV amplification due to  
 420 these input parameters according to the proposed  $A_{UV}$  parameterization is shown in Fig.10. The separate effects of  
 421 different factors can be seen in Fig. 10a and their total effects on different BAUVR types are shown in Fig.10b. One  
 422 can see a different role of geophysical factors at different altitudes: the prevalence of molecular scattering especially  
 423 at high altitudes while the extremely high surface albedo may play the most important role at the altitudes of its  
 424 abrupt increase (in our case -  $H=2$ km,  $A_S=1.48$ ). However, in our example the UV amplification due to surface  
 425 albedo decreases at the altitude higher than 2 km because of the reduction in multiple scattering. The UV altitude  
 426 amplification due to aerosol is more distinct and reaches 1.1-1.2 at high altitudes if there is a strong aerosol pollution  
 427  $AOD_{340}=0.8$  at  $H=0$  km. It is more pronounced for the Alpine-type AOD altitude dependence and in our example at  
 428  $H=2$  km it can be even higher than the  $A_M$  value (see Fig.10). The effects of ozone in UV amplification do not  
 429 exceed 1.1-1.2 at high altitudes depending on BAUVR type. We would like to emphasize that Fig.10 is only the  
 430 illustration of the application of the proposed  $A_{UV}$  parameterization for a given parameters altitude variations.  
 431 We implemented the proposed UV altitude parameterization in the developed on-line UV tool <http://momsu.ru/uv/>,  
 432 which had been developed for the simulation of erythemally-weighted irradiance and the UV resources over  
 433 Northern Eurasia (the PEEEX domain) at  $H=0$  km (Chubarova and Zhdanova, 2013). Using this program it is  
 434 possible to calculate UV irradiance and UV-resources for different atmospheric conditions at a given geographic  
 435 location and specified time. Based on the threshold for vitamin D and erythemally active irradiance the UV  
 436 resources are obtained for various skin types and open body fraction. According to the classification we consider  
 437 different categories of UV-deficiency, UV-optimum and UV-excess (Chubarova and Zhdanova, 2013). The  
 438 interactive on-line UV tool represents a client-server application where the client part of the program is the web-  
 439 page with a special form for the input parameters required for erythemally-weighted UV irradiance simulations. The  
 440 server part of the program consists of the web-server and the CGI-script, where the different input parameters are set  
 441 by a user or taken from the climatological data available at the same site. In addition, in this part of the program  
 442 erythemally-weighted irradiance is calculated, visualized and classified according to the proposed UV resources  
 443 categories. The proposed UV irradiance altitude parameterization has been also incorporated in the calculation  
 444 scheme with additional account for the changes in the atmospheric parameters with the height. This enables a user to  
 445 evaluate UV irradiance at any requested elevation above sea level taking into account for a variety of the altitude  
 446 dependent parameters.  
 447 Let us analyze the UV resources for skin type 2 and the open body fraction of 0.25 in the Alpine region  
 448 (approximately  $46^\circ$ N and  $7^\circ$ E) for winter and spring noon conditions. For these conditions the vitamin D threshold  
 449 is equal to  $100 \text{ J/m}^2$  and Minimal Erythral Dose -  $250 \text{ J/m}^2$ . According to our estimates on January, 15<sup>th</sup>, at  
 450  $H=0$  km for typical (climatological) ozone, aerosol and surface albedo conditions the noon UV deficiency (no  
 451 vitamin D generation) is observed with noon erythemally UV dose of about  $97.2 \text{ Jm}^{-2}$ , while at the same



452 coordinates at  $H$  higher 0.5 km up to  $H=4.8$  km (the highest point within the Alps, peak Mont Blanc) we obtain the  
453 UV optimum conditions with noon erythemal UV dose varying from  $100.6 \text{ Jm}^{-2}$  to  $122.9 \text{ Jm}^{-2}$ .  
454 However, for skin type 4 (Fitzpatrick, 1988) with vitamin D threshold of  $180 \text{ J/m}^2$  the noon UV deficiency is  
455 observed at all altitudes and even at high surface albedo  $S=0.9$  corresponding to the pure snow with UV dose of  
456  $154.4 \text{ Jm}^{-2}$ . The decrease in open body fraction for this skin type from 0.25 to 0.05, which could take place in  
457 frosty weather, provides 100% UV deficiency, when no vitamin D can be generated during the whole day at the  
458  $H=4.8$  km and  $S=0.9$ .  
459 On April, 15<sup>th</sup>, at  $H=0$  km and typical climatological conditions at this geographical point noon UV dose is about  
460  $437.7 \text{ Jm}^{-2}$ . This means that for the open body fraction of 0.4 the moderate UV excess is observed for skin type 2  
461 and the UV optimum – for skin type 4, when vitamin D threshold is  $112.5 \text{ Jm}^{-2}$  and MED threshold -  $450 \text{ Jm}^{-2}$ . At  
462 the altitude  $H = 2$  km the conditions are characterized by the moderate UV excess for both skin types 2 and 4 with  
463 UV dose of  $463.4 \text{ Jm}^{-2}$ . At the  $H=4$  km a high UV excess is observed for skin type 2 and the moderate UV excess -  
464 for skin type 4 with UV dose of  $532.4 \text{ Jm}^{-2}$ .  
465 Thus, the proposed altitude UV parameterization can be effectively used for accurate estimating the BAUVR at  
466 different altitudes with any altitude resolution for a variety of geophysical parameters over the PEEEX domain in  
467 Northern Eurasia. The accurate RT methods like Monte-Carlo, Discrete Ordinate method or others, of course, can be  
468 used instead for UV irradiance simulations, however, their application is very time-consuming and are not possible  
469 in some applications. The proposed approach is especially very useful for the application in different kinds of on-  
470 line UV tools, where it is not possible to use a lot of prescribed calculations for a wide set of different geophysical  
471 parameters or accurate UV modelling.  
472 The current state of the online interactive tool does not take into account for the skin orientation relative to the Sun  
473 and the geometry of the human body which can modify the results limited to UV irradiance simulations on  
474 horizontal surface (Hess and Koepke, 2008, Vernez et al., 2014). In this case UV dosimeters which have a spectral  
475 response almost identical to that of the UV-induced photobiological effect (Siani et al., 2008) is the most accurate  
476 way for evaluating the individual levels of UV exposure. The vitamin D production can be also affected by other  
477 factors such as obesity and age (Engelsen, 2010). However, these are the tasks for the future work.  
478 The combination of different altitude dependencies for main geophysical factors in the proposed parameterization  
479 allows a user to make a reliable altitude UV assessment. We should also emphasize that the proposed ozone and  
480 aerosol altitude dependencies in the troposphere were taken from the experimental data obtained in background  
481 conditions and, hence, should be applied only for these conditions. However, they can be easily substituted by any  
482 other altitude dependence of considering factors.  
483 With the application of the proposed method we can also reveal the effects of different geophysical factors on  
484 various types of BAUVR and estimate their comparative role in the altitude UV effects. And, of course, the  
485 parameterization can be also used in downscaling the UV results from the coarse resolution global chemistry-  
486 climate models for the regions located at high altitudes. The proposed method can be applied not only over the  
487 PEEEX domain but on a global scale over the world. However, more attention should be paid in this case to the  
488 evaluation of the particular altitude dependence of the different parameters.

## 489 5.Conclusions

490 The objective of this paper was to develop a flexible parameterization based on rigorous model simulations with  
491 account for generalized altitude dependencies of molecular density, ozone, and aerosols considering different

492 surface albedo conditions. We show that for the specified groups of parameters we can present the altitude UV  
493 amplification ( $A_{UV}$ ) for different BAUVR as a composite of independent contributions of UV amplification from  
494 different factors with the mean uncertainty of 1% and standard deviation of 3%. The parameterization takes into  
495 account for the altitude dependence of molecular number density, ozone content, aerosol loading, and spatial surface  
496 albedo. We also provide the generalized altitude dependencies of different parameters for evaluating the  $A_{UV}$ . Their  
497 validation against the accurate RT model (8 stream DISORT RT code) for different types of BAUVR shows a good  
498 agreement with maximum uncertainty of few percents and correlation coefficient  $r > 0.996$ . It was not possible, of  
499 course, to cover all the observed variety in the parameters. However, due to the proposed approach the parameter  
500 altitude dependencies can be easily substituted by a user.

501 Using this parameterization one can estimate the role of different atmospheric factors in the altitude UV variation.  
502 The decrease in molecular number density, especially at high altitudes, and the increase in surface albedo play the  
503 most significant role in  $A_{UV}$  growth. At high solar elevations the UV amplification due to aerosol at  $H=8.848$  km  
504 does not exceed 1.3 even when  $AOD_{340}=0.8$  at  $H=0$  km. The UV amplification due to aerosol calculated with the  
505 Alpine-type AOD altitude aerosol dependence is much more pronounced than that calculated with the Asian-type  
506 AOD altitude dependence, especially at relatively lower altitudes ( $H=2-4$  km). The UV amplification due to ozone  
507 does not exceed 1.20 and is higher at smaller solar elevations, especially, for vitamin-D-weighted irradiance.

508 This parameterization was applied to the on-line tool for calculating the UV resources (<http://momsu.ru/uv/>) over the  
509 PEEEX domain. Using this tool one can easily evaluate the UV conditions (UV deficiency, UV optimum or UV  
510 excess) at different altitudes for a given skin type and open body fraction. As an example, we analyzed the altitude  
511 UV increase and its possible effects on health considering different skin types and various open body fraction for  
512 January and April conditions in the Alpine region. We showed that even in clear sky conditions over the same  
513 geographical point (46°N, 7°E) in mid-January the UV optimum can be observed at altitudes higher than  $H=0.5$  km  
514 for skin type 2, while the noon UV deficiency are still remained at the altitudes up to  $H=4.8$  km for skin type 4 when  
515 open body fraction is 0.25. In mid-April the account for the altitude dependence at 4 km provides the changes from  
516 the moderate UV excess to a high UV excess for skin type 2 and from UV-optimum the moderate UV excess - for  
517 skin type 4 when open body fraction is 0.4.

518 This approach can be also used in downscaling the results of global chemistry-climate models with the coarse spatial  
519 resolution in mountainous domain and as a simple tool for different types of applications for personal purposes of  
520 users.

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688 **FIGURE CAPTIONS**

689 Fig. 1. Action spectra for erythema (CIE, 1998), vitamin D (CIE, 2006) and for eye damage (cataract) (Oriowo et.  
690 al. 2001) effects.

691 Fig.2. The comparison of  $A_{UV}$  amplification factor calculated from Eq.(4) as multiplication of  $A_M A_X A_{AOD} A_S$  with  
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694 Comment. The simulations were performed for different altitudes ( $H=0$  and  $H=5km$ ), aerosol optical depth ( $AOD_{340}= 0, 0.2,$   
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696 albedo ( $S=0, S=0.9$ ) and solar elevation ( $h=20^\circ$  and  $50^\circ$ ). For estimating the UV amplification we assume at  $H=0$  km the  
697 conditions with 350DU,  $AOD_{340}=0.4$ ,  $S=0\%$  and normalized the BAUVR at the altitude  $H=5km$  to the value calculated with  
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701 model simulations: TUV, 8- stream DISORT TUV model.  $h=50^\circ$ .  $X=300 DU$ .

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705 were recalculated to AOD at  $\lambda=340$  nm using the Angstrom exponent  $\alpha=1.0$ . See further details in the text.

706 Fig.5. UV amplification due to the surface albedo increase in mountainous areas according to different experimental  
707 data and model simulations. The error bars of model simulation relates to the different input parameters – altitude of  
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716 Fig.8. The comparison between the simulated UV amplification according to the proposed parameterization and the  
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718 elevation  $h=50^\circ$ . Clear sky conditions. Note: since we used the data of different field campaigns the ozone altitude  
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720 and  $X\sim 250 DU$  at  $H\sim 1-2 km$ .

721 Fig.9. Total UV amplification as a function of the altitude for different types of BAUVR in a variety of atmospheric  
722 conditions with  $S=0$  (a) and  $S=0.9$  (b). The model parameters at  $H=0$  km:  $X=250-350 DU$ ,  $AOD_{340}=0.2-0.4$ . The  
723 Alpine type of AOD altitude dependence according to the Eq. (15) was taken into account. Solar elevation  $h=50^\circ$ .

724 Fig.10. The UV amplification due to molecular  $A_M(Qery)$ ,  $A_M(QvitD)$ , ozone  $A_X(QvitD)$ ,  $A_X(Qery)$ , aerosol  $A_{AOD,fl(H)}$ ,  
725  $A_{AOD,f2(H)}$  for the Alpine  $f1(H)$  and Asian  $f2(H)$  types of altitude dependences, and surface albedo  $A_S(Qery)$ ,  $A_S(QvitD)$   
726 changes with the altitude (a) and their total altitude effect on  $A_{UV}$  for different types of BAUVR (b). At  $H=0$  km:  
727  $AOD_{340}=0.8$ ,  $X=250$  DU. The surface albedo has an abrupt change at 2 km from  $S=0$  to  $S=0.95$ . Solar elevation -  
728  $h=50^\circ$ .

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730 **LIST OF THE TABLES:**

731

732 Table 1. Relative molecular gradients  $G_{bio\_M(A=0)}$ (%/km) at different solar elevations and ozone content for different  
 733 types of BAUVR. Accurate model simulations. Zero surface albedo conditions. Determination coefficient  $R^2$  is  
 734 higher than 0.997 in all cases. The standard error of  $G_{bio\_M(A=0)}$  is given in the brackets at P=95% .

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h, solar elevation, degrees	erythemally-weighted irradiance	cataract-weighted irradiance	vitamin D-weighted irradiance	erythemally-weighted irradiance	cataract-weighted irradiance	vitamin D-weighted irradiance
	X=300 DU			X=500 DU		
10	4.5 (0.04)	3.8 (0.04)	1.8 (0.01)	4.8 (0.05)	3.9 (0.04)	0.8 (0.03)
20	6.4 (0.04)	6.9 (0.05)	7.1 (0.06)	6.0 (0.03)	6.6 (0.04)	6.8 (0.05)
30	6.7 (0.01)	7.2 (0.02)	7.8 (0.02)	6.1 (0.01)	7.0 (0.01)	7.8 (0.02)
40	6.4 (0.02)	6.8 (0.01)	7.3 (0.01)	5.8 (0.02)	6.6 (0.02)	7.4 (0.01)
50	6.0 (0.03)	6.2 (0.03)	6.7 (0.03)	5.5 (0.03)	6.1 (0.03)	6.8 (0.03)
60	5.7 (0.04)	5.8 (0.04)	6.2 (0.04)	5.3 (0.04)	5.7 (0.04)	6.4 (0.04)

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738 Table 2. The coefficients for calculating the  $r_{bio}$  values (Eq. 21) for different types of BAUVR. Model estimations.

739 The standard error of the coefficients is given in the brackets at P=95%.

	erythemally-weighted irradiance	cataract-weighted irradiance	vitamin D-weighted irradiance
$b$	-0.025(0.0002)	-0.025(0.0002)	-0.025(0.0002)
$c$	0.394(0.0009)	0.394(0.0009)	0.405(0.0008)
$R^2$	>0.99	>0.99	>0.99

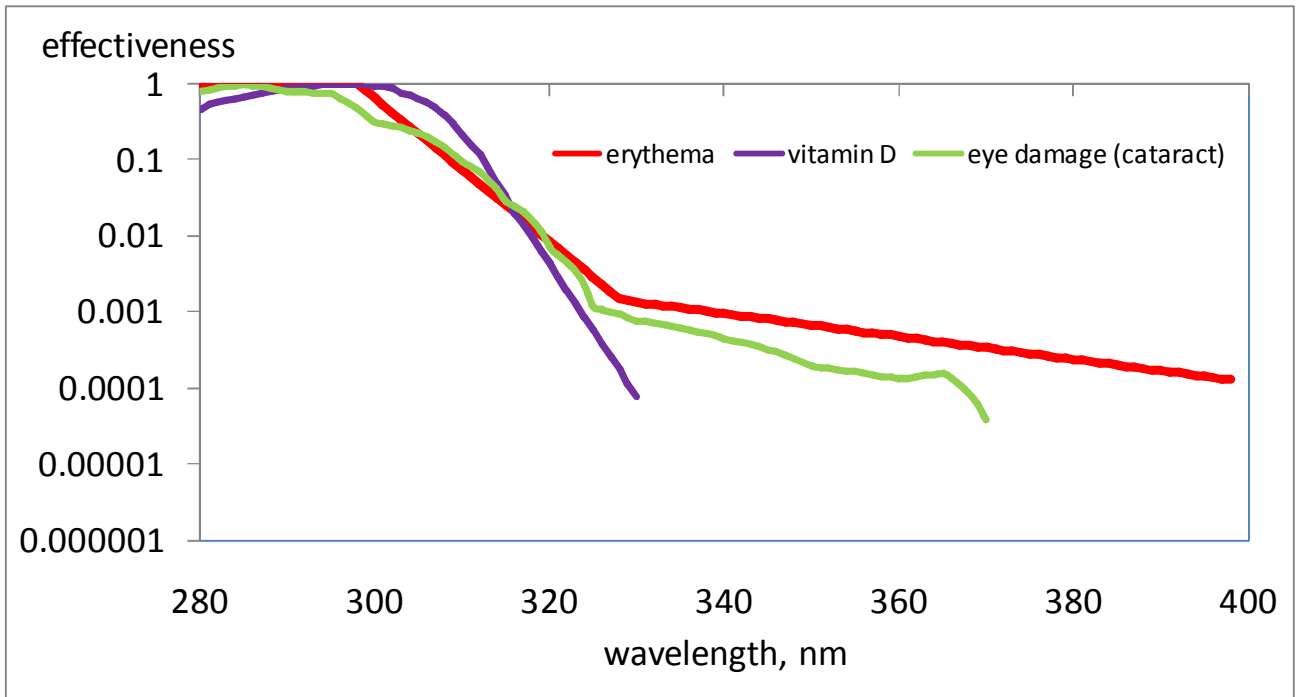
740 Note: the simulations were fulfilled for different combinations of input parameters at  $H=0$ :

741  $X=250-350$  DU,  $AOD_{340}=0.2-0.4$ ,  $S=0-0.9$ ,  $h=20-50^\circ$ .

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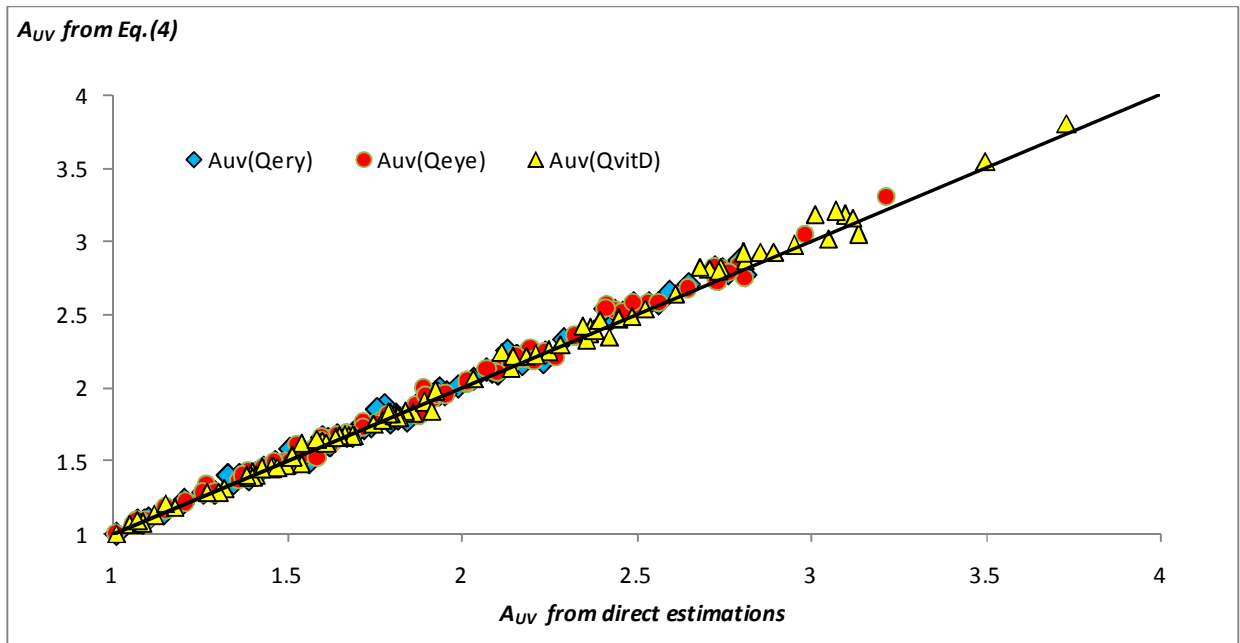
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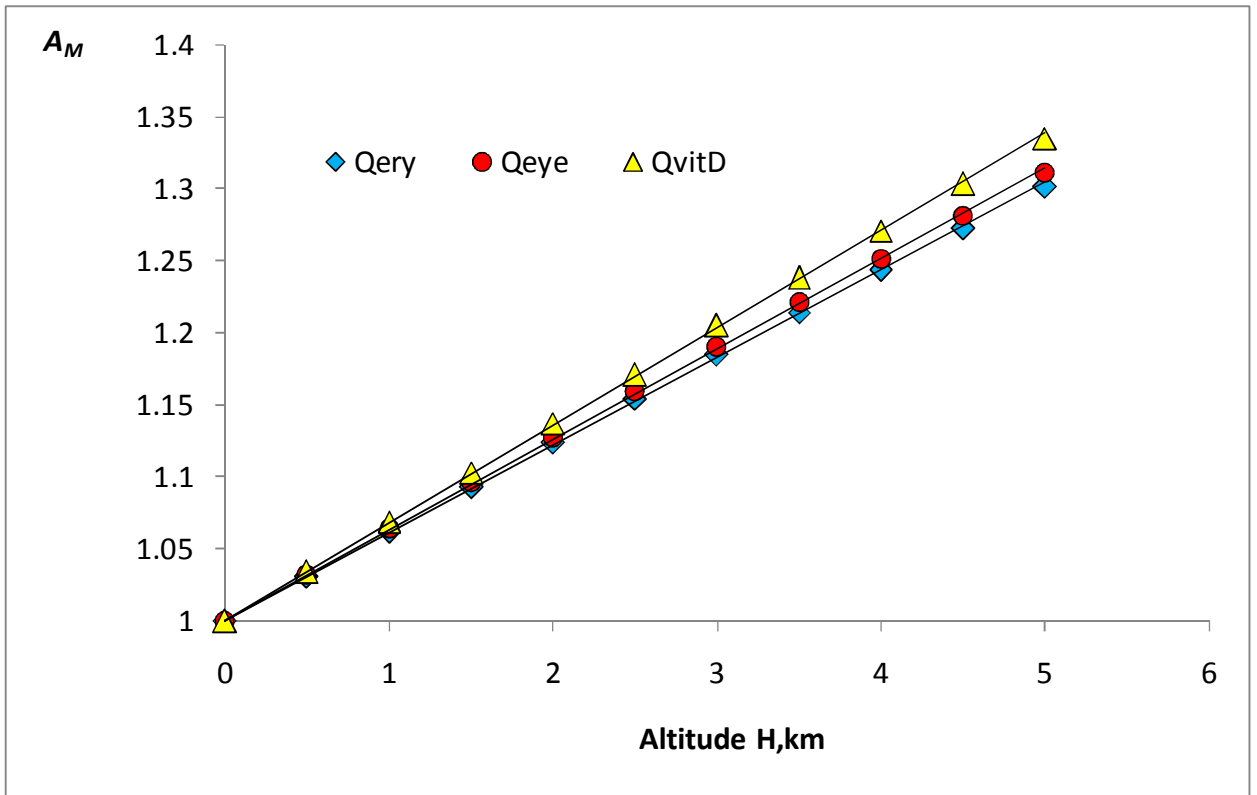
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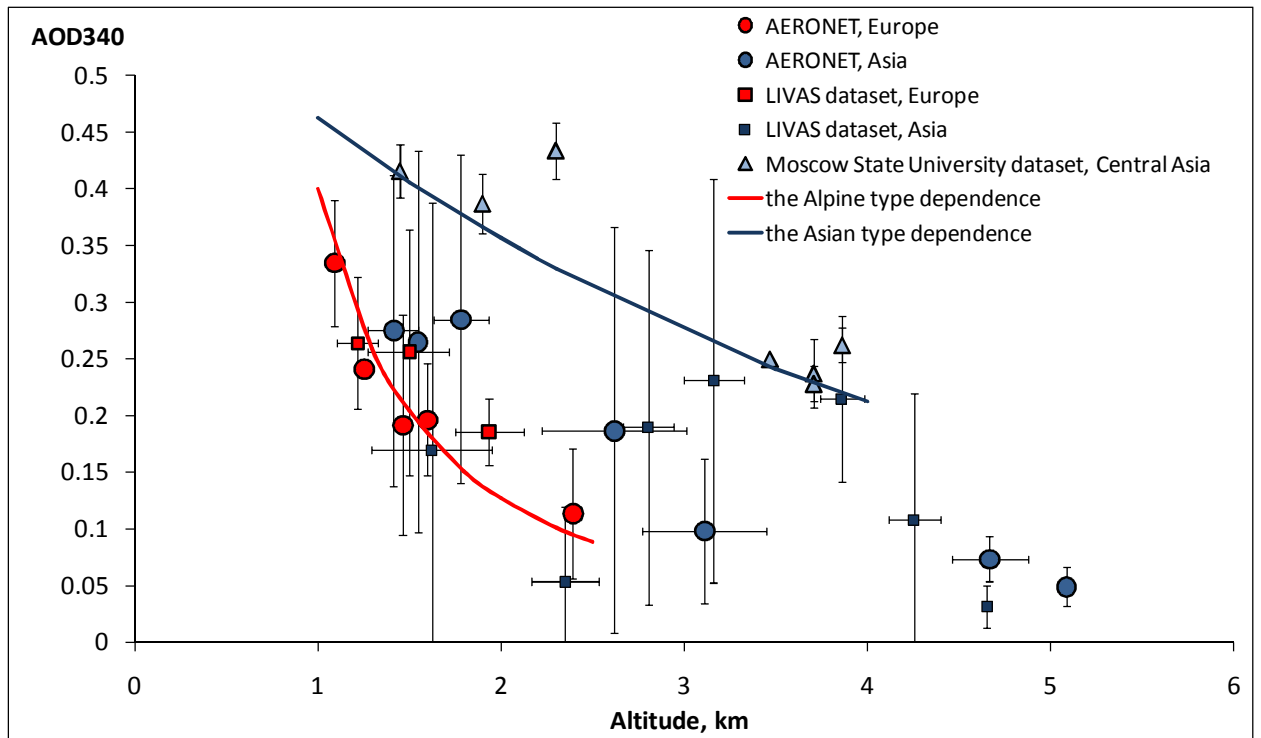
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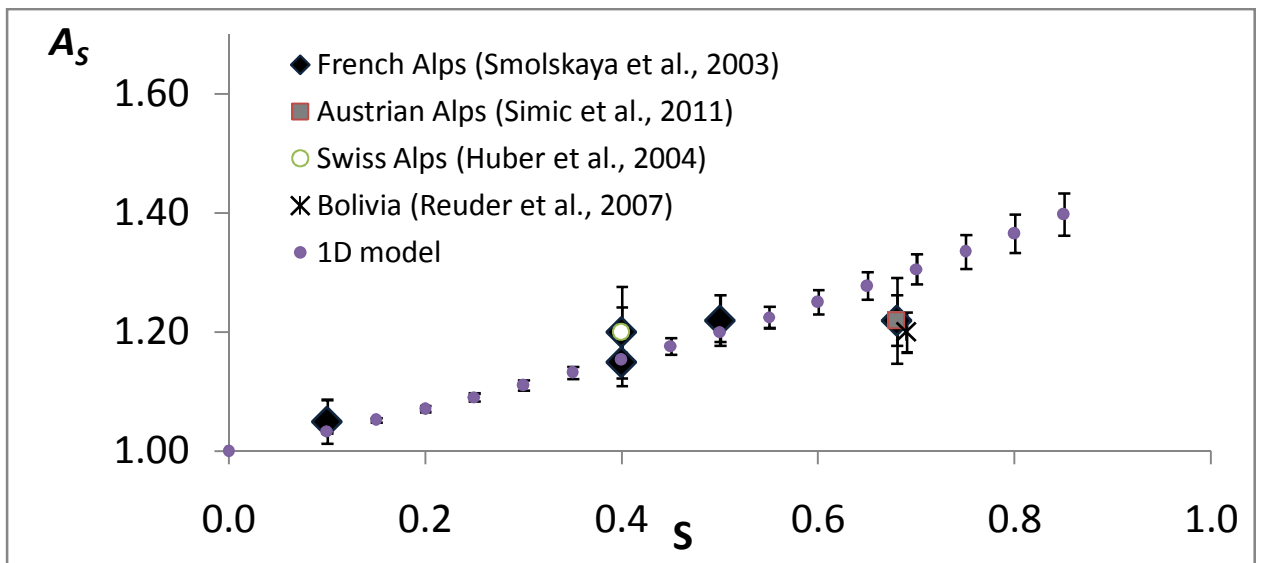
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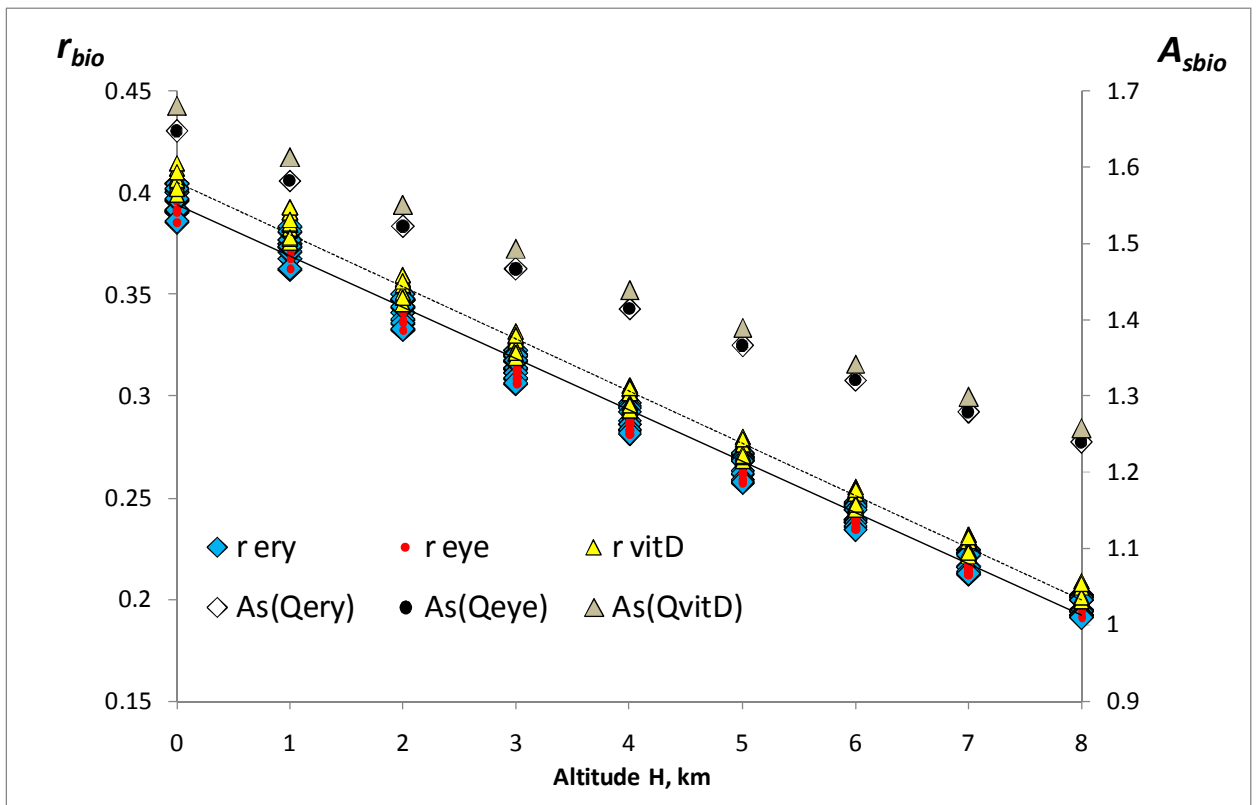
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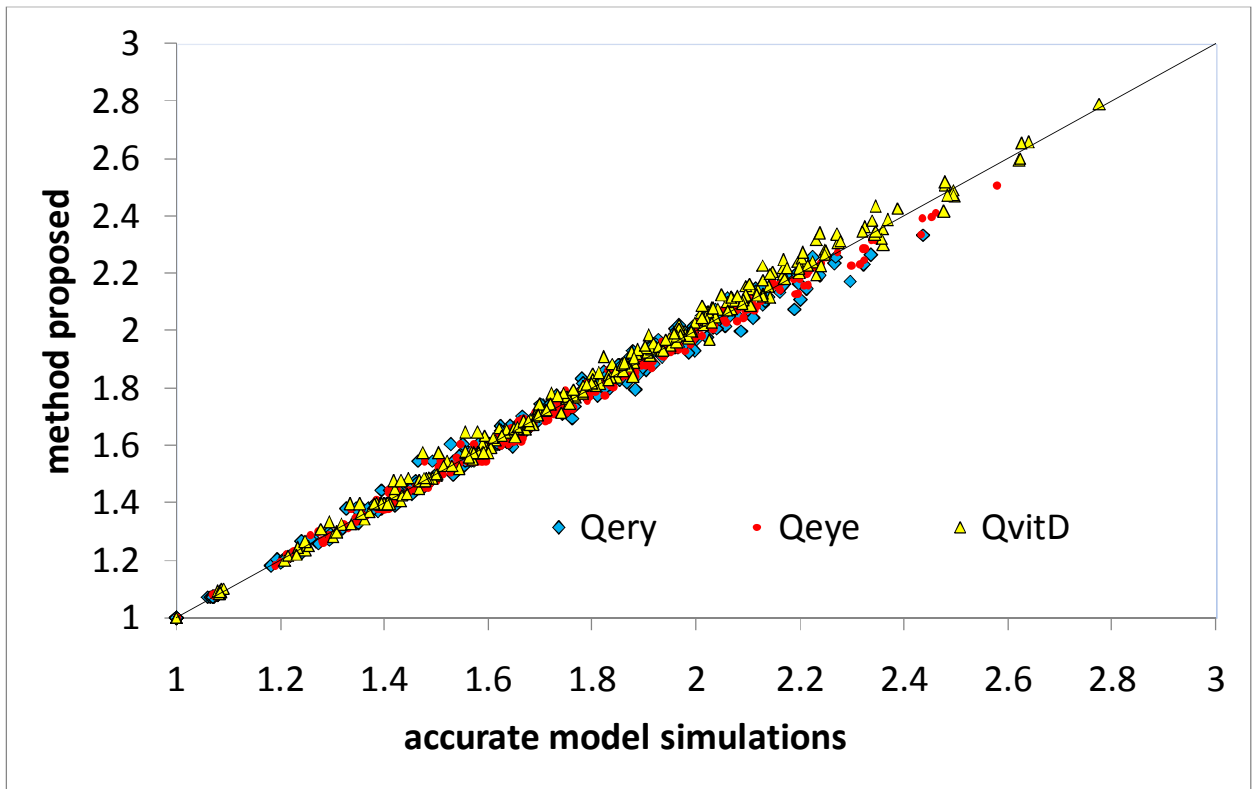
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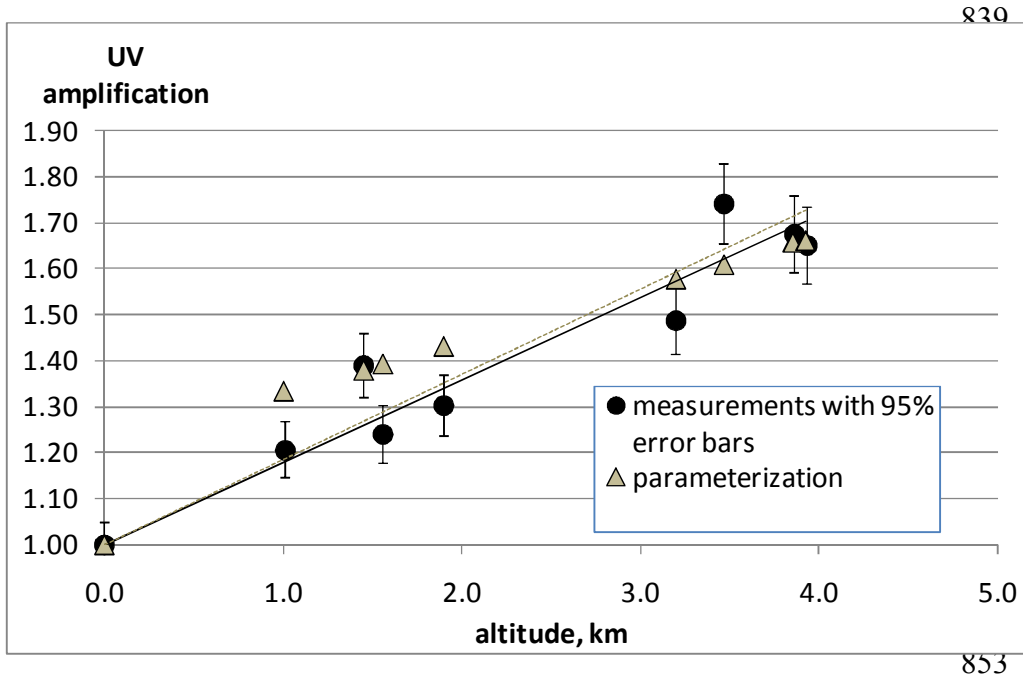
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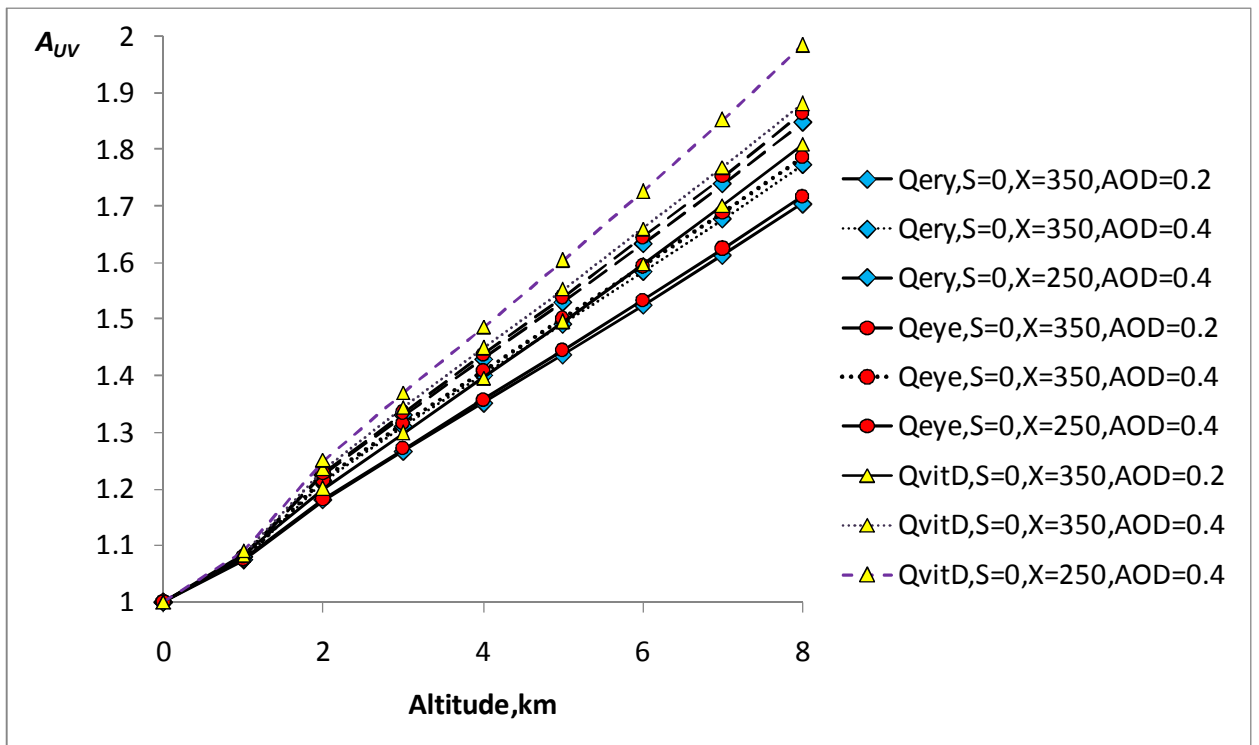
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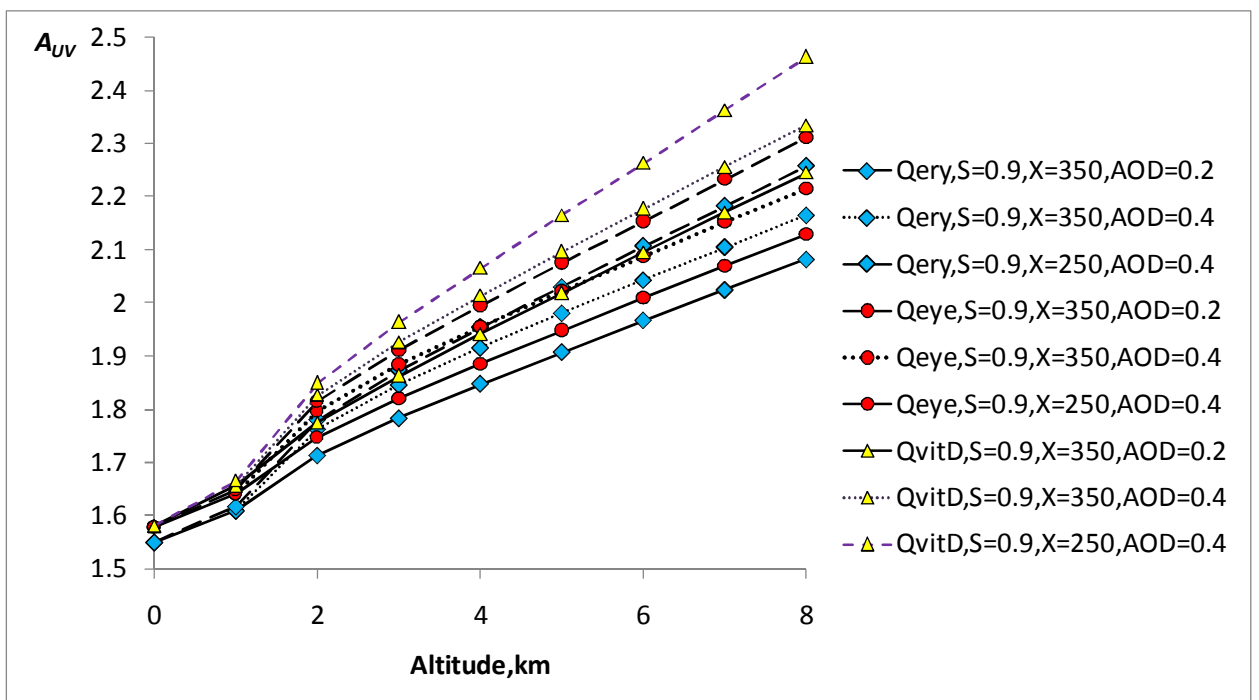
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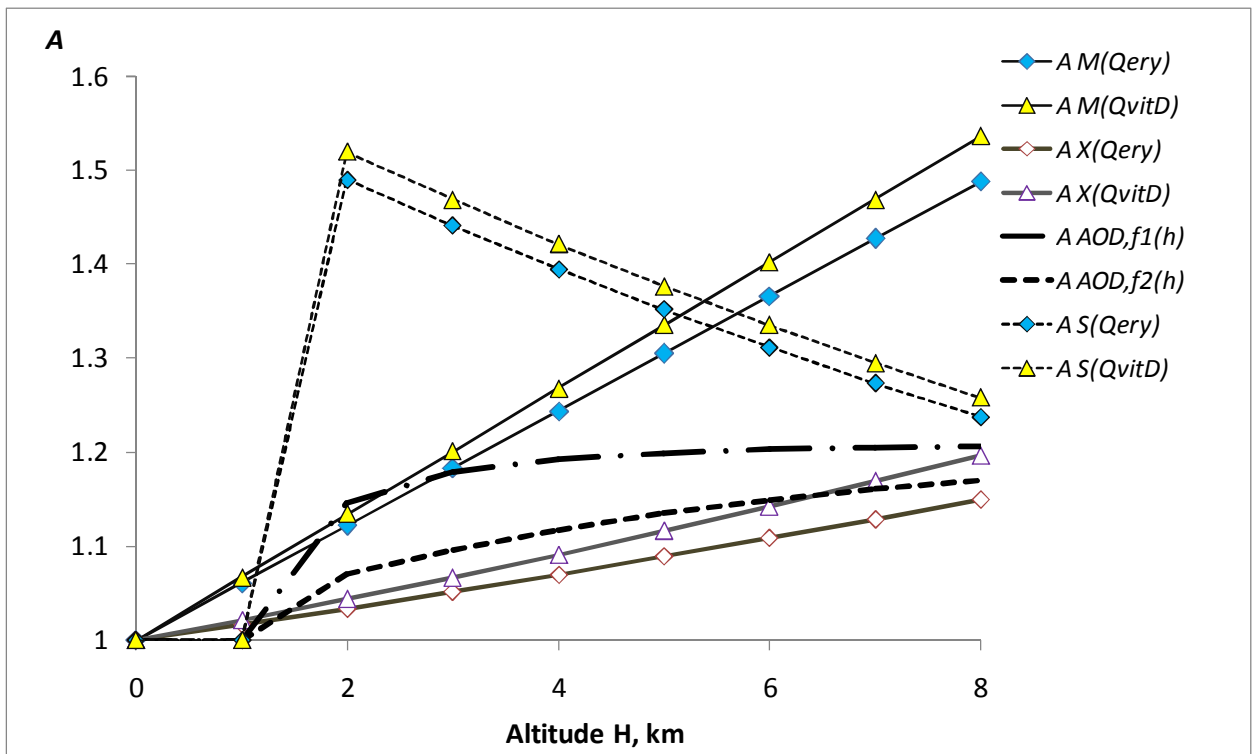
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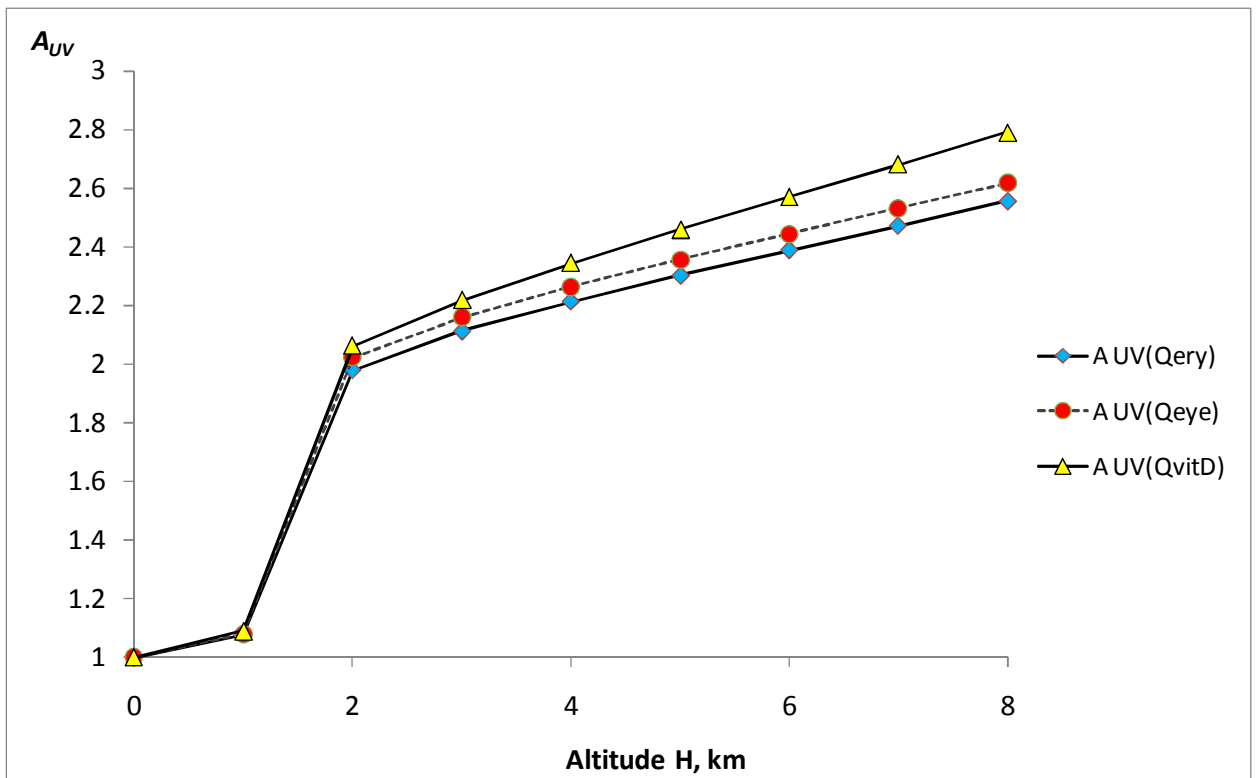
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881 Figure 10. The UV amplification due to molecular  $A_M(Qery)$ ,  $A_M(QvitD)$ , ozone  $A_X(QvitD)$ ,  $A_X(Qery)$ , aerosol  $A_{AOD,f1(H)}$ ,  
 882  $A_{AOD,f2(H)}$  for the Alpine  $f1(H)$  and Asian  $f2(H)$  types of altitude dependences, and surface albedo  $A_S(Qery)$ ,  $A_S(QvitD)$   
 883 changes with the altitude (a) and their total altitude effect on  $A_{UV}$  for different types of BAUVR (b). At  $H=0$  km:  
 884  $AOD_{340}=0.8$ ,  $X=250$  DU. The surface albedo has an abrupt change at 2 km from  $S=0$  to  $S=0.95$ . Solar elevation -  $h=50^\circ$ .