

# **“New insights into the atmospheric mercury cycling in Central Antarctica and implications at a continental scale” by H. Angot et al.**

## **Response to referee comments by Referee #1.**

We thank this anonymous referee for insightful questions and comments. We provide below a point-by-point reply to the comments (points raised by the referee in bold, changes made in the manuscript in red).

### 1. General comments

**The manuscript is well written and presents a novel dataset of atmospheric mercury and mercury in snow interstitial air covering slightly more than one annual cycle. Mercury measurements from Antarctica is generally scarce in particular good data covering more than the summer season, so this manuscript will definitely be a longed for addition to the pool of mercury data for the scientific community, both for experimentalists and modelers. The references used are recent and relevant. The structure of the result and discussion chapter is complicated. The text jumps between different environmental compartments (atmosphere and snow interstitial air) and seasons, which makes it ponderous to follow the discussion. The authors should consider re-arranging the sections within the results and discussion chapter.**

We agree with the referee regarding the structure of the results and discussion chapter. In the revised manuscript we have changed the structure as follows in order to avoid jumps between different environmental compartments (atmosphere and snow interstitial air) and seasons.

### 3. Results and Discussion

#### 3.1 Hg(0) concentrations in ambient air

##### 3.1.1 Spring

##### 3.1.2 Summertime

- a) Oxidation of Hg(0) in ambient air and Hg(II) deposition onto snowpack
- b) Multi-day depletion events of atmospheric Hg(0)
- c) Hg(0) diurnal cycle

##### 3.1.3 Fall

##### 3.1.4 Winter

#### 3.2 Hg(0)/Hg(II) redox conversions within the snowpack

##### 3.2.1 Sunlit period

##### 3.2.2 Winter

### 4. Implications at a continental scale

## 2. Scientific comments

**- Line 65-70: The manuscript text says that the Antarctic plateau was first considered to be chemically inactive for atmospheric species including Hg and a paper from 2008 is cited. The manuscript text further says that it turned out to be highly active, now citing a paper from 2001 and 2007. To me there is a lack of logic in this argument.**

We agree. In the revised manuscript we now refer to an older reference:

“The Antarctic plateau (...) was first considered to be chemically-inactive and a giant cold trap for atmospheric species, ~~including mercury~~ (e.g., Lambert et al., 1990). It turned out to be highly photochemically active (Davis et al., 2001; Grannas et al., 2007)...”

**- Line 131: Was mercury saturated manually? I assume you mean manually injecting saturated mercury.**

Yes indeed. This has been corrected in the revised manuscript.

**- Line 144-146: QA/QC; Interesting that you use internal standard on the 2600, what was used as internal standard? Why not use commercially available control samples?**

These questions have been addressed in the revised manuscript:

“The instrument was calibrated with the NIST SRM-3133 mercury standard. Quality assurance and quality control included the analysis of analytical blanks, replicates, and internal standards (Reference Waters for mercury: HG102-2 at 22 ng/L from Environment Canada).”

**- Line 202-204: This was an elegant way of defining seasons!**

Thank you.

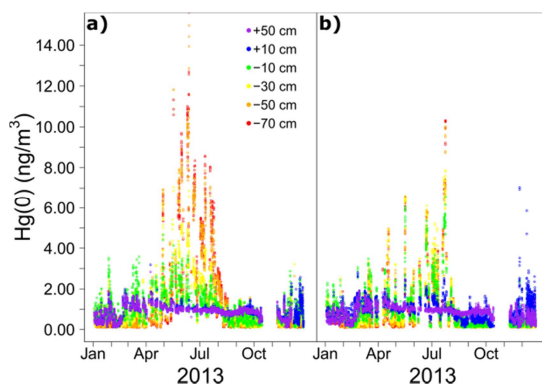
**- Line 211-212: Troll can hardly be called a coastal site as it is situated almost 250 km from the coast and at almost 1500 masl. However, due to its location Troll experiences air from both the Antarctic plateau and the southern Ocean, but it is not a coastal site.**

We agree that Troll can hardly be called a coastal site. This has been corrected in the revised manuscript:

“(...) Hg(0) concentrations are lower than annual averages reported at **near-coastal** or coastal Antarctic stations...”

**- Line 259-263: I like the figure showing the vertical distribution of Hg(0) concentrations, however I think it could also be interesting for the audience to present a figure showing time series of Hg(0) in SIA as well. The atmospheric Hg(0) time series seems well covered in other figures.**

A Figure displaying the annual variation of Hg(0) concentrations in the snow interstitial air collected at the various inlets of the two snow towers has been added in the revised manuscript:



“Figure 10: Annual variation of hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) in the snow interstitial air collected at the various inlets of the two snow towers: **a)** snow tower #1, **b)** snow tower #2. Note that we regularly experienced technical problems on snow tower #2 leading to missing values.”

**- Line 356 and throughout the paragraph: I am not quite sure but I believe it is general consensus regarding the use of the term “depletion event” being oxidation of Hg resulting from bromine photochemistry and high correlation with tropospheric ozone depletions. The multi-day low concentrations described in chapter 3.5 are a different mechanism and should consequently be called something else.**

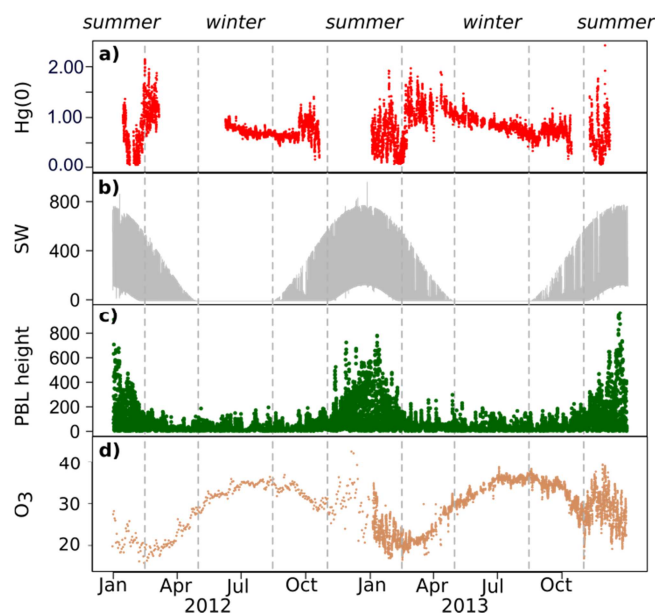
We are not sure that renaming a depletion of mercury specific to this site is warranted. A “depletion event” is just a depletion of Hg(0), and is not specific to the halogen chemistry. We do understand that some may associate it but “depletion event” is a more generic term than that.

**- Line 391 and throughout chapter 3.6: A mechanism for reduced Hg(0) during winter is proposed, and it is also mentioned that this reductions are not observed at Troll or Neumayer. Any thought on why this reaction mechanism does not occur at Troll or Neumayer? Do you believe this reaction mechanism occur throughout the Antarctic plateau?**

The reason why this reaction mechanism does not occur at Troll or Neumayer is unclear. This could be due to meteorological conditions on the Antarctic plateau (e.g., temperature, relative humidity, boundary layer dynamics). Further research is clearly needed. Our sense is that this reaction mechanism might occur throughout the Antarctic plateau, but the spatial distribution of Hg(0) measurements should be improved.

**- Additionally, as you have O<sub>3</sub> measurements it would be interesting to have them presented in fig 12 alongside Hg(0) since O<sub>3</sub> is suggested to be involved in the Hg(0) wintertime decrease.**

Figure 12 shows that – despite the overall decreasing trend in winter – Hg(0) concentration exhibited abrupt increases when moist and warm air masses from lower latitudes occasionally reached Concordia Station. Key parameters are therefore Hg(0), integrated water vapor, and temperature. We do not think that adding O<sub>3</sub> measurements here would be relevant. However, the annual variation of O<sub>3</sub> measurements in 2012 and 2013 has been added in Figure 3 of the revised manuscript.



“Figure 3: Annual variation in 2012 and 2013 of a) hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) at 500 cm and 25 cm above the snow surface in 2012 and 2013, respectively, b) downwelling shortwave (SW) radiation (in  $\text{W/m}^2$ ), c) planetary boundary layer (PBL) height (in m), and d) ozone ( $\text{O}_3$ , daily mean in 2012 and hourly mean in 2013) mixing ratios (in ppbv). The vertical dashed lines represent seasonal boundaries.”

- The authors should perhaps also have a look at a very recent paper by Nerentorp Mastromonaco et al., 2016 (Atmospheric Environment) where winter depletions are present and to check whether this may have any relevance for the winter decreasing trend observed in this study.

Thank you for the suggestion. Nerentorp Mastromonaco et al. (2016) observed Atmospheric Mercury Depletion Events (AMDEs) during Antarctic winter over sea ice areas and proposed a dark mechanism in the marine boundary layer. It should be noted that such events have never been reported on the Antarctic continent (coastal or inland stations). Given the distance of Concordia station from sea ice areas a similar dark mechanism seems unlikely.

### 3. Technical comments

- Line 412: Several characters in the equation disappear in print

Thanks for noticing that. We will make sure there is no such problem in the final version of the manuscript.

- Figure 6: It is difficult to tell the difference between summer and spring/fall colours in the figure and I would very much like to see something similar to box and whisker plots to be able to tell the concentration distribution at each height at each season.

The color for spring/fall values has been changed in the revised manuscript. It should now be easier to tell the difference between summer and spring/fall colors. The concentration distribution at each height at each season can be inferred from the Figure we have added in the revised manuscript (see comment line 259-263).

**- Figure 9: Hg(0) concentrations, which measurement height do the results represent?**

This has been added in the caption of the revised manuscript:

“Figure 7: Hourly (local time) mean variation, along with the 95% confidence interval for the mean, of: **a)** Hg(0) concentration (in  $\text{ng/m}^3$ ) **at 25 cm above the snow surface**, **b)** downwelling shortwave (SW) radiation (in  $\text{W/m}^2$ ) according to the MAR model simulations, **c)** temperature (in  $^\circ\text{C}$ ) at 3 m above the snow surface, **d)** wind speed at 3 m above the snow surface (in m/s), **e)** planetary boundary layer (PBL) height (in m) according to the MAR model simulations, **f)** friction velocity ( $u_*$ , in m/s), and **g)** Eddy diffusivity ( $K$ , in  $\text{m}^2/\text{s}$ ) in summer (red), fall (green), winter (blue), and spring (purple). **Note that the hourly mean variation of Hg(0) concentration in summer is similar at the three inlets of the meteorological tower**”.

**- Figure 11: Same comment as above, please indicate at which height the Hg(0) measurements are from.**

This has been added in the caption of the revised manuscript:

“Figure 6: Top: January and February 2012 cycle of: **a)** hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) **at 500 cm above the snow surface**, **b)** Integrated Water Vapor (IWV,  $\text{kg/m}^2$ ), **c)** Temperature (in  $^\circ\text{C}$ ) at 10 m above ground level, and **d)** ozone ( $\text{O}_3$ , daily mean) mixing ratios (ppbv). Hg(0) was low from 19 January to 8 February (period highlighted in red) while  $\text{O}_3$  showed no abnormal variability. Bottom: January and February 2013 cycle of: **e)** hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) **at 210 cm above the snow surface**, **f)** Integrated Water Vapor (IWV,  $\text{kg/m}^2$ ), **g)** Temperature (in  $^\circ\text{C}$ ) at 10 m above ground level, and **h)** ozone ( $\text{O}_3$ ) mixing ratio (ppbv). Hg(0), IWV, and temperature were low from 5 to 20 February (period highlighted in red) while  $\text{O}_3$  showed no abnormal variability. **Note that Hg(0) concentrations exhibited the same pattern at the three inlets of the meteorological tower from 5 to 20 February 2013**”.

**- Figure 12: same comment as above.**

This has been added in the caption of the revised manuscript:

“Figure 9: Year 2012 wintertime record of: **a)** hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) **at 500 cm above the snow surface**, **b)** Integrated Water Vapor (IWV,  $\text{kg/m}^2$ ), and **c)** Temperature ( $T$ ,  $^\circ\text{C}$ ) at 10 m above ground level. Hg(0), temperature, and IWV increased from June 12 to 15 (in red) suggesting transport of moist and warm air masses originating from lower latitudes.”

#### 4. References

Davis, D., Nowak, J. B., Chen, G., Buhr, M., Arimoto, R., Hogan, A., Eisele, F., Mauldin, L., Tanner, D., Shetter, R., Lefer, B., and McMurry, P.: Unexpected high levels of NO observed at South Pole, Geophysical research letters, 28, 3625-3628, 2001.

Grannas, A. M., Jones, A. E., Dibb, J., Ammann, M., Anastasio, C., Beine, H. J., Bergin, M., Bottenheim, J., Boxe, C. S., Carver, G., Chen, G., Crawford, J. H., Domine, F., Frey, M. M., Guzman, M. I., Heard, D. E., Helmig, D., Hoffmann, M. R., Honrath, R. E., Huey, L. G., Hutterli, M., Jacobi, H.-W., Klan, P., Lefer, B., McConnell, J. R., Plane, J. M. C., Sander, R., Savarino, J., Shepson, P. B., Simpson, W. R., Sodeau, J., Von Glasow, R., Weller, R., Wolff, E. W., and Zhu, T.: An overview of snow photochemistry: evidence, mechanisms and impacts, *Atmospheric Chemistry and Physics*, 7, 4329-4373, 2007.

Lambert, G., Ardouin, B., and Sanak, J.: Atmospheric transport of trace elements toward Antarctica, *Tellus B*, 42, 76-82, 10.1034/j.1600-0889.1990.00009.x, 1990.

Nerentorp Mastromonaco, M., Gårdfeldt, K., Jourdain, B., Abrahamsson, K., Granfors, A., Ahnoff, M., Dommergue, A., Méjean, G., and Jacobi, H.-W.: Antarctic winter mercury and ozone depletion events over sea ice, *Atmospheric Environment*, 129, 125-132, 2016.

## **“New insights into the atmospheric mercury cycling in Central Antarctica and implications at a continental scale” by H. Angot et al.**

### **Response to referee comments by Referee #2.**

We would like to thank the anonymous referee for its time and useful comments towards the improvement of our manuscript. We provide below a point-by-point reply to the comments (points raised by the referee in bold, changes made in the manuscript in red).

#### 1. General comments

**This paper describes year-round measurements of Hg(0) in the atmosphere and snow on the Antarctic plateau along with ancillary measurements. These comprise a novel data set that is very valuable for understanding the global atmospheric (and cryospheric) mercury cycle. Given the value of these data and the difficulty in duplicating them, I would encourage the authors to make the complete data set available in some capacity (e.g. as a supplementary file, or a link to a data repository) in order to aid modellers, etc., in using these measurements to advance further research.**

Mercury data reported in this paper are available upon request at [http://sdi.iaa.cnr.it/geoint/publicpage/GMOS/gmos\\_historical.zul](http://sdi.iaa.cnr.it/geoint/publicpage/GMOS/gmos_historical.zul). This has been added in the acknowledgments of the revised manuscript.

**The analysis and interpretation is largely sound, with a few gaps as identified in the comments below. I do agree with the first reviewer that the organization of the Results and Discussion could be improved. I recommend the publication of this paper after the minor issues discussed below have been addressed.**

We agree with the referee regarding the structure of the results and discussion chapter. In the revised manuscript we have changed the structure as follows in order to avoid jumps between different environmental compartments (atmosphere and snow interstitial air) and seasons.

### 3. Results and Discussion

#### 3.1 Hg(0) concentrations in ambient air

##### 3.1.1 Spring

##### 3.1.2 Summertime

- a) Oxidation of Hg(0) in ambient air and Hg(II) deposition onto snowpack
- b) Multi-day depletion events of atmospheric Hg(0)
- c) Hg(0) diurnal cycle

##### 3.1.3 Fall

##### 3.1.4 Winter

#### 3.2 Hg(0)/Hg(II) redox conversions within the snowpack

### 3.2.1 Sunlit period

### 3.2.2 Winter

## 4. Implications at a continental scale

### 2. Specific comments

**- I. 38-39: “according to observations at coastal Antarctic stations” is vague; this is used elsewhere in the paper (l. 488) and is not very enlightening. Can you summarize the evidence you are using to draw this conclusion? Perhaps in Section 3.7? There are references there but the observations are not described.**

“According to observations at coastal Antarctic stations” refers to the following sentence in Section 3.7 (Section 4 in the revised version of the manuscript, see previous comment regarding the structure of the Results and Discussion chapter):

“Conversely, low Hg(0) concentrations that were not correlated or anti-correlated with O<sub>3</sub> were observed at Neumayer and Troll (Temme et al., 2003; Pfaffhuber et al., 2012), while elevated Hg(II) concentrations (up to 0.33 ng/m<sup>3</sup>) were recorded at Terra Nova Bay in the absence of Hg(0)/O<sub>3</sub> depletion (Sprovieri et al., 2002)”.

In an attempt to clarify this point, we have added (1) a reference to this sentence in lines 482-484 of the revised manuscript and (2) a reference to this Section in the conclusion:

(1) “but can be sporadically observed elsewhere explaining the **above-mentioned** observations at Neumayer, Troll, or Terra Nova Bay (~~Temme et al., 2003; Sprovieri et al., 2002; Pfaffhuber et al., 2012~~)”.

(2) “According to observations at coastal Antarctic stations (**see section 4**), the reactivity observed at Concordia Station can be transported at a continental scale by strong katabatic winds”.

**- I. 53: “rapid deposition” is relative. You later describe a reservoir of gaseous Hg(II), which can hardly be expected if the deposition lifetime is very rapid.**

We agree. This has been corrected in the revised manuscript: “leading to the formation and subsequent **rapid** deposition”

**- Section 2.3: What is the estimated precision of your Hg(0) measurements?**

This has been specified in the revised version of the manuscript: “**Based on experimental evidence, the average systematic uncertainty for Hg(0) measurements is of ~ 10 % (Slemr et al., 2015)**”.

**- Section 3.1: (a) This is titled “seasonal variation” but mostly summarizes annual values and spatial/vertical differences.**



This title has been removed in the revised manuscript (see previous comment regarding the structure of the Results and Discussion chapter).

**(b) Can you identify what the  $\pm$  values are? Standard deviation? Confidence limits on the mean?**

$\pm$  values refer to standard deviations. This has been clarified in the revised manuscript: “In summer, the mean atmospheric Hg(0) concentration was  $0.69 \pm 0.35$  ng/m<sup>3</sup> (mean  $\pm$  standard deviation).”

**(c) What statistical test was used to determine that your values were lower than the Troll and Neumayer – Mann-Whitney as well?**

None. This would require the entire distribution of Hg(0) concentrations at Troll and Neumayer. A two-sample z-test for comparing two means could be used. However, this test assumes that the two populations are normally distributed.

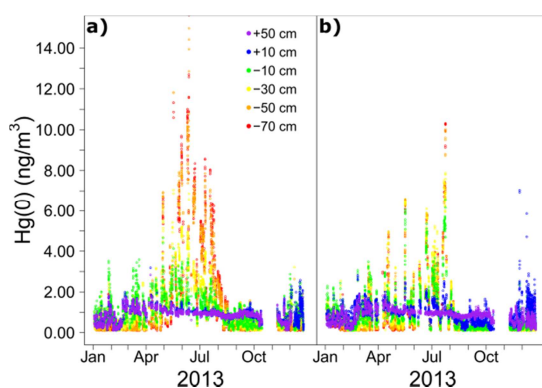
**- How did the 25 cm inlet met tower values compare with the 50 cm and 10 cm snow tower inlet values overall? Fig. 6 suggests there was some offset between the met tower and snow tower inlets, at least in winter and possible spring/fall. Is the sampling coverage the same? Could there be an effect of heated/non-heated lines, or the length of the sampling lines?**

Overall, the 25 cm inlet met tower values are lower than the 50 and the 10 cm snow tower inlet values. Values at the 50 and 10 cm snow tower inlets might be biased high due to contamination from the deeper inlets. Indeed, as mentioned in section 2.2 of the manuscript, sampling lines were purged continuously at 5 L/min on the met tower but intermittently at  $\sim$  2-3 L/min on the two snow towers.

**- Fig. 3: Why did you choose the 25 cm inlet to show? Can you add a time series or two (shallow/deep) for the snowpack data?**

The decision to show Hg(0) at the 25 cm inlet is arbitrary. There is little variation of Hg(0) with height on the meteorological tower. Displaying all of the Hg(0) data (i.e., at the three inlets of the meteorological tower) makes this Figure difficult to read.

A Figure displaying the annual variation of Hg(0) concentrations in the snow interstitial air collected at the various inlets of the two snow towers has been added in the revised manuscript:



“Figure 10: Annual variation of hourly-averaged Hg(0) concentrations (in ng/m<sup>3</sup>) in the snow interstitial air collected at the various inlets of the two snow towers: a) snow tower #1, b) snow tower #2. Note that we regularly experienced technical problems on snow tower #2 leading to missing values.”

- Fig. 4: I only see error bars on a few points. Are these the only ones with replicates? How many replicates were done in those cases? A line or two in the caption to explain this would be helpful.

All samples were analyzed in replicates of three. Standard error is most of the time smaller than the width of the dot explaining why you only see error bars on a few points. This has been clarified in the caption of the revised manuscript:

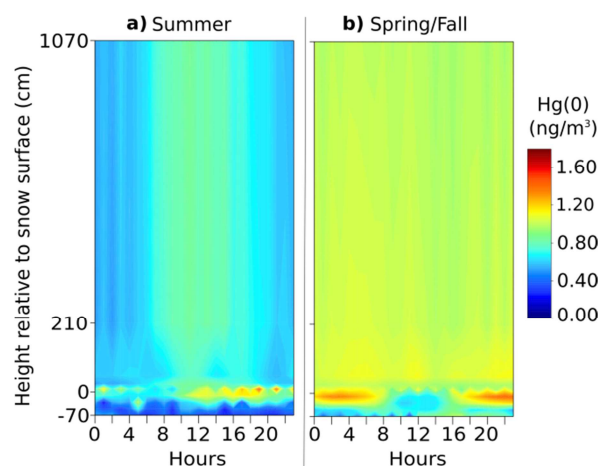
“Figure 5: Total mercury concentration (ng/L), along with standard errors, in surface snow samples collected weekly at Concordia Station from February 2013 to January 2014. Dark period (winter) highlighted in grey, sunlit period highlighted in yellow. Total mercury concentrations were elevated (up to 74 ng/L) in November-December 2013 (summer). All samples were analyzed in replicates of three. Standard errors are frequently smaller than the width of the dots.”

- Section 3.2: You refer to “continuous” oxidation in the summer. What do you mean by that? It is clear there is net oxidation, but I am not sure you have shown it is continuous.

Yes, indeed. The term “continuous oxidation” has been removed throughout the revised manuscript.

- Fig. 7: This figure is a nice attempt to visualize the data, but it is rather confusing. Does the shading represent some sort of unspecified interpolation? Do the top boxes represent the met tower and the bottom the snow tower? In that case, why do the top boxes extend down below zero and the bottom ones not go up to 50 cm? If not, why don't the top and bottom agree within the overlap region?

We agree that this figure, as it is, might be confusing. It has been modified in the revised version of the manuscript:



“Figure 12: Hourly (local time) mean atmospheric and interstitial air Hg(0) concentrations in a) summer, and b) spring/fall. The vertical axis is the height of measurement relative to the snow surface (in cm). Color contours show Hg(0) concentrations (in ng/m<sup>3</sup>). Concentrations at 25, 210, and 1070 cm above the snow surface were acquired on the meteorological tower while concentrations at 50, 10, -10, -30, -50 cm, and -70 cm were collected on snow tower #1. Data were cubic spline interpolated using software R.”

- **Fig. 8: I think this figure is not crucial to the paper, since it is only being used to support a suggested mechanism for a single extreme value. I think you can make that suggestion without an additional figure, though it is up to your discretion.**

We agree. This Figure has been removed in the revised manuscript.

- **Section 3.3.2: You mention a shift from oxidation to reduction at the beginning of winter, but it would be very helpful to see the time series of Hg(0) at depth (as mentioned above) – is it a sudden drop to a stable “winter” value, or is there a longer trend over the winter to accompany the atmospheric decline?**

As mentioned above, a Figure displaying the annual variation of Hg(0) concentrations in the snow interstitial air collected at the various inlets of the two snow towers has been added in the revised manuscript. It is a sudden drop and not a longer trend over the winter to accompany the atmospheric decline.

- **Section 3.4.1: Your summary is a bit confusing (ll. 339-343). I think you are saying by “continuous” that (i) has a week diurnal cycle and by “important” that (iii) has a strong diurnal cycle, resulting in the observed concentration pattern. Can you say this more clearly? I’m not sure what “important” means in (ii). Clearly it is important to the surface snow THg, but I’m not sure how this is related to the Hg(0) diurnal cycle.**

The summary has been modified in the revised manuscript:

“In summary, the observed summertime Hg(0) diurnal cycle in ambient air might be due to a combination of factors: i) ~~a continuous~~ **the intense** oxidation of Hg(0) in ambient air due to the high oxidative capacity on the plateau – **as evidenced by low mean Hg(0) concentrations (see section 3.2.1.a)**, ii) ~~important~~ subsequent Hg(II) deposition onto snowpack – **as evidenced by elevated total mercury levels in surface snow samples (see section 3.2.1.a)**, and iii) ~~important~~ emission of Hg(0) from the snowpack during convective hours. **Fig. 8 summarizes the processes that govern mercury exchange at the air/snow interface. Redox processes occurring within the snowpack are discussed in details in section 3.2.**”

- **Section 3.4.2: Where is this Hg(0) building up from? Presumably the snow, but it’s not mentioned. Why is the fall concentration higher than spring? The reservoir of Hg in the summer snow?**

Yes indeed, we believe that Hg(0) is building up due to emissions from the snowpack. This has been clarified in the revised manuscript: “We believe that the shallow boundary layer could cause Hg(0) concentrations in ambient air to build up to where they exceeded levels

recorded at lower latitudes in the Southern Hemisphere because Hg(0) – emitted from the snowpack – was dispersed into a reduced volume of air, limiting the dilution”.

The fall concentration is higher than spring indeed likely due to the reservoir of mercury in the summer snow. It should be noted that in both 2012 and 2013 mercury depletion events occurred at the end of summer likely leading to Hg(II) deposition on the snowpack.

**- Fig. 9g: What inlet is the Hg(0) cycle from?**

This has been added in the caption of the revised manuscript:

“Figure 7: Hourly (local time) mean variation, along with the 95% confidence interval for the mean, of: **a)** Hg(0) concentration (in ng/m<sup>3</sup>) at 25 cm above the snow surface, **b)** downwelling shortwave (SW) radiation (in W/m<sup>2</sup>) according to the MAR model simulations, **c)** temperature (in °C) at 3 m above the snow surface, **d)** wind speed at 3 m above the snow surface (in m/s), **e)** planetary boundary layer (PBL) height (in m) according to the MAR model simulations, **f)** friction velocity ( $u_*$ , in m/s), and **g)** Eddy diffusivity ( $K$ , in m<sup>2</sup>/s) in summer (red), fall (green), winter (blue), and spring (purple). Note that the hourly mean variation of Hg(0) concentration in summer is similar at the three inlets of the meteorological tower”.

**- Section 3.5: It’s a bit odd to refer to Fig. 10e first. I suggest you rearrange the figure to make this 10a.**

We agree. The Figure has been rearranged in the revised manuscript.

**- Section 3.6: (a) Given you have a single winter of data, and the decline is not seen at the other stations, can you eliminate instrument drift (e.g. trap poisoning)? Were the external calibrations before and after the winter consistent?**

The instrument failed in 2014 but the decreasing trend has also been observed at Concordia Station in 2015. The data are presented in a paper that will soon be submitted in *Atmospheric Chemistry and Physics* (Angot et al., in preparation). Additionally, the decreasing trend has also been observed at Dumont d’Urville and data are presented in a companion paper (Angot et al., 2016). These results give us confidence that the decline is not due to an instrument drift.

**(b) Why do you not include dry deposition of Hg(0) as a possible mechanism for this decrease? Given the low BL, what deposition flux would be needed to remove the observed amount of Hg(0)? How does this compare to other observations/calculations (Cobbett et al 2007, Zhang et al 2009)? I think it’s quite similar. It would also account for the gradient in the decrease (3.6.2).**

You are absolutely right. Dry deposition of Hg(0) is a possible mechanism for this decrease. This has been added in the revised manuscript:

“The observed declining trend could also be attributed to the dry deposition of Hg(0) onto the snowpack. The dry deposition velocity is defined as follows (Joffe, 1988), as the ratio between the deposition flux  $F$  (ng/m<sup>2</sup>/s) and the concentration  $C$  (ng/m<sup>3</sup>):

$$v_d = \frac{F}{C} \quad (2)$$

Denoting the height of the boundary layer  $h$  and the Hg(0) concentration at the beginning of winter  $C_0$ , the evolution of the concentration versus time is thus given by the following ordinary differential equation:

$$C = C_0 e^{-(v_d/h)t} \quad (3)$$

During winter ( $t = 107$  days), the Hg(0) concentration gradually decreased from  $C_0 \sim 1.03$  ng/m<sup>3</sup> to  $C \sim 0.73$  ng/m<sup>3</sup> at 25 cm above the snowpack, in a mixing layer of 25 m high. According to Equation (3) the associated dry deposition velocity is  $9.3 \cdot 10^{-5}$  cm/s. This result is in very good agreement with dry deposition velocities reported for Hg(0) over snow (Cobbett et al., 2007; Zhang et al., 2009)."

**(c) Speaking of which, you don't report in 3.1 if there are any seasonal differences in the three met tower inlets. Your discussion of the winter data suggests there would be.**

This has been added in the revised manuscript:

Lines 209-213: "No significant difference was observed between annual averages of Hg(0) concentrations measured at the three inlets of the meteorological tower in 2013 ( $p$  value =  $3.1 \cdot 10^{-14}$ , Mann-Whitney test). **It should be noted that Hg(0) concentrations at the three inlets were significantly different in winter only (see section 3.1.4).**"

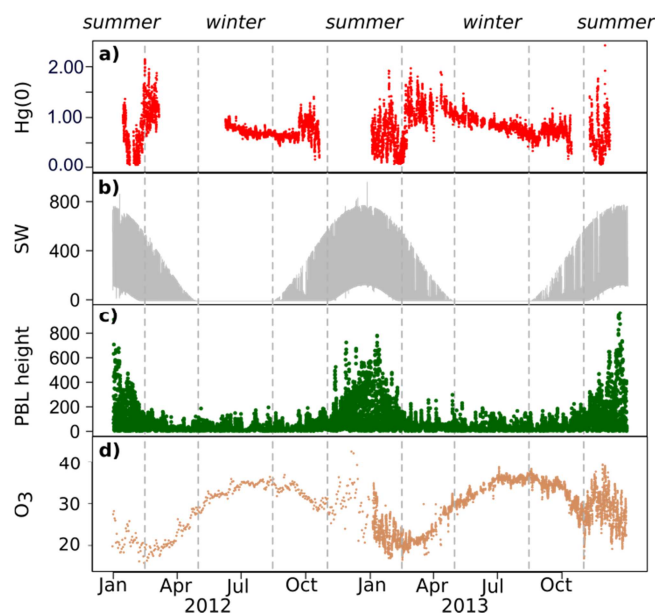
Lines 372-376: "In 2013, the height of measurement had a significant influence on the decline over time of Hg(0) concentrations (ANCOVA test,  $p$  value  $< 0.05$ ), with a steeper decrease at 25 cm than at 1070 cm. **Additionally, wintertime Hg(0) concentrations were significantly lower at 25 cm than at 1070 cm ( $p$  value  $< 0.05$ , Mann-Whitney test).**"

**(d) This section is poorly organized. I suggest removing the sub-sections since you basically discount the gas-phase reaction without doing so explicitly.**

Sub-sections have been removed in the revised version of the manuscript.

**(e) l. 402: Why don't you report your O<sub>3</sub> data instead of (or as well as) referring to another paper? Also, are there O<sub>3</sub> data at Troll or Neumayer that suggest that a winter reaction with O<sub>3</sub> would not also happen there?**

The annual variation of O<sub>3</sub> measurements in 2012 and 2013 has been added in Figure 3 of the revised manuscript:



“Figure 3: Annual variation in 2012 and 2013 of a) hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) at 500 cm and 25 cm above the snow surface in 2012 and 2013, respectively, b) downwelling shortwave (SW) radiation (in  $\text{W/m}^2$ ), c) planetary boundary layer (PBL) height (in m), and d) ozone ( $\text{O}_3$ , daily mean in 2012 and hourly mean in 2013) mixing ratios (in ppbv). The vertical dashed lines represent seasonal boundaries.”

The reason why the Hg(0) decline throughout winter does not occur at Troll or Neumayer is unclear. This could be due to meteorological conditions on the Antarctic plateau (e.g., temperature, relative humidity, boundary layer dynamics). Further research is clearly needed.

**- Section 3.7: A bit more detail about the observations that are attributed to transport from the plateau (ll. 457-462) would be helpful, as mentioned above.**

See response above.

**- Section 4: (a) I’d like to see a mention of the intriguing winter subsurface Hg(0) peak in here.**

This has been added in the conclusion of the revised manuscript:

“Additionally, Hg(0) concentrations increased with depth in the snow interstitial air in winter likely due to a dark reduction of Hg(II) species accumulated within the snowpack during the sunlit period.”

**(b) Maybe change “heating” to “snowpack ventilation” or “ventilation and heating” in l. 479.**

This has been changed in the revised manuscript: “Summertime Hg(0) concentration in ambient air exhibited a pronounced diurnal cycle likely due to large emissions from the snowpack as a response to daytime snowpack ventilation.”

**(c) In l. 481 I would change “likely” to “possibly”.... And do you really think gas-phase oxidation is even that likely? Your earlier discussion suggests not. You may need to add dry deposition of Hg(0) as well, depending what you find.**

Indeed, the decreasing trend observed in winter is most likely due to the dry deposition of Hg(0). This has been changed in the revised manuscript (abstract and conclusion).

### 3. Specific comments

**l. 32: change “never been observed” to “not been reported”**

Done.

**l. 46: change “contamination” to “contaminant”**

Done.

**l. 246: change “bound” to “bind”**

Done.

**l. 315: change “significant and daily” to “significant daily”**

Done.

**l. 316: change “all along” to “throughout”**

Done.

**l. 358: suggest changing “explosions” to “so-called ‘bromine explosions’” to avoid leaving unfamiliar readers with the impression there are actual explosions.**

We agree. This has been corrected in the revised manuscript.

**l. 378: change “these depletions of Hg(0)” to “the depletions of Hg(0) reported here”**

Done.

### 4. References

Angot, H., Dion, I., Vogel, N., Magand, O., Legrand, M., and Dommergue, A.: Atmospheric mercury record at Dumont d'Urville, East Antarctic coast: continental outflow and oceanic influences, *Atmospheric Chemistry and Physics Discussions*, 10.5194/acp-2016-257, in review, 2016.

Angot, H., Dastoor, A., De Simone, F., Gårdfeldt, K., Gencarelli, C. N., Hedgecock, I. M., Langer, C. S., Magand, O., Nerentorp Mastromonaco, M., Nordstrøm, C., Pfaffhuber, K. A., Pirrone, N., Ryjkov, A., Selin, H., Skov, H., Song, S., Sprovieri, F., Steffen, A., Toyota, K., Travnikov, O., Yang, X., and Dommergue, A.: Chemical cycling and deposition of

atmospheric mercury in Polar Regions: review of recent measurements and comparison with models, in preparation.

Cobbett, F. D., Steffen, A., Lawson, G., and Van Heyst, B. J.: GEM fluxes and atmospheric mercury concentrations (GEM, RGM and Hg(p)) in the Canadian Arctic at Alert, Nunavut, Canada (February–June 2005), *Atmospheric Environment*, 41, 6527-6543, <http://dx.doi.org/10.1016/j.atmosenv.2007.04.033>, 2007.

Joffre, S. M.: Modelling the dry deposition velocity of highly soluble gases to the sea surface, *Atmospheric Environment* (1967), 22, 1137-1146, [http://dx.doi.org/10.1016/0004-6981\(88\)90343-5](http://dx.doi.org/10.1016/0004-6981(88)90343-5), 1988.

Pfaffhuber, K. A., Berg, T., Hirdman, D., and Stohl, A.: Atmospheric mercury observations from Antarctica: seasonal variation and source and sink region calculations, *Atmospheric Chemistry and Physics*, 12, 3241-3251, 2012.

Sprovieri, F., Pirrone, N., Hedgecock, I. M., Landis, M. S., and Stevens, R. K.: Intensive atmospheric mercury measurements at Terra Nova Bay in antarctica during November and December 2000, *Journal of geophysical research*, 107, 4722, 2002.

Temme, C., Einax, J. W., Ebinghaus, R., and Schroeder, W. H.: Measurements of atmospheric mercury species at a coastal site in the antarctic and over the atlantic ocean during polar summer, *Environmental Science and Technology*, 37, 22-31, 2003.

Zhang, L., Wright, L. P., and Blanchard, P.: A review of current knowledge concerning dry deposition of atmospheric mercury, *Atmospheric Environment*, 43, 5853-5864, 2009.



# New insights into the atmospheric mercury cycling in Central Antarctica and implications at a continental scale

Hélène Angot<sup>1</sup>, Olivier Magand<sup>2, 1</sup>, Detlev Helmig<sup>3</sup>, Philippe Ricaud<sup>4</sup>, Boris Quennehen<sup>2, 1</sup>, Hubert Gallée<sup>2, 1</sup>, Massimo Del Guasta<sup>5</sup>, Francesca Sprovieri<sup>6</sup>, Nicola Pirrone<sup>7</sup>, Joël Savarino<sup>2, 1</sup>, Aurélien Dommergue<sup>1, 2</sup>

<sup>1</sup>Univ. Grenoble Alpes, Laboratoire de Glaciologie et Géophysique de l'Environnement (LGGE), 38041 Grenoble, France.

<sup>2</sup>CNRS, Laboratoire de Glaciologie et Géophysique de l'Environnement (LGGE), 38041 Grenoble, France.

<sup>3</sup>Institute of Arctic and Alpine Research (INSTAAR), University of Colorado, Boulder, CO 80309-0450, USA.

<sup>4</sup>CNRM/GAME, Météo-France/CNRS, 42 avenue de Coriolis, 31057 Toulouse, France.

<sup>5</sup>CNR-Istituto Nazionale di Ottica, Largo E. Fermi 6, Firenze, 50125, Italy.

<sup>6</sup>CNR-Institute of Atmospheric Pollution Research, Division of Rende, Italy.

<sup>7</sup>CNR-Institute of Atmospheric Pollution Research, Montelibretti, Rome, Italy.

Correspondence to: A. Dommergue (aurelien.dommergue@ujf-grenoble.fr)

Changes made in the revised manuscript are highlighted in red.

## Abstract

Under the framework of the GMOS project (Global Mercury Observation System) atmospheric mercury monitoring has been implemented at Concordia Station on the high-altitude Antarctic plateau (75°06'S, 123°20'E, 3220 m above sea level). We report here the first year-round measurements of gaseous elemental mercury (Hg(0)) in the atmosphere and in snowpack interstitial air on the East Antarctic ice sheet. This unique dataset shows evidence of an **continuous intense** oxidation of atmospheric Hg(0) in summer (24-hour daylight) due to the high oxidative capacity of the Antarctic plateau atmosphere at this period of the year. Summertime Hg(0) concentrations exhibited a pronounced daily cycle in ambient air with maximal concentrations around midday. Photochemical reactions and chemical exchange at the air/snow interface were prominent, highlighting the role of the snowpack on the atmospheric mercury cycle. Our observations reveal a 20 to 30% decrease of atmospheric Hg(0) concentrations from May to mid-August (winter, 24-h darkness). This phenomenon has **not been reported** elsewhere and **possibly results from the dry deposition of Hg(0) onto the snowpack**. We also reveal the occurrence of multi-day to weeklong atmospheric Hg(0) depletion events in summer, not associated with depletions of ozone, and likely due to a

35 stagnation of air masses above the plateau triggering an accumulation of oxidants within the  
36 shallow boundary layer. Our observations suggest that the inland atmospheric reservoir is  
37 depleted in Hg(0) in summer. Due to katabatic winds flowing out from the Antarctic plateau  
38 down the steep vertical drops along the coast and according to observations at coastal  
39 Antarctic stations, the striking reactivity observed on the plateau most likely influences the  
40 cycle of atmospheric mercury at a continental scale.

41

## 42 **1 Introduction**

43 Mercury biomagnifies in its methylated form in aquatic food webs to elevated levels in  
44 freshwater and marine fish, causing adverse health effects to wildlife and humans (Mason et  
45 al., 2012). In 2013, the Minamata Convention (UNEP, 2013) was adopted and opened for  
46 signature to reduce the exposure of populations to this worldwide **contaminant**. Gaseous  
47 elemental mercury (Hg(0)), the most abundant form of mercury in the atmosphere, is  
48 efficiently transported around the globe, and even remote areas receive significant inputs of  
49 anthropogenic mercury by long-range atmospheric transport, as recently reported in modeling  
50 and observational studies (Weiss-Penzias et al., 2007; Lin et al., 2010).

51 Hg(0) can be oxidized into highly reactive and water-soluble gaseous and/or particulate  
52 divalent species (Hg(II) and Hg(p), respectively) (Lin and Pehkonen, 1999) leading to the  
53 formation and subsequent **rapid** deposition of reactive mercury onto environmental surfaces  
54 (Hedgecock and Pirrone, 2004). Upon deposition mercury can be reemitted back to the  
55 atmosphere or may enter the food chain through the conversion of Hg(II) to its methylated  
56 form (Driscoll et al., 2013). Effects and toxicity of mercury depends on this complex cycle,  
57 which is still not fully understood, and are only indirectly related to regional and global  
58 emissions (Driscoll et al., 2013). A better understanding of atmospheric mercury chemistry  
59 will lead to improved global transport and deposition models and could help refine pollution-  
60 control strategies around the world.

61 New oxidation pathways, discovered in 1995 in the Arctic (Schroeder et al., 1998) and  
62 highlighting the influence of halogen radicals on Hg(0) oxidation in spring, changed our  
63 understanding of the mercury cycle. While the Arctic has been extensively monitored, there is  
64 still much to be learned from the Antarctic continent where studies are scarce (Dommergue et  
65 al., 2010), especially on the high altitude plateau (see Fig. 1). The Antarctic plateau – ice-  
66 covered area of ~ 7 million km<sup>2</sup> – was first considered to be chemically-inactive and a giant  
67 cold trap for atmospheric species, ~~including mercury~~ (e.g., Lambert et al., 1990). It turned out  
68 to be highly photochemically active (Davis et al., 2001; Grannas et al., 2007) during the  
69 sunlit period with oxidant concentrations approaching those of tropical or urban mid-latitude  
70 environments (Eisele et al., 2008; Kukui et al., 2014). Earlier studies (Brooks et al., 2008;  
71 Dommergue et al., 2012) – the only two mercury studies ever carried out on the high-altitude  
72 Antarctic plateau with modern instruments – also suggested, based on short-term observations  
73 (a few weeks) in summer, an intense reactivity of mercury on the plateau at the air/snow  
74 interface. In this context, and under the framework of the GMOS project (Global Mercury  
75 Observation System, [www.gmos.eu](http://www.gmos.eu)), atmospheric mercury was continuously monitored at

76 Concordia Station (see Fig. 1) since 2012 and, for the first time, Hg(0) has been monitored  
77 year-round in both the snow interstitial air and the overlying atmosphere in 2013. Given harsh  
78 weather conditions (see section 2.1), and technical and logistical limitations, presenting such a  
79 record is in itself an important achievement. The main objective of this study is to provide  
80 new insights into the year-round cycling of gaseous mercury on the Antarctic plateau.

81

## 82 **2 Experimental Section**

### 83 **2.1 Sampling site**

84 Year-round measurements of gaseous mercury were conducted in 2012 and 2013 at the  
85 French/Italian Concordia Station (75°06'S, 123°20'E, 3220 m above sea level), located on the  
86 Antarctic plateau, 1100 km away from the nearest coast of East Antarctica (see Fig. 1).  
87 Concordia Station is a regional topographic maximum on the plateau; the surface terrain  
88 slopes do not exceed 1% (Genthon et al., 2010). The air temperature ranges between -20 °C in  
89 summer and -70 °C in winter, with an annual mean value of -45 °C (Pietroni et al., 2012).  
90 There is permanent daylight in summer and permanent darkness in winter. Snow  
91 accumulation is ~ 10 cm/year (Genthon et al., 2013).

### 92 **2.2 Sampling instrumentation**

93 Instrumentation was located in a below-surface shelter at the edge of the “clean area”, 800 m  
94 away from the main camp and upwind with respect to the dominant wind direction (south  
95 west). In 2012, year-round measurements were performed in the atmospheric boundary layer  
96 at about 500 cm above the snow surface. In 2013, measurements were performed in both the  
97 atmosphere and in snowpack interstitial air for several trace gases including gaseous mercury  
98 and ozone (O<sub>3</sub>). Sampling instrumentation included one 10 m meteorological tower for above-  
99 surface gradient sampling and two multi-inlet snow sampling manifolds (“snow towers”) for  
100 measuring trace gases at various depths in interstitial air (Bocquet et al., 2007; Seok et al.,  
101 2009). The 10 m meteorological tower was installed ~ 15 m upwind of the underground  
102 instrument shelter. It accommodated three gas inlets at 1070 cm, 210 cm, and 25 cm above the  
103 snow surface (see Fig. 2a). Trace gas measurements were acquired on each snow tower at six  
104 height levels: 50 and 10 cm above the snow surface, and 10, 30, 50, and 70 cm below the  
105 snow surface (see Fig. 2b). Sampling lines were purged continuously at 5 L/min on the  
106 meteorological tower and intermittently at ~ 2-3 L/min on the snow towers. On each snow  
107 tower, inlets were fitted with a small glass fiber filter in PTFE housing (25 mm Acrodisc  
108 syringe filters, Pall Life Sciences, Ann Arbor, Michigan, USA) to prevent snow crystals from  
109 entering the PFA sampling lines. Sampling lines were inside insulation tubing and the  
110 temperature of the sampling lines was maintained at a level 5-10 °C warmer than the  
111 snowpack temperature with a heat trace to prevent water vapor from freezing and clogging the  
112 lines. An automatic sampling pattern was implemented: trace gases were collected  
113 sequentially from the uppermost inlets on the meteorological tower to deepest levels of the  
114 snow towers. Measurements were taken for 10 min from each inlet.

## 115 **2.3 Gaseous mercury measurements**

116 Measurements were performed using a Tekran 2537A analyzer (Tekran Inc., Toronto,  
117 Canada) based on the amalgamation of mercury onto a gold cartridge followed by thermal  
118 desorption and detection by an integrated cold vapor atomic fluorescence spectrometer  
119 (CVAFS) at 253.7 nm (Fitzgerald and Gill, 1979; Bloom and Fitzgerald, 1988). The presence  
120 of two gold cartridges allowed alternating sampling and desorption modes and thus a  
121 continuous analysis in the pre-filtered (0.45 µm PTFE filter) and unheated sample air stream.  
122 The sampling resolution was 5 min with a sampling flow rate of 0.8 L/min. Concentrations  
123 are expressed in nanograms per cubic meter at standard temperature and pressure (273.15 K,  
124 1013.25 hPa). Using both a 0.45 µm PTFE filter at the entrance inlet of the sample line, and  
125 an unheated ¼"PTFE sample line, we assume that only Hg(0) (vs. total gaseous mercury,  
126 defined as the sum of gaseous mercury species) was efficiently collected and subsequently  
127 analyzed in this study (Steffen et al., 2002; Temme et al., 2003; Steffen et al., 2008).

### 128 *Quality assurance and quality control procedures*

129 An automatic calibration step of the Tekran 2537A analyzer was carried out every 25 h with  
130 an internal mercury permeation source. External calibrations were performed twice a year **by**  
131 **manually injecting** saturated mercury vapor taken from a temperature-controlled vessel  
132 (Tekran 2505 mercury vapor calibration unit, Hamilton digital syringe). As described by  
133 Angot et al. (2014), bi-monthly to monthly routine maintenance operations were performed. A  
134 software program was developed at the LGGE (Laboratoire de Glaciologie et Géophysique de  
135 l'Environnement) in accordance with quality control practice in well-established North  
136 American networks (Steffen et al., 2012). Based on various flagging criteria (Munthe et al.,  
137 2011; D'Amore et al., 2015), it enabled rapid data processing in order to produce clean time  
138 series of Hg(0). The detection limit is estimated at 0.10 ng/m<sup>3</sup> (Tekran, 2011). **Based on**  
139 **experimental evidence, the average systematic uncertainty for Hg(0) measurements is of ~ 10**  
140 **% (Slemr et al., 2015).**

## 141 **2.4 Surface snow sampling and analysis**

142 Surface snow samples (first cm) were collected weekly from February 2013 to January 2014  
143 using acid cleaned PTFE bottles and clean sampling procedures. Upon collection, samples  
144 were stored in the dark at -20 °C. Field blanks, carried out by opening and closing a bottle  
145 containing mercury-free water, were regularly conducted. Surface snow samples and field  
146 blanks were analyzed for total mercury using a Tekran Model 2600. **The instrument was**  
147 **calibrated with the NIST SRM-3133 mercury standard.** Quality assurance and quality control  
148 included the analysis of analytical blanks, replicates, and internal standards (**Reference Waters**  
149 **for mercury: HG102-2 at 22 ng/L from Environment Canada**). The limit of quantification –  
150 calculated as 10 times the standard deviation of a set of 3 analytical blanks – amounted to 0.3  
151 ng/L with a relative accuracy of ± 8%.

## 152 **2.5 Ancillary parameters**

### 153 *Ozone*

154 Measurements were performed using a UV absorption monitor (Thermo Electron  
155 Corporation, Franklin, MA), model 49I in 2012 (Legrand et al., 2016) and model 49C in  
156 2013. In 2013, the instrument was calibrated against the National Oceanic and Atmospheric  
157 Administration Global Monitoring Division, Boulder, Colorado, standard.

### 158 *Air mass back trajectories*

159 Air mass back trajectories were computed using the Lagrangian model FLEXPART (Stohl et  
160 al., 1998; Stohl and Thomson, 1999; Stohl et al., 2005) run in the backward mode and driven  
161 by NCEP (National Center for Environmental Predictions) GFS (Global Forecast System)  
162 final meteorological fields. Simulations were done every day at 1200 UTC in 2012 and 2013.  
163 For each simulation, 20000 pseudo-particles were released by the model around the position  
164 of Concordia Station and tracked for 5 days in three layers of altitude (0-0.1, 0.1-4 and 4-10  
165 km above ground level). Simulations at an altitude of 4-10 km were computed in order to  
166 investigate the potential occurrence of upper troposphere/lower stratosphere intrusions. For  
167 each 1-h time step, the model produced a normalized particle residence time (in seconds)  
168 within an output grid of  $0.5 \times 0.5^\circ$ . The sum of the 5 days outputs provided potential emission  
169 sensitivities (PES, in seconds) for the three layers of altitude. PES in a particular grid cell is  
170 proportional to the particle residence time in that cell. It should be noted that, in Antarctica,  
171 the meteorological data driving the FLEXPART transport model rely on sparse  
172 measurements. Consequently, the trajectories calculated in this region are often associated  
173 with relatively high uncertainties.

### 174 *Height of the boundary layer and shortwave radiation*

175 The height of the boundary layer and downwelling shortwave radiation were calculated by the  
176 MAR regional atmospheric model (Modèle Atmosphérique Régional). MAR was developed  
177 at the LGGE for Polar Regions and the simulations have been evaluated against  
178 meteorological observations made at Concordia Station (Gallée and Gorodetskaya, 2010;  
179 Gallée et al., 2015).

### 180 *Meteorological data*

181 Temperature, wind speed and direction were recorded at six height levels on a 45 m tower.  
182 The general observation set up is described by Genthon et al. (2010).

### 183 *Ice precipitation*

184 A tropospheric depolarization LIDAR (Light Detection And Ranging) operating at 523 nm  
185 provided tropospheric profiles of aerosol and clouds every 5 min allowing detection of  
186 water/ice clouds, snow drift, diamond dust and pollution plumes.

### 187 *Tropospheric temperature and integrated water vapor*

188 A H<sub>2</sub>O Antarctica Microwave Stratospheric and Tropospheric Radiometers (HAMSTRAD)  
189 instrument was used for the detection of the 60-GHz oxygen and the 183-GHz water vapor  
190 lines allowing measurement of tropospheric temperature and water vapor profiles,  
191 respectively, together with integrated water vapor (IWV) every 7 min. The instrument is fully  
192 automated and a liquid nitrogen calibration is performed twice per year (Ricaud et al., 2015)

193 *Eddy diffusivity and friction velocity*

194 The Eddy diffusivity was calculated as follows (Xiao et al., 2014):

$$195 \quad K = k u_* z / \varphi_h \quad (1)$$

196 where  $k$  (set to 0.40) is the von Karman constant,  $u_*$  the friction velocity (m/s),  $z$  the  
197 measurement height (m), and  $\varphi_h$  the Obukhov stability function. According to Frey et al.  
198 (2013), the stability function was  $\varphi_h = 0.95 + 4.62 \frac{z}{L}$  for stable conditions above snow (King  
199 and Anderson, 1994), and  $\varphi_h = 0.95 \left(1 - 11.6 \frac{z}{L}\right)^{-0.5}$  for unstable conditions (Hoegstroem,  
200 1988).  $u_*$  and  $L$  (the Obukhov length (m)) were computed from the three-dimensional wind  
201 components ( $u$ ,  $v$ ,  $w$ ) and temperature measured by a Metek sonic anemometer mounted at 8  
202 m above the snow surface.

203

## 204 **3 Results and Discussion**

### 205 **3.1 Hg(0) concentrations in ambient air**

206 The record of atmospheric Hg(0) over the entire 2012-2013 period is displayed in Fig. 3a.  
207 Hg(0) concentrations ranged from below 0.10 to 2.30 ng/m<sup>3</sup>, with average values amounting  
208 to  $0.76 \pm 0.24$  ng/m<sup>3</sup> in 2012, and to  $0.81 \pm 0.28$  ng/m<sup>3</sup>,  $0.84 \pm 0.27$  ng/m<sup>3</sup>, and  $0.82 \pm 0.26$   
209 ng/m<sup>3</sup> in 2013 at 25, 210, and 1070 cm above the snow surface, respectively. No significant  
210 difference was observed between annual averages of Hg(0) concentrations measured at the  
211 three inlets of the meteorological tower in 2013 ( $p$  value =  $3.1 \cdot 10^{-14}$ , Mann-Whitney test). **It**  
212 **should be noted that Hg(0) concentrations at the three inlets were significantly different in**  
213 **winter only (see section 3.1.4).** These mean annual Hg(0) concentrations are lower than  
214 annual averages reported at **near-coastal** or coastal Antarctic stations (i.e.,  $0.93 \pm 0.19$  ng/m<sup>3</sup>  
215 for Hg(0) at Troll (Pffaffhuber et al., 2012) and  $1.06 \pm 0.24$  ng/m<sup>3</sup> for total gaseous mercury at  
216 Neumayer (Ebinghaus et al., 2002)).

217 The seasonal boundaries **are** defined according to the transitions in downwelling shortwave  
218 radiation (see Fig. 3b) as follows: winter from May to mid-August, spring from mid-August  
219 to October, summer from November to mid-February, and fall from mid-February to April.  
220 Unlike in winter, Hg(0) concentrations were highly variable during the sunlit period with  
221 concentrations ranging from below 0.10 ng/m<sup>3</sup> to 1.50-2.00 ng/m<sup>3</sup>, up to twice the average  
222 background levels recorded in the Southern Hemisphere mid-latitudes (Slemr et al., 2015).  
223 These seasonal features, in good agreement with observations at other Antarctic stations

224 (Ebinghaus et al., 2002; Pfaffhuber et al., 2012), suggest the existence of a photochemically-  
225 induced reactivity of atmospheric mercury during the sunlit period. **The mechanisms which**  
226 **cause the seasonal variation of Hg(0) concentrations are discussed in the following sections.**

### 227 **3.1.1 Spring**

228 First discovered in the Arctic (Schroeder et al., 1998), atmospheric Hg(0) depletion events  
229 (AMDEs) result from an oxidation by reactive bromine species released during springtime so-  
230 called **“bromine explosions”** in coastal regions (Durnford and Dastoor, 2011 and references  
231 therein) and are concurrent with tropospheric O<sub>3</sub> depletion events (Simpson et al., 2007).  
232 Despite the distance of Concordia Station from the coast (1100 km), a Hg(0) depletion event  
233 was observed on 11 September 2013 due to a maritime air transport event (see Fig. 4a).  
234 During this event, Hg(0) concentrations **dropped from 0.85 to 0.56 ng/m<sup>3</sup> and** exhibited a  
235 strong positive correlation with O<sub>3</sub> mixing ratios ( $\rho = 0.94$ ,  $p$  value =  $5.10^{-7}$ ).

### 236 **3.1.2 Summertime**

#### 237 a) Oxidation of Hg(0) in ambient air and Hg(II) deposition onto snowpack

238 In summer, the mean atmospheric Hg(0) concentration was  $0.69 \pm 0.35$  ng/m<sup>3</sup> (**mean  $\pm$**   
239 **standard deviation**). This means that Hg(0) concentrations are ~ 25% lower than levels  
240 recorded at the same period of the year at coastal Antarctic stations (Ebinghaus et al., 2002;  
241 Sprovieri et al., 2002; Pfaffhuber et al., 2012). Total mercury concentrations in surface snow  
242 samples were highly variable (median value: 4.8 ng/L, range: < detection limit – 73.8 ng/L,  
243 see Fig. 5) and were higher in summer (median value: 10.4 ng/L, range: 1.3 – 73.8 ng/L),  
244 suggesting that divalent mercury species were preferentially deposited onto the snowpack at  
245 this period of the year. The lower Hg(0) concentrations in ambient air along with high total  
246 mercury concentrations in surface snow samples suggest an **continuous intense** oxidation of  
247 Hg(0) in ambient air in summer, followed by the deposition of oxidation products on surface  
248 snow. This hypothesis is further supported by elevated oxidized mercury concentrations  
249 measured on the Antarctic plateau at South Pole in summer (0.10 – 1.00 ng/m<sup>3</sup>) by Brooks et  
250 al. (2008).

251 The oxidative capacity of the Antarctic plateau atmosphere is elevated in summer, as  
252 evidenced by several studies (Davis et al., 2001; Grannas et al., 2007; Eisele et al., 2008;  
253 Kukui et al., 2014), likely explaining this **continuous intense** oxidation of Hg(0) in ambient  
254 air. Among these oxidants, NO<sub>2</sub>, RO<sub>2</sub>, and OH are particularly abundant at Concordia Station  
255 in summer (Frey et al., 2013; Kukui et al., 2014) and a recent study provided as a first  
256 estimate a BrO mixing ratio of 2-3 pptv near the ground during sunlight hours (Frey et al.,  
257 2015). Given the current understanding of mercury oxidation and the lack of continuous  
258 halogens measurements, we were not able to identify the exact mechanism for the reactivity  
259 observed at Concordia Station. A two-step oxidation mechanism, favored at cold  
260 temperatures, is worth being considered further. The initial recombination of Hg(0) and Br is  
261 followed by the addition of a second radical (e.g., I, Cl, BrO, ClO, OH, NO<sub>2</sub>, or HO<sub>2</sub>) in  
262 competition with thermal dissociation of the HgBr intermediate (Goodsite et al., 2004; Wang  
263 et al., 2014). According to Dibble et al. (2012), HO<sub>2</sub>, NO, NO<sub>2</sub>, and NO<sub>3</sub> **bind** Hg(0) too

264 weakly to initiate its oxidation in the gas phase and reactions of the HgBr intermediate with  
265 NO<sub>2</sub>, HO<sub>2</sub>, ClO, and BrO are more important than with Br and OH. Further modeling or  
266 laboratory chamber studies investigating the fate of Hg(0) in the presence of various potential  
267 oxidants are needed to improve our understanding of the mechanisms.

## 268 b) Multi-day depletion events of atmospheric Hg(0)

269 From 19 January to 8 February 2012 and from 5 to 20 February 2013 we observed Hg(0)  
270 depletion events. **The mechanism seems however different from springtime AMDEs** (see  
271 section 3.1.1). While atmospheric Hg(0) concentrations dropped and remained low ( $0.39 \pm$   
272  $0.19 \text{ ng/m}^3$  from 19 January to 8 February 2012,  $0.41 \pm 0.21 \text{ ng/m}^3$  from 5 to 20 February  
273 2013) for several weeks (see Figs. 3a, 6a, and 6e), O<sub>3</sub> showed no abnormal variability (see  
274 Figs. 6d and 6h). These depletion events occurred as air masses stagnated over the Antarctic  
275 plateau (see Figs. 4b and 4c) according to our FLEXPART simulations. This stagnation of air  
276 masses is confirmed in 2013 (see Figs. 6f and 6g) by a decrease of temperature at 10 m a.g.l  
277 (from  $-29 \pm 3 \text{ }^\circ\text{C}$  in January to  $-43 \pm 4 \text{ }^\circ\text{C}$  during the Hg(0) depletion event) and a low  
278 integrated water vapor ( $0.40 \pm 0.13 \text{ kg/m}^2$  during the Hg(0) depletion event vs.  $0.77 \pm 0.20$   
279  $\text{kg/m}^2$  in January). In both 2012 and 2013, depletions of Hg(0) ended when air masses started  
280 moving out of the plateau (see Figs. 4d and 4e).

281 While previous studies attributed high Hg(II) concentrations in the Antarctic summer to  
282 subsiding upper tropospheric air (Holmes et al., 2006; Brooks et al., 2008), potential emission  
283 sensitivities suggest that **the depletions of Hg(0) reported here** were unlikely concomitant with  
284 upper troposphere/lower stratosphere intrusions (see Figs. 4b and 4c, PES at 4-10 km). This is  
285 also confirmed by stable O<sub>3</sub> mixing ratios. High altitude vertical profiles of Hg(0) should be  
286 carried out to rule out this hypothesis of subsiding upper tropospheric air. We suggest that  
287 these Hg(0) depletion events observed at Concordia Station result from processes occurring  
288 within the shallow boundary layer. Since O<sub>3</sub> was not depleted during these events, Hg(0)  
289 depletion cannot be accounted for by bromine oxidation alone. FLEXPART simulations along  
290 with integrated water vapor and temperature measurements indicate that these Hg(0) depletion  
291 events occurred as air masses stagnated over the Antarctic plateau. As highlighted in the  
292 previous section, the oxidative capacity is high in summer on the plateau (Davis et al., 2001;  
293 Grannas et al., 2007; Eisele et al., 2008; Kukui et al., 2014). This air mass stagnation might  
294 favor an accumulation of oxidants within the shallow boundary layer (< 300 m in average),  
295 leading to an oxidation of Hg(0) stronger than usual.

## 296 c) Hg(0) diurnal cycle

297 Based on a week of measurements made at Concordia Station in January 2009, Dommergue et  
298 al. (2012) reported that atmospheric Hg(0) exhibited a **significant daily** cycle with maximal  
299 concentrations around noon. We show that this daily cycle occurred **throughout** the summer,  
300 with low atmospheric Hg(0) concentrations ( $\sim 0.50 \text{ ng/m}^3$ ) when solar radiation was  
301 minimum and a maximum ( $\sim 0.80 \text{ ng/m}^3$ ) around noon (see Fig. 7a). Such a pronounced daily  
302 cycle has never been observed at other Antarctic stations (Dommergue et al., 2010;  
303 Pfaffhuber et al., 2012). Several studies showed that Hg(0) emission from the snowpack  
304 maximizes near midday (e.g., Steffen et al., 2002; Ferrari et al., 2005; Brooks et al., 2006;



305 Faïn et al., 2007; Ferrari et al., 2008; Johnson et al., 2008). As suggested by Durnford and  
306 Dastoor (2011), the noon emission does not necessarily reflect maximum concentrations of  
307 cryospheric Hg(0) around midday (Hg(0) concentration peaked in the afternoon at 10 cm  
308 below the snow surface, see section 3.2.1) and could be driven by ventilation generated by  
309 atmospheric thermal convection. Stable boundary layers are almost ubiquitous in Polar  
310 Regions due to radiation cooling (Anderson and Neff, 2008). However, convective boundary  
311 layers have been observed in summer at polar domes at Concordia Station (King et al., 2006)  
312 and Summit in Greenland (Cohen et al., 2007). Fig. 7 displays the hourly mean variation of  
313 several parameters. As illustrated by Figs. 7c and 7d, and in agreement with earlier  
314 observations (Argentini et al., 2005; Pietroni et al., 2012; Argentini et al., 2013), there was a  
315 strong diurnal cycle in near-surface temperature and wind speed in summer at Concordia  
316 Station. These observations are typical for locations where a convective boundary layer  
317 develops as a response to daytime heating (King et al., 2006), as can be seen in Fig. 7e. In a  
318 convective boundary layer, vertical mixing is enhanced during convective hours (Anderson  
319 and Neff, 2008), as shown in Figs. 7f and 7g by increasing values for the friction velocity ( $u_*$ ,  
320 indicative of the strength of the mixing processes in the surface layer (Neff et al., 2008)) and  
321 the Eddy diffusivity ( $K$ ). Similarly, several studies highlighted that the atmospheric  
322 turbulence at Concordia Station in summer influences the vertical flux and concentration  
323 profiles of various atmospheric species (Legrand et al., 2009; Dommergue et al., 2012;  
324 Kerbrat et al., 2012; Frey et al., 2013).

325 In summary, the observed summertime Hg(0) diurnal cycle in ambient air might be due to a  
326 combination of factors: i) ~~a continuous~~ **the intense** oxidation of Hg(0) in ambient air due to  
327 the high oxidative capacity on the plateau – **as evidenced by low mean Hg(0) concentrations**  
328 **(see section 3.2.1.a)**, ii) ~~important~~ **subsequent** Hg(II) deposition onto snowpack – **as**  
329 **evidenced by elevated total mercury levels in surface snow samples (see section 3.2.1.a)**, and  
330 iii) ~~important~~ **emission** of Hg(0) from the snowpack during convective hours. **Fig. 8**  
331 **summarizes the processes that govern mercury exchange at the air/snow interface. Redox**  
332 **processes occurring within the snowpack are discussed in details in section 3.2.**

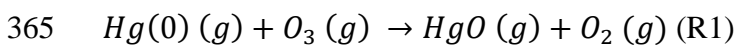
### 333 **3.1.3 Fall**

334 In fall, Hg(0) concentrations in ambient air no longer peaked around midday (see Fig. 7a) and  
335 were in average 67% higher than during the summer, exceeding levels recorded at lower  
336 latitudes in the Southern Hemisphere (Slemr et al., 2015). At this period of the year, the  
337 boundary layer lowered to ~ 50 m in average and no longer exhibited a pronounced diurnal  
338 cycle (see Figs. 3c and 7e). We believe that the shallow boundary layer could cause Hg(0)  
339 concentrations in ambient air to build up to where they exceeded levels recorded at lower  
340 latitudes in the Southern Hemisphere because Hg(0) – **emitted from the snowpack** – was  
341 dispersed into a reduced volume of air, limiting the dilution. Similarly, NO<sub>x</sub> mixing ratios are  
342 enhanced when the boundary layer is shallow (Neff et al., 2008; Frey et al., 2013). Elevated  
343 Hg(0) concentrations were also likely favored by the fact that oxidation in ambient air was  
344 weaker under lower solar radiation.

#### 345 **3.1.4 Winter**

346 While stable concentrations were expected in winter given the absence of photochemistry, our  
347 observations reveal a 20 to 30% decrease of atmospheric Hg(0) concentrations from May to  
348 mid-August (see Fig. 3a). Conversely, Hg(0) concentrations remained stable at Neumayer and  
349 Troll from late fall through winter (Ebinghaus et al., 2002; Pfaffhuber et al., 2012). This  
350 decreasing trend observed in winter might be due to several mechanisms, including gas-phase  
351 oxidation, ~~and~~ heterogeneous reactions, or dry deposition of Hg(0).

352 Several studies suggested the involvement of nitrate radicals in the night-time oxidation of  
353 Hg(0) (Mao and Talbot, 2012; Peleg et al., 2015). However, as previously mentioned, Dibble  
354 et al. (2012) indicated that NO<sub>3</sub> binds Hg(0) too weakly to initiate its oxidation in the gas  
355 phase. Another potential oxidant is O<sub>3</sub>, with this reactant reaching a maximum in the winter  
356 (see Fig. 3d). However, according to some theoretical studies (e.g., Hynes et al., 2009),  
357 reaction (R1) is unlikely to proceed as a homogeneous reaction. Several experimental studies  
358 confirmed the major product of reaction (R1) to be solid mercuric oxide, HgO (s) and not  
359 HgO (g) (e.g., Pal and Ariya, 2004; Ariya et al., 2009), suggesting that pure gas phase  
360 oxidation of elemental mercury by O<sub>3</sub> may not occur in the atmosphere. However, Calvert and  
361 Lindberg (2005) proposed an alternative mechanism that would make this reaction potentially  
362 viable in the atmosphere (Subir et al., 2011). The reaction may start with the formation of a  
363 metastable HgO<sub>3</sub> (g) molecule which then decomposes to OHgOO (g) and thereafter  
364 transforms to HgO (s) and O<sub>2</sub> (g).



366 As suggested by Subir et al. (2011), the influence of heterogeneous surfaces of water droplets,  
367 snow, ice and aerosols should be taken into account when attempting to describe mercury  
368 chemistry in the atmosphere. O'Concubhair et al. (2012) showed that freezing an acidic  
369 solution containing nitrite or hydrogen peroxide can oxidize dissolved gaseous mercury in the  
370 dark. Nitrous acid and hydrogen peroxide are present on the Antarctic plateau (Huey et al.,  
371 2004; Hutterli et al., 2004). As suggested by Dommergue et al. (2012), similar processes  
372 could occur in the snow or on surface hoar at Concordia Station in winter. In 2013, the height  
373 of measurement had a significant influence on the decline over time of Hg(0) concentrations  
374 (ANCOVA test,  $p$  value < 0.05), with a steeper decrease at 25 cm than at 1070 cm.  
375 Additionally, wintertime Hg(0) concentrations were significantly lower at 25 cm than at 1070  
376 cm ( $p$  value < 0.05, Mann-Whitney test). These results suggest that snowpack may act as a  
377 sink for mercury, enhancing the deposition rate due to heterogeneous reactions, through  
378 absorption of oxidation products, and/or physical sorption/condensation of Hg(0) on surface  
379 snow.

380 The observed declining trend could also be attributed to the dry deposition of Hg(0) onto the  
381 snowpack. The dry deposition velocity is defined as follows (Joffe, 1988), as the ratio  
382 between the deposition flux  $F$  (ng/m<sup>2</sup>/s) and the concentration  $C$  (ng/m<sup>3</sup>):

$$383 v_d = \frac{F}{C} \quad (2)$$

384 Denoting the height of the boundary layer  $h$  and the  $\text{Hg}(0)$  concentration at the beginning of  
385 winter  $C_0$ , the evolution of the concentration versus time is thus given by the following  
386 ordinary differential equation:

$$387 \quad C = C_0 e^{-(v_d/h)t} \quad (3)$$

388 During winter ( $t = 107$  days), the  $\text{Hg}(0)$  concentration gradually decreased from  $C_0 \sim 1.03$   
389  $\text{ng/m}^3$  to  $C \sim 0.73 \text{ ng/m}^3$  at 25 cm above the snowpack, in a mixing layer of 25 m high.  
390 According to Equation (3) the associated dry deposition velocity is  $9.3 \cdot 10^{-5} \text{ cm/s}$ . This result  
391 is in very good agreement with dry deposition velocities reported for  $\text{Hg}(0)$  over snow  
392 (Cobbett et al., 2007; Zhang et al., 2009).

393 In spite of the observed decreasing trend of  $\text{Hg}(0)$  concentrations in ambient air, total mercury  
394 concentrations in surface snow samples did not significantly increase over time in winter (see  
395 Fig. 5). Using a snow density of  $300 \text{ kg/m}^3$  a loss of  $0.30 \text{ ng/m}^3$  over a period of three months  
396 in a mixing layer of 25 m high would lead to a 2.5 ng/L increase in the first cm of the  
397 snowpack. Given the variability of chemical species deposition onto the snow surface, and the  
398 occurrence of either fresh snowfall or blowing snow, this 2.5 ng/L increase over a period of  
399 three months could not be detected in our weekly surface snow samples.

400 Despite the overall decreasing trend in winter,  $\text{Hg}(0)$  concentration exhibited abrupt increases  
401 when moist and warm air masses from lower latitudes occasionally reached Concordia  
402 Station. This is, for example, evidenced on 13 June 2012 by an increase of  $0.25 \text{ ng/m}^3$  of the  
403  $\text{Hg}(0)$  concentration, an increase of temperature at 10 m a.g.l. from  $-63$  to  $-26 \text{ }^\circ\text{C}$ , and a high  
404 integrated water vapor column (see Fig. 9).

## 405 **3.2 Hg(0)/Hg(II) redox conversions within the snowpack**

406 The 2013 record of  $\text{Hg}(0)$  in the snow interstitial air (SIA) is displayed in Fig. 10. Fig. 11  
407 depicts the mean  $\text{Hg}(0)$  concentration at various heights above and below the snow surface (in  
408 the interstitial air of the snow) for all seasons.

### 409 **3.2.1 Sunlit period**

410 During the sunlit period (summer, spring/fall),  $\text{Hg}(0)$  concentration peaked in the upper layers  
411 of the snowpack and then decreased with depth, with levels in the SIA dropping below  
412 atmospheric values.

413  $\text{Hg}(0)$  is generally produced in the upper layers of the snowpack – as the result of a  
414 photolytically initiated reduction of  $\text{Hg}(\text{II})$  (Lalonde et al., 2003) – and diffuses upward and  
415 downward. According to our observations,  $\text{Hg}(0)$  concentration peaked at  $\sim 10$  cm below the  
416 snow surface (see Fig. 11). Similarly, Brooks et al. (2008) reported  $\text{Hg}(0)$  concentrations  
417 peaking at a depth of 3 cm at South Pole. Below the top layer, the actinic flux decreases  
418 exponentially with depth (King and Simpson, 2001; Domine et al., 2008). The light  
419 penetration depth ( $e$ -folding depth) is the depth at which the actinic flux's magnitude is  $1/e$  of  
420 its incident value (Perovich, 2007). It is estimated that  $\sim 85\%$  of the photoreduction occurs in

421 the top two *e*-folding depths (King and Simpson, 2001). At Concordia Station, the *e*-folding  
422 depth is ~ 10 cm at 400 nm for the wind pack layers (France et al., 2011), which supports our  
423 observations. Reduced mercury can concurrently be reoxidized within the snowpack. Below  
424 the top layer, Hg(0) concentration in the SIA dropped with depth (see Fig. 11) suggesting that  
425 oxidation dominated in the deepest layers – in good agreement with observations within the  
426 snowpack at Kuujuarapik/Whapmagoostui, Québec, Canada (Dommergue et al., 2003) –  
427 leading to the formation of a Hg(II) reservoir.

428 The amount of Hg(0) emitted from the snowpack to the atmosphere (see section 3.1.2.c)  
429 depends on the balance of reduction and oxidation processes within the upper layers of the  
430 snowpack (Durnford and Dastoor, 2011). Fig. 12 depicts the hourly mean atmospheric and  
431 interstitial air Hg(0) concentrations. Our observations indicate that summertime Hg(0)  
432 concentration in the upper layers of the snowpack exhibited a diurnal cycle and peaked in the  
433 afternoon (see Fig. 12a). Conversely, in spring/fall, Hg(0) concentration reached a maximum  
434 at night and a minimum near midday in the upper layers of the snowpack (see Fig. 12b). The  
435 balance of reduction and oxidation processes within the upper layers of the snowpack  
436 suddenly shifted in summer. Similarly, Faïn et al. (2008) found that reduction dominated  
437 during summer and oxidation in spring in the upper layers of the snowpack at Summit,  
438 Greenland.

439 It is worth noting that Hg(0) concentration in the SIA was occasionally very high. For  
440 instance, on 24 February 2013, Hg(0) concentration reached 3.00 ng/m<sup>3</sup> at a depth of 10 cm.  
441 During this event, ice precipitation was observed at Concordia Station with depolarization  
442 values greater than 30% (data not shown). This suggests that the presence of ice crystals could  
443 enhance the dry deposition of Hg(II) species onto the snow surface leading to increased Hg(0)  
444 formation in the upper layers of the snowpack. Indeed, due to an elevated specific surface  
445 area, mercury-capture efficiency of ice crystals is high (Douglas et al., 2008). Unfortunately,  
446 due to a low sampling frequency of surface snow samples (weekly), total mercury  
447 concentrations cannot be used to study further the relationship between the occurrence of ice  
448 precipitation events and dry deposition of mercury species.

### 449 **3.2.2 Winter**

450 Contrarily to the sunlit period, Hg(0) concentration increased with depth in the SIA in winter  
451 (see Figs. 10 and 11). The average Hg(0) concentration amounted to 3.60 ng/m<sup>3</sup> at a depth of  
452 70 cm and was observed at a temperature of about -60 °C and not related to any change in  
453 atmospheric composition. Our observations are in agreement with earlier studies indicating  
454 that reduction of Hg(II) species is possible in the dark (Ferrari et al., 2004; Faïn et al., 2007;  
455 Ferrari et al., 2008). The production of Hg(0) might be due to the reduction of Hg(II) species  
456 accumulated in the deepest layers of the snowpack during the sunlit period (see section 3.2.1).  
457 This shift from oxidation to reduction in the deepest layers of the snowpack at the beginning  
458 of winter remains unexplained.

459

#### 460 **4 Implications at a continental scale**

461 Depletion events of atmospheric Hg(0) that have been observed in the Arctic and at various  
462 coastal Antarctic stations have been associated with O<sub>3</sub> depletions, where Hg(0) and O<sub>3</sub>  
463 concentrations are positively correlated (Simpson et al., 2007). Increases in both Hg(II) and  
464 Hg(p) have been reported in conjunction with decreases of Hg(0) (Lu et al., 2001; Lindberg  
465 et al., 2002; Aspino et al., 2005). Conversely, low Hg(0) concentrations that were not  
466 correlated or anti-correlated with O<sub>3</sub> were observed at Neumayer and Troll (Temme et al.,  
467 2003; Pfaffhuber et al., 2012), while elevated Hg(II) concentrations (up to 0.33 ng/m<sup>3</sup>) were  
468 recorded at Terra Nova Bay in the absence of Hg(0)/O<sub>3</sub> depletion (Sprovieri et al., 2002). The  
469 **continuous intense** oxidation of Hg(0) in summer (see section 3.1.2.a) and multi-day Hg(0)  
470 depletion events observed at Concordia Station in January/February (see section 3.1.2.b) are  
471 expected to result in the build-up of an inland atmospheric reservoir enriched in Hg(II) and  
472 depleted in Hg(0) in the summer. Due to strong katabatic winds flowing out from the  
473 Antarctic Plateau – generated by the negative buoyant force that develops in the stable cooled  
474 layer along the ice sheet slopes (Gallée and Pettré, 1998) –, a fraction of this inland  
475 atmospheric reservoir can be transported toward the coastal margin. The influence of the  
476 flows from the Antarctic plateau on coastal locations varies depending on the location. As  
477 demonstrated by Parish and Bromwich (1987) and Parish and Bromwich (2007), the volume  
478 of air moving off inland Antarctica toward the coastal margin displays significant spatial  
479 variability due to the topographic slope and orientation of the underlying ice sheets.  
480 Northward transport of air from the plateau is enhanced in a few locations called confluence  
481 zones – e.g., the broad region upslope from the Ross Ice Shelf at 175°E and the area near  
482 Adélie Land at 142°E – but can be sporadically observed elsewhere explaining the  
483 **mentioned** observations at Neumayer, Troll, or Terra Nova Bay (~~Temme et al., 2003;~~  
484 ~~Sprovieri et al., 2002; Pfaffhuber et al., 2012~~). Monitoring atmospheric mercury at a coastal  
485 station situated close to a confluence zone could provide new insights regarding the extent of  
486 the transport of reactive air masses from the Antarctic plateau. This topic is addressed in a  
487 companion paper (Angot et al., 2016). The Antarctic continent shelters unconventional  
488 atmospheric pathways of mercury reactivity both in winter and in summer. Its role should be  
489 taken into account in the modeling of the global geochemical cycle of mercury.

490

#### 491 **5 Conclusion**

492 Mean summertime atmospheric Hg(0) concentration was ~ 25% lower compared to values  
493 recorded at other Antarctic stations at the same period of the year, suggesting **an continuous**  
494 **intense** oxidation of atmospheric Hg(0) within the shallow boundary layer as a result of the  
495 high oxidative capacity of the Antarctic plateau atmosphere at this period of the year. This  
496 hypothesis is further supported by high total mercury concentrations in surface snow samples  
497 measured at the station (up to 74 ng/L). Our results confirm short-term observations by  
498 Brooks et al. (2008) and Dommergue et al. (2012) of chemical exchange at the air/snow  
499 interface. During the sunlit period, Hg(0) concentration peaked in the upper layers of the  
500 snowpack. Summertime Hg(0) concentration in ambient air exhibited a pronounced diurnal  
501 cycle likely due to large emissions from the snowpack as a response to daytime **snowpack**

502 **ventilation**. Our observations also reveal a decrease of atmospheric mercury concentrations in  
503 winter (24-h darkness) **possibly** due to the **dry deposition of Hg(0)**. Interestingly, this  
504 decreasing trend has never been observed elsewhere. **Additionally, Hg(0) concentrations**  
505 **increased with depth in the snow interstitial air in winter likely due to a dark reduction of**  
506 **Hg(II) species accumulated within the snowpack during the sunlit period**. Finally, we reveal  
507 the occurrence of multi-day to weeklong depletion events of Hg(0) in ambient air in summer,  
508 that are not associated with depletion of O<sub>3</sub>, and likely result from a stagnation of air masses  
509 on the plateau triggering an accumulation of oxidants in the shallow boundary layer. This  
510 behaviour is radically different from what is usually observed in the Arctic where only  
511 mercury depletion events that were associated with O<sub>3</sub> depletion (and with a Hg(0)/O<sub>3</sub>  
512 correlation) have been highlighted so far. According to observations at coastal Antarctic  
513 stations (**see section 4**), the reactivity observed at Concordia Station can be transported at a  
514 continental scale by strong katabatic winds. Our understanding of the atmospheric mercury  
515 chemistry on the Antarctic plateau is currently limited by the lack of continuous halogens  
516 measurements. Our findings point out new directions for future kinetic, observational, and  
517 modeling studies.

518

## 519 **Acknowledgements**

520 **Mercury data reported in this paper are available upon request at**  
521 **[http://sdi.ija.cnr.it/geoint/publicpage/GMOS/gmos\\_historical.zul](http://sdi.ija.cnr.it/geoint/publicpage/GMOS/gmos_historical.zul)**. We thank A. Barbero and  
522 the rest of the overwintering crew: S. Aubin, C. Lenormant, and R. Jacob. We also gratefully  
523 acknowledge M. Barret for the development of a QA/QC software program, **L. Bonato for the**  
524 **analysis of total mercury in surface snow samples**, D. Liptzin for the calculation of the  
525 Obukhov length and friction velocity, M. Legrand for the 2012 ozone data, C. Genthon for the  
526 meteorological data, E. Vignon for helpful discussion, and B. Jourdain and X. Faïn for their  
527 help in the field. This work contributed to the EU-FP7 project Global Mercury Observation  
528 System (GMOS – [www.gmos.eu](http://www.gmos.eu)) and has been supported by a grant from Labex  
529 OSUG@2020 (Investissements d'avenir – ANR10 LABX56), and the Institut Universitaire de  
530 France. BQ and PR acknowledge ANR CLIMSLIP and HAMSTRAD Program 910,  
531 respectively. Logistical and financial support was provided by the French Polar Institute IPEV  
532 (Program 1028, GMOstral and Program 1011, SUNITEDC), and a grant from the U.S.  
533 National Science Foundation (NSF, PLR#1142145). Computing resources for FLEXPART  
534 simulations were provided by the IPSL CICLAD/CLIMSERV mesocenter. Meteorological  
535 data were obtained thanks to LEFE/IMAGO programs CLAPA and GABLS4, IPEV program  
536 CALVA/1013, and Observatoire des Sciences de l'Univers de Grenoble (GLACIOCLIM  
537 observatory).

## References

- Anderson, P. S., and Neff, W. D.: Boundary layer physics over snow and ice, *Atmospheric Chemistry and Physics*, 8, 3563-3582, 2008.
- Angot, H., Barret, M., Magand, O., Ramonet, M., and Dommergue, A.: A 2-year record of atmospheric mercury species at a background Southern Hemisphere station on Amsterdam Island, *Atmospheric Chemistry and Physics* 14, 11461-11473, 2014.
- Angot, H., Dion, I., Vogel, N., Magand, O., Legrand, M., and Dommergue, A.: Atmospheric mercury record at Dumont d'Urville, East Antarctic coast: continental outflow and oceanic influences, *Atmospheric Chemistry and Physics Discussions*, 10.5194/acp-2016-257, in review, 2016.
- Argentini, S., Viola, A., Sempreviva, A. M., and Petenko, I.: Summer boundary-layer height at the plateau site of Dome C, Antarctica, *Boundary-Layer Meteorology*, 115, 409-422, 2005.
- Argentini, S., Petenko, I., Viola, A., Mastrantonio, G., Pietroni, I., Casasanta, G., Aristidi, E., and Ghenton, C.: The surface layer observed by a high-resolution sodar at Dome C, Antarctica, *Annals of geophysics*, 56, doi:10.4401/ag-6347, 2013.
- Ariya, P. A., Peterson, K., Snider, G., and Amyot, M.: Mercury chemical transformation in the gas, aqueous and heterogeneous phases: state-of-the-art science and uncertainties, in: *Mercury fate and transport in the global atmosphere*, edited by: Pirrone, N., and Mason, R. P., Springer, New York, 2009.
- Aspmo, K., Gauchard, P.-A., Steffen, A., Temme, C., Berg, T., Bahlmann, E., Banic, C., Dommergue, A., Ebinghaus, R., Ferrari, C., Pirrone, N., Sprovieri, F., and Wibetoe, G.: Measurements of atmospheric mercury species during an international study of mercury depletion events at Ny-Ålesund, Svalbard, spring 2003. How reproducible are our present methods?, *Atmospheric Environment*, 39, 7607-7619, 2005.
- Bloom, N. S., and Fitzgerald, W. F.: Determination of volatile mercury species at the picogram level by low temperature gas chromatography with cold-vapor atomic fluorescence detection, *Analytica Chimica Acta*, 208, 151-161, 1988.
- Bocquet, F., Helmig, D., and Oltmans, S. J.: Ozone in interstitial air of the mid-latitude seasonal snowpack at Niwot Ridge, Colorado, Arctic, Antarctic, and Alpine Research, 39, 375-387, 2007.
- Brooks, S., Saiz-Lopez, A., Skov, H., Lindberg, S. E., Plane, J. M. C., and Goodsite, M. E.: The mass balance of mercury in the springtime arctic environment, *Geophysical research letters*, doi: 10.1029/2005GL025525, 2006.
- Brooks, S. B., Arimoto, R., Lindberg, S. E., and Southworth, G.: Antarctic polar plateau snow surface conversion of deposited oxidized mercury to gaseous elemental mercury with fractional long-term burial, *Atmospheric Environment*, 42, 2877-2884, 2008.
- Calvert, J. G., and Lindberg, S. E.: Mechanisms of mercury removal by O<sub>3</sub> and OH in the atmosphere, *Atmospheric Environment*, 39, 3355-3367, 2005.
- Cobbett, F. D., Steffen, A., Lawson, G., and Van Heyst, B. J.: GEM fluxes and atmospheric mercury concentrations (GEM, RGM and Hg(p)) in the Canadian Arctic at Alert, Nunavut,

Canada (February–June 2005), *Atmospheric Environment*, 41, 6527-6543, <http://dx.doi.org/10.1016/j.atmosenv.2007.04.033>, 2007.

Cohen, L., Helmig, D., Neff, W. D., Grachev, A. A., and Fairall, C. W.: Boundary-layer dynamics and its influence on atmospheric chemistry at Summit, Greenland, *Atmospheric Environment*, 5044-5060, 2007.

D'Amore, F., Bencardino, M., Cinnirella, S., Sprovieri, F., and Pirrone, N.: Data quality through a web-based QA/QC system: implementation for atmospheric mercury data from the Global Mercury Observation System, *Environmental Science: Processes & Impacts*, 17, 1482-1491, 2015.

Davis, D., Nowak, J. B., Chen, G., Buhr, M., Arimoto, R., Hogan, A., Eisele, F., Mauldin, L., Tanner, D., Shetter, R., Lefer, B., and McMurry, P.: Unexpected high levels of NO observed at South Pole, *Geophysical research letters*, 28, 3625-3628, 2001.

Dibble, T. S., Zelic, M. J., and Mao, H.: Thermodynamics of reactions of ClHg and BrHg radicals with atmospherically abundant free radicals, *Atmospheric Chemistry and Physics*, 12, 10271-10279, 2012.

Domine, F., Albert, M., Huthwelker, T., Jacobi, H.-W., Kokhanovsky, A. A., Lehning, M., Picard, G., and Simpson, W. R.: Snow physics as relevant to snow photochemistry, *Atmospheric Chemistry and Physics*, 8, 171-208, 2008.

Dommergue, A., Ferrari, C. P., Gauchard, P.-A., and Boutron, C. F.: The fate of mercury species in a sub-arctic snowpack during snowmelt, *Geophysical research letters*, 30, 1621, 2003.

Dommergue, A., Sprovieri, F., Pirrone, N., Ebinghaus, R., Brooks, S., Courteaud, J., and Ferrari, C. P.: Overview of mercury measurements in the antarctic troposphere, *Atmospheric Chemistry and Physics*, 10, 3309-3319, 2010.

Dommergue, A., Barret, M., Courteaud, J., Cristofanelli, P., Ferrari, C. P., and Gallée, H.: Dynamic recycling of gaseous elemental mercury in the boundary layer of the antarctic plateau, *Atmospheric Chemistry and Physics*, 12, 11027-11036, 2012.

Douglas, T. A., Sturm, M., Simpson, W. R., Blum, J. D., Alvarez-Aviles, L., Keeler, G. J., Perovich, D. K., Biswas, A., and Johnson, K.: Influence of snow and ice crystal formation and accumulation on mercury deposition to the Arctic, *Environmental Science and Technology*, 42, 1542-1551, 2008.

Driscoll, C. T., Mason, R. P., Chan, H. M., Jacob, D. J., and Pirrone, N.: Mercury as a global pollutant: sources, pathways, and effects, *Environmental Science and Technology*, 47, 4967-4983, 2013.

Durnford, D., and Dastoor, A.: The behavior of mercury in the cryosphere: a review of what we know from observations, *Journal of geophysical research*, 116, doi:10.1029/2010JD014809, 2011.

Ebinghaus, R., Kock, H. H., Temme, C., Einax, J. W., Löwe, A. G., Richter, A., Burrows, J. P., and Schroeder, W. H.: Antarctic springtime depletion of atmospheric mercury, *Environmental Science and Technology*, 36, 1238-1244, 2002.



Eisele, F., Davis, D. D., Helmig, D., Oltmans, S. J., Neff, W., Huey, G., Tanner, D., Chen, G., Crawford, J. H., Arimoto, R., Buhr, M., Mauldin, L., Hutterli, M., Dibb, J., Blake, D., Brooks, S. B., Johnson, B., Roberts, J. M., Wang, Y., Tan, D., and Flocke, F.: Antarctic tropospheric chemistry (ANTCI) 2003 overview, *Atmospheric Environment*, 2008, 2749-2761, 2008.

Faïn, X., Grangeon, S., Bahlmann, E., Fritsche, J., Obrist, D., Dommergue, A., Ferrari, C., Cairns, W., Ebinghaus, R., Barbante, C., Cescon, P., and Boutron, C. F.: Diurnal production of gaseous mercury in the alpine snowpack before snowmelt, *Journal of geophysical research*, 112, doi:10.1029/2007JD008520, 2007.

Faïn, X., Ferrari, C., Dommergue, A., Albert, M., Battle, M., Arnaud, L., Barnola, J.-M., Cairns, W., Barbante, C., and Boutron, C. F.: Mercury in the snow and firn at Summit Station, Central Greenland, and implications for the study of past atmospheric mercury levels, *Atmospheric Chemistry and Physics*, 8, 3441-3457, 2008.

Ferrari, C. P., Dommergue, A., Boutron, C. F., Skov, H., Goodsite, M. E., and Jensen, B.: Nighttime production of elemental gaseous mercury in interstitial air of snow at Station Nord, Greenland, *Atmospheric Environment*, 38, 2727-2735, 2004.

Ferrari, C. P., Gauchard, P.-A., Aspmo, K., Dommergue, A., Magand, O., Bahlmann, E., Nagorski, S., Temme, C., Ebinghaus, R., Steffen, A., Banic, C., Berg, T., Planchon, F., Barbante, C., Cescon, P., and Boutron, C. F.: Snow-to-air exchanges of mercury in an arctic seasonal snowpack in Ny-Alesund, Svalbard, *Atmospheric Environment*, 39, 7633-7645, 2005.

Ferrari, C. P., Padova, C., Faïn, X., Gauchard, P.-A., Dommergue, A., Aspmo, K., Berg, T., Cairns, W., Barbante, C., Cescon, P., Kaleschke, L., Richter, A., Wittrock, F., and Boutron, C. F.: Atmospheric mercury depletion event study in Ny-Alesund (Svalbard) in spring 2005. Deposition and transformation of Hg in surface snow during springtime, *Science of the Total Environment*, 397, 167-177, 2008.

Fitzgerald, W. F., and Gill, G. A.: Subnanogram determination of mercury by two-stage gold amalgamation and gas detection applied to atmospheric analysis, *Analytical chemistry*, 51, 1714-1720, 1979.

France, J. L., King, M. D., Frey, M. M., Erbland, J., Picard, G., Preunkert, S., MacArthur, A., and Savarino, J.: Snow optical properties at Dome C (Concordia), Antarctica; implications for snow emissions and snow chemistry of reactive nitrogen, *Atmospheric Chemistry and Physics*, 11, 9787-9801, 2011.

Frey, M. M., Brough, N., France, J. L., Anderson, P. S., Traulle, O., King, M. D., Jones, A. E., Wolff, E. W., and Savarino, J.: The diurnal variability of atmospheric nitrogen oxides (NO and NO<sub>2</sub>) above the Antarctic Plateau driven by atmospheric stability and snow emissions, *Atmospheric Chemistry and Physics*, 13, 3045-3062, 2013.

Frey, M. M., Roscoe, H. K., Kukui, A., Savarino, J., France, J. L., King, M. D., Legrand, M., and Preunkert, S.: Atmospheric nitrogen oxides (NO and NO<sub>2</sub>) at Dome C, East Antarctica, during the OPALE campaign, *Atmospheric Chemistry and Physics*, 15, 7859-7875, 2015.

Gallée, H., and Pettré, P.: Dynamical constraints on katabatic wind cessation in Adélie Land, Antarctica, *Journal of the atmospheric sciences*, 55, 1755-1770, 1998.

- Gallée, H., and Gorodetskaya, I. V.: Validation of a limited area model over Dome C, Antarctic Plateau, during winter, *Climate Dynamics*, 34, 61-72, 2010.
- Gallée, H., Preunkert, S., Argentini, S., Frey, M. M., Genthon, C., Jourdain, B., Pietroni, I., Casasanta, G., Barral, H., Vignon, E., Amory, C., and Legrand, M.: Characterization of the boundary layer at Dome C (East Antarctica) during the OPALÉ summer campaign, *Atmospheric Chemistry and Physics*, 15, 6225-6236, 2015.
- Genthon, C., Town, M. S., Six, D., Favier, V., Argentini, S., and Pellegrini, A.: Meteorological atmospheric boundary layer measurements and ECMWF analyses during summer at Dome C, Antarctica, *Journal of geophysical research*, 115, D05104, 2010.
- Genthon, C., Six, D., Gallée, H., Grigioni, P., and Pellegrini, A.: Two years of atmospheric boundary layer observations on a 45-m tower at Dome C on the Antarctic plateau, *Journal of geophysical research: atmospheres*, 118, 3218-3232, 2013.
- Goodsite, M. E., Plane, J. M. C., and Skov, H.: A theoretical study of the oxidation of  $\text{Hg}^0$  to  $\text{HgBr}_2$  in the troposphere, *Environmental Science and Technology*, 38, 1772-1776, 2004.
- Grannas, A. M., Jones, A. E., Dibb, J., Ammann, M., Anastasio, C., Beine, H. J., Bergin, M., Bottenheim, J., Boxe, C. S., Carver, G., Chen, G., Crawford, J. H., Domine, F., Frey, M. M., Guzman, M. I., Heard, D. E., Helmig, D., Hoffmann, M. R., Honrath, R. E., Huey, L. G., Hutterli, M., Jacobi, H.-W., Klan, P., Lefer, B., McConnell, J. R., Plane, J. M. C., Sander, R., Savarino, J., Shepson, P. B., Simpson, W. R., Sodeau, J., Von Glasow, R., Weller, R., Wolff, E. W., and Zhu, T.: An overview of snow photochemistry: evidence, mechanisms and impacts, *Atmospheric Chemistry and Physics*, 7, 4329-4373, 2007.
- Hedgecock, I. M., and Pirrone, N.: Chasing quicksilver: modeling the atmospheric lifetime of  $\text{Hg}(0)$  in the marine boundary layer at various latitudes, *Environmental Science and Technology*, 38, 69-76, 2004.
- Hoegstroem, U.: Non-dimensional wind and temperature profiles in the atmosphere surface layer: a re-evaluation, *Boundary-Layer Meteorology*, 42, 55-78, 1988.
- Holmes, C. D., Jacob, D. J., and Yang, X.: Global lifetime of elemental mercury against oxidation by atomic bromine in the free troposphere, *Geophysical research letters*, 33, 2006.
- Huey, L. G., Tanner, D. J., Slusher, D. L., Dibb, J. E., Arimoto, R., Chen, G., Davis, D., Buhr, M. P., Nowak, J. B., Mauldin, L., Eisele, F. L., and Kosciuch, E.: CIMS measurements of  $\text{HNO}_3$  and  $\text{SO}_2$  at the South Pole during ISCAT 2000, *Atmospheric Environment*, 38, 5411-5421, 2004.
- Hutterli, M. A., McConnell, J. R., Chen, G., Bales, R. C., Davis, D. D., and Lenschow, D. H.: Formaldehyde and hydrogen peroxide in air, snow and interstitial air at South Pole, *Atmospheric Environment*, 38, 5439-5450, 2004.
- Hynes, A. J., Donohoue, D. L., Goodsite, M. E., and Hedgecock, I. M.: Our current understanding of major chemical and physical processes affecting mercury dynamics in the atmosphere and at the air-water/terrestrial interfaces, in: *Mercury fate and transport in the global atmosphere*, edited by: Pirrone, N., and Mason, R. P., Springer, New York, 427-457, 2009.

- Joffre, S. M.: Modelling the dry deposition velocity of highly soluble gases to the sea surface, *Atmospheric Environment* (1967), 22, 1137-1146, [http://dx.doi.org/10.1016/0004-6981\(88\)90343-5](http://dx.doi.org/10.1016/0004-6981(88)90343-5), 1988.
- Johnson, K. P., Blum, J. D., Keeler, G. J., and Douglas, T. A.: Investigation of the deposition and emission of mercury in arctic snow during an atmospheric mercury depletion event, *Journal of geophysical research*, 113, doi: 10.1029/2008JD009893, 2008.
- Kerbrat, M., Legrand, M. P., S., Gallée, H., and Kleffmann, J.: Nitrous acid at Concordia (inland site) and Dumont d'Urville (coastal site), east antarctica, *Journal of geophysical research*, 117, doi:10.1029/2011JD017149, 2012.
- King, J. C., and Anderson, P. S.: Heat and water vapor fluxes and scalar roughness lengths over an Antarctic ice shelf, *Boundary-Layer Meteorology*, 69, 101-121, 1994.
- King, J. C., Argentini, S. A., and Anderson, P. S.: Contrasts between the summertime surface energy balance and boundary layer structure at Dome C and Halley stations, Antarctica, *Journal of geophysical research*, 111, doi:10.1029/2005JD006130, 2006.
- King, M. D., and Simpson, W. R.: Extinction of UV radiation in Arctic snow at Alert, Canada (82°N), *Journal of geophysical research* 106, 12499-12507, 2001.
- Kukui, A., Legrand, M., Preunkert, S., Frey, M. M., Loisil, R., Gil Roca, J., Jourdain, B., King, M. D., France, J. L., and Ancellet, G.: Measurements of OH and RO<sub>2</sub> radicals at Dome C, East Antarctica, *Atmospheric Chemistry and Physics*, 14, 12373-12392, 2014.
- Lalonde, J. D., Amyot, M., Doyon, M.-R., and Auclair, J.-C.: Photo-induced Hg(II) reduction in snow from the remote and temperate experimental lakes area (Ontario, Canada), *Journal of geophysical research*, 108, doi:10.1029/2001JD001534, 2003.
- Lambert, G., Ardouin, B., and Sanak, J.: Atmospheric transport of trace elements toward Antarctica, *Tellus B*, 42, 76-82, 10.1034/j.1600-0889.1990.00009.x, 1990.
- Legrand, M., Preunkert, S., Jourdain, B., Gallée, H., Goutail, F., Weller, R., and Savarino, J.: Year-round record of surface ozone at coastal (Dumont d'Urville) and inland (Concordia) sites in east antarctica, *Journal of geophysical research*, 114, doi:10.1029/2008JD011667, 2009.
- Legrand, M. P., S., Savarino, J., Frey, M. M., Kukui, A., Helmig, D., Jourdain, B., Jones, A., Weller, R., Brough, N., and Gallée, H.: Inter-annual variability of surface ozone at coastal (Dumont d'Urville, 2004-2014) and inland (Concordia, 2007-2014) sites in East Antarctica, *Atmospheric Chemistry and Physics*, doi:10.5194/acp-2016-95, in press, 2016.
- Lin, C.-J., and Pehkonen, S. O.: The chemistry of atmospheric mercury: a review, *Atmospheric Environment*, 33, 2067-2079, 1999.
- Lin, C.-J., Pan, L., Streets, D. G., Shetty, S. K., Jang, C., Feng, X., Chu, H.-W., and Ho, T. C.: Estimating mercury emission outflow from East Asia using CMAQ-Hg, *Atmospheric Chemistry and Physics*, 10, 1856-1864, 2010.
- Lindberg, S. E., Brooks, S., Lin, C.-J., Scott, K. J., Landis, M. S., Stevens, R. K., Goodsite, M. E., and Richter, A.: Dynamic oxidation of gaseous mercury in the arctic troposphere at polar sunrise, *Environmental Science and Technology*, 36, 1245-1256, 2002.

- Lu, J. Y., Schroeder, W. H., Barrie, L. A., Steffen, A., Welch, H. E., Martin, K., Lockhart, L., Hunt, R. V., Boila, G., and Richter, A.: Magnification of atmospheric mercury deposition to polar regions in springtime: the link to tropospheric ozone depletion chemistry, *Geophysical research letters*, 28, 3219-3222, 2001.
- Mao, H., and Talbot, R.: Speciated mercury at marine, coastal, and inland sites in New England - Part 1: Temporal variability, *Atmospheric Chemistry and Physics*, 12, 5099-5112, 2012.
- Mason, R. P., Choi, A. L., Fitzgerald, W. F., Hammerschmidt, C. R., Lamborg, C. H., Soerensen, A. L., and Sunderland, E. M.: Mercury biogeochemical cycling in the ocean and policy implications, *Environmental Research*, 119, 101-117, 2012.
- Munthe, J., Sprovieri, F., Horvat, M., and Ebinghaus, R.: SOPs and QA/QC protocols regarding measurements of TGM, GEM, RGM, TPM and mercury in precipitation in cooperation with WP3, WP4 and WP5. GMOS deliverable 6.1, CNR-IIA, IVL. <http://www.gmos.eu>, last access: 3 March 2014, 2011.
- Neff, W., Helmig, D., Grachev, A. A., and Davis, D.: A study of boundary layer behavior associated with high NO concentrations at the South Pole using a minisodar, tethered balloon, and sonic anemometer, *Atmospheric Environment*, 42, 2762-2779, 2008.
- O'Concubhair, R., O'Sullivan, D. A., and Sodeau, J. R.: Dark oxidation of dissolved gaseous mercury in polar ice mimics, *Environmental Science and Technology*, 46, 2012.
- Pal, B., and Ariya, P. A.: Studies of ozone initiated reactions of gaseous mercury: kinetics, product studies, and atmospheric implications, *Physical Chemistry Chemical Physics*, 6, 572-579, 2004.
- Parish, T. R., and Bromwich, D. H.: The surface windfield over the Antarctic ice sheets, *Nature*, 328, 51-54, 1987.
- Parish, T. R., and Bromwich, D. H.: Reexamination of the near-surface airflow over the Antarctic continent and implications on atmospheric circulations at high southern latitudes, *Monthly Weather Review*, 135, 1961-1973, 2007.
- Peleg, M., Tas, E., Obrist, D., Matveev, V., Moore, C., Gabay, M., and Luria, M.: Observational evidence for involvement of nitrate radicals in nighttime oxidation of mercury, *Environmental Science and Technology*, 49, 14008-14018, 2015.
- Perovich, D. K.: Light reflection and transmission by a temperate snow cover, *Journal of glaciology*, 53, 201-210, 2007.
- Pfaffhuber, K. A., Berg, T., Hirdman, D., and Stohl, A.: Atmospheric mercury observations from Antarctica: seasonal variation and source and sink region calculations, *Atmospheric Chemistry and Physics*, 12, 3241-3251, 2012.
- Pietroni, I., Argentini, S., Petenko, I., and Sozzi, R.: Measurements and parametrizations of the atmospheric boundary-layer height at Dome C, Antarctica, *Boundary-Layer Meteorology*, 143, 189-206, 2012.
- Ricaud, P., Grigioni, P., Zbinden, R., Attié, J.-L., Genoni, L., Galeandro, A., Moggio, L., Montaguti, S., Petenko, I., and Legovini, P.: Review of tropospheric temperature, absolute

humidity and integrated water vapor from the HAMSTRAD radiometer installed at Dome C, Antarctica, 2009-14, *Antarctic Science*, 27, 598-616, 2015.

Schroeder, W. H., Anlauf, K. G., Barrie, L. A., Lu, J. Y., Steffen, A., Schneeberger, D. R., and Berg, T.: Arctic springtime depletion of mercury, *Nature*, 394, 331-332, 1998.

Seok, B., Helmig, D., Williams, M. W., Liptzin, D., and Chowanski, K. H., J.: An automated system for continuous measurements of trace gas fluxes through snow: an evaluation of the gas diffusion method at a subalpine forest site, Niwot Ridge, Colorado, *Biogeochemistry*, 95, 95-113, 2009.

Simpson, W. R., Von Glasow, R., Riedel, K., Anderson, P. S., Ariya, P. A., Bottenheim, J., Burrows, J. P., Carpenter, L. J., Friess, U., Goodsite, M. E., Heard, D. E., Hutterli, M., Jacobi, H.-W., Kaleschke, L., Neff, W., Plane, J. M. C., Platt, U., Richter, A., Roscoe, H. K., Sander, R., Shepson, P. B., Sodeau, J., Steffen, A., Wagner, T., and Wolff, E. W.: Halogens and their role in polar boundary-layer ozone depletion, *Atmospheric Chemistry and Physics*, 7, 4375-4418, 2007.

Slemr, F., Angot, H., Dommergue, A., Magand, O., Barret, M., Weigelt, A., Ebinghaus, R., Brunke, E.-G., Pfaffhuber, K. A., Edwards, G., Howard, D., Powell, J., Keywood, M., and Wang, F.: Comparison of mercury concentrations measured at several sites in the Southern Hemisphere, *Atmospheric Chemistry and Physics*, 15, 3125-3133, 2015.

Sprovieri, F., Pirrone, N., Hedgecock, I. M., Landis, M. S., and Stevens, R. K.: Intensive atmospheric mercury measurements at Terra Nova Bay in Antarctica during November and December 2000, *Journal of geophysical research*, 107, 4722, 2002.

Steffen, A., Schroeder, W., Bottenheim, J., Narayan, J., and Fuentes, J. D.: Atmospheric mercury concentrations: measurements and profiles near snow and ice surfaces in the Canadian Arctic during Alert 2000, *Atmospheric Environment*, 36, 2653-2661, 2002.

Steffen, A., Douglas, T., Amyot, M., Ariya, P. A., Aspö, K., Berg, T., Bottenheim, J., Brooks, S., Cobbett, F., Dastoor, A., Dommergue, A., Ebinghaus, R., Ferrari, C., Gardfeldt, K., Goodsite, M. E., Lean, D., Poulain, A. J., Scherz, C., Skov, H., Sommar, J., and Temme, C.: A synthesis of atmospheric mercury depletion event chemistry in the atmosphere and snow, *Atmospheric Chemistry and Physics*, 8, 1445-1482, 2008.

Steffen, A., Scherz, T., Oslon, M., Gay, D. A., and Blanchard, P.: A comparison of data quality control protocols for atmospheric mercury speciation measurements, *Journal of Environmental Monitoring*, 14, 752-765, doi: 10.1039/c2em10735j, 2012.

Stohl, A., Hittenberger, M., and Wotawa, G.: Validation of the Lagrangian particle dispersion model FLEXPART against large scale tracer experiments, *Atmospheric Environment*, 32, 4245-4264, 1998.

Stohl, A., and Thomson, D. J.: A density correction for Lagrangian particle dispersion models, *Boundary-Layer Meteorology*, 90, 155-167, 1999.

Stohl, A., Forster, C., Frank, A., Seibert, P., and Wotawa, G.: Technical note: the Lagrangian particle dispersion model FLEXPART version 6.2, *Atmospheric Chemistry and Physics*, 5, 2461-2474, 2005.

Subir, M., Ariya, P. A., and Dastoor, A.: A review of uncertainties in atmospheric modeling of mercury chemistry I. Uncertainties in existing kinetic parameters - fundamental limitations

and the importance of heterogeneous chemistry, *Atmospheric Environment*, 45, 5664-5675, 2011.

Tekran: Tekran 2537 mercury monitor detection limit. Summary of known estimates, Tekran Instruments Corp., Toronto, ON, Canada., 2011.

Temme, C., Einax, J. W., Ebinghaus, R., and Schroeder, W. H.: Measurements of atmospheric mercury species at a coastal site in the antarctic and over the atlantic ocean during polar summer, *Environmental Science and Technology*, 37, 22-31, 2003.

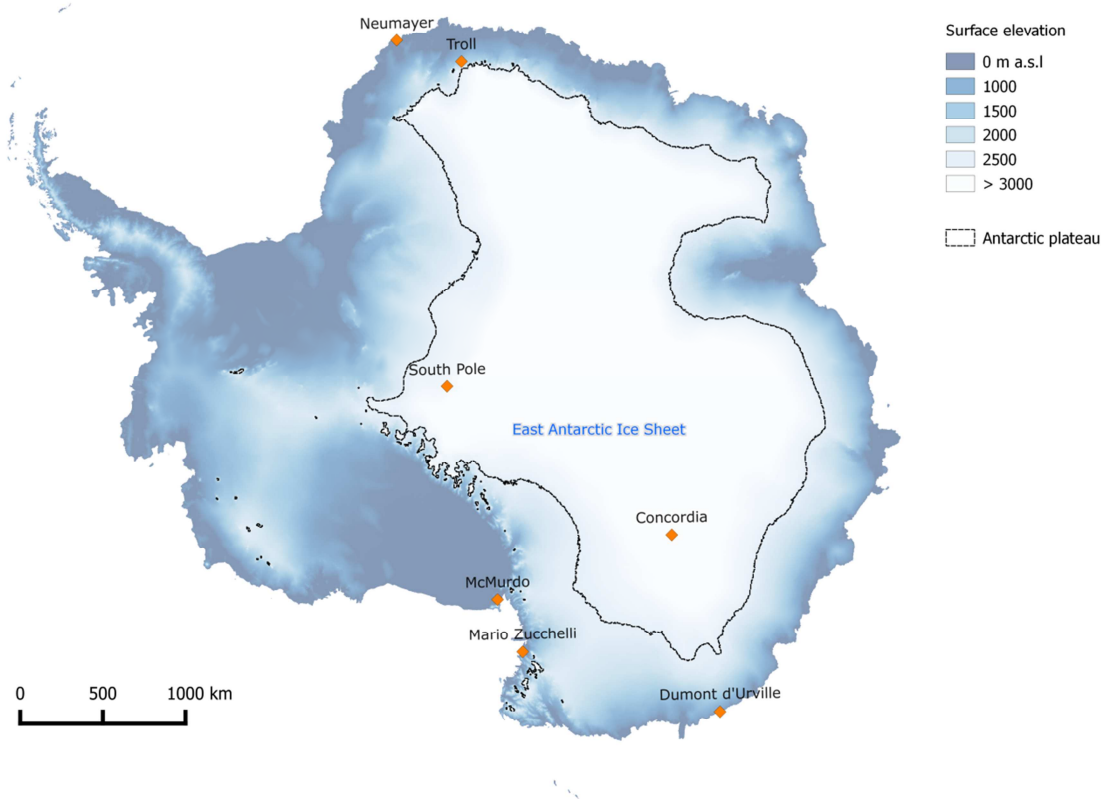
UNEP: Text of the Minamata Convention on Mercury for adoption by the Conference of Plenipotentiaries. [unep.org](http://www.unep.org/hazardoussubstances/Portals/9/Mercury/Documents/dipcon/CONF_3_Minamata%20Convention%20on%20Mercury_final%2026%2008_e.pdf). July 31, Available at: [http://www.unep.org/hazardoussubstances/Portals/9/Mercury/Documents/dipcon/CONF\\_3\\_Minamata%20Convention%20on%20Mercury\\_final%2026%2008\\_e.pdf](http://www.unep.org/hazardoussubstances/Portals/9/Mercury/Documents/dipcon/CONF_3_Minamata%20Convention%20on%20Mercury_final%2026%2008_e.pdf), last access: 27 March 2016, 2013.

Wang, F., Saiz-Lopez, A., Mahajan, A. S., Gomez Martin, J. C., Armstrong, D., Lemes, M., Hay, T., and Prados-Roman, C.: Enhanced production of oxidised mercury over the tropical pacific ocean: a key missing oxidation pathway, *Atmospheric Chemistry and Physics*, 14, 1323-1335, 2014.

Weiss-Penzias, P., Jaffe, D. A., Swartzendruber, P., Hafner, W., Chand, D., and Prestbo, E.: Quantifying Asian and biomass burning sources of mercury using the Hg/CO ratio in pollution plumes observed at the Mount Bachelor observatory, *Atmospheric Environment*, 41, 4366-4379, 2007.

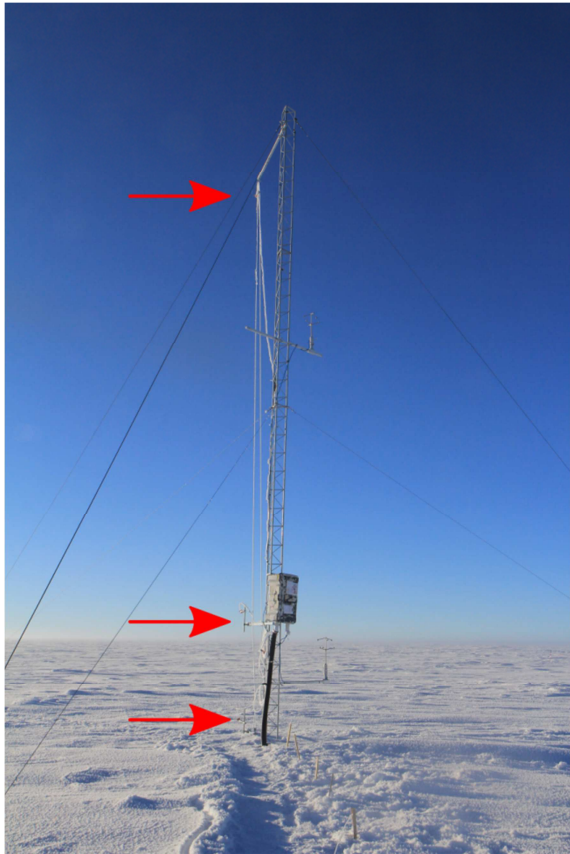
Xiao, W., Liu, S., Li, H., Xiao, Q., Wang, W., Hu, Z., Hu, C., Gao, Y., Shen, J., Zhao, X., Zhang, M., and Lee, X.: A flux-gradient system for simultaneous measurement of the CH<sub>4</sub>, CO<sub>2</sub>, and H<sub>2</sub>O fluxes at a lake-air interface, *Environmental Science and Technology*, 48, 14490-14498, 2014.

Zhang, L., Wright, L. P., and Blanchard, P.: A review of current knowledge concerning dry deposition of atmospheric mercury, *Atmospheric Environment*, 43, 5853-5864, 2009.



**Figure 1:** Map of Antarctica showing surface elevation (meters above sea level, m a.s.l) and the position of stations where atmospheric mercury measurements have been performed with modern on-line instruments. The black line shows the periphery of the high altitude plateau (> 2500 m a.s.l).

a)

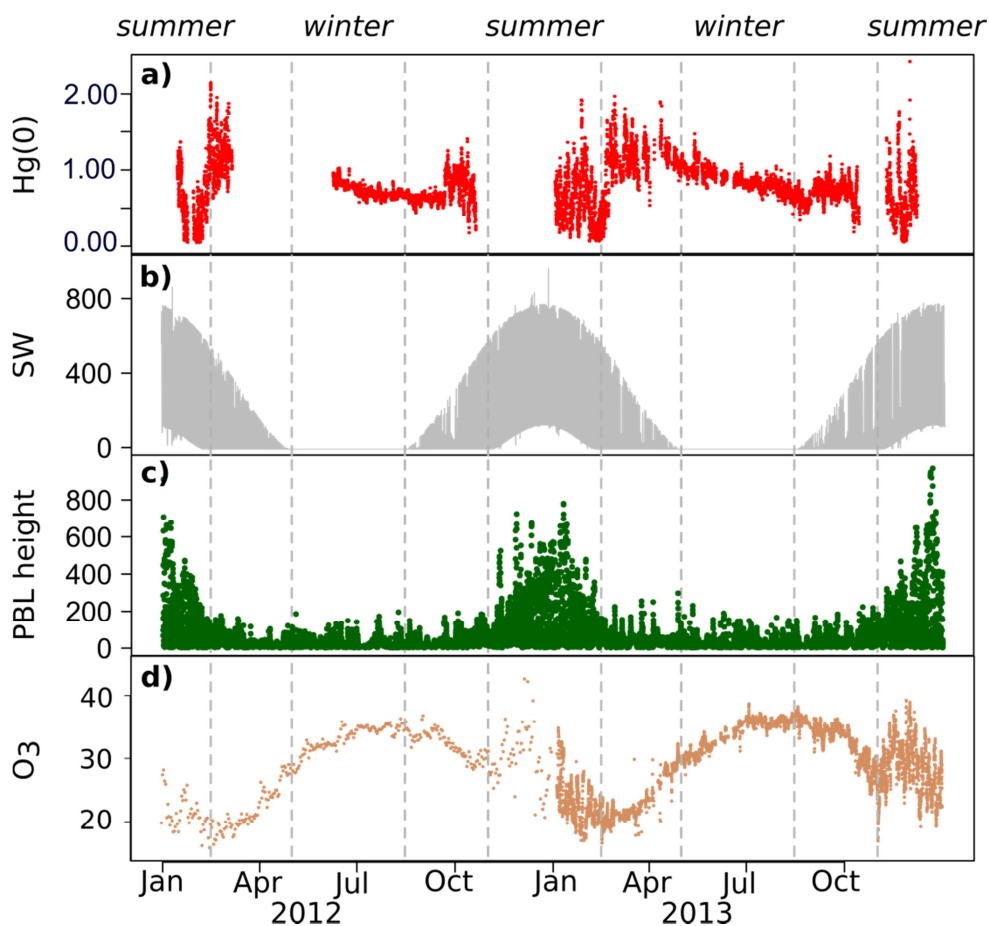


b)

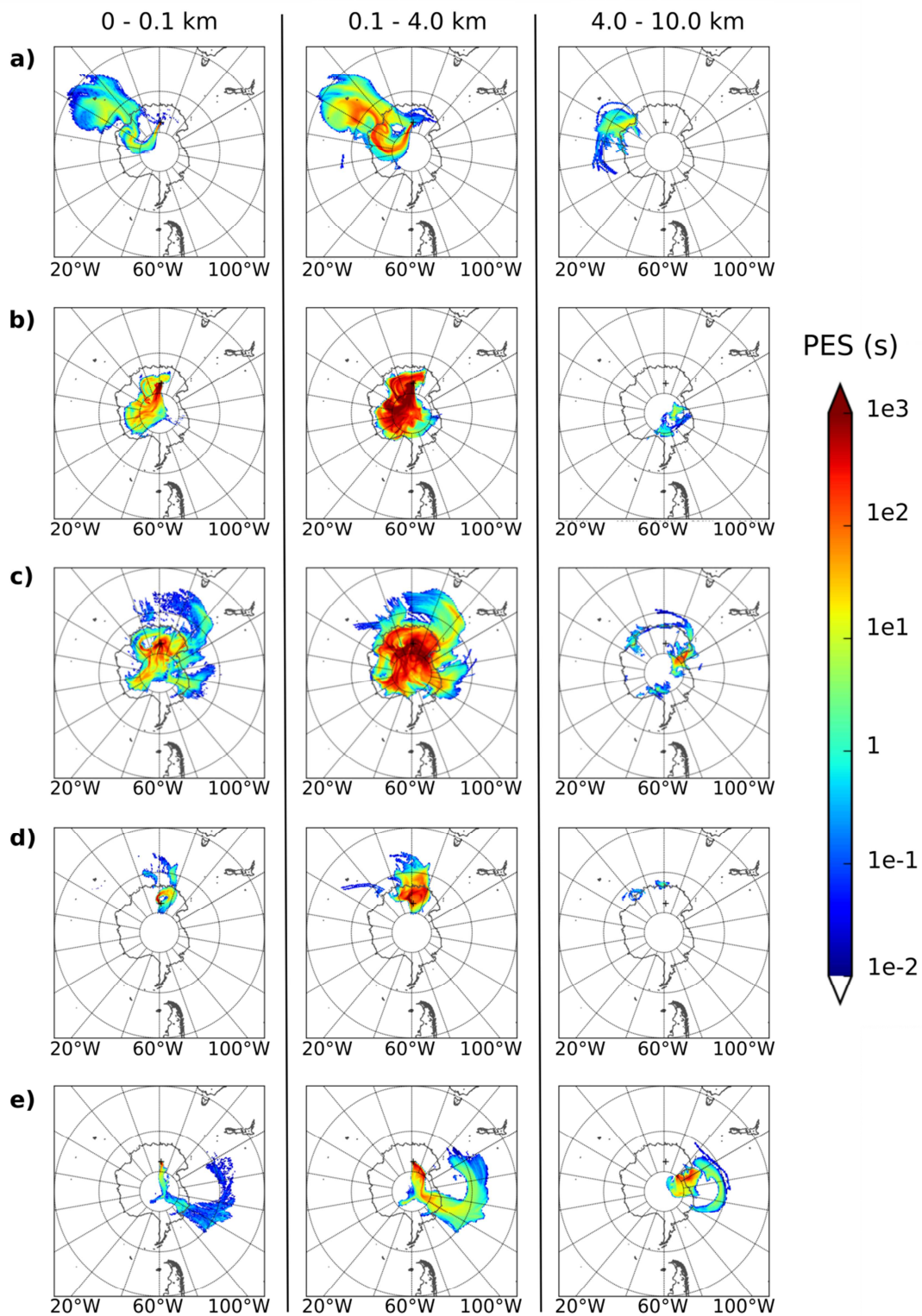


**Figure 2:** Photographs showing **a)** the meteorological tower with the three gas inlets (red arrows) at 1070 cm, 210 cm and 25 cm above the snow surface (photo credit: B. Jourdain), and **b)** one of the snow towers with the two sampling inlets above the snowpack at 50 and 10 cm (photo credit: D. Helmig).

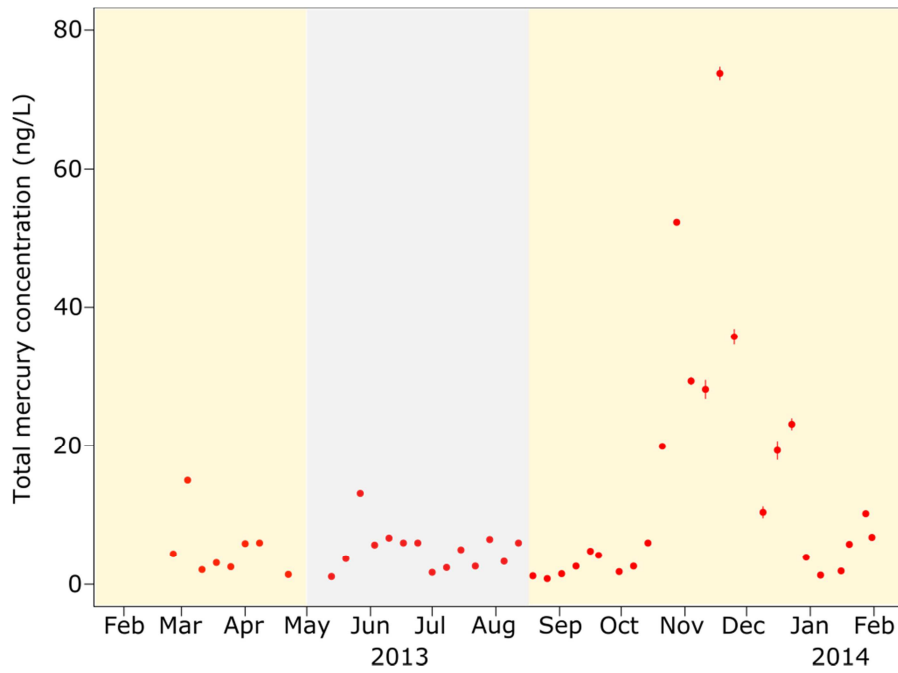




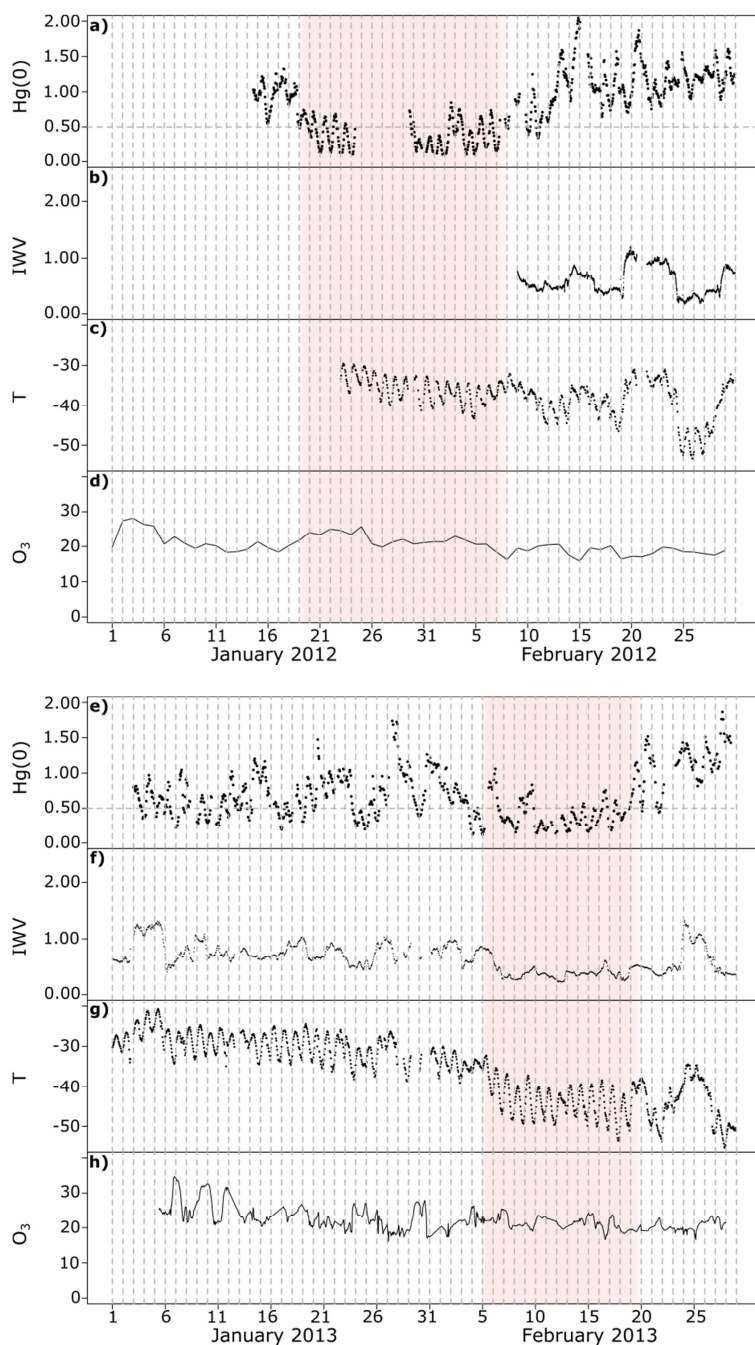
**Figure 3:** Annual variation in 2012 and 2013 of **a)** hourly-averaged Hg(0) concentrations (in  $\text{ng}/\text{m}^3$ ) at 500 cm and 25 cm above the snow surface in 2012 and 2013, respectively, **b)** downwelling shortwave (SW) radiation (in  $\text{W}/\text{m}^2$ ), **c)** planetary boundary layer (PBL) height (in m), and **d)** ozone ( $\text{O}_3$ , daily mean in 2012 and hourly mean in 2013) mixing ratios (in ppbv). The vertical dashed lines represent seasonal boundaries.



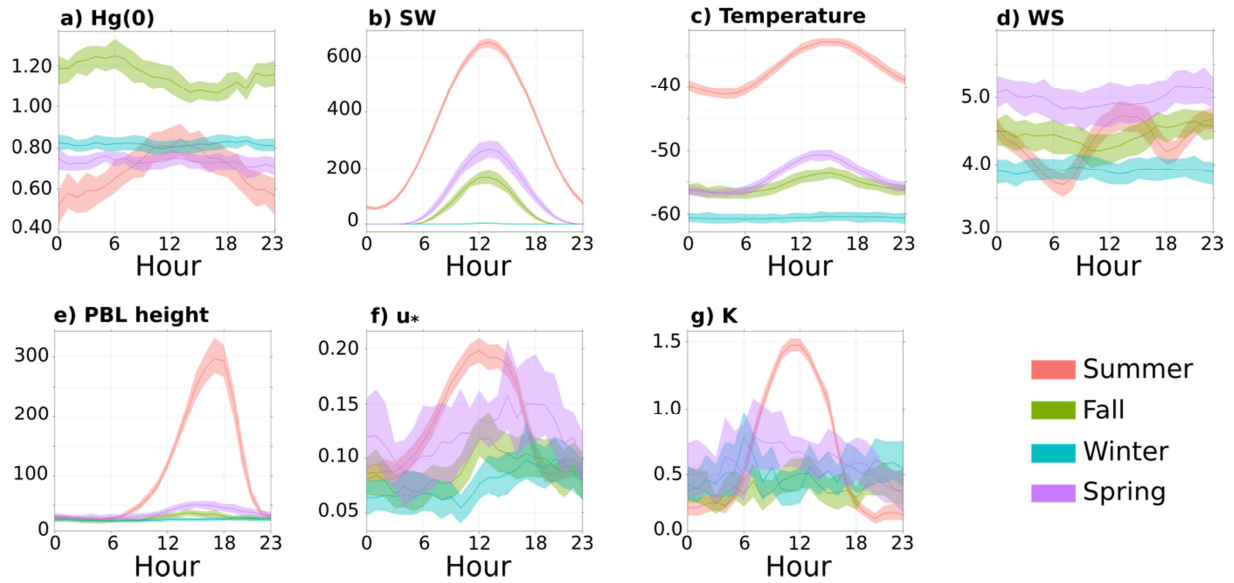
**Figure 4:** Back trajectories for the 3 layers of altitude colored according to the potential emission sensitivity (PES, in seconds) **a)** on 11 September 2013, **b)** from 19 January to 8 February 2012, **c)** from 5 to 20 February 2013, **d)** on 10 February 2012, and **e)** on 22 February 2013. Note that PES in a particular grid cell is proportional to the particle residence time in that cell.



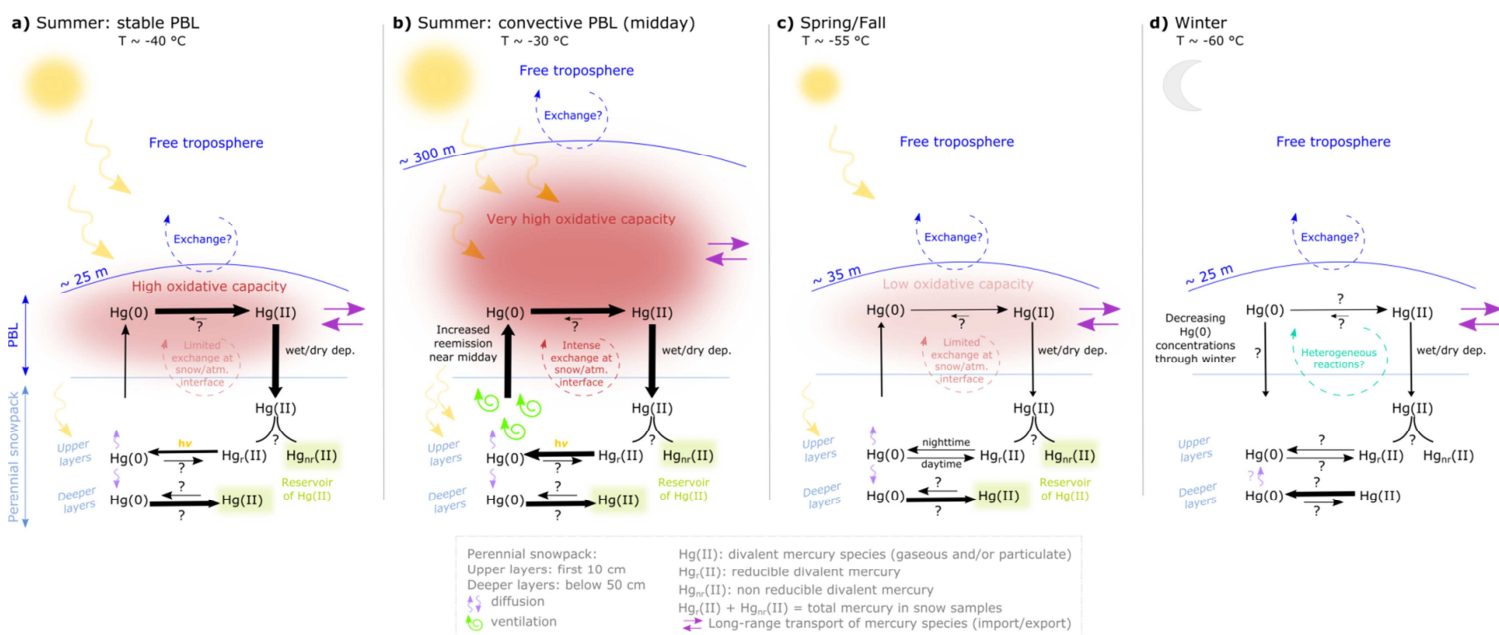
**Figure 5:** Total mercury concentration (ng/L), along with standard errors, in surface snow samples collected weekly at Concordia Station from February 2013 to January 2014. Dark period (winter) highlighted in grey, sunlit period highlighted in yellow. Total mercury concentrations were elevated (up to 74 ng/L) in November-December 2013 (summer). All samples were analyzed in replicates of three. Standard errors are frequently smaller than the width of the dots.



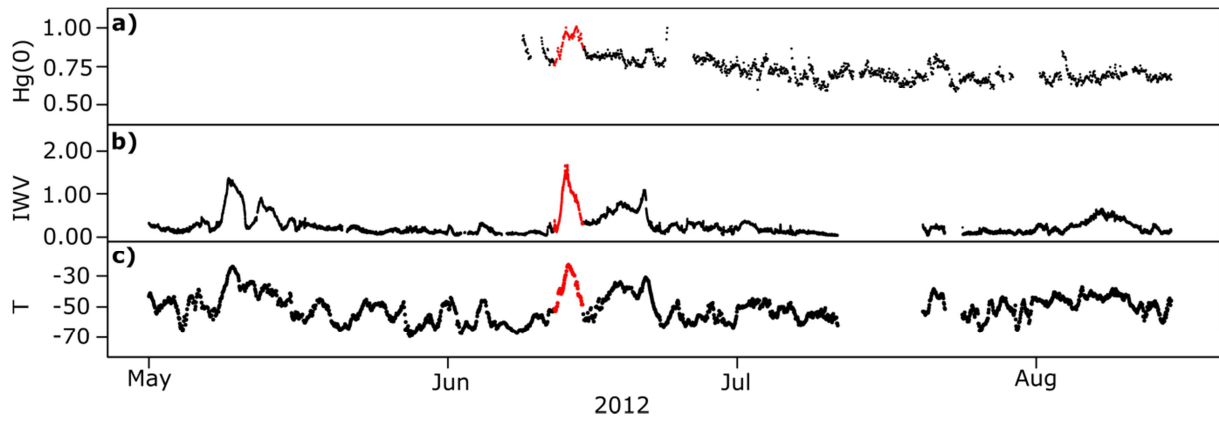
**Figure 6:** Top: January and February 2012 cycle of: **a)** hourly-averaged Hg(0) concentrations (in  $\text{ng}/\text{m}^3$ ) at 500 cm above the snow surface, **b)** Integrated Water Vapor (IWV,  $\text{kg}/\text{m}^2$ ), **c)** Temperature (in  $^{\circ}\text{C}$ ) at 10 m above ground level, and **d)** ozone ( $\text{O}_3$ , daily mean) mixing ratios (ppbv). Hg(0) was low from 19 January to 8 February (period highlighted in red) while  $\text{O}_3$  showed no abnormal variability. Bottom: January and February 2013 cycle of: **e)** hourly-averaged Hg(0) concentrations (in  $\text{ng}/\text{m}^3$ ) at 210 cm above the snow surface, **f)** Integrated Water Vapor (IWV,  $\text{kg}/\text{m}^2$ ), **g)** Temperature (in  $^{\circ}\text{C}$ ) at 10 m above ground level, and **h)** ozone ( $\text{O}_3$ , hourly mean) mixing ratio (ppbv). Hg(0), IWV, and temperature were low from 5 to 20 February (period highlighted in red) while  $\text{O}_3$  showed no abnormal variability. Note that Hg(0) concentrations exhibited the same pattern at the three inlets of the meteorological tower from 5 to 20 February 2013.



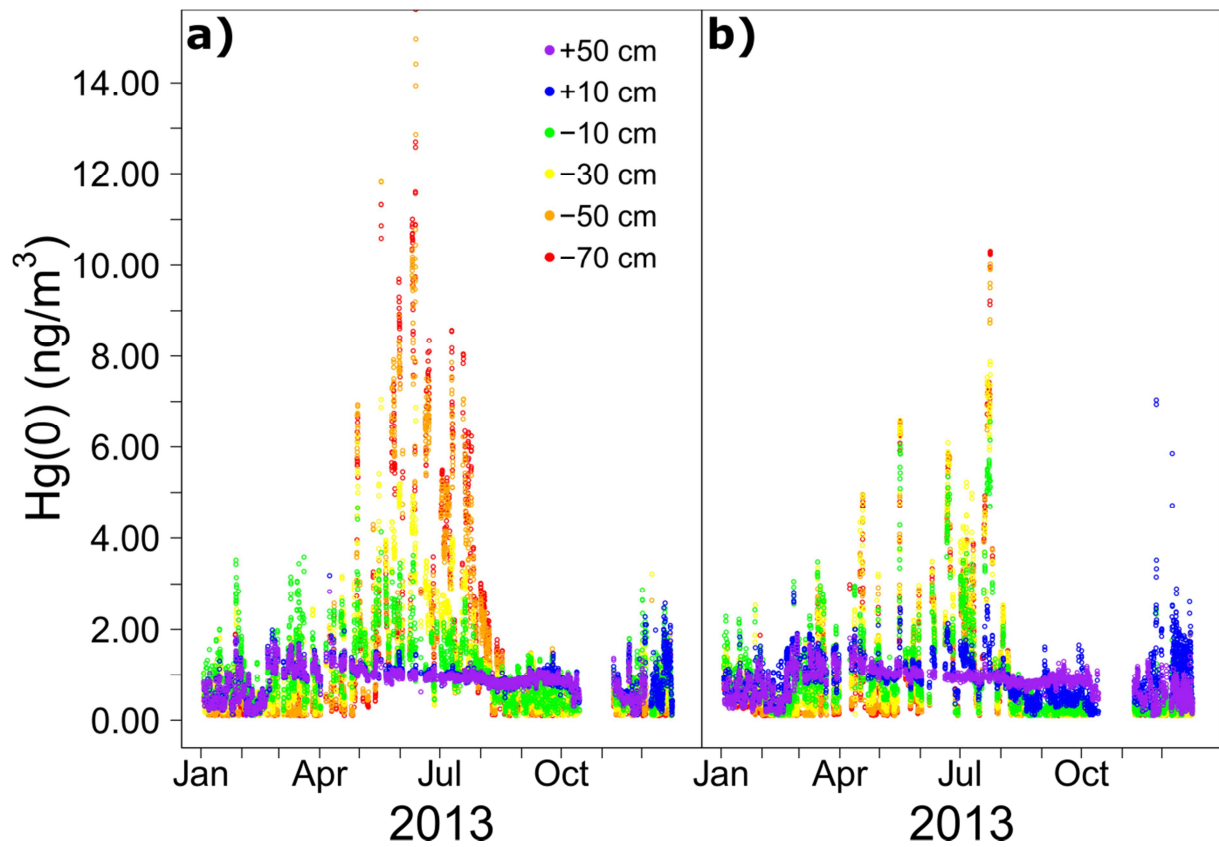
**Figure 7:** Hourly (local time) mean variation, along with the 95% confidence interval for the mean, of: **a)** Hg(0) concentration (in ng/m<sup>3</sup>) at 25 cm above the snow surface, **b)** downwelling shortwave (SW) radiation (in W/m<sup>2</sup>) according to the MAR model simulations, **c)** temperature (in °C) at 3 m above the snow surface, **d)** wind speed at 3 m above the snow surface (in m/s), **e)** planetary boundary layer (PBL) height (in m) according to the MAR model simulations, **f)** friction velocity ( $u_*$ , in m/s), and **g)** Eddy diffusivity ( $K$ , in m<sup>2</sup>/s) in summer (red), fall (green), winter (blue), and spring (purple). Note that the hourly mean variation of Hg(0) concentration in summer is similar at the three inlets of the meteorological tower.



**Figure 8:** Schematic diagram illustrating the processes that govern the Hg(0) budget at Concordia Station **a)** in summer under stable Planetary Boundary Layer (PBL) conditions, **b)** in summer under convective PBL conditions, **c)** in spring/fall, and **d)** in winter. In summer, Hg(0) is **continuously intensely** oxidized due to the high oxidative capacity of the boundary layer and a large amount of divalent mercury species deposit onto the snowpack. A fraction of deposited mercury can be reduced (the reducible pool, Hg<sub>r</sub>(II)) in the upper layers of the snowpack and subsequently reemitted to the atmosphere as Hg(0). Hg(0) emission from the snowpack maximizes near midday likely due to increased ventilation as a response to daytime **heating**. Oxidation of Hg(0) dominates in the deeper layers of the snowpack likely leading to the formation of a Hg(II) reservoir. In spring/fall, the balance of reduction and oxidation processes within the upper layers of the snowpack differs from summertime: oxidation dominates during the day, reduction at night. In winter, Hg(0) is produced in the deeper layers of the snowpack likely as a result of the reduction of Hg(II) species accumulated during the sunlit period. Ambient Hg(0) concentrations exhibit a 20 to 30% decrease through winter **possibly due to dry deposition of Hg(0)**.

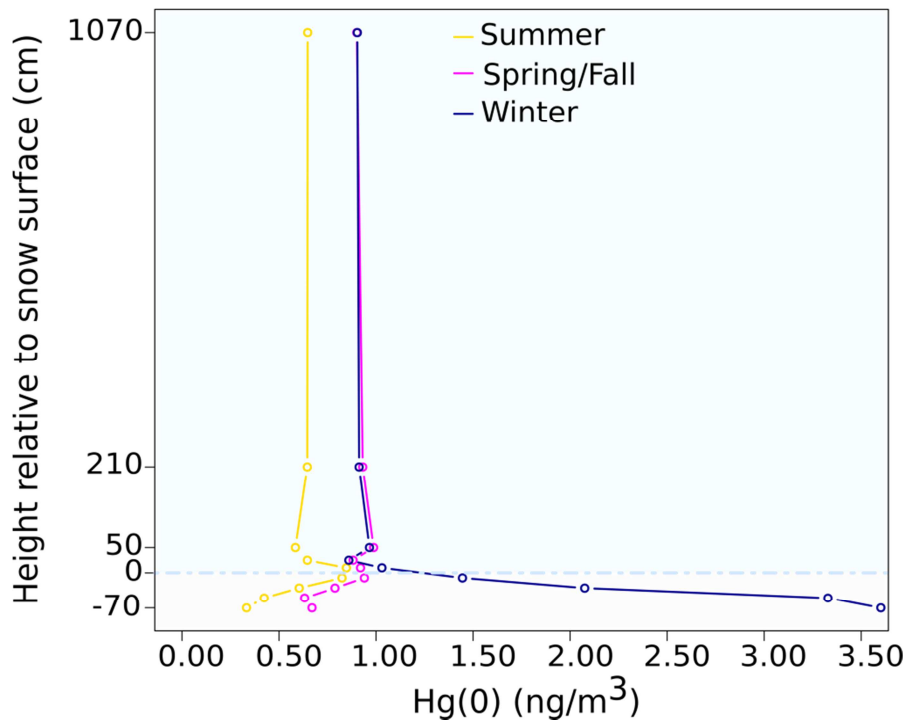


**Figure 9:** Year 2012 wintertime record of: **a)** hourly-averaged Hg(0) concentrations (in  $\text{ng}/\text{m}^3$ ) **at 500 cm above the snow surface**, **b)** Integrated Water Vapor (IWV,  $\text{kg}/\text{m}^2$ ), and **c)** Temperature (T,  $^{\circ}\text{C}$ ) at 10 m above ground level. Hg(0), temperature, and IWV increased from June 12 to 15 (in red) suggesting transport of moist and warm air masses originating from lower latitudes.

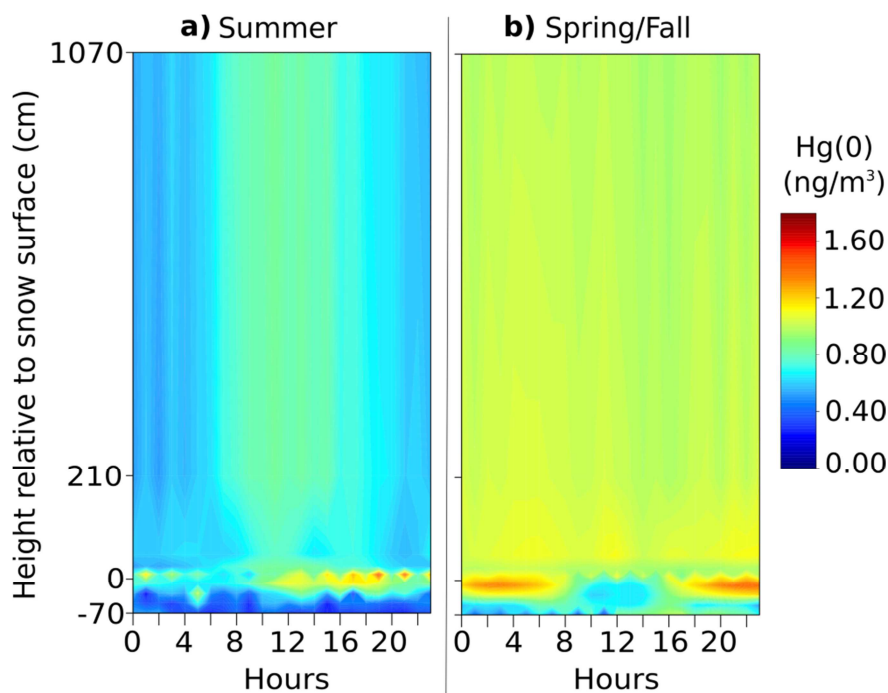


**Figure 10:** Annual variation of hourly-averaged Hg(0) concentrations (in  $\text{ng/m}^3$ ) in the snow interstitial air collected at the various inlets of the two snow towers: **a)** snow tower #1, **b)** snow tower #2. Note that we regularly experienced technical problems on snow tower #2 leading to missing values.





**Figure 11:** Mean Hg(0) concentration ( $\text{ng/m}^3$ ) measured at various heights above and below the snow surface (cm) at Concordia Station in summer (yellow), spring/fall (purple), and winter (dark blue) The horizontal light blue dashed line represents the snow surface.



**Figure 12:** Hourly (local time) mean atmospheric and interstitial air Hg(0) concentrations in **a)** summer, and **b)** spring/fall. The vertical axis is the height of measurement relative to the snow surface (in cm). Color contours show Hg(0) concentrations (in  $\text{ng}/\text{m}^3$ ). Concentrations at 25, 210, and 1070 cm above the snow surface were acquired on the meteorological tower while concentrations at 50, 10, -10, -30, -50 cm, and -70 cm were collected on snow tower #1. Data were cubic spline interpolated using software R.