# **Evidence for renoxification in the tropical marine boundary layer**

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### 13 **Abstract**

- We present two years of NO<sub>x</sub> observations from the Cape Verde Atmospheric Observatory
- located in the tropical Atlantic boundary layer. We find NO<sub>x</sub> mixing ratios peak around solar
- noon (at 20-30 pptV depending on season), which is counter to box model simulations that show
- 17 a midday minimum due to OH conversion of NO<sub>2</sub> to HNO<sub>3</sub>. Production of NO<sub>x</sub> via
- decomposition of organic nitrogen species and the photolysis of HNO<sub>3</sub> appear insufficient to
- 19 provide the observed noon-time maximum. A rapid photolysis of nitrate aerosol to produce
- 20 HONO and NO<sub>2</sub>, however, is able to simulate the observed diurnal cycle. This would make it the
- dominant source of NO<sub>x</sub> at this remote marine boundary layer site overturning the previous
- 22 paradigm of transport of organic nitrogen species such as PAN being the dominant source. We
- show that observed mixing ratios (Nov-Dec 2015) of HONO at Cape Verde (~3.5 pptV peak at
- solar noon) are consistent with this route for NO<sub>x</sub> production. Reactions between the nitrate
- 25 radical and halogen hydroxides which have been postulated in the literature appear to improve
- 26 the box model simulation of NO<sub>x</sub>. This rapid conversion of aerosol phase nitrate to NO<sub>x</sub> changes
- our perspective of the NO<sub>x</sub> cycling chemistry in the tropical marine boundary layer, suggesting a
- 28 more chemically complex environment than previously thought.

### 1 Introduction

- 2 The chemical environment in the remote marine boundary layer (MBL) is characterized by very
- low concentrations of nitrogen oxides ( $NO_x = NO + NO_2$ ) i.e. 10 to <100 pptV (Carsey et al.,
- 4 1997; Lee et al., 2009; Monks et al., 1998), high concentrations of water vapour and the presence
- of inorganic halogen compounds, resulting in net daytime ozone (O<sub>3</sub>) destruction (Dickerson et
- 6 al., 1999; Read et al., 2008; Sherwen et al., 2016; Vogt et al., 1999). This MBL loss of ozone
- 7 plays an important role in determining the global budget of ozone and the overall oxidizing
- 8 capacity of the region. Understanding the concentrations of NO<sub>x</sub> in these environments is thus
- 9 important for determining the global ozone budget, alongside wider atmospheric chemical
- impacts.
- NO<sub>x</sub> in the remote MBL has been attributed to a) long range transport and decomposition of
- species such as peroxy acetyl nitrates (PAN), organic nitrates, or HNO<sub>3</sub> (Moxim et al., 1996) b)
- shipping emissions (Beirle et al., 2004) c) a direct ocean source (Neu et al., 2008) and d) its
- direct atmospheric transport (Moxim et al., 1996). However, more recently the possibility of
- 15 'renoxification' by rapid nitrate photolysis on a variety of surfaces has garnered attention.
- Photolytic rate enhancements have been reported on aerosol nitrate (Ndour et al., 2009; Ye et al.,
- 2016b), urban grime (Baergen and Donaldson, 2013, 2016), natural and artificial surfaces (Ye et
- al., 2016a), and in laboratory prepared organic films and aqueous solutions (Handley et al., 2007;
- 19 Scharko et al., 2014; Zhou et al., 2003).
- 20 The oxidation of NO<sub>2</sub> to HNO<sub>3</sub> by OH is the predominant sink for NO<sub>x</sub> in the remote-MBL. NO<sub>x</sub>
- can also be converted into aerosol phase nitrate via the hydrolysis of N<sub>2</sub>O<sub>5</sub> (R2) (Evans, 2005)
- but this is a slow gas phase process in these low NO<sub>x</sub> environments. NO<sub>x</sub> can be returned
- 23 through HNO<sub>3</sub> photolysis (R3) or reaction with OH (R4) but in general these processes are again
- slow in the gas phase and so HNO<sub>3</sub> can deposit to the surface, be washed out by rain, or taken up
- by aerosol (R5).

$$26 \qquad \text{NO}_2 + \text{OH} + M \rightarrow \text{HNO}_3 + M \tag{R1}$$

27 
$$N_2O_5 + H_2O_{(aer)} \rightarrow 2HNO_{3(aer)}$$
 (R2)

$$1 HNO_3 + hv \rightarrow OH + NO_2 (R3)$$

$$2 \quad HNO_3 + OH \rightarrow NO_3 + H_2O \tag{R4}$$

3 
$$HNO_{3(g)} + aerosol \rightarrow HNO_{3(aer)}$$
 (R5)

- 4 More recently the production and subsequent hydrolysis of halogen nitrates (IONO<sub>2</sub>, BrONO<sub>2</sub>,
- 5 ClONO<sub>2</sub>) have been suggested to be a potentially important sink for NO<sub>x</sub> in the marine boundary
- 6 layer (Keene et al., 2007, 2009; Lawler et al., 2009; Pszenny et al., 2004; Sander et al., 1999)
- In this paper we investigate the budget of  $NO_x$  in the remote MBL using observations of  $NO_x$
- and HONO collected at the Cape Verde Atmospheric Observatory during 2014 and 2015. We
- 9 use a 0-D model of NO<sub>x</sub>, HO<sub>x</sub>, halogen, and VOC chemistry to interpret these observations and
- investigate the role that different NO<sub>x</sub> source and sink terms play.

# 11 2 Methodology

- 12 The Cape Verde Atmospheric Observatory (CVO), a WMO Global Atmospheric Watch (GAW)
- station, is located in the tropical North Atlantic (16.864, -24.868) on the island of São Vincente
- and is exposed to air travelling from the North East in the trade winds (Carpenter et al., 2010). In
- 15 general, the air reaching the station has travelled many days over the ocean since exposure to
- anthropogenic emissions, thus the station is considered representative of the remote marine
- boundary layer (Read et al., 2008). A large range of compounds are measured at the CVO
- (Carpenter et al., 2010), but we focus here on the NO and NO<sub>2</sub> continuous measurements,
- alongside HONO measurements that were made for 10 days in Winter 2015.

### 20 **2.1 NO and NO<sub>2</sub>**

- NO and NO<sub>2</sub> are measured by NO chemiluminescence (Drummond et al., 1985) coupled to
- 22 photolytic NO<sub>2</sub> conversion by selective photolysis at 385-395 nm as described by (Lee et al.,
- 23 2009; Pollack et al., 2011; Reed et al., 2016a, 2016b; Ryerson et al., 2000). A single
- 24 photomultiplier detector switches between 1 minute of chemiluminescent zero, 2 minutes of NO,
- and 2 minutes of NO<sub>x</sub> measurement.

- Air is sampled from a common 40 mm glass manifold (QVF) which draws ambient air from a
- 2 height of 10m above ground level. The manifold is downward facing into the prevailing wind at
- the inlet and fitted with a hood. The manifold is shielded from sunlight outside, and thermostated
- within the lab to  $30^{\circ}$ C to prevent condensation. Air is drawn down by centrifugal pump at ~ 750
- 5 L/min<sup>-1</sup> resulting in a sample flow speed of 10 m/s<sup>-1</sup> and a residence time to the NO<sub>x</sub> instrument
- of 2.3 seconds. Humidity and aerosol are reduced by two dead-end traps at the lowest points of
- 7 the manifold inside and outside the lab which are drained off regularly. Manifold sample flow,
- 8 humidity and temperature are recorded and logged continuously.
- 9 Air is sampled a 90° to the manifold flow through ¼ inch PFA tubing at 1 standard L per minute,
- being filtered through a 47mm, 0.22 μm mesh filter before entering the NO<sub>x</sub> analyser.
- The humidity of the sample gas is further reduced by a Nafion dryer (PD-50T-12-MKR, 11 Permapure), fed by a constant sheath flow of zero air (PAG 003, Eco Physics AG) which is also 12 filtered through a Sofnofil (Molecular Products) and activated carbon (Sigma Aldrich) trap. This 13 reduces sample humidity variability which affected NO sensitivity through chemiluminescent 14 quenching (Clough and Thrush, 1967) where sample humidity can vary from 60 to 90% 15 (Carpenter et al., 2010). Calibration for NO sensitivity and NO<sub>2</sub> converter efficiency occurs 16 every 73 hours by standard addition to ambient air as described by Lee et al., (2009); in this way 17 correction for humidity affecting sensitivity, and O<sub>3</sub> affecting NO<sub>2</sub> conversion efficiency are 18 unnecessary. Sensitivity drift between calibration is <2%, within the overall uncertainty of the 19 measurement. Zero air is also used to determine the NO<sub>2</sub> artifact signal which can arise when 20 NO<sub>x</sub> free air is illuminated at UV wavelengths due to photolysis of HNO<sub>3</sub> etc., adsorbed on the 21 walls of the photolysis cell (Nakamura et al., 2003; Pollack et al., 2011; Ryerson et al., 2000). 22 23 NO artifact correction is made by assuming it is equivalent to a stable night-time NO value in remote regions (Lee et al., 2009), away from any emission source, where NO should be zero in 24 the presence of O<sub>3</sub>. Reed et al., (2016b) showed that thermal interferences in NO<sub>2</sub> using this 25 technique may cause a bias in cold or temperate remote regions, but that in warm regions, such 26 as Cape Verde, the effect is negligible. Photolytic interferences such as BrONO<sub>2</sub> and HONO, and 27 inlet effects may also alter the retrieved NO or NO2 (Reed et al., 2016a, 2016b). These effects 28 29 are considered to be sufficiently small that the concentrations of NO and NO<sub>2</sub> can be determined

- within an accuracy of 5% and 5.9% respectively (Reed et al., 2016a, 2016b) at the (very low)
- levels present at CVO. The instrument having a zero count rate of  $\sim 1700$  Hz with 1  $\sigma$  standard
- deviation of that signal ~ 50 Hz this gives a precision of 7.2 pptV for 1 second data with typical
- 4 sensitivity over the measurement period of 6.9 cps/pptV. The resultant limits of detection for NO
- and NO<sub>2</sub> being 0.30 and 0.35 pptV when averaged over an hour.

## 6 **2.2 HONO**

- 7 Between 24<sup>th</sup> November and 3<sup>rd</sup> December 2015 a Long Path Absorption Photometer (LOPAP)
- 8 (Heland et al., 2001) was employed at CVO to provide an *in-situ* measurement of nitrous acid.
- 9 The LOPAP has its own thermostated inlet system with reactive HONO stripping to minimise
- losses so does not sample from the main lab manifold. The LOPAP inlet was installed on the
- 11 roof of a container lab ~ 2.5 m above ground level, unobstructed from the prevailing wind.
- 12 Calibration and operation of the LOPAP was carried out in line with the standard procedures
- described by Kleffmann and Wiesen, (2008). Specifically at CVO, the sampling conditions were
- set in order to maximise the sensitivity of the LOPAP, using a gas sampling flow rate of 2
- standard L per minute. A two point calibration was performed using a standard solution of nitrite
- 16 (NO<sub>2</sub>-) at concentrations of 0.8 and 10 µg L<sup>-1</sup>. To account for instrument drift, baseline
- measurements using an overflow of high-purity  $N_2$  were performed at regular intervals (8 hours).
- The detection limit of the LOPAP  $(2\sigma)$  was calculated by the variability during a typical baseline
- measurement under N<sub>2</sub> and was found to be 0.2 pptV. The relative error of the LOPAP was
- 20 conservatively set to 10% of the measured concentration.

### 2.3 Box Model

- 22 We use the Dynamically Simple Model of Atmospheric Chemical Complexity (DSMACC) box
- 23 model (Emmerson and Evans, 2009) to interpret the observed NO<sub>x</sub> measurements. We focus on
- 24 the summer season (June, July, and August) as this has the largest data coverage (N=153) and is
- out of dust season which extends through winter and spring (Carpenter et al., 2010; Fomba et al.,
- 26 2014; Ridley et al., 2014) and coincides with the lowest NO<sub>v</sub> mixing ratios (Carpenter et al.,
- 27 2010). The model is run for day 199 at 16.864°N, -024.868°W. We initialize the model with the
- mean observed H<sub>2</sub>O, CO, O<sub>3</sub>, VOCs (Carpenter et al., 2010; Read et al., 2012), 100 μm<sup>2</sup>/cm<sup>3</sup>

- aerosol surface area (Carpenter et al., 2010). We also initialise the model with 1.5 pptV of I<sub>2</sub> and
- 2.5 pptV of Br<sub>2</sub> to provide ~1.5 pptV of IO and ~2.5 pptV BrO during the day, consistent with
- the levels measured over 9 months at the CVO during 2007 (Mahajan et al., 2010; Read et al.,
- 4 2008). We use the average diurnal cycle of the measured HONO concentrations, described
- 5 above. Solar radiation at this location in the tropics shows little seasonal variation, hence
- 6 photolysis rates are similar in summer and autumn. This measurement period was also free of
- dust influence. We assume clear sky conditions for photolysis. The meteorological parameters
- 8 pressure, temperature, relative humidity, and boundary layer height are set to median values
- 9 reported by Carpenter et al., (2010). Boundary layer height is fixed at 713m as no overall
- seasonal or diel pattern is evident in boundary layer height at Cape Verde (Carpenter et al.,
- 11 2010). This is expected at a site representative of the marine boundary layer, which has almost
- 12 no island effects (except for very rare instances of wind outside the northwesterly sector, which
- are excluded). Thus, we discount any influence from boundary layer height changes on the
- 14 diurnal cycles presented.
- 15 The unconstrained model is run forwards in time until a stable diurnal cycle is attained; ~ 3 days.
- A full description of the model chemistry is provided in the supplementary material. The base
- case chemistry has only gas phase sources plus gas phase and deposition sinks for NO<sub>x</sub> as
- described in the supplementary material.

### 19 **Results and discussion**

# 20 3.1 Diurnal cycles in NO<sub>x</sub> and HONO

- Figure 1 shows the measured mean diurnal cycles of NO, NO<sub>2</sub>, NO<sub>x</sub>, and O<sub>3</sub> observed in each
- season (Meteorological Spring, Summer, Autumn, and Winter) during 2014 and 2015. Every
- season shows a strong diurnal cycle in NO, peaking after solar noon at around ~13:00 to 14:00.
- 24 The diurnal cycle of NO<sub>2</sub> is much less pronounced but also exhibits weak maxima in the early
- 25 afternoon. Overall this leads to a maximum in NO<sub>x</sub> during the day. This behaviour is consistent
- throughout the year and air mass, though not necessarily on a "day-to-day" basis.

- 1 The observed diurnal cycle in NO<sub>x</sub> is hard to explain with conventional chemistry. The increase
- 2 in night time NO<sub>x</sub> suggests a continuous source but the maximum around noon suggests a
- photolytic source. Given the predominant NO<sub>x</sub> sink is reaction with OH to form HNO<sub>3</sub>, it would
- 4 be expected that there would be a minimum in NO<sub>x</sub> during the day rather than a maximum.
- 5 Similar observations have been reported previously (Monks et al., 1998) at the Cape Grim
- 6 Baseline Air Pollution station (-40.683, 144.670), a comparably remote site in the southern
- 7 hemisphere, and during the Atlantic Stratocumulus Transition Experiment (ASTEX) cruise
- 8 (~29°N, 24°W) which reported similar daytime NO<sub>x</sub> production (Carsey et al., 1997). The
- 9 observed behaviour in the CVO NO<sub>x</sub> was historically attributed to atmospheric thermal
- decomposition of NO<sub>v</sub> species (Lee et al., 2009).
- Figure 2 shows the average diurnal cycle at CVO of measured HONO concentrations. The data
- exhibits a strong daytime maximum peaking at noon local time (Solar noon ~13:20) and reaching
- 13 near zero at night, indicating a photolytic source. HONO is lost through deposition, photolysis
- and reaction with OH, whilst night time build-up often observed (Ren et al., 2010; VandenBoer
- et al., 2014; Zhou et al., 2002), here HONO appears to reach a steady state concentration of
- ~0.65 pptV throughout the night. This pseudo steady state behaviour of nocturnal HONO has
- previously been reported in the polluted marine boundary layer by Wojtal et al., (2011), albeit
- 18 reporting much higher HONO mixing ratios.
- 19 Daytime production of HONO is similarly hard to reconcile if its formation by the homogeneous
- 20 OH + NO reaction (or other gas-phase HO<sub>x</sub>-NO<sub>x</sub> chemistry, e.g. Li et al., (2014)). With NO
- 21 mixing ratios below 5 pptV, OH measured peaking at ~ 0.25 pptV during the RHaMBLe
- campaign (Carpenter et al., 2010; Whalley et al., 2010) and a maximum noontime jHONO of 1.2
- $\times 10^{-3} \text{ s}^{-1}$ , a steady state HONO mixing ratio of ~ 0.04 pptV is found  $(k_{\text{(OH + NO)}} = 7.4 \times 10^{-12})$
- 24 mol.cm<sup>-3</sup> s<sup>-1</sup>). An additional daytime source of HONO must be present to explain the observed
- 25 concentrations.
- 26 Both the long-term NO<sub>x</sub> and the short-term HONO observations made at CVO are cannot be
- 27 explained with purely gas phase chemistry. Both datasets show daytime maxima indicative of a

- photolytic source of NO<sub>x</sub> and HONO, whereas gas phase chemistry would predict minima in
- 2 NO<sub>x</sub> during daytime and two orders of magnitude less HONO.

# 3 3.2 Box modelling of $NO_x$ sources

- 4 Using the box model (section 2.3) we explore the observed diurnal variation in NO<sub>x</sub> and
- 5 understand the role of different processes. Classically, the predominant source of NO<sub>x</sub> in remote
- 6 regions is considered to be the thermal decomposition of compounds such as peroxyacetyl nitrate
- 7 (PAN) which can be produced in regions of high NO<sub>x</sub> and transported long distances (Fischer et
- al., 2014; Jacobi et al., 1999; Moxim et al., 1996). We consider a source of PAN which descends
- 9 from the free troposphere and then thermally decomposes to NO<sub>2</sub> in the warm MBL. The main
- sink of NO<sub>x</sub> is conversion to HNO<sub>3</sub>, which is slightly counterbalanced by a slow conversion of
- 11 HNO<sub>3</sub> back into NO<sub>x</sub> through gas phase photolysis or reaction with OH. Figure 3 shows the
- model with a source of PAN which results in mixing ratios of 5-8 pptV, consistent with the few
- measurements made in the marine boundary layer, most notably by Jacobi et al., (1999) who
- measured levels from <5 to 22 pptV in the tropical Atlantic, and Lewis et al., (2007) who
- reported PAN mixing ratios of ~10 pptV in the remote mid-Atlantic MBL.
- 16 It is evident from the base case model results shown in Fig. 3 that the model fails to calculate the
- 17 NO<sub>x</sub> and HONO diurnal cycles. Modelled NO<sub>x</sub> concentrations do increase during the night,
- consistent with the observations, but the model's minimum for NO<sub>x</sub> occurs during the day when
- the observations show a maximum. The modelled and measured HONO is also shown in Fig. 3,
- both peaking during midday with observations reaching 3.5 pptV whilst the model simulates
- 21 only ~ 0.2 pptV underestimating HONO at all times. It is clear that long-range transport and
- thermal decomposition of NO<sub>v</sub> species such as PAN alone cannot explain the NO<sub>x</sub> diurnal at
- 23 Cape Verde. A PAN-type continuous thermal decomposition forming NO<sub>x</sub> would be inconsistent
- 24 with the diurnal maximum in NO<sub>x</sub> which is observed. The NO<sub>x</sub> source necessary to support a
- 25 noon time maximum would have to show a strong day-time maximum to counter the strong
- diurnal in the sink.
- 27 This need for a diurnal cycle in the NO<sub>x</sub> source also suggests that the shipping source of NO<sub>x</sub> is
- unlikely to explain the diurnal cycle. The dominant source of ship  $NO_x$  in the region occurs from

- the large container ships which pass the region on their way to South America or the Cape of
- 2 Good Hope. It is unlikely that these emissions are systematically higher during the day than
- during the night and thus are unlikely to explain the observed diurnal signal.
- 4 There have been a number of studies which have identified much faster photolysis of nitrate
- 5 within and on aerosol, than for gas phase nitric acid. These include studies using real-world
- 6 natural and artificial surfaces (Baergen and Donaldson, 2013; Ye et al., 2016a, 2016b),
- 7 laboratory substrates such as organic films and aqueous acid solutions (Handley et al., 2007;
- 8 Scharko et al., 2014; Zhou et al., 2003), aerosol nitrate (Ndour et al., 2009; Ye et al., 2016b), and
- 9 a model estimate (Cohan et al., 2008). These studies have found that particulate nitrate
- photolysis rates can be up to ~3 orders of magnitude greater than gas phase HNO<sub>3</sub> photolysis in
- marine boundary layer conditions (Ye et al., 2016b). There is also broad agreement between
- different studies on the main photolysis product being nitrous acid (HONO) with NO<sub>2</sub> as a
- secondary species. The product ratio appears dependent on aerosol pH, with HONO production
- only occurring at low pH (Scharko et al., 2014). This is shown in reaction (R6) as particulate
- nitrate (p-NO<sub>3</sub>) photolysing to HONO and NO<sub>2</sub> in a ratio x:y.

16 
$$p-NO_3 + hv \rightarrow xHONO + yNO_2$$
 (R6)

- 17 There is also evidence that the photolysis rate is positively correlated with relative humidity
- 18 (Baergen and Donaldson, 2013; Scharko et al., 2014). As such, particulate nitrate photolysis rates
- 19 appear to increase with increasing aerosol acidity and relative humidity. With the CVO site
- 20 experiencing relative humidity of 79 % on average (Carpenter et al., 2010) and aerosol
- containing a significant acidic fraction (Fomba et al., 2014), particulate nitrate photolysis could
- have a role to play in the NO<sub>x</sub> budget at Cape Verde.
- In order to explore the implications for Cape Verde NO<sub>x</sub> chemistry, we re-ran the base model
- removing the PAN source but including particulate nitrate (p-NO<sub>3</sub>) photolysis (R6) leading to
- 25 HONO and NO<sub>2</sub> production, scaled to the gas phase photolysis of HNO<sub>3</sub>. This parameterisation
- 26 nominally represents photolysis of nitrate within and on aerosol, however conceptually includes
- 27 any additional surface production of HONO and NO<sub>2</sub>. We use an aerosol phase concentration of
- nitrate of 1.1 μg m<sup>-3</sup> (equivalent to 400 pptV), which is the mean concentration found in PM10

- aerosol at Cape Verde, with little apparent seasonal variability (Fomba et al., 2014; Savarino et
- al., 2013). The branching ratio of HONO to  $NO_2$  production from reaction 6 (x and y) was set to
- 2:1 in line with the findings of Ye et al., (2016b). We scale the p-NO<sub>3</sub> photolysis rate to gas
- 4 phase HNO<sub>3</sub> photolysis by factors of 1, 10, 25, 50, 100, and 1000. The study of Ye et al., (2016b)
- 5 describes enhancements of up to ~300 fold. The impact on the summer months is shown in Fig.
- 6 4.
- 7 Including the photolysis of aerosol nitrate changes both the mean concentration and diurnal cycle
- 8 of NO<sub>x</sub> significantly. The diurnal NO<sub>x</sub> is now flat or peaks during the daytime, more consistent
- 9 with observations. We find the best approximation is achieved when the rate of particulate nitrate
- 10 photolysis is ~10 times that of HNO<sub>3</sub> which is broadly consistent with laboratory based
- observations (Zhou et al., 2003). A wide variability of p-NO<sub>3</sub> photolysis rates on different
- surfaces are reported (Ye et al., 2016a), thus the photolysis of nitrate is uncertain and likely to be
- variable with aerosol composition. In all particulate nitrate photolysis-only scenarios, depicted in
- Fig. 4 and Fig. 5, it is evident that p-NO<sub>3</sub> photolysis alone doesn't give the observed increase in
- night time NO<sub>x</sub> observations. Conversely the PAN only scenario is insufficient to sustain daytime
- NO<sub>x</sub>. It is therefore likely that the actual source of NO<sub>x</sub> is a combination of PAN entrainment
- from the free troposphere and particulate nitrate photolysis.
- 18 Combining the free-tropospheric source of PAN, and the photolysis of particulate nitrate at a rate
- of 10 times the gas phase HNO<sub>3</sub> photolysis (Fig. 5) results in a model simulation with roughly
- 20 twice as much NO<sub>x</sub> both at night and during daylight but a roughly flat diurnal profile. Simulated
- 21 HONO peaks at local noon, similar to the observations though underestimates the mid-day peak.
- 22 Nocturnal HONO mixing ratios agree with observations being non-zero at ~0.5 pptV.
- 23 Introduction of an additional source of NO<sub>x</sub> is able to roughly produce a flat diurnal cycle,
- 24 though is not able to simulate a definite peak of NO<sub>x</sub> during daytime. With the addition of a
- source and no change in sinks for NO<sub>x</sub> this is unsurprising and leads to relative over estimation
- of NO<sub>x</sub> particularly at night. This is therefore likely that one or more NO<sub>x</sub> sinks are absent from
- 27 the base simulation which must be explored further.

### 3.3 NO<sub>x</sub> sinks

- Aside from loss to HNO<sub>3</sub> directly through reaction with OH (R1) NO<sub>x</sub> is also lost to nitrate by
- 2 reaction with halogen oxides (XO) forming a halogen nitrates (R7) (Keene et al., 2009). Read et
- al., (2008) showed how halogen oxides mediate ozone formation and loss at Cape Verde thus
- 4 also exerting an indirect effect on NO<sub>x</sub>.

5 
$$XO + NO_2 + M \rightarrow XONO_2 + M$$
 (R7)

- 6 Figure 6 shows the rates of production and loss analysis for NO<sub>x</sub> in this simulation with both
- 7 PAN thermal decomposition and particulate nitrate photolysis. The largest net source of NO<sub>x</sub>
- 8 after net sinks (such as halogen nitrate cycling) are removed is nitrate photolysis to HONO and
- 9 NO<sub>2</sub>. The major net sink of NO<sub>x</sub> is the formation of nitric acid by reaction of NO<sub>2</sub> with OH. –
- However, uptake of HNO<sub>3</sub> onto aerosol, and subsequent rapid (compared to gas phase HNO<sub>3</sub>)
- photolysis acts to balance this.
- The pronounced drop in modelled NO<sub>2</sub> at sunrise is due to production of halogen nitrates (R7)
- when HOX rapidly photolyses to produce XO which can then react with NO<sub>2</sub> to produce
- 14 XONO<sub>2</sub>. XO is formed quickly and spikes in concentration leading to the rapid loss of NO<sub>2</sub>. This
- feature is not observed in the  $NO_x$  observations during any season.
- The diagnostics in Figure 6 show the role of the different sinks of NO<sub>x</sub>. In that simulation these
- are dominated by the gas phase reaction between NO<sub>2</sub> and OH but with the rapid formation and
- subsequent hydrolysis of BrONO<sub>2</sub> and IONO<sub>2</sub> (R8) playing a major role (Sander et al., 1999).
- 19 The uptake coefficient (γ) of halogen nitrates onto aerosol therefore could have a strong
- 20 influence on the NO<sub>x</sub> diurnal.

21 
$$XONO_2 + H_2O_{(aer)} \rightarrow HNO_{3(aer)} + X^+ + OH^-$$
 (R8)

- We perform a sensitivity analysis on the effect of the uptake coefficients on the NO<sub>x</sub> and XO
- diurnals. We do this in a particulate nitrate photolysis only simulation, without PAN, to isolate
- 24 the effect of XONO<sub>2</sub> hydrolysis on nitrate-NO<sub>x</sub> cycling. Figure 7 shows the effect of changing  $\gamma$
- of XONO<sub>2</sub> (X = Br, I) within recommended ranges (Burkholder et al., 2015; Saiz-Lopez et al.,
- 26 2008) on Saharan dust and sea salt the predominant coarse mode aerosol by mass at Cape
- 27 Verde (Carpenter et al., 2010; Fomba et al., 2014), ranging from 0.02 to 0.8.

- 1 Increasing the γ of XONO<sub>2</sub> from 0.02 (the low end of recommended values) to 0.1 results in
- 2 small changes to both the NO<sub>x</sub> and XO diurnals. The loss of NO<sub>x</sub> at sunrise becomes more
- 3 pronounced whereas the XO diurnals become slightly more 'square' or 'top-hat' as per the
- observations of Read et al., (2008). Increasing the  $\gamma$  to the upper extreme ( $\gamma = 0.8$ ) results in a
- 5 spike in BrO at sunrise, which consumes the majority of NO<sub>2</sub> though formation of BrONO<sub>2</sub>. No
- 6 combination of uptake coefficients can completely reproduce the characteristic XO diurnals due
- to poor constraints on heterogeneous halogen chemistry (Abbatt et al., 2012) in addition to gaps
- 8 in understanding of gas phase halogen chemistry (Simpson et al., 2015).
- 9 The effect on the  $NO_x$  diurnal of changing  $\gamma$  is clear in that greater uptake coefficients
- recommended by Burkholder et al., (2015) result in objectively worse simulation of both the NO<sub>x</sub>
- and XO diurnals. It is therefore likely that information is lacking from the  $XO NO_x$  chemistry
- scheme as it is currently known.

### 13 3.4 HOI/HOBr - NO<sub>x</sub> chemistry

- Recently, IO recycling by reaction with NO<sub>3</sub> has been proposed by Saiz-Lopez et al., (2016) who
- calculated that the reaction (R9) of HOI + NO<sub>3</sub> producing IO and HNO<sub>3</sub> has a low enough
- activation energy and fast enough rate constant to be atmospherically relevant in the troposphere.

17 HOI + NO<sub>3</sub> 
$$\rightarrow$$
 IO + HNO<sub>3</sub> :  $k = 2.7 \times 10^{-12} (300/T)^{2.66}$  (R9)

- This mechanism provides a route to nitric acid, and thus particulate nitrate at night, whilst also
- leading to nocturnal IO production leading to loss of NO<sub>2</sub> by IONO<sub>2</sub> formation.
- 20 Including this new reaction and re-running the model leads to a diurnal profile of IO much more
- 21 representative of the observations. This however introduces a more pronounced loss of NO<sub>x</sub> at
- sunrise and sunset, and also results in NO<sub>x</sub> peaking during the day which fits better with the
- observations as shown in Fig. 8. HONO is still underestimated during daytime though nocturnal
- values agree well.
- 25 The inclusion of this HOI + NO<sub>3</sub> reaction reproduces the general NO<sub>x</sub> and O<sub>3</sub> diurnals more
- 26 closely than without i.e. a daytime maximum in NO<sub>x</sub>. There are also effects on the halogen oxide

- behaviour. The simulated BrO has a flatter profile, which more closely matches the observations.
- 2 However, modelled IO is now non-zero at night and the sunrise build-up and sunset decay still
- 3 occurs more abruptly than the observations.
- 4 Although the NO<sub>x</sub> and O<sub>3</sub> diurnals are reproduced more closely with this new chemistry, there is
- 5 still disagreement with the observed NO<sub>x</sub> diurnal at sunrise and sunset especially indicating a
- 6 missing reaction or reactions. To best approximate the observed diurnal behaviour an analogous
- 7 HOBr + NO<sub>3</sub> night time reaction (R10) was introduced with a rate 10 times that of HOI + NO<sub>3</sub>
- 8 as calculated by Saiz-Lopez et al., (2016b)

9 HOBr + NO<sub>3</sub> 
$$\rightarrow$$
 BrO + HNO<sub>3</sub> :  $k = 2.7 \times 10^{-11} (300/T)^{2.66}$  (R10)

- 10 This results in an improved reproduction of the observed NO<sub>x</sub> diurnal, Fig. 9. This is a purely
- speculative representation in order to reproduce the observed NO<sub>x</sub> diurnal and highlights how
- some mechanistic knowledge of NO<sub>x</sub>-halogen-aerosol systems is still missing.
- With HOX + NO<sub>3</sub> chemistry included in the model as in Fig. 9, significant loss of NO<sub>x</sub> at sunrise
- and sunset is eliminated and agreement is improved over any previous simulation. Greater
- 15 HONO production is also simulated, with up to ~ 3.0.pptV predicted in line with the
- observations shown in Fig. 2. Halogen oxide modelled diurnal cycles remain broadly consistent
- with observations. Diagnosis of the net production and loss terms for  $NO_x$  reveal that nitrate
- photolysis to HONO or NO<sub>2</sub> contribute ~ 80% of all NO<sub>x</sub> with decomposition of PAN
- contributing the remainder. Major net sinks of NO<sub>x</sub> are shown to be reaction with halogen
- 20 hydroxides and OH to form HNO<sub>3</sub>. Nitric acid is then taken up on surfaces and recycled to NO<sub>x</sub>
- 21 through photolysis to  $NO_2$  and HONO
- 22 The improvement can be better understood by diagnosing the modelled NO<sub>v</sub> distribution. In Fig.
- 23 10 the distribution of PAN, IONO<sub>2</sub>, BrONO<sub>2</sub>, N<sub>2</sub>O<sub>5</sub>, NO<sub>3</sub> and particulate nitrate (p-NO<sub>3</sub>) is
- shown for the base case scenario (where entrained PAN is the sole source of NO<sub>x</sub> in the MBL),
- 25 for the particulate nitrate photolysis case including HOI + NO<sub>3</sub> chemistry, and the same but also
- including HOBr + NO<sub>3</sub> chemistry. The major feature changing through the different simulations
- is the magnitude and shape of the BrONO<sub>2</sub> diurnal. From the base run to the inclusion of HOI +

- 1 NO<sub>3</sub> chemistry and particulate nitrate photolysis a major increase in BrONO<sub>2</sub> mixing ratio is
- 2 expected at sun rise and sun set. It is this rapid production of BrONO<sub>2</sub> which consumes NO<sub>x</sub>
- resulting in the sharp dips at these times not seen in the observations. In the HOBr &  $HOI + NO_3$
- 4 and particulate nitrate photolysis case these features are eliminated and halogen nitrates do not
- 5 spike at sunrise or sunset. Nitrate is shown to be conserved by hydrolysis of halogen nitrates on
- surfaces and uptake of nitric acid. This cycling leads to a NO<sub>x</sub> diurnal profile which is more
- 7 representative of the observations.
- 8 In models which included nitrate photolysis a strong diurnal cycle in particulate nitrate presents
- 9 which is depleted during the day and recycles at night being conserved overall. The daily average
- concentration remains constant however in line with integrating filter sample study of Fomba et
- 11 al., (2014).
- Unsurprisingly, the inclusion of  $HOX + NO_3$  chemistry results in lower mixing ratios of  $NO_3$  at
- night. In all cases N<sub>2</sub>O<sub>5</sub> (in black) is effectively zero at all times due to very low NO<sub>x</sub> mixing
- ratios in this pristine environment and the relatively high ambient temperatures (24.5 °C) where
- the  $N_2O_5$  lifetime is ~ 3 s<sup>-1</sup>. This precludes  $N_2O_5$  channels to  $NO_x$  (and ultimately nitrate),
- 16 consistent with the experimental findings of Savarino et al., (2013) at Cape Verde who found
- isotope ratios which were incompatible with high production rates of HNO<sub>3</sub> from N<sub>2</sub>O<sub>5</sub>
- hydrolysis, and concluded that N<sub>2</sub>O<sub>5</sub> and nitryl compound (ClNO<sub>2</sub>, BrNO<sub>2</sub>) levels in this region
- are very low. This is consistent with our own and other studies modelling the pristine marine
- boundary layer at Cape Verde of Sommariva and Von Glasow, (2012). This is in contrast with
- 21 more polluted regions where N<sub>2</sub>O<sub>5</sub> has been shown to be a route to NO<sub>x</sub> and ClNO<sub>2</sub> (Kim et al.,
- 22 2014).
- 23 The agreement in modelled and observed NO<sub>x</sub> improves and the modelled values fall within the
- 24 error of the observations. Additionally the approximate BrO diurnal is achieved without the
- characteristic 'horns', however replicating IO observations is still problematic.
- 26 The effect of dramatically changing NO<sub>x</sub> diurnal could be expected to have an effect on OH and
- 27 HO<sub>2</sub> mixing ratios. The difference between the base model case, where PAN decomposition is

- the dominant daytime source, and the final model where the NO<sub>x</sub> is more accurately described by
- 2 particulate nitrate photolysis and HOX + NO<sub>3</sub> chemistry is shown in Fig. 11.
- In the case of OH the change from the base model to the final model is an increase of 3.3% at the
- 4 maximum, for HO<sub>2</sub> the increase is a more significant 8.6% (or 1.7 pptV), however this is well
- 5 within the uncertainty of measured values (Whalley et al., 2010). Figure 11 shows that even with
- dramatic changes in the NO<sub>x</sub> simulation, the OH and HO<sub>2</sub> changes very little comparatively
- 7 despite increased daytime HONO production.
- 8 From these simulations it would appear that the photolysis of surface adsorbed nitrate may be the
- 9 dominant source of NO<sub>x</sub> into the marine boundary layer around Cape Verde. Photolysis of
- aerosol nitrate, or nitrate in solution would be capable of producing a diurnal cycle in NO<sub>x</sub> which
- was consistent with the observations when HOX + NO<sub>3</sub> chemistry is considered also. Whilst
- agreement between model and observation is improved there is a clear gap in understanding the
- halogen-NO<sub>x</sub>-aerosol system in the remote marine boundary layer.

### 4 Conclusions

- Fast aerosol nitrate photolysis is shown to be likely the primary source of  $NO_x$  in the remote
- tropical Atlantic boundary layer. A 0-D model replicated the observed halogen, O<sub>3</sub>, OH, NO<sub>x</sub> and
- 17 HONO levels when including particulate nitrate photolysis at a rate of ~10 times that of gas
- phase nitric acid, consistent with previous laboratory measurements. Model optimisation shows
- 19 that this new source of daytime NO<sub>2</sub> is compatible with observations and currently known
- 20 chemistry at night and at mid-day, but that at sunrise and sunset there is disagreement due to the
- treatment of halogen oxides at these times. Recently suggested halogen hydroxide + nitrate
- 22 radical chemistry may provide better agreement between model and observation if theoretical
- reactions can be substantiated.

### 1 Data Availability

- 2 We thank the NASA Jet Propulsion Laboratory (Burkholder et al., 2015) for providing
- 3 comprehensive rate and uptake coefficient data for atmospheric compounds, which can be found
- 4 at http://jpldataeval.jpl.nasa.gov.
- 5 All data used in this work is available from the British Atmospheric Data Centre (BADC)
- 6 http://badc.nerc.ac.uk and is included as a .csv file in the supplementary information. The
- 7 DSMACC model is available from https://github.com/barronh/DSMACC and a full description
- 8 of the model can be found in the supplementary information.
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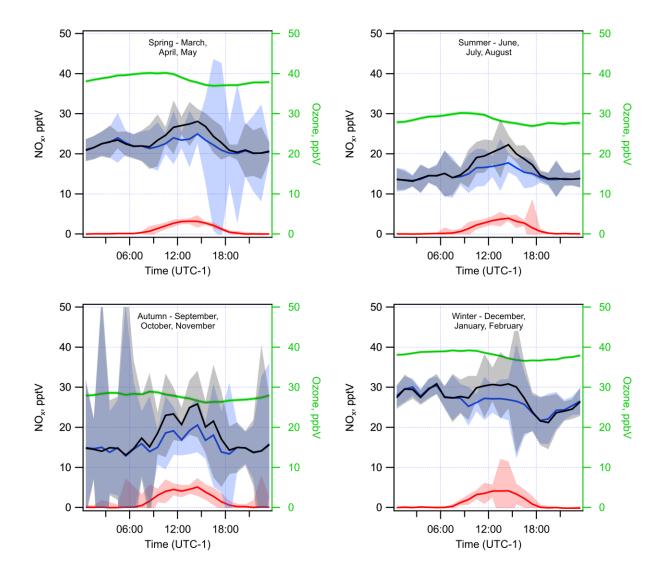


Figure 1. The observed seasonal diurnal cycles in NO, NO<sub>2</sub>, NO<sub>x</sub>, and O<sub>3</sub> at the CVO GAW station during 2014 and 2015. NO is shown in red, NO<sub>2</sub> in blue, NO<sub>x</sub> in black, and O<sub>3</sub> in green. Shaded areas indicate the standard error of data.

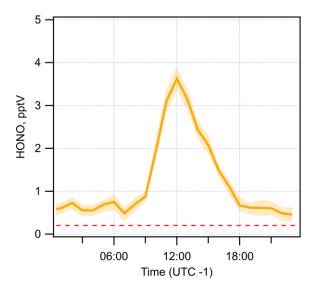


Figure 2. The observed average HONO diurnal measured at CVO during 24<sup>th</sup> November – 3<sup>rd</sup>
December 2015. Shaded area indicates standard deviation and cumulative error of data. Dashed red line shows the HONO limit of detection.

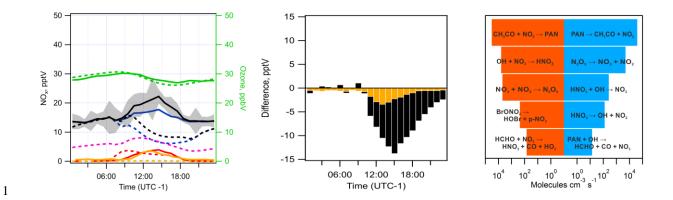


Figure 3. Left shows the measured (solid lines) and modelled (dashed)  $NO_x$  and HONO diurnal behaviour at the CVO GAW station where the dominant source of  $NO_x$  is a source of PAN descending from the upper troposphere having been transported from polluted regions. Shaded areas are standard error of the observations ( $NO_x$  N = 153 HONO, N = 10). $O_3$  – green;  $NO_x$  – black;  $NO_2$  – blue; NO – red; HONO – yellow; PAN – pink. Centre is the difference between measured and modelled  $NO_x$  (black) and HONO (Yellow). Right shows the rates of production and loss of NO and  $NO_2$  from sources listed in descending order of contribution over a 24 hour period accounting for >95% of the total.

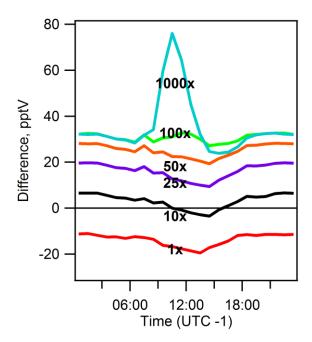


Figure 4. The difference between measured and modelled of  $NO_x$  at CVO during summer months when photolysis of nitrate is considered. The rate of particulate nitrate photolysis has been scaled to the rate of  $HNO_3$  photolysis by factors of 1, 10, 25, 50, 100, and 1000.

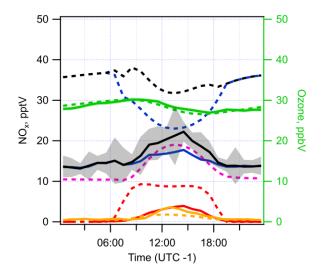
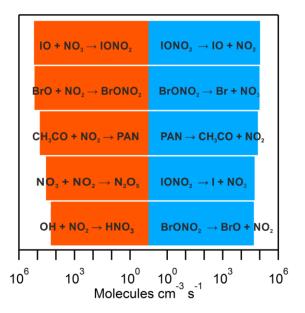


Figure 5. The modelled diurnal profile of  $NO_x$  at CVO during summer months when photolysis of nitrate (set at  $10\times$  the gas phase  $HNO_3$  photolysis) and a tropospheric PAN source are considered. Shaded areas for  $NO_x$  are the standard error of the observation.  $O_3$  – green;  $NO_x$  – black;  $NO_2$  – blue; NO – red; HONO – yellow; PAN – pink.



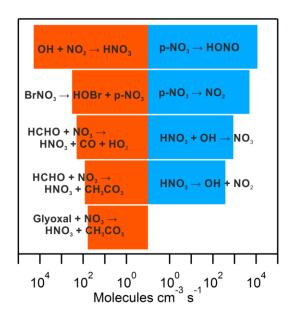


Figure 6. Left is the total production and loss analysis for NO<sub>x</sub> of the combined model of particulate nitrate photolysis and PAN decomposition over 24 hours. Right is the same analysis discarding the major balanced sinks of fast cycling short lived species.

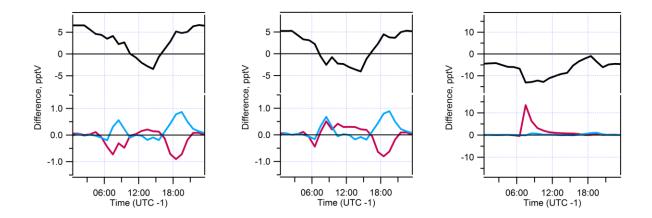


Figure 7. Sensitivity analysis of the effect of changing reactive uptake co-efficients ( $\gamma$ ) of halogen nitrates , XONO<sub>2</sub> (X = Br, I), on NO<sub>x</sub> (top) and XO (bottom) diurnal behaviour during summer months at CVO. The difference between measured and modelled values is plotted. Particulate nitrate photolysis is set at 10 times the rate of gaseous HNO<sub>3</sub>.

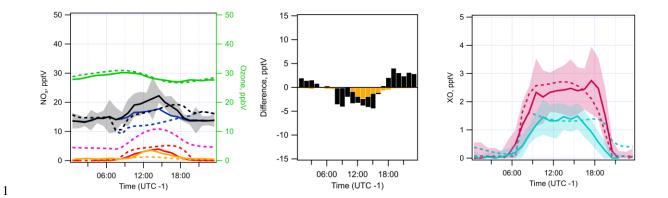


Figure 8. Left is the modelled  $NO_x$  and HONO diurnal cycle for the CVO site during summer months with the inclusion of night time HOI chemistry. Centre shows the difference between measured and modelled values of  $NO_x$  (black) and HONO (yellow). Right is the observed (adapted from Read et al., (2008)) and modelled IO and BrO. Observations are solid lines whilst modelled values are shown dashed. Shaded areas are standard error of the observation.  $O_3$  – green;  $NO_x$  – black;  $NO_2$  – blue; NO – red; HONO – yellow; PAN – pink; IO – turquoise; BrO – purple.

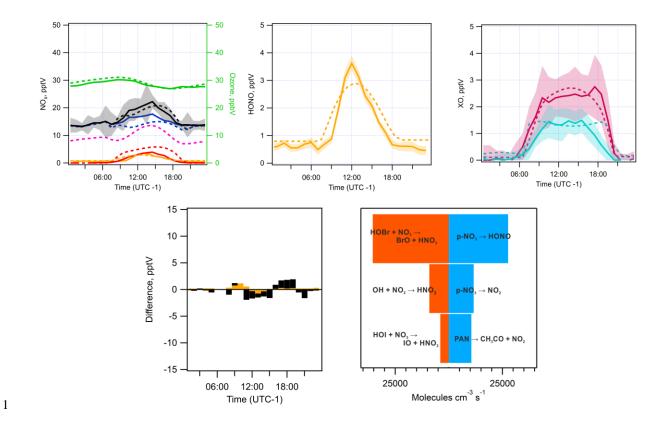


Figure 9.  $NO_x$ ,  $O_3$ , HONO at CVO during summer months compared to model values when HOI and HOBr +  $NO_3$  is included in the simulation. Top left:  $NO_x$ ,  $O_3$ , HONO, and PAN diurnal cycles. Top middle: expanded view of modelled and measured HONO. Top right: observed (adapted from Read et al., (2008)) and modelled IO and BrO. Bottom left: difference between modelled and measured  $NO_x$  and HONO. Bottom right: the net production and loss analysis for  $NO_x$  in this simulation.  $O_3$  – green;  $NO_x$  – black;  $NO_2$  – blue; NO – red; HONO – yellow; PAN – pink; IO – turquoise; PAN – purple. Measured values are solid lines, modelled values are dashed.

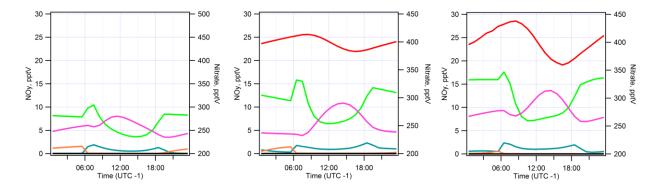


Figure 10. Shown are  $NO_y$  diurnals for the CVO site during summer months in the base scenario (left), with  $HOI + NO_3$  chemistry included (centre), and with  $HOI \& HOBr + NO_3$  chemistry included (right).  $BrONO_2 = green$ ,  $IONO_2 = teal$ , PAN = pink,  $NO_3 = orange$ ,  $N_2O_5 = black$ ,  $p-NO_3$  - red.

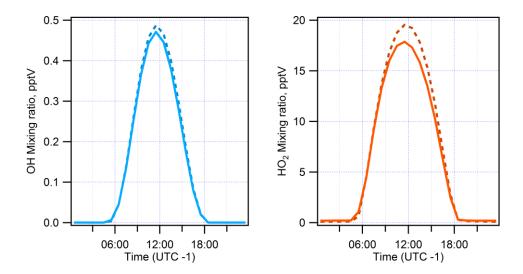


Figure 11. Modelled OH (left) and  $HO_2$  (right) mixing ratios comparing the base case model where PAN decomposition is the dominant source of  $NO_x$  in the remote MBL (solid lines), with the final model where the dominant source of  $NO_x$  is particulate nitrate photolysis and  $HOX + NO_3$  chemsitry is included (dashed lines).