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photodissociation of
peroxy acetyl nitrate**

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Near-IR photodissociation of peroxy acetyl nitrate

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Abstract

Measurements of the C-H overtone transition strengths combined with estimates of the photodissociation cross sections for these transitions suggest that near-IR photodissociation of peroxy acetyl nitrate (PAN) is less significant ($J_{\text{near-IR}} \approx 3 \times 10^{-8} \text{ s}^{-1}$ at noon) in the lower atmosphere than competing sinks resulting from unimolecular decomposition and ultraviolet photolysis. This is in contrast to the photochemical behavior of a related peroxy nitrate, pernitric acid (PNA), that undergoes rapid near-IR photolysis in the atmosphere with $J_{\text{near-IR}} \approx 10^{-5} \text{ s}^{-1}$ at noon (Roehl et al., 2002). This difference is attributed to the larger binding energy and larger number of vibrational degrees of freedom in PAN, which make $4\nu_{\text{CH}}$ the lowest overtone excitation with a high photodissociation yield (as opposed to $2\nu_{\text{OH}}$ in PNA).

1. Introduction

Atmospheric photochemistry is normally driven by electronic state excitations in the ultraviolet range of the solar spectrum. However, photodissociation processes resulting from near-IR direct overtone excitations within the ground electronic state can also be significant in special cases. Donaldson et al. (1997) examined several atmospheric molecules featuring small binding energies and strong overtone transitions, and identified HO_2NO_2 (pernitric acid or PNA) as the most likely candidate for an efficient overtone photodissociation. They argued that the dissociation energy threshold for PNA, $D_0(\text{HO}_2\text{-NO}_2) = 7970 \pm 280 \text{ cm}^{-1}$ (Zabel, 1995), is sufficiently low to permit its overtone photolysis via the second OH stretching overtone transition at $E(3\nu_{\text{OH}}) = 10090 \text{ cm}^{-1}$. Subsequent experimental study of near-IR photodissociation cross sections of PNA by Roehl et al. (2002) showed that PNA overtone dissociation is even more important than Donaldson et al. (1997) initially surmised. Roehl et al. (2002) found that PNA can photolyze with a substantial yield not only upon $3\nu_{\text{OH}}$ but also upon $2\nu_{\text{OH}}$ excitation at 6900 cm^{-1} , which lies as much as 1070 cm^{-1} below the PNA formal dis-

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sociation energy threshold. These striking results have recently been independently confirmed by Sinha's research group (Matthews et al., 2004). Such efficient under-threshold photolysis can be attributed to the collisional activation and internal energy contributions to the dissociation process, just as in the well-known cases of ultraviolet photodissociation of NO₂ (Pitts et al., 1964; Roehl et al., 1994) and O₃ (Finlayson-Pitts and Pitts, 2000). The combined 2ν_{OH} and 3ν_{OH} PNA atmospheric photodissociation rate amounts to ≈10⁻⁵ s⁻¹, making near-IR photolysis the dominant sink process for PNA in the upper troposphere and lower stratosphere (Roehl et al., 2002).

This letter examines possible atmospheric significance of near-IR photolysis for another important peroxy nitrate, CH₃C(O)OONO₂ (peroxy acetyl nitrate or PAN). PAN is a key odd-nitrogen component of photochemical smog (Finlayson-Pitts and Pitts, 2000), and it serves both as a temporary reservoir and as a global transporter for NO_x. PAN often accounts for the largest fraction of all odd-nitrogen compounds (NO_y) in the atmosphere, especially at high latitudes (Ridley et al., 2003; Stroud et al., 2003). The primary known sinks for PAN include unimolecular decomposition, which is the dominant channel in the lower atmosphere, and ultraviolet photolysis, which becomes prevalent in the colder upper tropospheric layers. Reaction of PAN with OH is too slow to be of relevance (DeMore et al., 1997). Similar to other peroxy nitrates (Kirchner et al., 1999; Zabel, 1995), PAN has a low dissociation energy, D₀(CH₃C(O)OO–NO₂) = 9820±240 cm⁻¹ (Bridier et al., 1991; Zabel, 1995), which falls between the 3ν_{CH} (E≈8550 cm⁻¹) and 4ν_{CH} (E≈11 100 cm⁻¹) vibrational states of the molecule. To evaluate the efficiency of near-IR photochemistry of PAN, knowledge of both the intensity of the CH stretching overtones and photodissociation quantum yields are required. Although the strength of CH stretching overtone transitions are well-known for many hydrocarbons (Hanazaki et al., 1985; Kjaergaard et al., 1991; Nakagaki and Hanazaki, 1986), no information on the cross sections and photodissociation quantum yields for PAN is presently available. In what follows, we show that these quantities can be estimated from statistical models and comparisons with experimental data on related molecules, and we use them to predict the near-IR photodissociation rates for PAN

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under typical atmospheric conditions.

2. Results and discussion

2.1. Definitions and units

Subject to the approximations described below, band-specific atmospheric photodissociation rates, J_{band} , can be calculated in the following way:

$$J_{\text{band}} = \int_{\text{band}} F(\bar{\nu}) \sigma(\bar{\nu}) \phi(\bar{\nu}) d\bar{\nu} \approx \langle F(\bar{\nu}) \rangle \times S_{\text{band}} \times \phi_{\text{band}} \quad (1)$$

$F(\bar{\nu})$ is the frequency-dependent solar flux measured in $\frac{\text{photons}}{\text{cm}^2 \text{s cm}^{-1}}$, $\langle F(\bar{\nu}) \rangle$ is the average value of $F(\bar{\nu})$ over the absorption band, $\sigma(\bar{\nu})$ is base- e molecular absorption cross section in $\text{cm}^2/\text{molecule}$, $\phi(\bar{\nu})$ is photodissociation quantum yield, $\bar{\nu}$ is wave number in cm^{-1} , and $S_{\text{band}} = \int_{\text{band}} \sigma(\bar{\nu}) d\bar{\nu}$ is the integrated absorption cross section

(in $\text{cm}/\text{molecule}$) for the band in question. The last equality applies only if $F(\bar{\nu})$ varies weakly with frequency, which is usually a good approximation for the narrow near-IR bands unless they are close to strong water or carbon dioxide absorptions. The quantity ϕ_{band} is best interpreted as the ratio between the integrated photodissociation and absorption cross sections for the band. The paragraphs below deal with estimation of S and ϕ for the CH stretching overtone transitions of PAN.

2.2. PAN overtone absorption cross-sections

Integrated absorption cross sections for PAN fundamentals have been reported by Gaffney et al. (1984). However, their analysis has subsequently been questioned and corrected by Tsalkani and Toupance (1989) for the four strongest infrared bands of PAN below 2000 cm^{-1} . To resolve the resulting ambiguity of the intensity data for

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the remaining bands not considered by Tsalkani and Toupance (1989), we have measured relative intensities of the CH stretching fundamental and overtone transitions and determined absolute transition intensities by relation to Tsalkani and Toupance (1989) data for NO stretching vibrations of PAN. Pure gaseous PAN samples were synthesized as described by Gaffney et al. (1984), and gas-phase spectra of PAN were recorded with an FTIR spectrometer at 1 cm^{-1} resolution (Fig. 1). Depending on the examined spectral range, we used either a single-pass 20 cm cell containing 0.1–1 Torr partial pressure of PAN, or a multipass cell with an effective path length of 360 cm containing the equilibrium vapor pressure of PAN at room temperature (about 4.5 Torr). The multipass configuration was required to observe the second CH stretching overtone with an acceptable signal-to-noise ratio (Fig. 1).

The ratios of PAN band strengths were obtained from multiple spectra. The absorbances ($A = \sigma(\bar{\nu}) \times [\text{PAN}] \times \text{pathlength}$) for the bands in question were always kept small to avoid deviations from Beer-Lambert law. The value of $S(\nu_{\text{CH}})$ was measured relative to the band strengths of NO_2 stretches reported by Tsalkani and Toupance (1989). Our value, $S(\nu_{\text{CH}}) = 1.57 \times 10^{-18}\text{ cm/molecule}$, is 40% lower than that reported by Gaffney et al. (1984). This agrees with the conclusions of Tsalkani and Toupance (1989), who showed that data from Gaffney et al. (1984) overestimated band intensities of PAN by 5 to 35%. The CH stretching overtone band strengths were always measured relative to the closest available CH stretching transition at lower frequency. Our results $S(2\nu_{\text{CH}})/S(\nu_{\text{CH}}) = 0.096 \pm 0.015$ and $S(3\nu_{\text{CH}})/S(2\nu_{\text{CH}}) = 0.085 \pm 0.030$ are consistent with the generally observed trend of the 10-fold drop of overtone intensity per vibrational quantum (we define $S(2\nu_{\text{CH}})$ and $S(3\nu_{\text{CH}})$ as the total integrated absorption cross section for the $5600\text{--}6300\text{ cm}^{-1}$ and $8300\text{--}9000\text{ cm}^{-1}$ ranges, respectively).

Integrated intensity for the $4\nu_{\text{CH}}$ overtone of PAN could not be reliably determined from our FTIR spectra, but it can be deduced from comparison with spectra of other molecules bearing an acetyl group. Indeed, the observed spectra of the first and second CH overtones of PAN (Fig. 1) closely resemble the spectra of acetone and acetaldehyde (Hanazaki et al., 1985). The orientational site splitting pattern arising from

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the inequivalence of CH bonds in the methyl group (Hanazaki et al., 1985; Kjaergaard et al., 1991) is similar for all three molecules. The integrated intensities (per methyl group) for the ν_{CH} fundamental, $2\nu_{\text{CH}}$ first overtone, and $3\nu_{\text{CH}}$ second overtone bands are also quite similar (Table 1). Assuming that the PAN overtone intensity decreases with vibrational quantum number by the same factors as the average of the corresponding acetone and acetaldehyde values, $S(4\nu_{\text{CH}})/S(3\nu_{\text{CH}}) = 0.059$, we can estimate $S_{\text{PAN}}(4\nu_{\text{CH}})$ as 7.6×10^{-22} cm/molecule. Based on the good uniformity of the experimentally available band strength ratios for these three molecules, we do not expect the estimated value of $S_{\text{PAN}}(4\nu_{\text{CH}})$ to deviate from the true band strength by more than a factor of two.

2.3. Photodissociation quantum yields

Photodissociation quantum yields for CH overtone excitations of PAN have been calculated with the Multiwell suite (Barker, 2001) using a statistical approach. The calculation involves shifting the Boltzmann internal energy distribution of the molecule upward by the photon energy, and solving the coupled master equations for discretized internal energy states. Excited molecules can undergo either a unimolecular decomposition with the microcanonical rate constant $k_{\text{uni}}(E)$ or collision-induced internal energy change with the rate that depends on pressure and internal energy. The fraction of molecules that end up decomposing by the time internal energy in the surviving molecules is nearly thermalized ($\approx 10^3$ collisions) is taken as the photodissociation yield.

The exponential collision model built into the Multiwell was used, with the probability of downward collisions taken as $e^{(E_{\text{final}} - E_{\text{initial}})/\Delta}$, where $\Delta = 45 \text{ cm}^{-1} + 0.005 \times E_{\text{initial}}$, a weak collision model (Barker, 2001). Vibrational frequencies for PAN were taken from Miller et al. (1999). Lennard-Jones collision parameters were the same as used by Bridier et al. (1991). The NO_2 and $-\text{C}-\text{O}-\text{O}$ moieties were treated as hindered rotors in the transition state. The CH_3 group was treated as a free rotor for both reactant and

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transition state. The initially-selected RRKM parameters were fine-tuned to reproduce i) high pressure limit of the experimental unimolecular decomposition rate (Bridier et al., 1991; Zabel, 1995); ii) the equilibrium constant for PAN decomposition (Bridier et al., 1991) at all experimentally available temperatures. The final choice of parameters for the calculation could reproduce both experimental data sets within 30%.

The predictive power of the calculation was first tested on PNA because photodissociation quantum yields have already been measured for this molecule (Roehl et al., 2002). Figure 2 compares the experimental results with our statistical calculations for the $\nu_{\text{OH}} + 2\nu_{\text{OOH}}$ combination band, $2\nu_{\text{OH}}$ overtone, and $2\nu_{\text{OH}} + \nu_{\text{OOH}}$ combination band of PNA (listed here in the order of increasing excitation energy). Required RRKM parameters for PNA were taken from experiment and from Chen and Hamilton (1996) and Zabel (1995). The simulation reproduces both the magnitude and the temperature dependence of the photodissociation quantum yields with good accuracy. To achieve this level of agreement, we had to reduce the dissociation energy threshold of PNA by 200 cm^{-1} from the accepted experimental value, which is well within the error bars for $D_0(\text{HO}_2\text{--NO}_2) = 7970 \pm 280 \text{ cm}^{-1}$ (Zabel, 1995).

A curious side result of this simulation is the prediction of unity quantum yield for the $2\nu_{\text{OH}} + \nu_{\text{OOH}}$ excitation of PNA. The experimental data suggest a yield that is slightly below unity for this transition. This discrepancy may be a reflection of not entirely statistical behavior of PNA at these excitation energies. It may also reflect deficiencies of the chosen weak collision model. However, the experimental data for $2\nu_{\text{OH}} + \nu_{\text{OOH}}$ and $\nu_{\text{OH}} + 2\nu_{\text{OOH}}$ bands are relatively noisy (Fig. 2) making it difficult to draw more definite conclusions. Matthews et al. (2004) assumed a photodissociation quantum yield of one for the $2\nu_{\text{OH}} + \nu_{\text{OOH}}$ band in their work.

The predicted dissociation yields for PAN excitation at 8700 cm^{-1} (i.e. within the blue wing of the $3\nu_{\text{CH}}$ band) are shown in Fig. 3 for several atmospherically relevant temperatures. The most interesting observation is that the photodissociation yields for PAN display a remarkably broad pressure dependence. In contrast, the photodissociation yield for the $2\nu_{\text{OH}}$ overtone transition of PNA is fairly constant below $P =$

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1 atm. This important difference can be attributed to the substantially larger number of vibrational degrees of freedom and larger binding energy in PAN resulting in lower microcanonical decomposition rates. The collisional deactivation and excitation rates can, therefore, easily compete with unimolecular decomposition even when the system is placed above the dissociation energy. In fact, the calculated PAN dissociation yields do not reach unity under typical atmospheric pressures even for excitation lying as much as 1000 cm^{-1} above the dissociation threshold (Fig. 4). In contrast, PNA is predicted to fall apart with 100% efficiency when excited so high above the dissociation energy threshold at pressures below one atmosphere.

2.4. Near-IR photodissociation rates for atmospheric PAN

Simulated PAN photodissociation yields for a standardized atmosphere are given in Table 2. The effective photodissociation yield for $3\nu_{\text{CH}}$ band (estimated as an average of yields calculated for 8450, 8550, and 8650 cm^{-1} excitations) is very small compared to the $4\nu_{\text{CH}}$ yield (corresponding to $11\,100\text{ cm}^{-1}$ excitation). Using the transition strength data from Table 1 and a typical near-IR solar flux of $5 \times 10^{14}\text{ photons cm}^{-2}\text{ s}^{-1}\text{ nm}^{-1}$ (Finlayson-Pitts and Pitts, 2000), we estimate the total near-IR photodissociation rate for PAN as $J_{\text{near-IR}} \approx 3 \times 10^{-8}\text{ s}^{-1}$ with a fairly weak altitude dependence (Table 2).

Because of the very low contribution of the $3\nu_{\text{CH}}$ band of PAN to the total near-IR photolysis rate (<5% above 6 km), the predicted rate is relatively insensitive to the calculated photodissociation yields. For example, an order of magnitude increase in $\phi(3\nu_{\text{CH}})$ would translate to less than 50% increase in the total photolysis rate. The $\phi(4\nu_{\text{CH}})$ values are already above 0.8; adjusting them to the maximal value of unity would only produce a marginal increase in the rate. The confidence in the predicted rate is, therefore, dominated by the uncertainty in $S(4\nu_{\text{CH}})$ believed to be accurate to within a factor of two (Table 1). Given these uncertainties, a very conservative upper limit for the near-IR PAN photodissociation rate under typical atmospheric conditions can be set at $J_{\text{near-IR}}(\text{PAN}) < 10^{-7}\text{ s}^{-1}$. Using similar reasoning, one can also show that

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$J_{\text{near-IR}}$ is insensitive to poor atmospheric transmittance at wavelengths corresponding to $3\nu_{\text{CH}}$ transition (the $4\nu_{\text{CH}}$ spectral range is clean).

This limit is to be contrasted with a characteristic near-IR photolysis rate for PNA, $J_{\text{near-IR}}(\text{PNA}) \approx 10^{-5} \text{ s}^{-1}$, which is higher by two orders of magnitude. The large difference in the photodissociation rates for PAN and PNA reflects the disparity in absorption cross sections for the overtone excitations responsible for their near-IR photodissociation activity. Whereas the under-threshold $2\nu_{\text{OH}}$ excitation accounts for the major fraction of $J_{\text{near-IR}}(\text{PNA})$, the $3\nu_{\text{CH}}$ band of PAN is but a minor contributor (<5% above 6 km) to $J_{\text{near-IR}}(\text{PAN})$. Because the overtone absorption strengths go down by roughly one order of magnitude with each vibrational quantum, the jump in two vibrational quanta between the primary photolytic excitations in PAN($4\nu_{\text{CH}}$) and PNA($2\nu_{\text{OH}}$) can be held responsible for the observed difference in the $J_{\text{near-IR}}$ values. In addition, CH_3 overtone intensities in acetyl groups are known to be systematically lower than those of other kinds of CH_3 groups (Nakagaki and Hanazaki, 1986). They are also considerably lower than intensities of typical OH stretching overtones (e.g. compare $S_{\text{PNA}}(2\nu_{\text{OH}}) = (9.5 \pm 1.9) \times 10^{-19} \text{ cm}^2/\text{cm}^{-1}$ (Roehl et al., 2002) with $S_{\text{PAN}}(2\nu_{\text{CH}}) = 1.51 \times 10^{-19} \text{ cm}^2/\text{cm}^{-1}$).

Although the predicted near-IR photodissociation rate of PAN is quite low (the 24-hour average of $1.5 \times 10^{-8} \text{ s}^{-1}$ translates into an atmospheric lifetime of 770 days), it becomes strongly competitive with the thermal decomposition rate (Zabel, 1995) above 6 km (Table 2 and Fig. 5). It is also higher than the upper limit for the rate of OH + PAN reaction ($k_{\text{OH+PAN}}[\text{OH}] < 10^{-8} \text{ s}^{-1}$ for the troposphere (DeMore et al., 1997)). Comparison with the UV photolysis rate of PAN is less straightforward because UV radiation is attenuated by the atmosphere to a much greater extent than near-IR at high solar zenith angle (SZA). For example, at 15 km altitude, $J_{\text{UV}}(\text{PAN}) \approx 6 \times 10^{-7} \text{ s}^{-1}$ at SZA = 0° but it drops to $\approx 6 \times 10^{-8} \text{ s}^{-1}$ at SZA = 86° (solar fluxes for this calculation are taken from Finlayson-Pitts and Pitts, 2000 and PAN absorption cross sections are from DeMore et al., 1997). On the contrary, the estimated near-IR photolysis rate ($\approx 3 \times 10^{-8} \text{ s}^{-1}$) is reduced by only 35% upon this SZA change. Consequently, the near-IR photolysis of PAN can potentially become comparable to the UV photolysis rate at very

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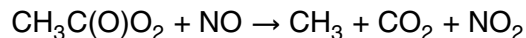
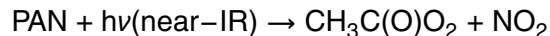
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high SZA, providing a small contribution to the twilight production of HO_x radicals at high altitudes (Wennberg et al., 1999).

Figure 5 displays 24-hour averaged rates of UV photolysis, near-IR photolysis, and thermal decomposition for PAN. The UV photolysis rates correspond to conditions found at high and middle latitudes near equinox as described in Salawitch et al. (2002) (8 May 1997 with latitudes between 65° N and 70° N and 25 September 1993 at 35° N). Near-IR rates are estimated by assuming day lengths of 18 h and 12 h, respectively, and a near-IR solar flux of 5×10^{14} photons cm⁻² s⁻¹ nm⁻¹. The thermal decomposition rate for PAN is calculated from data of Zabel (1995). Figure 5 shows that $\langle J_{\text{near-IR}}(\text{PAN}) \rangle$ is approximately 5–10% of $\langle J_{\text{UV}}(\text{PAN}) \rangle$ at high tropospheric altitudes. At lower altitudes, both photolysis channels are dwarfed by the thermal decomposition. We conclude that near-IR photolysis is not likely to significantly affect the average atmospheric lifetime of PAN. Given that it is hard to quantify the uncertainties of the predicted photodissociation yields and absorption cross sections of PAN, however, the near-IR photolysis of PAN in the atmosphere cannot be completely discounted at this stage. More experimental research in this area is clearly needed.

2.5. Experiments on near-IR photolysis of PAN

We have attempted to measure photodissociation cross sections of PAN in the $3\nu_{\text{CH}} - 4\nu_{\text{CH}}$ range using the action spectroscopy approach similar to that used for PNA by Roehl et al. (2002). PAN/N₂ mixture at 1–5 Torr and 250–300 K was irradiated with a tunable pulsed near-IR laser in the presence of a small amount of NO and O₂. A dye laser was employed in a fluorescence scheme to monitor the OH produced in the mixture following the photolysis laser pulse. If the near-IR laser were to photolyze PAN molecules in the gas flow, a momentary increase in the OH fluorescence signal would be observed assuming that the following reaction sequence runs to completion:



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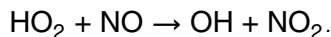
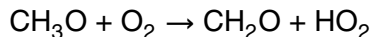
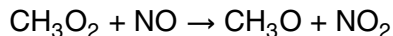
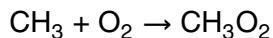
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5 Despite numerous attempts under various reaction conditions, we have not been able to observe any near-IR laser induced increase in the OH signal. Detailed kinetic simulations of our experimental conditions predict a signal-to-noise ratio reduction of $\approx 10^2$ for $3\nu_{\text{CH}}$ band of PAN compared to $3\nu_{\text{OH}}$ band of PNA (which could be successfully observed in this apparatus, Roehl et al., 2002) assuming unity photodissociation yield in both cases. This large penalty is comprised of i) a factor of 2 reduction in the integrated absorption cross section; ii) factor of 5 increase in the effective band width for the CH_3 vs. OH second overtone; iii) a factor of ~ 10 reduction in the efficiency of conversion of $\text{CH}_3\text{C}(\text{O})\text{O}_2$ into OH in the chemical sequence above, as opposed to that of HO_2 into OH in the case of PNA photodissociation. The signal-to-noise ratio in our best $3\nu_{\text{OH}}$ photodissociation spectrum of PNA ($\approx 10^3$, Roehl et al., 2002) allows us to place an upper limit of 10% on the photodissociation yield of the $3\nu_{\text{CH}}$ state of PAN at 1–5 Torr and 250–300 K, which is in rough agreement with simulations presented here. The predicted signal-to-noise ratio for the $4\nu_{\text{CH}}$ band of PAN is < 0.5 , consistent with the failure to experimentally detect it.

20 We are currently designing an experiment that will allow us to measure $\text{CH}_3\text{C}(\text{O})\text{O}_2$ following pulsed near IR laser excitation of PAN/ N_2 mixtures using a CIMS (chemical ionization mass spectrometry) technique. Preliminary tests show positive results using UV photons to dissociate PAN. The results of this work will be published elsewhere.

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Due to its unique combination of binding energy and molecular size, PNA may well turn out to be the only peroxy nitrate that is controlled by near-IR photolysis (Roehl et al., 2002) in the atmosphere. The larger binding energies and higher degrees of molecular complexity make peroxy acetyl nitrates, like PAN, considerably more resistant to direct overtone photolysis in the near-IR range of the solar spectrum. Specifically, our analysis predicts a near-infrared photodissociation rate of $\approx 3 \times 10^{-8} \text{ s}^{-1}$ for PAN at noon, largely independent of altitude. The near-IR photolysis of PAN can become competitive with UV photolysis under twilight illumination conditions, but even in the coldest upper tropospheric layers it is calculated to contribute only 5–10% to the 24-hour averaged removal rate of PAN.

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Table 1. Experimental integrated absorption cross sections (base e) for the CH-stretching overtones of acetone, acetaldehyde, and PAN. The units of $S(\text{band})$ are $\text{cm}^2/\text{molecule cm}^{-1}$ ($1 \text{ cm}^{-2} \text{ atm}^{-1}$ at $298 \text{ K} = 4.063 \times 10^{-20} \text{ cm/molecule}$). The integrated absorption cross sections are given per CH_3 group; i.e. the acetone values have been divided by 2.

Band	Acetone/2 ^a	Acetaldehyde ^a	PAN
(ν_{CH})	3.04E-18	2.02E-18	1.57E-18
2(ν_{CH})	1.28E-19	1.51E-19	1.51E-19 ^b
3(ν_{CH})	1.60E-20	1.63E-20	1.28E-20 ^c
4(ν_{CH})	1.09E-21	8.21E-22	7.56E-22 ^d

^a From Hanazaki et al. (1985).

^b Measured to be 9.6% of $S(\nu_{\text{CH}})$ in this work.

^c Measured to be 8.5% of $S(2\nu_{\text{CH}})$ in this work.

^d Assumed value based on comparison with acetone and acetaldehyde (5.9% of $S(3\nu_{\text{CH}})$).

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Table 2. Calculated photodissociation quantum yields for PAN overtone transitions under representative atmospheric conditions. The $J_{\text{near-IR}}$ column lists the total near-IR photodissociation rates ($J_{3\nu\text{CH}} + J_{4\nu\text{CH}}$) calculated from the integrated absorption cross sections from Table 1 and a solar flux of $5 \times 10^{14} \text{ photons cm}^{-2} \text{ s}^{-1} \text{ nm}^{-1}$. The last column is the fractional contribution of under-threshold dissociation via $3\nu_{\text{CH}}$ to the total rate (i.e. $J_{3\nu\text{CH}}/J_{\text{near-IR}}$).

Altitude (km)	T (K)	P (Torr)	$\phi(3\nu_{\text{CH}})$	$\phi(4\nu_{\text{CH}})$	$J_{\text{near-IR}}$ (s^{-1})	$3\nu_{\text{CH}}$ contribution (%)
0	288	760	0.0041	0.80	2.8E-8	14.8
3	269	526	0.0027	0.82	2.8E-8	9.3
6	249	354	0.0016	0.85	2.7E-8	5.5
9	230	230	0.0009	0.88	2.8E-8	3.0
12	217	143	0.0007	0.92	2.9E-8	2.2
15	217	90	0.0011	0.97	3.1E-8	3.1

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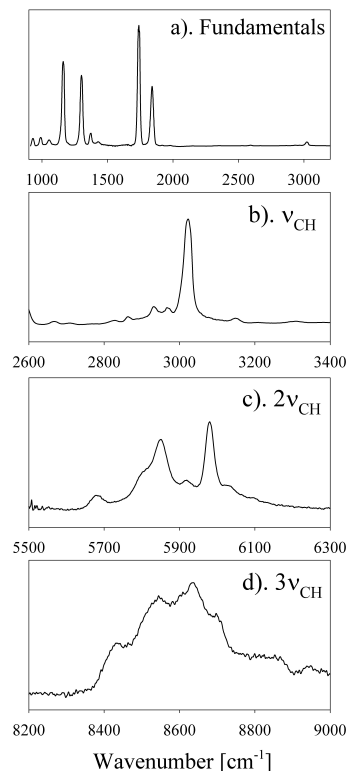


Fig. 1. IR spectrum of PAN recorded in this work: **(a)** fundamental range displaying well-known PAN bands; **(b)** CH stretching fundamental range; **(c)** first CH stretching overtone range; **(d)** second CH stretching overtone range. The CH-stretching overtones have a complicated band structure resulting from the inequivalence of CH bonds in the methyl group. Experimentally measured ratios between integrated absorption cross sections are $S(\nu_{\text{CH}}):S(2\nu_{\text{CH}}):S(3\nu_{\text{CH}}) = 1:0.096:0.0082$. $S(\nu_{\text{CH}})$ is measured to be 1.57×10^{-18} cm/molecule (base e) in this work.

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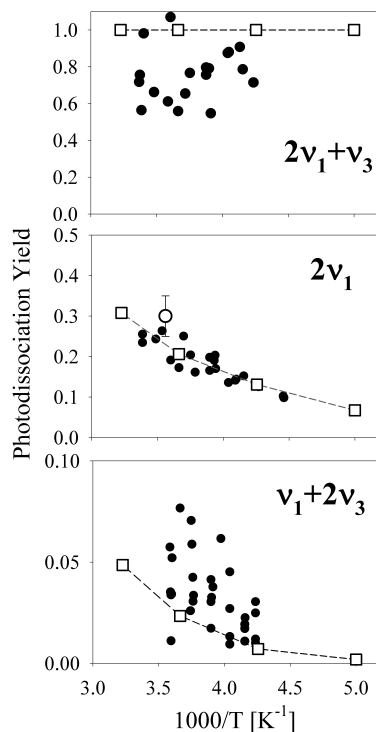


Fig. 2. Experimental photodissociation quantum yields for different vibrational overtone excitations of PNA measured by Roehl et al. (2002) (filled circles) compared to the statistical simulation results (open squares) of this work. Both simulations and data correspond to $P = 5$ Torr. The open circle point in the $2\nu_1$ panel comes from Matthews et al. (2004) and it corresponds to $T = 298$ K, $P = 0.1$ Torr. Parameters for our statistical simulation were chosen to reproduce the $2\nu_1$ photodissociation yields.

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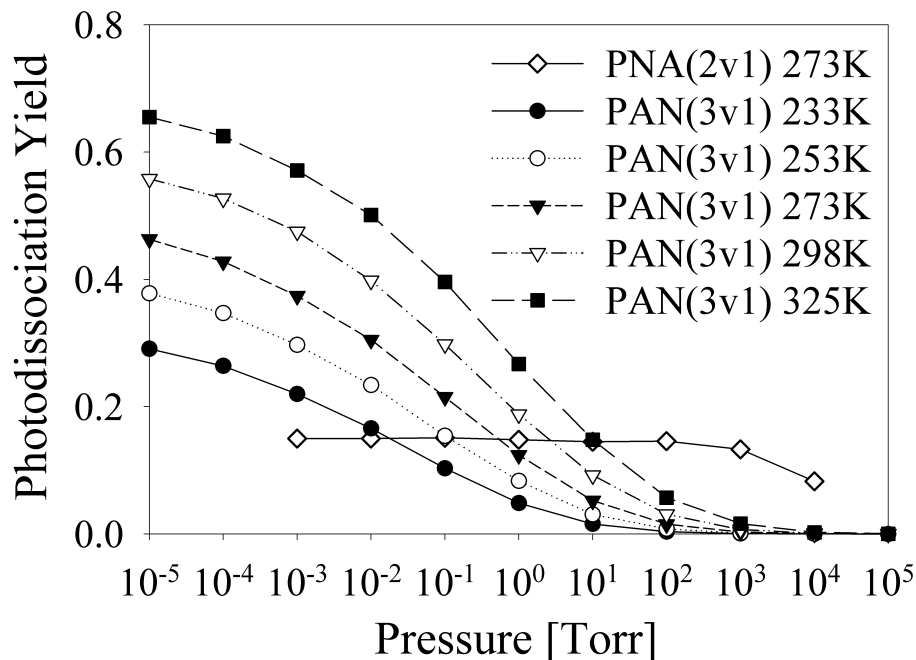


Fig. 3. Pressure dependence of the predicted photodissociation quantum yields for 8700 cm^{-1} excitation of PAN (representative of the $3\nu_{\text{CH}}$ excitation) at different temperatures. Calculated photodissociation yields for $2\nu_{\text{OH}}$ band of PNA at 273 K are also shown for comparison. In contrast to the PNA case, photodissociation quantum yields for PAN photolysis at 8700 cm^{-1} strongly depend on pressure under typical atmospheric conditions.

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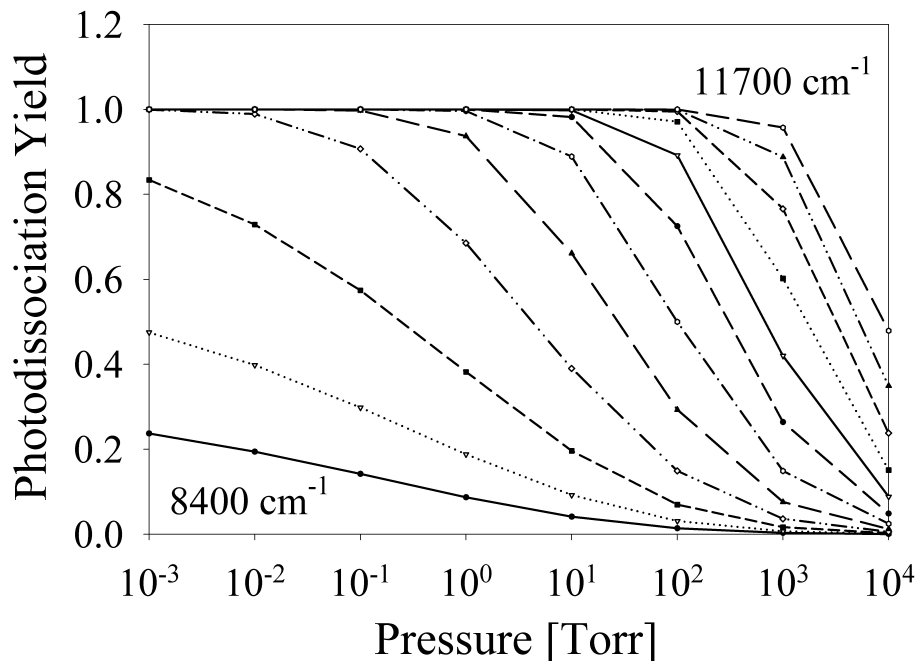


Fig. 4. Pressure dependence of the predicted photodissociation quantum yields of PAN at 298K at different excitation frequencies. The frequencies cover the range 8400 cm^{-1} (bottom curve) to 11700 cm^{-1} (top curve) in 300 cm^{-1} increments. Even at the highest excitation energy, the yield is still noticeably different from unity at typical atmospheric pressures.

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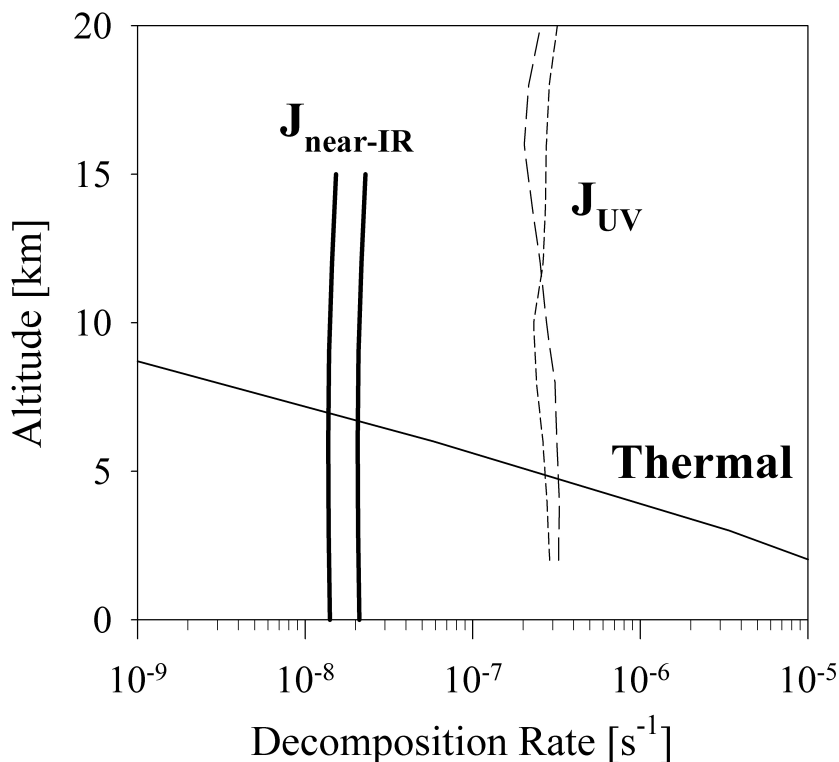


Fig. 5. Dependence of the predicted 24 hour-averaged near IR photodissociation rates (thick solid lines), UV photolysis rates (dashed lines), and unimolecular decomposition rate (thin solid line) on altitude. The UV photolysis rates were calculated for actual conditions of 8 May, 67° N (short dash) and 25 September, 35° N (long dash). Two near IR photolysis curves correspond to 18-hour (8 May) and 12-hour (25 September) day lengths with an average solar flux of 5×10^{14} photons $\text{cm}^{-2} \text{s}^{-1} \text{nm}^{-1}$. The near IR channel is seen to play a minor role (<10%) at all altitudes.

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