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# Measurements of the sum of HO<sub>2</sub>NO<sub>2</sub> and CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> in the remote troposphere

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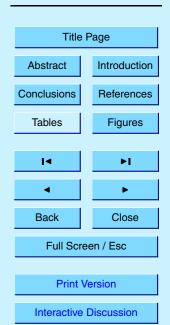
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#### **Abstract**

The chemistry of peroxynitric acid (HO<sub>2</sub>NO<sub>2</sub>) and methyl peroxynitrate (CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub>) is predicted to be particularly important in the upper troposphere where temperatures are frequently low enough that these compounds do not rapidly decompose. At temperatures below 240 K, we calculate that about 20% of NO<sub>v</sub> in the mid and polar latitude upper troposphere is HO<sub>2</sub>NO<sub>2</sub>. Under these conditions, the reaction of OH with HO<sub>2</sub>NO<sub>2</sub> is estimated to account for as much as one third of the permanent loss of hydrogen radicals. During the Tropospheric Ozone Production about the Spring Equinox (TOPSE) campaign, we used thermal dissociation laser-induced fluorescence (TD-LIF) to measure the sum of peroxynitrates ( $\Sigma PNs = HO_2NO_2 + CH_3O_2NO_2 + PAN + PPN$ + ...), aboard the NCAR C-130 research aircraft. We infer the sum of HO<sub>2</sub>NO<sub>2</sub> and CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> as the difference between ΣPN measurements and gas chromatographic measurements of the two major peroxy acyl nitrates, peroxy acetyl nitrate (PAN) and peroxy propionyl nitrate (PPN). Comparison with NO<sub>v</sub> and other nitrogen oxide measurements confirms the importance of HO<sub>2</sub>NO<sub>2</sub> and CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> to the reactive nitrogen budget and shows that current thinking about the chemistry of these species is approximately correct. The temperature dependence of the inferred concentrations corroborates the contribution of overtone photolysis to the photochemistry of peroxynitric acid.

#### 1. Introduction

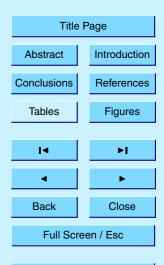
Reactions of hydrogen radicals ( $HO_x \equiv OH + HO_2$ , and  $RO_2$ ) and nitrogen radicals ( $NO_x \equiv NO + NO_2$ ) affect atmospheric composition and climate by regulating global tropospheric ozone and the Earth's oxidative capacity. Peroxynitrates ( $RO_2NO_2$ ) link these chemical families, thereby influencing tropospheric ozone production and the abundance of OH in both the troposphere and stratosphere. Non-acyl peroxynitrates such as  $HO_2NO_2$  and  $CH_3O_2NO_2$  are increasingly important at temperatures below

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240 K, because the molecules are very weakly bound ( $\sim$ 95 kJ/mol) and rapidly dissociate at higher temperature. Consequently these molecules are most important in the stratosphere and upper troposphere.

Knowledge of the chemistry of alkyl RO<sub>2</sub>NO<sub>2</sub> species is limited because they are 5 difficult to isolate in the laboratory or to observe directly in the atmosphere. Remote sensing measurements of peroxynitric acid concentrations in the stratosphere show that HO<sub>2</sub>NO<sub>2</sub> mixing ratios peak at approximately 200 ppt near 27 km altitude (Rinsland et al., 1996; Rinsland et al., 1986; Sen et al., 1998). The only other measurements are from the Antarctic surface where Slusher et al. (2001) used chemical ionization mass spectrometry to observe HO<sub>2</sub>NO<sub>2</sub> mixing ratios on the order of 20 ppt, showing they were nearly equal to nitric acid mixing ratios. Recent OH and HO2 measurements made in the lower stratosphere at high solar zenith angles (SZA) indicate that there is a source of HO, that dissociates following absorption of low energy photons (Salawitch et al., 1994; Wennberg et al., 1999). Overtone photodissociation of HO<sub>2</sub>NO<sub>2</sub>, as suggested by Donaldson et al. (1997), can explain these observations and some of the temporal variability seen in remote sensing of HO<sub>2</sub>NO<sub>2</sub> (Salawitch et al., 2002). These observations and calculations of HO<sub>2</sub>NO<sub>2</sub> abundances and their effect on stratospheric HOx have been the impetus for revisiting the long-wavelength photolysis and kinetics of peroxynitrates in the laboratory (Knight et al., 2002; Roehl et al., 2002; Zhang et al., 2000).

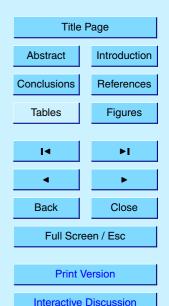
Here we describe the first in situ measurements of these peroxynitrates in the free troposphere. The measurements were made during the Tropospheric Ozone Production about the Spring Equinox (TOPSE) experiment (Atlas et al., 2003). Analysis along chemical coordinates (Cohen et al., 2000), in this case temperature and  $NO_2$ , and consideration of the  $NO_y$  budget are used to evaluate current understanding of photolytic and thermal decomposition rates of  $HO_2NO_2$  and  $CH_3O_2NO_2$ .

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#### 2. Instrumentation

NO<sub>y</sub> and its components were measured from the NCAR C-130 aircraft during the TOPSE experiment. Ridley et al. measured NO, NO<sub>2</sub> and NO<sub>y</sub> by conversion to NO followed by O<sub>3</sub> + NO chemiluminescence (Ridley et al., 2000). The NO<sub>2</sub> measurements were precise to ±4 ppt and accurate to 5%. C1–C4 alkyl nitrates were quantified by GC analysis of whole air samples (Blake et al., 2003). In our analysis, the sum of C1–C4 alkyl nitrate observations was linearly interpolated to a 1 min time base. Nitric acid was measured using a mist chamber followed by ion chromatography (Talbot et al., 1990). PAN and PPN were measured in situ by the NCAR Gas Chromatography with Electron Capture Detection (GC-ECD) instrument (The behavior of PANs and the balance of NO<sub>y</sub> during TOPSE, Flocke et al., in prep). The PAN and PPN measurements represent 1–3 s samples with a precision of ±5 ppt at 20 ppt, ±3 ppt at 100 ppt, ±5–7 ppt at 500 ppt. The accuracy of the PAN and PPN measurements is 10%. Because they were made with the same instrument, we assume their uncertainties are correlated, that is, if necessary, both could be brought to the "true" value by multiplying by a single constant. During a 5 min interval, two PAN measurements and one PPN are obtained.

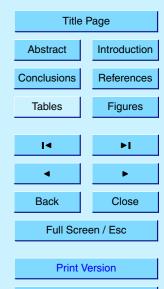
two-channel thermal dissociation laser-induced fluorescence LIF) instrument was used for observations of total peroxynitrates (ΣPNs  $\equiv$  PAN + PPN + HO<sub>2</sub>NO<sub>2</sub> + CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> + ...) and of NO<sub>2</sub>. TD-LIF is described in detail in Day et al. (2002). The specific instrument configuration used during TOPSE is presented by Thornton (2002). Briefly, TD-LIF relies on a heated inlet to dissociate RO2NO<sub>2</sub> to yield NO<sub>2</sub> and RO<sub>2</sub>. The resultant NO<sub>2</sub> is quantified along with ambient NO<sub>2</sub> using laser-induced fluorescence (LIF). By operating the heated inlet in parallel with an inlet maintained at ambient temperature, SPN mixing ratios can be derived from the difference between the NO<sub>2</sub> signals of the two channels. The first third of the campaign was used to test and improve some aspects of instrument performance. High time resolution measurements of NO<sub>2</sub> and ΣPNs were obtained simultaneously and continuously during the final nineteen flights of the TOPSE experiment. The

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precision of the measurements for a 1 min average in each channel under optimal conditions was  $\pm 20$  ppt which is dominated by uncertainty in the instrument zero. Under some conditions, usually after steep altitude changes, the apparent precision is  $\pm 50$  ppt. The accuracy of channel 1 is estimated to be 10% and the accuracy of channel 2 to be 15%. The precision of  $\Sigma PNs$  is given by:

$$\sigma(\Sigma PNs) = [(\sigma_{channel 1})^2 + (\sigma_{channel 2})^2]^{1/2}$$
(1)

and for a typical scenario  $\sigma(\Sigma PNs)$  is  $\pm 28$  ppt for 1 min averaging. The accuracy is estimated to be 15% since the accuracy of the two channels is largely correlated and most of the signal is in channel 2. While TD-LIF  $NO_2$  measurements were used for calculating  $\Sigma PN$ , in the model calculations below we use the NCAR photolysis-chemiluminescence measurements of  $NO_2$  because of their superior precision in the low concentration ranges encountered during TOPSE.

Measurements suggest that PAN is the dominant contribution to  $\Sigma$ PNs and that PPN is the only other major acyl peroxynitrate in the air sampled during TOPSE. In order to infer abundances of non-acyl peroxynitrates (hereafter referred to as  $\Sigma$ PNs) we subtract speciated measurements of PAN and PPN obtained by the NCAR GC-ECD instrument from the TD-LIF measurements of  $\Sigma$ PNs.

$$\Delta PNs = \Sigma PNs - PAN - PPN. \tag{2}$$

We averaged 1 min of TD-LIF data about each PAN measurement and interpolated PPN to generate a dataset of all species coincident in time.

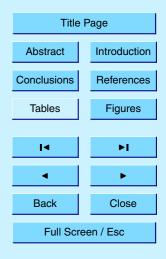
The uncertainty in  $\Delta PNs$  has components associated with both the precision and accuracy of the TD-LIF and GC measurements. In what follows, we do not draw any conclusions based on individual measurements, rather our conclusions are based on averages over a large enough sample that uncertainty in mean quantities is entirely a function of the instruments' accuracies. The techniques are thought to be accurate to 10% (GC) and 15% (LIF) and for both techniques one single multiplicative correction (if it could be known) would bring the observed values to the "true" values (to within

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the precision uncertainty). The accuracy of these instruments depends on a chain of calibration that connects to NIST traceable reference standards of NO or NO<sub>2</sub>, on the accuracy of flow controllers used to dilute these standards and on titration reactions. None of these factors depend on environmental parameters. Thus the associated accuracy in the average value of a combined quantity like ΔPNs adds in quadrature, since the accuracy of the GC and LIF measurements are assumed to be uncorrelated, resulting in a combined accuracy of 18%. The random error associated with an individual ΔPNs measurement is obtained by standard propagation of error based on Eq. (2);

$$\sigma(\Delta PNs) = [(\sigma_{\Sigma PNs})^2 + (\sigma_{PAN})^2 + (\sigma_{PPN})^2]^{1/2}, \tag{3}$$

which results in an uncertainty of  $\pm 29$  ppt.

Figure 1a shows the ratio of  $\Sigma PNs/(PAN+PPN)$  versus humidity for data that calculations suggest have negligible amounts of  $\Delta PNs$  ( $\Delta PNs$  less than 10% of PAN). In contrast to the expectations outlined above, Fig. 1 shows that the two techniques differ by 35% at high relative humidity and agree at low humidity. The absolute difference is larger than expected and the accuracy is not observed to be independent of environmental parameters. One or both instruments are not accurate to within the 10 or 15% quoted.

Neither technique is known to have a humidity-dependent bias or interference of this magnitude. Applying a constraint that

$$\Delta PNs \ge 0$$
 (4)

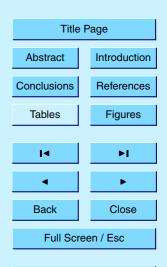
permits the derivation of a constant that implicitly couples the calibration constants for the two instruments and corrects for this difference as a function of humidity. A correction factor, C=(1+0.0025\*%RH), was generated that brings the average (PAN + PPN)/ $\Sigma$ PNs near one when  $\Delta$ PNs are insignificant. Since we have no reason to apply the correction to one or the other instrument, the TD-LIF data was multiplied by the correction factor C, and the GC data was divided by this factor. This removed the humidity bias (Fig. 1b). Throughout the remainder of this paper we use these adjusted

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datasets. Applying the factor C to the two data sets alters them by 12% on average (+12%  $\Sigma$ PNs and -12% PAN and PPN). Using Eq. (2) with the corrected datasets provides 1433 measurements of  $\Delta$ PNs, which calculations suggest are predominantly  $HO_2NO_2$  and  $CH_3O_2NO_2$ . We estimate the uncertainty in  $\Delta$ PNs is 25%.

Measurements of  $NO_y$  and comparison to the sum of individual  $NO_y$  compounds  $(\Sigma NO_y)$  provide another means to evaluate the relative calibration of the GC-ECD, TD-LIF and  $NO_y$  techniques, as well as our ability to infer  $\Delta PNs$  from the adjusted difference of the two sets of measurements. Figure 2 demonstrates the importance of  $HO_2NO_2$  in the reactive nitrogen budget and the ability of TD-LIF  $\Delta PN$  measurements to account for this compound. The ratio  $\Sigma NO_{yi}/total\ NO_y$ , with the above correction applied to each peroxynitrate dataset, is plotted against fraction of  $NO_y$  that is predicted by a model to be  $HO_2NO_2$ . The ratio where  $\Sigma NO_{yi} = NO_x + HNO_3 + \Sigma ANs + PAN + PPN$  (gray circle) decreases from a value of 1.2 where  $\Delta PNs$  are insignificant to close to 0.8 where  $HO_2NO_2$  contributes nearly half of total  $NO_y$ . Conversely, the  $NO_y$  budget using TD-LIF measurements  $\Sigma NO_{yi} = NO_x + HNO_3 + \Sigma ANs + \Delta PN$  (black square) suggests that our measurements account for the predicted contribution of  $HO_2NO_2$  to the reactive nitrogen budget. If the modelled  $\Delta PNs$  are added to the PAN and PPN measurements, the average  $\Sigma NO_{yi}/total\ NO_y$  ratio over the whole campaign is 1.2.

### 3. Model description and comparison to observations

Peroxynitrates are formed when an  $RO_2$  radical (where R = H, alkyl, or acyl group) reacts with  $NO_2$  (R5). In the remote atmosphere,  $HO_2$  and  $CH_3O_2$  make up a large fraction of  $RO_2$ . Sinks of  $HO_2NO_2$  and  $CH_3O_2NO_2$  include thermal decomposition (R6), photolysis (R7), and for  $HO_2NO_2$ , reaction with OH (R8). Note that the products given for (R8) have not been measured. Models all assume that water is a product of (R8) and consequently that (R8) is a sink for  $HO_x$  radicals.

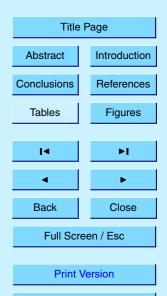
$$RO_2 + NO_2 + M \rightarrow RO_2NO_2 + M \tag{5}$$

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$$RO_2NO_2 + M \rightarrow RO_2 + NO_2 + M \tag{6}$$

$$RO_2NO_2 + h\nu \rightarrow RO_2 + NO_2 \tag{7a}$$

or 
$$\rightarrow RO + NO_3$$
 (7b)

$$HO_2NO_2 + OH \rightarrow H_2O + NO_2 + O_2.$$
 (8)

5 Uncertainties associated with this chemistry are large. For example, for the thermal decomposition of HO<sub>2</sub>NO<sub>2</sub> the JPL panel recommends an uncertainty factor of 5 at 298 K and of 13 at 230 K (Sander et al., 2003). Recent laboratory measurements of the UV (Knight et al., 2002) and IR (Roehl et al., 2002) cross sections of peroxynitric acid greatly improve our understanding of its photolysis rate, (J≡R7), although the temperature and pressure dependencies of the quantum yield require further investigation. We know little about the photolysis of CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub>.

We use an instantaneous steady state model representing (R5)-(R8) to predict the abundances of HO<sub>2</sub>NO<sub>2</sub> and CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub>.

$$[HO_2NO_2]_{SS} = \frac{k_5[HO_2][NO_2][M]}{k_6[M] + J + k_8[OH]}$$
(9)

$$[CH_3O_2NO_2]_{SS} = \frac{k_{5'}[CH_3O_2][NO_2][M]}{k_{6'}[M] + J'}.$$
(10)

Measurements of NO<sub>2</sub>, the spectrally resolved actinic flux, temperature, and pressure are used in the calculations. As a result of the low mixing ratios of NO<sub>2</sub> (50% of the observations have NO<sub>2</sub> below 11 ppt) the precision of the model calculations are roughly proportional to the precision of the NO<sub>2</sub> measurements, for example ±40% at 10 ppt NO<sub>2</sub>. Although OH and HO<sub>2</sub> were measured during the TOPSE campaign, the observations were too infrequent to use in the model. Steady state concentrations of OH, HO2, and CH3O2 were calculated using the photochemical model described by Cantrell et al. (2003), which included the steady state equations for HO<sub>2</sub>NO<sub>2</sub> and 5696

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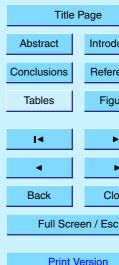
Introduction

References

**Figures** 

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CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub>. Four individual calculations in which radicals and reservoirs were solved self-consistently were used to investigate the main sources of uncertainty in the kinetics of HO<sub>2</sub>NO<sub>2</sub>. Rate constants used in the model are from the JPL-2000 recommendations except for the self-reaction of HO<sub>2</sub> (Christensen et al., 2002) and for the reaction <sub>5</sub> HO<sub>2</sub> + O<sub>3</sub> and OH + O<sub>3</sub> for which we use JPL-97 recommendations (see Lanzendorf et al., 2001). Photolysis rates were determined using measured actinic flux and made use of the recent molecular data for H<sub>2</sub>CO (Smith et al., 2002) and HO<sub>2</sub>NO<sub>2</sub> (Knight et al., 2002; Roehl et al., 2002). The UV cross section of CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> is assumed to be equal to that of HO<sub>2</sub>NO<sub>2</sub>, and we assume it has no infrared component because it lacks the OH group required for significant IR overtone photodissociation. In the first calculation (Model A), the IR photolysis rate of  $HO_2NO_2$  is set at  $1 \times 10^{-5}$  s<sup>-1</sup>, slightly higher than the recommendation of  $8.3 \times 10^{-6}$  s<sup>-1</sup> from Roehl et al., which was given for a system with no albedo. In a second calculation (Model B), the IR photolysis rate for HO<sub>2</sub>NO<sub>2</sub> was set to zero. In two separate calculations, the rate of thermal decomposition of HO<sub>2</sub>NO<sub>2</sub> was varied to the maximum (Model C) and minimum (Model D) limits given in JPL-2003. Model A, with standard thermal decomposition and IR photolysis, calculates peak HO<sub>2</sub>NO<sub>2</sub> and CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> mixing ratios during the last 19 flights of the TOPSE experiment that range as high as 300 ppt and 100 ppt respectively. It predicts  $HO_2NO_2$  reaching as much as 50% of measured  $NO_v$  and  $CH_3O_2NO_2$  as much as 15% of measured NO<sub>v</sub>.

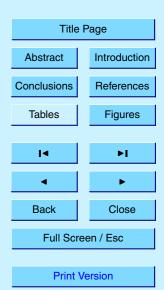
Figure 3 depicts the temperature dependence of the thermal and photochemical sinks of these two peroxynitrates. The photochemical sinks (R7 and R8) are nearly independent of temperature.  $HO_2NO_2$  has a slower thermal decomposition rate and a faster photochemical loss than  $CH_3O_2NO_2$  because of its large infrared photodissociation cross section and because its reaction with OH also contributes to photochemical loss. For  $HO_2NO_2$ , thermal decomposition becomes faster than the photochemical sinks above 240 K, while the crossing point for  $CH_3O_2NO_2$  is near 220 K. One consequence of this is that the lifetime of  $HO_2NO_2$  (~12 h in sunlit conditions) is independent of temperature for the colder conditions sampled during TOPSE.

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The instantaneous steady state model is valid only when the lifetime of the HO<sub>2</sub>NO<sub>2</sub> or CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub> is short compared to the rate of change of its sources and sinks. This is not always true for these peroxynitrates under the coldest conditions sampled during TOPSE. However, many of the data points were collected at high latitudes during the late spring, where the photochemistry has a fairly shallow diurnal profile, making the model steady state predictions reasonable estimates. To help avoid any bias due to being strongly out of steady state, model results in which the lifetime of either compound exceeded 12 h were discarded, unless there were more than 20 h of sunlight per day at that location.

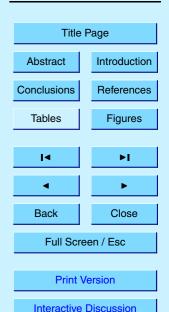
In the lower panel of Fig. 3, the ΔPN concentrations inferred from our observations are displayed versus temperature. The mixing ratios below approximately 250 K are larger than at warmer temperatures. The wide range of concentrations at each temperature derives partly from variation in the abundance of HO<sub>2</sub>, CH<sub>3</sub>O<sub>2</sub> and NO<sub>2</sub> radicals in the airmass. As an example, the large symbols are the mean  $\Delta PNs$  within a 5 K bin for NO<sub>2</sub> mixing ratios above (black squares) and below (open circles) 11 ppt, which was the median NO<sub>2</sub> during the last 19 flights of TOPSE. On average, higher values of ΔPNs were observed when NO<sub>2</sub> mixing ratios were higher than average and vice versa. In addition to being proportional to NO<sub>2</sub>, below 250 K ΔPNs are calculated to depend strongly on the IR photolysis of HO<sub>2</sub>NO<sub>2</sub>, which has only recently been recognized as large enough to affect its atmospheric mixing ratio. Our data provides direct observational evidence for this term. In the upper panel of Fig. 4 we show the difference between the  $\Delta PN$  observations and model B (no IR photolysis) and in the lower panel we show the difference between  $\Delta PNs$  and model A. The relative significance of IR photolysis increases strongly as a function of decreasing temperature and dominates over thermal decomposition as a sink of HO<sub>2</sub>NO<sub>2</sub> below 235 K. At 230 K, the model without IR photolysis has, on average, 200 ppt more HO<sub>2</sub>NO<sub>2</sub> than the observed ΔPNs. There are specific events where modeled ΔPNs exceed the measurements by 400 ppt without the infrared sink of HO<sub>2</sub>NO<sub>2</sub>.

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The model A peroxynitrate ( $HO_2NO_2 + CH_3O_2NO_2$ ) predictions are compared to  $\Delta PN$  measurements in a time series during a TOPSE flight on 22 May (Fig. 5 upper panel). This flight between Thule and Winnipeg sampled a wide range of tropospheric conditions and encountered air of mixed stratospheric-tropospheric character ( $O_3 > 100 \, \mathrm{ppb}$ ) near the end of the flight.  $HNO_3$  and  $NO_x$  make significant contributions to  $NO_y$  in this particular airmass, which nonetheless has quite high PAN mixing ratios. The  $\Delta PN$  measurements are generally scattered around the model predictions and tend to be highest under the coldest conditions as expected based on the chemistry described above. During this flight,  $NO_x$  and  $\Delta PNs$  are nearly equal.

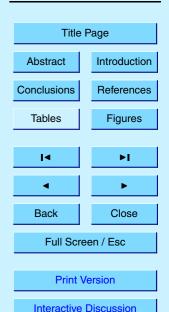
The figures above show that  $\triangle PNs$  vary as expected in response to two of the individual terms in the steady-state equations (J and NO<sub>2</sub>) and vary along a flight track as the steady state calculation suggests they should. Figure 6 shows a direct comparison of calculations of ΔPNs and the observations. Within the statistical variance, the observations and model agree. However model predictions exceed the observations at higher mixing ratios and the fact that nearly all calculated concentrations above 150 ppt are higher than observations suggests the effect is statistically reliable. If the steady state approximation is correct, then this could indicate an error in the photochemistry of HO<sub>2</sub>NO<sub>2</sub> or the thermal decomposition of CH<sub>3</sub>O<sub>2</sub>NO<sub>2</sub>. We investigated the factor of 5–15 uncertainty in the HO<sub>2</sub>NO<sub>2</sub> thermal decomposition rate at temperatures in the 220-298 K range. Model calculations (C, D) that incorporated the most extreme values for rates of thermal decomposition predicted ΔPN concentrations that differed at most by 100 ppt, with the greatest effect occurring between 235 and 250 K. Both models were less consistent with the observations than Model A, but because HO2NO2 is largest where thermal decomposition is slow compared to photolysis we are unable to define substantially more precise limits on the rate of thermal decomposition than recommended by the JPL panel. We examined other possible effects but found no single explanation that would bring the model-measurement comparison into agreement.

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#### 4. Discussion

In the upper troposphere, several studies (Folkins et al., 1997; Jaegle et al., 2000; Jaegle et al., 2001) have described calculations indicating that reaction of OH with  $HO_2NO_2$  is a dominant sink for  $HO_x$ . Analysis of  $HO_x$  observations by Faloona et al. (2000) suggest that a model in which the  $HO_2NO_2$  equilibrium constant is reduced by a factor of five, or with  $HO_2NO_2$  formation suppressed entirely, provides a more accurate representation of upper tropospheric  $HO_x$  observations than the standard model. We estimate that inclusion of the IR photolysis of  $HO_2NO_2$  should have nearly the same effect on  $HO_2NO_2$  concentrations at the altitudes where most of the observations described by Faloona et al. were collected. The main consequence is a decrease in the modelled abundance of  $HO_2NO_2$ . However, the IR photolysis will have a distinctly different temperature dependence than the suggested change to the equilibrium constant. For the coldest conditions sampled during TOPSE, between 220 K and 240 K, Model A, which includes IR photolysis, predicts ~10% more  $HO_x$  than Model B, and also results in slightly different  $HO_x$  partitioning.

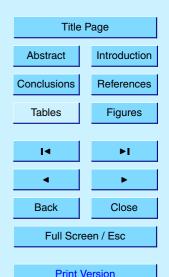
Another interesting aspect of  $HO_2NO_2$  and  $CH_3O_2NO_2$  chemistry is their behavior as a short-term reservoir for HOx and  $NO_x$  radicals. Under sunlit conditions, the radicals and alkyl peroxynitrates should reach steady state, yet once the sun sets, formation diminishes and the only remaining sink is thermal decomposition. During 12 h of darkness in the upper troposphere at 240 K, all the  $CH_3O_2NO_2$  and over half of the  $HO_2NO_2$  will have thermally dissociated, releasing on the order of tens to hundreds ppt of  $CH_3O_2$  and  $HO_2$ . The fate of these radicals during the night will depend strongly on the abundance of  $NO_x$  and  $O_3$ . This fact underscores the importance of making measurements at night to challenge our understanding of radical sources and sinks.

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#### 5. Conclusions

We describe observations of  $\Delta PNs$  and demonstrate that they are  $HO_2NO_2$  and  $CH_3O_2NO_2$ . The observations show that during the spring high latitude conditions sampled during the TOPSE experiment, the contribution of  $\Delta PNs$  to  $NO_y$  is highly temperature dependent, with  $\Delta PNs$  on average 30% of  $NO_y$  at 230 K and 7% of  $NO_y$  at 250 K. The magnitude and observed temperature dependence of the mixing ratios is consistent with recent laboratory evidence that the J value for  $HO_2NO_2$  has a large IR component. Prior analyses of the  $NO_y$  budget in the upper troposphere that did not take into account  $HO_2NO_2$  or the new J value should be revisited.

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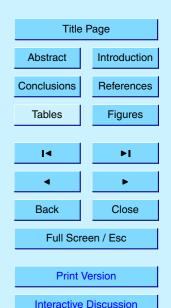
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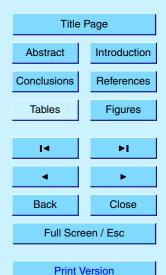


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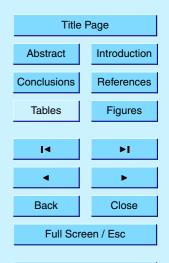
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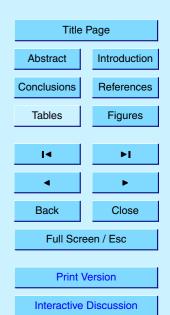
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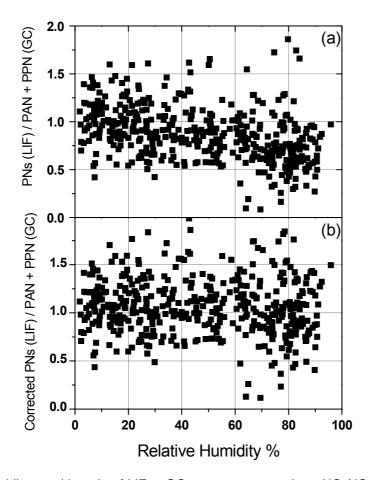
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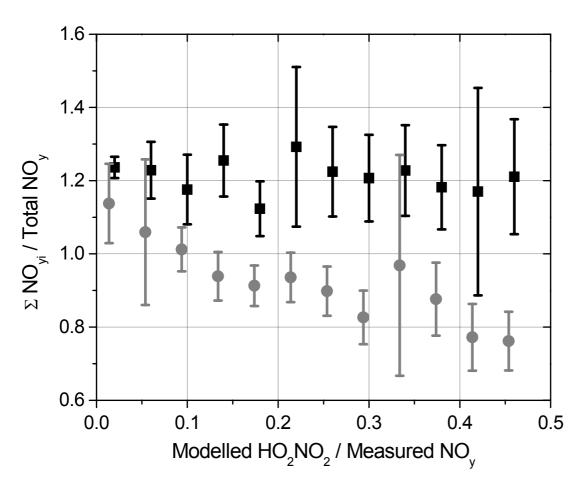
**Fig. 1. (a)** Humidity trend in ratio of LIF to GC measurements where  $HO_2NO_2$  is predicted be less than 10% of PAN **(b)** Ratio after the humidity-dependent correction has been applied.

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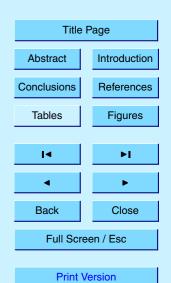


**Fig. 2.**  $\Sigma NO_{yi}/NO_y$  vs calculated  $HO_2NO_2/NO_y$ , using  $\Sigma NO_{yi} = NO_x + HNO_3 + \Sigma ANs + PAN + PPN (grey) and <math>\Sigma NO_{yi} = NO_x + HNO_3 + \Sigma ANs + \Sigma PNs$  (black). The data were averaged within ±.04 of  $HO_2NO_2/NO_y$  and the error bars represent twice the standard deviation of the mean.

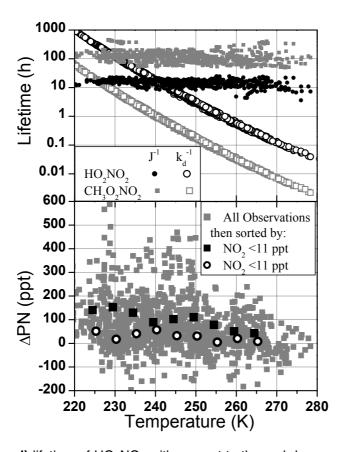
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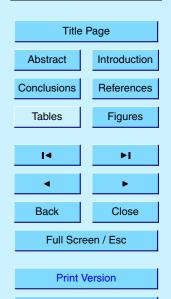


**Fig. 3.** (Upper panel) lifetime of  $HO_2NO_2$  with respect to thermal decomposition (large black circle) and the sum of photolysis and reaction with OH (large black dot) and of  $CH_3O_2NO_2$  with respect to thermal decomposition (large grey square) and photolysis (small grey square). (**Lower panel**)  $\Delta PN$  measurements (small grey square) as a function of temperature, averaged into separate temperature bins for  $NO_2 < 11$  ppt (big black circle) and for  $NO_2 > 11$  ppt (large black square).

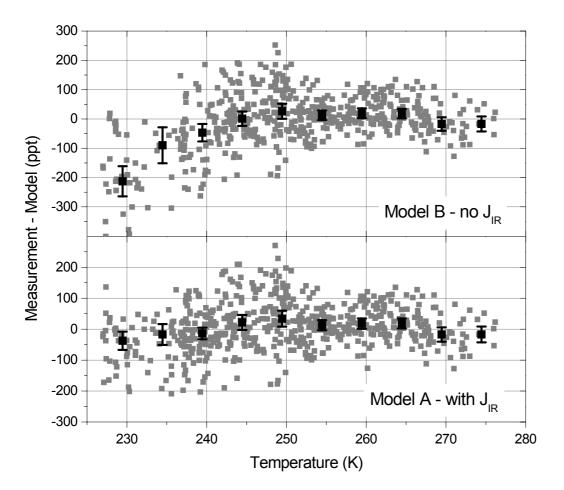
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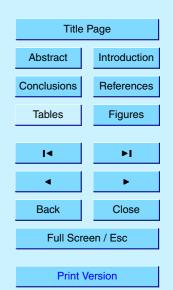


**Fig. 4.** Difference between measured and modelled  $\Delta PNs$  for a model without IR photolysis (**upper panel**) and a model including IR photolysis for  $HO_2NO_2$  (**lower panel**). The black squares represent the median of each 5 K bin and the error bars are twice the standard deviation of the mean.

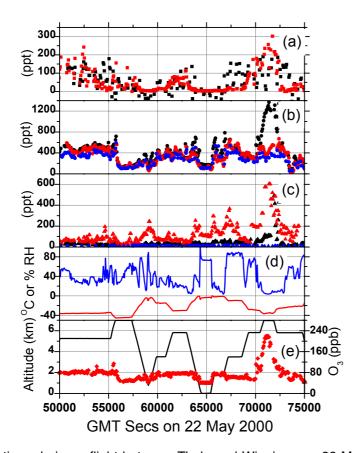
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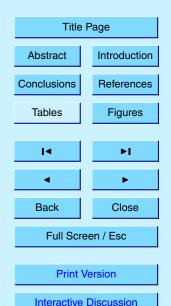


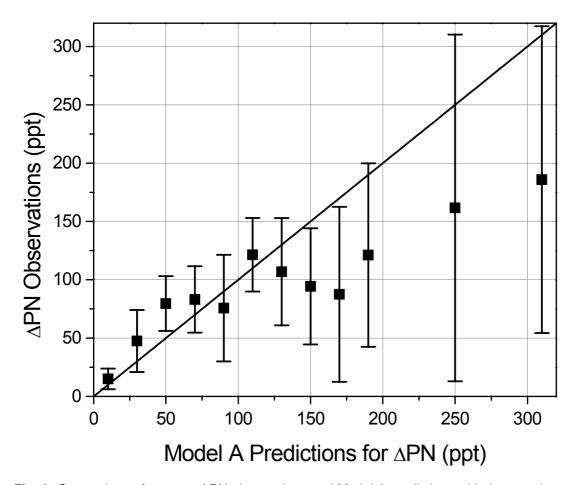
**Fig. 5.** Observations during a flight between Thule and Winnipeg on 22 May 2000, **(a)** ΔPN measurements (small black square) and predictions from Model A (small red square), **(b)** NO $_{y}$  (black dot), PAN + PPN (blue dot), ΣPN (red dot), **(c)** HNO $_{3}$  (red triangle), NO $_{x}$  (black triangle), ΣAlkyls (blue triangle), **(d)** temperature (red line), % relative humidity (blue line), **(e)** altitude (black line), O $_{3}$  (red diamond).

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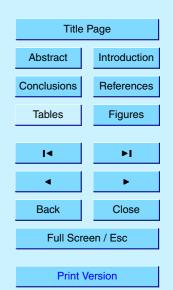


**Fig. 6.** Comparison of average  $\Delta$ PN observations and Model A predictions with the error bars representing twice the standard deviation on the mean and a 1:1 line shown.

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