

Transport of Antarctic stratospheric strongly dehydrated air into the troposphere observed during the HALO-ESMVal campaign 2012

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Abstract

Dehydration in the Antarctic winter stratosphere is a well-known phenomenon that is **annually observed by satellites and** occasionally observed by balloon-borne measurements. However, in-situ measurements of **dehydrated air masses** in the Antarctic vortex are very rare. Here, we present detailed observations with the in-situ and GLORIA remote sensing instrument payload aboard the German aircraft HALO. Strongly dehydrated air masses down to 1.6 ppmv of water vapor were observed as far **northward** as 47° S **in an altitude** between 12 and 13 km in **the lowermost stratosphere**. The dehydration can be traced back to individual ice formation events **above the Antarctic Peninsula and Plateau**, where ice crystals sedimented out and water vapor was irreversibly removed. Within these dehydrated stratospheric air masses, filaments of moister air reaching down to the tropopause are detected with the high resolution limb sounder, GLORIA. Furthermore, dehydrated air masses are observed with GLORIA in the Antarctic **lowermost stratosphere** down to 7 km. With the help of a backward trajectory analysis, a **mid-latitude** origin of the moist filaments in the vortex can be identified, while the dry air masses **down to 7 km** have stratospheric origins. Antarctic stratosphere/troposphere exchange (STE) and **transport of dehydrated air masses into the troposphere is investigated**. Further, it is shown that the exchange process can be attributed to several successive Rossby wave events in combination with an isentropic **exchange** of air masses across the **thermal tropopause**. **The transport into the troposphere is caused by air masses that are detached from the PV structure by Rossby wave breaking events and subsequent transported diabatically across the dynamical tropopause**. Once transported to the troposphere, air masses with stratospheric origin **can** reach near-surface levels within **several days**.

1 Introduction

Antarctic stratospheric dehydration occurs regularly every winter and spring in the very isolated **and very cold** southern hemispheric polar vortex

(e.g. Kelly et al., 1989; Schoeberl et al., 1992; Vömel et al., 1995; Nedoluha et al., 2002; Jimenez

The reason is that the temperatures in the Antarctic stratosphere are regularly sufficiently low for ice particles to form. These ice particles will sediment with a rate of 2.5 km d^{-1} depending on their size and will thus transport water to lower altitudes. If dehydration extends down to the tropopause, water is removed from the stratosphere and changes the stratospheric water vapor budget of the lower Antarctic vortex drastically. After the break-up of the Antarctic vortex, redistribution and exchange between vortex air and air masses outside the vortex lead to a return to typical stratospheric background water vapor mixing ratios of about 4–5 ppmv (Nedoluha et al., 2007). Rehydration with redistribution of water from the solid into the gas-phase occurs at lower altitudes where ice crystals sublimate caused by sub-saturation of water vapor at higher temperatures. In the Arctic region, stratospheric dehydration is also possible but much less frequent than in the Southern Hemisphere (e.g. Schiller et al., 2002; Khaykin et al., 2013).

Dehydration of Antarctic vortex air masses is observed frequently by satellite measurements (e.g. Nedoluha et al., 2002; Jimenez et al., 2006; Schoeberl and Dessler, 2011). However, satellite measurements in the lower stratosphere and upper troposphere (UT/LS) do not offer the necessary horizontal and vertical resolution to be suitable for process studies in the tropopause region. Balloon borne measurements launched from the Antarctic continent also observed dehydration and the corresponding rehydration (e.g. Vömel et al., 1995). Beside these balloon soundings, only three aircraft campaigns have been conducted where Antarctic stratospheric air masses were probed in-situ; two US campaigns with the ER-2 aircraft (AAOE in 1987; Tuck et al., 1989, and ASHOE in 1994; Tuck et al., 1997) and one European campaign with the Geophysica aircraft (APE-GAIA in 1999; Giovanelli et al., 2005). Airborne in-situ measurements in the Antarctic vortex are very rare, since the flight range of research aircraft is limited. This paper focuses on one measurement flight directly into the Antarctic vortex that was performed as part of the ESMVal (Earth System Model Validation; Schlager, 2014) campaign in September 2012 with the German research aircraft HALO (High Altitude and Long range; Krautstrunk and Giez, 2012). The ESMVal campaign was embedded in the TACTS (Transport And Composition in the Upper

Troposphere/Lowermost Stratosphere; Engel et al., 2013) campaign which took place just before and after the ESMVal flights in August and the end of September 2012. A mutual scientific objective of both campaigns was to investigate the transition region between the troposphere and stratosphere. HALO was equipped with in-situ instruments as well as the high resolution limb sounder, GLORIA (Gimballed Limb Observer for Radiance Imaging of the Atmosphere). This combination of high-precision in-situ measurements together with sophisticated remote sensing observations in the UT/LS is an outstanding attribute of both campaigns. Another objective of the ESMVal flights was to **obtain a full meridional cross section of atmospheric measurements for global chemistry model evaluation. The Antarctica flight was performed in an effort to penetrate as far south as possible and reach the stratospheric vortex. During this flight, dehydrated air masses were measured in-situ far north up to 45° S between 12 and 13 km. In contrast, Aura-MLS and POAM 3 satellite measurements (Nedoluha et al., 2002; Jimenez et al., 2006; Schoeberl and Dessler, 2011) do not show dehydrated air masses as far north (not beyond 57° S) and as low in the stratosphere and upper troposphere as was observed during this ESMVal flight. This is mostly caused by the poor resolution of satellites in the UT/LS region.**

In general, the process of dehydration is well understood. Relatively **little** is known about the fate of the dehydrated air masses **at the bottom of the stable vortex. Tuck et al. (1997) found out that dehydrated air masses descend downwards through the bottom of the vortex down to the thermal tropopause. Normally, the air within the Antarctic polar vortex is highly isolated with a weak exchange of trace gases across the vortex edge driven by stratospheric planetary Rossby waves propagating from the troposphere and Rossby wave breaking (RWB) events. However, the RWB events mainly disturb the bottom of the polar vortex, where the vortex edge barely exists as an isolating mechanism. Mixing and transport of dehydrated air masses from the Antarctic lower stratosphere towards mid-latitudes is also indicated by previous aircraft mission with the ER-2 (Tuck, 1989). Thus, the Antarctic lowermost stratosphere consists of dehydrated air masses frequently mixed with moist mid-latitude air (Tuck et al., 1997; Lee et al., 2002). The vortex edge in the Southern Hemisphere is typically less strongly disturbed than in the Arctic due to less wave activity**

(Schoeberl et al., 1992). For instance, Khosrawi et al. (2006) presented a case study where at the edge of a midlatitude deep intrusion in the northern hemisphere also dehydrated air masses were observed.

Nevertheless, Antarctic stratospheric air masses can also be exchanged with tropospheric air masses (stratosphere/troposphere exchange (STE)) and can be transported deep into the troposphere creating a dry tongue reaching down to the Earth's surface. This air is rich in ozone and reactive nitrogen due to the stratospheric origin and can influence the chemical composition of the Antarctic troposphere (Stohl and Sodemann, 2010; Mihaikova and Kirkwood, 2013). The thermal tropopause, for instance, is characterized by a low temperature lapse rate (≤ 2 K/km) and coincides with a change in the static stability. Thus, the thermal tropopause separates the unstable troposphere from the stable stratosphere. However, the thermal tropopause in the Antarctic region is formed under unique climate conditions (Evtushevsky et al., 2011). The upper troposphere and lower stratosphere are generally very cold with a very weak vertical temperature gradient during winter and spring. Therefore, the thermal tropopause in the Antarctic region is rather poorly defined (Stohl and Sodemann, 2010). As a consequence, transport across the thermal tropopause can be either along isentropes crossing the thermal tropopause or diabatic with reduction in potential temperature. In addition, Zangl and Hoinka (2001) showed that the definition of the thermal tropopause in the Antarctic winter is not very meaningful in contrast to the potential vorticity (PV)-based tropopause definition.

Stohl and Sodemann (2010) identify two general processes that transport stratospheric air masses down to near-surface levels. The first process might involve katabatic winds at the surface over the Antarctic Plateau caused by the high topography that create a general downwelling above the Antarctic continent as suggested by Roscoe (2004). The second process is driven by mid-latitude cyclones on the poleward side of the jet stream that support RWB events and a corresponding stratospheric intrusion as reported by Ndarana et al. (2012). Downward transport could be either very fast along tilted isentropes on time scales of 1-5 days (Škerlak et al., 2014) or slower by radiative cooling and reduction in potential temperature. In the latter case, air masses can potentially descend from the ther-

mal tropopause to near-surface heights within 10 days (descent rate of 5 mm s^{-1}) or longer as reported by van de Berg et al. (2007). Rossby wave induced stratospheric intrusions, such as tropopause folds, occur more often further north and in the midlatitudes than directly above the Antarctic continent (Sprenger et al., 2003; Škerlak et al., 2014). However, the frequency, the seasonality, and the process behind tropopause folds and stratospheric intrusions in the Antarctic region is still under debate (Stohl and Sodemann, 2010; Ndarana et al., 2012; Mihalikova and Kirkwood, 2013).

The following research questions will be answered, based on the observations of the ES-MVal flight (13 September 2012): 1) Can we confirm dehydration in the lowermost stratosphere and at the bottom of the Antarctic vortex? 2) How deep do dehydrated air masses from the vortex descend into the troposphere? 3) Where and how have Antarctic dehydrated stratospheric air masses been transported?

To answer the questions above the study presented here is structured as follows: in Sect. 2, a brief overview of the different instruments and data used is given. The in-situ and remote sensing measurements across the Antarctic polar vortex are described in Sect. 3. Furthermore, Sect. 4 includes an extensive trajectory case study: First, dehydration in the in the Antarctic lowermost stratosphere is analyzed. Second, the transport of dehydrated air masses and the question how deep below the thermal tropopause dehydrated air masses can be found is investigated.

2 Instrumentation and meteorological data

The German research aircraft, HALO, deployed during TACTS and ESMVal has a long flight endurance of up to 12 h. This allows air masses in the Antarctic vortex to be sampled without landing in Antarctica. Altogether, a set of nine in-situ instruments for measuring trace gases and one remote sensing instrument were installed aboard HALO. For the study presented here, we use the water vapor data of FISH and HAI as well as measurements from TRIHOP for methane and FAIRO for ozone. The remote sensing instrument GLORIA, provides cross-sections of trace gases that are approximately parallel to the flight track of the aircraft.

GLORIA's high vertical resolution makes it particularly suited for the investigation of small-scale structures.

In addition to the aircraft measurements, satellite observations from CALIPSO and meteorological **reanalysis** data from ECMWF are used for further interpretation of the observed situation. The instruments and meteorological data are described in the following subsections. **The thermal tropopause definition in this study based on the WMO definition (WMO, 1957) using the lowest height at which the temperature lapse rate decreases to ≤ 2 K/km and remains below this value for a range of at least 2 km.**

2.1 FISH – total water measurements (in-situ H₂O)

The airborne Lyman- α photofragment fluorescence hygrometer FISH (Fast In-situ Stratospheric hygrometer) is a well-established closed-path instrument for measuring water vapor in the range of 1 to 1000 ppmv (Zöger et al., 1999). The FISH is especially built to measure the low water vapor mixing ratios prevailing in the upper troposphere and lower stratosphere. It is regularly calibrated on the ground to a reference frost point hygrometer (MBW DP30) and had an accuracy of $\pm(6\% + 0.4 \text{ ppmv})$ during TACTS and ESMVal campaigns (Meyer et al., 2015). The air supply for the measuring cell is provided by a forward facing inlet, which also samples ice crystals if present (H₂O_{cond}). Together, the evaporated crystals and the gas-phase water (H₂O_{gas}) sum up to a total water measurement.

2.2 HAI – total water measurements (in-situ H₂O)

The Hygrometer for Atmospheric Investigation (HAI) (Buchholz, 2014) is a fast, airborne, in-situ hygrometer capable of multi-phase water detection that employs direct Tunable Diode Laser Absorption Spectroscopy (dTDLAS) in combination with a special data retrieval method, which allows absolute water vapor measurements without any sensor calibration (Buchholz et al., 2013). This data retrieval was recently successfully validated via a direct side by side comparison with the German metrological water vapor scale (Buchholz et al., 2014). HAI is a 2×2 , multi-channel TDLAS spectrometer which realizes, in a unique

concept, a simultaneous gas phase and total water measurement via a parallel water detection in an open-path cell, outside the aircraft fuselage, and in extractive (i.e. gas sampling) closed-path cells inside the aircraft. To enlarge the dynamic range to 1–40 000 ppmv, both cell types are simultaneously probed by two laser wavelengths, at 1.4 μm for the lower to mid troposphere and at 2.6 μm for the upper troposphere and the stratosphere. The ranges of the 1.4 and 2.6 μm spectrometer channels overlap and therefore provide a redundant, independent detection between 10 and about 5000 ppmv. The metrologically calculated spectrometer uncertainties are $\pm(4.3\% + 3 \text{ ppmv})$ at 1.4 μm and $\pm(5.9\% + 0.4 \text{ ppmv})$ at 2.6 μm . In this study, data from the two closed-path spectrometers are used, i.e. total water measurements similar to FISH.

2.3 TRIHOP – methane measurement (in-situ CH_4)

The TRIHOP instrument is a three channel Quantum Cascade Laser Infrared Absorption spectrometer capable of the subsequent measurement of CO , CO_2 , N_2O and CH_4 . During TACTS/ESMVal the instrument was calibrated in-situ against secondary standards of compressed ambient air which are traceable against NOAA standards. Integration time for each species was 1.5 s at a duty cycle of 8 s, which finally limits the temporal resolution of the measurements. For the ESMVal flight on 13 September 2012, TRIHOP CH_4 -data achieved a precision (2σ) of 9.5 ppbv and accuracy of 13.5 ppbv, respectively, without any corrections applied.

2.4 FAIRO – ozone measurement (in-situ O_3)

FAIRO is a new accurate ozone instrument developed for use on board the HALO aircraft. It combines two techniques, the UV photometry (light absorption of O_3 at $\lambda = 250\text{--}260 \text{ nm}$) with high accuracy and chemiluminescence detection with high measurement frequency. A UV-LED is used as a light source for the UV photometer, which can be controlled well (in contrast to Hg lamps) for constant light emission. The 1-sigma precision is 0.08 ppbv at a measurement frequency of 4 s and a cuvette pressure of 1 bar and the total uncertainty

is 2%. The chemiluminescence detector shows a measurement frequency of 12.5 Hz and a high precision of 0.05 ppbv (at 10 ppbv absolute, a measurement frequency of 5 Hz, and a pressure of 1 bar) (Zahn et al., 2012).

2.5 GLORIA (aircraft remote sensing)

The Gimballed Limb Observer for Radiance Imaging of the Atmosphere (GLORIA) combines a two-dimensional focal plane detector array with a Fourier-transform spectrometer to capture about 6000 infrared limb spectra simultaneously. This enables remote sensing observations with high vertical and horizontal resolution to resolve small scale structures (see Riese et al., 2014). The spectral sampling can be switched between 0.625 and 0.0625 cm^{-1} at the cost of an increased acquisition time in case of higher spectral resolution (Friedl-Vallon et al., 2014). The gimbal mount and inertial altitude control system allows GLORIA to maintain a steady pointing on a moving aircraft; it can also be used to point the instrument at a range of azimuth angles with respect to the aircraft, covering about 70°. This allows for tomographic measurement patterns and 3-D reconstruction of fine-scale filamentary structures (Ungermann et al., 2011; Kaufmann et al., 2014). The vertical location of retrieved quantities approximately follows the tangent points of the measurement geometry (parabola curve through the atmosphere). Thus, while the location of quantities retrieved at flight level lies in the vicinity of the aircraft, the locations of quantities at lower altitudes are **horizontally** several tens to hundreds of kilometers away **from the flightpath**. **Assuming a flight altitude of 13 km the water vapor observation at 12 km is horizontally 113 km away from the aircraft while at 8 km it is already 250 km.** Concerning the water vapor product used in this paper, **GLORIA achieved a vertical and horizontal resolution of about 500 m and 100 km, respectively. The water vapor precision and total error (accuracy + precision) amounts to about 0.2 ppmv and 1.0 ppmv in the stratosphere.** Ungermann et al. (2015) showed that GLORIA water vapor at flight level agrees fairly well with the in-situ FISH measurements, within error bars. The deviations to FISH during the ESMVal/TACTS flights are mostly less than 0.4 ppmv.

2.6 CALIPSO (satellite remote sensing)

The Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observation (CALIPSO) satellite is one of five satellites in the NASA A-train constellation. CALIPSO completes 14.55 orbits per day with an inclination of 98.2° and thus delivers good coverage above the polar regions. Besides one wide field camera and an imaging infrared radiometer, the Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) is aboard CALIPSO. The lidar operates with two wavelengths (532, 1064 nm) with additional polarization-sensitivity, providing high-resolution vertical backscatter profiles of aerosols and clouds. In this study we use the CALIPSO Lidar Level 2 Polar Stratospheric Cloud (PSC) data product, which is described in Pitts et al. (2009) and Pitts et al. (2011). This data product provides a PSC composition scheme on a daily basis for all nighttime orbits with a resolution of 5 km horizontally by 180 m vertically.

2.7 ECMWF (meteorological data)

Global meteorological reanalysis ERA-Interim data (Dee et al., 2011) of the ECMWF (European Centre for Medium-Range Weather Forecasts) are used to **interpret** the observations. The meteorological fields have a resolution of $1^\circ \times 1^\circ$ with 60 vertical levels from the surface (1000 hPa) up to 0.1 hPa. Every 6 h a full global dataset is available. The trajectories used in Sect. 4 are based on the horizontal wind fields and diabatic heating rates; additional parameters like temperature, pressure etc. are added to the trajectories from the ERA-Interim data (**for details see Sec. 4**).

3 ESMVal flight on 13 September 2012

3.1 Meteorological situation/flight pattern

The **general dynamical situation of the Antarctic vortex is shown in Figure 1a. High PV values of < -9 PVU denotes the Antarctic vortex at the 350 K isentrope which corresponds**

approximately to the flight level. The vortex edge is marked with a blue line according to the definition of Nash et al. (1996). The vortex itself is shifted about 10° towards Africa indicating a wave one pattern caused by Rossby wave activity. This results in high PV values up to 47° S, where the aircraft in-situ measurements confirm the presence of Antarctic vortex air as will be shown later (see Section 3.2). In the lowermost stratosphere below the vortex, represented by the 310 K isentrope (see Figure 1b), the PV structure with high PV values around -6 to -3 PVU is more jagged but mostly connected. Here, a strong Rossby wave activity is visible and also RWB events associated with isolated PV structures are discernible. A blob with high PV at 60° to 47° S and 10° to 30° E indicates an ongoing RWB event. These air masses were probed during the ESMVal flight on 13 September 2012. The flight was performed from Cape Town (South Africa) heading towards Antarctica until 65° S at an altitude of 12–13 km (see Fig. 2). After a dive from 12.5 km down to 3 km to sample the transition from the polar stratosphere to the troposphere, the return flight took place at 12.5 km and rose up to 15 km shortly before descending back to Cape Town. A hexagonal flight pattern was included in the flight path back to Cape Town to enable the tomographic study of trace gases with GLORIA. The ECMWF wind field (black lines) as well as the strongly increasing potential vorticity (PV, see Fig. 3) show that the vortex edge (yellow dots, **determined with the Nash criterion**) was far north, up to 47° S **down to the 340 K level**. The color code in Fig. 2a represents the in-situ measured water vapor mixing ratios from FISH along the flight path. Directly south of the vortex edge the water vapor mixing ratios decrease rapidly to below 2.5 ppmv (purple dots along the flight path).

The latitudinal cross section of **PV from ECMWF** along the flight path is shown in Fig. 2b. The thermal tropopause (**cyan** dots) stays almost constant at around 10 km and rises up to 12 km further poleward. The PV and **also the equivalent latitude (not shown)** reveal high values within the vortex and thus **clearly mark** the vortex air masses. However, high **PV values of -2 PVU** show up far below the thermal tropopause down to 5 km altitude in the latitude range from 60 – 45° S. **The difference compared to a tropopause fold at the mid-latitude jet is that the air masses directly originate from the Antarctic vortex and were strongly dehydrated. Thus, the air masses have a different chemical composition than the "normal" air**

masses in the mid-latitude lowermost stratosphere. This indicates that these air masses originate from higher altitudes and were transported through the thermal tropopause and form a stratospheric intrusion with a low PV-based tropopause. Interestingly, the isentropes between 310 K and 320 K intersect the thermal tropopause on the poleward side of the jet core. Therefore, air masses can be potentially transported along these tilted isentropes to the low altitudes across the thermal tropopause into the low PV structure. The flight path crosses the air masses with high PV in the stratosphere before and after the dive. The dive down to 3.5 km altitude was performed on the poleward side of the air masses with high PV below the thermal tropopause.

3.2 Observations

The time series of the in-situ measurements are shown in Fig. 3, where the measurements of the water vapor instruments are visible in the upper panel and the measurements of the tracers for defining stratospheric/vortex air masses are shown in the lower panel. After the ascent of HALO near Cape Town, the water vapor mixing ratio decreases to the typical stratospheric value of 4 ppmv, from 06:30 until 08:15 UTC. During this time, ozone and PV (i.e. the stratospheric tracers with low values in the troposphere) show an increase compared to the values before 06:30 UTC, while methane (tropospheric tracer with low values in the stratosphere) shows a slight decrease from 1790 to 1740 ppbv. Both indicates that the aircraft entered the stratosphere. Henceforth, methane and water vapor further decrease to 1650 ppbv and 3 ppmv respectively on the same flight level starting at 08:15 UTC, and simultaneously ozone and PV further increase. This is the time when HALO penetrated the Antarctic vortex (marked in blue in Fig. 3). After ascending to a higher flight level (08:35 UTC), water vapor decreases to slightly above 2 ppmv and stays there until the southernmost point was reached and the dive started at 11:00 UTC. This low water vapor in the time between 08:15 and 11:00 UTC coincides well with high PV values (see Fig. 2b). Subsequently, the strong increase in methane and decrease in ozone and PV indicate the penetration of the aircraft into the Antarctic troposphere at 11:00 UTC. After returning to the previous flight level (11:45 UTC), the vortex signatures with low values of water vapor and

methane as well as high values of ozone and PV show up again. However, at this point ($\sim 12:30$ UTC) the water vapor mixing ratio reaches the lowest value of 1.6 ± 0.5 ppmv together with a strong maximum in ozone and PV. The low water vapor below 4 and even below 2 ppmv indicates strongly dehydrated air masses.

The saturation mixing ratio with respect to ice within the vortex is three to four times larger than the measured water vapor mixing ratio. Thus, the sampled air masses were clearly sub-saturated with relative humidities of 25–33 % and it is unlikely that ice particles remained in the probed air masses. In addition, HAI and FISH show no peaks in the total water vapor time series (based on 1 Hz data), which would indicate the presence of ice particles on the flight level. Especially, the water vapor mixing ratios around 2 ppmv during the stratospheric flight legs (06:30 to 11:00 and 11:45 to 15:45 UTC) are very smooth and show no significant variations. Thus, both water vapor instruments, FISH and HAI, measured only gas-phase water vapor, and no cloud particles were present in the vortex air. Both FISH and HAI observed these low water vapor mixing ratios independently.

In order to show the fairly good agreement of both water vapor instruments FISH and HAI based on 1 Hz data at the observed low water vapor values, we choose two flight legs for the comparison (leg one: 09:45–10:45 UTC and leg two: 13:00–14:00 UTC). During flight leg one, both hygrometers have an absolute difference of 0.21 ppmv and a mean relative difference of -14.9% ($\pm 10.5\%$, 1σ). During flight leg two, a difference of 0.26 ppmv and a mean relative difference of -5.9% ($\pm 7.3\%$, 1σ) is found. These rather small differences are consistent with the uncertainties of both hygrometers, which are ± 0.45 and ± 0.5 ppmv at the 2 ppmv level for FISH and HAI, respectively. The HAI data measured before 09:10 UTC (marked with gray in Fig. 3) are influenced by an untypical memory effect of the measurement cell during ascent, caused by a valve which was opened too late, but only when flying in the upper troposphere. As a result, the “wet” measurement cells were not sufficiently dried off with the high mass flows during take-off and ascent. This also explains the asymmetry in the deviations to FISH before and after the dive. In addition, the agreement of GLORIA water vapor observations at flight level with the two in-situ instruments, FISH and HAI, is remarkably good, considering that GLORIA is a remote sensing instrument.

The GLORIA instrument measures quasi-vertical profiles along the flight path by viewing the atmosphere on the right side of HALO. Here, we focus on the H_2O retrieval product, which is shown in Fig. 4. We consider only the times from 08:00 to 14:00 UTC to focus on a time period when GLORIA observed vortex air (the whole dataset is from 06:30 to 15:30 UTC). The dive is noticeable as the white area between 11:00 and 12:00 UTC, where GLORIA did not measure in order to prevent condensation of tropospheric water vapor on the cold instrument. The dry vortex air masses are clearly visible between 08:00 and 13:00 UTC at flight level (black solid line in Fig. 4), but also below, down to altitudes of 7 km. As mentioned in Sect. 2.5, the quantities retrieved from GLORIA are approximately placed on a parabola following the tangent points through the atmosphere (e.g. 250 km horizontal distance to flight path at 9 km altitude). The dry region just before the dive was measured by GLORIA west of the flight path on the way towards Antarctica, while the air masses after the dive are measured east of the flightpath on the way back to Cape Town. Thus, the dehydrated air masses below the thermal tropopause cover a large region, having a dimension of at least 500 km horizontally at 9 km altitude.

The thermal tropopause is also derived from the GLORIA retrieved temperature profiles and is marked with black dots. Interestingly, low water vapor mixing ratios around and below 2 ppmv are observed beneath the thermal tropopause. Especially in the time range from 09:00 to 10:30 UTC and from 12:00 to 13:30 UTC, very dry air masses with water vapor mixing ratios of 2–3 ppmv can be found far below the thermal tropopause, down to 7 km. This indicates that dehydrated stratospheric air masses could have been transported through the thermal tropopause. The dynamic tropopause, which is between the -4 and -2 PVU isoline, is somewhat lower than the thermal tropopause at ~ 7 km in the time range from 12:00 to 13:30 UTC and indicates a proceeding stratospheric intrusion. In Sect. 4.2, we analyze this transport process with the help of air mass backward trajectories. In addition, no water vapor values less than 6 ppmv are found below the -4 PVU isoline in ECMWF data (not shown here). This shows, at least for this situation, that the transport process of dry stratospheric air masses to these low altitudes we observed here is not captured by the ECMWF meteorological analysis (see Sect. 4).

Embedded in the dry vortex air masses, small filaments (marked in Fig. 4 with red boxes) of enhanced water vapor are visible. These filaments may indicate rehydration layers, where sedimented ice particles from upper layers have sublimated and resulted in layers of enhanced water vapor as described in Khaykin et al. (e.g. 2013) for the Arctic. However, trajectory reverse domain filling (RDF) (Sect. 4.2) showed that these air masses could not be attributed to rehydration. Also the in-situ measurements of FISH/HAI showed no evidence of a rehydration layer during the dive from ~ 12.5 to 3 km and back. We have analyzed these somewhat moister air masses in more detail, as discussed in the next section.

4 Trajectory based case study of 13 September 2012

Detailed investigation of the observed Antarctic dehydration is performed using air mass trajectories calculated with the tool CLaMS-traj which is part of the Chemical Lagrangian Model for the Stratosphere (CLaMS) (McKenna et al., 2002; Konopka et al., 2007). The trajectories are driven by the ERA-Interim reanalysis horizontal winds and diabatic heating rates (Dee et al., 2011; Ploeger et al., 2010). **To resolve transport processes in the troposphere influenced by the orography and transport processes in the stratosphere where adiabatic horizontal transport dominates, the hybrid σ - θ coordinate ζ is used (Mahowald et al., 2002) . Therefore in the stratosphere and in the UTLS, potential temperature θ is employed as the vertical coordinate of the model and the cross-isentropic velocity $\dot{\theta} = Q$ is deduced from the ERA-Interim forecast total diabatic heating rates Q , including the effects of all-sky radiative heating, latent heat release and diffusive heating as described by Ploeger et al. (2010) . In the tropospheric region defined by the condition $\sigma = 0.3$, the vertical model coordinate smoothly transforms into an orography-following $\sigma = p/p_s$ -coordinate (p - pressure, p_s - surface pressure), with the vertical velocity transforming into the corresponding $\dot{\sigma}$ (Pommrich et al., 2014) . Besides temperature, pressure, and humidity along the trajectories, the potential vorticity (PV) and thermal tropopause height (WMO, <2 K/km lapse rate criterion) are **derived** from the ECMWF meteorological fields. The temperatures from ECMWF along the flight path agree very well with the measured temperatures from the**

HALO aircraft (not shown here). The mean deviation on each flight level is mostly around 0.4 K, and the temperatures in ECMWF are slightly lower than the measured temperatures.

In the following we will use the trajectory analysis to investigate the time scale and location of potential dehydration events. In a second step we will analyze the trajectory data for possible exchange into the Antarctic troposphere.

4.1 History of dehydrated air masses along flight path

In this section, we investigate the history of the in-situ measured dehydrated air masses. For that purpose, the trajectories are calculated every 10 s along the flight path, each reaching 50 days backwards in time. For further analysis, we selected all trajectories between 12:20–12:29 UTC (in total 54), where the lowest water vapor mixing ratios were measured. However, the other trajectories with measured low water vapor mixing ratios between 2 and 3 ppmv show very similar history. Figure 5 shows the median saturation mixing ratio with respect to ice (orange line) and the frequency distribution (gray scale color-code) of every 1 h trajectory time step of all trajectories calculated by ClAMS along the 50 days back in time. The trajectories show the same history in saturation mixing ratio from 13 September until 14 August, where the **trajectory start to diverge**. This indicates that the trajectories have the same path with the same temperature history for at least 30 days. After this time, the trajectories split up into two branches, indicating a bifurcation point and associated mixing on 14 August. The trajectories within the branch with lower saturation mixing ratios of 2 to 3 ppmv below the median line stay together on the same path until the beginning of August. The trajectories before the 1 August have different saturation mixing ratios indicating the dispersion of air masses, which limits the interpretation of the trajectory calculations.

To show the presence of ice particles and subsequent dehydration, CALIPSO data (see Sect. 2.6) are used. A distance criterion is used to match the air mass trajectories with the CALIPSO data. Every trajectory at a specific time which coincides with the CALIPSO orbit inside a square of $\pm 5^\circ$ latitude and $\pm 5^\circ$ longitude is checked for the presence of ice in an altitude range of ± 0.2 km. If more than 50 % of all trajectories at one time coincide with

a CALIPSO ice observation, the corresponding time periods are marked with a blue bar (see Fig. 5).

Figure 5 illustrates the coincidence of low saturation mixing ratios along the trajectories with observed ice formation (CALIPSO) in the considered air masses. The trajectories pass several cold regions associated with low saturation mixing ratio where ice formation is initiated. If the ice cloud and supersaturation exist long enough (typically around 0.5–1 day, Nedoluha et al., 2002) to allow the ice particles to grow to sizes of several microns (diameter of 10 to 20 μm), sedimentation of these ice particles begins and causes irreversible dehydration. The sedimented ice particles with these sizes fall with velocities of around 2.5 km d^{-1} (see Müller and Peter, 1992) until warmer temperatures in lower regions causes them to evaporate. With each pass of a cold region, more and more water vapor is removed from the air mass. As stated by Nedoluha et al. (2002), air masses must undergo several periods with dehydration before reaching the 1 to 2 ppmv mixing ratio level, which is typically observed by satellites within the vortex in the Antarctic spring. This periodic behavior is confirmed by Fig. 5, where trajectories undergo four events with low saturation mixing ratios and simultaneous ice cloud observations by CALIPSO. The first and the second period (1 and 9 August) reveal rather high saturation mixing ratios of around 3 ppmv. In every later period, colder temperatures and lower saturation mixing ratios are necessary to initiate additional ice formation and dehydration. **This is the case during** the third and the fourth period on 16 and 29 August where CALIPSO observed ice show lower saturation mixing ratios of 1.5 and 1.2 ppmv, respectively. The last two events will have dried out the air masses to the final state of about 1.6 ppmv, which is observed by the FISH and HAI instruments. After this strong dehydration, the temperatures became warmer in the late Antarctic spring and summer and no additional dehydration took place according to the simulations.

The recurring low saturation mixing ratios and corresponding low temperatures are primarily caused by gravity waves, which are induced by the high topography of the Antarctic continent (Stohl and Sodemann, 2010). Especially, the first three ice formation events occurred just as the trajectories passed the Antarctic Peninsula (northernmost part), which is a strong source of gravity waves (e.g. Ern et al., 2011; Hoffmann et al., 2013). Only the last

freezing event with the lowest supersaturation occurs above the Antarctic Plateau (central part) in the north-east of the continent.

After freezing, the ice particles can even reach the troposphere due to sedimentation, especially if dehydration takes place at low altitudes (11.5–12.5 km) as **revealed by the trajectories**. The lack of rehydration signatures in the GLORIA measurements (see Fig. 4) also indicates the permanent removal of water vapor from the stratosphere, as will be further discussed in Sect. 4.2. The air masses from which the ice particles originate and the air masses where they **potentially sublimate in the troposphere** behave very differently in dynamical perspective (**not shown**). Due to the length of time since the last dehydration on 28 August (16 days before measurement) it is very unlikely that rehydrated air masses would still be found below the observed dehydrated air masses in the troposphere. In addition, the in-situ measurements taken during the dive confirm the absence of any observed rehydration features.

4.2 History of air masses observed by GLORIA

The trajectories for analyzing the transport of dehydrated air masses observed by GLORIA are calculated **70** days backward in time from each tangent point (in total 30 000 single trajectories). In addition, the trajectories are also calculated **14** days forward to show where the air masses are further transported. With this set of trajectories, the GLORIA measurements can be interpreted and the associated transport pathways can be investigated.

In the first subsection, the dehydrated air masses observed with the in-situ instruments and GLORIA in the lowermost stratosphere are analyzed and the origin of moist filaments is investigated. The second subsection contains the trajectory analysis of dry air masses found to be below the thermal tropopause and investigate the further transport.

4.2.1 Vortex air masses

From the in-situ ozone and methane measurements, it is visible that the HALO aircraft has penetrated the Antarctic vortex (see Sect. 3.2). The identification of vortex air masses in the

GLORIA trace gas cross-sections is performed using the tangent point trajectories. Figure 6 shows the amount of time that the trajectory for each GLORIA tangent point stayed within the vortex before the measurement took place. To calculate this so-called vortex indicator, the vortex border needs to be determined at each trajectory time step with the Nash criterion (Nash et al., 1996). Then, the time that each individual trajectory remains within the vortex border is counted backwards from the time of measurement. This procedure is limited to the preceding 50 days. The blue color in Fig. 6 shows the trajectories that stay in the vortex for at least the last 50 days. In contrast, air masses that were never in the vortex are indicated by a deep red color. This includes tropospheric air, but also stratospheric air masses north of 48° S. Interestingly, **strong gradients between former vortex air (blue) and extra-vortex air (red) are evident** between 09:00 and 11:00 UTC and also between 11:45 and 12:20 UTC, which exactly follows the measured thermal tropopause (black dots). Here, the **lower vortex boundary defined by (Nash et al., 1996) along the trajectories coincide well with the WMO thermal tropopause.**

In contrast, the air masses that spent between 5 and 40 days in the vortex (marked by light blue and reddish colors) indicate mixing of **mid latitude** air into the vortex. Even at the vortex edge and in the core of the vortex itself, small filaments are apparent. As indicated in Sect. 3.2, some filaments with elevated amounts of water vapor were also observed with GLORIA. Two of these filaments observed around 13:00 UTC between 10 to 12 km altitude can be found with this Nash PV criterion showing vortex residence times of 5 days (filament 6) and 35 days (filament 7), respectively. This shows that these filaments are likely freshly mixed in from outside the vortex and thus could potentially contain higher water vapor content compared to the dehydrated vortex air and therefore weaken the hypothesis of rehydration signatures.

To show that all observed filaments (red boxes in Fig. 4) are caused by mixing, reverse domain filling (RDF) is applied to all trajectories, as already stated in Sect. 3.2 (similar as in Beuermann et al., 2002). For the RDF calculation, only advection and no mixing along the trajectories is assumed. In each time step of the RDF method, the ECMWF water vapor field

is interpolated onto the locations of the trajectories and then projected onto the observed time-altitude grid of the GLORIA observations.

Figure 7 shows ECMWF water vapor mixing ratios five days prior to the measurement time as projected by the trajectories. The red boxes highlight the position of the water vapor filaments observed by GLORIA (see Fig. 4). With some small deviations in time and altitude, almost all of the observed filaments in the vortex can be reproduced by the RDF method. Note, that the RDF method cannot produce the effect of rehydration. This indicates that the water vapor filaments observed by GLORIA were not generated by rehydration, but are likely the signatures of moister **mid-latitude** air masses that were mixed into the vortex during the preceding 5 days. Continuing the RDF method further back in time, the filament 7 (≈ 12 km) at 13:00 UTC emerges 35 days before observation and indicates the origin to be another mixing event. The time of in-mixing of the filament 7 was already visible in our analysis in Fig. 6. Thus, it seems that these filaments originate from several mixing events, not from a single mixing event only. Furthermore, very dry air masses observed in the time between 09:00 to 10:30 and 12:00 to 13:30 UTC down to 7 km (see Sect. 3.2) can also be partly reproduced with the RDF method, indicating that transport is the main reason for the occurrence of dry air below the thermal tropopause. The origin of these air masses and the transport mechanism are analyzed and discussed in the next subsection.

4.2.2 Transport of dehydrated air masses across the **thermal tropopause**

The weak vertical temperature gradient **and the poor definition of** the thermal tropopause **is** implied by the tropopause heights derived from GLORIA, which have a broad scatter (± 0.5 km) at some places. In addition, the PV gradient **along isentropes (i.e. isentropes following the PV isolines) as well as the vertical gradient** at the transition between troposphere and stratosphere within the vortex (08:30–13:30 UTC) is small. The PV isolines (-4 and -2) in the vortex have a larger spacing than at the edge (jet region), where the vertical distance between the isolines decreases (08:15 and 13:30 UTC in Fig. 6). As a consequence, **the weak isentropic PV gradient along the sloping isentropes facilitate a transport to lower altitudes below the location of the thermal tropopause. Subsequent diabatic processes may**

lead to STE and irreversible exchange with tropospheric air. Especially in the case of planetary wave breaking events that produce stratospheric intrusions, large amounts of air can be transported into the troposphere **across the tropopause**. Figure 8 shows the time difference between **the trajectory initialization time and the first thermal tropopause crossing**. This time is determined by counting the trajectory time steps backward in time to wherever the difference in altitude between current tropopause and trajectory is positive. Air parcels having stayed for the last 70 days **below the thermal tropopause** (dark red colors) can clearly be assumed to be tropospheric nature. **In addition, low PV values (indicator for tropospheric air masses) can be attributed to these air masses as discernible from Figure 2 b.**

The stratospheric and vortex air masses did not cross the **thermal** tropopause (dark blue colors). The air masses below the measured **thermal** tropopause are obviously more patchy and contain more trajectories originating from the troposphere (reddish colors). However, some freshly **transported** stratospheric air masses are also discernible below the thermal tropopause (light blue colors), down to 7 km. The green framed areas in Fig. 8 mark all the trajectories having tropopause crossing times between 0 and 70 days and where GLORIA observations reveal **a** water vapor content of less than 3 ppmv. Indeed, there is a very good correlation between the light blue air masses, which crossed the **thermal** tropopause between 2 and 6 days before, and the green framed areas marking the dehydrated air masses. In addition to the low water vapor content in the green framed areas, higher values of ozone were also observed by GLORIA, which **corroborates** the stratospheric origin of these air masses (not shown here).

Figure 9 shows the median **altitude (a), PV (b), latitude (c) and equivalent latitude (d)** distributions of all trajectories within the green framed areas from Fig. 8 for the period of **2 weeks preceding and the 2 weeks following** the observation. **In addition, two RWB events are marked where subsets** of air masses

(green and red line) are detached and separated from the rest (blue line). The strong Rossby wave activity is made visible by comparing the median latitude and equivalent latitude just a few days before the observation (8 to 17 September) as seen in Fig. 9c and d. The latitude is oscillating from near **70** down to 80° S, up to **50**° S, and then back to **70**° S,

while the equivalent latitude in this time frame stays constant at around 70° S. This implies no PV change and only an isentropic latitudinal displacement due to planetary Rossby waves.

On 10 September, on the point furthest south at latitudes around 80° S (see Fig. 9c), the back trajectories split up into two different branches, indicating a bifurcation point caused by partly breaking of the Rossby wave. Later, at the point of most northward displacement (around 50° S) a second RWB event is identifiable; thenceforward parts of the air masses are separated from the rest. The air masses associated with the first RWB event (green line) stay at a similar altitude and PV level for at least up to the 14 September. At this day the PV starts to increase further to more tropospheric values, while the altitude starts later to decrease from the 17 September on rapidly within 4 days to an altitude of less than 3 km with a simultaneous decrease in latitude to 30° S. The previous PV change before the altitude change indicate the point in time, where air masses detach the lowermost stratosphere and become more tropospheric. The trajectories associated with the second RWB event (red line) show a similar behavior but shifted by about 3 days. So it is not possible to match one specific event to the transport of all air masses which were observed below the thermal tropopause with low water vapor content (green framed areas in Fig. 8). This implies that one Rossby wave event is not the sole reason for the downward transport. Parts of the air masses (950 trajectories) within the stratospheric intrusion did not leave the lowermost stratosphere within the two weeks. Indeed, if one looks into the following weeks, additional Rossby wave and RWB events occurred which brings more parts of the remaining air masses (950 trajectories) into the troposphere.

Thus, Rossby wave breaking events drive the mixing process together with isentropic transport through the thermal tropopause. Additionally, RWB events detach air masses from the PV structure and can facilitate a wave-driven, secondary circulation that transports warmer air masses from around the vortex edge to the colder inner vortex. The change in PV is most likely not caused by cloud processes due to the dryness of these air masses. However, despite the dryness, the slight enhancement of ozone and the residual water vapor enable radiative cooling and a reduction in potential temper-

ature i.e. descent of air masses. Also the potential general subsidence in the Antarctic troposphere caused by the katabatic surface winds could lead to adiabatic heating of these air masses (Roscoe, 2004) . Once transported into the troposphere, air masses will be transported down to near-surface level within few days as suggested already by Stohl and Sodemann (2010) .

The vortex becomes more and more unstable during the Antarctic spring and one may assume that this **mixing** event is already a first sign of the vortex break-up. Haigh and Roscoe (2009) reported about recent Antarctic final warming events occurring between the beginning of December to late January at the 100 hPa level. Therefore, the observed in-mixing of vortex air masses occurs long (at least two months) before the final warming and break-up of the vortex. It is therefore independent of the final warming event **to a certain extent** and seems to be a **frequently occurring** process in the Antarctic winter and spring.

Finally, the reason for the large downward transport is a mixture of strong Rossby wave activity, the presence of a **permeable** thermal tropopause, **RWB events** and the additional radiative cooling with corresponding subsidence of the air masses.

5 Conclusions

Detailed observations of dehydration and **STE** in the Antarctic UT/LS are very rare. In this study, **we used** high resolution in-situ **and remote sensing measurements aboard the HALO aircraft in combination with an extensive trajectory case study to answer the following questions:**

1. **Can we confirm dehydration in the lowermost stratosphere and at the bottom of the Antarctic vortex ? We observed** strongly dehydrated air in the lower vortex **at potential temperatures of 350 to 360 K (\approx 12-13 km)** with the hygrometers FISH and HAI. The final dehydration can be traced back to several ice formation events **occurring at 350 to 370 K, in a time frame of several weeks**, where water is sedimented out of the observed air masses. **Finally, dehydration occurs directly in the lower vortex and in the lowermost stratosphere below the vortex.**

2. **How deep do dehydrated air masses from the vortex descend into the troposphere ?** Dehydrated air masses are also observed with the high resolution remote sensing instrument GLORIA down to the thermal tropopause, as well as below this point, down to 7 km in altitude. **We confirm that dehydrated air masses with water vapor mixing ratio <3 ppmv cover large regions of the Antarctic lowermost stratosphere down to the -2 PVU isoline and thus dominate at least for this observed case the transition region between the upper troposphere and lower stratosphere.** But besides that, small filaments of enhanced water vapor are observed by GLORIA in the vortex air **and the lowermost stratosphere at 330 to 350 K (≈ 10 -12.5 km).** These filaments can be clearly assigned to in-mixing of moister air masses into the dry vortex at several points in time, indicating that this transport process occurs frequently.
3. **Where and how have Antarctic dehydrated stratospheric air masses been transported ?** With an extensive trajectory case study, dry air masses below the thermal tropopause **in the south hemisphere polar and mid-latitude region** are determined to have stratospheric origin. **The enhanced permeability of the Antarctic thermal tropopause in winter and spring allowing isentropic downward transport is confirmed (e.g. Zangl and Hoinka , 2001; Roscoe, 2004; Evtushevsky et al., 2011) .** It is shown that **parts of these dry air masses are transported further down to near surface levels within few days. This transport into the troposphere is caused by several Rossby wave breaking events, where they detach air masses from the PV structure.** Once the air masses have been transported into the **Antarctic troposphere, most likely** radiative cooling causes their subsidence to near surface levels.

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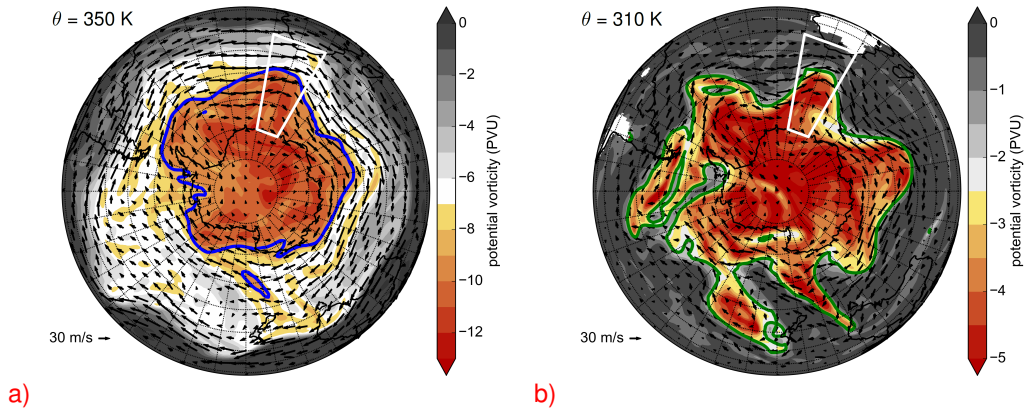


Figure 1. Dynamical situation with color-coded PV from ECMWF data on 13 September 2012 at 12 UTC on two different potential temperature levels: a) 350 K b) 310 K. The black arrows depict the wind vectors. The blue line represents the vortex edge at 350 K (Nash et al., 1996) and the green line the -2 PVU isoline at the 310 K level. The region of the ESMVal flight is marked by a white box.

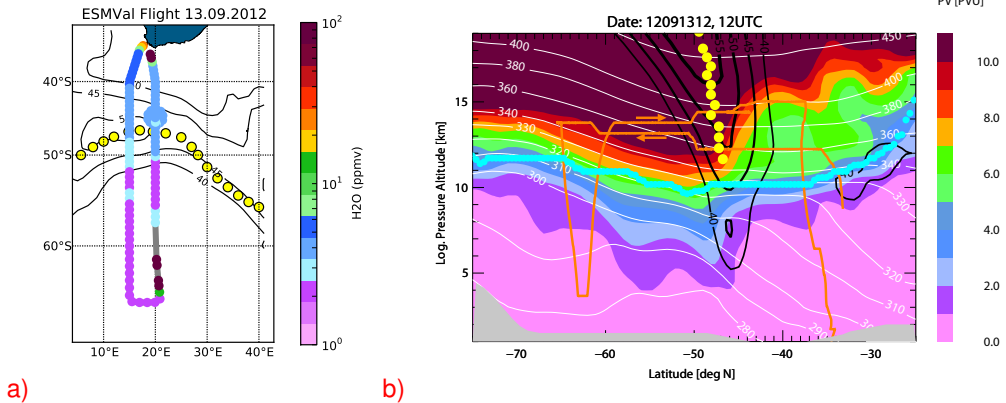


Figure 2. Flight pattern of the ESMVal flight on 13 September 2012. The black contours illustrate the horizontal westerly wind from ECMWF data. Yellow dots represent the vortex edge derived from the Nash criterion (Nash et al., 1996) based on ECMWF data. **(a)** Horizontal map of water vapor mixing ratios measured by FISH (5 min averaged data) is color coded on the flight path (gray color indicates data gaps due to the dive). **(b)** Meridional cross-section of PV along the flight path (orange line) calculated from ECMWF data. Cyan dots represent the ECMWF thermal tropopause.

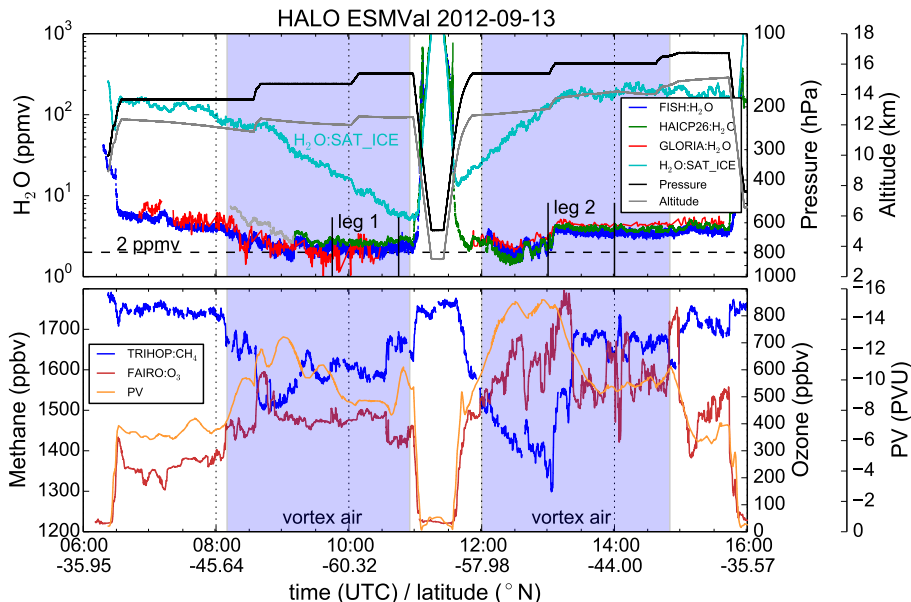


Figure 3. Timeseries of ESMVal Antarctica flight on 13 September 2012: the upper panel shows FISH (blue), HAI (green) and GLORIA (red) water vapor measurements, water vapor saturation mixing ratio with respect to ice (cyan, derived from HALO temperature), pressure (black), and altitude (gray) at flight level. The lower panel shows the tracer observations ozone (red), and methane (blue) and interpolated ECMWF PV values (orange). Time ranges within the vortex are marked with bluish shadows according to the tracer measurements. Time ranges leg 1 and leg 2 are marked for hygrometer intercomparison of FISH and HAI (see text).

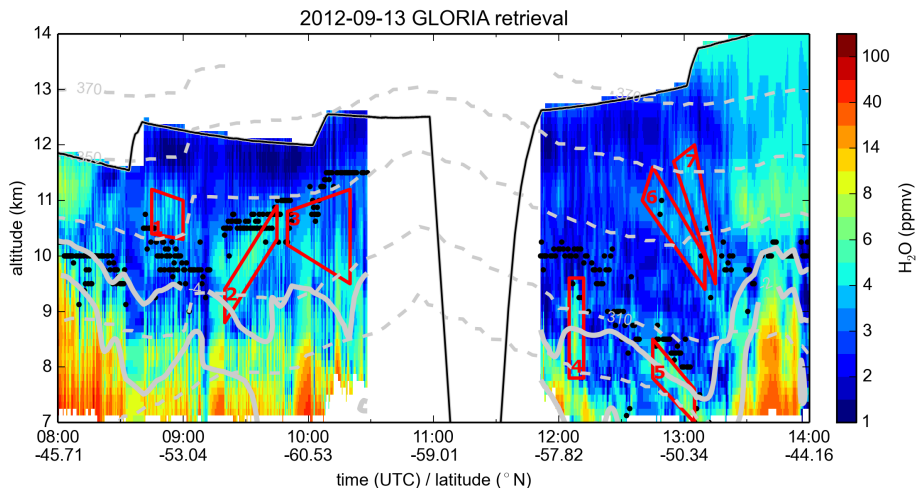


Figure 4. GLORIA time series of the H_2O retrieval product (13 September 2012). The solid black line and the black dots represent the flight path and thermal tropopause derived from the GLORIA temperature measurements, respectively. The dashed gray lines and the solid lines marks isolines of potential temperature (290–370 K, $\Delta\theta = 20$ K) and PV (-2 and -4 PVU), respectively. The red boxes marks filaments (1–7) with slightly enhanced water vapor mixing ratios in or below the vortex.

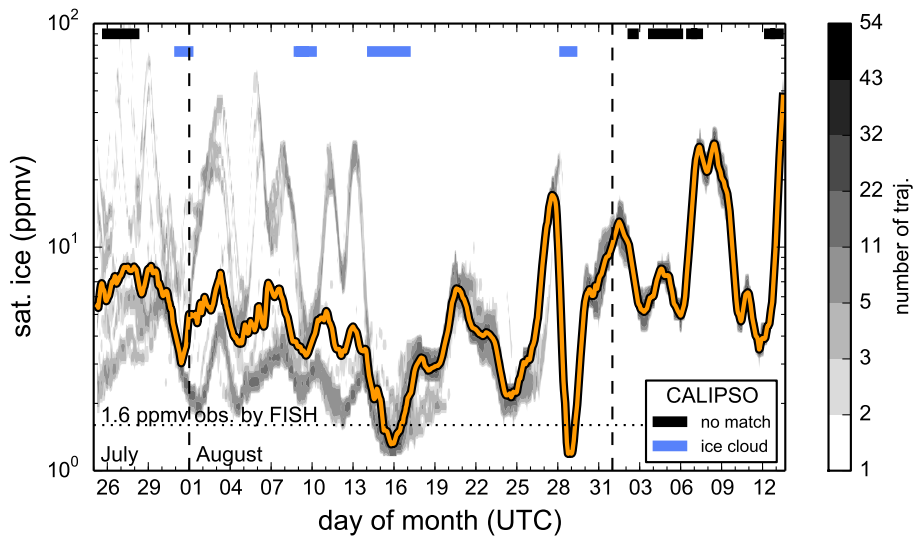


Figure 5. Calculated frequency distribution (number, gray shadow) and median (orange line) of ice saturation mixing ratio along trajectories 50 days back in time starting at the flight path on 13 September 2012 (12:20 to 12:29 UTC, 54 trajectories in total). If more than 50 % of trajectories have a corresponding CALIPSO observations of ice or no ice, the time is marked with blue or white, respectively. Otherwise, it is marked in black, indicating no CALIPSO match.

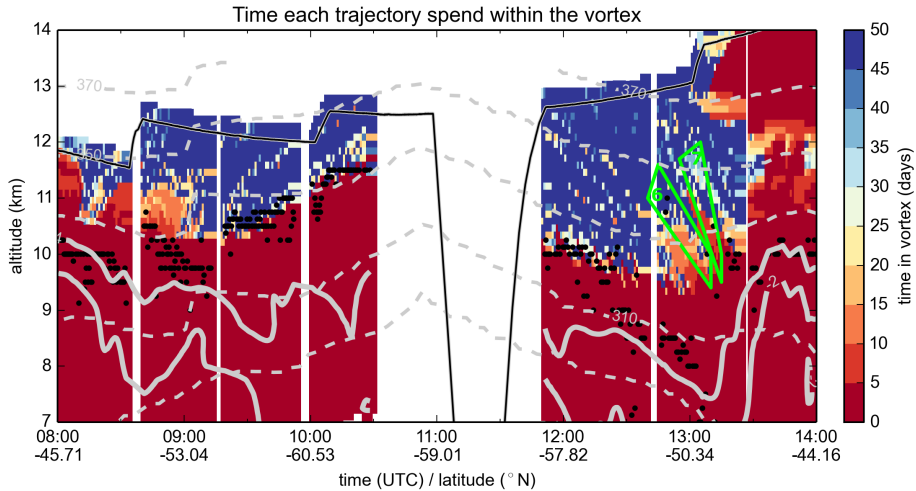


Figure 6. The time each GLORIA tangent point trajectory has spent in the vortex before the measurement. The green boxes mark two of the filaments (6, 7) observed by GLORIA (see Fig. 4). The flight path, tropopause height, potential temperature, and PV isolines are the same as in Fig. 4.

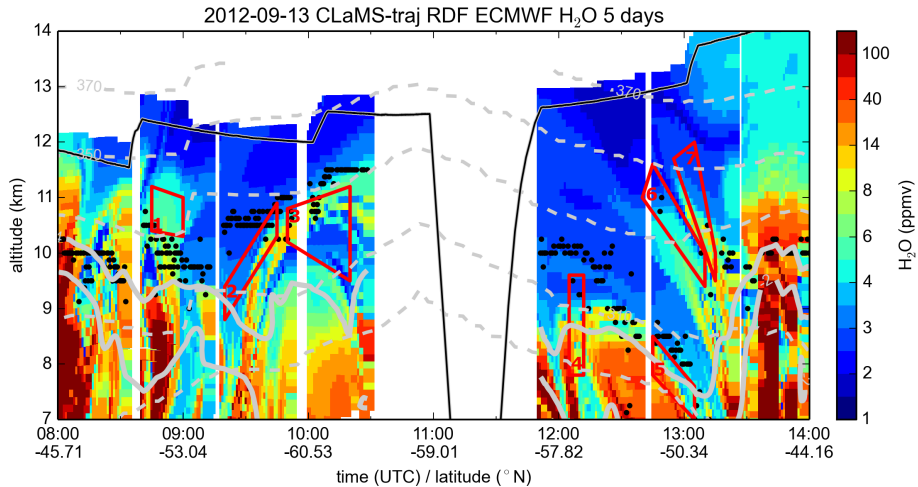


Figure 7. RDF (Reverse Domain Filling) of each GLORIA tangent point trajectory with ECMWF water vapor of 5 days prior the measurement time. Red boxes show the location of the observed filaments. The flight path, tropopause height, potential temperature, and PV isolines are the same as in Fig. 4.

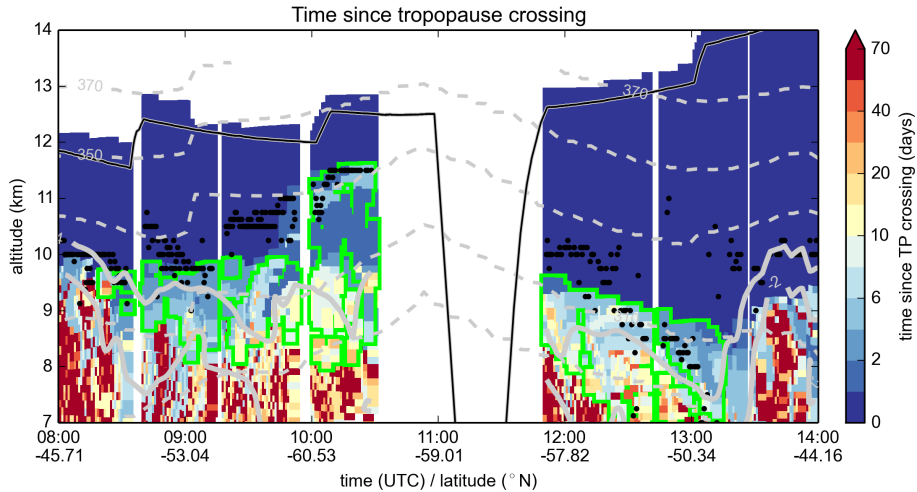


Figure 8. The color code denotes the time since trajectories passed the thermal tropopause from stratosphere to troposphere. Bluish colors denote trajectories which recently passed the thermal tropopause, reddish colors indicate tropopause crossings longer ago. The green framed areas mark the trajectories where GLORIA observations show water vapor below 3 ppmv. The flight path, tropopause height, potential temperature, and PV isolines are the same as in Fig. 4.

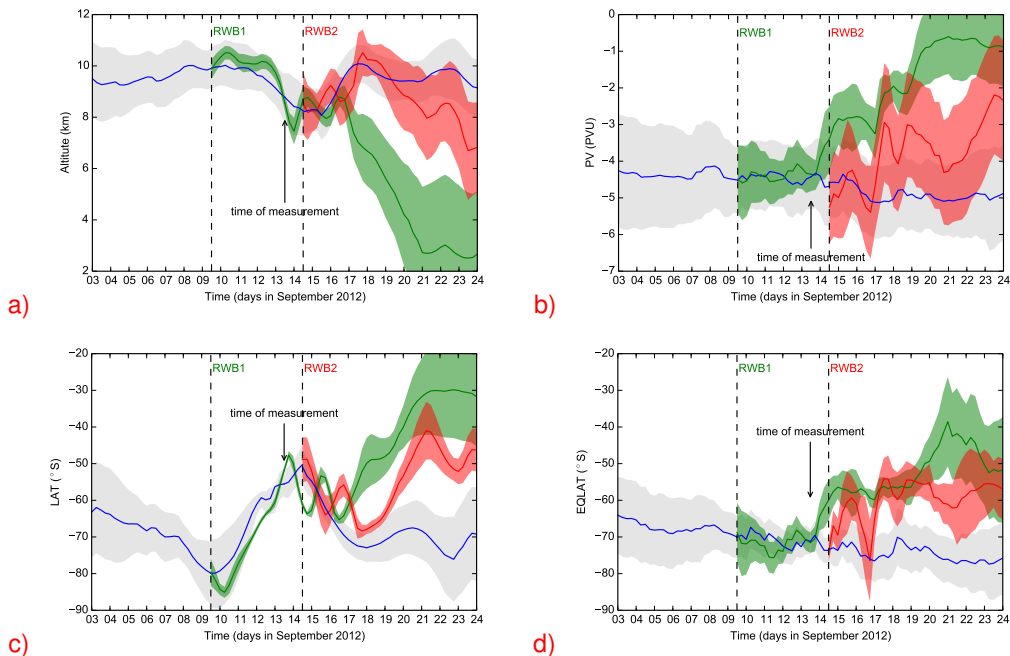


Figure 9. Air parcel properties based on trajectory analyses for all green framed trajectories from Fig. 8 for the period of 2 weeks before to 2 weeks after the measurement. On the 3 of September the median of all 1400 trajectories is represented by the blue line with the respective standard deviation as gray shadow. At RWB event 1 a subset of 250 trajectories (green line, median) is separated from the rest (from that moment on the blue line represents 1150 traj.). At RWB event 2 a subset of 200 trajectories (red line, median) is separated from the rest (from that moment on the blue line represents 950 traj.): (a) median of potential temperature (θ), (b) median altitude above mean sea level, (c) median latitude, (d) median equivalent latitude. The gray, green and red shaded areas marks the respective SD of the trajectories.