

The authors have provided a detailed account of ice particle residue (IPR), hydrometeor, and aerosol measurements at the Jungfraujoch station during the winter of 2013 using single-particle measurements and a series of different inlets. Results show the differentiation in IPR type dependent on cloud and meteorological conditions, and aerosol source, in comparison to the background aerosol composition. The authors found that IPR were generally enriched in organic material at higher temperatures, and partly internally mixed with black carbon. Minerals and lead were also observed to be dominant in the IPR, but at lower temperatures. Although these results provide another useful reference for ice nucleation measurements in the literature, the manuscript suffers from a few issues discussed below. After these issues are addressed, I believe this manuscript is suitable for publication in ACP.

Changes in the manuscript are highlighted as follows:

Referee#1 → yellow

Referee#2 → green

Referee#3 → red

Additional changes → blue

General concerns:

The new findings are not emphasized or discussed enough. The authors demonstrated that mineral dust serves as ice nuclei (IN), however the finding that OC was found on most of the IPR is interesting and a somewhat novel finding. Could it be that the organic material is biological in origin? This is hinted at in the discussion, but should be emphasized.

It could be that a part of the organic material is of biological origin. Most of the organic material shows C_n -Fragmentation ($n = 1-3$) and the fragmentation of aliphatic hydrocarbons ($C_nH_{2n\pm 1}$). In some minor cases we found oxidized organic carbon and nitrogen-containing carbon, which may indicate biogenic origin. However, due to the lack of anion spectra we cannot unambiguously classify those particles as biological particles. It is also possible and more probable (winter time and long range transport), that this organic material results from a coating by secondary organic. Therefore, all particles with an organic pattern are summarized to the OC cluster and are not classified in more detail.

Could they be organic particles that contributed to riming versus ice nucleation? This could most certainly happen during the residue drying process and should be noted.

It is expected that ice particles with a size range of 20 μm and smaller grow only by water vapor diffusion and there is a negligible probability that they scavenge aerosol particles by impaction and riming. Thus, the released residual particles by the Ice-CVI can be interpreted as the original ice nuclei. This has been discussed already in the Ice-CVI method paper (Mertes et al., 2007) and is supported by the references cited therein.

Discussing these IPR in more detail in addition to focusing on their sources using HYSPLIT analysis would be beneficial and enable the uniqueness of the results to stand out.

The general region of air mass origin is discussed, however, there seems to be a dearth of discussion on the potential sources of the IPR. What types of possible sources are found in each of the different transport regions, such as Figs 3 and 6.

During the cloud events the air masses arrived at the measurement platform from western direction, from North America over the Atlantic Ocean and France. Therefore, particle composition is influenced by different sources (marine, biological and anthropogenic sources) and aging processes. However, also local emissions like residential heating emissions, diesel/fuel exhaust, cooking emissions, etc., cannot be ruled out because of the large skiing area that is situated in the Swiss Plateau. It is therefore hard to say which source influenced which event mostly because all events have more or less the same air mass origin. Because of the low counting statistics during the cloud events it is not possible to further analyze the chemical composition of the ice particle residuals for their origin.

Page 16 line 27-30

The abstract is essentially a list of results. The authors could revise to provide more context as to why their observations and measurements are important on a broader scale. For instance, the abstract could be started with a few sentences on the motivation for the study. The remainder of the sentences are predominantly composed of results, but can the authors provide a few sentences on why the findings are important? What do they mean in general? Closing with a more general statement on their importance would help as well.

We extended the abstract according to the reviewer's suggestions. Page 1 line 24-27.

Please define acronyms before using them. Once, defined no need to redefine (i.e., JFJ is defined twice). Also, please define the acronym for the field campaign (i.e., CLACE).

The acronym of CLACE was defined. Page 1 line 28.

Where interstitial aerosols investigated? It has been shown that these types can vary quite drastically from ice residues and clear aerosol. It would be interesting to comment on this if they were indeed measured. If not, please clarify.

No, interstitial aerosol was not investigated. In section 2.1 it is described which types of aerosol with which inlet was measured.

The contribution of residues from rime droplets to the full ice crystals is not discussed. Surely, some if not most of the ice crystals analyzed have been rimed, which is an efficient process in a mixed-phase cloud system. The authors need to address the contribution of rimed species to the original ice nucleus residue.

As stated above and in Mertes et al. (2007), small ice crystals grow by water vapor diffusion until they reach a size of about 50 μm . Thus, the ice crystals sampled by both Ice-CVI and ISI (upper cut-off: 20 μm) have not been subject to riming (see also Kupiszewski et al. (2014)).

Section 2.1 Page 4 line 31 – Page 5 line 2.

Specific comments:

Page 1 line 26: 200-900 nm is referring to IPR after drying, correct? Please confirm in text.

Yes, the size range 200 – 900 nm is referring to IPR after drying. But particularly, this size range holds for all particles measured with the ALABAMA. The sampling efficiency of the ALABAMA is the best in this size range. For better understanding we rearrange this sentence.

Page 1 line 29 – 32

Page 1, line 28: Out of curiosity, how did the clear air temperature compare?

The temperature mentioned in the abstract is the temperature range over the whole campagne. During non-cloud episodes the temperature range was between -6.7 °C and -27.3 °C. There are no differences regarding the temperatures between cloud and non-cloud events.

Page 16 line 10 – 12.

Page 1, line 29: State that “background aerosol” refers to cloud free air aerosol. Were interstitial aerosol measured? Also, reiterate that both composition and size were determined for background aerosol.

During this campaign no interstitial aerosol was measured. We modified the abstract accordingly.

Page 1 line 29 – 32

Page 2, lines 6-9: Instead of providing ranges for some species and averages for other, be consistent with provide either an average with a standard deviation or range for each type.

We provide ranges the IPR to express the range of the measured values that were obtained using the ISI and the Ice-CVI (Figure 8). The value for the background aerosol refers to the whole background aerosol data set as given in Figure 8.

Page 2, lines 17-18: Saying aerosols affect solar and thermal radiation is vague. Perhaps use the term radiation budget, or explain these a little more.

We changed the text accordingly.

Page 2 line 21 – 22.

Page 2, line 28: It would be benefit the authors to explain the Wegener-Bergeron-Findeisen process here, and that the riming of ice crystals from supercooled droplets is what causes enhanced precipitation.

An explanation of the Wegener-Bergeron-Findeisen process was added.

Page 2 line 33 – Page 3 line 7.

Page 3, line 4: What do the authors mean by important? Technically, biological particles are the most efficient, but dust is the most abundant. Please clarify.

It was clarified in the manuscript. Because of its abundance mineral dust is the most important ice nuclei but biological particles the most efficient ice nuclei (Hoose et al., 2010).

Page 3 line 13 – 23.

Page 4, line 16: What temperature was the inlet heated at? Wouldn't there be an issue with low molecular weight organics or nitrate/sulfate evaporating from the residues in the inlet?

The temperature of the total inlet is regulated to 20 °C (Weingartner et al., 1999). Sulfate is unlike to evaporate, but evaporation of nitrate is possible. However, due to technical problems we could not measure anions, therefore we did not detect sulfate and nitrate. It is also possible that volatile organics evaporate from the residues. However, since our analysis is not quantitative in terms of particle mass, and a lot of organic material is detected in the spectra, this effect will only play a minor role, if any.

The temperature was added. Page 4 line 28.

Page 5, line 20: Is this referring to the inlet system of the ALABAMA or the OPC?

The size range from 0.25 to 52 µm refers to the size channel boundaries recorded by the OPC. We rearranged this sentence to make it more clearly.

Page 6 line 3 – 5.

Page 5, lines 17-27: The size ranges defined do not correspond to that of the abstract. Why is only 200-900 nm presented? How can 200 nm be achieved if the lens only measures down to 250 nm? What are the two values for transmission efficiency for?

The ALABAMA covers a size diameter range from approx. 100 – 3000 nm (with best transmission from 200 to 900 nm) and the Sky-OPC from 250 – 52000 nm. This section was rearranged for clarification.

Page 5/6 line 33 – 5 and Page 6 line 11 – 13.

Page 6, line 2: Why are only supercooled droplets >5 µm removed?

The traditional CVI downstream of the droplet pre-impactor segregates particles smaller than 5 µm in diameter. Thus, the supercooled droplets < 5 µm are removed from the sample flow by the CVI, so that it is sufficient to use a size cut of > 5 µm for the droplet pre-impactor to exclude all supercooled droplets. This is not explicitly mentioned in the indicated text passage and is now reworded to make this point more clear.

Page 6 line 24 – 32.

Page 6, line 4: How do the ice crystals not stick to the surface of the plates? Does the ice from the supercooled droplets freeze fast enough? Is there any loss of ice crystals to the chilled plate?

The idea of such a droplet pre-impaction for the Ice-CVI design was mainly inspired by the work of Tenberken-Pötzsch et al. (2000). They showed using a cooled impactor with several stages that the drop and ice particle residuals are found on the first and the last stages, respectively. Mertes et al. (2007) did lab tests that showed a 100 % sticking efficiency for supercooled drops on the cooled impaction stages and no significant changes within the

measurement uncertainty in the ice particle concentration before and behind the impactor. The latter is the consequence of low velocities that avoid ice particle shattering (which would otherwise lead to an increase of ice particles) and a readily bounce off of ice particles from ungreased surfaces (which would otherwise lead to a decrease of ice particles), which was also observed by Murphy et al. (2004). Comparison with directly measured ice particle and ice particle residue concentration indicate at least that both are in the same order of magnitude. More detailed information is now given in the respective text passage.

Page 6 line 24 – 32.

Page 7, line 17: Rename section to “Additional instruments” or “Supplement instruments”

The headline was changed into “Supplement instruments”.

Page 7, line 21: Why was 20 chosen as the threshold, and is this unit correct? Shouldn't it be something along the lines of g m^{-3} ?

The unit is mg m^{-3} . The typo was corrected.

Page 8 line 21.

Page 7, lines 30 and on: No need to introduce SEM if there are no results discussed in the current manuscript.

Results of the ESEM are presented in Figure 10. The size of the collected IPR on impactors is measured by ESEM and compared to the sizes of the IPR analyzed by ALABAMA.

Page 8, line 23: What was the percentage of residues that were classified out of all residues measured?

All measured ice residues were classified. The spectra which could not be classified as one of the seven particle types were sorted into the “others” cluster.

Page 9, lines 22 and on: The BioMinSal type presented here is very similar to the residues observed by Creamean et al. (2013). This is a great finding in that this type has been observed elsewhere. The authors should reference Creamean et al. (2013) and perhaps provide a brief discussion on their similarities.

Creamean et al. (2013) report the finding of dust particles mixed with biological particles. However, the Bio/Min/Sal cluster includes spectra, which show only metal ions. The source of these metals is unclear and can be attributed to dust or biological or salt particles (see explanation in section 2.6). The particles in the cluster are not necessarily (ad even unlikely) a mixture of all three compounds. To make this statement more clear we decided to rename the cluster to Bio/Min/Sal.

Page 10, line 14: Make the title more proactively descriptive, i.e., “Ice particle residue composition and potential aerosol sources”.

The section title was changed into “Ice particle residual composition and potential aerosol sources”.

Page 10, lines 22-25: The event classification is somewhat vague. More information is needed to differentiate these events. What were the temperature ranges of each? What was the wind direction? Etc?

The Ice-CVI events were classified according to the sample efficiency and properties of the Ice-CVI. Due to closer consideration of the (++)- and (+-)-events and taking into account that the ALABAMA did not measure the whole cloud event at the Ice-CVI we get a new classification with more (++)-events which are now called (++)/(+-)-sampling-periods (Table 1). The average temperature, wind direction and wind speed of each event were added.

Page 12 line 16 – 26.

Page 11, line 3: So, the ISI measures IN and CCN, correct? This is what I gathered from the methods. How do the authors eliminate contribution from CCN/droplets? This should be clarified.

No, the ISI measures ice particle residuals (IPR) obtained after the ice crystals sampled through the inlet are evaporated – it does not measure CCN. Droplets are evaporated within the inlet itself in the droplet evaporation unit, a chamber with ice-coated inner walls. Due to the differences in saturation vapour pressures over ice and water, liquid droplets evaporate in the chamber, while ice crystals are affected to a much lesser degree. The CCN released in this process are subsequently removed together with the interstitial aerosols using a pumped counterflow virtual impactor (PCVI), which separates particles based on their inertia, with only larger (aerodynamic diameter > ca. 5 μm) particles (i.e. the ice crystals) transmitted through the PCVI. As the ISI is composed of a number of components, we refer the reader to Kupiszewski et al. (2014), where the operating principle of the inlet and all the components are described in detail.

Page 11, line 18: What were the initialization parameters for HYSPLIT (starting altitude, times initialized, hours back, etc)? These should be provided in the text.

The initialization parameters were added in the text. The starting altitude was 3580 m asl; initialization time: end of the current event; hours back: 72 h.

Page 13/14 line 30 – 1.

Page 11, lines 18-31: It would be helpful to provide the average temperature or wind direction for the events Fig 3 (or any figures with pie chart comparisons for that matter).

We added average temperature and wind direction and speed of each event in Table 1.

Page 12, line 8: Were the lead artifacts a large percentage of the particles (provide the percentage)?

We assume that particles originating from contamination of mechanical re-suspension of lead containing particles from the surface of the impaction plate produce large homogeneous lead-rich particles. This is explained in detail in the paper by Worringer et al. (2015). Our measurements do not show any spectra containing only lead. Therefore, we assume that the sampled lead containing particles are no artifacts.

Page 14 line 18 – 25.

Page 12, line 15: Why were these events chosen? A sentence or two describing the reasoning for showing these warranted.

The six Ice-CVI-sampling-periods were chosen because of the (++) classification (no secondary ice or snowflakes sampled) and the need for sufficient particle mass spectra.

The ISI event was chosen because only during this event a sufficient number of mass spectra was recorded. This is now explained in the text.

Page 14 line 28 – 31.

Page 13, line 19: Clarify that these are referring to background aerosol measurements.

The CPC measurements were done simultaneously during cloud conditions at the total-inlet. This is now stated clearly in the text.

Page 16 line 2 – 4.

Page 13: More discussion is needed on why the events are so different. Is it due to temperature, wind direction, air mass source, etc.

In section 3.2 we already discussed in detail the possible impact of temperature and air mass origin (source and differences in altitude) on the chemical composition of these seven different events. Further discussion would result in pure speculation and does not seem appropriate here.

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In-situ single submicron particle composition analysis of ice residuals from mountain-top mixed-phase clouds in Central Europe

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Abstract

Ice nucleating particles have a high impact on ice crystal formation in clouds and thereby on precipitation and the global radiation budget. Because the chemical composition of these particles plays an important role at the nucleating processes, investigation of the chemical composition of ice nucleating particles is a topic of current research. This paper presents results from the “INUIT/CLACE 2013” (Cloud and Aerosol Characterization Experiment) field campaign at the high alpine research station Jungfraujoch in January/February 2013. The chemical composition of ice particle residuals (IPR) was measured in orographic, convective and non-convective clouds. The submicron particles were analyzed by a single particle mass

1 spectrometer (ALABAMA) under ambient conditions characterized by temperatures between
2 $-28\text{ }^{\circ}\text{C}$ and $-4\text{ }^{\circ}\text{C}$ and wind speed from 0.1 m s^{-1} to 21 m s^{-1} . Additionally, background aerosol
3 particles in cloud free air were investigated. The IPR were sampled from mixed-phase clouds
4 with two inlets which selectively extract small ice crystals in-cloud, namely the Counterflow
5 Virtual Impactor (Ice-CVI) and the Ice Selective Inlet (ISI). The IPR as well as the aerosol
6 particles were classified into seven different particle types: (1) black carbon, (2) organic
7 carbon, (3) black carbon internally mixed with organic carbon, (4) minerals, (5) one particle
8 group (termed “Bio/Min/Sal”) that may contain biological particles, minerals, or salts, (6)
9 industrial metals, and (7) lead containing particles. For any sampled particle population it was
10 determined by means of single particle mass spectrometer how many of the analyzed particles
11 belonged to each of these categories. Accordingly, about 30 % of the IPR and roughly 43 %
12 of the background particles contained organic carbon. The measured fractions of minerals in
13 the IPR composition varied from 10 % to 14 %, while the values for the “Bio/Min/Sal” group
14 were between 20 % and 54 %. Two percent to 20 % of the IPR contained organic carbon
15 mixed with black carbon. Both inlets delivered similar results of the chemical composition
16 and of the particle size distribution, although lead was found only in the IPR sampled by the
17 Ice-CVI. The results show that the ice particle residual composition varies substantially
18 between different cloud events, which indicates the influence of different meteorological
19 conditions, such as origin of the air masses, temperature and wind speed.

20

21 **1 Introduction**

22 Ice formation in clouds influences the life-time of clouds, the formation of precipitation, and
23 the Earth’s solar and thermal radiation budget (Lohmann and Feichter, 2005). There are two
24 major pathways by which ice is formed in the atmosphere: homogeneous and heterogeneous
25 freezing. The spontaneous formation of ice inside droplets at temperatures of lower $-37\text{ }^{\circ}\text{C}$
26 and without any catalysts is called homogenous freezing (Cantrell and Heymsfield, 2005) and
27 occurs in high clouds at a saturation similar to that of liquid water (Koop et al., 2000).
28 Heterogeneous freezing needs particulate catalysts to initiate freezing at temperatures higher
29 than $-37\text{ }^{\circ}\text{C}$ in mixed-phase clouds. These catalysts are called ice nucleating particles (INP).
30 There are four different heterogeneous ice formation modes (Pruppacher and Klett, 2010):
31 deposition nucleation, condensation freezing, immersion freezing and contact freezing.
32 Because heterogeneous freezing starts at higher temperatures than homogeneous freezing, ice
33 crystals, which are formed by ice nucleating particles grow faster (by diffusion) and perhaps
34 precipitate first (DeMott et al., 1998). Another reason for faster formation of precipitation is

1 the Wegener-Bergeron-Findeisen process, which plays an important role on cloud formation
2 and precipitation, especially in mixed-phase clouds. In mixed-phase clouds, where
3 supercooled droplets and ice crystals exist, the supercooled droplets evaporate and the water
4 vapor freezes on the ice crystals. Because of the lower saturation vapor pressure over ice than
5 over water the ice crystals will grow at the expense of cloud droplets. This results in a quick
6 freezing of the supercooled water cloud. These iced clouds have a shorter lifetime than
7 supercooled water clouds, because the precipitation formation via the ice phase is more
8 efficient than in warm clouds (Lohmann and Feichter, 2005).

9 Ice formation is very selective because of the low number concentration of particles that can
10 act as INP. Roughly about one in 10^5 atmospheric particles can act as INP (Rogers et al.,
11 1998; DeMott et al., 2010). The ice nucleation is induced at some point on the surface of the
12 nucleating particle and the ice nucleation ability of different types of aerosol particles is
13 usually described either by an ice active-site density approach or by classical nucleation
14 theory involving a contact angle (see e.g. Steinke et al., 2015, and references therein). Several
15 laboratory, field and model studies (DeMott et al., 2003a; DeMott et al., 2003b; Cantrell and
16 Heymsfield, 2005; Kamphus et al., 2010; Hoose et al., 2010; Hartmann et al., 2011; Murray et
17 al., 2011; Hoose and Möhler, 2012; Murray et al., 2012; Atkinson et al., 2013; Diehl et al.,
18 2014) have shown that certain mineral dusts belong to the most important ice nucleators in the
19 atmosphere because of their high abundance. On the other hand, biological particles like
20 bacteria, pollen and spores (e.g. Diehl et al., 2001; von Blohn et al., 2005; Möhler et al., 2007;
21 Pratt et al., 2009; Prenni et al., 2009; Diehl and Wurzler, 2010) are the most efficient ice
22 nuclei (Hoose et al., 2010). Also soot (e.g. Cozic et al., 2008a; Pratt and Prather, 2010; Pratt et
23 al., 2010) and effloresced salts (Abbatt et al., 2006; Wise et al., 2012) have been found to act
24 as ice nuclei. There is also some evidence that glassy organics (Froyd et al., 2010; Murray et
25 al., 2010), porous particles (Adler et al., 2013) and lead-containing particles (Cziczo et al.,
26 2009; Ebert et al., 2011) have favorable ice nucleation qualities.

27 Direct measurements of the chemical composition of single INP or IPR (Ice particle residuals
28 = particles remaining after the ice of cloud ice particles is evaporated) are important to better
29 understand ice formation in mixed-phase clouds, to improve the prediction of precipitation,
30 and to evaluate the influence of anthropogenic aerosol on these processes. Measurements of
31 INP/IPR require instrumentation that can separate the INP/IPR from the background or
32 interstitial aerosol particles and supercooled drops and measure the chemical composition on-
33 line. In several field studies a combination of a CFDC (Continuous Flow Diffusion Chamber;

1 Chen et al., 1998), a CVI (Counterflow Virtual Impactor; Cziczo et al., 2003; Mertes et al.,
2 2007) which separates the ice crystals from background aerosol and a Single Particle Mass
3 Spectrometer (SPMS; DeMott et al., 2003a; Cziczo et al., 2003; Cziczo et al., 2004a; Cziczo
4 et al., 2004b; Pratt et al., 2009; Corbin et al., 2012) to measure ice nucleating particles were
5 deployed. Previous measurements of INP and IPR (on-line and off-line) conducted at the
6 Jungfraujoch showed that particles consisting of mineral components dominate the ice
7 particles number (Kamphus et al., 2010), but also particles containing black carbon (Mertes et
8 al., 2007; Cozic et al., 2008a) and lead (Cziczo et al., 2009; Ebert et al., 2011) were found to
9 be enriched in IPR.

10 Here we report on measurements conducted with the single particle instrument ALABAMA
11 (Aircraft-based Laser Ablation Aerosol Mass Spectrometer; Brands et al., 2011) to analyze
12 the chemical composition of IPR sampled by two unique different ice sampling inlets. The
13 goal of these measurements was to investigate the chemical composition of IPR in ambient
14 mixed-phase clouds in upward transported boundary layer air and free tropospheric air at the
15 high-alpine research station Jungfraujoch (JFJ; 3580 m asl) in Central Europe. Additionally,
16 background aerosol was measured by the use of a total inlet during cloud free conditions.

17

18 **2 Experimental**

19 **2.1 Measurement location and meteorological conditions**

20 The intensive field campaign INUIT-JFJ/CLACE 2013 in January/February 2013 at the High
21 Alpine Research Station Jungfraujoch (JFJ, Swiss Alps; Sphinx Laboratory, 3580 m asl;
22 7°59'2''E, 46°32'53''N) took place in the framework of the DFG (Deutsche
23 Forschungsgemeinschaft)-funded research unit INUIT (Ice Nuclei research UnIT) and the
24 Swiss National Science Foundation-funded *Interaction of aerosols with Clouds and Radiation*
25 *project* (see also overview paper by Schneider et al., this issue). The IPR were sampled out of
26 mixed-phase clouds by the Ice-CVI (Ice Counterflow Virtual Impactor; Mertes et al., 2007)
27 and the ISI (Ice Selective Inlet; Kupiszewski et al., 2014) and fed into the ALABAMA single
28 particle mass spectrometer (Brands et al., 2011). During cloud free times the ALABAMA was
29 connected to the heated total inlet (20 °C; Weingartner et al., 1999) to sample the background
30 aerosol particles. The Ice-CVI and the total inlet were located in a central position on the roof
31 of the Sphinx Laboratory and the ISI was located at the railing at the eastern edge of the roof.
32 A short description of the instruments is given in Sections 2.2 to 2.5. Both the Ice-CVI and
33 ISI sample only ice crystals smaller than 20 μm. Such small ice crystals have grown only by

1 water vapor diffusion and have an age of less than about 20 seconds (Fukuta and Takahashi,
2 1999). It is therefore very likely that these ice crystals have formed in the vicinity of the inlets
3 and that the IPR extracted from such fresh ice crystals corresponds to the original IPN. The
4 meteorological conditions during the campaign along with basic information on the
5 encountered cloud types are summarized in Table 1.

7 **2.2 ALABAMA (Aircraft-based Laser Ablation Aerosol MAss spectrometer)**

8 The ALABAMA (Brands et al., 2011) consist of three sections: aerosol inlet, sizing region
9 and desorption/ionization region. The aerosol inlet is an aerodynamic lens (Liu-type; Liu et
10 al., 1995b; Liu et al., 1995a; Liu et al., 2007; Kamphus et al., 2008) which focuses aerosol
11 particles to a narrow beam by utilizing 6 apertures with decreasing diameter. The particles are
12 separated from ambient air and accelerated to $50 - 100 \text{ m s}^{-1}$, while entering the vacuum-
13 system. For optimal sampling conditions the pressure in front of the aerodynamic lens should
14 be between 0.5 hPa and 5 hPa. For this, a critical orifice is used, which limits the flow to
15 $80 \text{ cm}^3/\text{min}$ under standard working conditions and which reduces the pre-pressure of the lens
16 to 3.8 hPa. During this campaign, the critical orifice was constructed by a constricted O-ring
17 such that also under Jungfraujoch conditions (ambient pressure around 630 hPa) the lens
18 pressure was regulated to best operating conditions. A skimmer separates the first and the
19 second pumping stages. Two continuous wave detection lasers, which are orthogonal to the
20 particle beam, are located in the second pumping stage. As a modification of the instrument
21 described in Brands et al. (2011), two UV laser diodes (Blu-Ray laser; InGaN, 405 nm) are
22 now used. The particles pass through the laser beams and the scattered light is reflected by an
23 elliptical mirror and detected by a photomultiplier tube (PMT). The vacuum aerodynamic
24 particle size (DeCarlo et al., 2004) can be determined from the velocity of the particles by
25 calibration with particles of known size. In this section the ablation laser (pulsed Nd-YAG-
26 Laser, 266 nm; 6 – 8 mJ per pulse, 5.2 ns per pulse, max. 21 Hz) is triggered. If one particle
27 passes through both laser beams, the in-house designed and built electronic control system
28 sends out a trigger signal to the ablation laser, the laser shoots and ionizes the particles. The
29 ions are separated in the Z-shaped bipolar time-of-flight mass spectrometer (TOFWERK AG,
30 Switzerland) by their mass-to-charge ratio. In principle both signals (for positive and negative
31 ions) can be detected using a multichannel plate (MCP). However, for this study only positive
32 ions were available due to technical issues. The average mass resolution $m/\Delta m$ of the Z-ToF
33 mass spectrometer ranges between 100 for low m/z values and 200 for larger ions (Brands et
34 al., 2011). The ALABAMA covers a size diameter range from 100 to 3000 nm (see Fig. 10)

1 with the most efficient size diameter range being between 200 nm and 900 nm. Particles in the
2 size diameter range of approximately 400 nm were ionized at best.

3 Additionally the particle size distribution was measured by a Sky-OPC (Optical Particle
4 Counter; Grimm 1.129) with size diameter range from 0.25 to 52 μm . The Sky-OPC is
5 integrated into the ALABAMA instrument upstream of the aerodynamic lens. Due to the size
6 limitation by the inlet system a maximum particle size of approximately 3 μm was measured.
7 Both instruments sampled through ¼" stainless steel tubes with different lengths (ISI =
8 370 cm; Ice-CVI = 126 cm; total = 261 cm). Particle losses inside the sampling tube were
9 calculated with the Particle Loss Calculator (von der Weiden et al., 2009). In the size range
10 between 200 nm and 500 nm the transmission efficiency is about 99 % for all three inlets. The
11 larger the particles and the longer the tubes, the higher are the losses inside the tubes. For a
12 size of 3 μm the transmission efficiency is between 90 % (tube to the Ice-CVI with a length of
13 126 cm) and 45 % (tube to the ISI with a length of 370 cm). The switching between the three
14 different inlets was performed manually.

16 2.3 Ice-CVI (Ice Counterflow Virtual Impactor)

17 The Ice-CVI is designed to sample IPR from mixed-phase clouds. A detailed description and
18 instrumental characterization is provided in Mertes et al. (2007) thus it is only briefly
19 described here. The Ice-CVI uses an omnidirectional inlet, which removes particles larger
20 than 20 μm from the aspired air. The separation of the particles larger than 20 μm out of the
21 sample flow is done with a virtual impactor (VI) downstream of the inlet. After that the
22 supercooled droplets ($> 5 \mu\text{m}$) are removed by a pre-impactor (PI). The PI consists of
23 impaction plates colder than 0 °C. The supercooled droplets freeze on these plates upon
24 contact while the ice crystals bounce off and remain in the sample flow. Subsequently, a
25 counterflow virtual impactor (CVI, Mertes et al., 2005) rejects interstitial aerosol particles,
26 and potentially remaining supercooled drops and ice particles smaller than 5 μm . Thus, all
27 supercooled drops are pre-segregated by the cut-off behavior of PI and CVI and only small ice
28 crystals, i.e. those with aerodynamic size diameters between 5 μm and 20 μm (cut-off
29 properties of VI and CVI) remain in the sample flow. A transmission efficiency of the PI
30 concerning supercooled drops and ice particles of about 0 % and 100 %, respectively, as
31 expected from Tenberken-Pötzsch et al. (2000) and Murphy et al. (2004), was shown by
32 Mertes et al. (2007).

1 For complete sublimation, the small ice particles are injected into particle-free and dry carrier
2 air inside the CVI. The aerosol particles that remain after evaporation/sublimation of the ice
3 are termed IPR. The walls of the evaporation tube as well as the carrier air are at room
4 temperature. The IPR can be transferred to different aerosol instruments for physico-chemical
5 characterization. The sampling principle of the CVI leads to an enrichment of the collected
6 particles. This enrichment (factors between 5 and 10) is given by the ratio of the velocity
7 upstream and downstream the CVI inlet tip.

8 A CPC (Condensation Particle Counter, model 3010, TSI Inc.) was connected downstream of
9 the Ice-CVI at a suitable position for measuring to determine the total particle number
10 concentration of the IPR. Significant non-zero particle number concentrations indicate the
11 presence of ice particles in mixed-phase clouds.

12

13 **2.4 ISI (Ice Selective Inlet)**

14 Extracting IPR from mixed-phase clouds was also possible with the ISI (detailed description
15 in Kupiszewski et al., 2014). Small ice crystals (the IPR within are assumed to be
16 representative of the original ice nuclei) are segregated from larger ice crystals, supercooled
17 droplets and interstitial aerosol particles by a combination of four inlet components. The first
18 component is the omnidirectional inlet, which is shielded from above **in order to prevent** large
19 precipitating particles entering the inlet. Subsequently, a custom-made cyclone (BGI Inc.)
20 removes ice crystals, with aerodynamic diameters of ca. 20 μm (D_{50} cut size, i.e., the particle
21 diameter at which 50 % of incoming particles are removed from the sample flow) and larger.
22 Further downstream the removal of supercooled droplets takes place in a custom built ice-
23 coated droplet evaporation chamber using the Wegener-Bergeron-Findeisen (WBF) process.
24 **The WBF process is used by deploying a chamber with ice-coated inner walls within the ISI,**
25 **thus providing an environment in which there is saturation with respect to ice. This in turn**
26 **means that there is sub-saturation with respect to water (due to the difference in saturation**
27 **vapour pressures over water and ice), which leads to evaporation of liquid droplets inside the**
28 **chamber and transfer of water vapour from the droplets to the ice-coated walls of the**
29 **chamber. The design of the chamber is described in detail in Sect. 2.2 of Kupiszewski et al.**
30 **(2014).** The key difference to the Ice-CVI is that here the separation of supercooled droplets
31 from ice crystals is conducted in the airborne state via evaporation. Finally, the interstitial
32 aerosol and the cloud condensation nuclei (CCN) released during droplet evaporation are
33 removed by a pumped counterflow virtual impactor ($D_{50} \approx 5 \mu\text{m}$). The extracted ice crystals

1 with aerodynamic diameter between 5 μm and 20 μm are heated and the released IPR are
2 transferred into the laboratory for physical and chemical characterization using state of the art
3 single particle aerosol instrumentation (single particle soot photometer (SP2), ALABAMA
4 single particle mass spectrometer, Wideband Integrated Bioaerosol Sensor (WIBS) and a
5 Grimm optical particle counter). The ISI additionally allows for simultaneous counting, sizing
6 and imaging of the hydrometers contained in the cloud with the use of two WELAS (white
7 light aerosol spectrometer; Heim et al., 2008; Rosati et al., 2015) aerosol sensor system (Palas
8 GmbH) and a Particle Phase Discriminator (PPD; Kaye et al., 2008; Vochezer et al., 2015).
9 These are placed behind the cyclone (first WELAS OPSS) as well as behind the droplet
10 evaporation unit (second WELAS OPSS and PPD). The ISI, together with the downstream
11 aerosol instrumentation, thus provides number size distribution of sampled hydrometers and
12 aerosol particles, information on ice crystal shape and surface roughness, as well as number
13 size distribution and chemical composition of ice residual particles.
14 The ISI enriches the collected particles inside the pumped counterflow virtual impactor with
15 an enrichment factor between 1 and 10.

16

17 **2.5 Supplement instruments**

18 **Particle Volume Monitor (PVM; Gerber, 1996):** The PVM measured the liquid water
19 content (LWC = total liquid water present in a cloud) and the surface area of cloud droplets
20 (e.g., Wendisch, 1998). Cloud droplets cross a laser beam and the scattered light is measured.
21 Elevated LWC values are used to identify cloud events, using a threshold value of 20 mg m^{-3} .
22 During the measurement campaign two PVMs were operated by PSI (Paul Scherrer Institute)
23 and the University of Manchester.

24

25 **Cloud Droplet Probe (CDP):** The CDP is a forward-scattering spectrometer for
26 determination of particles sizes and concentrations of cloud droplets. Counting and sizing of
27 the particles is done by detecting light scattered by the particles (for more details see for
28 example Lance et al., 2010).

29

30 **Scanning electron microscopy (SEM):** Ice particle residuals (IPR) were collected by a two
31 stage impactor system (50 % cut-off aerodynamic diameters 1.0 and 0.1 μm , respectively) on
32 transmission electron microscopy grids and elemental boron substrates and analyzed by
33 scanning electron microscopy (SEM, FEI Quanta 200 FEG, FEI Eindhoven, the Netherlands)
34 and energy-dispersive X-ray fluorescence (EDX, EDAX, Tilburg, The Netherlands) to

1 characterize the particles with regard to their chemical composition, morphology, size,
2 internal mixing state and electron beam stability (volatility). (For further details see Kandler
3 et al., 2011). Here we only present the size distribution of IPR measured by SEM and
4 compare them to the IPR sizes obtained by the ALABAMA and the Sky-OPC. Detailed
5 results on the chemical composition analysis of IPR, INP, and background aerosol by SEM
6 are given in Worringer et al., 2015.

7

8 **2.6 ALABAMA data evaluation**

9 The data evaluation was done with the software package CRISP (Concise Retrieval of
10 Information from Single Particles; Klimach, 2012) based on the software IGOR Pro (Version
11 6, WaveMetrics) and following the procedure described in Roth (2014). This software
12 package includes mass calibration, conversion of the mass spectra into ‘stick’ spectra by
13 integration over the peak width, and various procedures to sort the mass spectra into groups or
14 clusters of similar spectra. This sorting can be done manually, by means of cluster algorithms
15 (*fuzzy c-means* or *k-means*), or by searching for certain parameters (e.g., ion signals at a
16 certain m/z ratio).

17 The Clustering depends to some degree on the interstitial settings of the algorithm and its
18 parameters. The start parameters are *Normalization type and time*, *Preprocessing type*,
19 *Initialisation type*, the *number of clusters* and the *start cluster difference*. Using the *fuzzy c-*
20 *means* algorithm (e.g. Bezdek et al., 1984; Hinz et al., 1999; Huang et al., 2013) the
21 parameters *fuzzifier* and *fuzzy abort* have to be set accordingly. The cluster parameters are
22 explained in detail in the appendix.

23 As only the cations were detected during the INUIT-JFJ/CLACE 2013 campaign the grouping
24 of the clusters into different particle types by using certain marker peaks was difficult in some
25 cases. At the end, seven different particles types were extracted: black carbon (BC), organic
26 carbon (OC), internal mixture of black carbon and organic carbon (BC/OC), lead containing
27 particles, industrial metals, **Bio/Min/Sal** (containing possibly bioparticles, minerals or salts),
28 and minerals. The average mass spectra of each particle type are shown in Figure 1. The
29 association of the mass spectra to the individual particle types is explained in detail below.
30 The error limits of the number of mass spectra per particle types was estimated using counting
31 statistics.

32 *Organic fraction: Black carbon (BC) and organic carbon (OC)*

1 The clusters categorized as BC are due to the presence of the typical fragmentation of C_n^+
2 (Dall'Osto et al., 2004; Pratt and Prather, 2010) of higher mass-to-charge ratios (see Fig. 1a).
3 In comparison to the black carbon, the organic carbon cluster (Fig. 1b) shows different
4 fragmentations of aromatics ($[C_4H_3]^+$ (m/z 51); $[C_5H_3]^+$ (m/z 63); $[C_6H_5]^+$ (m/z 77); $[C_7H_7]^+$
5 (m/z 91) (Dall'Osto and Harrison, 2006; Pratt and Prather, 2010)), amines ($[CHN]^+$ (m/z 27);
6 $[C_3H_8N]^+$ (m/z 58) (Pratt and Prather, 2010)) or other organic material ($[C_2H_3]^+$ (m/z 27);
7 $[C_3H]^+$ (m/z 37); $[CH_3O/CHNO]^+$ (m/z 43); $[C_4H_2/C_3N]^+$ (m/z 50); $[C_3H_7O]^+$ (m/z 59);
8 $[C_{11}H_{10}/C_8H_{16}NO]^+$ (m/z 142) (Pratt and Prather, 2010)). Black carbon internally mixed with
9 organic (BC/OC) shows both fragmentation types of $[C_n]^+$ and $[C_nH_mNO]^+$ (Corbin et al.,
10 2012; Fig. 1c).

11 Closer consideration of the three different carbon clusters (BC, BC/OC, OC) shows that
12 potassium is present in nearly every mass spectrum in the background aerosol. In the
13 background aerosol 45282 mass spectra including black or organic carbon were measured;
14 92 % of all these mass spectra show potassium. The BC-type includes the mass spectra
15 without potassium (20 %; only 1 % in the OC/BC-type and 6 % in the OC-type). The same
16 observation holds for the IPR composition measured using the Ice-CVI (Section 3.1): The
17 particle types including organic carbon show more spectra including potassium (91 % of all
18 OC and OC/BC). In contrast to that the particle type including organic carbon measured while
19 using the ISI show more spectra without a high potassium peak (77 %). Previous
20 investigations show that the presence of high potassium peak in the mass spectra in
21 combination with organic carbon is a marker for biomass burning (e.g. Pratt and Prather,
22 2010; Twohy et al., 2010; Corbin et al., 2012) or biogenic aerosol (e.g. Trimborn et al., 2002;
23 Pratt and Prather, 2010). However, because only cations were measured during this
24 experiment, spectra with a potassium peak mixed with BC and OC cannot clearly assigned to
25 biomass burning or biological particles.

26 *Lead containing particles*

27 Previous measurements at the Jungfraujoch showed that lead-containing particles can be
28 found in IPR (Cziczo et al., 2009; Ebert et al., 2011). The cluster of lead containing particles
29 shows the typical isotope pattern of lead (m/z 208, 207, 206, 204; Fig. 1d).

30 *Industrial metals*

31 The cluster classified as 'industrial metals' (Fig. 1e) shows metal signals from anthropogenic
32 (urban/industrial) emissions like $[Cr]^+$ (m/z 52, 53, 54, 50), $[Co]^+$ (m/z 59), $[V]/[VO]^+$ (m/z

1 51/67), [Fe]⁺ (m/z 54, 56), [Mn]⁺ (m/z 55), [Ni]⁺ (m/z 58, 60, 61, 62) and [Zn]⁺ (m/z 64, 66,
2 68) (de Foy et al., 2012). Chromium and nickel containing particles might also be due to
3 contamination from stainless steel tubes. However, due to the low flow speed and laminar
4 flow conditions in our sampling tubes we can exclude abrasion of material by collision of
5 aerosol particles with the flow tube walls. Possible sources for steel contaminations particles
6 are only opening and closing of the valves. Thus, detection of such “industrial metals”
7 particles within a few seconds after opening/closing of a valve indicates contamination
8 artifacts. However, during these time periods no particles of the type “industrial metals” were
9 detected and therefore no particles were excluded as contamination.

10 *Bio/Min/Sal cluster*

11 The spectra in this cluster type contain only metal cations and can therefore not be
12 unambiguously be assigned to a certain particle type. The spectra in the *Bio/Min/Sal* cluster
13 show a mixture or single signals of [Na]⁺ (m/z 23), [Al]⁺ (m/z 27), [K]⁺ (m/z 39, 40), [Ca]⁺
14 (m/z 40) and/or [Fe]⁺ (m/z 54, 56). An example of an average spectrum of this particle type is
15 shown in Figure 1f.

16 The interpretation of this particle cluster is difficult. The signals are most likely cations of
17 different kinds of salts (K₂SO₄ or KNO₃ are more conceivable than sea salt), minerals or
18 primary biological particles, but without corresponding anion an unambiguously assignment
19 to one particle type is not possible. Previous investigations of biological aerosol have shown
20 that metal signals like [Na]⁺ (m/z 23), [K]⁺ (m/z 39) or [Ca]⁺ (m/z 40) dominate the cation
21 spectra (e.g. Frank et al., 2011; Pratt and Prather, 2010; Ferguson et al., 2004). During
22 laboratory studies with test aerosol we also observed mainly single metal signals of sodium
23 and potassium in biological particles, but as well in mineral dust particles. Also iron (Fe⁺ m/z
24 54, 56), which is a typical marker for mineral dust, was observed to occur in some biological
25 particles (maybe due to hemoglobin or due to other ions with an m/z of 56). For these reasons
26 the cluster “*Bio/Min/Sal*” cannot be further differentiated in this study.

27 *Minerals*

28 Typically marker peaks of minerals are: [Li]⁺ (m/z 6, 7), [Mg]⁺ (m/z 24, 25), [Al]⁺ (m/z 27),
29 [Si]/[SiO]⁺ (m/z 28, 29, 30/44), [Ca]⁺ (m/z 40), [Ti]/[TiO]⁺ (m/z 48, 46, 47, 49, 50/64) and/or
30 [Fe]⁺ (m/z 54, 56) (Trimborn et al., 2002; Dall'Osto et al., 2004; Dall'Osto et al., 2010)
31 Clusters containing at least three of the above marker peaks were classified as *minerals* (Fig.
32 1g).

1 *Others*

2 The cluster *others* includes all particle mass spectra, which are very noisy or could not be
3 unambiguously assigned to any of the above cluster types.

4

5 **3 Results and Discussion**

6 During the INUIT-JFJ/CLACE 2013 campaign we measured 71064 background aerosol
7 particles during a measurement time of 217 h and 1803 ice particle residuals during a
8 measurement time of 342 h: 1659 spectra behind the Ice-CVI during 273 h and 144 spectra
9 behind the ISI during 69 h.

10 **3.1 Ice particle residual composition and potential aerosol sources**

11 The ALABAMA was used to analyze ice crystals extracted from mixed-phase clouds by two
12 different inlets: ISI and Ice-CVI. The times during which cloud particles were sampled by the
13 ISI were identified using the cloud liquid water content (LWC)-data measured with two PVM
14 instruments (PSI and University of Manchester) and a CDP (University of Manchester) while
15 the times when the Ice-CVI sampled cloud particles were identified using the ice residual
16 concentration measured with the CPC behind the Ice-CVI. According to the Ice-CVI sampling
17 capability, which is described in section 2.3, the cloud event periods where the ALABAMA
18 was coupled to the Ice-CVI were classified. This classification resulted in (++)-sampling-
19 periods, when there was hardly any precipitation or blown snow, so that small secondary ice
20 particles created by crystal-crystal collisions can be neglected and thus INP containing
21 primary ice particles are sampled. For (+)-sampling-periods, the sampling of small secondary
22 ice particles, that do not contain a particle that served as INP, cannot be excluded, since
23 precipitation and blown snow occurred. As a consequence, the ice particle residues analyzed
24 during (++)-sampling-periods are attributed to INP whereas for (+)-sampling-periods this
25 relation is not that clear and might be only valid for particles that are larger than at least
26 300 nm. Table 1 shows the identified cloud events along with sampling time, number of
27 acquired mass spectra, sampling-period classification, and weather conditions and cloud type.
28 The meteorological situation during the whole campaign was mainly influenced by low
29 pressure systems with low temperatures.

30 The occurrence of artifacts during the measurements using the Ice-CVI as well as the ISI is
31 described in Worringer et al. (2015). Aluminum oxide is an artifact of the measurements with
32 the Ice-CVI and silicon oxide is an artifact of the measurements with the ISI. The impaction

1 plates of the Ice-CVI contain aluminum. When the ice crystals bounce on the impactation plates
2 of the Ice-CVI, some aluminum oxide particles are removed from these plates and can be
3 carried with the air stream to the analysis instrument. The ISI was calibrated with silicon
4 oxide particles, which are remained in the system and occurred during the measurements as
5 artifacts. In our dataset it is difficult to distinguish between artifacts and a real component of a
6 particle. Therefore the peaks of m/z 27 (Al^+) and m/z 43 (AlO^+) with respect of the Ice-CVI
7 measurements and the peak of m/z 28 (Si^+) with respect to the ISI measurements were
8 ignored.

9 ~~In total, we measured 329 spectra during the (++) events (61 h measurement time) and 1335~~
10 ~~spectra during the (+-) events (160 h measurement time) by using the Ice-CVI. Behind the ISI~~
11 ~~we measured 144 spectra during 35 h measurement time.~~

12 Figure 2 shows the results of the clustering algorithm for all ice particle residual
13 measurements. The Ice-CVI data were analyzed separately for all **(++)-sampling-periods** (left
14 pie chart) and for all **(+)-sampling-periods** (middle pie chart). The right pie chart shows the
15 results for all ISI measurements. **This analysis shows that the metal containing particle groups**
16 **(Bio/Min/Sal, industrial metals, and minerals) are the dominating fractions in IPR**
17 **composition measured by using the Ice-CVI during the (+)-sampling-periods (49 %) and the**
18 **ISI (66 %).** Interestingly, during the **(++)-sampling-periods metal containing particles**
19 **represent only a smaller fraction (40 %).** We also detected lead containing particles, but only
20 while sampling through the Ice-CVI **(7 % in the (++) cases and 12 % in the (+-) cases).**
21 During the **(++)-sampling-periods** measured behind the Ice-CVI, black carbon internally
22 mixed with organic carbon (20 %) is one of the main fractions. Together with OC **(26 %),**
23 these two particle types represent **46 %** of all IPR. It is worth noting that there are more
24 similarities between the Ice-CVI and ISI during **(+)-sampling-periods than the (++)-**
25 **sampling-periods.** A reason for that can be that the ISI measurements are also influenced by
26 secondary ice or incoming snowflakes. During the **(+)-sampling-periods** as well as during the
27 **ISI-sampling-periods** precipitation occurred. On the other hand, the measurements at the ISI
28 and the Ice-CVI were not done simultaneously, such that **also** differences in the air mass
29 origin could play a role here.

30 **Backward trajectories were compiled using CRISP (Klimach, 2012) with access to HYSPLIT**
31 **(Hybrid Single Particle Lagrangian Integrated Trajectory Model, National Oceanic and**
32 **Atmospheric Administration; Draxler and Rolph, 2012; Rolph, 2014), which was run using**
33 **GDAS (Global Data Assimilation System) meteorological dataset (starting altitude: 3580 m**

1 asl.; hours back: 72 h; initialize time: end of the current event). The geographical origin of the
2 air masses and change of altitude during the transport can be used to infer whether the air
3 masses are influenced by surface emissions or mainly characterized by free tropospheric air
4 and long distance transport. During all measurement episodes (total, Ice-CVI, ISI) the air
5 masses approached the measurement station from slightly different directions (Fig. 3). The air
6 masses during the (++)-sampling-periods came primarily from north-western directions from
7 North America over France, while during the (+)-sampling-periods the air masses arrived
8 more from the North, over Scandinavia and Germany. The origin of the air masses during the
9 measurements using the ISI are similar to those of the (++)-sampling-periods but with a
10 tendency towards southern directions (Fig. 3). During non-cloud episodes air masses
11 approached the measurement location from all directions.

12 Previous measurements at the JFJ showed that there is an enrichment of mineral dust and
13 black carbon in IPR composition (Mertes et al., 2007; Cozic et al., 2008b; Kamphus et al.,
14 2010). An enrichment of black carbon is not observed in our measurements. Cziczo et al.
15 (2009) found that lead-containing particles are also enriched in the IPR compared to the
16 background aerosol. Our measurements show also that mineral dust (measured behind the ISI
17 and during the (+) sampling periods) and lead (only measured behind the Ice-CVI) were two
18 of the dominating fractions in the IPR composition. Worrigen et al. (2015) report on SEM
19 analyses of particles sampled during the same project behind the ISI and the Ice-CVI. They
20 found lead in two states, namely homogeneous lead-rich particles and small lead inclusions in
21 particles of other material. While the small lead inclusions are not considered as potential
22 artifacts, the large lead-rich particles might be artifacts from mechanical re-suspension of lead
23 containing particles from the surface of the impaction plates of the Ice-CVI. In our case the
24 lead-containing particles are internally mixed with organics, minerals or other metals,
25 therefore we interpret these particles as real atmospheric particles and not as artifacts.

26

27 **3.2 Event-to-event variability**

28 Figure 4 shows the IPR chemical composition measured during seven different events. These
29 sampling-periods were chosen, because these are (++)-sampling-periods and contamination of
30 e.g. incoming snowflakes are excluded and during these sampling-periods the most particles
31 were measured (except for the ISI measurements). Although the air masses sampled during all
32 seven events have very similar origin and history of altitude (Fig. 5a and 5b) the chemical
33 composition for each event is different from the others. All air masses during these events

1 arrived rising from a western direction to the measurement platform, such that the chemical
2 composition of these seven events was probably influenced by local emissions. During Event
3 3 and 4 the air masses arrived from a south-western direction. This is reflected in the chemical
4 composition: Both events show organic carbon as the main fraction and the amount of lead
5 containing particles is comparable (13 % Event 3 and 15 % Event 4). The differences between
6 Event 3 and 4 are the absence of BC in Event 4 and of minerals in Event 3. The chemical
7 composition of Event 12 is also comparable to Event 4 and Event 6. In this case also the
8 dominating fractions are OC and Bio/Min/Sal. The absence of lead containing particles (as
9 well as in Event 6) and the presence of 2 % BC/OC in Event 12 or 4 % in Event 6 are the
10 main differences between Event 12, 6 and Event 4.

11 Another reason for similarities and differences between these 7 events can be the temperature
12 (Fig. 6). Events 3, 4 and 6 have the highest fraction of OC (nearly 50 % of the particles) and
13 Event 6 additionally a high amount of metal containing particles (37 %; industrial metals,
14 minerals and Bio/Min/Sal). Event 3 and 4 have also the highest temperatures (Event 3:
15 average temperature $\bar{T} = -7.8$ °C; Event 4: $\bar{T} = -8.6$ °C). In contrast to that, Event 12 has the
16 highest amount of metal containing particles (65 %) and has an average temperature of about
17 $\bar{T} = -17$ °C. Event 6 has also a high amount of metal containing particles and the lowest
18 average temperature of about $\bar{T} = -21$ °C. Event 4 has also a high amount (41 %) of 'metal'
19 components but by closer consideration of the Bio/Min/Sal-type it can be shown that this type
20 consists mostly of metal ions together with organic ions in Event 4 and more single metal ion
21 signals during Event 12 and 6. This high amount of metal components measured at lower
22 temperatures is also comparable to Event 2 and 5. 62 % of the chemical composition of Event
23 2 and 51 % of Event 5 consist of metal components. Therefore the data indicate that organic
24 material dominates the IPR composition at higher temperatures and metal components from
25 mineral or salty origin dominate at lower temperatures. This finding is surprising, because
26 previous studies found that organics are ice active only at temperatures below -40 °C (Hoose
27 and Möhler, 2012). However, the organic signal in the mass spectra might also originate from
28 bioparticles, which would explain the ice formation at higher temperatures, because
29 bioparticles have been reported in the literature by many authors to be ice active at higher
30 temperatures (e.g. Möhler et al., 2008; Hoose and Möhler, 2012). Another source can be soil
31 organic matter. Tobo et al. (2014) shows that organic-rich particles are more important than
32 mineral particles for the ice-nucleating ability of agricultural soil dust at temperatures higher
33 than -36 °C. An exception is Event 1. During this event the second lowest temperatures ($\bar{T} = -$
34 19.2 °C) of all events prevailed, but a high relative amount of organic components was

1 measured (80 % BC/OC and OC). A reason for this high relative amount of organic material
2 can be local emissions. At this time the total particle concentration (measured inside the cloud
3 using the total inlet) shows a short period with high concentrations ($> 2000 \text{ cm}^{-3}$, see Fig. 7),
4 which indicates local emissions.

6 3.3 Comparison between background aerosol particle composition and IPR

7 During non-cloud phases the ALABAMA was connected to the total inlet to measure
8 background aerosol particles. Figure 8 shows the averaged composition of all background
9 aerosol particles along with the IPR composition measured behind the Ice-CVI during the
10 (++)-sampling-periods and behind the ISI. The temperature range during the background
11 aerosol measurements was between $-27 \text{ }^{\circ}\text{C}$ and $-6 \text{ }^{\circ}\text{C}$, which is comparable to the temperature
12 range of the cloud events. In comparison to the IPR composition, the measurements show a
13 high fraction of black carbon (17 %; Fig. 8) and only a minor fraction of BC/OC (4 %). The
14 fraction of the particle type ‘BC/OC’ in the background aerosol is comparable to that in the
15 IPR sampled by the ISI. However, also the aerosol particles are dominated by the particle
16 types ‘organic carbon’ (43 %) and ‘Bio/Min/Sal’ (30 %). ~~Previous high mountain-top~~
17 ~~measurements at the Storm Peak Laboratory (SPL; 3200 m asl.) in northern Colorado show~~
18 ~~also organic as the major component in the background aerosol (DeMott et al., 2003a ;Cziczo~~
19 ~~et al., 2004a).~~ The differences as compared to the IPR compositions are the negligibly small
20 amount ($< 0.1 \%$) of lead-containing particles and the much smaller amount of minerals
21 (about 1 %). This is partly in contrast to other measurements. Previous measurements with the
22 ATOFMS at the Jungfraujoch as well as airborne measurements over North America show 5
23 and 10 % lead-containing particles in the background aerosol (Murphy et al., 2007) but also
24 only a small amount of mineral dust or fly ash measured at the Storm Peak Laboratory (SPL;
25 3200 m asl.) in northern Colorado (DeMott et al., 2003a). Data from the SPL show also
26 organic as the major component in the background aerosol (DeMott et al., 2003a ;Cziczo et
27 al., 2004a). A part of the organic material detected in the IPR as well as in the background
28 aerosol may be due to could be local anthropogenic emissions. In the vicinity of the
29 measurement station is a large and popular skiing region. Additionally, cooking emissions,
30 diesel exhaust or residential heating may also play a role.

31 ~~Previous investigations show that the presence of a high potassium peak in the mass spectra in~~
32 ~~combination with organic carbon is a marker for biomass burning (e.g. Pratt and Prather,~~
33 ~~2010; Twohy et al., 2010; Corbin et al., 2012) or biogenic aerosol (e.g. Trimborn, 2002; Pratt~~
34 ~~and Prather, 2010). Closer consideration of the three different carbon clusters (BC, BC/OC,~~

1 OC) shows that potassium is present in nearly every mass spectrum. In the background
2 aerosol 45282 mass spectra including black or organic carbon were measured: 92 % of all
3 these mass spectra show potassium. The BC type includes the mass spectra without potassium
4 (20 %; only 1 % in the OC/BC type and 6 % in the OC type). The same observation holds for
5 the IPR composition: The particle types including organic carbon measured by the use of the
6 Ice-CVI as well as the ISI show more spectra including potassium (94 % of all OC and
7 OC/BC spectra while using the Ice-CVI and 98 % while using the ISI). Because of this results
8 spectra with a potassium peak cannot clearly assigned to biomass burning or biological
9 particles.

11 3.4 Influence of the air mass trajectories on the aerosol composition

12 The chemical composition of the background (out-of-cloud) aerosol has been shown in Figure
13 8. The high variability of the measured IPR might be due to the different aerosol properties
14 when air masses arrive from different geographic origin and/or different altitudes at the
15 Jungfrauoch. Therefore, we investigate here how air masses origin influences the aerosol
16 properties. Because of the saddle position of the JFJ, there are only two possibilities of how
17 the air masses can approach the measurement platform (Fig. 6). Predominantly, the air masses
18 arrive from northwestern direction over the Swiss plateau (wind direction of approx. 315°).
19 Less frequently, the approach was from a southeastern direction from the inner Alps (approx.
20 135°, via the Aletsch Glacier). According to the direction from which air masses approach the
21 measurement platform, the chemical composition is influenced predominantly by
22 anthropogenic and organic emissions or by mineral dust sources. To assess the influence of
23 origin and the average altitude of the air masses during their transport to the Jungfrauoch, we
24 compare two time periods with aerosol measurements using the total inlet (11.02. and
25 19.02.2013). The chosen time periods exhibit nearly the same relative humidity, potential wet
26 bulb temperature and temperature (Fig. 6).

27 Fig. 9 shows the chemical composition of the aerosol particles for both time periods, along
28 with corresponding back-trajectories. The comparison shows that more particles containing
29 BC (internally mixed with ammonium or the Bio/Min/Sal-type) and the Bio/Min/Sal-type are
30 observed in the air masses arriving from a northwestern direction (11.02.2013), while the
31 measurements during the SE-event show more OC containing particles (internally mixed with

1 ammonium or the Bio/Min/Sal-type). The fraction of particles measured containing only
2 potassium and pure organic carbon is very similar for both events.

3 The air masses not only approached locally from different directions, but also had different
4 long-range origins (see back trajectories Fig. 9). The air masses measured on 19.02.2013
5 (NW-event) approached from the north over Germany and have remained at approximately
6 the same altitude as the Jungfrauoch during the last 72 hours. Thus, these air masses have
7 been in the free troposphere for at least three days prior to the measurement. The chemical
8 composition of the measured particles is dominated by black carbon internally mixed with
9 ammonium and by the Bio/Min/Sal-type. In contrast to the 19.02.2013, the air masses on
10 11.02.2013 (SE-event) approached over southern France and the western Po Valley, have
11 been lifted from low altitudes and clearly influenced by the boundary layer. These air masses
12 contain higher amounts of OC containing particles. This shows that the vertical uplifting
13 brings anthropogenic emissions (e.g., from the Po Valley) up to the altitude of the
14 measurement station, while the air masses measured during the NW-event contain a higher
15 amount of sea salt or mineral dust (both contained in the Bio/Min/Sal-type) and aged,
16 processed BC particles.

17

18 3.5 Size distribution of IPR and background aerosol

19 Ice-CVI and ISI both extract small ice particles in the size range from 5 μm to 20 μm from
20 mixed-phase clouds (Mertes et al., 2007; Kupiszewski et al., 2014) while the upper cut-off of
21 the total inlet was 40 μm for wind speeds up to 20 m/s (Weingartner et al., 1999). In Figure 10
22 is shown the size distributions of IPR measured with the ALABAMA (a, c, e) and the Sky-
23 OPC (b, d, f) compared with ESEM measurements (a and b). Figure 10a shows the size
24 distributions of IPR measured by ALABAMA and ESEM applying the Ice-CVI. Figure 10e
25 the size distributions measured by ALABAMA and ESEM using the ISI. In Figure 10b and d,
26 the corresponding size distributions measured with the Sky-OPC are shown. Figure 10c
27 shows the size distributions measured by ALABAMA and Fig. 10d measured by the Sky-
28 OPC using the ISI. Fig. 10e and f show the size distribution measured by ALABAMA and the
29 Sky-OPC using the total inlet. The measurement times for Ice-CVI and ISI are given in Table
30 1; the measurements using the total inlet were done during all cloud-free periods. For all three
31 inlets (total, Ice-CVI, ISI) most of the detected IPR measured with the ALABAMA and also
32 with the Sky-OPC is in the size range of 300–650 nm. The IPR size distributions measured
33 by ALABAMA and SEM behind the Ice-CVI are in very good agreement (Fig. 10a). In

1 contrast, the IPR size distribution measured behind the ISI are somewhat different. (Fig. 10c):
2 Both distributions show two maxima, but the positions of the maxima are not the same. The
3 ALABAMA IPR distribution has a primary maximum in the size bin 0.4–0.5 μm and a
4 secondary maximum in the size bin 1.3–1.6 μm . The SEM IPR distribution both maxima are
5 shifted to larger sizes. The first maximum is in the 1–1.5 μm interval and the second
6 maximum is in the 2–2.5 μm interval. However, total count rates are very low, so this
7 difference should not be overemphasized. Furthermore, SEM measured geometric diameters,
8 while the sizes determined by ALABAMA refer to the vacuum aerodynamic diameter. SEM
9 and ALABAMA measure the size distribution in different ways: ALABAMA measures
10 vacuum aerodynamic diameter and SEM geometric diameter. To compare both data sets the
11 SEM data were converted to the aerodynamic diameter. The vacuum aerodynamic diameter
12 was calculated according to DeCarlo et al. (2004) with the assumption that the projected area
13 diameter from the electron microscopy analysis equals the volume-equivalent diameter.
14 Furthermore, an average dynamic shape factor of 1.2 was used for the calculations. This
15 seems to be justified, as most non-extremely shaped particles show low variation in shape
16 factor (Schneider et al., 2006; Zelenyuk et al., 2006), except for soot with high fractal
17 dimension (Slowik et al., 2004) which was not commonly encountered in the present work.
18 The density ratios required for the calculation were estimated from the major composition of
19 the particle according to the values from Weinbruch et al. (2014) (Table S2). The IPR size
20 distributions measured by ALABAMA and SEM behind the Ice-CVI show the same behavior
21 (Fig. 10a) with shifting the SEM graph to larger sizes. Both distributions show a slow
22 increase at small size ranges and a higher number of larger particles. The ALABAMA IPR
23 distribution has a primary maximum in the size bin 300–400 μm and a secondary maximum
24 in the size bin 500–600 μm . The SEM size distribution shows a single maximum at about
25 900 μm . However, total count rates are very low, so this difference should not be
26 overemphasized. For all three inlets (total, Ice-CVI, ISI) most of the detected particles
27 measured with the ALABAMA are in the size range of 300–800 nm. In contrast to the size
28 distribution measured with the ALABAMA (and also ESEM for the Ice-CVI), the size
29 distributions measured with Sky-OPC (b, d, f) do not show a maximum but only a decreasing
30 number concentration of particles with increasing particle diameter for all three inlet types.
31 The difference to the ALABAMA size distributions is a result of the detection and ionization
32 efficiency of the ALABAMA which is the highest around 400 nm. The IPR size distributions
33 measured with the ESEM and Sky-OPC (10b) show both a decreasing number of particles at
34 larger size ranges. In contrast to the size distribution measured with the Sky-OPC the size

1 distribution measured with ESEM shows a lower number of particles at smaller size ranges.
2 The Sky-OPC size distribution measured behind the ISI (10d) shows an unexplained sudden
3 decrease above 3 μm . Figure 10f shows that the background aerosol particles (Sky-OPC) have
4 a size range of up to approximately 20 μm but with highest number concentration in the
5 smaller range ($< 1 \mu\text{m}$). The relative amount of larger particles is ($d > 1 \mu\text{m}$) clearly elevated
6 for the IPR, both behind the ISI and Ice-CVI (10b and d) in comparison to the background
7 particles (10f). Note that the measurements of background particles (out of cloud) and IPR
8 were conducted by definition at different times, such that the absolute amount of particles
9 cannot be compared.

10 As mentioned above, relatively more IPR larger than 1 μm were observed compared to the
11 background aerosol particles. The percentage of the background aerosol particles larger than
12 1 μm (vacuum aerodynamic diameter) measured by ALABAMA was only 3 %, while 12 -
13 13 % of the IPR behind the Ice-CVI and ISI were larger than 1 μm . Therefore, the chemical
14 composition of these particles has been analyzed separately:

15 Figure 11 shows the chemical composition of the IPR and background aerosol particles larger
16 than 1 μm . The most frequent particle types found in the IPR larger than 1 μm are OC, lead
17 containing particles, Bio/Min/Sal, and minerals for the Ice-CVI, and Bio/Min/Sal and
18 minerals for the ISI. Lead containing particles are found in the IPR sampled by the Ice-CVI to
19 a similar percentage (14 %) as in all IPR (7 %, see Figs. 2 and 8). Interestingly, the
20 background aerosol supermicron particles are dominated by the Bio/Min/Sal-type and OC.
21 The finding strengthens the assumption that the metal ions detected in these mass spectra are
22 likely metal cations of primary aerosol like sea salt, minerals or biological particles, which are
23 mainly found in the coarse mode (e.g., Seinfeld and Pandis, 2006). Such large primary
24 particles are also expected to act efficiently as INP, and it is therefore not surprising that they
25 are found in the supermicrom IPR. More surprising is the large fraction of OC and BC/OC
26 particles detected in the supermicron IPR. These large particles containing organic material
27 appear to be even better INP than the supermicron particles assigned to the Bio/Min/Sal
28 category. It has recently been suggested that porous and glassy organic particles under certain
29 conditions play an important role as INP (Murray et al., 2010; Adler et al., 2013), and this
30 may explain the results presented here. But also the origin of these supermicron organic
31 particles remains unclear. Primary biological material as a source for these particles is
32 regarded to be unlikely, because the marker peaks observed in cations of biological particles
33 (like K^+ , Na^+ , Ca^{2+}) were not observed in these particle spectra assigned to the OC and BC/OC

1 types. However, particle classification of single particle mass spectra is never absolutely
2 certain, and future work may reveal more insight into the nature and origin of these particles.

3 The particle composition of the particle smaller than 1000 nm closely resembled the
4 composition of the whole data set, because the particle smaller than 1000 nm dominate the
5 analyzed particles by number. Only the BC/OC-type is increased in the particles smaller than
6 1000 nm measured while using the Ice-CVI.

7 The size distribution allow also for an estimation of the ice active site (e.g., Vali et al., 2014)
8 density of particles larger than 1 μm compared to those smaller than 1 μm : Assuming that
9 each IPR corresponds to at least one ice active site on the original aerosol particle (otherwise
10 it would not have acted as an INP), we can estimate the ice active site density for the particles
11 analyzed by ALABAMA, separated for particles smaller than 1 μm and larger than 1 μm . The
12 number of ice active sites per IPR surface (ice active site density) is $2.0 \cdot 10^{-7} \text{ nm}^{-2}$ for
13 $d > 1 \mu\text{m}$ and $1.6 \cdot 10^{-6} \text{ nm}^{-2}$ for particles with $d < 1 \mu\text{m}$. Relating the number of ice active sites
14 to the total aerosol surface yields $7.1 \cdot 10^{-10} \text{ nm}^{-2}$ for $d > 1 \mu\text{m}$ and $4.3 \cdot 10^{-9} \text{ nm}^{-2}$ for $d < 1 \mu\text{m}$.
15 In both cases the ice active site density is larger for particles smaller than 1 μm . This is an
16 important finding and explains the high number of IPR found in the size range below 1 μm .
17 However, it should be noted that the assumption that each IPR corresponds to one ice active
18 site may underestimate the number of ice active sites. In cases where more than one ice active
19 site is found on the particles, also only one ice crystal will form and one IPR will remain.
20 Thus the estimated number yields a lower limit for the ice active sites.

22 4 Summary and Conclusions

23 This paper presents results of the single particle measurements during the INUIT-JFJ/CLACE
24 2013 campaign conducted at the Jungfraujoch in January/February 2013. In agreement with
25 previous measurements at the Jungfraujoch (Chou et al., 2011; Kamphus et al., 2010; Mertes
26 et al., 2007), it was found that larger particles are more efficient ice nuclei. Specific analysis
27 of only those particles $> 1 \mu\text{m}$ showed that the chemical composition of these larger particles
28 is dominated by organic material and the "Bio/Min/Sal" type, both in the IPR as well as in the
29 background aerosol. Besides, the "minerals" group was found to be one of the main fractions
30 of the IPR measured behind the Ice-CVI. From the finding that IPR have a higher relative
31 number of larger particles than the background aerosol particles we conclude and confirm a
32 better ice nucleation ability for the larger, primary aerosol particles such as minerals, salt and
33 biological particles.

1 In general, IPR are enriched in organic material, partly also internally mixed with BC.
2 Additionally, minerals and lead were found in IPR as has already been observed in previous
3 measurements conducted at the JFJ (Mertes et al., 2007; Cozic et al., 2008b; Cziczo et al.,
4 2009; Kamphus et al., 2010; Ebert et al., 2011). In comparison to the composition of the IPR
5 measured using the Ice-CVI, the measured composition using the ISI is dominated by
6 minerals, and no lead containing particles were found, while 7 % of the IPR sampled using
7 the Ice-CVI included lead. In contrast to Cozic et al. (2008) we found no enrichment of pure
8 black carbon in the IPR composition. Only the chemical composition of the background
9 aerosol particles shows a high fraction of black carbon. These results agree with previous
10 investigations (DeMott et al., 1999; Dymarska et al., 2006; Hoose and Möhler, 2012)
11 describing black carbon as a good ice nucleus but only at lower temperatures and its ice
12 nucleating ability appears to be reduced if coatings by organic material or sulfuric acid are
13 present. During the INUIT-JFJ/CLACE 2013 campaign the temperatures were never lower
14 than -30 °C; see Fig. 6. In the background aerosol, the main particle fraction is also organic
15 carbon. The high amount of organic material in the background aerosol as well as in the IPR
16 motivates further investigation of different organic material by single particle mass
17 spectrometry to better characterize the organic fraction. While there have been previous
18 observations that organic matter is of general importance for ice nucleation ability (DeMott et
19 al., 2003a; Cziczo et al., 2004b), recent laboratory studies have shown the importance of
20 porous and glassy organic matter for ice nucleation (Murray et al., 2010; Adler et al., 2013).
21 The investigation of the influence of the temperature and the origin and altitude of the air
22 masses on the chemical composition of IPR and aerosol particles indicate that organic
23 material dominates the IPR composition at higher temperatures and metal components from
24 mineral or salt at lower temperatures which agrees with previous findings (e.g. Möhler et al.,
25 2008; Hoose and Möhler, 2012). It was illustrated that the origin (geographic and altitude) of
26 the air masses influenced the chemical composition of the aerosol particles. For the non-cloud
27 aerosol, the particles sampled in air masses characterized by free tropospheric air and long
28 distance transport were found to be dominated by BC and single metal ions from salt, mineral
29 or biological sources (Bio/Min/Sal-type). These particle types were probably partly mixed
30 with secondary inorganic components, because NH_4^+ was detected in some spectra. In contrast
31 to that, the particles in air masses which are influenced by surface emissions are dominated by
32 organic carbon and potassium.

33

34 **Appendix A: Parameters underlying the clustering algorithm**

1 This section describes the cluster parameters of the software package CRISP. The chosen
2 parameters are listed in Table A.

3 *Normalization type and time*

4 To compensate for differences in the ionization efficiency, each mass spectrum is normalized
5 to the sum of the signal intensities with preservation of the relative relationship of the signals.
6 By this, the influence of absolute signal intensity is decreased. Absolute signal intensity is not
7 needed because the laser ablation method does not measure quantitatively. The point at which
8 the normalization is applied in the analysis sequence can be chosen. In case of measuring both
9 polarities the spectra of both polarities are first merged, and normalization can be applied
10 before, after or before and after concatenation.

11

12 *Preprocessing type*

13 Preprocessing (before normalization) can be applied to increase mass signals with lower
14 intensity with respect to signals with a higher intensity, such that the presence of a certain m/z
15 has more significance than the signal intensity in the calculation of the distance to the
16 reference spectrum. Typically the square root of the signal intensities is used.

17

18 *Initialisation type*

19 This parameter defines the way the reference mass spectra are selected. Two options are
20 available:

- 21 - *find different startcluster*: different spectra are randomly selected according to certain
22 criteria
- 23 - *random*: reference spectra are chosen randomly

24 Reference spectra can also be determined manually, however, this last option is not very
25 useful for field data because the number of particle types is unknown.

26

27 *Number of cluster*

28 The result of the clustering is highly dependent on the pre-selected number of clusters. In the
29 case that the number of clusters is too small, some cluster types cannot be found, while in the
30 case that the number is too large, too many similar spectra are assigned to different clusters.
31 In the fuzzy c-means algorithm all non-matched spectra are collected in a rest cluster.

32

33 *Cluster difference*

1 This parameter determines the distance between the spectra which are chosen as start
 2 references at 'find different startcluster'. The *cluster difference* is between 1 and 0, where a
 3 *cluster difference* of 1 means that the chosen startclusters are identically. This distance results
 4 from the Pearson correlation and affects the number of the resulting clusters. The number of
 5 clusters increases with increasing similarity of the references and is constant at a value of 0.7
 6 (Roth, 2014). If the *cluster difference* is too small, the distinctions between the clusters are too
 7 large, resulting in too few cluster types. Thus, it is recommended to use more similar
 8 reference mass spectra and to summarize those clusters that show similar mass spectra at the
 9 end of the clustering.

10

11 *Fuzzifier and fuzzy abort*

12 The *fuzzifier* (also called *fuzzy weighting* exponent) is a dimension of fuzziness of the
 13 classification of the cluster ($1 \leq \textit{fuzzifier} \leq \infty$; e.g. Huang et al., 2012). The bigger the
 14 *fuzzifier*, the more inaccurate is the assignment to the references, because the spectra cannot
 15 exactly be assigned to one cluster. The distance d_{ij} between spectra i and reference j is
 16 weighted with:

$$17 \quad d_{ij}^{\frac{1}{z-1}} \quad (1)$$

18 where z is the *fuzzifier*. When the preprocessing option of CRISP is used, it has been found
 19 that the result of the clustering is independent of the *fuzzifier* value (Roth, 2014).

20 During the clustering, each mass spectrum is compared to the reference spectra in terms of
 21 distance, membership and correlation. If these parameters fulfill the pre-selected criteria, the
 22 mass spectrum is sorted into the cluster represented by the reference and the average spectrum
 23 is calculated again. Subsequently all mass spectra are compared again with the reference
 24 spectra. This procedure is repeated until the criterion of abort (*fuzzy abort*) is reached, such
 25 that the clustering stops when the value of the difference of all reference spectra between two
 26 iterations is smaller than the value defined by *fuzzy abort*.

27 For the data evaluation of the current dataset (both IPR and background aerosol particles) a
 28 mixture of manual and automatic clustering was used. First, an automatic clustering was done
 29 with the following parameters:

30

31 Table A: Chosen clustering parameter.

Preprocessing type	Power each m/z
Preprocessing power	0.5
Normalization type	Sum
Cluster difference	0.7 (0.6)
Distance	Correlation
Fuzzifier	1.2
Fuzzy abort	0.0001

1
2 As already mentioned the result of the clustering is highly dependent on the chosen initial
3 number of clusters. If this initial number of clusters is too small, some cluster types may not
4 be found. Therefore we chose a high number of clusters. In the case of the background aerosol
5 (71064 spectra) we started the first clustering with 200 clusters. Each of the resulting 200
6 clusters was manually inspected, and based on its ion signatures it was assigned to one
7 particle type. Those spectra that cannot be sorted by the algorithm into one of the 200 clusters
8 (membership coefficient below a pre-defined threshold) are assigned to a “rest cluster”
9 (12407 spectra in the case of the background aerosol). This rest cluster was clustered again
10 with a reduced condition for the cluster difference and with an initial number of clusters of 25.
11 The number of cluster was still too large after the clustering of the “rest cluster”, so it was
12 further analyzed. First, the remaining spectra were sorted by the use of specific marker peaks
13 and the resulting clusters were then clustered again with changing the fuzzyfier from 1.2 to
14 1.3 and a cluster difference of 0.7. After that, the resulting clusters were manually inspected
15 and sorted (with the criterion $r^2 \geq 0.7$). By this, we could reduce the number of spectra in the
16 rest cluster from 12407 to 3243. Due to this sorting procedure based on both automatic
17 clustering and manual cluster inspection one particle type typically contains different clusters
18 with different fragmentations. For example, the particle type “organic carbon” includes
19 clusters showing aliphatic hydrocarbon, aromatic carbon, nitrogen and oxidized carbon
20 fragmentation patterns. In order to restrict the result to a reasonable number of particle types,
21 we summarized these clusters to one particle type.

22

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6

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13

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34

1 Table 1: Cloud events during which measurements were taken behind Ice-CVI or ISI, along with the ALABAMA sampling time, the number of the
 2 measured spectra, the evaluation of the Ice-CVI-sampling-periods and the weather conditions and cloud formation during these events. The times of
 3 each cloud event based ^{a)} on the particle counter operated behind the Ice-CVI or ^{b)} on the liquid water content measured by two PVM instruments
 4 and a CDP outside the laboratory.

Event	Sampling time (local time)	Inlet	Number of spectra	Classification of the Ice CVI- sampling- periods	Weather conditions & clouds
1	^{a)} 22.01.2013 01:48 – 12:02	CVI	174	++	Warm front occlusive low pressure system; $\bar{T} = -19.2\text{ °C}$; wind direction $w_d = 325^\circ$; wind speed $\bar{v} = 8.1\text{ ms}^{-1}$
2	^{a)} 27.01.2013 16:27 – 28.01.2013 09:44	CVI	78	++	Occlusive frontal system, Ns; $\bar{T} = -12.7\text{ °C}$; $w_d = 318^\circ$; $\bar{v} = 8.4\text{ ms}^{-1}$
3	^{a)} 1. 29.01.2013 04:30 – 13:57 2. 29.01.2013 19:27 – 19:52	CVI	20 3	++ ++	Unstable warm front; first Ns than Cb; $\bar{T} = -7.8\text{ °C}$; $w_d = 320^\circ$; $\bar{v} = 12.6\text{ ms}^{-1}$
4	^{a)} 30.01.2013 16:14 – 31.01.2013 01:03	CVI	39	++	Squall line at cold front; Cb; $\bar{T} = -8.6\text{ °C}$; $w_d = 323^\circ$; $\bar{v} = 16.9\text{ ms}^{-1}$
5	^{a)} 1. 01.02.2013 17:22 – 02.02.2013 01:30 2. 02.02.2013 01:30 – 13:07 3. 02.02.2013 17:58 – 03.02.2013 01:00 4. 03.02.2013 01:00 – 10:45	CVI	134 701 209 73	++ + + ++	31.01.: cellular convection; Cu; $\bar{T} = -12.5\text{ °C}$; $w_d = 324^\circ$; $\bar{v} = 17.7\text{ ms}^{-1}$ 01.02. – 02.02.: frontal wave; Ns and Cb; $\bar{T} = -13.7\text{ °C}$; $w_d = 319^\circ$; $\bar{v} = 13.4\text{ ms}^{-1}$ Until 03.02. cellular convection; Cb; $\bar{T} = -23.1\text{ °C}$; $w_d = 308^\circ$; $\bar{v} = 9.7\text{ ms}^{-1}$

	b) 1. 31.01.2013 18:55 – 21:42 2. 01.02.2013 01:53 – 11:05	ISI	9 18		
6	a) 1. 05.02.2013 19:06 – 23:00 2. 05.02.2013 23:00 – 06.02.2013 04:20 3. 06.02.2013 04:20 – 05:00	CVI	24 42 3	++ + ++	Squall line at postfrontal convergence line; Cb; $\bar{T} = -19.8^{\circ}\text{C}$; wd = 331°; $\bar{v} = 10.5 \text{ ms}^{-1}$
7	a) 1. 06.02.2013 08:33 – 11:00 2. 06.02.2013 11:00 – 12:30 3. 06.02.2013 12:30 – 14:56 4. 06.02.2013 15:44 – 15:53 5. 06.02.2013 23:32 – 07.02.2013 04:30 6. 07.02.2013 04:30 – 08:45 7. 07.02.2013 08:45 – 09:49 8. 07.02.2013 20:29 – 08.02.2013 05:50 9. 08.02.2013 05:50 – 09:20	CVI	26 4 15 3 8 6 2 31 2	+ ++ + + ++ + ++ + ++	06.02: Convective dominating low-pressure vortex; Cb; $\bar{T} = -24.3^{\circ}\text{C}$; wd = 310°; $\bar{v} = 8.0 \text{ ms}^{-1}$ From 07.02. on: cellular convection; Cu; $\bar{T} = -20.8^{\circ}\text{C}$; wd = 324°; $\bar{v} = 7.6 \text{ ms}^{-1}$
8	a) 1. 09.02.2013 01:09 – 02:30 2. 09.02.2013 02:30 – 09:00 3. 09.02.2013 09:00 – 09:41	CVI	4 32 3	++ + ++	Diamond dust and cellular convection at convergence line; Ci, Cu (Cb); $\bar{T} = -26.7^{\circ}\text{C}$; wd = 316°; $\bar{v} = 7.2 \text{ ms}^{-1}$
9	a) 10.02.2013 18:10 – 18:24	CVI	2	++	Approaching warm front; Ci, Cs, As; $\bar{T} = -15.7^{\circ}\text{C}$; wd = 324°; $\bar{v} = 9.5 \text{ ms}^{-1}$
10	a) 1. 11.02.2013 08:03 – 10:31 2. 11.02.2013 15:08 – 12.02.2013 06:11	CVI	1 11	++ ++	Warm front; Ns; $\bar{T} = -18.7^{\circ}\text{C}$; wd = 227°; $\bar{v} = 6.6 \text{ ms}^{-1}$

11	^{a)} 12.02.2013 16:07 – 13.02.2013 01:33	CVI	8	++	Back of a low with northeast stream; Ns, Ac, As; $\bar{T} = -20.1^{\circ}\text{C}$; wd = 319°; $\bar{v} = 7.9 \text{ ms}^{-1}$
12	^{b)} 15.02.2013 01:17 – 09:10	ISI	114		Low pressure system; As, Cu, Cb, Ac; $\bar{T} = -16.9^{\circ}\text{C}$; wd = 322°; $\bar{v} = 7.7 \text{ ms}^{-1}$
13	^{a)} 19.02.2013 16:46 – 20.02.2013 16:47	CVI	1	++	Cold front from North; Ac, Ns, Sc, Cu; $\bar{T} = -14.3^{\circ}\text{C}$; wd = 321°; $\bar{v} = 8.2 \text{ ms}^{-1}$
14	^{b)} 24.02.2013 06:57 – 24.02.2013 09:31	ISI	3		Approaching warm front with precipitation; Cs, As, Ns; $\bar{T} = -20.4^{\circ}\text{C}$, wd = 150°; $\bar{v} = 8.5 \text{ ms}^{-1}$

1

2 Ac: altocumulus

3 As: altostratus

4 Cb: cumulonimbus

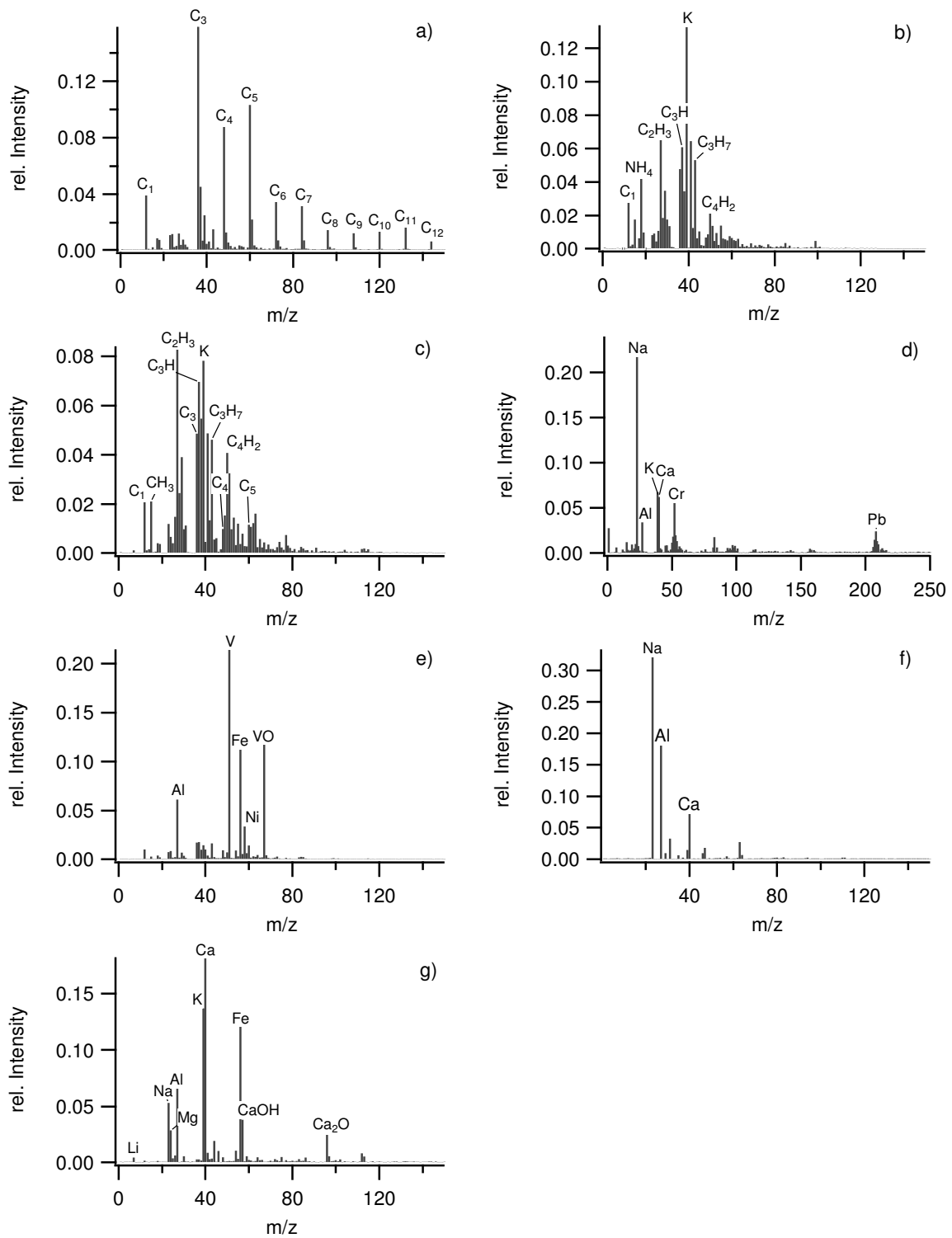
5 Ci: cirrus

6 Cs: cirrostratus

7 Cu: cumulus

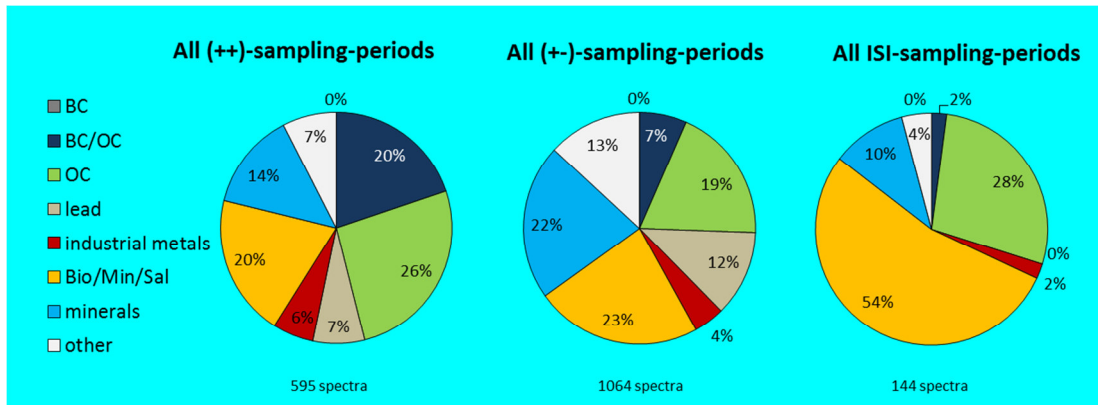
8 Ns: nimbostratus

9 Sc: stratocumulus



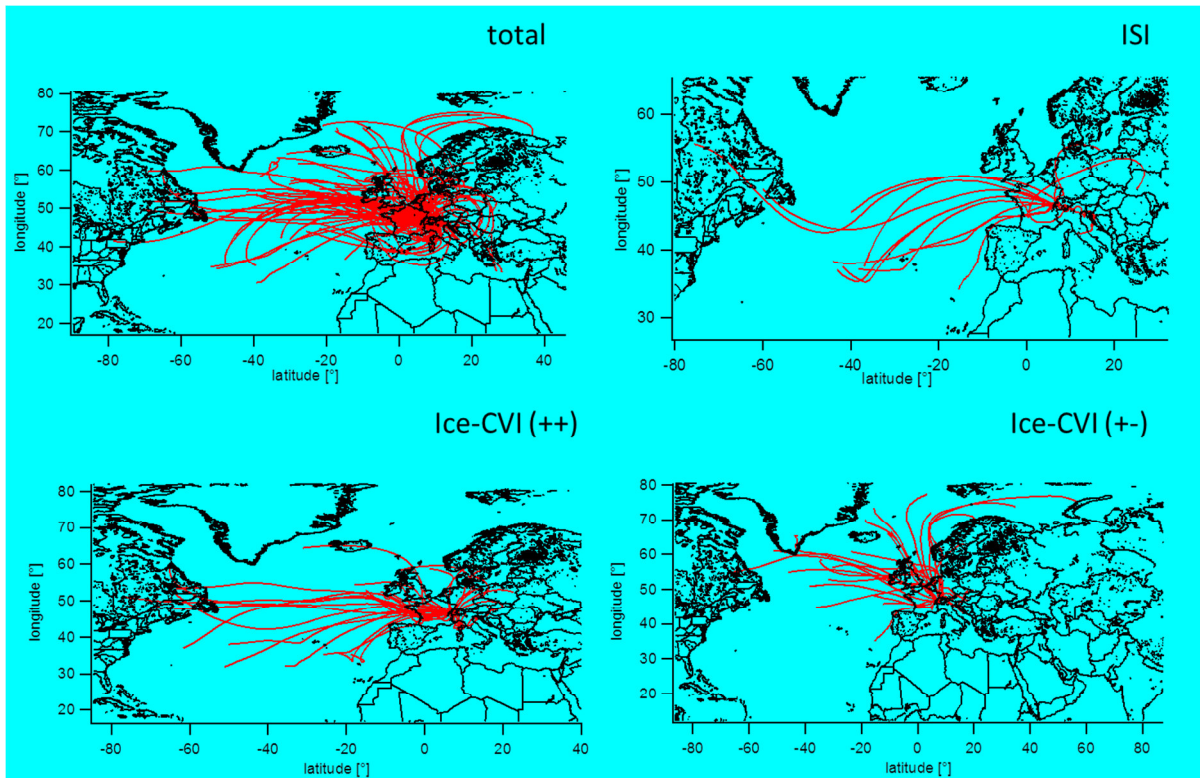
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Figure 1: Average mass spectra of each particle type. a) black carbon; b) organic carbon; c) black carbon internally mixed with organic carbon; d) lead containing particles; e) industrial metals; f) **Bio/Min/Sal**; g) minerals



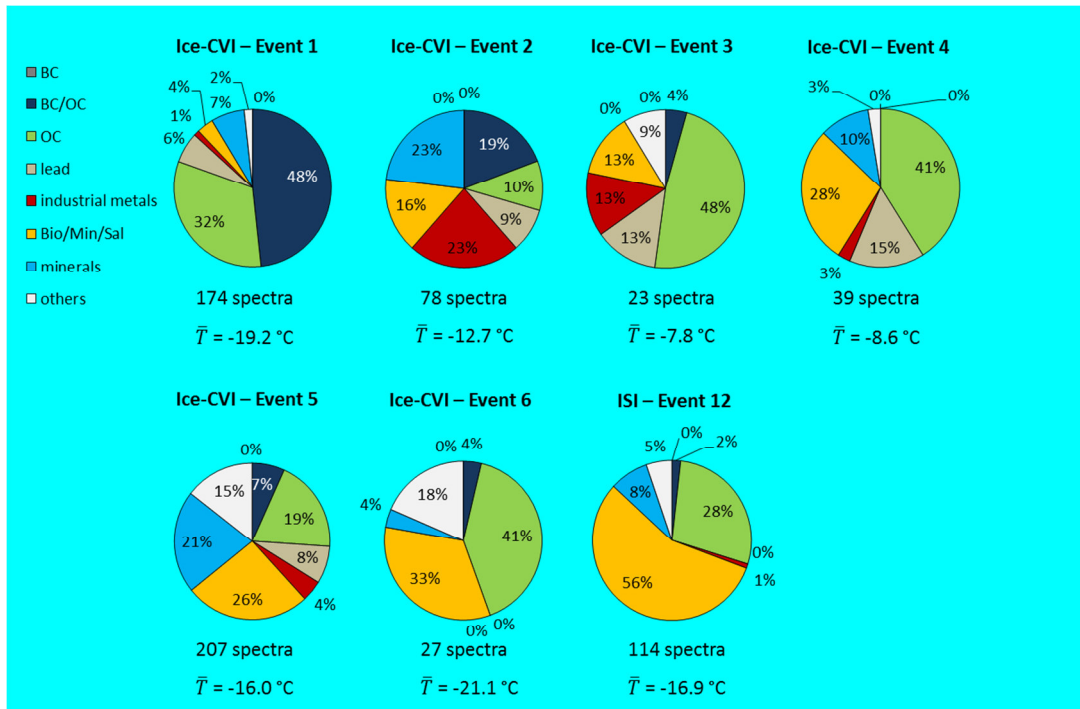
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Figure 2: Results of the IPR composition analysis. Left and middle: measurements behind the Ice-CVI (left: all (++)-sampling-periods; middle: all (+-)-sampling-periods). Right: IPR composition measured behind the ISI.



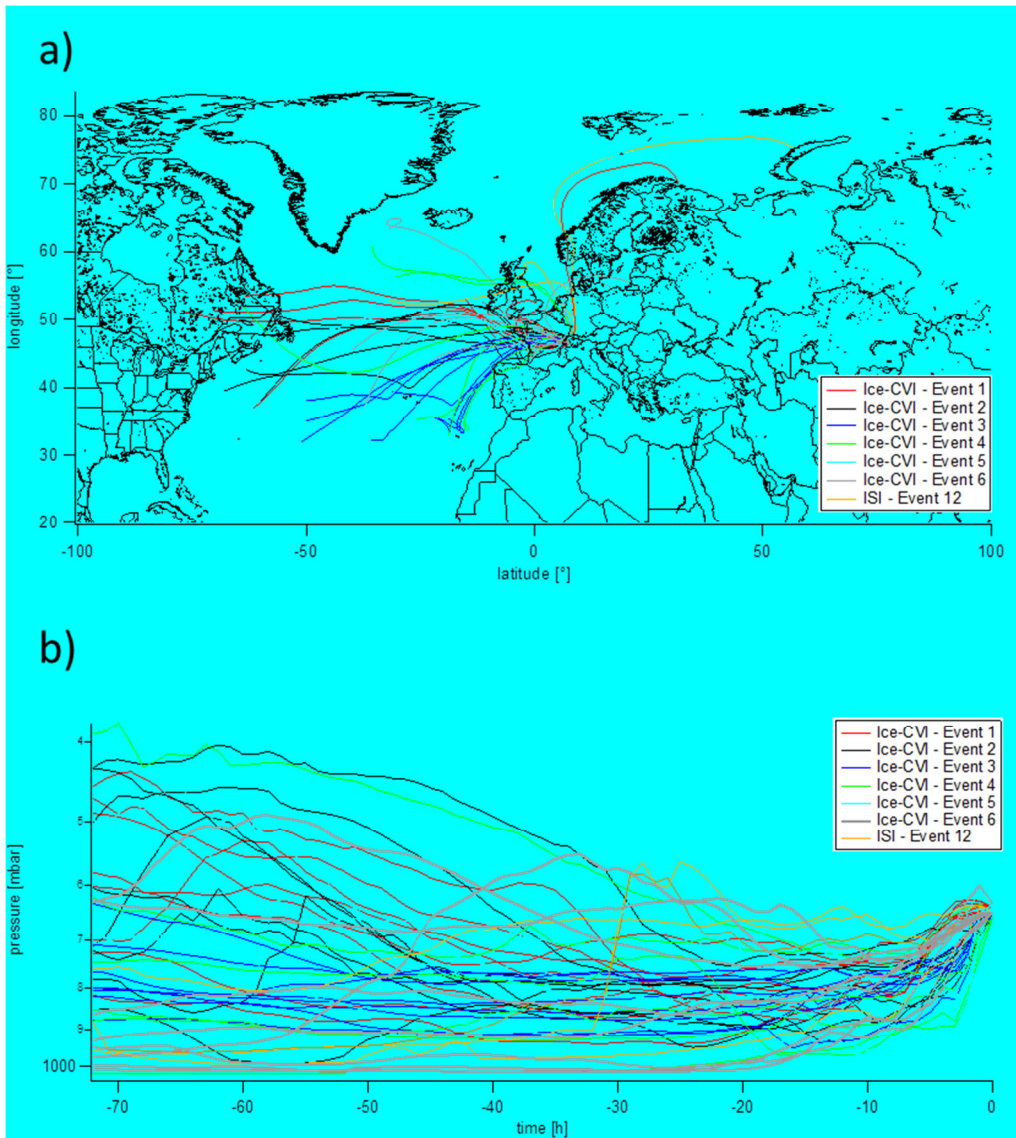
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Figure 3: 72 h backward trajectories calculated with CRISP/HYSPLIT (one trajectory every 6 h)



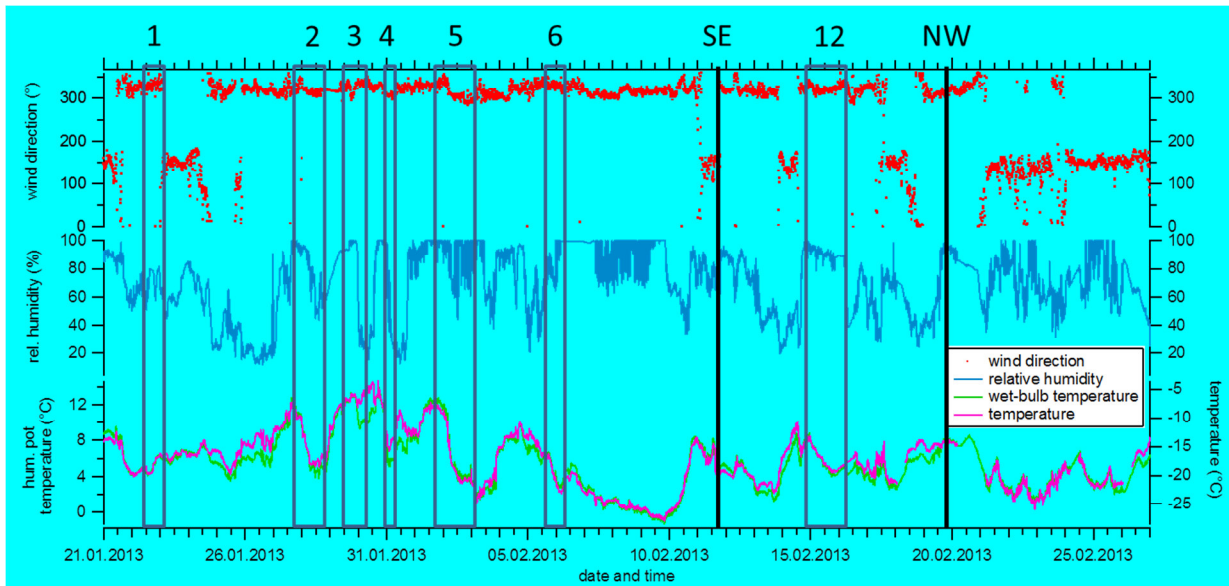
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Figure 4: Comparison of the composition analysis of seven events measuring behind the Ice-CVI (all (++)-sampling-periods, Event 1: 22.01.2013 01:48 – 12:02; Event 2: 27.01.2013 16:27 – 28.01.2013 09:44; Event 3: 3.1 29.01.2013 04:30 – 13:57; 3.2 29.01.2013 19:27 – 19:52; Event 4: 30.01.2013 16:14 – 31.01.2013 01:03; Event 5: 5.1 01.02.2013 17:22 – 02.02.2013 01:30; 5.4 03.02.2013 01:00 – 10:45 and Event 6: 6.1 05.02.2013 19:06 – 23:00; 6.3 06.02.2013 04:20 – 05:00) and the ISI (Event 12: 15.02.2013 01:17 – 09:10) together with the corresponding average temperature during each event.



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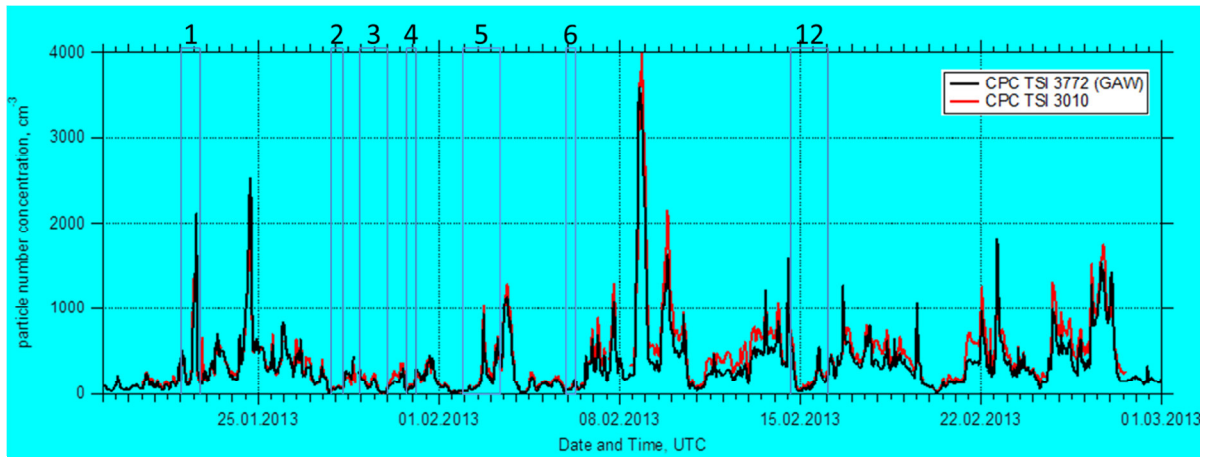
Figure 5: a) 72 h backward trajectories (one trajectory every three hours) calculated with CRISP for each event. b) The altitudes of air masses reaching the Jungfraujoeh as a function of time, converted to pressure units.



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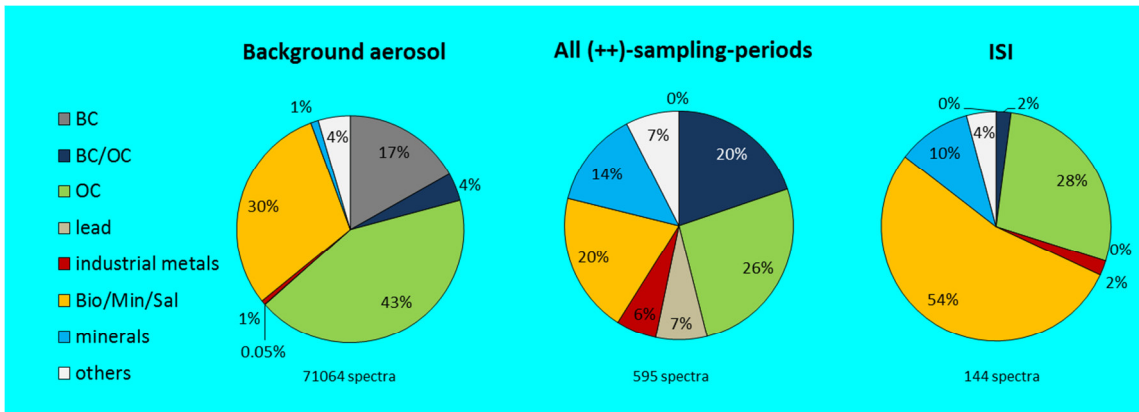
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3 Figure 6: Wind direction, relative humidity, potential wet-bulb temperature and temperature
 4 over the whole measurement period (data from MeteoSwiss at the JFJ station). Black lines
 5 denote the events with different air mass origin (discussed in section 3.4). Events highlighted
 6 with blue bars denote the 7 different events from Section 2.1.1.



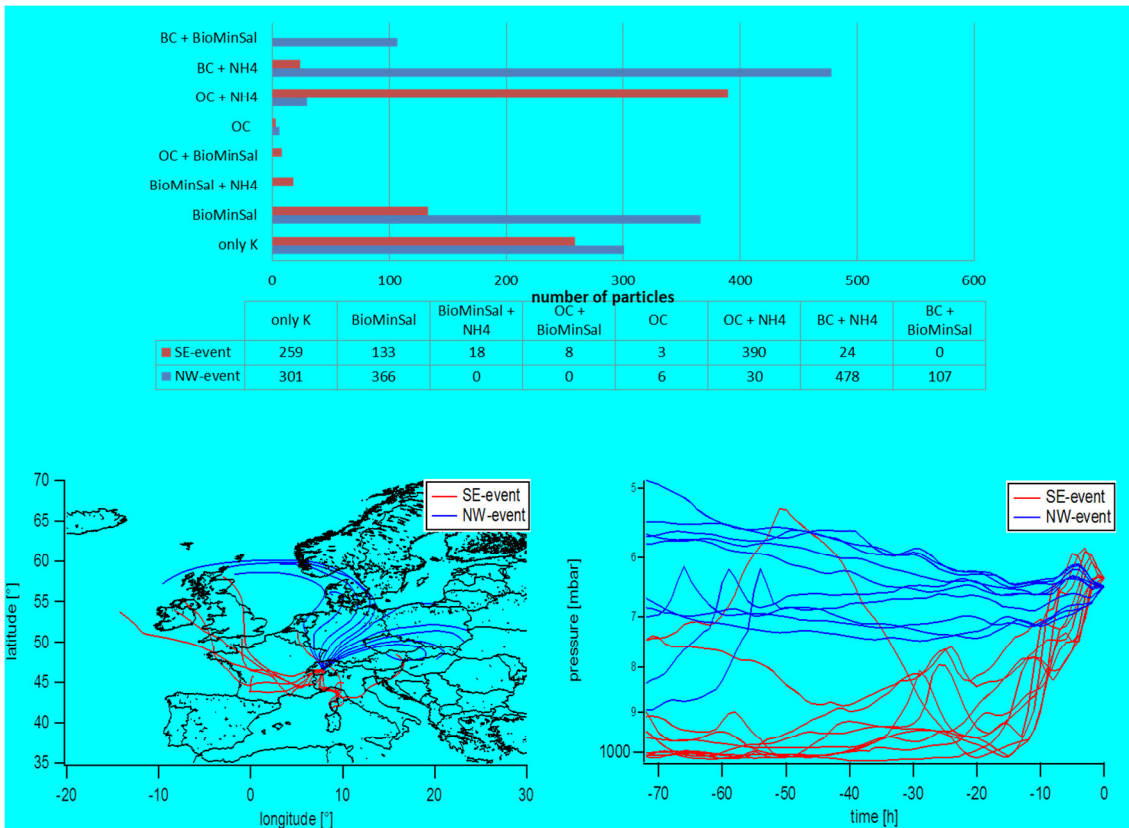
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Figure 7: Particle number concentrations measured during the INUIT-JFJ/CLACE 2013 campaign by two different condensation particle counters (CPC): The GAW-CPC which is located directly behind the total inlet, and a TSI 3010 that was operated with a much longer sampling line (approx. 7 m) next to one of the aerosol mass spectrometers. Short elevated concentrations indicate local emissions. Events highlighted with blue bars denote the 7 different events from Section 2.1.1



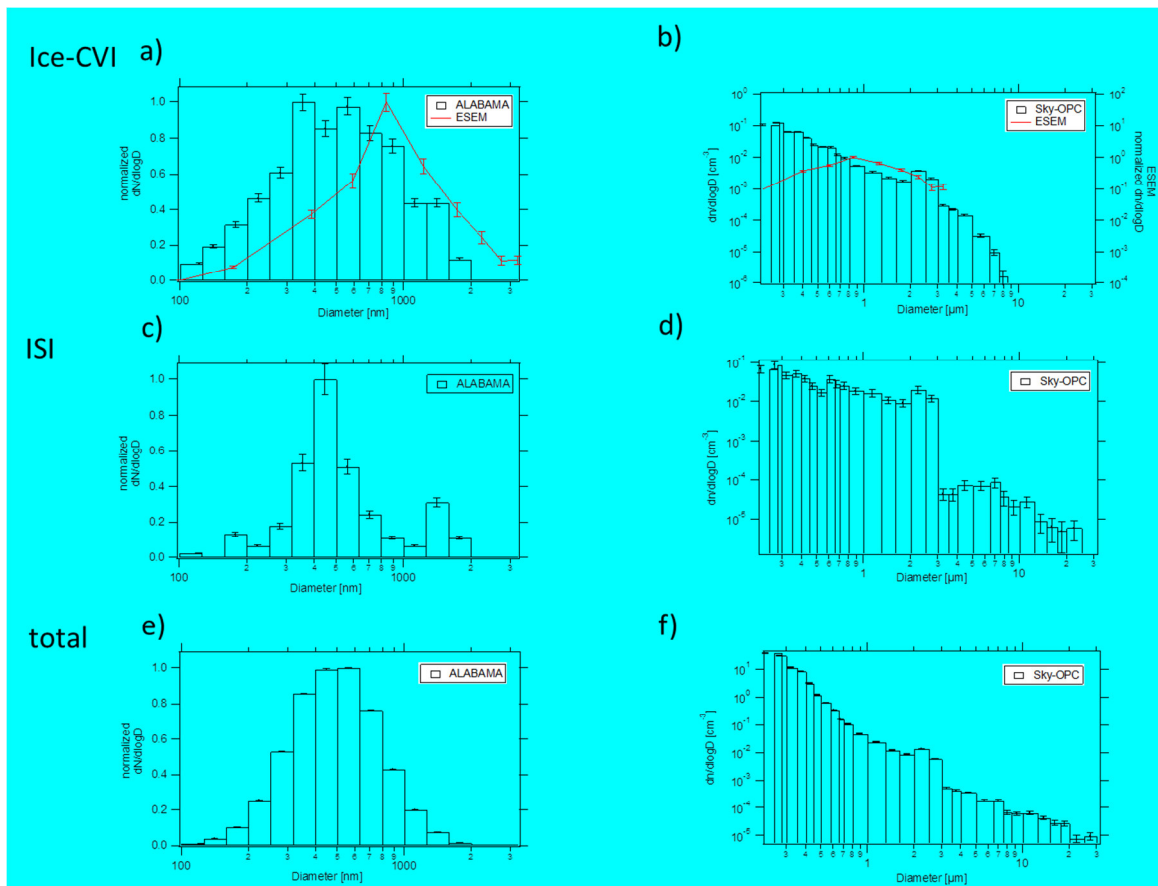
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Figure 8: Comparison between the chemical composition of background aerosol particles (left) and IPR measured behind Ice-CVI (middle) and ISI (right).



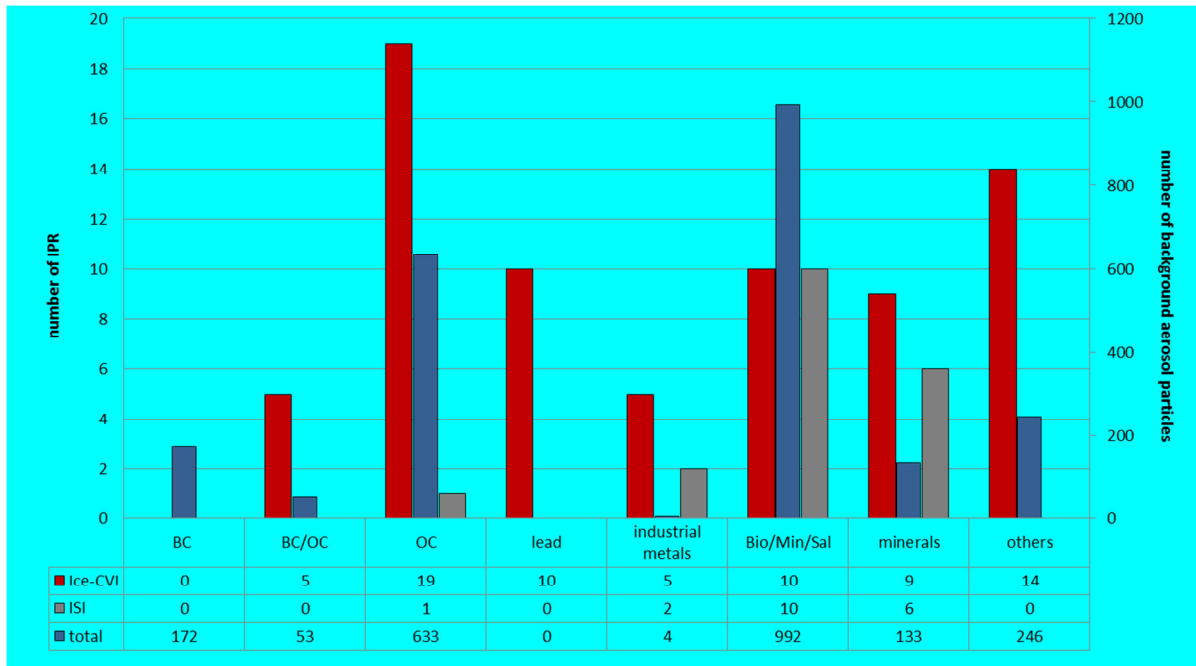
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Figure 9: Comparison of two events measured during non-cloud episodes. The northwestern event (NW; blue) was on 19.02.2013 14:00 – 16:14 and the southeastern event (SE; red) was on 11.02.2013 10:32 – 14:27. 72 h back trajectories were calculated with one trajectory every 2 hours.



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Figure 10: Measured size distribution of the IPR and background aerosol particles with the ALABAMA (bars; a, c, e) and Sky-OPC (bars; b, d, f) compared with the off-line SEM analysis of IPR samples (line; a and b). The error bars for the ALABAMA data result from counting statistics (averaged errors; $\Delta_{\text{Ice-CVI}} = 5.5\%$; $\Delta_{\text{ISI}} = 8.3\%$; $\Delta_{\text{total}} = 4\%$). The error of the OPC data results from Gaussian propagation of uncertainty, including counting statistics, the manufacturer-given error of the OPC of 3%, and the error of the enrichment factor (4% for the Ice-CVI and 20% for the ISI). The error of the ESEM dataset was determined by counting statistics. Note the different axis scaling. The y-axis is linear in a), c), e) and logarithmic in b), d), f). The x-axis range is 100 nm – 3250 nm in a), c), e) and 200 nm – 30 μm in b), d), f).



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Figure 11: Chemical composition of IPR and background aerosol particles larger than 1000 nm measured behind the Ice-CVI (red), ISI (grey) and total inlet (blue). The error bars results from counting statistics.