



Transport of Antarctic dehydrated air into the troposphere

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Transport of Antarctic stratospheric strongly dehydrated air into the troposphere observed during the HALO-ESMVal campaign 2012

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Abstract

Dehydration in the Antarctic winter stratosphere is a well-known phenomenon that is occasionally observed by balloon-borne and satellite measurements. However, in-situ measurements of dehydration in the Antarctic vortex are very rare. Here, we present detailed observations with the in-situ and GLORIA remote sensing instrument payload aboard the new German aircraft HALO. Strongly dehydrated air masses down to 1.6 ppmv of water vapor were observed as far north as 47° S and between 12 and 13 km in altitude, which has never been observed by satellites. The dehydration can be traced back to individual ice formation events, where ice crystals sedimented out and water vapor was irreversibly removed. Within these dehydrated stratospheric air masses, filaments of moister air reaching down to the tropopause are detected with the high resolution limb sounder, GLORIA. Furthermore, dehydrated air masses are observed with GLORIA in the Antarctic troposphere down to 7 km. With the help of a backward trajectory analysis, a tropospheric origin of the moist filaments in the vortex can be identified, while the dry air masses in the troposphere have stratospheric origins. The transport pathways of Antarctic stratosphere/troposphere exchange are investigated and the irrelevant role of the Antarctic thermal tropopause as a transport barrier is confirmed. Further, it is shown that the exchange process can be attributed to several successive Rossby wave events in combination with an isentropic interchange of air masses across the weak tropopause and subsequent subsidence due to radiative cooling. Once transported to the troposphere, air masses with stratospheric origin are able to reach near-surface levels within 1–2 months.

1 Introduction

Antarctic stratospheric dehydration occurs regularly every winter and spring in the very isolated southern hemispheric polar vortex (e.g. Vömel et al., 1995; Kelly et al., 1989; Schoeberl et al., 1992; Nedoluha et al., 2002; Jimenez et al., 2006). The reason is that

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and LOnG range; Krautstrunk and Giez, 2012). The ESMVal campaign was embedded in the TACTS (Transport And Composition in the Upper Troposphere/Lowermost Stratosphere; Engel et al., 2013) campaign which took place just before and after the ESMVal flights in August and the end of September 2012. A mutual scientific objective of both campaigns was to investigate the transition region between the troposphere and stratosphere. Here, we want to note that HALO was equipped with in-situ instruments as well as the high resolution limb sounder, GLORIA (Gimballed Limb Observer for Radiance Imaging of the Atmosphere). This combination of high-precision in-situ measurements together with sophisticated remote sensing observations in the UT/LS is an outstanding attribute of both campaigns. Another objective of the ESMVal flights was to get a full meridional cross section of atmospheric measurements for global chemistry model evaluation. Hence, the Antarctica flight was performed in an effort to get as far south as possible and reach the stratospheric vortex. During this flight, dehydrated air masses were measured in-situ quite far north up to 45° S between 12 and 13 km. In contrast, Aura-MLS and POAM 3 satellite measurements (Nedoluha et al., 2002; Jimenez et al., 2006; Schoeberl and Dessler, 2011) do not show dehydrated air masses as far north (not beyond 57° S) and as low in the stratosphere and upper troposphere as was observed during this ESMVal flight.

In general, the process of dehydration is well understood. Relatively less is known about the fate of the dehydrated air masses. The air within the Antarctic polar vortex is highly isolated with a weak exchange of trace gases across the vortex edge driven by stratospheric planetary Rossby waves propagating from the troposphere and Rossby wave breaking (RWB) events. However, these mainly disturb the bottom of the polar vortex. Normally, the vortex edge in the Southern Hemisphere is less strongly disturbed than in the Arctic due to less wave activity (Schoeberl et al., 1992). In the Arctic and in midlatitudes, intrusion of stratospheric air into the troposphere seems to occur more frequent and was discussed by Khosrawi et al. (2006) where at the edge of the deep intrusion dehydration was observed. Nevertheless, Antarctic stratospheric air masses can be vertically transported through the thermal tropopause directly into

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the troposphere and dry the troposphere down to the Earth's surface. In addition, this air is mostly rich in ozone and reactive nitrogen due to the stratospheric origin and can influence the chemical composition of the Antarctic troposphere (Stohl and Sodemann, 2010; Mihalikova and Kirkwood, 2013). Stohl and Sodemann (2010) identify two general processes that transport stratospheric air masses down to near-surface levels. The first process consists of katabatic winds over the Antarctic Plateau caused by the high topography that create a general downwelling above the Antarctic continent as reported by Roscoe (2004). The second process is driven by mid-latitude cyclones on the poleward side of the jet stream that support RWB events and a corresponding stratospheric intrusion as reported by Ndarana et al. (2012). Rossby wave induced stratospheric intrusions, such as tropopause folds, occur more often further north and in the midlatitudes than directly above the Antarctic continent (James et al., 2003). Once an air mass is in the troposphere, the mean cooling rates cause a reduction in potential temperature and the air mass will descend from the tropopause to near-surface heights within 10 days (descent rate of 5 mm s^{-1}) as reported by van de Berg et al. (2007). However, the frequency, the seasonality, and the process behind tropopause folds and stratospheric intrusions in the Antarctic region is still under debate (Stohl and Sodemann, 2010; Ndarana et al., 2012; Mihalikova and Kirkwood, 2013).

The study presented here is structured as follows: in Sect. 2, a brief overview of the different instruments and data used is given. The in-situ and remote sensing measurements across the Antarctic polar vortex are described in Sect. 3, where it is also shown that dehydration occurs directly in the transition region between the upper troposphere and lower stratosphere (UT/LS). Furthermore, Sect. 4 includes a case study of observed stratosphere/troposphere exchange of dehydrated air masses combined with an extensive trajectory analysis showing how deep below the thermal tropopause dehydrated air masses can be found.

2 Instrumentation and meteorological data

The new German research aircraft, HALO, deployed during TACTS and ESMVal has a long flight endurance of up to 12 h. This enables air masses in the Antarctic vortex to be sampled without landing in Antarctica. Altogether, a set of nine in-situ instruments for measuring trace gases and one remote sensing instrument were installed aboard HALO. For the study presented here, we use the water vapor data of FISH and HAI as well as measurements from TRIHOP for methane and FAIRO for ozone. The remote sensing instrument, GLORIA, provides cross-sections of trace gases that are fairly parallel to the flight track of the aircraft. GLORIA's high vertical resolution makes it particularly suited for the investigation of small-scale structures.

In addition to the aircraft measurements, satellite observations from CALIPSO and meteorological data from ECMWF are used for further interpretation of the observed situation. The instruments and meteorological data are described in the following subsections.

2.1 FISH – total water measurements (in-situ H₂O)

The airborne Lyman- α photofragment fluorescence hygrometer FISH (Fast In-situ Stratospheric hygrometer) is a well-established closed-path instrument for measuring water vapor in the range of 1 to 1000 ppmv (Zöger et al., 1999). The FISH is especially built to measure the low water vapor mixing ratios prevailing in the upper troposphere and lower stratosphere. It is regularly calibrated on the ground to a reference frost point hygrometer (MBW DP30) and had an accuracy of $6\% \pm 0.4$ ppmv during TACTS and ESMVal campaigns (Meyer et al., 2015). The air supply for the measuring cell is provided by a forward facing inlet, which also samples ice crystals if present (H₂O_{cond}). Together, the evaporated crystals and the gas-phase water (H₂O_{gas}) sum up to a total water measurement.

atmosphere). Thus, while the location of quantities retrieved at flight level lies in the vicinity of the aircraft, the locations of quantities at lower altitudes are several tens to hundreds of kilometers away. Concerning the water vapor product used in this paper, Ungermaann et al. (2014) showed that GLORIA water vapor at flight level agrees fairly well with the in-situ FISH measurements, within error bars. The deviations to FISH during the ESMVal/TACTS flights are mostly less than 0.4 ppmv.

2.6 CALIPSO (satellite remote sensing)

The Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observation (CALIPSO) satellite is one of five satellites in the NASA A-train constellation. CALIPSO completes 14.55 orbits per day with an inclination of 98.2° and thus delivers good coverage above the polar regions. Besides one wide field camera and an imaging infrared radiometer, the Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) is aboard CALIPSO. The lidar operates with two wavelengths (532, 1064 nm) with additional polarization-sensitivity, providing high-resolution vertical backscatter profiles of aerosols and clouds. In this study we use the CALIPSO Lidar Level 2 Polar Stratospheric Cloud (PSC) data product, which is described in Pitts et al. (2009) and Pitts et al. (2011). This data product provides a PSC composition scheme on a daily basis for all nighttime orbits with a resolution of 5 km horizontally by 180 m vertically.

2.7 ECMWF (meteorological data)

Global meteorological reanalysis ERA-Interim data (Dee et al., 2011) of the ECMWF (European Centre for Medium-Range Weather Forecasts) are used to facilitate the interpretation of the observations. The meteorological fields have a resolution of 1° × 1° with 60 vertical levels from the surface (1000 hPa) up to 0.1 hPa. Every 6 h a full global dataset is available. The trajectories used in Sect. 4 are based on the horizontal wind fields and diabatic heating rates; additional parameters like temperature, pressure etc. are added to the trajectories from the ERA-Interim data.

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crosses the air masses with high equivalent latitude in the stratosphere before and after the dive. The dive down to 3.5 km altitude was performed on the poleward side of the high equivalent latitude tropospheric air masses.

3.2 Observations

5 The time series of the in-situ measurements are shown in Fig. 2, where the measurements of the water vapor instruments are visible in the upper panel and the measurements of the tracers for defining stratospheric/vortex air masses are shown in the lower panel. After the ascent of HALO near Cape Town, the water vapor mixing ratio decreases to the typical stratospheric value of 4 ppmv, from 06:30 until 08:15 UTC. During this time, ozone and PV (i.e. the stratospheric tracers with low values in the troposphere) show an increase compared to the values before 06:30 UTC, while methane (tropospheric tracer with low values in the stratosphere) shows a slight decrease from 1790 to 1740 ppbv. Both indicates that the aircraft entered the stratosphere. Henceforth, methane and water vapor further decrease to 1650 ppbv and 3 ppmv respectively on the same flight level starting at 08:15 UTC, and simultaneously ozone and PV further increase. This is the time when HALO penetrated the Antarctic vortex (marked in blue in Fig. 2). After ascending to a higher flight level (08:35 UTC), water vapor decreases to slightly above 2 ppmv and stays there until the southernmost point was reached and the dive started at 11:00 UTC. This low water vapor in the time between 08:15 and 11:00 UTC coincides well with high equivalent latitude (see Fig. 1b). Subsequently, the strong increase in methane and decrease in ozone and PV indicate the penetration of the aircraft into the Antarctic troposphere at 11:00 UTC. After returning to the previous flight level (11:45 UTC), the vortex signatures with low values of water vapor and methane as well as high values of ozone and PV show up again. However, at this point (~ 12:30 UTC) the water vapor mixing ratio reaches the lowest value of 1.6 ± 0.5 ppmv together with a strong maximum in ozone and PV. The low water vapor below 4 and even below 2 ppmv indicates strongly dehydrated air masses.

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trieval product, which is shown in Fig. 3. We consider only the times from 08:00 to 14:00 UTC to focus on the air masses where GLORIA observed vortex air (the whole dataset is from 06:30 to 15:30 UTC). The dive is noticeable as the white area between 11:00 and 12:00 UTC, where GLORIA did not measure in order to prevent condensation of tropospheric water vapor on the cold instrument. The dry vortex air masses are clearly visible between 08:00 and 13:00 UTC at flight level (black solid line in Fig. 3), but also below, down to altitudes of 7 km. As mentioned in Sect. 2.5, the quantities retrieved from GLORIA are approximately placed on a parabola following the tangent points through the atmosphere (e.g. 250 km horizontal distance to flight path at 9 km altitude). The dry region just before the dive was measured by GLORIA in the westerly direction to the flight path on the way towards Antarctica, while the air masses after the dive are measured in the easterly direction on the way back to Cape Town. Thus, the dehydrated air masses below the thermal tropopause seem to cover a large region, having a dimension of at least 500 km horizontally at 9 km altitude.

The thermal tropopause is also derived from the GLORIA retrieved temperature profiles and is marked with black dots. Interestingly, low water vapor mixing ratios around and below 2 ppmv are observed beneath the thermal tropopause. Especially in the time range from 09:00 to 10:30 UTC and from 12:00 to 13:30 UTC, very dry air masses with water vapor mixing ratios of 2–3 ppmv can be found far below the thermal tropopause, down to 7 km. This indicates that dehydrated stratospheric air masses could have been transported through the thermal tropopause into the troposphere. The dynamic tropopause, which is between the -4 and -2 PVU isoline, is somewhat lower than the thermal tropopause at ~ 7 km in the time range from 12:00 to 13:30 UTC and indicates a proceeding stratospheric intrusion. In Sect. 4.2, we analyze this transport process with the help of air mass backward trajectories. In addition, no water vapor values less than 6 ppmv are found below the -4 PVU isoline in ECMWF data (not shown here). This shows, at least for this situation, that the transport process of dry stratospheric air masses into the troposphere is not captured by the ECMWF meteorological analysis (see Sect. 4).

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temperatures in lower regions causes them to evaporate. With each pass of a cold region, more and more water vapor is removed from the airmass. As stated by Nedoluha et al. (2002), air masses must undergo several periods with dehydration before reaching the 1 to 2 ppmv mixing ratio level, which is typically observed by satellites within the vortex in the Antarctic spring. This periodic behavior is confirmed by Fig. 4, where trajectories undergo four events with low saturation mixing ratios and simultaneous ice cloud observations by CALIPSO. The first and the second period (1 and 9 August) reveal rather high saturation mixing ratios of around 3 ppmv. In every later period, colder temperatures and lower saturation mixing ratios are necessary to initiate additional ice formation and dehydration. So the third and the fourth period on 16 and 29 August where CALIPSO observed ice show lower saturation mixing ratios of 1.5 and 1.2 ppmv, respectively. The last two events will have dried out the air masses to the final state of about 1.6 ppmv, which is observed by the FISH and HAI instruments. After this strong dehydration, the temperatures became warmer in the late Antarctic spring and summer and no additional dehydration took place according to the simulations.

The recurring low saturation mixing ratios and corresponding low temperatures are primarily caused by gravity waves, which are induced by the high topography of the Antarctic continent (Stohl and Sodemann, 2010). Especially, the first three ice formation events occurred just as the trajectories passed the Antarctic Peninsula (northernmost part), which is a strong source of gravity waves (e.g. Ern et al., 2011; Hoffmann et al., 2013). Only the last freezing event with the lowest supersaturation occurs above the Antarctic Plateau (central part) in the north-east of the continent.

After freezing, the ice particles can even reach the troposphere due to sedimentation, especially if dehydration takes place at low altitudes (11.5–12.5 km) as observed during the Antarctic ESMVal flight with HALO. The lack of rehydration signatures in the GLORIA measurements (see Fig. 3) also indicates the permanent removal of water vapor from the stratosphere, as will be further discussed in Sect. 4.2. The air masses from which the ice particles originate and the air masses where they sublimate behave very differently in dynamical perspective. Due to the length of time since the last

the measured thermal tropopause (black dots). Here, the fundamental role of the thermal tropopause as a vertical transport barrier for stratosphere/troposphere exchange emerges.

In contrast, the air masses that spent between 5 and 40 days in the vortex (marked by light blue and reddish colors) indicate mixing of outside air into the vortex. Even at the vortex edge and in the core of the vortex itself, small filaments are apparent. As indicated in Sect. 3.2, some filaments with elevated amounts of water vapor were also observed with GLORIA. Two of these filaments observed around 13:00 UTC between 10 to 12 km altitude can be found with this Nash PV criterion showing vortex residence times of 5 days (filament 6) and 35 days (filament 7), respectively. This shows that these filaments are likely freshly mixed in from outside the vortex and thus could potentially contain higher water vapor content compared to the dehydrated vortex air and therefore weaken the hypothesis of rehydration signatures.

To show that all observed filaments (red boxes in Fig. 3) are caused by mixing, reverse domain filling (RDF) is applied to all trajectories, as already stated in Sect. 3.2 (similar as in Beuermann et al., 2002). For the RDF calculation, only advection and no mixing along the trajectories is assumed. In each time step of the RDF method, the ECMWF water vapor field is interpolated onto the locations of the trajectories and then projected onto the observed time-altitude grid of the GLORIA observations.

Figure 6 shows ECMWF water vapor mixing ratios five days prior to the measurement time as projected by the trajectories. The red boxes highlight the position of the water vapor filaments observed by GLORIA (see Fig. 3). With some small deviations in time and altitude, almost all of the observed filaments in the vortex can be reproduced by the RDF method. Note, that the RDF method cannot produce the effect of rehydration. This indicates that the water vapor filaments observed by GLORIA were not generated by rehydration, but are likely the signatures of moister air masses that were mixed into the vortex during the preceding 5 days. Continuing the RDF method further back in time, the filament 7 (≈ 12 km) at 13:00 UTC emerges 35 days before observation and indicates the origin to be another mixing event. The time of in-mixing

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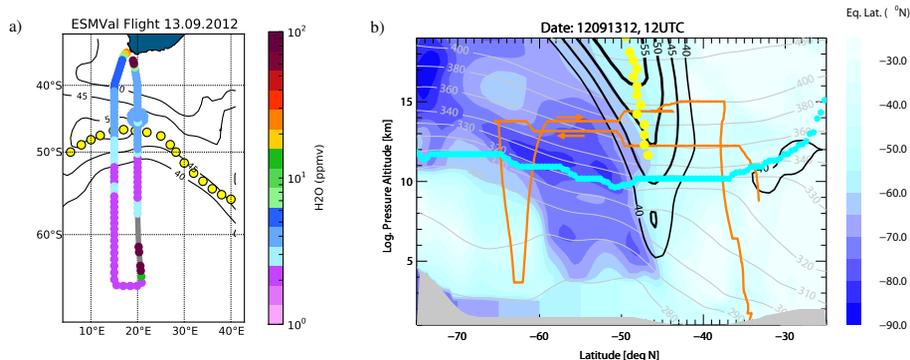


Figure 1. Flight pattern of the ESMVal flight on 13 September 2012. The black contours illustrate the horizontal westerly wind from ECMWF data. Yellow dots represent the vortex edge derived from the Nash criterion (Nash et al., 1996) based on ECMWF data. **(a)** Horizontal map of water vapor mixing ratios measured by FISH (5 min averaged data) is color coded on the flight path (gray color indicate data gaps due to the dive). **(b)** Meridional cross-section of equivalent latitude along the flight path (orange line) calculated from ECMWF data. Blue dots represent the ECMWF thermal tropopause.

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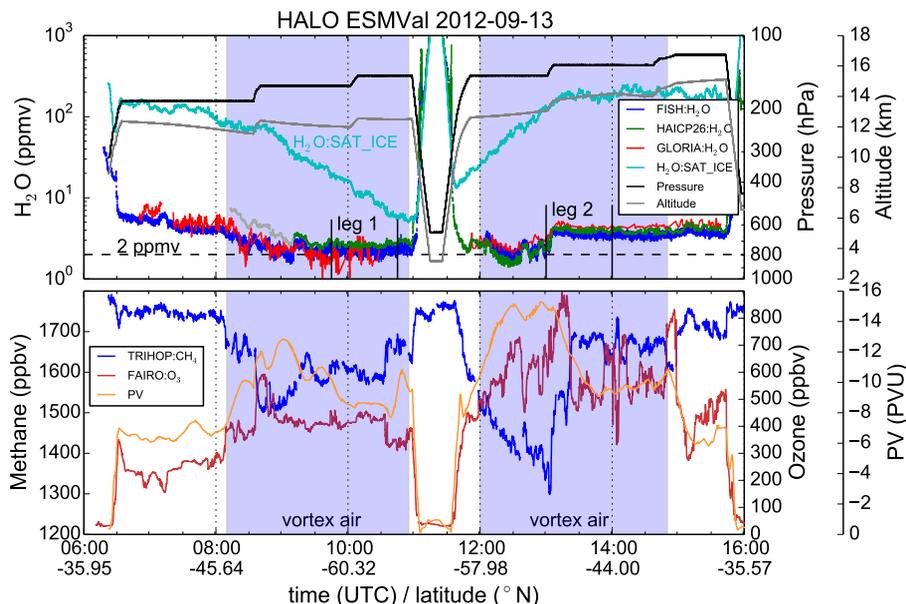


Figure 2. Timeseries of ESMVal Antarctica flight on 13 September 2012: the upper panel shows FISH (blue), HAI (green) and GLORIA (red) water vapor measurements, water vapor saturation mixing ratio with respect to ice (cyan, derived from HALO temperature), pressure (black), and altitude (gray) at flight level. The lower panel shows the tracer observations ozone (red), and methane (blue) and interpolated ECMWF PV values (orange). Time ranges within the vortex are marked with bluish shadows according to the tracer measurements. Time ranges leg 1 and leg 2 are marked for hygrometer intercomparison of FISH and HAI (see text).

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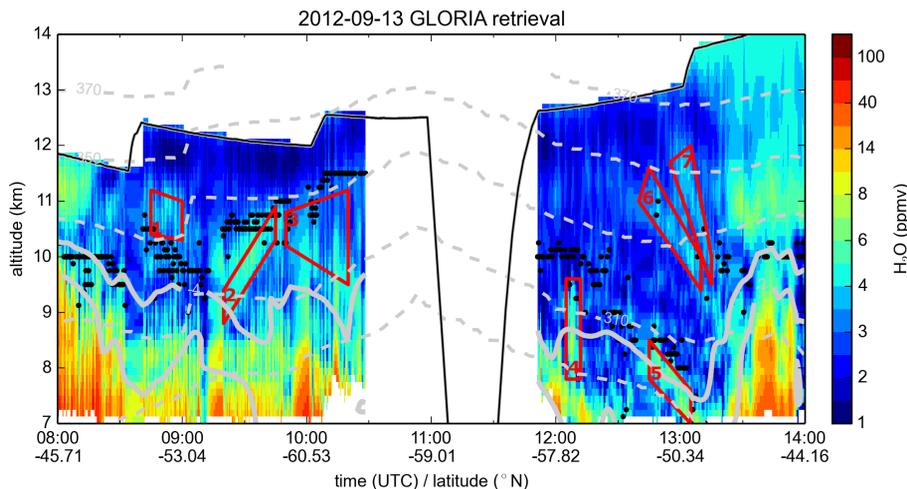


Figure 3. GLORIA time series of the H_2O retrieval product (13 September 2012). The solid black line and the black dots represent the flight path and thermal tropopause derived from the GLORIA temperature measurements, respectively. The dashed gray lines and the solid lines marks isolines of potential temperature (290–370 K, $\Delta\theta = 20$ K) and PV (−2 and −4 PVU), respectively. The red boxes marks filaments (1–7) with slightly enhanced water vapor mixing ratios in or below the vortex.

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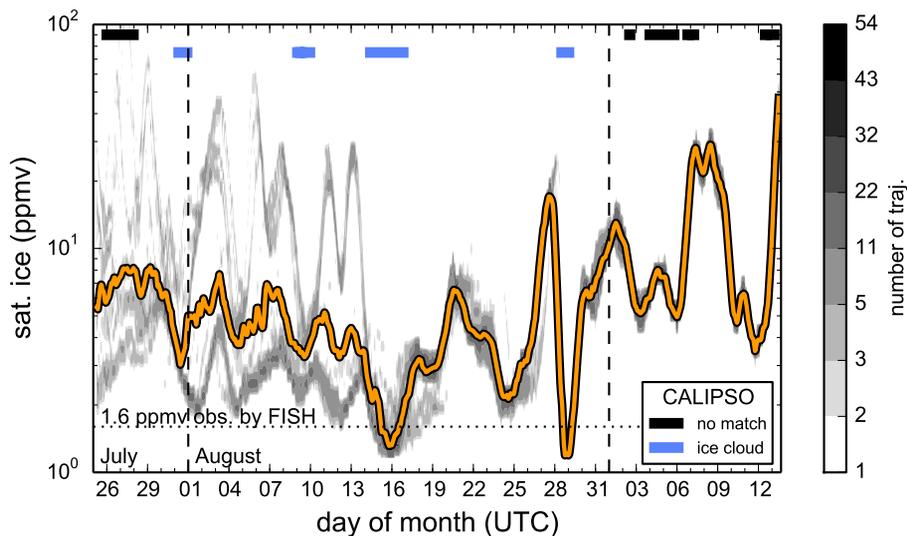


Figure 4. Calculated frequency distribution (number, gray shadow) and median (orange line) of ice saturation mixing ratio along trajectories 50 days back in time starting at the flight path on 13 September 2012 (12:20 to 12:29 UTC, 54 trajectories in total). If more than 50% of trajectories have a corresponding CALIPSO observations of ice or no ice, the time is marked with blue or white, respectively. Otherwise, it is marked in black, indicating no CALIPSO match.

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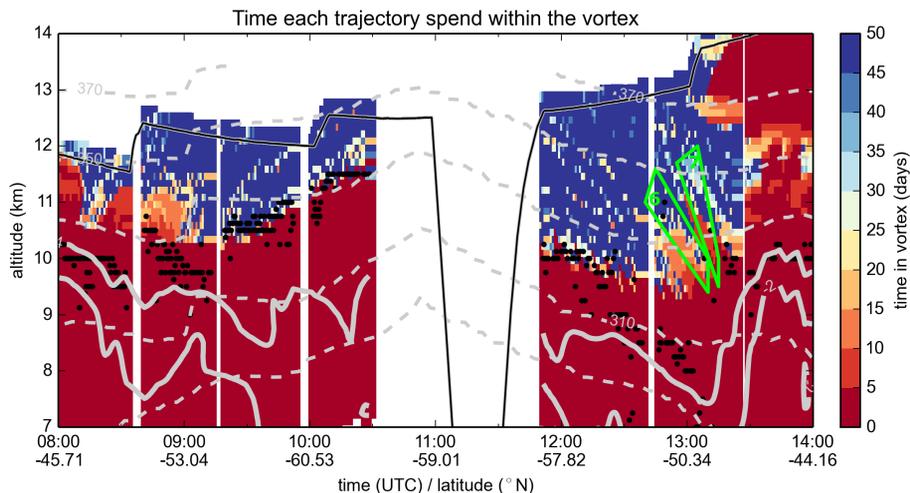


Figure 5. The time each GLORIA tangent point trajectory has spent in the vortex before the measurement. The green boxes mark two of the filaments (6, 7) observed by GLORIA (see Fig. 3). The flight path, tropopause height, potential temperature, and PV isolines are the same as in Fig. 3.

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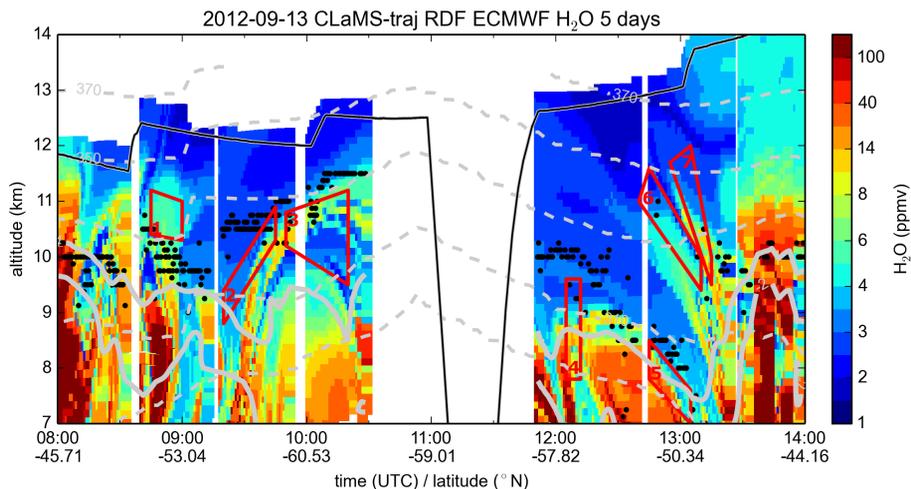


Figure 6. RDF (Reverse Domain Filling) of each GLORIA tangent point trajectory with ECMWF water vapor of 5 days prior the measurement time. Red boxes show the location of the observed filaments. The flight path, tropopause height, potential temperature, and PV isolines are the same as in Fig. 3.

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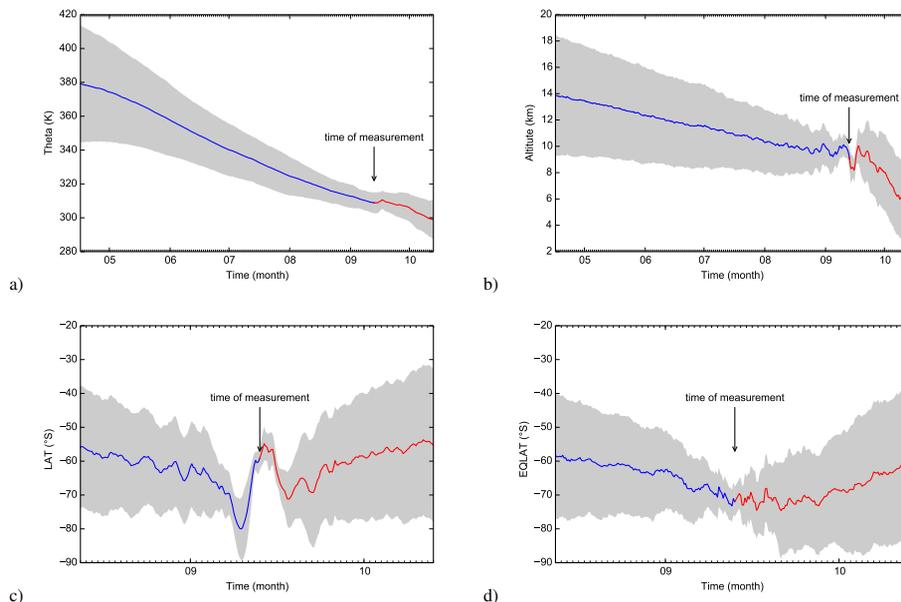


Figure 8. Air parcel properties based on trajectory analyses for all green framed trajectories from Fig. 7 for the period of 5 months before (blue lines) to 1 month after the measurement (red lines): **(a)** median of potential temperature (θ), **(b)** median altitude above mean sea level, **(c)** median latitude from 1 month before to 1 month after observation, **(d)** median equivalent latitude from 1 month before to 1 month after observation. The gray shaded areas marks the SD of the 1400 trajectories.

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