Anonymous Referee #1

This paper describes the construction of an important and comprehensive long-term dataset for key trace gases in the stratosphere. It is important to have a peer-reviewed and citable description of datasets like this GOZCARDS dataset. The manuscript serves this purpose very well, and I support publication in ACP.

My main criticism concerns the length of the manuscript. Over 100 journal pages, almost 30 figures, plus appendix and supplement are a tall order. I wish the authors would have made the effort to be more concise. At this point it is probably difficult to make the paper much shorter, but I have some suggestions:
Response: This main criticism has been addressed by significantly reducing the size of the paper, following most of the reviewer comments/suggestions; we certainly appreciate all the reviewer efforts regarding suggestions to help address the length issue.

We decided to present the merging methodology in this paper but also included some original analyses (after having released the data for community analyses) in order to make it more relevant for ACP. We also include useful details such as data screening procedures, not always described well enough in past papers, in order to assist anyone potentially wanting to duplicate this dataset (and for completeness) – but now relegate most of this to the Appendix. We also believe that ACP papers (and this paper in particular) should not deal primarily with methodology - so we were not inclined to remove everything of interest regarding early sample results using GOZCARDS. One can get criticized the other way sometimes and get suggestions to try publishing in AMT instead, if one deals only or mainly with methodology (or techniques or validation). While this paper is still somewhat long because of the dual goal (methodology and some results + the description of a multi-species effort), we feel that at this point, a 23% reduction in the main text (and our cutting out 3.5 Figures out of the original 27) without eliminating most newer discussions/results is a reasonable compromise.

1 Major comments

GOZCARDS HCl sample results: Given recently published results, e.g. Mahieu et al. 2014, I feel that it is not necessary to discuss the HCl variations at various levels in such detail. I think Figures 7 and 8 already present all the relevant information. Figures 9 and 10, to me, are not needed (a look at Fig. 7 gives almost the same information). The discussion on pages 5865 last paragraph to 5868 first paragraph is also not necessary. I think it provides no really new or essential information, and is not necessary for the main topic of the paper, which is description of the overall dataset. Moving the important citations to the early part of section 3.3 would be enough, in my opinion.
Response: We have made significant cuts here. However, showing that the HCl behavior is turning around in the later years is useful and new (and in our view, informative for ACP audience, and as a brief update to the Mahieu et al. work, which mainly looked at columns - and stopped in 2011). We believe that it is most useful to show (keep) the 2 left panels of Fig. 10 to show the non-linearity and connection to surface total chlorine changes. So we took out two thirds of Figs. 9 & 10 (combined) [we are just left with the new Fig. 9] and shortened the text significantly (by about 45% - see the new discussion before Sect. 4).

GOZCARDS H_2O sample results, pages 5873 to 5876: This reads almost like a review paper. Again, I think many readers will be familiar with the overarching issues, and would appreciate a briefer and more concise discussion. Figures 15 and 17 already give a very good description of long and shorter term water vapour variations, even without much text. Is Figure 16 really needed? Are all the (confusing and overall similar) thin and dashed lines needed in a dataset description paper?

Response: Agreed, we have significantly shortened this H_2O discussion (by almost 40% for section 4.3) and we simplified Fig. 16 (now with no thin lines), without taking it out altogether, as one can refer to this as a useful diagnostic for testing model variability.

GOZCARDS Ozone sample results: Since GOZCARDS is primarily a stratospheric ozone profile dataset, I wonder if the long comparison and discussion with the Ziemke and Chandra 2012 column dataset is really that helpful. As the authors show, this dataset seems to definitely have a problem, and generally TOMS and OMI are not recommended for long-term trend analyses any more. It seems to me that this part of the current paper dates back a few years. There are several more recent analyses of ozone profile trends (Tummon et al. ACPD 2014, Harris et al. ACPD 2015, WMO 2014). I feel it would be much better to tie in the GOZCARDS data with these publications, and drop the comparison with the Ziemke and Chandra column data. At the same time, the discussion could be shortened, and number of figures in this part could be reduced.

Response: We think that cutting out the column ozone discussion completely would remove a significant analysis as we have not really seen published work on this subject, despite what some may view or characterize as "common knowledge" regarding TOMS/OMI column data time series analyses. This ozone discussion may "date back" a bit, but it is new evidence nevertheless (many papers discuss "old data"), and the ozone topic is the primary reason to tie this paper to the SI^{2}N special issue (on ozone only). The reviewer states "as the authors show", and we did feel we needed to show something on this topic, as we did not see such a mention (to the ZC12 work) in the WMO Report, for example. Unless specific published references already exist to clearly make our point and/or to imply that other column series are superior to TOMS/OMI, we plan to retain the flavor of this ozone-related "sample result". For brevity, however, we remove Fig. 23 and Fig, 25 and retain only Fig. 24 (Fig. 22 in the update); this preserves our point regarding comparison issues for GOZCARDS stratospheric column relative changes versus those provided by ZC12, but more succinctly. Overall, the Sect. 5.3 text has been cut by 37%.

Also, adding something really new to the ozone profile trends discussions at this point would likely require substantial efforts and a separate manuscript, given the assessments and discussions provided for GOZCARDS ozone profiles in WMO (2014) and in several other references – as mentioned in our ozone section. We think that briefly referencing/highlighting recent trend studies using GOZCARDS is sufficient, as there are a lot of details in those papers.

Summary and conclusions could also be shortened. There, and in several other places, it seems to me that the manuscript resulted from adding newer references to an older previous manuscript version, and discussing them later / separately. E.g. Section 5.5.2 about diurnal ozone variations seems like a tag-on coming from more recent interest and publications on this topic. The reference to Mahieu et al. 2014 paper on page 5893 restarts the discussion again, making the paragraph unnecessarily long.

Response: Agreed. We have shortened this summary significantly; the whole summary section was cut by 31%.

The water vapour summary on pages 5894 and 5895 is also very long. It should be shortened and focused.
Response: Agreed. We have shortened this summary significantly; the whole summary section was cut by 31%.

Overall, the manuscript would benefit greatly from a thorough attempt at removing redundacies and shortening plus focusing on the main topics: Data set generation and sample results.
Response: Agreed. We have shortened the main text significantly (by 23% - and deleted 3.5 Figures), also following another set of referee comments – but keeping in mind our introductory response (sample results are viewed as a useful contribution to ACP).

2 Minor comments
page 5851, lines 15 to 17: I find this sentence confusing and hard to understand. Maybe change to: In the upper stratosphere, however, SAGE II is not used after June 2000, when the (NCEP temperature-dependent) data conversion from a density/altitude to a mixing ratio/pressure grid exhibits a shift in the Southern Hemisphere. Instead HALOE data are used as a transfer reference from SAGE II before 2000 to the datasets starting in 2004.
Response: Agreed, but this portion and the whole Abstract were shortened (per other referee comments), so the upper stratospheric details are now left out of the Abstract.

page 5852, line 15: Here, and in most places, WMO 2011 should be supplemented by WMO 2014, or replaced by it. WMO 2014 was published in December 2014.
Response: Agreed; in places, as appropriate, we have added the WMO (2014) reference (although in the first mention listed here, we removed the WMO reference for brevity).

page 5854, line 6: Probably good to add "converted to mixing ratio versus pressure using their enclosed NCEP temperature profiles, and" after "were"
Response: Agreed; we essentially used the above additional text.

page 5854, line 8: Add a typical number of MLS profiles per day.
Response: Agreed, done (about 3500 profiles).
By the way, also related to the next paragraph: How is the standard error of the mean calculated in Figure A1(b)? From the Appendix and the Figure caption I get the impression that it is calculated by standard deviation divided by the square root of the number of profiles. This assumes that all the profiles are uncorrelated. This may be true for the few and far spread occultation profiles, but for the many close emission profiles it is probably not true.
Response: Yes, this is how it is calculated (standard definition); the correlation issue is not trivial to accurately estimate, and the main result (many more data points from emission measurements) still would lead to these relatively much smaller errors (probably within a factor of two or better). Correlation occurs for some MLS bands (radiances) more than others but mainly along-track; in particular, the MLS measurement calibration (which can introduce some correlation) occurs on a much shorter timescale (a few minutes/profiles) than the typical profile

separation in a monthly 10° zonal average, which mixes in many uncorrelated data points from separate orbits.

page 5854, lines 15, 16: cross-sections (UV, visible, near IR, far IR, microwave) could also be an issue.
Response: Agreed, but this is implicitly included in the retrievals and their assumptions; if we were to produce all the potential reasons, this list would be quite long (e.g., see the MLS validation/characterization papers). Thus, and for brevity, we did not try to list all error sources.

page 5855, 1st paragraph: Maybe add WMO 2014 as well?
Response: Agreed, done.

page 5856: I am missing a mention of SAGE screening. Maybe mention the discussion in the Appendix here.
Response: The main data screening procedures have now been collected/summarized in Table A1 (Appendix A). We found a few species-specific details to be more appropriate for the species sections, but Table A1 is mentioned in these sections also.

page 5860+5861: I find this whole discussion lengthy and difficult to understand. Could it not be shortened? Is it not simply like this: First AURA-MLS and ACE-FTS are shifted to have the same average level in 2004 and 2005. This level is equal to their combined average level. Then HALOE (which has fewer data in 2004 and 2005) is also shifted to this level.
Response: No, this is not the way we adjust the third dataset and this will not provide the same answer as our procedure, as we wish to give equal weight to each of the three datasets (also note that the number of points changes with latitude and HALOE can often have more points than ACE-FTS, especially in the tropics). Shifting the 3rd dataset to the average of the first two is not the same as weighting them all three equally to get an average reference that depends on all three series. We describe our iterative procedure, which does use this equal weighting philosophy (since the first 2 series give a temporary merged result, then weighted by 2/3 with the 3rd series weighted by 1/3 for the last adjustments). Since a simplification was sought here, text cuts (by close to 30%) were implemented for the 1st part of Sect. 3.2 (Fig. 1 description of the HCl and H_2O merging process), with little reduction in the main message. However, an oversimplification would be dangerous, for what is an essential description of the process. The Appendix has a mathematical explanation that should help, and this process can be carefully tested on sample datasets (and we have done this - also when considering the simplifications offered by the reviewers). The iterative process is what we decided upon as the best path, and it does take a few sentences to explain this. Note that this explanation paragraph is not a major portion of the full paper (or even the HCl section), but we did shorten this section. We also added some clarification words in the caption to Figure 1, which we hope can help.

page 5865, line 23: With the addition of the Mahieu et al. 2014 references, and one or two additional sentences, section 3.3 could end right here. The rest of the section, until page 5868 line 15, is not needed, in my opinion.
Response: We made some significant cuts here. However, showing that the HCl behavior is turning around in the later years is useful and new (and in our view, informative for ACP audience, and as a brief update to the Mahieu et al. work, which mainly looked at columns - and

stopped in 2011). We believe that it is most useful to show (keep) the 2 left panels of Fig. 10 to show the non-linearity and connection to surface total chlorine changes. Thus, we took out two thirds of Figs. 9 and 10 (combined) and shortened the text accordingly; we are left with the new Fig. 9.

page 5868, around line 25: Is it certain that these low H_2O values are wrong? Is there a reference for that? Could H_2O from the gas phase have gone into Pinatubo aerosols, which consist of a sulfuric acid + water mixture? I have not done the numbers, so maybe this just shows my ignorance.
Response: The water vapor averages are lower than usual and other tropical measurements from UARS MLS (for 22 hPa) do not typically dip lower than 3 ppmv (including 3 std. devs. from the monthly means, which are above 3.5 ppmv). The Pinatubo-related stratospheric SO_2 amounts are quite small compared to the 0.5 to 1 ppmv needed for these water vapor differences in 1992. In particular, Read et al. (GRL, June 1993) indicate that enhanced lower stratospheric SO_2 profile values (measured by UARS MLS) were only of order 10 ppbv in Sep. 1991. Thus, conversion to sulfuric acid aerosols would not be expected to deplete water vapor substantially on a large scale, even if shortly after the eruption it may have had some (more localized) impact. However, we added the following sentence: "While this method may exclude some good data points, the lowest values (< 3 ppmv) do get screened out; such outliers are not corroborated by 22 hPa UARS MLS data (with most values > 3 ppmv).

page 5873, line 15 to page 5876, line 6: This is very lengthy, half a water vapour review paper. I think it should be shortened substantially. A few key messages, no literature review.
Response: Agreed, we eliminated most references and shortened this significantly (section 4.3 is now almost 40% shorter), also in response to comments from another referee.

page 5877, lines 1 to 7: I am missing clarity here. There are two aspects: 1.) SAGE retrieval needs density profiles. For V6.2 these came from NCEP operational, for V7 these came from MERRA. The effect of the different sources on SAGE ozone (number density vs. altitude) is very small 2.) Conversion to mixing ratio vs. pressure needs temperature/ pressure profiles. There, a large artefact arises in the SH upper stratosphere when NCEP is used instead of MERRA. If GOZCARDS would use MERRA here (and not the NCEP that comes with SAGE V6.2) the problem would be much smaller. I think this needs to be made clearer.
Response: Agreed, we have reworded the relevant sentences (before Sect. 5.1.1).

page 5878, line 20: Give some references for SAGE-HALOE comparisons, e.g. E. Remsberg et al., and others.
Response: While the data versions in the past have often differed slightly, we do not know or find which Remsberg reference the referee might be thinking of; but we did add the SPARC (1998 Report) reference, providing overall agreement with our own results (shown in the Supplementary material).

page 5879, line 27: I don't think SAGE ozone drifts. What drifts is NCEP / the conversion to mixing ratio versus pressure. Please correct.
Response: Agreed, we clarified this further by rearranging the sentences and rewording (see discussion of Fig. 17, before Sect. 5.2).

page 5881, line 17 to page 5887 line 21: Two very long subsections (review papers) that present lots of references, but very little really new information. Please shorten substantially.
Response: Agreed, we shortened this substantially (also per comment of another referee). These combined two ozone sections are now shorter by about 28% (5.2.2 is shorter by nearly 20% and 5.3 by 37%). Also, two ozone Figures were deleted from the "sample results" section (5.3).

Anonymous Referee #2

This manuscript describes the GOZCARDS data set comprised of several key stratospheric constituents. The goal of GOZCARDS is to combine data from multiple satellites into a consistent record suitable for long-term analysis. This paper describes the methodology used in constructing the data set, as well as provides analysis examples for each of the primary constituents. Continued long-term stratospheric time series analysis depends on our ability to combine data from multiple instruments, and as such documentation and validation of the methods used to construct this data set are very important to the research community. The manuscript is appropriate for publication in ACP with some notable revisions.

General Comments:

Overall the analysis and approach is good. However, I fear the length of the manuscript (27 figures, plus appendix and supplemental material) makes the paper in its current form very difficult to digest. I feel some of the analysis confuses the main point of the paper, which is the methods used to construct the data set, and the consistency between the source data sets used by GOZCARDS which directly affects the uncertainty of the final product. While I understand shortening the manuscript at this point is easier said than done, I suggest the authors consider whether some of the detailed analysis in particular can be removed, and possibly put into a separate manuscript. Though I'm not sure it would help significantly with length, in several sections the manuscript could be written more concisely. I will make a few suggestions in my specific comments, but general suggestions include writing in present active tense as much as possible, limiting qualitative descriptors, especially when the actual results follow in parentheses, and using tables as much as possible.

Response: This main criticism has been addressed by significantly reducing the size of the paper, following most of the reviewer comments/suggestions). However, we note that this paper is really the culmination of a 5-year effort to produce the datasets and to provide some results and overview of the possibilities, which are enhanced as well by other recent (or upcoming) papers analyzing ozone or HCl variations, for example. The paper has been shortened, largely following the suggestions from the reviewers and we certainly appreciate all the reviewer efforts regarding suggestions to help address the length issue.

We decided to present the merging methodology in this paper but also included some original analyses (<u>after</u> having released the data for community analyses) in order to make it more relevant for ACP. We also include useful details such as data screening procedures, not always described well enough in past papers, in order to assist anyone potentially wanting to duplicate this dataset (and for completeness) – but now relegate most of this to the Appendix. We also believe that ACP papers (and this paper in particular) should not deal primarily with methodology - so we were not inclined to remove everything of interest regarding early sample results using GOZCARDS. One can get criticized the other way sometimes and get suggestions to try publishing in AMT instead, if one deals only or mainly with methodology (or techniques or validation). While this paper is still somewhat long because of the dual goal (methodology and some results + the description of a multi-species / multi-year effort), we feel that at this point, a

23% reduction in the main text (and our cutting out 3.5 Figures out of the original 27) without eliminating most newer discussions/results is a reasonable compromise.

For example, Tables 1 and 2 are very informative, and easy to refer back to as needed. Can the general data screening information be put into a table? It is convenient to have all the filtering information in one place, but there are many details that are covered in the relevant publications that are not needed in the narrative. The authors might put the filtering information in a table, sticking just to the actual filtering and not including all the details as to what each filter is doing, for example: UARS MLS precision > 0, MMAF_STAT = "G", "t", "T", Quality=4, per Livesey et al (2003).
Conversely, if the authors do want to describe what each filter means (precision > 0 ensures only a negligible contribution to the retrieval from a priori, as an example), they might consider putting it in the Appendix as the SAGE filtering is done. Either way the information is presented in one place, but not in the main portion of the paper. There are other bits of information that seem to be repeated multiple times, such as the ACE_FTS data processing problems in 2010, that can be noted in the tables or in the data provenance information, rather than repeatedly in the text.
Response: Agreed, we have put the general screening in a Table (Table A1) in the Appendix but have kept some small discussion in the main text for the less well documented screening issues.

The description of the merging process seems overly complicated. For HCl and H_2O, if I understand correctly, the reference level is the average between mean HALOE values from August 2004-Nov 2005 (all that exist), mean ACE-FTS values over the same time period (all that exist), and mean AURA MLS values, but only averaged over the months that ACE-FTS has data. However, since the figures and discussion are keyed to a multi-step process, I do not recommend any large changes, just look for opportunities to simplify the text when possible. Also in the wording, keep in mind that the multi-instrument mean used as reference is not necessarily the correct answer, it is just a common reference. If AURA MLS were perfect, for example, it would still be adjusted in this procedure. In later discussion when referring to data sources that are biased low or high, I believe it is important to clearly state that they are biased high or low relative to the reference level.
Response: Since a simplification was sought here, text cuts (by close to 30%) were implemented for the 1st part of Sect. 3.2 (Fig. 1 description of the HCl and H_2O merging process), with little reduction in the main message. However, an oversimplification would be dangerous, for what is an essential description of the process. The Appendix has a mathematical explanation that should help, and this process can be carefully tested on sample datasets (we have done this - also when considering incorrect simplifications offered by the reviewers). The iterative process is what we decided upon as the best path, and it does take a few sentences to explain this. Note that this explanation paragraph is not a major portion of the full paper (or even the HCl section), but we did shorten this section. We also added some clarification words in the caption to Figure 1, which we hope can help.

Specific Comments/Wording Suggestions (many are minor and not mandatory)

p5850. General in abstract, don't need abbreviations unless you are using the abbreviation multiple times in the abstract (i.e. don't need MERRA and ESDR).

Response: Agreed, we deleted the terms "ESDR" and "MERRA" from the Abstract.

L9-L10: I'm not sure about the word debiased, it implies the full bias is removed, but in fact only the relative bias is removed. What about "To merge ozone records we first adjust each time series by its mean offset with respect to SAGE II during periods of measurement overlap to remove relative bias, and then average. For other species we use a multi-instrument mean computed during overlap periods as the reference."

Response: Agree to the "relative to SAGE II" modification (simple), but not on the 2nd simplification, which uses more words but does not state accurately enough what is done (we prefer to leave the details for the main text rather than oversimplify). Also, our slight revision is shorter than the above suggestion.

L17-18: and lower stratospheric Aura MLS data

Response: Agreed, done.

L20: On 6-8 year time scales [in general work to consolidate space by limiting the use of parentheses].

Response: Agreed, done.

L25-p5851 L2: this could be shortened: "Short-term tendencies of lower stratospheric and column HCl vary, with increases from 2005-2010 at northern mid-latitudes and deep tropics, but decreases after 2011 in northern mid-latitudes and increases at southern mid-latitudes."

Response: Agreed, done (even slightly more succinctly).

p5851 L17 I think you can end the sentence at reference and the meaning is clear [saving little bits of space]. L17: remove next sentence, this is confusing and not needed in the abstract.

Response: Agreed, done.

p5852 L10: of independent ground-based data (as in analysis of independent...)

Response: Agreed, done.

L19-20: A slow recovery of the ozone layer towards pre-1985 levels is expected...

Response: Agreed, done.

L22: High quality ozone and other constituent datasets are needed to document past variability and to constrain... [suggestion to reduce word count].

Response: Agreed, done.

p5853 L8: here is an example of how one might consolidate the words a bit: ... on a common latitude/pressure grid, using the following high quality satellite limb sounding instruments:[move list starting on line 15, and rest of paragraph, up.] Then change last sentence to "All source data sets still have shortcomings or imperfections, but in creating the GOZCARDS Earth Systems Data Record (ESDR) we maintain the integrity of the original data and do not arbitrarily

disregard data, nor do we attempt to fill in spatial or temporal gaps in the record." The sentence on ways to fill in data is beyond scope of paper and is not needed.
Response: Agreed, done.

P5854 L7: see Sect. 5.1.1 L8: profiles in two narrow latitude bands every day.
Response: Agreed, done.
L11-14: consolidate here [you've just introduced the source files and grid, so no need to refer back to these ideas]: Next we merge the GOZCARDS source data files by computing average biases between different source files during periods of overlap, and then adjusting each source file to a common reference to remove the relative biases. Non-zero biases ...
Response: Agreed, consolidation was done.
L16: such as differences in radiance measurement system or wavelength range, the retrieval algorithm, or in the vertical, temporal or spatial resolution. [no need to refer to Level 1 or Level 2, especially if you have to then define what you mean]
Response: Agreed, done.
L17: remove sentence starting "A useful reference" and start next sentence as Toohey et al. (2013) studied ...
Response: Agreed, done.
L21: '... vs. averages of output sampled at sub-orbital track locations'
Response: Agreed, done.

P5855 L4-5: remove clause "these analyses ... discussed later"
Response: Agreed, done.
L8: SI2N initiative (....) [remove 'which stands for'] L18: replace 'main' with 'source' [since source data sets already defined]
Response: Agreed, done.
L19-20: Section 6 briefly mentions GOZCARDS N2O, HNO3 and temperatures...
Response: Agreed, done.

P5856-5857 L5: As mentioned before, the filtering done according to various data teams can be summarized either in a table or in the appendix. **The special filtering for ACE-FTS is appropriate for the text body.** Also there is no discussion of SAGE here, or reference to the appendix that I saw. I suggest moving this information out of the main text body, but if it is not moved, SAGE should be mentioned.
Response: Agreed; as mentioned earlier, most of the data screening is now described in Table A1 (Appendix).

P5857 ACE-FTS: this discussion was a little confusing, and could be simplified. "When analyzing ACE-FTS data we found it necessary to remove occasional large outlier values that

could significantly impact the monthly zonal means. Our outlier screening procedure ... for each year of data."
Response: Agreed, done.

L12: What do the authors mean by independent zonal means from the ACE-FTS team? Is this a data set that the ACE-FTS filtered using their own method and it compares well? I'm not clear on how the zonal means supplied by the team are independent.
Response: Yes, the ACE-FTS team used a slightly different method; we added a brief clarification for this sentence.
L17: remove 'in the tropics' because the limits are extended in all latitude bands L17: remove parentheses around good.
Response: Agreed, done, and with slightly more consolidation.
L21-22: such data sparseness can increase trend uncertainties, for example.[I don't follow the rest of this sentence, comparisons with AURA gave confidence in trend results from ACE-FTS?]
Response: To remove confusion and length, and as this is not critical information, we have largely cut this out.
L24: add this last filter to the table or appendix, as it is standard.
Response: Agreed, this is now in Table A1 (Appendix).
L28: have been set to -999 in the GOZCARDS files.
Response: Agreed, done.

Section 3.1: This section is a bit confusing. First extra screening is covered, then the band deterioration on MLS, then validation studies, then more screening, more on the band deterioration and then more validation studies. Suggest adding "Froidevaux et al. (2008) found anomalously high MLS HCl vs. aircraft data at 147 hPa at low latitudes; these values are not included in the GOZCARDS HCl data set." to the screening info after the HALOE aerosol screening. Then summary of validation work. Then 'Finally, AURA MLS HCl data are not recommended for trend analysis at pressures < 10 hPa (reference?) even if monthly ... measurements. Aura MLS switched to backup band 14 after primary HCl band 13 showed signs of rapid degradation in early 2006. For pressures 10 hPa and above, the long-term band 14 data are considered robust, but drift in band 14 HCl in the upper stratosphere led us to not include AURA MLS HCl upper stratospheric source data in the GOZCARDS record.
Response: Agreed, we have rearranged and shortened this section.

Section 3.2: [the first two sentences don't add new information, but it would be helpful in Table 2 to note that the AURA MLS is included only from 10 hPa and below] 'Although not included in the final merged product, we do use the 2004-2005 absolute Aura HCl measurements at pressures less than 10 hPa to compute the offsets for the ACE-FTS and HALOE source files in a consistent manner at all pressure levels. Figure 1 ...
Response: Agreed, done.

P5860 L14: remove phrase in () L18: remove "too"
Response: Agreed, done.
L29: the average of the months when data from both instruments exist (i.e. at ACE-FTS temporal sampling).
Response: Agreed, done.

P5861 L4: we weight L11: thus removing any relative bias
Response: Agreed, done.
L15-16: checked against independent data sets from two institutions? This needs more detail, or to be removed. Did the authors compare to two other merged data sets?
Response: No, this was an independent computational check within our team (among co-authors at different places), but in the interest of shortening the text, we have removed this detail.
L29: is biased low relative to the reference mean and needs to be increased by the offset value.
Response: Agreed, done.

P5862 L7: detailed examples of upper and lower stratospheric offsets
Response: Agreed, done.
L 18-20: This seems a general data information statement, and should be moved up to the ACE-FTS screening section, if it needs to be mentioned explicitly.
Response: Agreed, done.

P5863 L17: ... merged values is given by the range of available ...
Response: Agreed, done.
L18: remove sentence 'We have made such a calculation....'
Response: Agreed, done.
L23: in relation to the reference values
Response: We have changed to: "biased significantly more than others, in a relative sense."

P5864 L4 ... comparisons; although not an official product, users can readily...
Response: Agreed, done.
L6-16: this seems a more general description of the GOZCARDS files. Should some or all of this information be covered elsewhere, before getting into individual constituent discussion? Maybe at the end of the data screening/binning section.
Response: Agreed, done.

Section 3.3 This section is very long, and seems like a good place to make some larger cuts. As a reader interested in the GOZCARDS data, I am most interested in GOZCARDS comparisons with other results. I would suggest at the very least removing Figure 10. The authors can simply state, as they have, that analysis showed the short-term drifts in the period from 2006-2011 to be

most hemispherically asymmetric. I would suggest ending this section at P5866 L27 (end sentence with GOZCARDS lower stratospheric HCl trends agree quite well.) This shows what GOZCARDS can do, which is what users need from this paper.
Response: We made some significant cuts here (Sec. 3.3 is cut by 33% per word count). However, showing that the HCl behavior is turning around in the later years is useful and new (and in our view, informative for ACP audience, and as a brief update to the Mahieu et al. work, which mainly looked at columns - and stopped in 2011). We believe that it is most useful to show (keep) the 2 left panels of Fig. 10 to show the non-linearity and connection to surface total chlorine changes. Thus, we took out two thirds of Figs. 9 and 10 (combined) and shortened the text accordingly; we are just left with the new Fig. 9.

P5865:
L1: Remove sentence 'The GOZCARDS HCl ..." this was covered already, and is in the Tables.
Response: Agreed, done.
L4: ... documented by satellite-based ...
Response: Agreed, done.
L12: ... for the one and a half year band 13 AURA MLS data record ...
Response: Agreed, done.

P5866: L2 less negative trends.
Response: Agreed, done.
L13: Is putting Fig 9a, 9b and 9c in parentheses required? It seems a bit odd, but it might be a journal requirement. If not, I would suggest removing the ().
Response: The Journal did this – but we have reworded this in the new version.
L12-13: remove 'For the results'
Response: Agreed, done.
L17: and both positive and negative values
Response: Agreed, done.

P5868 Section 4.1: consider writing in active voice. '... we screened HALOE H2O data for high aerosol extinction values, closely following the screening used in the Stratospheric ... '
Response: Agreed, done.

P5869 L7: (Version 20 or VPMC) could be used ...
Response: Agreed, done.
L16-17: do the authors mean the resulting UARS MLS source data span 1991-1993?
Response: Yes, we clarified this.
L20: The validation results here are confusing. Is it necessary to talk about V2.2 and V3.3 AURA MLS validation? Sticking to V3.3 might limit some of the confusion.

Response: Yes, we have limited this more to the V3.3 discussion, although not exclusively so, for proper referencing of important past validation work.

p5871 L3: Sentence #2, this sentence could be removed, as the reason given for not including SAGE H2O seems good enough, that speculation to retrieval problems after the volcano are not needed.
Response: Agreed, done.
L9-L16: Do the authors mean care should be taken because there are missing bands in the GOZCARDS? Since space is at a premium in this paper, I would consider taking this paragraph out. Most users should be aware of sampling issues when averaging without full coverage and it's already been made clear that missing values are not filled in GOZCARDS.
Response: A user came to us with this specific (sort of "blind plotting") result/issue, and thus, we feel that it is worthwhile to include a note especially for the lower altitude region (not to have users, even a minority of them, rediscover the wheel in this case). The merged dataset does not readily provide flags or latitudinal coverage information for the source datasets, and users would need to do more work to track this particular issue down (maybe they should, agreed…). We prefer to keep this cautionary note (latitudinal inhomogeneity issues in early/occultation datasets), given also the interest in H_2O at lower altitudes; however, we shortened this text significantly.
L18: remove "(mentioned earlier)"
Response: Agreed, done.

P5872 L10: with a large drift in the difference time series.
Response: Agreed, done.
L15: if the two time series
Response: Agreed, done.
L16: consider removing this sentence 'The main point here... ' [this is understood]
Response: Agreed, done.
L19: end sentence at ACE-FTS data [don't need to mention again data ending in 2010]
Response: Agreed, done.

P5873 Section 4.3 This section is well written, but it reads like a completely new paper, and possibly could be part of a new manuscript. For the purposes of this paper however it seems far too detailed, especially considering the overall length of the paper. One suggestion would be to use the first sentence of the section, then skip to Fig 15 through the end of the section.
Response: Agreed, the literature review has been removed, as it is not that directly relevant or necessary. Overall, Sect. 4.3 is almost 40% shorter now.

P5876 L14: remove commas around "is needed"
Response: Agreed, done.

P5877 L12 were discarded L24: It is known
Response: Agreed, done.

P5880 L 11: remove ’as described in Sect 5.1’
Response: Agreed, done.
L13: remove sentence ’The monthly means ... ’ I think the construction of monthly means after screening is applied follows from section 5.1 and doesn’t need to be restated.
Response: Agreed, done.

P5882
L9-11: Isn’t the sentence starting ’Biases’ at the end of line 9 repeating what was said in the line above? It might be possible to shorten the diurnal section, just noting that the data are simply not co-located in time (sunset to 1:30am is a large time range). Solving this issue really requires model output which isn’t feasible for this work (at least this version).
Response: Yes, the redundant sentence was removed and some trimming of the rest was implemented although the point is that the relative bias adjustments include diurnal effects and one does not really need model output (with their own uncertainties) to do such adjustments, thus taking care of diurnal sampling effects, at least to first order. All datasets are basically adjusted to the SAGE II average local times (sunset and sunrise combined). Also, MLS (and ACE-FTS to a large extent) do not have significantly different (local time) sampling patterns year after year (see also the stability issue below).
L21: Only at 3.2 hPa and above, correct?
Response: Yes, we clarified this.
L18-27: This paragraph is a little confusing. Why is stability of measurements coming up now (assuming the authors are not still considering diurnal ozone variations). Is this not a consideration for all constituents?
Response: This was meant to focus on the diurnally-changing region. We clarified this, but also removed the last two sentences for brevity - as the stable sampling patterns of various instruments should be known.

P5886 L3 - end of section: The Ziemke Chandra results have always been anomalous, and it is probably not worthwhile pursuing this comparison. Updated column ozone comparisons could be made with SBUV MOD (which do not agree with ZC12); SBUV is considered the better "trend-quality" data set. However, given space constraints, it is not clear that a column ozone analysis is needed. The authors could easily end the section after the review of the GOZCARDS profile comparisons that have already been done.
Response: We think that cutting out the column ozone discussion completely would remove a significant analysis as we have not really seen published work on this subject, despite what some may view or characterize as “common knowledge” regarding TOMS/OMI column data time series analyses. This ozone discussion is new evidence nevertheless and it is the primary reason to tie this paper to the SI^{2}N special issue (which is on ozone, not on other species). Unless

specific published references already exist to clearly make our point and/or to imply that other column series are superior to TOMS/OMI, we plan to retain the flavor of this ozone-related "sample result". For brevity, however, we remove Fig. 23 and Fig, 25 and retain only Fig. 24 (Fig. 22 in the update); this preserves our point regarding comparison issues for GOZCARDS stratospheric column relative changes versus those provided by ZC12, but more succinctly. Overall, the Sect. 5.3 text has been cut by 37%.

P5888 L23: We now briefly mention the N_2O, HNO_3 and temperature GOZCARDS records that were part of the delivery...
Response: Agreed, done.

P5889 L6 end sentence after data record [again don't need to repeat this information about ACE-FTS if it is mentioned in the general data screening section and/or Tables].
Response: Agreed, done.
L11-14: wording is confusing. Suggest 'Until then the GOZCARDS N_2O record will include AURA MLS N_2O though the end of 2012 only, to avoid the discontinuities resulting from the shift to 190 GHz band N_2O in the current V3.3 data.
Response: Yes, we have reworded and shortened this discussion.
L19-20: Is this sentence referring to the first validation studies mentioned? If so, combine these sentences. 'Validation results for the first few years of AURA MLS and ACE-FTS N_2O shows agreement mostly within 5% in the stratosphere (Lambert et al., 2007; Strong et al., 2008)'.
Response: We have shortened this brief validation discussion.
L25: is illustrated in Fig. S15, showing generally highly correlated fields and insignificant drifts.
Response: Agreed, done.

P 5890 L5 remove parentheses.
Response: Agreed, done.
L19: can remove 'the quality of'
Response: Agreed, done.

P5891 L2: and 2012 agree mostly within 10-15%
Response: Yes, modified slightly.
L15-18: remove parentheses.
Response: Agreed, done.

Summary and Conclusions: This section is also quite long, and will have to be changed to reflect changes made, particularly in the summary sections for each constituent. In the interest of space, a higher level summary, focusing primarily on the data set methodology, would suffice.
Response: We have shortened this summary section significantly (a total cut of 31%); however, we believe that sample results have their place in such a summary for ACP (and for many people who have or take time to read mostly Abstract and/or Summary).

Global OZone Chemistry And Related Datasets for the Stratosphere (GOZCARDS): methodology and sample results with a focus on HCl, H_2O, and O_3

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Abstract

We describe the publicly available data from the Global OZone Chemistry And Related Datasets for the Stratosphere (GOZCARDS) project, and provide some results, with a focus on hydrogen chloride (HCl), water vapor (H_2O), and ozone (O_3). This dataset is a global long-term stratospheric Earth System Data Record, consisting of monthly zonal mean time series starting as early as 1979. The data records are based on high quality measurements from several NASA satellite instruments and ACE-FTS on SCISAT. We examine consistency aspects between the various datasets. To merge ozone records, the time series are debiased relative to SAGE II values by calculating average offsets versus SAGE II during measurement overlap periods, whereas for other species, the merging derives from an averaging procedure during overlap periods. The GOZCARDS files contain mixing ratios on a common pressure/latitude grid, as well as standard errors and other diagnostics; we also present estimates of systematic uncertainties in the merged products. Monthly mean temperatures for GOZCARDS were also produced, based directly on data from the Modern-Era Retrospective analysis for Research and Applications.

The GOZCARDS HCl merged product comes from HALOE, ACE-FTS and lower stratospheric Aura MLS data. After a rapid rise in upper stratospheric HCl in the early 1990s, the rate of decrease in this region for 1997-2010 was between 0.4 and 0.7%/yr. On 6-8 yr timescales, the rate of decrease peaked in 2004-2005 at about 1%/yr, and has since levelled off, at ~0.5%/yr. With a delay of 6-7 years, these changes roughly follow total surface chlorine, whose behavior versus time arises from inhomogeneous changes in the source gases. Since the late 1990s, HCl decreases in the lower stratosphere have occurred with pronounced latitudinal variability at rates sometimes exceeding 1-2%/yr. Recent short-term tendencies of lower stratospheric and column HCl vary substantially, with increases from 2005-2010 for northern mid-latitudes and deep tropics, but decreases (increases) after 2011 at northern (southern) mid-latitudes.

For H_2O, the GOZCARDS product covers both stratosphere and mesosphere, and the same instruments as for HCl are used, along with UARS MLS stratospheric H_2O data (1991-1993). We display seasonal to decadal-type variability in H_2O from 22 years of data. In the upper mesosphere, the anti-correlation between H_2O and solar flux is now clearly visible over two full solar cycles. Lower stratospheric tropical H_2O has exhibited two periods of increasing values,

53 followed by fairly sharp drops (the well-documented 2000-2001 decrease and a recent drop in
54 2011-2013). Tropical decadal variability peaks just above the tropopause. Between 1991 and
55 2013, both in the tropics and on a near-global basis, H_2O has decreased by ~5-10% in the lower
56 stratosphere, but about a 10% increase is observed in the upper stratosphere and lower
57 mesosphere. However, such tendencies may not represent longer-term trends.

58 For ozone, we used SAGE I, SAGE II, HALOE, UARS and Aura MLS, and ACE-FTS data to
59 produce a merged record from late 1979 onward, using SAGE II as the primary reference. Unlike
60 the 2 to 3% increase in near-global column ozone after the late 1990s reported by some,
61 GOZCARDS stratospheric column O_3 values do not show a recent upturn of more than 0.5 to
62 1%; continuing studies of changes in global ozone profiles, as well as columns, are warranted.

63 A brief mention is also made of other currently available, commonly-formatted GOZCARDS
64 satellite data records for stratospheric composition, namely those for N_2O and HNO_3.

65 ## 1 Introduction

66 The negative impact of anthropogenic chlorofluorocarbon emissions on the ozone layer,
67 following the early predictions of Molina and Rowland (1974), stimulated interest in the trends
68 and variability of stratospheric ozone, a key absorber of harmful ultraviolet radiation. The
69 discovery of the ozone hole in ground-based data records (Farman et al., 1985) and the
70 associated dramatic ozone changes during southern hemisphere winter and spring raised the level
71 of research and understanding regarding the existence of new photochemical processes (see
72 Solomon, 1999). This research was corroborated by analyses of aircraft and satellite data (e.g.,
73 Anderson et al., 1989; Waters et al., 1993), and of independent ground-based data. Global total
74 column ozone averages in 2006-2009 were measured to be smaller than during 1964-1980 by
75 ~3%, and larger more localized decreases over the same periods reached ~6% in the southern
76 hemisphere midlatitudes (WMO, 2011). Halogen source gas emissions have continued to
77 decrease as a result of the Montreal Protocol and its amendments. Surface loading of total
78 chlorine peaked in the early 1990s and subsequent decreases in global stratospheric HCl and ClO
79 have been measured from satellite-based sensors (Anderson et al., 2000; Froidevaux et al., 2006;
80 Jones et al., 2011) as well as from the ground (e.g., Solomon et al., 2006, Kohlhepp et al., 2012).
81 A slow recovery of the ozone layer towards pre-1985 levels is expected (WMO, 2011; 2014).

High quality long-term datasets for ozone and related stratospheric species are needed to document past variability and to constrain global atmospheric models. The history of global stratospheric observations includes a large suite of satellite-based instruments, generally well-suited for the elucidation of long-term global change. A review of differences between past and ongoing satellite measurements of atmospheric composition has been the focus of the Stratosphere-troposphere Processes And their Role in Climate (SPARC) Data Initiative; results for stratospheric H_2O and O_3 intercomparisons have been described by Hegglin et al. (2013) and Tegtmeier et al. (2013), respectively, to be followed by a report on many other species. Systematic biases reported in these papers mirror past validation work.

Under the Global OZone Chemistry And Related Datasets for the Stratosphere (GOZCARDS) project, we have created monthly zonally averaged datasets of stratospheric composition on a common latitude/pressure grid, using high quality data from the following satellite instruments: the Stratospheric Aerosol and Gas Experiments (SAGE I and SAGE II), the Halogen Occultation Experiment (HALOE) which flew aboard the Upper Atmosphere Research Satellite (UARS), the UARS Microwave Limb Sounder (MLS), the Atmospheric Chemistry Experiment Fourier Transform Spectrometer (ACE-FTS) on SCISAT, and Aura MLS. Table 1 provides characteristics of the original datasets; validation papers from the instrument teams and other related studies give a certain degree of confidence in these data. However, the existence of validation references does not imply that there are no caveats or issues with a particular measurement suite. In this project, we have strived to optimize data screening and mitigate some undesirable features, such as the impact of outlier values or the effects of clouds or aerosols. All source data sets still have imperfections, but in creating the GOZCARDS Earth System Data Record (ESDR) we maintain the integrity of the original data and do not arbitrarily disregard data, nor do we typically attempt to fill in spatial or temporal gaps in the record.

Based on original profiles from the various instruments, GOZCARDS “source” monthly zonal mean values were derived. After data screening, monthly average profiles were created by vertical interpolation onto the GOZCARDS pressure levels, followed by binning and averaging into monthly sets. In order to accomodate the lower vertical resolution of some limb viewers, such as UARS MLS, the GOZCARDS pressure grid was chosen as

$$p(i) = 1000 \times 10^{-\frac{i}{6}} \text{(hPa)} \quad (1)$$

with i varying from 0 to a product-dependent top; this grid width corresponds to ~2.7 km. The high resolution SAGE O_3 profiles were converted to mixing ratio versus pressure using their associated NCEP temperature profiles, and smoothed vertically onto this grid. Given the sampling of solar occultation instruments, which usually provide 15 sunrise and 15 sunset profiles in two narrow latitude bands every day (versus the denser sampling from MLS, with almost 3500 profiles/day), we used 10°-wide latitude bins (18 bins from 80°S-90°S to 80°N-90°N) to construct monthly zonal means. Next, we merged the GOZCARDS source data by computing average relative biases between source datasets during periods of overlap, and then adjusting each source dataset to a common reference to remove relative biases. Non-zero biases always exist between data from different instruments for various reasons, such as systematic errors arising from the signals or the retrieved values, different vertical resolutions, or sampling effects. Toohey et al. (2013) studied sampling biases from a large suite of satellite-based stratospheric profiling instruments, based on simulations using fully-sampled model abundance averages versus averages of output sampled at sub-orbital locations. Larger sampling errors arise from occultation than from emission measurements, which often sample thousands of profiles per day. Toohey et al. (2013) found that sampling biases reach 10-15%, notably at high latitudes with larger atmospheric variability. Sofieva et al. (2014) have also discussed sampling uncertainty issues for satellite ozone datasets.

We have observed very good correlations between GOZCARDS and other long-term ozone data, such as the Stratospheric Water vapor and OzOne Satellite Homogenized (SWOOSH) data (S. Davis, personal communication, 2012) and homogenized Solar Backscatter Ultraviolet (SBUV) data. Dissemination of trend results arising from analyses of GOZCARDS and other merged ozone datasets was planned as part of WMO (2014) and the SI^{2}N (Stratospheric Processes And their Role in Climate (SPARC), International Ozone Commission (IOC), Integrated Global Atmospheric Chemistry Observations (IGACO-O_3), and the Network for the Detection of Atmospheric Composition Change (NDACC)) initiative. Profile trend results have been provided by Tummon et al. (2015), Harris et al. (2015), as well as Nair et al. (2013, 2015).

This paper starts with a discussion of data screening issues (Sect. 2 and Appendix A), and then describes the GOZCARDS data production methodology, followed by some atmospheric

results for HCl (Sect. 3), H_2O (Sect. 4), and O_3 (Sect. 5). We provide specific diagnostics that indicate generally good correlations and small relative drifts between the source datasets used to create the longer-term GOZCARDS merged time series. Section 6 briefly mentions GOZCARDS N_2O and HNO_3, as well as temperatures derived from Modern-Era Retrospective analysis for Research and Applications (MERRA) fields. The version of GOZCARDS described here is referred to as ESDR version 1.01 or ev1.01.

2 GOZCARDS source data and data screening

Data provenance information regarding the various measurements used as inputs for GOZCARDS is provided in Appendix A (Sect. A.1).

2.1 GOZCARDS data screening and binning

The screening of profiles for GOZCARDS has largely followed guidelines recommended by the various instrument teams and/or relevant publications, and we have documented these issues and procedures in Appendix A (Table A1). Unless otherwise noted, we only provide monthly means constructed from 15 or more (good) values in a given latitude/pressure bin. For ACE-FTS data, we also found it necessary to remove occasional large outlier values that could significantly impact the monthly zonal means. Our outlier screening removed values outside 2.5 times the standard deviation, as measured from the medians in each latitude/pressure bin, for each year of data. This was deemed close to optimum by comparing results to Aura MLS time series, which typically are not impacted by large outliers, and to ACE-FTS zonal means screened (in a slightly different way) by the ACE-FTS instrument team. Up to 5% of the profile values in each bin in any given month were typically discarded as a result, but the maximum percentage of discarded values can be close to 10% for a few months of ACE-FTS data, depending on year and species. Moreover, because of poor ACE-FTS sampling, the threshold for minimum number of good ACE-FTS values determining a monthly zonal mean was allowed to be as low as 10 for mid- to high latitudes, and as low as 6 for low latitudes (bins centered from 25°S to 25°N). Zonal mean data from ACE-FTS become too sparse in some years if such lower threshold values are not used. Finally, no v2.2 ACE-FTS data are used after September 2010 (or after December 2009 for

ozone) because of a data processing problem that affected this data version; a newly reprocessed ACE-FTS dataset was not available before we made the GOZCARDS data public.

Placing profiles on a common pressure grid is straightforward when pressures are present in the original files, as is the case for most data used here. Also, the vertical resolutions are similar for most of the instruments used for GOZCARDS. The UARS MLS, HALOE, and Aura MLS native pressure grids are either the same as or a superset of the GOZCARDS pressure grid, so these datasets were readily sampled for the construction of monthly means. For ACE-FTS, pressures are provided along with the fixed altitude grid, and we used linear interpolation versus log(pressure) to convert profiles to the GOZCARDS grid. More details are provided later for SAGE I and SAGE II O_3, for which density versus altitude is the native representation.

The binning of profiles occurs after the screened values are averaged (in each latitude/pressure bin). Note that for the species discussed here, sunset and sunrise occultation values in the same latitude bin during a given month are averaged together. Negative monthly means are set to -999.0 in the GOZCARDS files; while negative mixing ratios smaller (in absolute value) than the associated standard errors can in theory be meaningful, negative monthly means are unlikely to be very useful scientifically. Quantities other than mixing ratios are provided in the netCDF GOZCARDS files, which are composed of one set of individual yearly files for all source datasets, and one set of yearly files for the merged products. The main quantities are monthly averages, plus standard deviations and standard errors. The GOZCARDS source files also provide the number of days sampled each month as well as minimum and maximum values for the source datasets. Other information includes average solar zenith angles and local solar times for individual sources. Finally, formulae for monthly standard deviations of the merged data are given in Appendix A, where sample time series of the standard deviations and standard errors (not systematic errors) for both source and merged data are also shown.

3 GOZCARDS HCl

3.1 GOZCARDS HCl source data records

We used HCl datasets from HALOE, ACE-FTS and Aura MLS to generate the monthly zonal mean source products for GOZCARDS HCl. In addition to the procedures mentioned before, a

first-order aerosol screening was applied to the HALOE HCl profiles: all HCl values at and below a level where the 5.26 μm aerosol extinction exceeds 10^{-3} km^{-1} were excluded. Regarding Aura MLS HCl, Froidevaux et al. (2008a) found anomalously high values versus aircraft data at 147 hPa at low latitudes; these values are not used in the production of the merged HCl product. Also, the ongoing standard MLS HCl product is retrieved using band 14 rather than band 13, which targeted HCl for the first 1.5 years after launch, but started deteriorating rapidly after Feb. 2006. As the remaining lifetime for band 13 is expected to be short, this band has been turned on only for a few days since Feb. 2006. MLS HCl data are not recommended for trend analyses at pressures < 10 hPa. However, for pressures ≥ 10 hPa, band 14 HCl is deemed robust, because of the broader emission line in this region, in comparison to the measurement bandwidth.

Past validation studies have compared MLS HCl (v2.2), ACE-FTS (v2.2) and HALOE (v19) datasets using coincident pairs of profiles; such work was described by Froidevaux et al. (2008a) for MLS HCl validation and by Mahieu et al. (2008) for ACE-FTS HCl validation. The MLS version 3.3/3.4 HCl data used here (see Livesey et al., 2013) compare quite well with v2.2 HCl (average relative biases are within 5%). HALOE HCl values were found to be biased low by ~10-15% relative to both MLS and ACE-FTS, especially in the upper stratosphere; this low bias versus other (balloon- and space-based) measurements had been noted in past HALOE validation studies (Russell et al., 1996). Also, HALOE (v19) and ACE-FTS (v2.2) HCl data tend to lose sensitivity and reliability for pressures less than ~0.4 hPa.

3.2 GOZCARDS HCl merged data records

Although Aura MLS HCl data for pressures less than 10 hPa do not contribute to the time dependence of the merged HCl product, the 2004-2005 absolute Aura HCl measurements in this region are used to compute the offsets for the ACE-FTS and HALOE zonal mean source data in a consistent manner versus pressure. Figure 1 illustrates the merging process for HCl at 32 hPa for the 45°S latitude bin (covering 40°S-50°S). Given that there exists very little overlap between the three sets of measurements in the same months in 2004 and 2005, especially in the tropics, a simple 3-way averaging of the datasets would lead to significant data gaps. Our methodology is basically equivalent to averaging all three datasets during this period, and we use Aura MLS as a transfer dataset. This was done by first averaging ACE-FTS and Aura MLS data, where the

datasets overlap, and then including the third dataset (HALOE) into the merging process with the temporary merged data. As the HALOE HCl values are generally lower than both the MLS and ACE-FTS values, the merged HCl dataset is generally further away from HALOE than it is from either ACE-FTS or Aura MLS. The top left panel in Fig. 1 shows GOZCARDS source data for HALOE, ACE-FTS, and Aura MLS during the overlap period, from August 2004 (MLS data start) through November 2005 (HALOE data end). The top right panel illustrates the result of step 1 in the merging procedure, with the temporary merged data values (orange) resulting from the adjustment of ACE-FTS and Aura MLS values to the mean reference (black dashed line); this reference is simply the average of the two series for all months when both values exist. The middle left panel shows step 2, namely the values (brown) obtained from merging HALOE values with the temporary merged values from step 1; the temporary merged values are weighted by 2/3 and HALOE values by 1/3 (giving the black dashed line as mean reference), so this is equivalent to averaging the three datasets with a weight of 1/3 each. A simple mathematical description of the above procedure is provided in Appendix A. The middle right panel shows the source data along with the final merged values during the overlap period, whereas the bottom panel shows the full time period, after the additive offsets are applied to the whole source series, thus removing relative biases; the three adjusted series are then averaged together wherever overlap exists, to obtain the final merged dataset. We tested this procedure by using one or the other of the two occultation data as the initial one (for step 1) and the results were not found to differ appreciably. We also found that the use of multiplicative adjustments generally produces very similar results as additive offsets. Some issues were found on occasion with multiplicative offsets, when combining very low mixing ratios, but additive offsets can also have drawbacks if the merged values end up being slightly negative, notably as a result of changes that modify the already low HCl values during Antarctic polar winter. This occurs on occasion as additive offsets tend to be weighted more heavily by larger mixing ratios found during non-winter seasons; as a result, we decided not to offset lower stratospheric HCl source datasets in the polar winter at high latitudes for any of the years. Further specifics and procedural details regarding the merging of HCl data are summarized in Appendix A.

In Fig. 2, we display the offsets that were applied to the three HCl source datasets as a result of the merging process in each latitude/pressure bin; a positive value means that a dataset is

biased low relative to the reference mean and needs to be increased by the offset value. These offsets show that in general, ACE-FTS and Aura MLS HCl values were adjusted down by 0.1-0.2 ppbv (a decrease of about 2-10%), while HALOE HCl was adjusted upward by 0.2-0.4 ppbv. Offset values tend to be fairly constant with latitude and the sum of the offsets equals zero. The generally homogeneous behaviour versus latitude is a good sign, as large discontinuities would signal potential issues in the merging (e.g., arising from large variability or lack of sufficient statistics). Figure S1 provides more detailed examples of upper and lower stratospheric offsets versus latitude, including standard errors based on the variability in the offsets during the overlap period (error bars provide an indication of robustness). Another indication of compatibility between datasets is provided by a comparison of annual cycles. Figure S2 provides average annual cycle amplitudes obtained from simple regression model fits to HALOE, ACE-FTS, and Aura MLS series over their respective periods. While there are a few regions where noise or spikes exist (mainly for ACE-FTS), large annual amplitudes in the polar regions occur in all time series; this arises from HCl decreases in polar winter, followed by springtime increases.

A more detailed analysis of interannual variability and trend consistency is provided from results in Fig. 3, which shows an example of ACE-FTS and Aura MLS time series. We have used coincident points from these time series to compare the deseasonalized anomalies (middle panel in Fig. 3) from both instrument series; correlation coefficient values (R values) are also computed. Very good correlations are obtained and no significant trend difference between the anomalies (bottom panel in Fig. 3) exists for ACE-FTS versus Aura MLS HCl. A view of these correlations and drifts at all latitudes/pressures is provided in Fig. 4, where the top panel gives R values for deseasonalized anomalies, and the bottom panel gives the ratio of the difference trends over the error in these trends. The results in Fig. 4 confirm that there are significant trend differences between the upper stratospheric HCl time series from ACE-FTS and Aura MLS (as a reminder, we did not use Aura MLS HCl for pressures less than 10 hPa). Fig. 4 also shows very low correlation coefficients from the deseasonalized HCl series in the uppermost stratosphere, because Aura MLS HCl exhibits unrealistically flat temporal behavior, whereas ACE-FTS HCl varies more. In the lower stratosphere, there is generally good agreement between the ACE-FTS and Aura MLS HCl time series, with R values typically larger than 0.7 and difference trend to

error ratios smaller than 1.5. The few low R values for 100 hPa at low latitudes likely reflect more infrequent ACE-FTS sampling and some (possibly related) outlier data screening issues.

Figure S3 illustrates GOZCARDS merged 46 hPa HCl variations versus time; there is clearly a much more complete global view (with no monthly gaps) after the launch of Aura MLS. Gaps at low latitudes in 1991 and 1992 are caused by post-Pinatubo aerosol-related issues in the HALOE record, and gaps in later years arise from the decrease in coverage from UARS. In the upper stratosphere, there are more gaps compared to 10 hPa and below, as a result of the much poorer tropical coverage from ACE-FTS and the elimination of MLS data in this region.

An indication of systematic errors in the merged values is given by the range of available monthly mean source data. For each bin, we compute the ranges of monthly means above and below the merged values that include 95% of the available source data monthly means. These error bars are not usually symmetric about the merged values, especially if one dataset is biased significantly more than others, in a relative sense. We did not have enough datasets here to consider a more statistical approach (such as the standard deviations among source datasets). Figure 5 shows the result of such a systematic error calculation at 46 hPa for the 35°S latitude bin. The lower shaded region range gives the lower bound, determined by HALOE data, and the upper limit of the grey shading originates from ACE-FTS data. Figure 6 shows contour plots of these estimated systematic errors in HCl. These are fairly conservative error bars; however, even the source data averages at the 95% boundaries have their own systematic errors (rarely smaller than 5%), so our estimates do not really encompass all error sources. Error bars representing a range within which 95% of the source data values reside (see Figs. 5 and 6) can be a useful guide for data users or model comparisons; alhough this is not an official product, users can readily calculate such ranges (or we can provide these values).

3.3 GOZCARDS HCl sample results and discussion

Stratospheric HCl is important because it is the main reservoir of gaseous chlorine and it can be used to follow the chlorine budget evolution over the past decades. This includes a significant increase before the mid-1990s as a result of anthropogenic chlorofluorocarbon (CFC) production, followed by a slower decrease as a result of the Montreal Protocol and subsequent international

agreements to limit surface emissions that were correctly predicted to be harmful to the ozone layer (Molina and Rowland, 1974; Farman et al., 1985).

In Fig. 7, we provide an overview of the HCl evolution since 1991, based on GOZCARDS average merged HCl for 3 different latitude regions at 4 pressure levels, from the upper stratosphere to the lower stratosphere. In the upper stratosphere (at 0.7 hPa shown here), the rapid early rise in HCl was followed by a period of stabilization (1997-2000) and subsequent decreases. Rates of decrease for stratospheric HCl and total chlorine have been documented using satellite-based upper stratospheric abundances, which tend to follow tropospheric source gas trends with a time delay of order 6 years, with some uncertainties in the modeling of this time delay and related age of air issues (Waugh et al., 2001; Engel et al., 2002; Froidevaux et al., 2006). As summarized in WMO (2011), the average rate of decrease in stratospheric HCl has typically been measured at -0.6 %/yr to -0.9 %/yr, in reasonable agreement with estimated rates of change in surface total chlorine; see also the HCl upper stratospheric results provided by Anderson et al. (2000) for HALOE, Froidevaux et al. (2006) for the one and a half year band 13 Aura MLS data record, and Jones et al. (2011) and Brown et al. (2011) for a combination of HALOE and ACE-FTS datasets. The WMO (2011) summary of trends also includes results from column HCl data at various NDACC Fourier transform infrared (FTIR) measurement sites; see Kohlhepp et al. (2012) for a comprehensive discussion of ground-based results, showing some scatter as a function of latitude. Figure 7 demonstrates that a global-scale decline in mid- to lower stratospheric HCl is visible since about 1997. We also notice that at 68 hPa in the tropics, the long-term rate of change appears to be near-zero or slightly positive. In addition, there are shorter-term periods in recent years when an average increasing “trend” would be inferred rather than a decrease; in particular, see the northern hemisphere from 2005 through 2012 at 32 hPa.

We created deseasonalized GOZCARDS merged monthly zonal mean HCl data at different latitudes and we show in Fig. 8 the linear rate of change that results from simple fits through such series. The long-term trends (1997 - 2013 for lower and 1997 - 2010 for upper stratosphere) are generally negative and between about -0.5%/yr (upper stratosphere) and -1%/yr (lower stratosphere). Some separation between northern and southern hemisphere results is observed in the lower stratosphere, with less negative trends in the northern hemisphere. Also, the scatter increases from 68 hPa to 100 hPa, where some positive trends occur at low latitudes; however,

we have less confidence in the 100 hPa results, given the larger scatter and errors (and smaller abundances) in that region. Without trying to assign exact linear trends from these simple analyses, we observe considerable latitudinal variability in lower stratospheric HCl short-term behavior, especially after 2005. Such lower stratospheric changes in HCl have been captured in column HCl FTIR data (Mahieu et al., 2013, 2014). The latter reference shows that total column (FTIR) results and GOZCARDS lower stratospheric HCl trends agree quite well, and the authors imply that a relative slowdown in the northern hemispheric circulation is responsible for observed recent changes in the lower stratosphere. However, we note (Fig. 7) that changes in lower stratospheric HCl appear to be fairly short-term in nature, with an apparent reversal in behavior occurring at both northern and southern midlatitudes since 2011 (e.g., at 32 hPa). Lower stratospheric changes are distinct from the upper stratospheric long-term decrease, which we expect to continue, as long as the Montreal Protocol and its amendments are followed and total surface chlorine keeps decreasing.

In Figure 9, we provide simple rates of upper and lower stratospheric change in HCl for 6-yr sliding time periods (e.g., a 2004 value means a 2001-2006 average) for various latitudes. These results indicate that there has been an acceleration in the rate of decrease of upper stratospheric HCl between 2000 and 2004, followed by a period with somewhat smaller rates of change. This is roughly in agreement with curves showing the rates of change for surface total chlorine based on National Oceanic and Atmospheric Administration (NOAA) surface data (Montzka et al., 1999), as shown in Fig. 9 (top panel) with the Earth System Research Laboratory Global Monitoring Division data, time shifted by 6 or 7 years to account for transport delays into the upper stratosphere. Chlorine source gases have indeed shown a reduction in their rate of decrease during the second half of the past decade, as discussed by Montzka et al. (1999) and summarized in WMO (2011, 2014). Reasons include the initial rapid decrease in methyl chloroform, slower rates of decrease from the sum of CFCs in recent years, and increases in hydrochlorofluorocarbons (HCFCs). The lower stratospheric HCl behavior in Fig. 9 (bottom panel) shows rates of change in partial column density between 68 hPa and 10 hPa. These changes show more variability with latitude than in the upper stratosphere for short (6-yr) time periods, and a hemispheric asymmetry exists, peaking in 2009, when positive tendencies are seen in the northern hemisphere, as opposed to decreases in the south (Mahieu et al., 2014). These

results do not depend much on whether 6-yr or 8-yr periods (not shown) are used, but longer periods smooth out the rates of change; interannual variations, such as those arising from the quasi-biennial oscillation (QBO), will affect short-term results. Temporal patterns in the upper and lower stratosphere are qualitatively similar, and rates of change in surface emissions will impact both regions, but carefully disentangling this from changes in dynamics or in other species (e.g., CH_4) that can affect chlorine partitioning will require more analyses and modeling.

4 GOZCARDS H_2O

4.1 GOZCARDS H_2O source data records

We used water vapor datasets from HALOE, UARS MLS, ACE-FTS, and Aura MLS to generate the monthly zonal mean source products for GOZCARDS H_2O. In addition to the data screening procedures mentioned in Appendix A, we screened HALOE H_2O data for high aerosol extinction values, closely following the screening used for merged H_2O in the Stratospheric Water vapor and OzOne Satellite Homogenized (SWOOSH) dataset (S. Davis, personal communication, 2012). This method (see Fig. S4) screens out anomalous HALOE H_2O values that occurred mainly in 1991-1992, when the aerosol extinction near 22 hPa exceeded $5x10^{-4}$ km^{-1}; for pressure levels at and below 22 hPa, we have excluded the corresponding H_2O values. While this method may exclude some good data points, the lowest values (< 3 ppmv) do get screened out; such outliers are not corroborated by 22 hPa UARS MLS data (with most values > 3 ppmv). Also, for upper mesospheric HALOE data, care should be taken during high latitude summer months, as no screening was applied for the effect of polar mesospheric clouds (PMCs). High biases (by tens of percent) in H_2O above ~70 km have been shown to occur as a result of PMCs in the HALOE field of view (McHugh et al., 2003). Indeed, monthly means larger than 8-10 ppmv are observed in GOZCARDS H_2O merged data and in HALOE source data for pressures less than ~0.03 hPa. A more recent HALOE data version (V20 or VPMC) could be used to largely correct such PMC-related effects, although this was not implemented for GOZCARDS H_2O. Aura MLS and ACE-FTS measurements, obtained at longer wavelengths than those from HALOE, do not yield such large H_2O values; a rough threshold of 8.5 ppmv could also be used (by GOZCARDS data users) to flag the pre-2005 merged dataset.

UARS MLS stratospheric H_2O for GOZCARDS was obtained from V6 (or V600) H_2O data. This data version is identical to the original prototype (named V0104) from Pumphrey (1999), who noted that UARS MLS H_2O often exhibits drier values (by 5-10%) than HALOE H_2O (see also Pumphrey et al., 2000). The resulting UARS MLS H_2O source data span the period from Sep. 1991 through April, 1993; a significant fraction of this dataset in the tropics at 100 hPa is flagged bad, as a result of diminishing sensitivity.

Summarizing past validation results, SPARC WAVAS (2000) analyses pointed out the existence of a small low bias in HALOE stratospheric H_2O versus most other measurements, except for UARS MLS. Lambert et al. (2007) showed agreement within 5-10% between Aura MLS version 2.2 H_2O and other data, including ACE-FTS H_2O. From the mid-stratosphere to the upper mesosphere, excellent agreement between ground-based data from the Water Vapor Millimiter-wave Spectrometer and H_2O profiles from Aura MLS and ACE-FTS has been demonstrated by Nedoluha et al. (2007, 2009, 2011). Changes from MLS v2.2 to v3.3 led to an increase of 0.2-0.3 ppmv in stratospheric H_2O (Livesey et al., 2013). Recent comparisons by Hurst et al. (2014) of MLS v3.3 H_2O data versus Cryogenic Frost point Hygrometer time series above Boulder show excellent overall agreement, indicating that systematic uncertainties for lower stratospheric MLS data may be as low as ~5%; this reinforces MLS H_2O validation work by Read et al. (2007) and Voemel et al. (2007). Aura MLS stratospheric H_2O v3.3 values are slightly larger (by up to ~5%) than the multi-instrument average from a number of satellite datasets, as discussed in satellite intercomparisons by Hegglin et al. (2013), who observed only small disagreements in interannual variations from various series for pressures less than 150 hPa.

4.2 GOZCARDS H_2O merged data records

The merging process for H_2O is nearly identical to the method used for HCl. The main difference is an additional step that merges UARS MLS data with the already combined datasets from HALOE, ACE-FTS, and Aura MLS, by simply adjusting UARS MLS values to the average of the previously merged series during the early (1991-1993) overlap period; see Fig. S5 for an illustration. Typically, this requires an upward adjustment of UARS MLS H_2O data, as these values are biased low versus most other datasets; nevertheless, the short but global record from UARS MLS helps to fill the time series. After considering the channel drift issues for SAGE II

H_2O (and following past advice from the SAGE II team itself), we decided to use caution and did not include that dataset for GOZCARDS, as trend results could be affected. Other minor procedural merging details or issues for H_2O are included in the Supplement. Also, data users should be aware of anomalous effects arising in merged average series from non-uniform latitudinal sampling when no MLS data exist, in regions with large latitudinal gradients, as for H_2O at 147 hPa, the largest pressure for merged GOZCARDS H_2O. Latitudinal averages can be biased in certain months and month-to-month variability is increased because of relatively poor global sampling (in this region) prior to Aug. 2004, after which Aura MLS data are used.

In Fig. 10, we display the average offsets that were applied to the four H_2O source datasets; these offsets follow previously known relative data biases. For example, low biases in UARS MLS H_2O, especially in the mesosphere, were discussed by Pumphrey (1999) and the UARS MLS offsets (see Fig. 10) correct that dataset upward. The application of offsets derived for HALOE and UARS MLS raises the H_2O time series from these instruments, whereas negative offsets lower the H_2O source data from ACE-FTS and Aura MLS. As we found for HCl, the offset values generally display small variations versus latitude and are therefore fairly stable systematic adjustments to the time series. Figure S6 displays the amplitudes of the fitted annual cycles for HALOE, ACE-FTS, and Aura MLS. As for HCl, similar patterns emerge for these datasets. Wintertime descent into the polar vortex regions is responsible for large annual cycles at high latitudes, especially in the mesosphere; also, the seasonal impact of dehydration in the lower stratospheric Antarctic region causes a large annual cycle in Aura MLS high southern latitude data. Figure 11 provides some statistical information, as done for HCl in Sect. 3.2, regarding the correlations and trend differences between ACE-FTS and Aura MLS. There are a few regions with noisier relationships. While slow increases in H_2O are generally observed by both instruments in the stratosphere and mesosphere, the tropical region near 0.1 hPa shows a slight decreasing trend for the ACE-FTS points, thus leading to larger discrepancies; it is not clear what the source of these discrepancies is. While the tropical ACE-FTS data are generally sampled with a significantly lower temporal frequency, the same applies for all pressure levels; however, a few outlier points can have a much larger impact when sampling is poorer. There are also a few other spots, such as near 65°S and 65°N and near 5 hPa, with a large drift in the difference time series; this may be caused by a combination of poorer sampling by ACE-FTS

and higher atmospheric variability, which can lead to more scatter. At the highest latitudes in the lower stratosphere, the observed slope differences are more within error bars, but the larger variability means that a longer record is needed to determine if the time series trend differently. The merged dataset tends to be much closer to Aura MLS in terms of trends because there are many more months of Aura MLS than ACE-FTS data; the overall impact of ACE-FTS data on the merged H_2O series is fairly small.

Figure S7 provides a visual representation of the merged GOZCARDS H_2O fields at 3 hPa and 68 hPa, respectively. Well-known features are displayed in these plots, given the good global coverage in the post-2004 period in particular. In the upper stratosphere, descent at high latitudes during the winter months leads to larger H_2O values, and low latitude QBO features are also observed. In the lower stratosphere, one observes dehydration evidence at high southern latitudes in the winter months, as well as a low latitude seasonal "tape recorder" signal; this phenomenon is driven by tropopause temperatures and has been measured in satellite data since the early 1990s (Mote et al., 1996; Pumphrey, 1999). A vertical cross-section of this lower stratospheric tropical (20°S to 20°N) tape recorder in GOZCARDS merged H_2O for 1991-2013 is shown in Fig. 12; periods of positive anomalies alternate with negative anomalies, including the post-2000 lows, as well as the most recent decreases in 2012-2013 (see also next section).

As we discussed for HCl, we have estimated systematic errors for the merged H_2O product. This is illustrated by the contour plots in Fig. 13; these ranges encompass at least 95% of the monthly mean source data values from HALOE, UARS MLS, ACE-FTS, and Aura MLS above or below the merged series. These errors typically span 5 to 15% of the mean between 100 and 0.1 hPa; errors larger than 30% exist in the tropical upper troposphere (147 hPa), and large values in the upper mesosphere arise from the low bias in UARS MLS H_2O.

4.3 GOZCARDS H_2O sample results and discussion

Stratospheric H_2O variations have garnered attention because of the radiative impacts of water vapor in the UTLS and the connection to climate change (Solomon et al., 2010), as well as the stratospheric chemical significance of H_2O oxidation products. Individual water vapor datasets have been used here to produce a merged stratospheric H_2O record spanning more than two decades. We do not attempt here to characterize trends or to imply that recent tendencies will

carry into the next decade or two. Rather, as variability is also of interest to climate modelers, we focus here on observed decadal-type (longer-term) variability in stratospheric water vapor.

Figure 14 illustrates monthly, annual, and longer-term changes in stratospheric water vapor, based on GOZCARDS merged H_2O; this shows the well-known H_2O minimum in the lower tropical stratosphere as well as an increases in the upper stratosphere (as a result of methane oxidation). As we know from past studies (e.g., Randel et al., 2004), medium- to long-term changes in H_2O are large-scale in nature. However, lower stratospheric H_2O variations are more accentuated at low latitudes, in comparison to near-global (60°S-60°N) results. It has long been known (e.g., from the *in situ* balloon-borne measurements of Kley et al., 1979) that the hygropause is typically located a few km higher than the thermal tropopause. We observe that the tropical stratosphere is drier at 68 hPa than at 100 hPa (near the tropopause). According to the 22-year GOZCARDS data record, annually-averaged H_2O values in the tropics (20°S-20°N) have varied between about 3.2 and 4.2 ppmv at 68 hPa. The rapid drop between 2000 and 2001 is observed at 100 and 68 hPa, with some dilution of this effect at higher altitudes. There is a clear difference in long-term behavior between the upper stratosphere, where changes in methane should have the clearest influence, and the lower stratosphere, where cold point temperatures and dynamical changes have a significant impact. To first-order, the last few years show ~10% larger values in the upper stratosphere than in the early 1990s, while the opposite holds in the lowest stratospheric region, where a decrease of order 10% is observed over the same period. Figure 14 also shows that month-to-month and seasonal variations are usually somewhat larger than the long-term changes in the lower stratosphere, most notably at 100 hPa.

In order to provide longer-term variability diagnostics for water vapor, we show in Fig. 15 the minimum to maximum spread in annual averages (tropics and mid-latitudes) from Fig. 14 for the 22-yr period. We observe that the tropical variability is largest just above the tropopause (here this means at the 68 hPa GOZCARDS level), where it reaches ~27% (1 ppmv). Such diagnostics of variability should be useful for comparisons to various chemistry climate models.

The longer-term variability in water vapor increases above the stratopause and reaches close to 30% in the uppermost mesosphere, as seen in Fig. 16(a); this plot shows the monthly and annual near-global (60°S-60°N) H_2O variations at 0.01 hPa. Large seasonal changes in this region are driven by vertical advection associated with the mesospheric circulation, with each

hemisphere's summertime peaks contributing to the maxima (two per year) in these near-global averages; such seasonal variations were compared to model results by Chandra et al. (1997), based on the first few years of HALOE H_2O data. The strong upper mesospheric variability in annual-mean H_2O is known from previous studies of ground-based and satellite H_2O data (Chandra et al., 1997; Nedoluha et al., 2009; Remsberg, 2010), and this region is where the solar (Lyman α) influence on H_2O is strongest. Figure 16(b) displays the near-global variations in annual upper mesospheric H_2O from 0.1 to 0.01 hPa. We clearly see increased variability in the uppermost mesosphere, and decreases in the mixing ratios as a result of H_2O photodissociation.

5 GOZCARDS ozone

A number of discussions relating to signs of ozone recovery have been presented before (Newchurch et al., 2003; Wohltmann et al., 2007; Yang et al., 2008; Jones et al., 2009; Hassler et al., 2011; Salby et al., 2011, 2012; Ziemke and Chandra, 2012; Gebhardt et al., 2014; Kuttipurath et al., 2013; Kirgis et al., 2013; Nair et al., 2013, 2015; Shepherd et al., 2014, Frith et al., 2014). While there are some indications of small increases in O_3 in the past 10-15 years, further confirmation of an increase in global O_3 and its correlation with column increases is needed in order to more clearly distinguish between long-term forcings, notably from the 11-yr solar cycle, slow changes in halogen source gases, temperature changes, and shorter-term variability. Continuing, good long-term ozone datasets are clearly needed for such studies.

5.1 GOZCARDS ozone source data records

We used ozone datasets from SAGE I, SAGE II, HALOE, UARS MLS, ACE-FTS, and Aura MLS to generate the monthly zonal mean source products for GOZCARDS. Due to time constraints, we did not use the newer SAGE II version 7 ozone (see Damadeo et al., 2013) as part of the GOZCARDS merged dataset. Our studies indicate that there are systematic differences of only a few percent between SAGE II V6.2 and V7 O_3 on their native coordinates (number density versus altitude). However, these 2 versions exhibit some differences if the data are converted to mixing ratios on pressure surfaces. These differences result mainly from different temperatures (and their trends) between MERRA and analyses from the National Centers for Environmental Prediction (NCEP), used by SAGE II V7 and V6.2 retrievals,

548 respectively. The main differences between MERRA and NCEP temperatures occur in the upper
549 stratosphere for time periods before 1989 and after mid-2000 (see further details in Sect. 5.2).

550 5.1.1 Treatment of SAGE ozone profiles

551 Both SAGE I and SAGE II used solar occultations during satellite sunrise and sunset to measure
552 vertical profiles of ozone, along with other composition data and aerosol extinction (McCormick
553 et al., 1989; Cunnold et al., 1989). It takes about 1 month for SAGE I and II to provide near
554 global coverage (about 80°N to 80°S), with some dependence on season. The SAGE I
555 measurements started in February 1979 and stopped in November 1981, while SAGE II provided
556 data between October 1984 and August 2005. In the middle of July 2000, SAGE II had a
557 problem in its azimuth gimbal system. Although this was corrected by November 2000, the
558 instrument operation was switched to a 50% duty cycle, with either sunrise or sunset occultations
559 occurring in monthly alternating periods, until the end of the mission.

560 It is known that there were altitude registration errors in SAGE I (V5.9) data (Veiga et al.,
561 1995; Wang et al., 1996). To correct this problem, an empirical altitude correction method based
562 on Wang et al. (1996) had been applied to SAGE I (V5.9) data; these corrected SAGE I V5.9
563 profiles, which had been evaluated in previous trend studies (e.g. SPARC, 1998; WMO, 2003),
564 were used to create the GOZCARDS SAGE I product (denoted as version V5.9_rev). We did not
565 use reprocessed version 6.1 SAGE I data (L. W. Thomason, personal communication, 2012)
566 because the altitude registration problems had not been completely fixed and new altitude
567 correction criteria should be derived and validated.

568 Ozone data screening details for the original SAGE I and SAGE II datasets are provided in
569 Appendix A. The number density profiles were converted to mixing ratios on pressure levels by
570 using NCEP temperature and pressure data provided with each profile. Derived ozone profiles
571 were then interpolated to fixed pressure levels on the following grid:

572 $$p(i) = 1000 \times 10^{-\frac{i}{30}} (\text{hPa}) \quad i = 0, 1, 2, ... \tag{2}$$

573 Ozone values at each of the 5 levels centered on every GOZCARDS pressure level were then
574 averaged (weighted by pressure) to derive mixing ratios at each GOZCARDS pressure level. By
575 doing this, the SAGE profiles were smoothed to a vertical resolution comparable to that of the

other satellite instruments used in this GOZCARDS work. Monthly zonal means were then computed for the SAGE ozone datasets on the GOZCARDS-compatible grid.

5.1.2 Comparisons of ozone zonal means

Ozone differences between SAGE II and other satellite data are shown in Fig. S8. Zonal mean differences between SAGE II and HALOE are generally within 5% for 1.5 to 68 hPa at mid-latitudes, and for 1.5 to 46 hPa in the tropics; relative biases are larger outside those ranges and increase to ~10% near the tropopause and also near 1 hPa. This good level of agreement was demonstrated in the past (e.g., SPARC, 1998). SAGE II data show better agreement with UARS and Aura MLS in the upper stratosphere and lower mesosphere, within 5% up to 0.68 hPa and for latitudes outside the polar regions. Aura MLS O_3 compares better with SAGE II data than does UARS MLS in the tropics for pressures larger than 68 hPa; the high bias in UARS MLS O_3 at 100 hPa has been discussed previously (Livesey et al., 2003). There are no months that include both SAGE II and ACE-FTS data in the northern hemisphere tropics (see the gap in Fig. S8, bottom right panel), largely due to the poorer coverage from ACE-FTS in the tropics. ACE-FTS O_3 shows the largest positive bias (greater than 10%) with respect to SAGE II, for pressures less than 1.5 hPa. The high bias in upper stratospheric ACE-FTS ozone has been mentioned in past validation work using ACE-FTS data (e.g., Froidevaux et al., 2008b; Dupuy et al., 2009). The biases shown here are also consistent with recent O_3 intercomparison studies from a comprehensive array of satellite instruments by Tegtmeier et al. (2013). It has been known for some time that the HALOE and SAGE II ozone datasets, which govern the main variations of the GOZCARDS merged ozone values before 2005, agree quite well (within 5%) in absolute value, and also in terms of temporal trends (Nazaryan et al., 2005), and versus ozonesondes (mostly above ~20 km or ~50 hPa). Larger percentage differences occur in the lowest region of the stratosphere at low latitudes, and especially in the upper troposphere, where HALOE values become significantly smaller than SAGE II data, which are already biased low (by ~50%) versus sondes (Wang et al., 2002); see also Morris et al. (2002), as well as results of SAGE II and HALOE comparisons versus solar occultation UV-Visible spectrometer measurements from long duration balloons (Borchi et al., 2005). We should note here that in this GOZCARDS merging work, we have largely avoided the upper tropospheric region.

605 Zonal mean differences between SAGE II and Aura MLS show some latitudinal structure
606 between 1 and 3 hPa, with larger (5-10%) biases in the southern hemisphere, especially for 0 to
607 30°S (see Fig. S8). There are no such features between SAGE II and HALOE or UARS MLS.
608 We found that this results from anomalous NCEP temperatures after 2000, which affect SAGE II
609 data converted from number density/altitude to GOZCARDS VMR/pressure coordinates.
610 Figure 17 shows an example of the ozone series from SAGE II and other satellite data for 10°S to
611 20°S from 1 to 6.8 hPa. At 1 hPa, the SAGE II ozone values (converted to mixing ratios) drift
612 relative to HALOE and are elevated after mid-2000; this can be attributed to abnormal NCEP
613 temperature trends compared to MERRA and HALOE during the same time period (for detailed
614 views, see Figs. S9 and S10). Similar features are found down to pressures near 3 hPa. These
615 issues relating to anomalous upper stratospheric NCEP temperature trends were noted by
616 McLinden et al. (2009). Because such artifacts are confirmed by using either MERRA or
617 HALOE temperatures, we decided not to include in the merging process any SAGE II O_3 values
618 after June 30, 2000 for pressures equal to or less than 3.2 hPa. SAGE II ozone is not significantly
619 affected by the conversion to mixing ratio/pressure coordinates at 4.6 and 6.8 hPa (Fig. 17).

620 ## 5.2 GOZCARDS ozone merged data records

621 ### 5.2.1 Methodology for GOZCARDS merged ozone

622 Ozone measurements from SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS and ACE-FTS,
623 were used to establish a near-continuous monthly zonal mean record from late 1979 through
624 2012 for the GOZCARDS merged O_3 product (ESDR version 1.01). The SAGE II dataset was
625 used as a reference standard, since it has the longest period of measurements and has been
626 extensively validated. A GOZCARDS ozone merged data record is constructed by combining
627 these measurements after removing systematic biases with respect to SAGE II. This is done by
628 applying additive offsets to all other instrument series, as determined from average differences
629 between monthly zonal means and SAGE II during overlap time periods. The merged data are
630 then derived by averaging all available adjusted datasets. Because there are gaps in overlap
631 between SAGE II and ACE-FTS monthly mean data in some latitudes (Fig. S7), and as SAGE II
632 ozone VMRs obtained from the vertical grid transformation were affected by anomalous NCEP
633 temperatures after mid-2000 for pressures smaller than or equal to 3.2 hPa, a two-step approach

634 is used to generate the merged product. First, SAGE II data are used as reference for pressures
635 larger than 3.2 hPa to adjust HALOE, UARS MLS and Aura MLS based on overlapping months
636 between 1991 and Nov. 2005; see the method overview schematic in Fig. 18. For $p \leq 3.2$ hPa,
637 SAGE II O_3 is still used as a reference through June 2000, and HALOE and UARS MLS data are
638 adjusted accordingly. This eliminates the effect of anomalous NCEP temperatures on SAGE II
639 ozone and leads to more accurate offsets based on HALOE values, after they have been adjusted
640 to SAGE II. Adjusted HALOE data (HALOE* in Fig. 18) are then used as a reference to derive
641 estimated offsets for Aura MLS O_3, using the overlap period with HALOE from Aug. 2004 to
642 Nov. 2005. In step 2, a new reference value is derived by averaging all available data from
643 SAGE II, HALOE*, UARS MLS* and Aura MLS*. This value is used to adjust ACE-FTS ozone
644 based on all overlapping months between March 2004 and Nov. 2005. By including Aura MLS
645 in the dataset created in step 1, we obtain more complete spatial and temporal coverage than
646 possible with SAGE II and HALOE, and ensure that there are overlapping months between this
647 combined dataset and ACE-FTS source data. At the end of step 2, the final merged ozone is
648 derived by averaging the temporary merged dataset from step 1 with the adjusted ACE-FTS data.

649 **5.2.2 Further considerations regarding GOZCARDS merged ozone data**

650 Even in the absence of diurnal variations, measurements from occultation sensors can yield
651 larger sampling errors than those from densely-sampled emission measurements (Toohey et al.,
652 2013). Diurnal changes in ozone can affect data comparisons and could impact data merging.
653 Recently, Sakazaki et al. (2013) presented diurnal changes measured by the Superconducting
654 Submillimeter-Wave Limb-Emission Sounder (SMILES) and Parrish et al. (2014) analyzed
655 ground-based microwave O_3 profile variations versus local time in conjunction with satellite
656 data. Ozone diurnal variations range from a few percent in the lower stratosphere to more than
657 10% in the upper stratosphere and lower mesosphere (see also Ricaud et al., 1996; Haefele et al.,
658 2008; Huang et al., 2010). SAGE II and other occultation instruments observe ozone at local
659 sunrise or sunset, and the retrieved values are generally closer to nighttime values in the upper
660 stratosphere and mesosphere. To characterize systematic differences between satellite data,
661 coincident profiles with small differences in space and time are most often used; an example of
662 mean differences and standard deviations between SAGE II and Aura MLS using both
663 coincident profile and zonal mean methods is provided in Fig. S11. SAGE II and coincident

Aura MLS nighttime O_3 values agree within ~5% between 0.46 and 100 hPa, except in the tropical lower stratosphere where comparisons are noisier. Differences between zonal mean SAGE II and Aura MLS data are very close to differences from averaged coincident values, except for pressures less than 2 hPa, where differences increase from a few to ~10% at 0.3 hPa, consistent with what one expects from the diurnal cycle. Although zonal mean differences are likely to be less representative of "true" differences, by combining SAGE II with Aura MLS data adjusted by zonal mean biases, we provide a series adjusted to the average of sunrise and sunset, as measured by SAGE II. If Aura MLS data were adjusted by biases obtained using the coincident method, an upper stratospheric offset of more than several percent and artificial trends due to such a diurnal cycle effect could be introduced. The use of long-term datasets with consistent sampling should be an advantage for trend detection, even in a region with diurnal changes. Also, our avoidance of SAGE II upper stratospheric O_3 after mid-2000 mitigates potential artifacts arising from changing SAGE II sunrise/sunset sampling patterns over time.

Figure 19 displays the average ozone offsets obtained from the biases relative to SAGE II. A high bias in upper stratospheric ACE-FTS O_3 relative to other datasets is evident from the negative ACE-FTS offsets (as large as 25%). Most of the other instrument offsets are in the 5-10% range; lowering O_3 from UARS MLS, HALOE, and Aura MLS in the lower mesosphere is required to match SAGE II. Sampling differences and data sparseness may be mostly responsible for larger offsets at high latitudes; in these regions, the merged data are less amenable to long-term analyses because of data gaps and larger variability (especially prior to 2004).

As shown in the Supplement (Fig. S12), we observe strong similarities (e.g., peaks at midlatitudes near 10 and 1.5 hPa) in the O_3 annual cycle amplitude patterns from SAGE II, HALOE, ACE-FTS, and Aura MLS over their respective measurement periods. Middle stratospheric peaks are a result of the annual cycle in oxygen photolysis, whereas temperature variations drive the annual cycle in the upper stratosphere (Perliski et al., 1989). This sort of comparison provides some reassurance regarding the consistency of various datasets. Figure 20 provides diagnostics similar to those given for HCl and H_2O, namely correlation coefficients and significance ratios for the slopes of the deseasonalized anomaly time series from SAGE II versus HALOE as well as from ACE-FTS versus Aura MLS (for 1992 through 1999, and 2005 through 2009, respectively). These diagnostic results for ACE-FTS and Aura MLS are of a quality that is

694 comparable to the HALOE/SAGE II results; poorer fits occur mostly at high latitudes and in the
695 upper stratosphere. Poorer correlations at upper altitude appear largely tied to a decrease in the
696 amount of valid data in this region (especially at high latitudes), coupled with a relatively small
697 variability. For regions with poorer agreement between ACE-FTS and Aura MLS, we often see
698 small variability in the series from Aura MLS but larger changes (scatter) in the ACE-FTS series.
699 Larger differences in trends between SAGE II and HALOE were noted by Nazaryan et al. (2005)
700 at low latitudes near 50 km; this is also indicated by our simple linear fits (not shown here) to the
701 GOZCARDS source datasets from these two instruments and the existence of poorer agreements
702 in Fig. 20 for the slope of the difference series in that region. The existence of good correlations
703 in interannual ozone variations between a large number of satellite measurements was discussed
704 by Tegtmeier at el. (2013). Regarding temporal drifts, Nair et al. (2012) have shown that small
705 drifts (mostly within about ±0.5%/yr for the 20-35 km region) exist between most of the datasets
706 from six ozone lidar sites and coincident HALOE, SAGE II, and Aura MLS measurements;
707 similar results were obtained by Kirgis et al. (2013). Other recent studies (in particular, by
708 Hubert et al., 2015) corroborate the very good stability of the datasets used for GOZCARDS,
709 which relies most heavily on O_3 data from SAGE II and Aura MLS. While we feel justified in
710 the use of the longer-term time series chosen for GOZCARDS O_3, data users should still note the
711 existence of a few regions with poorer correlations or trend agreement (and, therefore, larger
712 uncertainties) between different satellite ozone datasets, as indicated in Fig. 20. Long-term
713 merged datasets from GOZCARDS and other sources should undergo continued scrutiny from
714 the community, as done recently for trends by Tummon et al. (2015) and Harris et al. (2015).
715 Sample cross-sectional views of two slices through the GOZCARDS merged O_3 field are
716 provided in the Supplement (Fig. S13). Figure 21 shows estimated systematic errors from our
717 calculation of the 95% ranges for the monthly mean source data used here, both above and below
718 the merged values. In this case, as SAGE II is used as a reference dataset, the applied offsets
719 (Fig. 20) correlate quite well with this plot depicting the ranges about SAGE II values. Minimum
720 error bars can be slightly lower than 5% for the middle stratosphere at low latitudes, where ozone
721 values are largest. This view of systematic error bars is consistent with results by Tegtmeier et al.
722 (2013), based on the larger set of data analyzed for the SPARC Data Initiative. They also found

723 that the regions with lowest errors (scatter) are in the middle stratosphere at low to mid-latitudes,
724 where most monthly mean satellite data fit within ±5% of the multi-instrument mean.

725 ## 5.3 GOZCARDS ozone sample results and discussion

726 Nair et al. (2013) used regression analyses to compare profile trend results from GOZCARDS
727 merged O_3 at northern midlatitudes versus a combined O_3 dataset from lidar and coincident
728 satellite data at the Observatoire de Haute Provence (OHP), France. They showed that good
729 consistency exists for the decreasing ozone time period, from the early 1980s to 1997, and for the
730 upper stratospheric increase since 1997, but some differences exist in the lower stratosphere
731 during this second time period, when the GOZCARDS results show a near-zero trend in
732 comparison to small positive trends from the combined (and more localized) dataset. The above
733 results for the declining time period agree broadly with earlier work (for the 1979-1997 period)
734 by Jones et al. (2009). Gebhardt et al. (2014) analyzed ozone profile trends from SCIAMACHY
735 on ENVISAT, and compared this to trends from Aura MLS, Optical Spectrograph and InfraRed
736 Imager System (OSIRIS) on the Odin satellite, and sondes; their results include the detection of
737 localized ozone increases in the mid-stratosphere at low latitudes; see also Bourassa et al. (2014),
738 who analyzed merged SAGE II and (OSIRIS) observations for 1984-2013, as well as results
739 from Kyrölä et al. (2013) on combined SAGE II and Global Ozone Monitoring by Occultation of
740 Stars (GOMOS) records for 1984-2012, and Eckert et al. (2014), who investigated ENVISAT
741 MIPAS trends for 2002-2012. The shortness of data records since 1997, coupled with relative
742 variability and potential drifts between various measurements may explain some differences in
743 recent trend results, notably for the post-1997 period. More comprehensive analyses from the
744 SI^2N initiative have focused on an intercomparison of profile changes from a variety of datasets,
745 including GOZCARDS and other merged records (Tummon et al., 2015; Harris et al., 2015).
746 Here, we investigate ozone column results based on the global GOZCARDS data, given the
747 work by Ziemke and Chandra (2012), hereafter referenced as ZC12. These authors analyzed total
748 column and stratospheric column data from satellites, and their analyses yielded a rather strong
749 near-global (60°S-60°N) average ozone increase since 1998. Their stratospheric columns
750 depend on the convective-cloud differential (CCD) method and use Total Ozone Mapping
751 Spectrometer (TOMS) and Ozone Monitoring Instrument (OMI) column data over convective

clouds near the tropopause (see also Ziemke et al., 2005). In Fig. 22, we compare changes in 60°S-60°N ZC12 column ozone data (J. Ziemke, personal communication, 2013) to changes in GOZCARDS O_3 columns above 68 hPa for that region; note that GOZCARDS values do not provide for a continuous long-term time series down to pressures of 100 hPa or more in the SAGE I years (1979-1981). To eliminate biases between stratospheric columns as calculated using the CCD methodology and the GOZCARDS fixed bottom pressure approach, we reference all stratospheric columns to the 1980 total column value. These column series include SAGE I data and are linearly interpolated between 1981 and 1984, when no GOZCARDS source datasets exist. We observe that relative changes in GOZCARDS columns follow the ZC12 curves within a few DU in the downward phase until about 1992, but the 1992-1997 decrease in total columns does not compare very well. Some of this discrepancy may occur because total columns capture a stronger decrease from levels below 68 hPa, not fully represented in GOZCARDS. Focusing on the late period (from Aura MLS and ACE-FTS), we also show the GOZCARDS columns above 68 hPa, referenced to 2007 instead of 1980. There is a good match in the variations between GOZCARDS and ZC12 columns during 2005-2010, in agreement with the fact that very good correlations were obtained by ZC12 between Aura MLS columns and stratospheric column data from the CCD technique. ZC12 values for stratospheric and total columns are in good agreement, although the stratospheric values have gaps when not enough data were present for near-global estimates. The increase in ZC12 data from 1997 to 1998 is not matched very well by GOZCARDS; this is also true if we remove the 11-yr solar cycle from both datasets (not shown here), as done by ZC12. We also note that recent analyses by Shepherd et al. (2014), who used a chemistry-climate model constrained by meteorology to investigate causes of long-term total column O_3 variations, show a partial return, in 2010, towards 1980 ozone column values, but not nearly as much as implied by ZC12. We note that long-term halogen source gas reductions that have occurred since the mid-1990s should only lead to column ozone increases of a few DU since 1997 (Steinbrecht et al., 2011). It is likely that the discrepancies seen here lie in the various datasets and their merging; for example, it would be worthwhile to check if homogenized SBUV column O_3 data show results that are substantially different from those of ZC12. Discrepancies could also arise from differences in ozone column calculations or coverage, because of different methodologies, grids or sampling to properly determine near-global results.

6 Other GOZCARDS data records

We now briefly mention the N_2O, HNO_3, and temperature GOZCARDS records that were part of the delivery for public dissemination in 2013. For N_2O and HNO_3, the somewhat simpler merging procedure consisted of averaging the source datasets from ACE-FTS and Aura MLS over the overlap time period (Aug. 2004 through Sep. 2010) to obtain the additive offsets for each of the two individual records. We then simply used the correspondingly-adjusted and averaged series to create the merged results; this procedure is the same as we described for the first step in the HCl (or H_2O) merging process.

6.1 N_2O

This data set starts in August 2004, when the Aura MLS data record began; the only dataset after Sep. 2010 is the Aura MLS N_2O (version 3.3) data record. Because of degradation in the main target MLS N_2O band (near 640 GHz) after the first few months of 2013, the N_2O standard MLS product is being reprocessed for the whole Aura MLS period using an alternate measurement band; currently, there are no official GOZCARDS N_2O data after 2012.

Excellent agreement (mostly within 5%) exists between stratospheric ACE-FTS and Aura MLS N_2O profiles (see Lambert et al., 2007; Strong et al., 2008; Livesey et al., 2013). Plots showing the average offsets applied to both MLS and ACE-FTS N_2O series as a function of latitude and pressure are provided in Fig. S14. These plots are in agreement (in magnitude and in sign) with the above-referenced studies; the two datasets yield typical offsets (one half of the average differences) of less than 5%. Also, very good temporal agreement between these two time series (for 2004-2010) is illustrated by the quality of the N_2O diagnostic information displayed in Fig. S15, showing generally highly correlated fields and insignificant drifts.

Figure 23 shows sample contour plots for the N_2O merged field (2004-2012); as seen from the bottom panel (100 hPa), wintertime descent brings low N_2O values down at high latitudes inside the polar vortices. N_2O is a conserved tracer in the lower stratosphere and its variations near the tropopause have implications regarding age of air. Variations in upper stratospheric N_2O are clearly affected by seasonal and dynamical effects; this is evident from the striking semi-annual, annual and QBO-related patterns displayed in Fig. 23 for the 6.8 hPa level (top panel).

6.2 HNO_3

As for N_2O, we merged the HNO_3 data from ACE-FTS (version 2.2) and Aura MLS (version 3.3) from Aug. 2004 onward, and included only the adjusted MLS dataset after Sep. 2010. The average offsets applied to MLS and ACE-FTS time series as a function of latitude and pressure for HNO_3 are provided in Fig. S16. The typical offsets (one half of the average differences) for HNO_3 are less than ~10% (and less than 0.5 ppbv). Despite somewhat larger percent absolute differences than for N_2O between Aura MLS and ACE-FTS HNO_3, there is very good agreement as a function of time between these two datasets in the stratosphere. This is illustrated by the HNO_3 diagnostic information provided in Fig. S17; the poorest correlations are obtained at or below the tropical tropopause.

Comparisons of v3.3 Aura MLS and v2.2 ACE-FTS nitric acid profiles have shown good agreement (see also Livesey et al., 2013), as the MLS HNO_3 v3.3 values are now generally larger than in v2.2, for which validation results were provided by Santee et al. (2007). Wolff et al. (2008) also compared MLS (v2.2) and ACE-FTS (v2.2) coincident profiles, and obtained similar results; in addition, they demonstrated that very good agreement exists between the HNO_3 profiles from ACE-FTS and coincident profiles from MIPAS on Envisat. Also, comparisons between Aura MLS HNO_3 (v3.3) profiles and wintertime HNO_3 profiles retrieved by a Ground-based Millimeter-wave Spectrometer (GBMS) in Thule, Greenland, during the first 3 months of 2010, 2011, and 2012 show agreement mostly within 10-15% (Fiorucci et al., 2013).

Figure 24 (top two panels) displays the HNO_3 fields at 46 hPa from the UARS MLS period (1991-1997) as well as from the 2004-2013 period, for which a merged GOZCARDS product was produced, based on Aura MLS and ACE-FTS source datasets. Also shown (bottom two panels) are time series for 45°N and 32 hPa from both these periods; the bottom right panel includes the source and merged time series. We have performed additional investigations (not shown here) which lead us to believe that small upward adjustments to the UARS MLS HNO_3 values (by about 10%) are needed to better cross-correlate these datasets across the two distinct time periods; such relative biases are within the expected systematic errors. This is based on a consideration of ground-based Fourier Transform infrared column HNO_3 data covering the full time period, as well as past GBMS HNO_3 profile retrievals. Also, Aura MLS and ACE-FTS

HNO_3 data match ground-based and other correlative data quite well, and typically better than the intrinsically poorer quality UARS MLS HNO_3 data. However, obtaining an optimum global set of adjustments for the UARS MLS nitric acid field will be limited by the number of sites with such ground-based data as well as by the different vertical resolutions for these datasets versus MLS. More collaborative work regarding such analyses is needed in order to find the optimum adjustments to help tie together these two time periods for this species. Although we did not deliver the UARS MLS HNO_3 source data files for GOZCARDS, we could provide these monthly zonal mean series upon request, keeping the above caveats in mind.

6.3 Temperature

Finally, in terms of the initial set of delivered GOZCARDS products, and for the convenience of stratospheric composition data users, we have used temperatures (T) from the Modern-Era Retrospective Analysis for Research and Applications (MERRA) to produce a monthly mean GOZCARDS temperature data set from 1979 onward. MERRA is a NASA Goddard reanalysis (Rienecker et al., 2011) for the satellite era using Goddard Earth Observing System Data Assimilation System version 5 (GEOS-5); T is from the DAS 3d analyzed state MAI6NVANA, version 5.2 files (such as MERRA300.prod.assim.inst6_3d_ana_Nv.20110227.hdf). Data from four daily MERRA files (for 00, 06, 12, and 18 hr UT) were averaged to provide daily mean temperature fields (appropriate for a mean time of 09 hr). Vertical interpolation was performed onto the GOZCARDS pressure grid, which, for temperature, covers 30 pressures levels from 1000 hPa to 0.0147 hPa. Averaged values were stored for the 10° GOZCARDS latitude bins, and daily results were binned to create the GOZCARDS monthly temperature data set (version 1.0).

7 Summary and conclusions

We have reviewed the GOZCARDS project's production of merged data records of stratospheric composition, mainly for HCl, H_2O, and O_3, using carefully screened satellite data, starting in 1979 with SAGE I and continuing through Aura MLS and ACE-FTS data periods. The source data have a high degree of maturity and we have reinforced our confidence in their usefulness through investigations of various diagnostics (offsets, annual cycles, correlations and trend differences of deseasonalized series). These records are publicly available as GOZCARDS ESDR version 1.01 and can be referenced using DOI numbers (Froidevaux et al., 2013b,

868 Anderson et al., 2013, and Wang et al., 2013, for the above species, respectively). The other
869 GOZCARDS data records also have references, namely Schwartz et al. (2013) for the MERRA-
870 based temperature records, and Froidevaux et al. (2013c, 2013d) for N_2O and HNO_3,
871 respectively. Table 2 provides a summary of the GOZCARDS monthly mean datasets. Yearly
872 netCDF files are available for public access (http://mirador.gsfc.nasa.gov). The merging
873 methodology follows from a determination of mean biases (for each pressure level and 10°
874 latitude bin) between monthly mean series, based on the overlap periods. For ozone, SAGE II
875 data are the chosen reference, whereas for other species, the merging basis is equivalent to an
876 average of the datasets during the periods of overlap. The merged data files contain the average
877 offset values applied to each source data time series, along with standard deviations and standard
878 errors. The GOZCARDS README document (Froidevaux et al., 2013a) provides more details
879 about data file quantities, including local time and solar zenith angle information, and a list of
880 days with available data. We also display here estimated systematic errors about the merged
881 values; we find that mixing ratio errors are typically within 5% to 15% and are consistent with
882 the magnitude of observed relative biases.

883 The GOZCARDS HCl merged record in the upper stratosphere enables long-term tracking of
884 changes in total stratospheric chlorine. The long-term increase in HCl prior to the late 1990s, and
885 the subsequent gentler decrease in the 21st century, are delayed manifestations of changes in the
886 sum of the surface source gas abundances as a result of regulations from the Montreal Protocol
887 and its amendments. From 1997 to 2010, the average rate of change in upper stratospheric HCl
888 (50°S to 50°N) was about -0.4 to -0.7%/yr (with the smaller rates of decrease after 2003). In the
889 lower stratosphere, where Aura MLS data are weighted heavily, recent short-term variations
890 have shown a flattening out and, in particular for northern midlatitudes and at 50-70 hPa for the
891 deep tropics, a significant reversal and increasing trend (see also Mahieu et al., 2014), compared
892 to the decrease from the late 1990s to about 2004. However, lower stratospheric HCl tendencies
893 appear to be reversing again in recent years (2011-2014), with decreases at northern
894 midlatitudes and some increasing tendencies at southern midlatitudes. In the future, we expect to
895 see long-term global HCl decreases in both the upper and lower stratosphere.

896 For water vapor, we have used data from the same instruments as for HCl, with the same
897 methodology, except for the addition of 1991-1993 UARS MLS data. The H_2O data record

898 shows large mesospheric variations that are anti-correlated with the solar flux over the past two
899 11-yr solar cycles. Net long-term trends in lower stratospheric H_2O are quite small if one
900 considers the past 22 years, but there has been considerable interannual variability, including the
901 steep drop from 2000 to 2001, as mentioned in past work. While H_2O tendencies have been
902 generally positive after 2001, the 68 and 100 hPa levels show some steep decreases (by 0.5-0.8
903 ppmv) from 2011 to 2013 (see also Urban et al., 2014). Over the past 22 years, long-term global
904 H_2O increases of order 10% are observed in the upper stratosphere and lower mesosphere,
905 whereas a decrease of nearly 10% has occurred in the lower stratosphere (near 70-100 hPa).
906 However, there is no regular monotonic change on decadal timescales, especially in the tropical
907 lower stratosphere, where fairly sharp decreases followed by steadier increases may be a
908 recurrent pattern (see also Fueglistaler, 2012); this complicates the detection of any small
909 underlying trend. As one might expect from the well-documented temperature influence on the
910 tropical lower stratosphere, H_2O variability (based on maximum minus minimum yearly
911 averages) is largest in the tropics and just above the tropopause. More accurate studies of
912 seasonal to decadal water vapor variability will be enabled by continuing such merged H_2O
913 datasets in the future. A reduction in model spread for stratospheric H_2O is likely easier to
914 achieve than tighter upper tropospheric model results; for the upper troposphere, see the
915 data/model comparisons (H_2O and ice water content) by Jiang et al. (2012).

916 For ozone, we have used measurements from SAGE I, SAGE II, HALOE, UARS MLS, Aura
917 MLS and ACE-FTS to produce a merged record starting in 1979, after adjusting the series to
918 SAGE II. We observed temporal drifts in the SAGE II series, after conversion to the
919 GOZCARDS mixing ratio/pressure grid, as a result of the NCEP temperature data used in this
920 conversion, mostly in the upper stratosphere after June 2000 (see also McLinden et al., 2009). To
921 mitigate this issue, we used HALOE upper stratospheric O_3 as a reference for July 2000 to
922 November 2005, after adjusting the HALOE series to SAGE II. The resulting GOZCARDS
923 merged O_3 data for northern midlatitudes have been used in regression analyses (Nair et al.,
924 2013) to reveal decreases in the whole stratosphere for 1984-1996. Nair et al. (2015) extended
925 this work and found increasing trends in upper stratospheric GOZCARDS O_3 since 1997, but no
926 significant positive trends in the lower stratosphere. Other studies of GOZCARDS O_3 profile
927 trends have been discussed as part of the WMO (2014) and SI^2N assessments (Tummon et al.,

928 2015; Harris et al., 2015). Here, we looked at the consistency of column data between
929 stratospheric GOZCARDS O_3 and work by Ziemke and Chandra (2012), who noted that a fairly
930 rapid change ("recovery'') in near-global ozone columns from TOMS and OMI could be inferred
931 since the mid-1990s. We show that the similarly analyzed GOZCARDS column data does not
932 show an upturn of more than 0.5-1% since that period. Reasons for these differences could
933 include data coverage or merging-related issues in either dataset, or inaccuracies in globally-
934 averaged stratospheric columns. A recent global total ozone study (Shepherd et al., 2014) also
935 points to less of a return towards 1980 levels than implied by ZC12.

936 We also briefly described the creation of N_2O and HNO_3 GOZCARDS data records, based on
937 Aura MLS and ACE-FTS. The agreement between these two instruments' datasets for these
938 species was shown to be generally very good. For HNO_3, UARS MLS HNO_3 source datasets in
939 the GOZCARDS format are available from the authors. However, a small upward adjustment (of
940 order 10%) to the UARS MLS values is likely needed based on our preliminary work comparing
941 these series to HNO_3 column results from FTIR measurements. More detailed work should help
942 determine if global adjustments can indeed be made to UARS MLS HNO_3 data; lacking this, one
943 should ensure that error bars reflect likely biases that can affect the continuity between HNO_3
944 datasets before and after 2000, given the multi-year gap in satellite coverage for this species.

945 **There is a Supplement related to this article.**

946 *Acknowledgements.* Work at JPL was performed under contract with the National Aeronautics
947 and Space Administration (NASA). We dedicate this work to the memory of Professor Derek
948 Cunnold (Georgia Institute of Technology) who was a member of the original NASA
949 MEaSUREs (Making Earth System Data Records for Use in Research Environments)
950 GOZCARDS proposal. The GOZCARDS data generation could not have been possible without
951 the past work from instrument teams for SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS,
952 and ACE-FTS, and related data usage documentation. At JPL, we thank Joe Waters for his
953 leadership role in making MLS instruments and datasets possible and Bill Read for his key role;
954 thanks to Vince Perun for MERRA-related work, and to Brian Knosp and Robert Thurstans for
955 database and computer management assistance. We also thank Kaley Walker and Ashley Jones
956 for comments regarding ACE-FTS data, Gloria Manney and William Daffer for help in making

957 the original ACE-FTS data profiles available, and Joe Zawodny and Larry Thomason for their
958 contributions and comments regarding SAGE data. We acknowledge the work of the GMAO
959 team responsible for MERRA data used to generate the GOZCARDS temperatures, specifically,
960 Steven Pawson and Jianjun Jin for discussions and cross-checks regarding temperature data. We
961 acknowledge Jerry Ziemke for the ozone column data (from Ziemke and Chandra, 2012), and
962 Sean Davis for discussions on data usage and screening, and the creation of long-term series. For
963 early HNO_3-related work connecting ground-based data to MLS datasets, we thank Giovanni
964 Muscari and Irene Fiorucci. We are thankful for the NOAA Earth System Research Laboratory
965 (ESRL) Global Monitoring Division (GMD) website information and data on total surface
966 chlorine. We obtained solar flux data for the Ottawa/Penticton sites from the NOAA National
967 Geophysical Data Center (NGDC) website (www.ngdc.noaa.gov), for which we also
968 acknowledge the National Research Council of Canada.

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Appendix A

A.1. GOZCARDS data provenance

The general origin of the datasets is summarized here. Data coverage from limb sounders (including the instruments used here) is displayed nicely in the work by Toohey et al. (2013).

SAGE I

SAGE I was launched February 18, 1979, aboard the Applications Explorer Mission-B (AEM-B) satellite. SAGE I was a sun photometer using solar occultation (Chu and McCormick, 1979), and it collected a global database for nearly three years on stratospheric aerosol, O_3, and NO_2. For more information, the reader is referred to http://sage.nasa.gov/SAGE1.

SAGE II

SAGE II was launched aboard the Earth Radiation Budget Satellite (ERBS) in October 1984 and its data gathering period ended in August 2005. During each sunrise and sunset, SAGE II measured stratospheric aerosols, O_3, NO_2, and H_2O via solar occultation. This long dataset has proven very valuable in determining past ozone trends. For more information on and data access to the (V6.2) dataset used for GOZCARDS, the reader is referred to http://sage.nasa.gov/SAGE2.

HALOE

Since its launch on September 12, 1991 from the Space Shuttle Discovery until November 2005, UARS HALOE collected profiles of atmospheric composition and temperature. HALOE (Russell et al., 1993) used solar occultation to measure vertical profiles of O_3, HCl, HF, CH_4, H_2O, NO, NO_2, temperature, aerosol extinction, and aerosol composition and size distribution. More information and access to the HALOE data can be obtained from http://haloe.gats-inc.com and http://disc.sci.gsfc.nasa.gov/UARS/data-holdings/HALOE. For GOZCARDS purposes, we have used Version 19 HALOE netCDF data files available at http://haloe.gats-inc.com.

UARS MLS

This instrument observed the Earth's limb in microwave emission using three radiometers, at frequencies near 63, 183 and 205 GHz (Waters, 1993; Barath et al., 1993), providing unique

daily global information on stratospheric ClO, along with other profiles, including O_3, H_2O, HNO_3, temperature, and cloud ice water content. The stratospheric H_2O data ceased on April 15, 1993, after the failure of the 183 GHz radiometer. After March 15, 1994, measurements became increasingly sparse in order to conserve the life of the MLS antenna scan mechanism and UARS power. Data exist until July 28, 1999, although for GOZCARDS, only data through mid-June 1997 are used, as data sparseness and degradation of the 63 GHz radiometer led to less 'trend-quality' data after this. Sampling patterns follow the alternating yaw cycles imposed on MLS by the precessing UARS orbit; MLS measurements were obtained continuously for all latitudes between 34°S and 34°N, with higher latitudes covered in either the northern or southern hemisphere with a roughly 36-day cycle. Livesey et al. (2003) provide more information on the UARS MLS instrument, retrievals, and results. For data access, the reader is directed to the relevant Goddard Earth Sciences and Information Services Center (GES DISC) data holdings at http://disc.sci.gsfc.nasa.gov/UARS/data-holdings/MLS. L3AT data files were used as the basis for the production of the GOZCARDS UARS MLS monthly source datasets.

ACE-FTS

ACE-FTS is the primary instrument onboard the SCISAT satellite, launched on August 12, 2003. It is a high spectral resolution (0.02 cm^{-1}) Michelson interferometer operating from 2.2 to 13.3 μm (750-4400 cm^{-1}); see Bernath et al. (2005) for an overview of the ACE mission. The instrument can simultaneously measure temperature and many trace gases (including all the species mentioned here for GOZCARDS), thin clouds, and aerosols, using the solar occultation technique. ACE-FTS data version 2.2, along with the version 2.2 update for ozone, were used here for GOZCARDS. For access to the public ACE-FTS datasets, with a routine measurement start date of March 2004, the reader is directed to http://www.ace.uwaterloo.ca.

Aura MLS

MLS is one of four instruments on NASA's Aura satellite, launched on July 15th 2004. Aura MLS is a greatly enhanced version of the UARS MLS experiment, providing better spatial coverage, vertical resolution, and vertical range, along with more continuous data over its lifetime (and with ongoing measurements at the time of writing). The instrument includes

radiometers at 118, 190, 240, and 640 GHz, and a 2.5 THz module (Waters et al., 2006). Aura MLS provides measurements of many chemical species, cloud ice, temperature and geopotential height. Continuous measurements have been obtained since August 2004, with the exception of OH, for which sparser measurements exist since August 2010, in order to preserve the life of the THz module. For more information and access to the Aura MLS datasets, the reader is referred to http://disc.sci.gsfc.nasa.gov/Aura/data-holdings/MLS. For GOZCARDS, we use the currently recommended Aura MLS data versions (version 2.2/2.3 for ozone and 3.3/3.4 for other species).

A.2. Calculation details for the iterative merging procedure

Given three time series, the merging procedure that we use first combines two out of the three time series, $y_1(i)$ and $y_2(i)$ (where index i represents time for each monthly mean value in a given latitude/pressure bin). We first obtain the temporary merged series $m_1(i)$ via:

$$m_1(i) = (1/2)\,(y_1(i) + y_2(i)) \quad (1)$$

with the average offsets for $y_1(i)$ and $y_2(i)$ being $(1/(2\,n_{12}))\,\Sigma\,(y_1(i) - y_2(i))$ and -1 times this value, respectively; n_{12} is the number of overlapping data points between the two time series. Then, we merge together the time series $m_1(i)$ and $y_3(i)$, keeping the weightings equal for all 3 time series (1/3 for each), so that we calculate the new merged time series $m(i)$ via:

$$m(i) = w_m\, m_1(i) + w_3\, y_3(i) = (1/3)\,(y_1(i) + y_2(i) + y_3(i)) \quad (2)$$

which will hold if the weights are $w_m = 2/3$ and $w_3 = 1/3$ (given equation (1) for $m_1(i)$). The average reference value (to which the adjustments of $m_1(i)$ and $y_3(i)$ in the 2nd step are made) is given by $(1/n_m)\,\Sigma\,(\,(2/3)\,m_1(i) + (1/3)\,y_3(i))$, where n_m represents the number of (overlapping pairs of) data values used in step 2. For the HCl and H_2O data merging procedure, we always use the Aura MLS time series as one of the first two series involved in the initial merging step, for example as $y_1(i)$, in order to maximize the overlap between the first two series and obtain more robust offset values. Then, we use the 3rd time series; the order used for HALOE and ACE-FTS (i.e., whether we use HALOE or ACE-FTS for y_2 or y_3) makes very little difference.

Calculation of the standard deviation for the merged data values

The average and standard deviation (square root of variance) for each y_k value (i.e. for each monthly zonal mean in a particular lat/p bin) are calculated from equations (3) and (4) below:

$$\bar{y}_k = \frac{1}{n_{yk}} \sum_j y_{kj} \tag{3}$$

and, for the variance,

$$\sigma_{yk}^2 = \frac{1}{n_{yk} - 1} \sum_j (y_{kj} - \bar{y}_k)^2 \tag{4}$$

where index "*j*" corresponds to individual data values within a month, index *k* represents a given instrument (data source), and *n* is the total number of data values for a given bin and source (instrument) time series point in time (or month). Each value $\overline{y_k}$ above is a monthly average (although we also use instead the simpler notation y_k), with standard deviation about the mean σ_{yk}. Now, given the merged series *u(i)* (where index *i* runs over a large number of months), the standard deviation of each merged data point (for a given month) can be obtained by considering the original datasets y_{kj} that were used to construct ***u*.** Specifically, we have the variance for the merged dataset

$$\sigma_u^2 = \frac{1}{n_u - 1} \sum_j (u_j - u_{ref})^2 \tag{5}$$

where u_{ref} is the merged value (which is not necessarily chosen to be the average value $\bar{u}$) and the u_j values represent the union of adjusted data values that make up the merged product, with the index j for this combined dataset covering all values (up to the total n_u) obtained from the original source values y_{kj}. In practice, we do not keep track of the individual data values that went into making the averages for the series y_k that are being merged, and we need to obtain σ_u based solely on the values $\bar{y}_k$, σ_{yk}, and the original number of points for each dataset y_k, namely n_{yk}. If we consider all the original values, we have a combined dataset with n_u points, such that $n_u = \sum_k n_{yk}$. Now, expanding equation (5), we get

$$(n_u - 1)\, \sigma_u^2 = \sum_j (u_j^2 + u_{ref}^2 - 2\, u_{ref}\, u_j) \tag{6}$$

or

$$(n_u - 1)\ \sigma_u^2 = \sum_j u_j^2 + n_u\, u_{ref}^2 - 2\, u_{ref} \sum_j u_j \qquad (7)$$

Expanding (4) for each individual dataset y_k, we get

$$(n_{yk} - 1)\ \sigma_{yk}^2 = \sum_j y_{kj}^2 + \bar{y}_k{}^2 - 2\bar{y_k} \sum_j y_{kj} \qquad (8)$$

which leads to

$$\sum_j u_j^2 = \sum_{k,j} y_{kj}^2 = \sum_k (n_{yk} - 1)\, \sigma_{yk}^2 \quad + \sum_k n_{yk}\ \overline{y_k}^2, \qquad (9)$$

so that extracting the variance from equation (7) now leads to

$$\sigma_u^2 = \frac{1}{(n_u - 1)} \left(\sum_k (n_{yk} - 1) \sigma_{yk}^2 + \sum_k n_{yk}\ \overline{y_k}^2 + n_u\, u_{ref}^2 - 2 u_{ref} \sum_k n_{yk}\, \overline{y_k} \right) \qquad (10)$$

The adjusted time series are obtained from the original series y_k as Y_k, and we can write Equation (4) in the same manner for the Y_k data values, namely

$$\sigma_{Yk}^2 = \frac{1}{n_{yk} - 1} \sum_j (Y_{kj} - \overline{Y_k})^2 \qquad (11)$$

with $\sigma_{Yk} = \sigma_{yk}$ as the adjustments (offsets) are performed in an additive manner; if these adjustments were performed using multiplicative factors, those factors would also have to be considered in a multiplicative way to get the new σ_{Yk} values. We can thus write (10) for the adjusted datasets as:

$$\sigma_u^2 = \frac{1}{(n_u - 1)} \left(\sum_k (n_{yk} - 1) \sigma_{yk}^2 + \sum_k n_{yk}\ \overline{Y_k}^2 + n_u\, U_{ref}^2 - 2 U_{ref} \sum_k n_{yk}\, \overline{Y_k} \right) \qquad (12)$$

Equation (12) for the standard deviation of the merged dataset simplifies if the original datasets are adjusted to exactly the same reference value ref ($\overline{Y_k} = ref$) and the merged value U_{ref} is also equal to that value, as the sum of the last 3 terms in Eq. (10) (with Y_k replacing y_k) then reduces to $n_u\, ref^2 + n_u\, ref^2 - 2\, n_u\, ref^2$, which is zero. In this case, one obtains

$$\sigma_u^2 = \frac{1}{(n_u - 1)} \left(\sum_k (n_{yk} - 1) \sigma_{yk}^2 \right) \quad (13)$$

However, in general, one should use equation (12) for the standard deviation of the merged dataset, given the adjusted datasets $\overline{Y_k}$ and the merged (or reference) value U_{ref}. Also, we often use a merged value equal to the average of the original data (over a given overlap period), so that

$$U_{ref} = \frac{1}{n_y} \sum_k \overline{y_k} \quad (14)$$

where n_y is the total number of datasets (y_k), as opposed to having the merged value place more weight on the larger datasets (e.g., for emission-type measurements versus occultation-type), in which case one would consider using $U_{ref} = \frac{1}{n_u} \sum_k n_{yk} \overline{y_k}$. For ozone, we use a particular dataset (SAGE II ozone) as the primary reference, but equation (12) can be used to obtain the standard deviation for the merged dataset (about the SAGE II reference) in that case also. While it is useful to have the formalism above for obtaining the merged dataset standard deviation σ_u, we often find significant differences between the standard deviations of various datasets, so that this effect will have the greatest influence on the results, as opposed to the impact of the last 3 terms in the summation (in (12)). Finally, it is easy to test equation (12) (and we have done so) by using synthetic series and calculating the standard deviation of the combined set. In reality, the standard deviations of the time series monthly mean values are typically larger for MLS than for ACE-FTS, mainly because of the more complete sampling of variability from the daily global measurements acquired by MLS. Sample plots for standard deviations and standard errors in the case of HCl are shown in Fig. A1. As expected, merged standard deviations follow the standard deviations from HALOE HCl before Aug. 2004 and those from MLS HCl after this time. However, the merged standard errors for the MLS time period follow the smaller MLS standard errors, because these values vary inversely with the square root of the number of values sampled, and are therefore made smaller by the significantly larger daily and monthly MLS sampling rate and coverage.

A.3. Procedural details for GOZCARDS HCl, H_2O, and O_3

Data screening procedures for the GOZCARDS source datasets, following previously described methods, are provided (with references) in Table A1, along with certain species-related specifics. Other GOZCARDS data characteristics and details are provided below for each species.

A.3.1. HCl

- The vertical data range for valid HCl merged values is between 0.46 hPa and 147 hPa (inclusive), as a result of data sparseness or data quality issues outside these ranges.
- At 147 hPa, no merged HCl values exist for latitude bins from 35°S to 35°N inclusive, because of unrealistically large Aura MLS HCl values in this region; also, there is not enough data at this level to provide a meaningful product from HALOE and ACE-FTS data alone.
- Because of occasional small negative merged values during southern hemisphere polar winter, we did not apply HCl data offsets in the lower stratosphere for the 65°S through 85°S bins from June through September and for pressures larger than or equal to 15 hPa. For vertical continuity purposes, we applied this method to all lower stratospheric pressure levels, although the small negative merged values only occurred in a small fraction of cases.
- As Aura MLS and ACE-FTS data exist in the 85°N and 85°S bins, but there are no HALOE measurements, we simply extended the offsets from the adjacent bins (at 75°N and 75°S) to these two bins to obtain a merged record after 2004 that exhibits continuity versus latitude.
- At 100 hPa, we used HCl offsets from the 5°S bin for the 5°N bin, as there was insufficient data from the combined data in the latter bin to calculate meaningful offsets. This procedure seems reasonable, given that the time series in these two adjacent tropical latitude bins (during years outside the 2004/2005 overlap period) look continuous and stable enough to justify identical adjustments in both bins and to avoid a data gap in the merged series at 5°N.

A.3.2. H_2O

- The vertical data range for valid H_2O merged values is between 0.01 hPa and 147 hPa (inclusive). While H_2O data exist at 147 hPa for low latitudes, more careful work would be needed to extend the merged data globally in such a region.

- Users should keep in mind the PMC-related caveats mentioned in Sect. 4 for summer at high latitudes in the upper mesosphere, prior to the end of the HALOE dataset (Nov. 2005).
- As for HCl, we could not use our standard merging procedure at the two most poleward latitude bins; we simply extended the offsets from the adjacent bins (at 75°N and 75°S) to these polar bins to obtain a merged record after 2004 that exhibits continuity versus latitude.
- Also as for HCl, at 100 hPa, we used H_2O offsets from the 5°S bin for the 5°N bin, as there was insufficient data from the combined datasets in the latter bin to calculate meaningful offsets and merge the datasets. This procedure avoids a data gap in the merged series at 5°N.

A.3.3. O_3

- The vertical range for valid O_3 merged data is from 0.2 hPa to 215 hPa (inclusive), with the lower altitude bound varying with latitude; the merged product at 147 and 215 hPa has valid data only for the 35° to 85° latitude bins, with values mostly larger than ~ 0.1 ppmv. The upper troposphere is more of a merging challenge, given smaller abundances, more difficult measurements, and a larger impact from different instrument resolutions. Also, while we suggest (see main text) that GOZCARDS merged ozone data should not be subject to a large impact from diurnal variations, the highest altitude region should be treated with caution.

- SAGE I monthly mean source data are used for the merged dataset in the tropical bins (25°S to 25°N) from 1 through 68 hPa only and, at higher latitudes, from 1 through 100 hPa only.
- We omitted the use of UARS MLS at 100 hPa for low latitudes (from 25°S to 25°N), as these monthly values are biased quite high and also exhibit too large a seasonal cycle amplitude, in comparison to HALOE and SAGE II data; this appears to relate to a UARS MLS artifact.

- Since there is no (monthly) overlap between SAGE II and HALOE versus UARS MLS or Aura MLS in the 85°N and 85°S latitude bins, the same offsets as for 75°N and 75°S (respectively) are applied for these bins, in order to minimize discontinuities.

- Because of discontinuities that appeared in merged O_3 at high latitudes above the stratopause, particularly in the 75°S bin, we flagged merged values for 75° and 85° (N and S) as bad, for pressures less than 1 hPa. This issue could be the result of a few bad data points or not enough data overlap. To minimize artifacts, we left the resolution of this issue for future investigations; also, the reduced amount of occultation data at these high latitudes makes the

1624 usefulness of a merged product with poorly sampled seasonal changes somewhat marginal
1625 (for certain years at least, the number of monthly values drops significantly at high latitudes).
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1628 **Table 1.** Characteristics of instrument datasets used to create GOZCARDS ESDRs (version ev1.01).

Instrument and Data Versions	Platform	Type of measurement	Time period (GOZCARDS source files)	Vertical Resolution (km)	Retrieved quantity and stratospheric vertical grid spacing
SAGE I V5.9_rev O_3	AEM-2	Solar occultation VIS/UV and near-IR	Feb. 1979 - Nov. 1981	1	Density on altitude grid 1 km spacing
SAGE II V6.2 O3	ERBS	Solar occultation VIS/UV and near-IR	Oct. 1984 - Aug. 2005	0.5 - 1	Density on altitude grid 0.5 km spacing
HALOE V19	UARS	Solar occultation mid-IR	Oct. 1991 - Nov. 2005	2.5	Volume Mixing Ratio on pressure grid with 30 levels per decade (LPD) change in p
MLS V5 O_3 V6 H_2O	UARS	Limb emission microwave / sub-mm	Oct. 1991 - June 1997 (May 1993 end for strat. H_2O)	**H_2O** 3 - 4 (strat.) 5 - 12 (mes.)	Volume Mixing Ratio on pressure grid with 6 LPD in stratosphere
				O_3 3.5 - 5 (strat.) 5 - 8 (mes.)	6 LPD in stratosphere
ACE-FTS V2.2 (V2.2 update for O_3)	SCISAT	Solar occultation mid-IR	Mar. 2004 through Sep. 2010 (2009 for O_3)	3 - 4	Volume Mixing Ratio on 1 km grid spacing (height and p provided)
MLS V3.3 V2.2 O_3	Aura	Limb emission microwave / sub-mm	Aug. 2004 through 2012	**HCl** 3 - 5	Volume Mixing Ratio on pressure grid with 6 LPD
				H_2O 3 - 4 (p > 0.1 hPa) 5 - 9 (0.1-0.01 hPa)	12 LPD
				O_3 3	6 LPD

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Table 2. Products and instrument source data making up the available GOZCARDS data records.

Merged Products and pressure range	**Source Datasets** (and years used)
HCl 147 – 0.5 hPa	HALOE (1991-2005), ACE-FTS (2004-2010), Aura MLS (2004 onward) Note: MLS data for p < 10 hPa not used for merged time series
H_2O 147 – 0.01 hPa	HALOE (1991-2005), UARS MLS (1991-1993), ACE-FTS (2004-2010), Aura MLS (2004 onward)
O_3 215 – 0.2 hPa	SAGE I (1979-1981), SAGE II (1984-2005), HALOE (1991-2005), UARS MLS (1991-1997), ACE-FTS (2004-2009), Aura MLS (2004 onward)
HNO_3 215 – 1 hPa	ACE-FTS (2004-2010), Aura MLS (2004 onward)
N_2O 100 – 0.5 hPa	ACE-FTS (2004-2010), Aura MLS (2004 onward)
Temperature 1000 – 0.015 hPa	GMAO MERRA (1979 onward)

Table A1. Data screening procedures and related references used for the source dataset generation.

Instrument	Data Screening Issue / Method	Reference
SAGE I (O3)	Aerosol interference issue: Remove values at altitudes below which the 1 μm extinction > 10^{-3} km^{-1}.	L. Thomason (personal communication, 2012)
SAGE II (O3)	Remove entire profile if any error error value exceeds 10% of VMR (for 30 to 50 km altitude); this occurred mainly in 1993 & 1994 (''short events'').	Wang et al. (2002)
	Use aerosol extinctions and extinction ratios to remove data affected by clouds or by aerosols (from Mt. Pinatubo).	
	Remove anomalously low values resulting from very small SAGE II transmittances (errors are capped at 300% as a flag).	
	Remove profiles under high beta angle conditions.	See also Wang et al. (1996)
HALOE	Remove cloud-contaminated values. Also remove profiles that may contain artifacts from faulty trip angle or constant lockdown angle registration.	Hervig and McHugh (1999) haloe.gats-inc.com/user_docs/index.php
	Remove aerosol contamination (O_3 and HCl).	Bhatt et al. (1999)
UARS MLS	Use screening guidelines based on instrument status, retrieval quality flags, and sign of precision values.	Livesey et al. (2003)
Aura MLS	Use screening guidelines based on instrument status, retrieval quality and convergence flags, and sign of precision values.	Livesey et al. (2013)
ACE-FTS	Remove occultations listed as bad.	databace.scisat.ca/validation/data_issues.php
	Remove data when error value > VMR or error value < 10^{-4} xVMR.	K. Walker (personal communication, 2012)
	Use a data screening procedure (see Sect. 2.1) to identify and remove the largest outliers.	
	V2.2 data after Sep. 2010 (2009 for ozone) are not used because of a data processing issue.	

Fig. 1. Merging procedure illustration for HCl. Top left panel shows the HCl monthly mean source data during the overlap period (Aug. 2004 - Nov. 2005) for HALOE, ACE-FTS, and Aura MLS. Top right panel illustrates step 1 in the merging procedure, with the temporary merged data values (orange) resulting from the adjustment of ACE-FTS and Aura MLS values to the mean reference indicated by the black dashed line (time mean of co-located ACE-FTS/Aura MLS points). Also, the cyan dashed line is the mean of the ACE-FTS points and the red dashed line is the mean of MLS points co-located with ACE-FTS. Middle left panel shows step 2 results, namely the merged values arising from merging HALOE data with the temporary merged data; the black dashed line is the new average reference value, obtained from a 2/3 and 1/3 weighting of the dashed orange (mean of orange points co-located with HALOE) and dashed blue line (mean of HALOE) values, respectively. Middle right panel shows all the source data and the final merged values during the overlap period. Bottom panel shows the source and merged time series from 1991 through 2012 after the calculated additive offsets are applied to the whole source datasets, which are then merged (averaged) together wherever overlap between instruments exists.

Fig. 2. Offsets applied to the HCl source datasets (top panels for HALOE, middle panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Nov. 2005) used here to compute the average offsets.

Fig. 3. Example of HCl time series analyses for 50°N-60°N and 32 hPa. (a) HCl monthly mean source data from ACE-FTS and Aura MLS; the MLS dots are filled when there is time overlap with ACE-FTS, and open if no such overlap exists. Simple linear fits are shown as colored lines for ACE-FTS and for Aura MLS (orange line for all red dots and red line for filled red dots only). Correlation coefficient values (R values) for the two time series are provided in the title. (b) Deseasonalized anomalies for both ACE-FTS and Aura MLS, with corresponding linear fits (and R values). (c) Difference of deseasonalized anomalies (ACE-FTS minus Aura MLS), with linear fit.

Fig. 4. Latitude/pressure contours of time series diagnostics obtained from analyses illustrated in Fig. 3 for HCl from Aura MLS and ACE-FTS. Top panel: Correlation coefficient for the deseasonalized time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the error in this slope.

Fig. 5. Illustration of GOZCARDS HCl monthly averages with systematic error estimates (grey shading) at 46 hPa for 30°S-40°S; see text for the meaning of this shaded region. The source data from HALOE, Aura MLS, and ACE-FTS are shown in different colors (see legend), along with the merged values.

Fig. 6. Systematic error estimates for GOZCARDS HCl. One error (left panels) is relevant for values lower than (below) the merged values, and one (right panels) for values larger than the merged values; the

1674 top panels give the error estimates in ppbv, and the bottom panel errors are expressed as percent of the
1675 average merged values over the relevant time periods (see text). These error bars provide a range within
1676 which 95% of the source data values lie.

1677 **Fig. 7.** Time series of the GOZCARDS monthly-averaged merged HCl abundance for 3 different latitude
1678 bin averages (see color legend in panel (a)) for (a) 0.7 hPa, (b) 10 hPa, (c) 32 hPa, and (d) 68 hPa.

1679 **Fig. 8.** The average rate of change (percent per year) for HCl as a function of pressure for different
1680 latitude bin averages (see legend) for time periods corresponding to the appropriate GOZCARDS HCl
1681 values (see text) in the upper stratosphere (Jan. 1997 - Sep. 2010) and lower stratosphere (Jan. 1997 -
1682 Dec. 2012). Deseasonalized monthly data were used to obtain a long-term trend for these time periods;
1683 two-sigma error bars are shown.

1684 **Fig. 9.** Rates of change for GOZCARDS HCl (connected open circles) are given as a function of latitude
1685 in 10° latitude bins for sliding 6-year periods centered on Jan. 1 of each year (e.g., the 1998 point is an
1686 average for data from 1995 through 2000, and the 2011 point is for data from 2008 through 2013). (a) is
1687 for changes in upper stratospheric HCl at 0.7 hPa and (b) is for the change in the integrated HCl column
1688 between 68 hPa and 10 hPa. The two additional curves in (a) represent the rates of change in the
1689 estimated surface total chlorine from NOAA data (green is for a 6-year time shift, and purple for a 7-year
1690 time shift, to account for transport time to the upper stratosphere); see text for more details. Error bars
1691 indicate twice the standard errors in the means.

1692 **Fig. 10.** Offsets applied to the H_2O source datasets as a function of latitude and pressure, similar to
1693 Fig. 2 for HCl.

1694 **Fig. 11.** Latitude/pressure contours of time series diagnostics for H_2O from Aura MLS and ACE-FTS;
1695 this is similar to Fig. 4 for HCl.

1696 **Fig. 12**. A depiction of the "tape recorder" evolution for tropical water vapor abundances from 147 to
1697 10 hPa for October 1991 through December 2013. This plot was produced from GOZCARDS merged
1698 H_2O time series anomalies (differences from the long-term means) for the average of the 4 tropical bins
1699 covering 20°S to 20°N.

1700 **Fig. 13.** Systematic error estimates for GOZCARDS H_2O (similar to Fig. 6 for HCl).

1701 **Fig. 14.** Variations in stratospheric water vapor from the GOZCARDS H_2O merged data records (1992
1702 through 2013) averaged from (a) 60°S to 60°N and (b) 20°S to 20°N. Monthly average values and annual
1703 averages are shown by thin and thick lines (connecting similarly-colored dots), respectively, for the
1704 pressure levels indicated in the plot legend.

Fig. 15. Stratospheric water vapor variability on decadal timescales for 1992 through 2013 for tropical (20°S-20°N in black) and mid-latitude (20°N-60°N in red and 20°S-60°S in blue) zonal means, based on the GOZCARDS merged H_2O data record. The variability is expressed here as the difference between maximum and minimum annual average abundances, from 100 to 1 hPa, in ppmv (left panel) and percent (right panel).

Fig. 16. (a) Variations in upper mesospheric (0.01 hPa) water vapor mixing ratios averaged from 60°S to 60°N for Oct. 1991 through Dec. 2013, based on the GOZCARDS merged H_2O data records. Monthly average values and annual averages are shown by connected brown dots and connected black dots, respectively. (b) GOZCARDS merged H_2O annual averages (connected filled symbols) from 60°S to 60°N for 1992 through 2013 at pressure levels between 0.1 and 0.01 hPa. A time series of annually-averaged Lyman α solar flux values (open circles), scaled to arbitrary units, is also displayed (see text).

Fig. 17. Time series of monthly zonal mean O_3 for 10°S - 20°S between 1 hPa and 6.8 hPa (with pressure values given by ″pre″) from SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS, and ACE-FTS, all color-coded following the legend in top left panel.

Fig. 18. Schematic diagram describing the creation of the merged GOZCARDS monthly zonal mean ozone data record from various satellite datasets. Instruments represented in red inside the boxes are used as a reference. Instruments whose measurements have already been adjusted to a reference are indicated with a "*" superscript. AMLS refers to Aura MLS and UMLS to UARS MLS. See text for more details.

Fig. 19. Offsets applied to the O_3 source datasets, similar to Fig. 2 for HCl.

Fig. 20. Latitude/pressure contours of time series diagnostics for O_3 from Aura MLS and ACE-FTS; this is similar to Fig. 4 for HCl. The correlation coefficients (R values) and slope trend diagnostics are provided for HALOE versus SAGE II in the top two panels (for 1993-1999 as the trend issue for converted SAGE II data occurs after mid-2000 and to avoid Pinatubo-related data gaps before 1993) and for ACE-FTS versus Aura MLS in the bottom two panels (for 2005-2009).

Fig. 21. Systematic error estimates for GOZCARDS O_3 (similar to Fig. 6 for HCl).

Fig. 22. Near-global (60°S to 60°N) results for average column ozone (total and stratospheric, from *Ziemke and Chandra,* 2012) compared to GOZCARDS O_3 columns above 68 hPa. Stratospheric columns are offset to better match the total column values, in order to more easily compare relative variations

versus time; the black dots and red crosses are referenced to the 1980 total column values, while the cyan curves are referenced to 2007 to better illustrate the fits in the later years.

Fig. 23. Time evolution (Aug. 2004 through 2012) versus latitude of GOZCARDS merged N_2O (ppbv) at (a) 6.8 hPa and (b) 100 hPa.

Fig. 24. Sample results display the time evolution of satellite-retrieved HNO_3 (ppbv) for two different periods, 1992-1997 in (a) and (c) versus 2004-2013 in (b) and (d). Panels (a) and (b) are contour plots at 46 hPa from UARS MLS global data and the merged GOZCARDS global data after 2004, respectively; (c) and (d) show time series at 32 hPa and for the 40°N-50°N latitude bin, with (a) from UARS MLS data, and (d) from ACE-FTS, Aura MLS, and the merged combination (between the two source data sets).

Fig. A1. Illustration of the standard deviations (in (a)) and standard errors (in (b)) for monthly mean GOZCARDS HCl (source and merged records) at 46 hPa for 30°S-40°S. Source data from HALOE, Aura MLS, and ACE-FTS are given by the filled colored dots (see legend); each standard deviation is simply obtained from the range of values measured during the month. The large open brown circles give standard deviations for the merged HCl product; this Appendix provides the formulae to calculate these quantities.

Fig. S1: Illustration of the latitudinal dependence of the HCl offsets for HALOE, ACE-FTS, and Aura MLS at two pressure levels (top panel for 0.46 hPa, bottom panel for 46 hPa). Error bars represent twice the standard error in the derived offsets (based on variability during the overlapping period). Larger standard error values indicate that there were either fewer points of overlap or larger offset variability (standard deviations); we found that both of these factors contribute.

Fig. S2: Latitude/pressure contours of the fitted mean annual amplitudes (ppbv) from HCl time series for HALOE, ACE-FTS, and Aura MLS, based on their respective measurement periods (see text).

Fig. S3: Time evolution (Oct. 1991 through 2013) versus latitude of GOZCARDS merged HCl (ppbv) at 46 hPa.

Fig. S4: HALOE sunrise measurements of H_2O versus the 3.46 μm extinction coefficient for 1992, 1993, and 1999 at 22 hPa. The green vertical line represents the aerosol extinction value ($5x10^{-4}$ km^{-1}) used to screen anomalous HALOE H_2O values. It is apparent that anomalously low H_2O values occurred in 1992 when the 3.46 μm aerosol extinction exceeded about $5x10^{-4}$ km^{-1}. These artifacts were confined to 1991 and 1992; for these years, and for pressure levels at and below 22 hPa, the corresponding H_2O data values

were excluded. This screening method eliminates about 10% of the global (lower stratospheric) measurements in 1992.

Fig. S5: Merging procedure illustration for H_2O at 5°N and 22hPa. This is similar to Fig. 2 (for HCl), but an additional step is illustrated for the end of this procedure, whereby stratospheric H_2O data from UARS MLS are adjusted to the early portion of the merged time series that was obtained after the 2nd step; this adds more coverage (more brown dots in the bottom panel for 1991-1993).

Fig. S6: Latitude/pressure contours of the fitted mean annual amplitudes (ppmv) from H_2O time series for HALOE, ACE-FTS, and Aura MLS, based on their respective measurement periods.

Fig. S7: Time evolution (Oct. 1991 through 2013) versus latitude of GOZCARDS merged H_2O (ppmv) at 3.2 hPa (top panel) and 68 hPa (bottom panel).

Fig. S8: Monthly zonal mean ozone differences (%) between SAGE II and (a) HALOE, (b) UARS MLS (UMLS for short), (c) Aura MLS (AMLS for short), and (d) ACE-FTS during their respective overlap periods. Differences are expressed (in percent) as 100 x [(SAGE II - Other) / (Other)]. Shaded areas indicate negative values.

Fig. S9: Monthly zonal mean temperature differences between NCEP (used by SAGE II) and HALOE temperatures relative to MERRA for 10°S - 20°S between 1 and 6.8 hPa, per color-coding indicated in bottom left panel; ''pre'' represents the pressure value. From 1 to 2.1 hPa, differences between NCEP and MERRA are generally within ± 4K before mid-2000. After that time, NCEP temperatures show a sharp increase and are systematically higher than MERRA values by 5 to 10K. However, this divergence and trend are not seen in HALOE temperatures. NCEP temperatures between 3.2 and 6.8 hPa are smaller than MERRA after mid-2000; negative trends (versus MERRA) also occur in the HALOE data at these levels.

Fig. S10: Relative trends (K/decade) in zonal mean temperature differences for NCEP – MERRA and HALOE – MERRA (color-coded as in Fig. S9) in the upper stratosphere. NCEP temperatures show positive trends versus MERRA of ~2-5 K/decade between 2.1 and 1 hPa for all latitudes. However, HALOE temperatures show no significant trends versus MERRA, except at 1.5 hPa in the southern hemisphere. For pressures between 3.2 and 6.8 hPa, the temperature analyses are not conclusive; although NCEP values show negative trends of ~2-3 K/decade versus MERRA, they agree with HALOE.

Fig. S11: Mean differences and standard deviations (horizontal bars) between SAGE II and Aura MLS ozone in three different latitude bins: 20°S to 60°S (left panel), 20°S to 20°N (middle panel), and 20°N to 60°N (right panel). Results based on monthly zonal mean and coincident profiles (see text for coincidence criteria) during overlap periods are shown in red and blue, respectively. To choose collocated profiles,

coincidence criteria of ±1° in latitude and ±8° in longitude were used; the time difference criterion was chosen as 12 hours, but only nighttime measurements from Aura MLS were used.

Fig. S12: Latitude/pressure contours of the fitted mean annual amplitudes (ppmv) from O_3 time series for SAGE II, HALOE, ACE-FTS, and Aura MLS, based on their respective measurement periods.

Fig. S13: Illustration of the time evolution of the GOZCARDS merged O_3 data field versus latitude at 68 hPa (top panel) and versus pressure for the 40°N-50°N latitude bin (bottom panel).

Fig. S14: Offsets applied to the N_2O source datasets (top panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Sep. 2010) used here to compute the average offsets.

Fig. S15: Latitude/pressure contours of time series diagnostics derived from Aura MLS and ACE-FTS N_2O data comparisons (and obtained from analyses similar to those illustrated in Fig. 6 for HCl). Top panel: Correlation coefficient for the deseasonalized time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the error in this slope.

Fig. S16: Offsets applied to the HNO_3 source datasets (top panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Sep. 2010) used here to compute the average offsets.

Fig. S17: Latitude/pressure contours of time series diagnostics derived from Aura MLS and ACE-FTS HNO_3 data comparisons (and obtained from analyses similar to those illustrated in Fig. 6 for HCl). Top panel: Correlation coefficient for the deseasonalized time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the error in this slope.

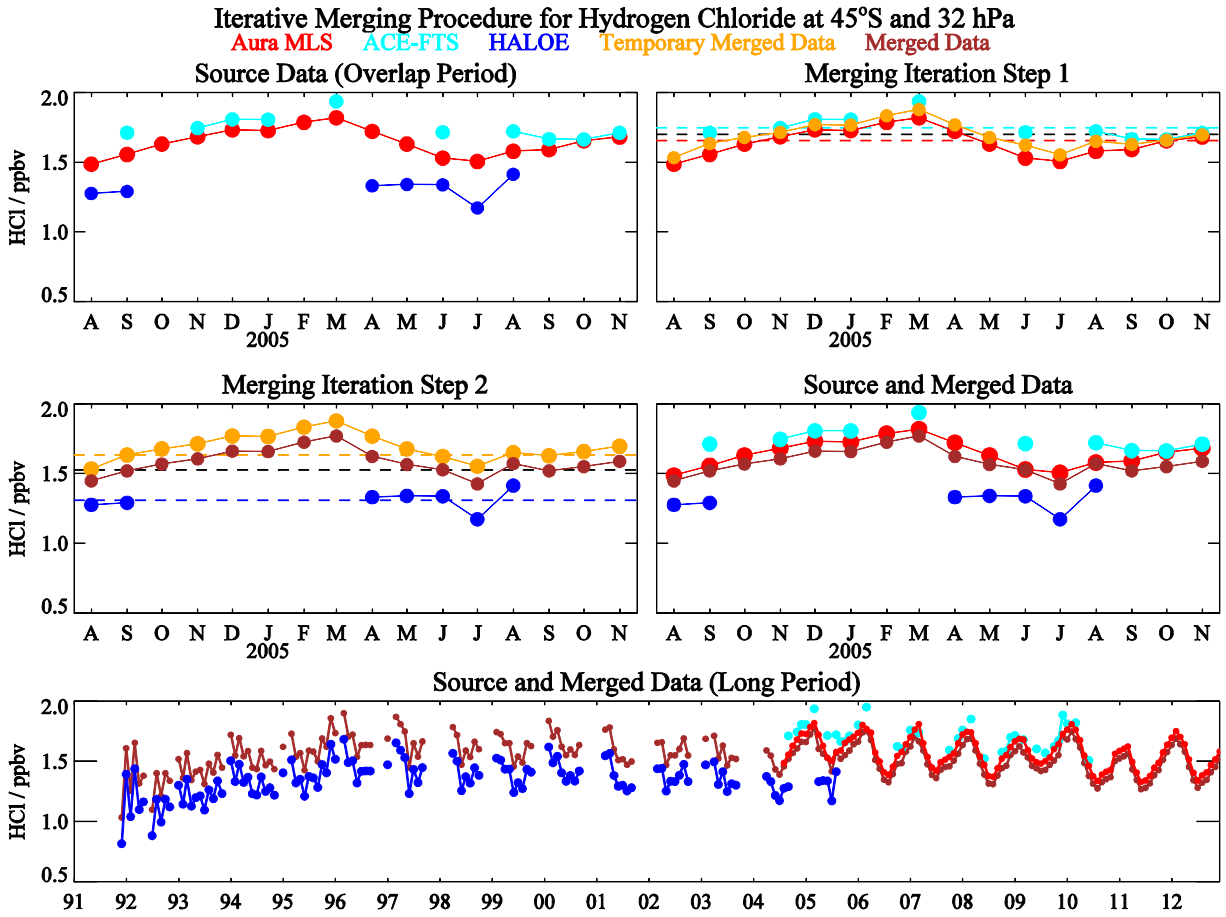


Fig. 1. Merging procedure illustration for HCl. Top left panel shows the HCl monthly mean source data during the overlap period (Aug. 2004 - Nov. 2005) for HALOE, ACE-FTS, and Aura MLS. Top right panel illustrates step 1 in the merging procedure, with the temporary merged data values (orange) resulting from the adjustment of ACE-FTS and Aura MLS values to the mean reference indicated by the black dashed line (time mean of co-located ACE-FTS/Aura MLS points). Also, the cyan dashed line is the mean of the ACE-FTS points and the red dashed line is the mean of MLS points co-located with ACE-FTS. Middle left panel shows step 2 results, namely the merged values arising from merging HALOE data with the temporary merged data; the black dashed line is the new average reference value, obtained from a 2/3 and 1/3 weighting of the dashed orange (mean of orange points co-located with HALOE) and dashed blue line (mean of HALOE) values, respectively. Middle right panel shows all the source data and the final merged values during the overlap period. Bottom panel shows the source and merged time series from 1991 through 2012 after the calculated additive offsets are applied to the whole source datasets, which are then merged (averaged) together wherever overlap between instruments exists.

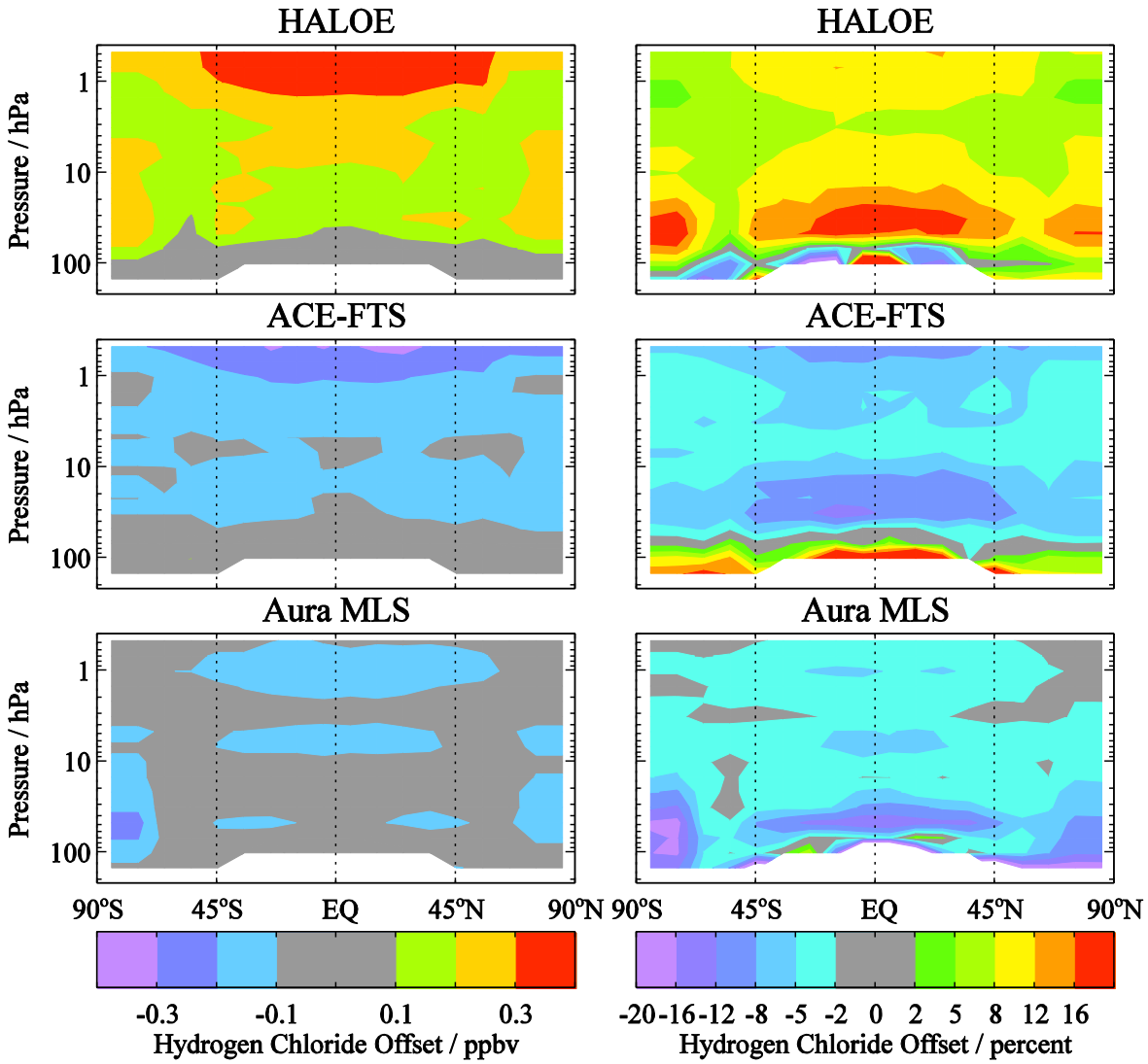


Fig. 2. Offsets applied to the HCl source datasets (top panels for HALOE, middle panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Nov. 2005) used here to compute the average offsets.

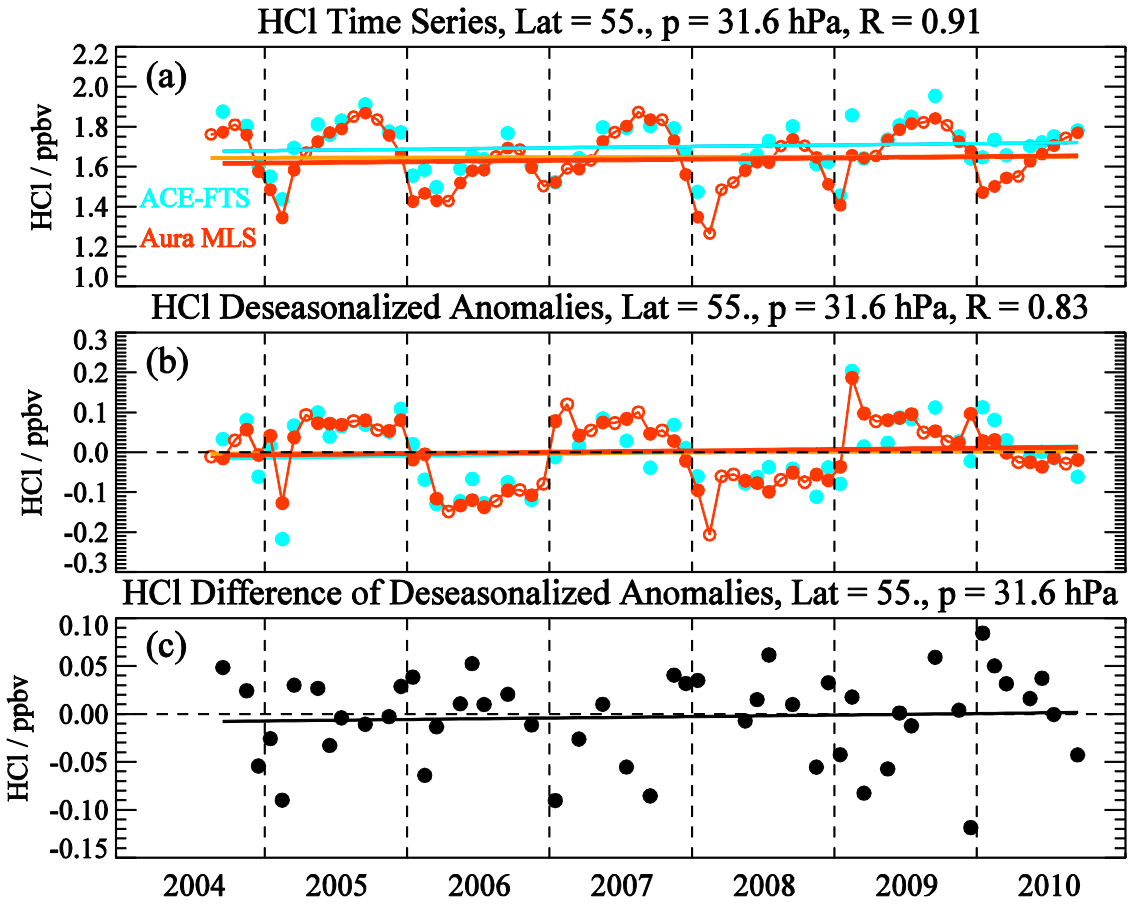
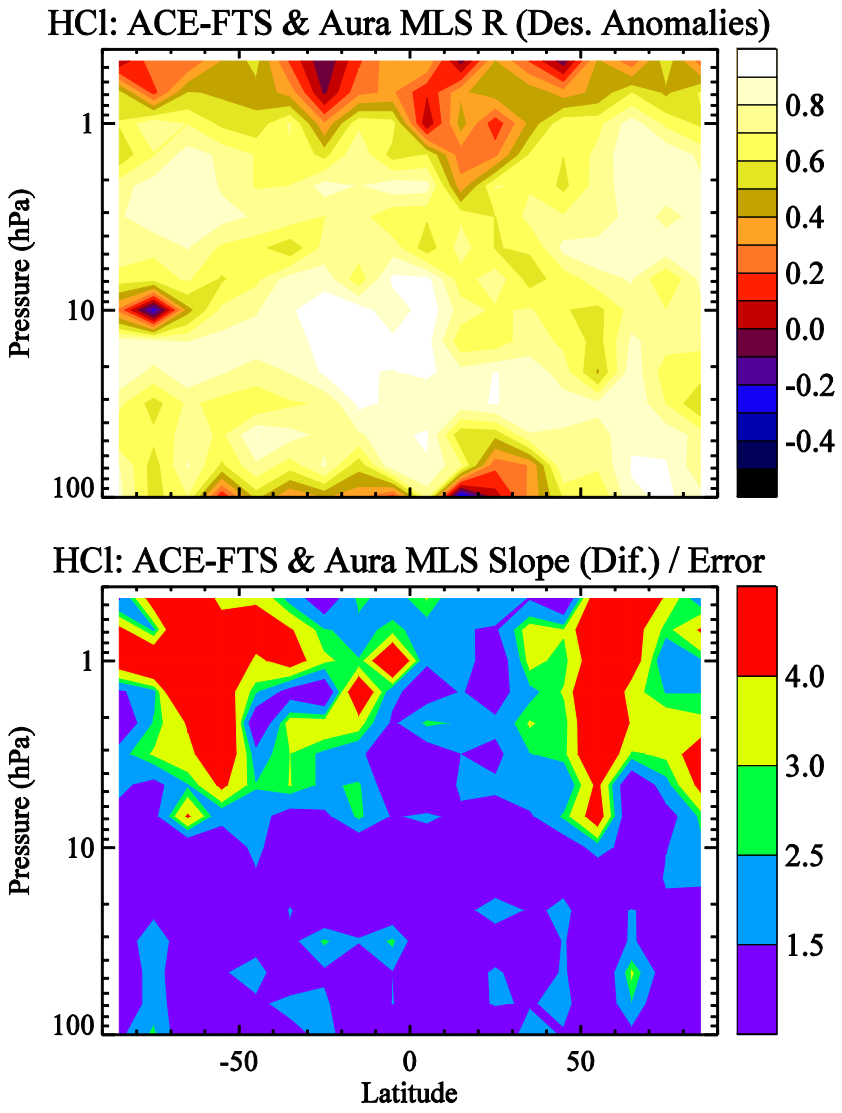


Fig. 3. Example of HCl time series analyses for 50°N-60°N and 32 hPa. (a) HCl monthly mean source data from ACE-FTS and Aura MLS; the MLS dots are filled when there is time overlap with ACE-FTS, and open if no such overlap exists. Simple linear fits are shown as colored lines for ACE-FTS and for Aura MLS (orange line for all red dots and red line for filled red dots only). Correlation coefficient values (R values) for the two time series are provided in the title. (b) Deseasonalized anomalies for both ACE-FTS and Aura MLS, with corresponding linear fits (and R values). (c) Difference of deseasonalized anomalies (ACE-FTS minus Aura MLS), with linear fit.



35

36 **Fig. 4.** Latitude/pressure contours of time series diagnostics obtained from analyses illustrated in
37 Fig. 3 for HCl from Aura MLS and ACE-FTS. Top panel: Correlation coefficient for the deseasonalized
38 time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the
39 error in this slope.

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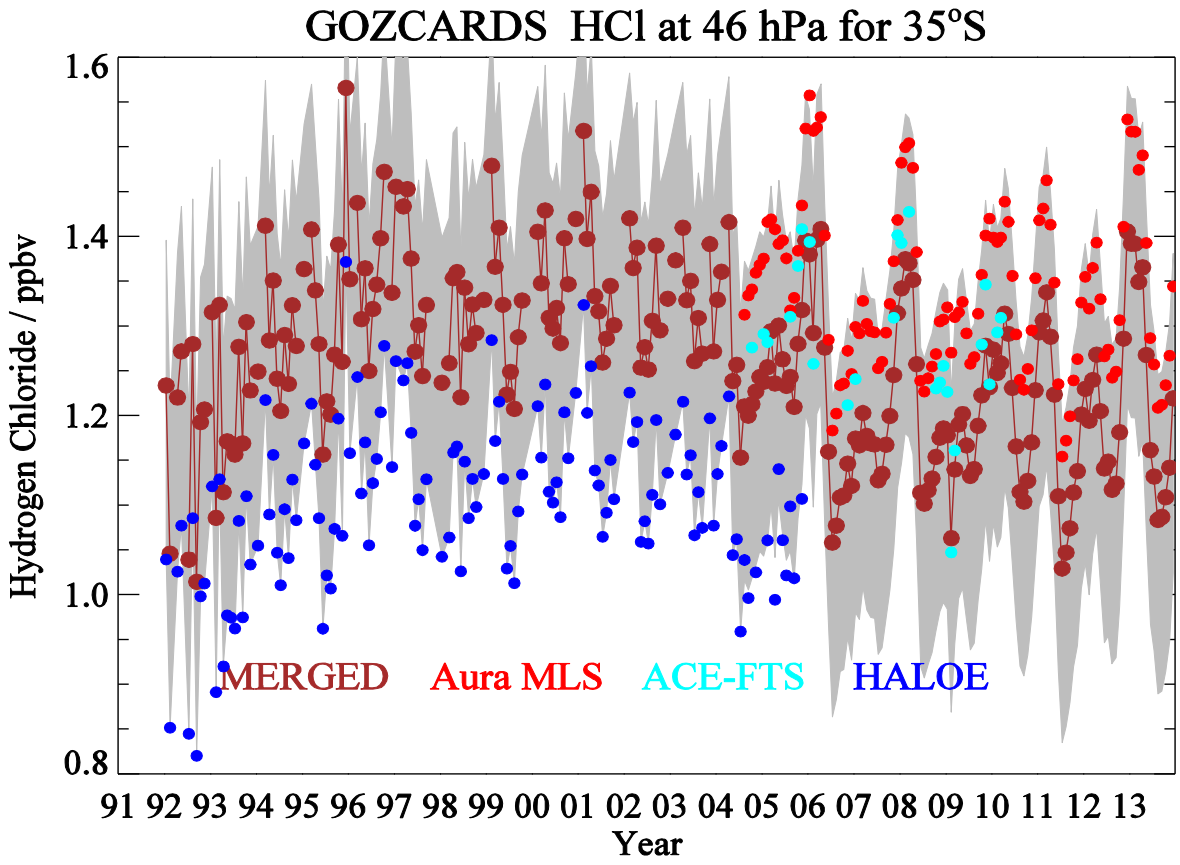


Fig. 5. Illustration of GOZCARDS HCl monthly averages with systematic error estimates (shown as grey shading) at 46 hPa for 30°S-40°S; see text for the meaning of this shaded region. The source data from HALOE, Aura MLS, and ACE-FTS are shown in different colors (see legend), along with the merged values.

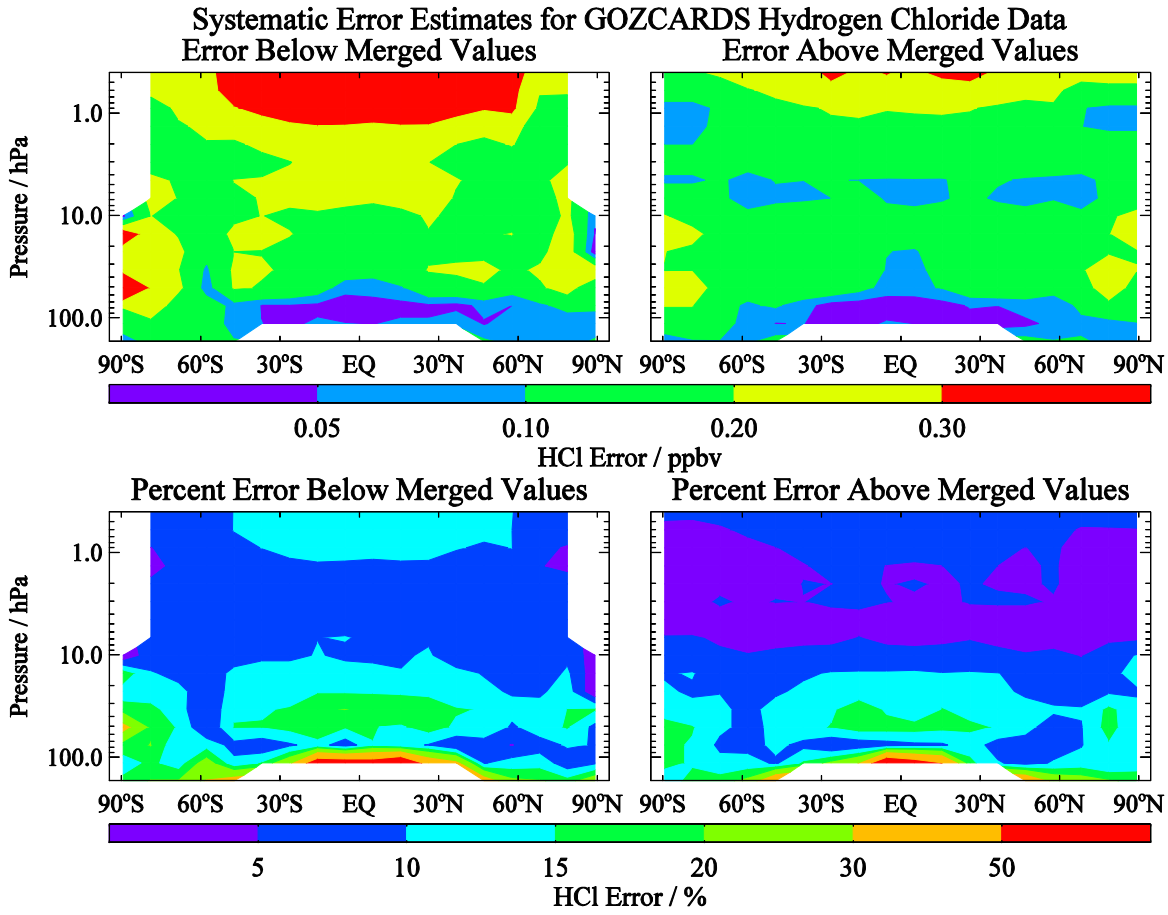


Fig. 6. Systematic error estimates for GOZCARDS HCl. One error (left panels) is relevant for values lower than (below) the merged values, and one (right panels) for values larger than the merged values; the top panels give the error estimates in ppbv, and the bottom panel errors are expressed as percent of the average merged values over the relevant time periods (see text). These error bars provide a range within which 95% of the source data values lie.

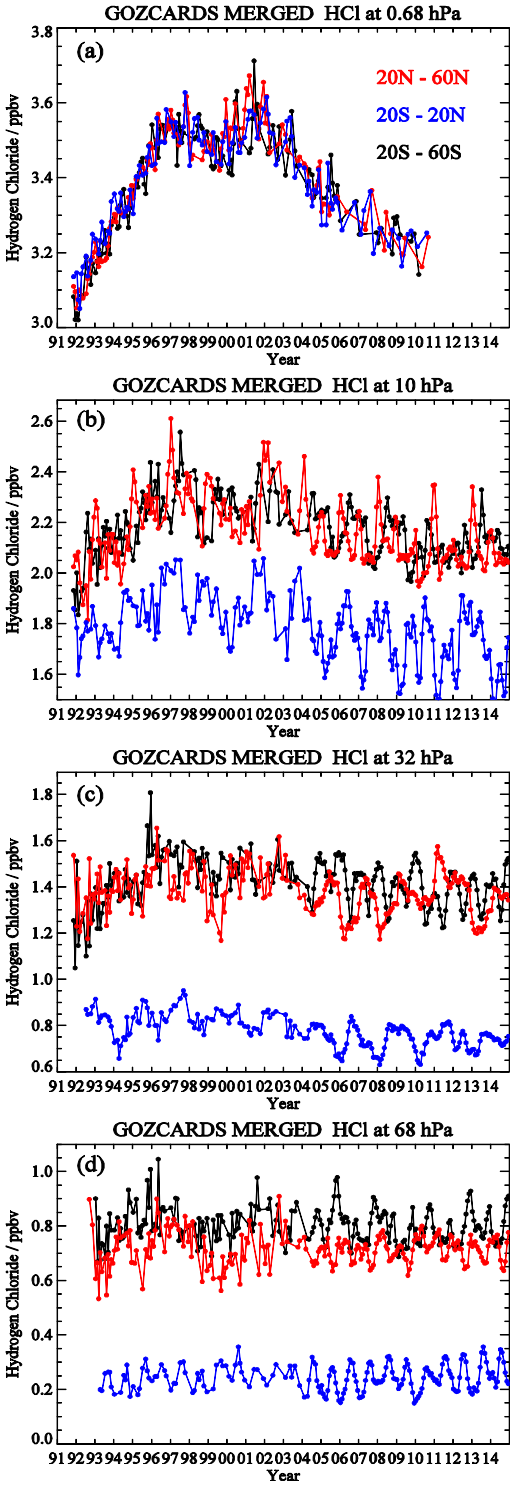


Fig. 7. Time series of the GOZCARDS monthly-averaged merged HCl abundance for 3 different latitude bin averages (see color legend in panel (a)) for (a) 0.7 hPa, (b) 10 hPa, (c) 32 hPa, and (d) 68 hPa.

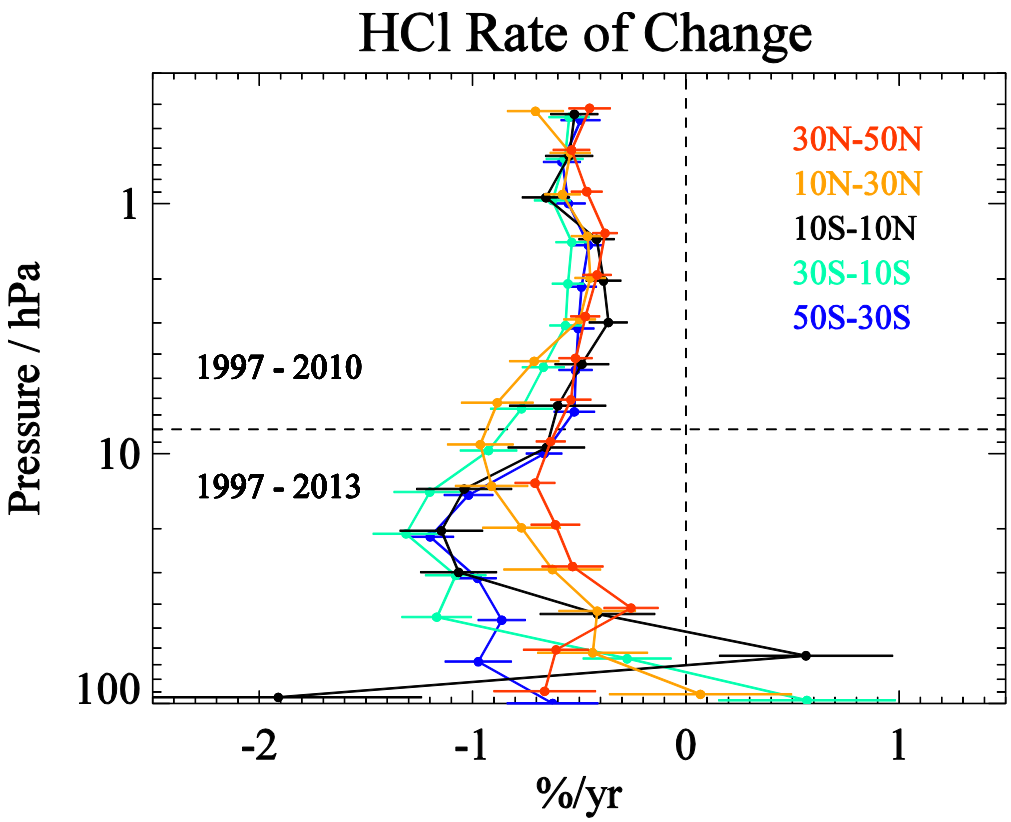


Fig. 8. The average rate of change (percent per year) for HCl as a function of pressure for different latitude bin averages (see legend) for time periods corresponding to the appropriate GOZCARDS HCl values (see text) in the upper stratosphere (Jan. 1997 - Sep. 2010) and lower stratosphere (Jan. 1997 - Dec. 2012). Deseasonalized monthly data were used to obtain a long-term trend for these time periods; two-sigma error bars are shown.

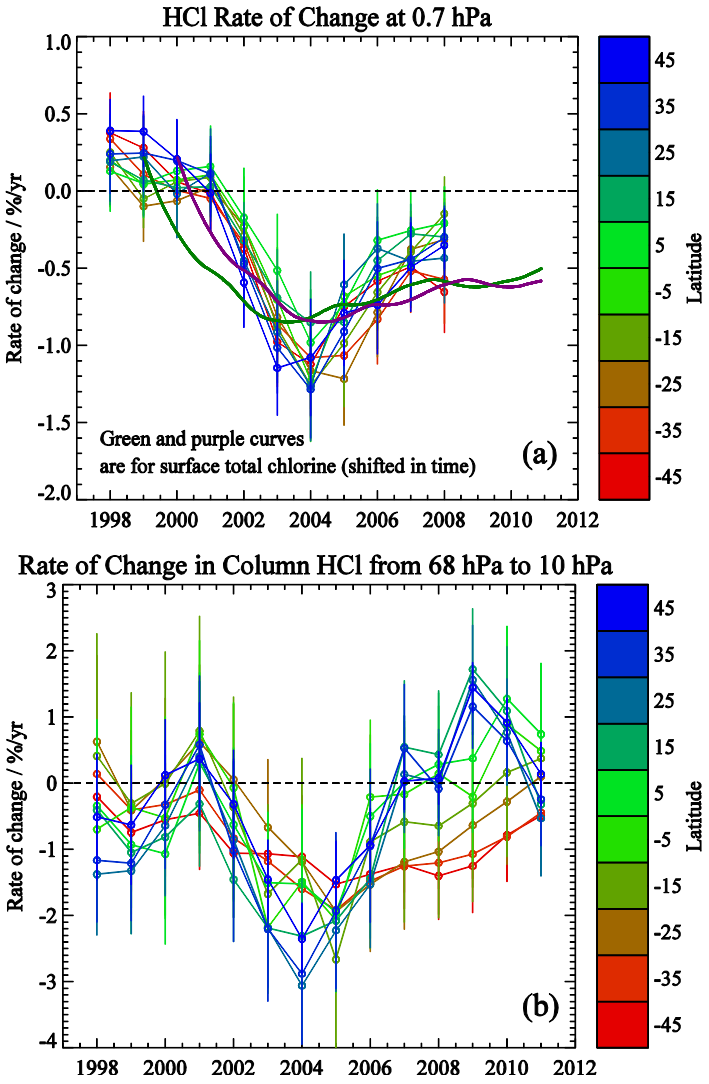
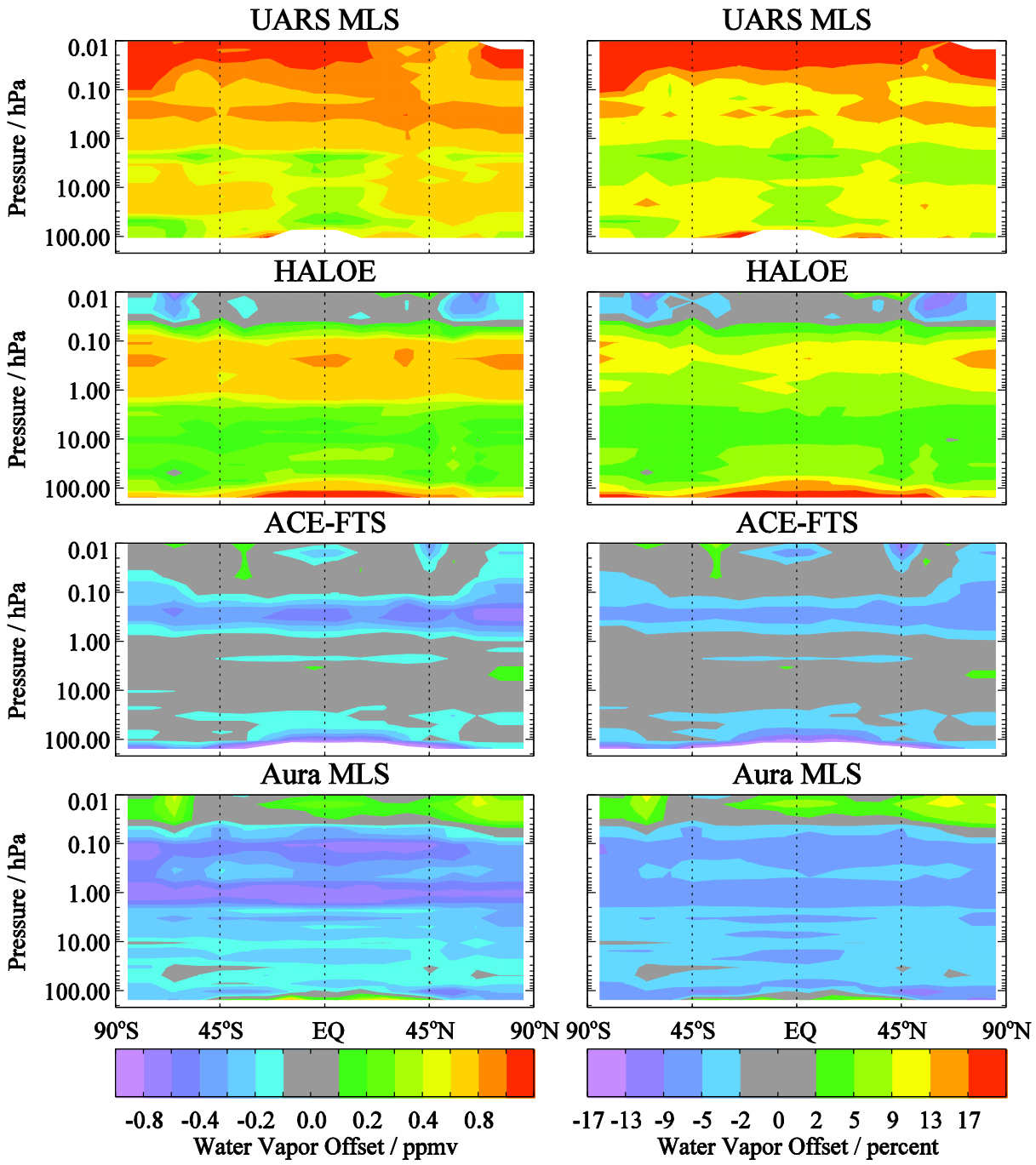


Fig. 9. Rates of change for GOZCARDS HCl (connected open circles) are given as a function of latitude in 10° latitude bins for sliding 6-year periods centered on Jan. 1 of each year (e.g., the 1998 point is an average for data from 1995 through 2000, and the 2011 point is for data from 2008 through 2013). (a) is for changes in upper stratospheric HCl at 0.7 hPa and (b) is for the change in the integrated HCl column between 68 hPa and 10 hPa. The two additional curves in (a) represent the rates of change in the estimated surface total chlorine from NOAA data (green is for a 6-year time shift, and purple for a 7-year time shift, to account for transport time to the upper stratosphere); see text for more details. Error bars indicate twice the standard errors in the means.



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106

107 **Fig. 10.** Offsets applied to the H_2O source datasets as a function of latitude and pressure, similar to

108 Fig. 2 for HCl.

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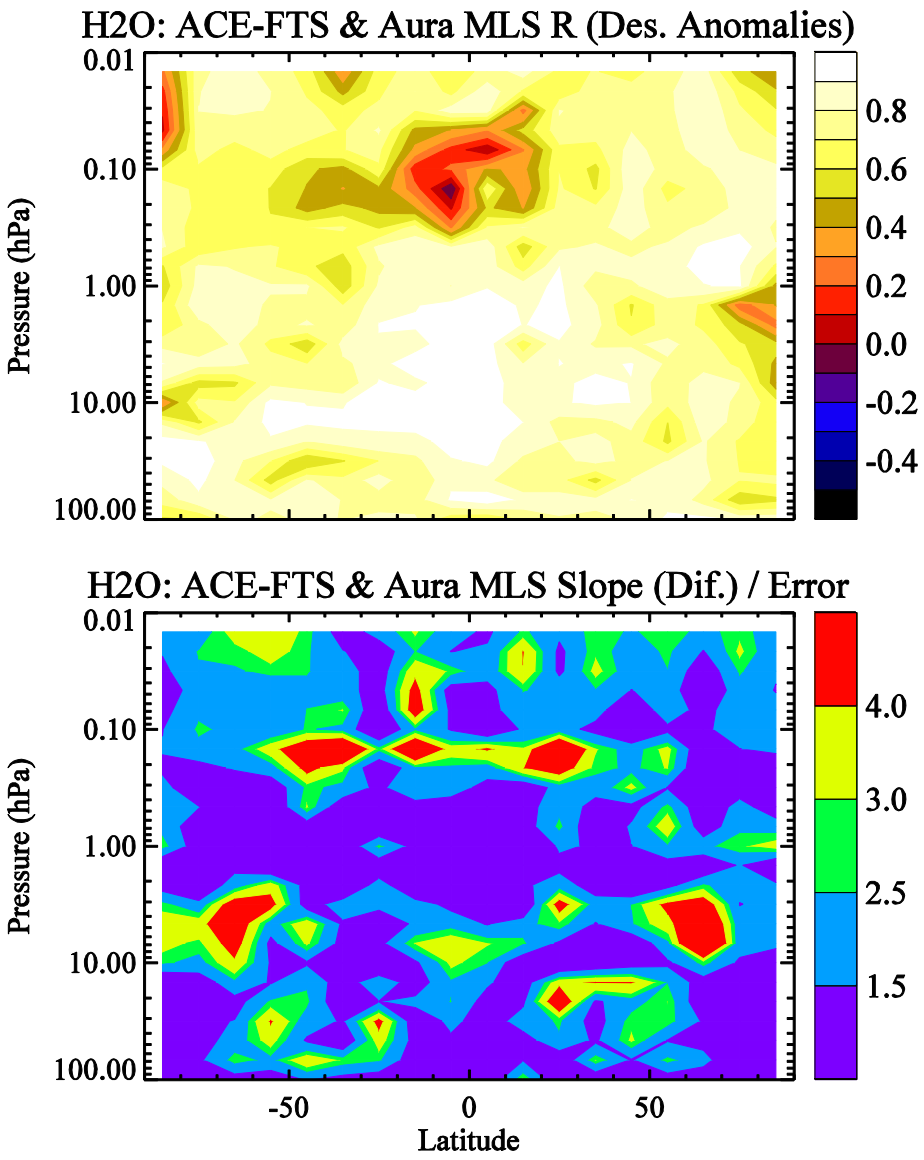

Fig. 11. Latitude/pressure contours of time series diagnostics for H_2O from Aura MLS and ACE-FTS; this is similar to Fig. 4 for HCl.

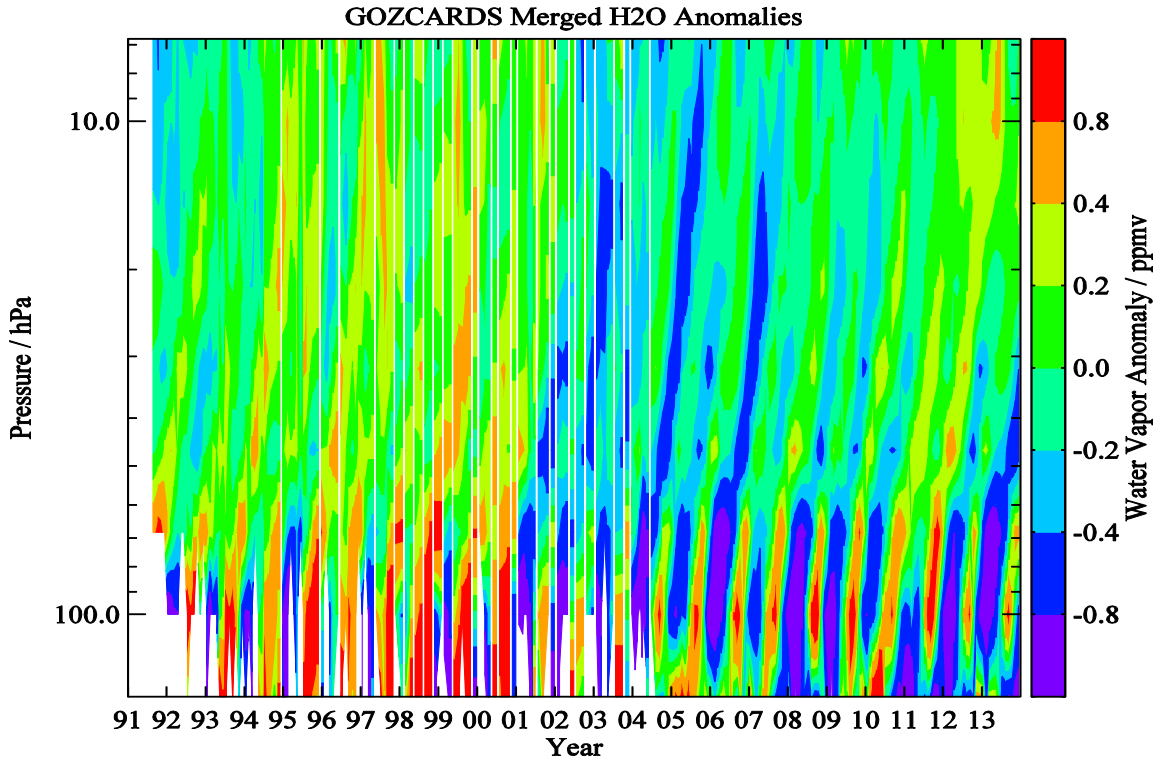

Fig. 12. A depiction of the "tape recorder" evolution for tropical water vapor abundances from 147 to 10 hPa for October 1991 through December 2013. This plot was produced from GOZCARDS merged H_2O time series anomalies (differences from the long-term means) for the average of the 4 tropical bins covering 20°S to 20°N.

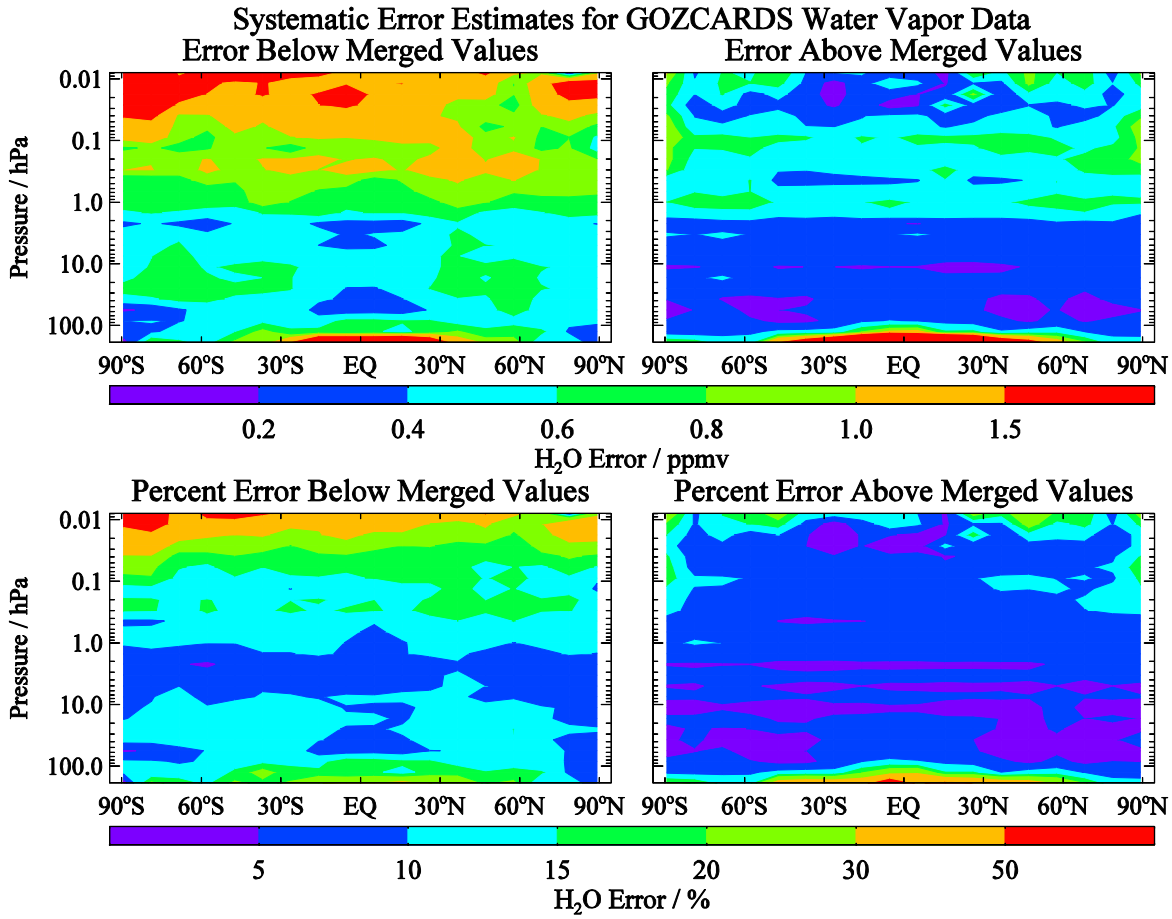

Fig. 13. Systematic error estimates for GOZCARDS H_2O (similar to Fig. 6 for HCl).

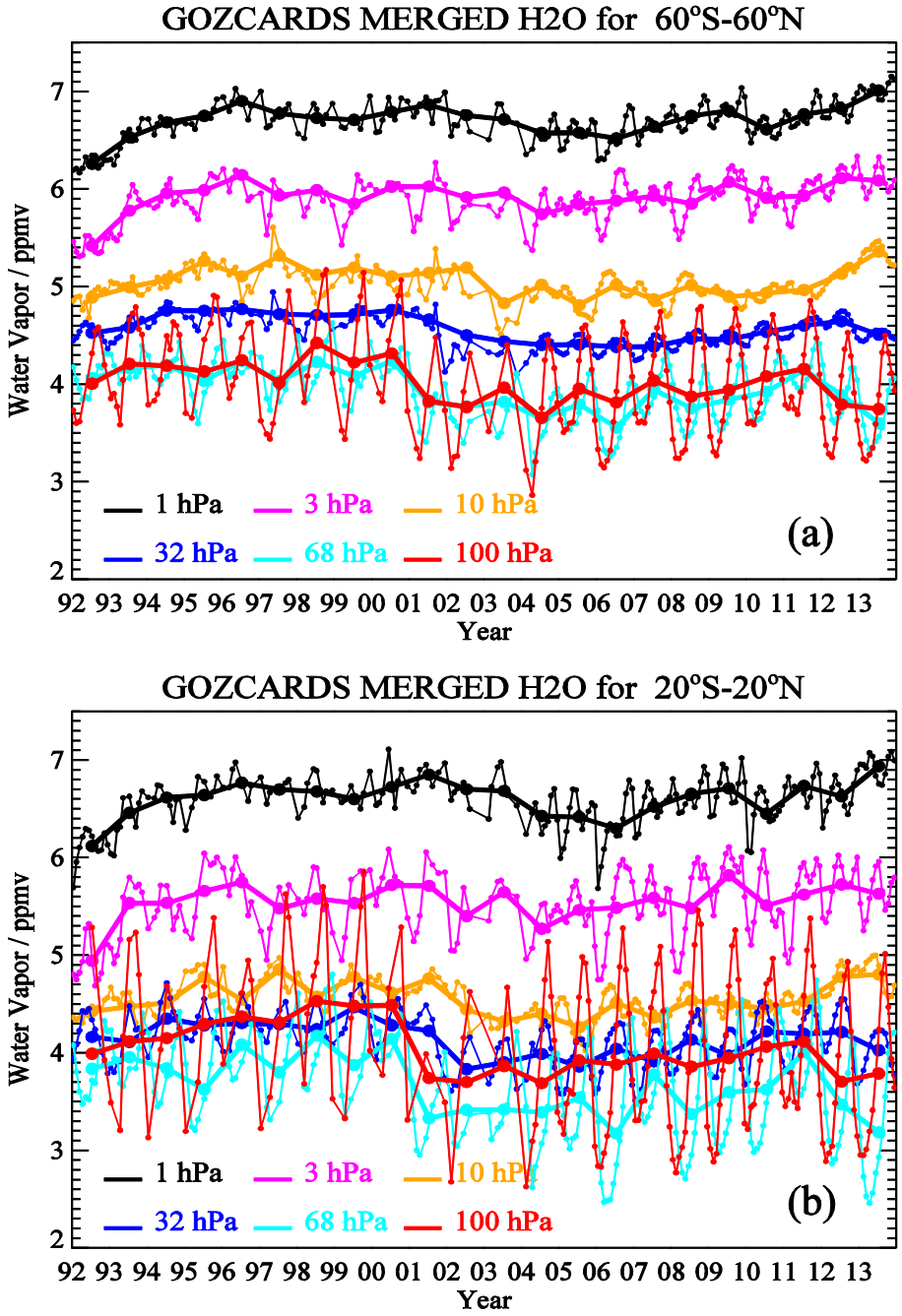

Fig. 14. Variations in stratospheric water vapor from the GOZCARDS H_2O merged data records (1992 through 2013) averaged from (a) 60°S to 60°N and (b) 20°S to 20°N. Monthly average values and annual averages are shown by thin and thick lines (connecting similarly-colored dots), respectively, for the pressure levels indicated in the plot legend.

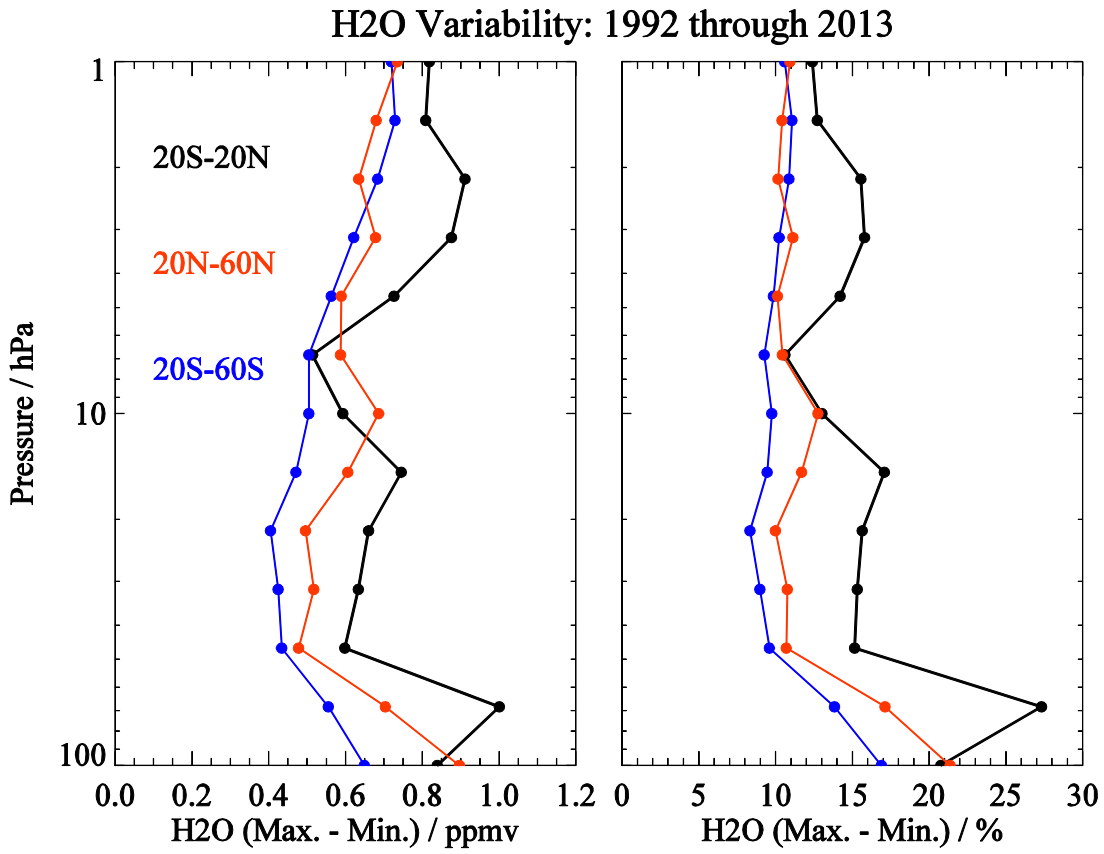

Fig. 15. Stratospheric water vapor variability on decadal timescales for 1992 through 2013 for tropical (20°S-20°N in black) and mid-latitude (20°N-60°N in red and 20°S-60°S in blue) zonal means, based on the GOZCARDS merged H_2O data record. The variability is expressed here as the difference between maximum and minimum annual average abundances, from 100 to 1 hPa, in ppmv (left panel) and percent (right panel).

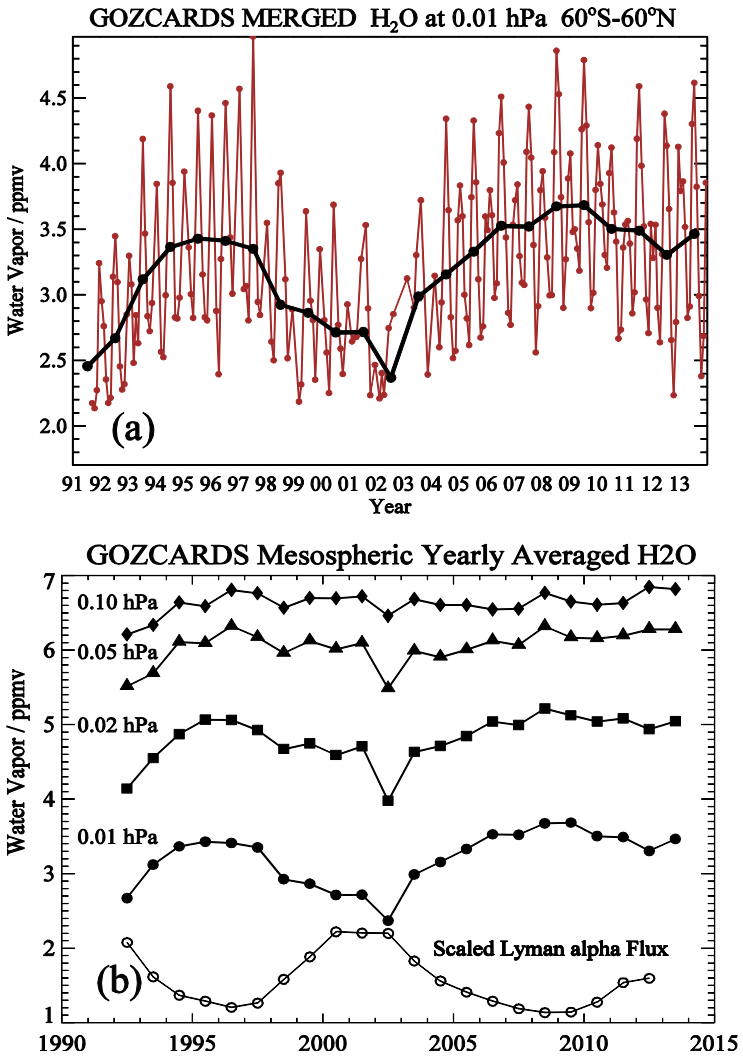

Fig. 16. (a) Variations in upper mesospheric (0.01 hPa) water vapor mixing ratios averaged from 60°S to 60°N for Oct. 1991 through Dec. 2013, based on the GOZCARDS merged H_2O data records. Monthly average values and annual averages are shown by connected brown dots and connected black dots, respectively. (b) GOZCARDS merged H_2O annual averages (connected filled symbols) from 60°S to 60°N for 1992 through 2013 at pressure levels between 0.1 and 0.01 hPa. A time series of annually-averaged Lyman α solar flux values (open circles), scaled to arbitrary units, is also displayed (see text).

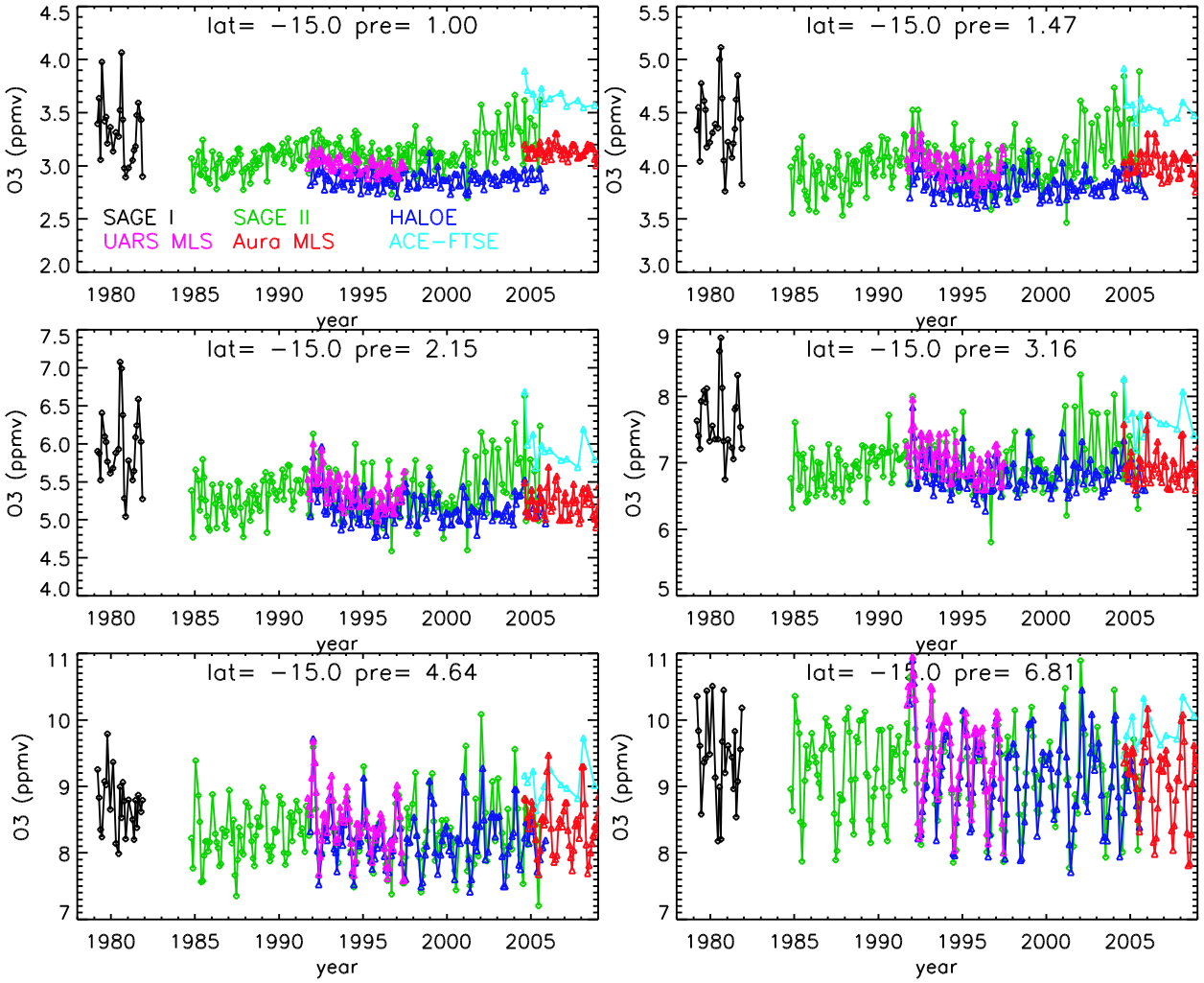

Fig. 17. Time series of monthly zonal mean O_3 for 10°S - 20°S between 1 hPa and 6.8 hPa (with pressure values given by "pre") from SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS, and ACE-FTS, all color-coded following the legend in top left panel.

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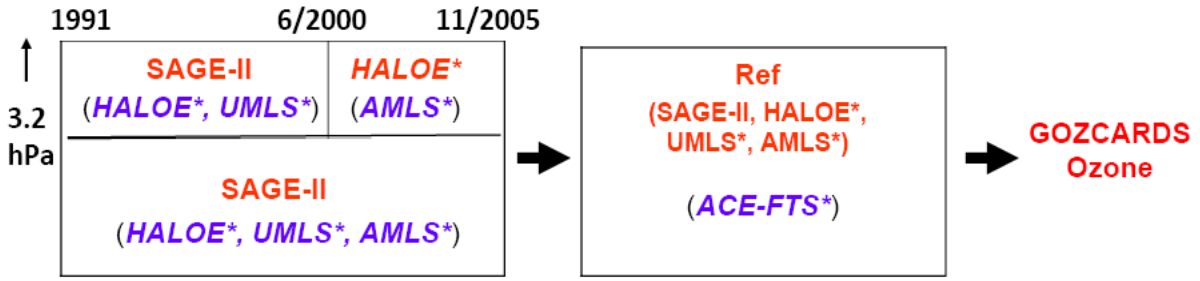

176

177 **Fig. 18.** Schematic diagram describing the creation of the merged GOZCARDS monthly zonal mean
178 ozone data record from various satellite datasets. Instruments represented in red inside the boxes are used
179 as a reference. Instruments whose measurements have already been adjusted to a reference are indicated
180 with a "*" superscript. AMLS refers to Aura MLS and UMLS to UARS MLS. See text for more details.

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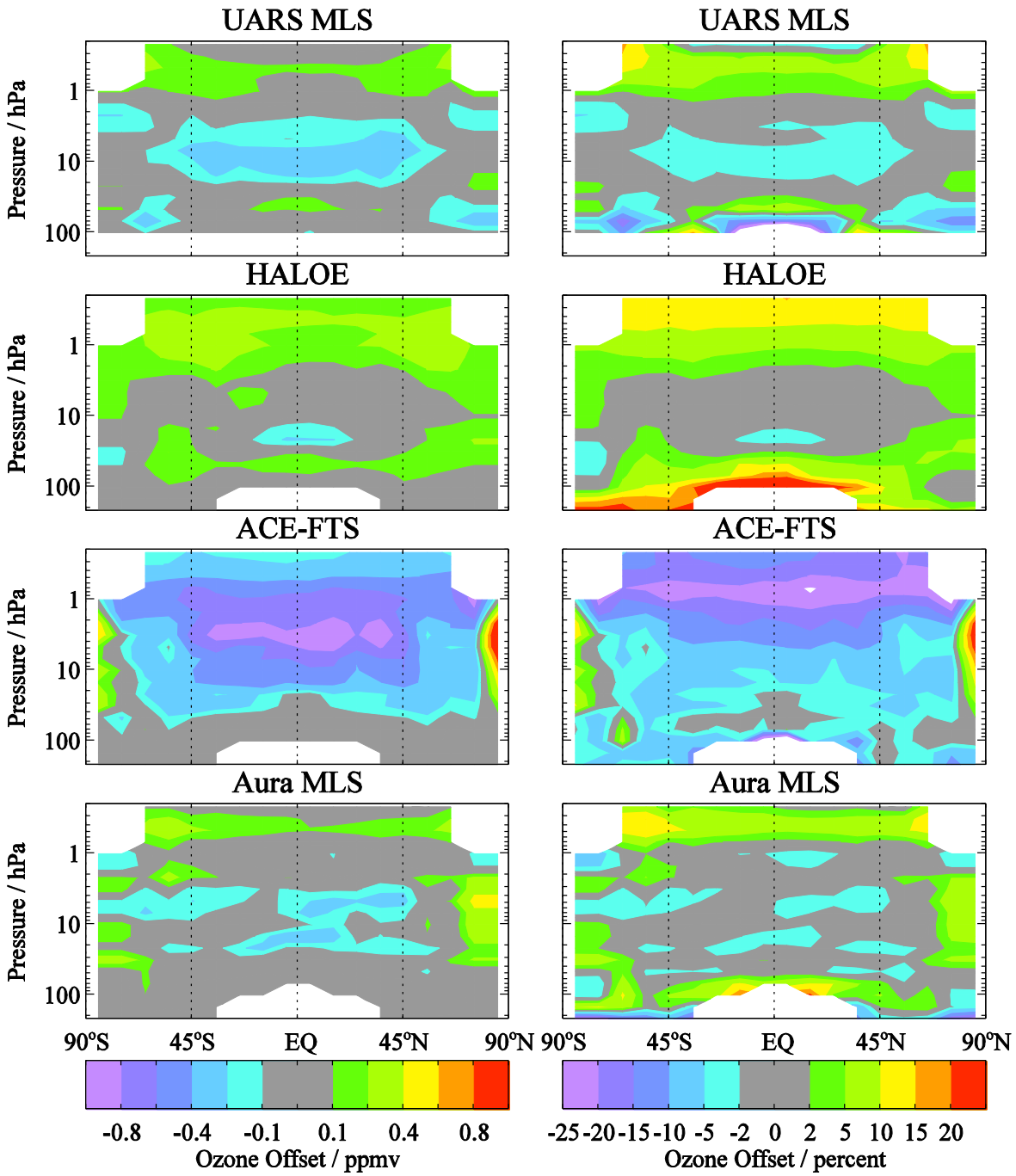

186

187 **Fig. 19.** Offsets applied to the O_3 source datasets, similar to Fig. 2 for HCl.

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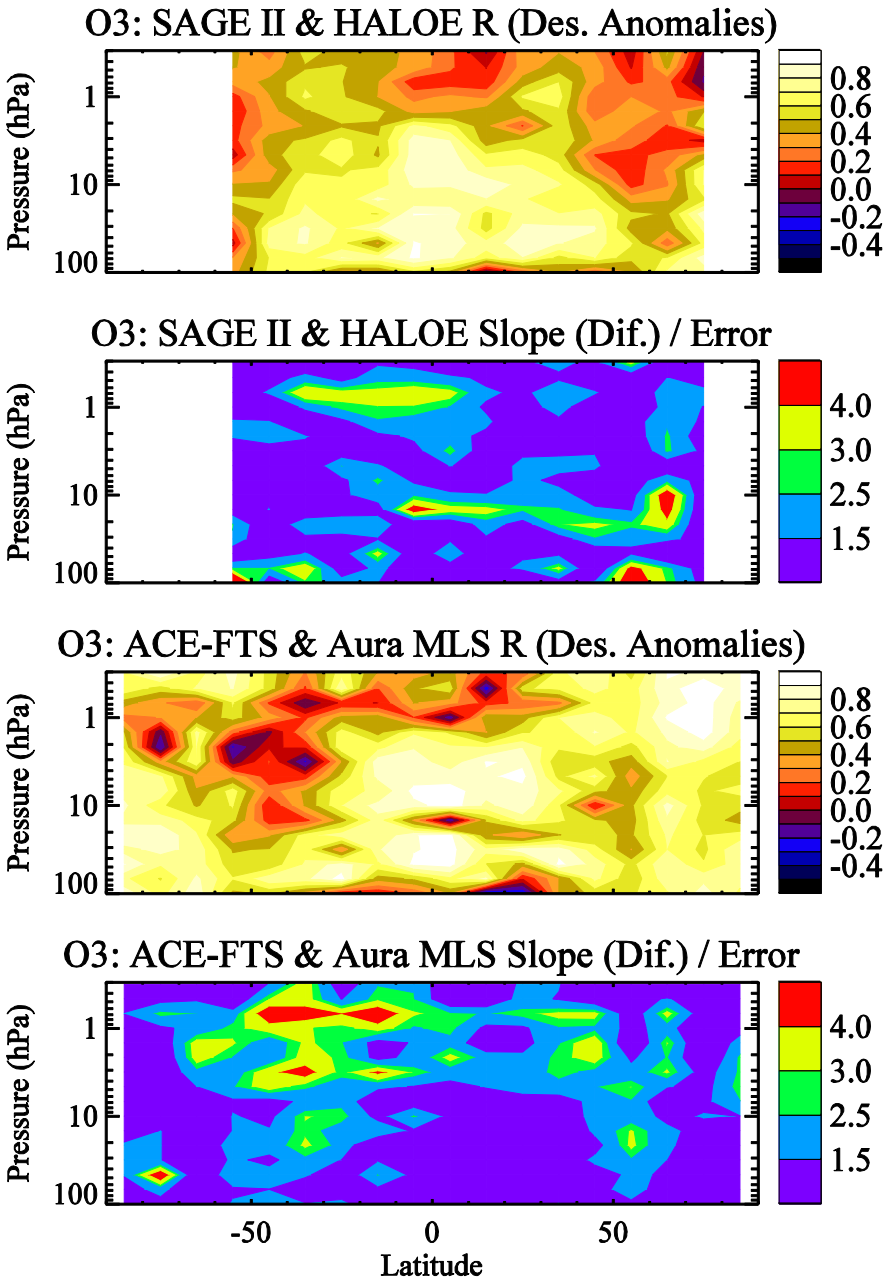


Fig. 20. Latitude/pressure contours of time series diagnostics for O_3 from Aura MLS and ACE-FTS; this is similar to Fig. 4 for HCl. The correlation coefficients (R values) and slope trend diagnostics are provided for HALOE versus SAGE II in the top two panels (for 1993-1999 as the trend issue for converted SAGE II data occurs after mid-2000 and to avoid Pinatubo-related data gaps before 1993) and for ACE-FTS versus Aura MLS in the bottom two panels (for 2005-2009).

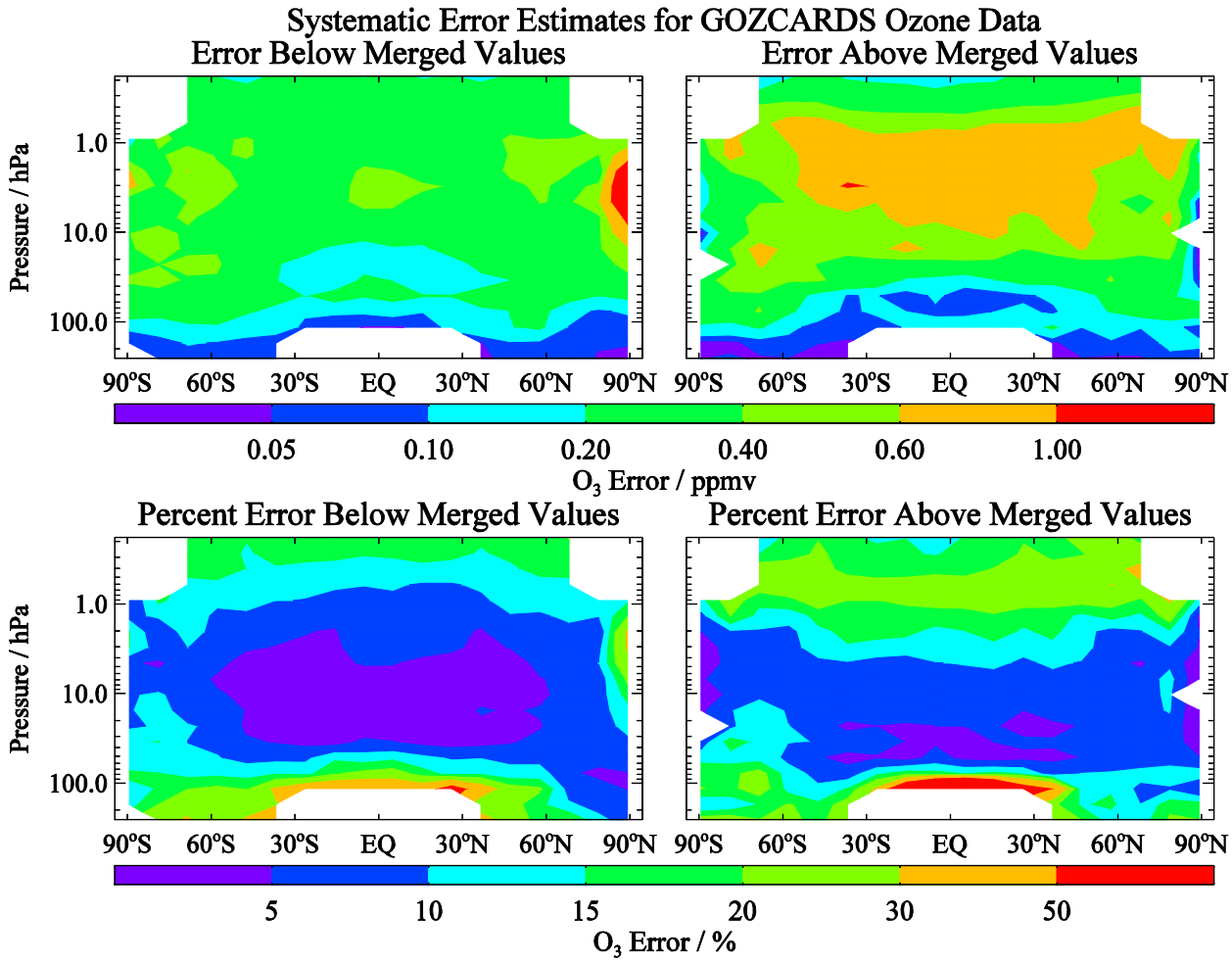


Fig. 21. Systematic error estimates for GOZCARDS O_3 (similar to Fig. 6 for HCl).

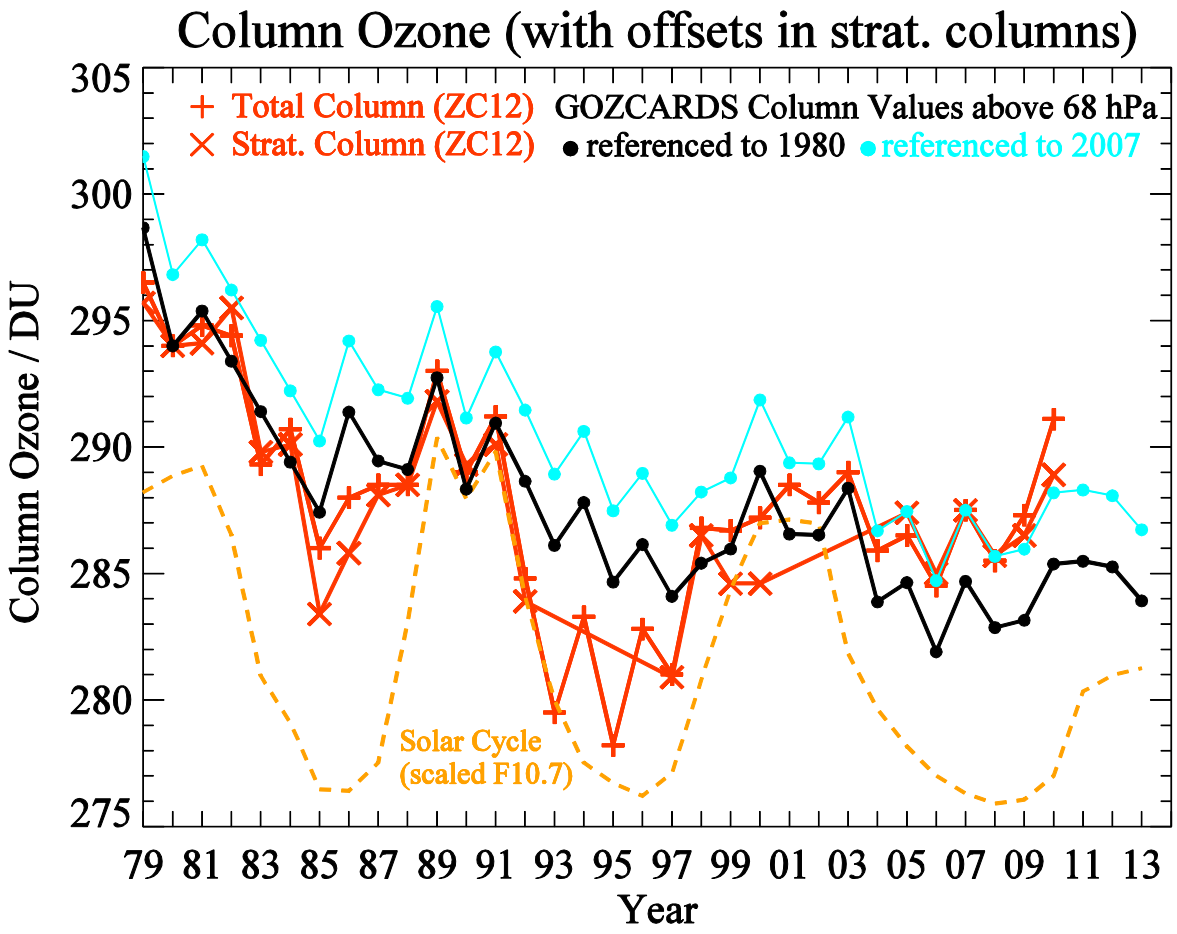
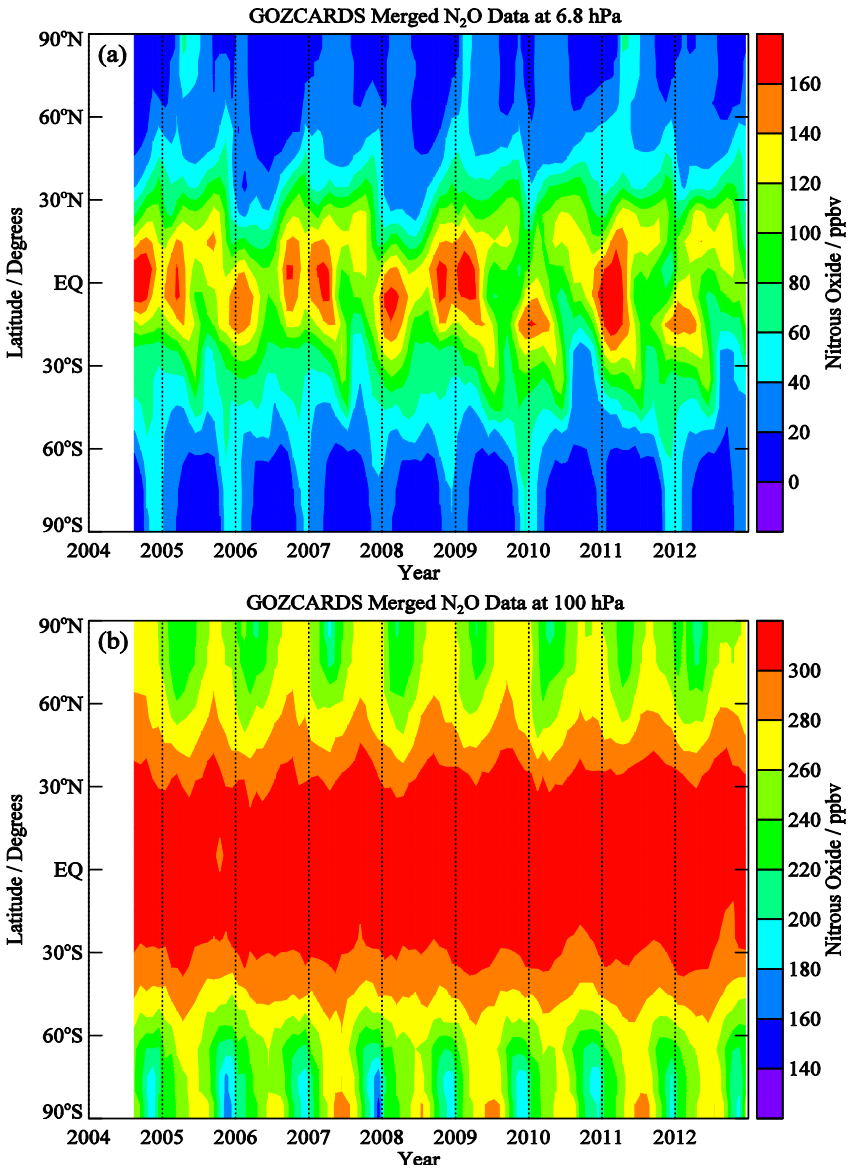


Fig. 22. Near-global (60°S to 60°N) results for average column ozone (total and stratospheric, from *Ziemke and Chandra*, 2012) compared to GOZCARDS O_3 columns above 68 hPa. Stratospheric columns are offset to better match the total column values, in order to more easily compare relative variations versus time; the black dots and red crosses are referenced to the 1980 total column values, while the cyan curves are referenced to 2007 to better illustrate the fits in the later years.



219

220 **Fig. 23.** Time evolution (Aug. 2004 through 2012) versus latitude of GOZCARDS merged N_2O (ppbv) at
221 (a) 6.8 hPa and (b) 100 hPa.

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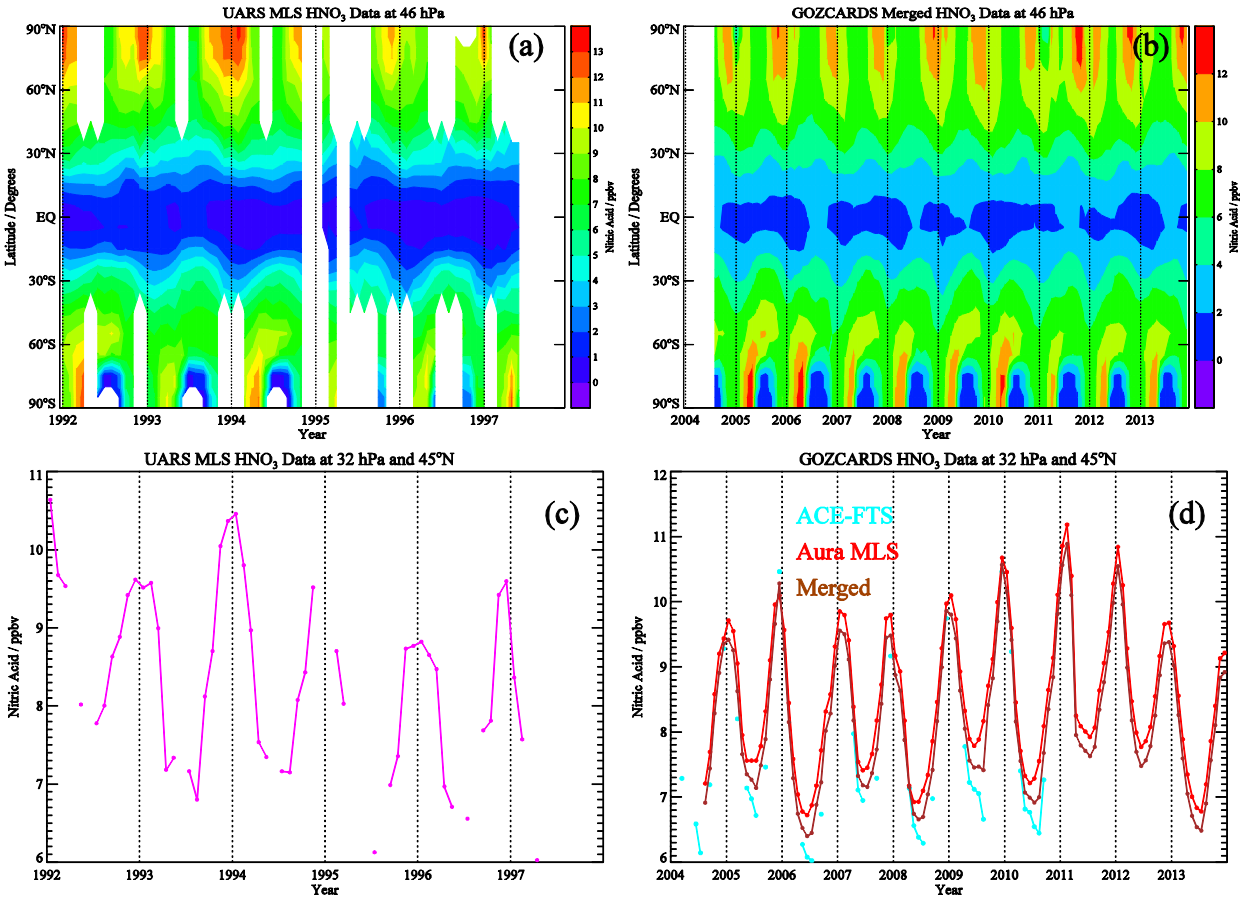

Fig. 24. Sample results display the time evolution of satellite-retrieved HNO_3 (ppbv) for two different periods, 1992-1997 in (a) and (c) versus 2004-2013 in (b) and (d). Panels (a) and (b) are contour plots at 46 hPa from UARS MLS global data and the merged GOZCARDS global data after 2004, respectively; (c) and (d) show time series at 32 hPa and for the 40°N-50°N latitude bin, with (a) from UARS MLS data, and (d) from ACE-FTS, Aura MLS, and the merged combination (between the two source data sets).

Global OZone Chemistry And Related Datasets for the Stratosphere (GOZCARDS): methodology and sample results with a focus on HCl, H_2O, and O_3

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Abstract

We describe the publicly available dataset from the Global OZone Chemistry And Related Datasets for the Stratosphere (GOZCARDS) project, and provide some results, with a focus on hydrogen chloride (HCl), water vapor (H_2O), and ozone (O_3). This dataset is a global long-term stratospheric Earth System Data Record (ESDR), consisting of monthly zonal mean time series starting as early as 1979. The data records are based on high quality measurements from several NASA satellite instruments and ACE-FTS on SCISAT. We examine consistency aspects between the various datasets. To merge ozone records, the time series are debiased by calculating average offsets with respect to SAGE II during periods of measurement overlap, whereas for other species, the merging derives from an averaging procedure based on overlap periods. The GOZCARDS files contain mixing ratios on a common pressure/latitude grid, as well as standard errors and other diagnostics; we also present estimates of systematic uncertainties in the merged products. Monthly mean temperatures for GOZCARDS were also produced, based directly on data from the Modern-Era Retrospective analysis for Research and Applications (MERRA).

The GOZCARDS HCl merged product comes from HALOE, ACE-FTS and (for the lower stratosphere) Aura MLS data. After a rapid rise in upper stratospheric HCl in the early 1990s, the rate of decrease in this region for 1997-2010 was between 0.4 and 0.7%/yr. On shorter timescales (6 to 8 years), the rate of decrease peaked in 2004-2005 at about 1%/yr, and has since levelled off, at ~0.5%/yr. With a delay of 6-7 years, these changes roughly follow total surface chlorine, whose behavior versus time arises from inhomogeneous changes in the source gases. Since the late 1990s, HCl decreases in the lower stratosphere have occurred with pronounced latitudinal variability at rates sometimes exceeding 1-2%/yr. There has been a significant reversal in the changes of lower stratospheric HCl abundances and columns for 2005-2010, in particular at northern midlatitudes and in the deep tropics, where short-term increases are observed. However, lower stratospheric HCl tendencies appear to be reversing after about 2011, with (short-term) decreases at northern midlatitudes and some increasing tendencies at southern midlatitudes.

For GOZCARDS H_2O, covering the stratosphere and mesosphere, the same instruments as for HCl are used, along with UARS MLS stratospheric H_2O data (1991-1993). We display seasonal to decadal-type variability in H_2O from 22 years of data. In the upper mesosphere, the anti-

correlation between H_2O and solar flux is now clearly visible over two full solar cycles. Lower stratospheric tropical H_2O has exhibited two periods of increasing values, followed by fairly sharp drops, the well-documented 2000-2001 decrease, and another recent decrease in 2011-2013. Tropical decadal variability peaks just above the tropopause. Between 1991 and 2013, both in the tropics and on a near-global basis, H_2O has decreased by ~5-10% in the lower stratosphere, but about a 10% increase is observed in the upper stratosphere and lower mesosphere. However, recent tendencies may not hold for the long-term, and the addition of a few years of data can significantly modify trend results.

For ozone, we used SAGE I, SAGE II, HALOE, UARS and Aura MLS, and ACE-FTS data to produce a merged record from late 1979 onward, using SAGE II as the primary reference for aligning (debiasing) the other datasets. ~~Other adjustments were needed in the upper stratosphere to circumvent temporal drifts in SAGE II O_3 after June 2000, as a result of the (temperature-dependent) data conversion from a density/altitude to a mixing ratio/pressure grid.~~ Unlike the 2 to 3% increase in near-global column ozone after the late 1990s reported by some, GOZCARDS stratospheric column O_3 values do not show a recent upturn of more than 0.5 to 1%; continuing studies of changes in global ozone profiles, as well as ozone columns, are warranted.

A brief mention is also made of other currently available, commonly-formatted GOZCARDS satellite data records for stratospheric composition, namely those for N_2O and HNO_3.

1 Introduction

The negative impact of anthropogenic chlorofluorocarbon emissions on the ozone layer, following the early predictions of Molina and Rowland (1974), stimulated interest in the trends and variability of stratospheric ozone, a key absorber of harmful ultraviolet radiation. The discovery of the ozone hole in ground-based data records (Farman et al., 1985) and the associated dramatic ozone changes during southern hemisphere winter and spring raised the level of research and understanding regarding the existence of new photochemical processes (see Solomon, 1999). This research was corroborated by analyses of aircraft and satellite datasets (e.g., Anderson et al., 1989; Waters et al., 1993), and by independent ground-based data. Global total column ozone averages in 2006-2009 were measured to be smaller than during 1964-1980 by ~3%, and larger more localized decreases over the same periods reached ~6% in the southern

hemisphere midlatitudes (WMO, 2011). Halogen source gas emissions have continued to decrease as a result of the Montreal Protocol and its amendments. Surface loading of total chlorine peaked in the early 1990s (WMO, 2011), and subsequent decreases in global stratospheric HCl and ClO have been measured from satellite-based sensors (Anderson et al., 2000; Froidevaux et al., 2006; Jones et al., 2011) as well as from the ground (e.g., Solomon et al., 2006, Kohlhepp et al., 2012). A slow recovery of the ozone layer is expected between the late 1990s and several decades from now, towards pre-1985 levels (WMO, 2011); the robust determination of a long-term global trend requires a sufficiently long and accurate data record. It is desirable to use high quality datasets for ozone and related stratospheric species for a robust documentation of past variations and as constraints for global atmospheric models.

The history of global stratospheric observations includes a large suite of satellite-based instruments, generally well-suited for the elucidation of long-term global change. A review of differences between past and ongoing satellite measurements of atmospheric composition has been the focus of the Stratosphere-troposphere Processes And their Role in Climate (SPARC) Data Initiative (DI); results for stratospheric water vapor and ozone intercomparisons have been published by Hegglin et al. (2013) and Tegtmeier et al. (2013), respectively, to be followed by a larger report on intercomparisons of multiple species. Systematic biases reported in these recent papers tend to mirror past validation work. However, these investigations have not pursued data merging aspects or the creation of long-term records.

Under the Global OZone Chemistry And Related Datasets for the Stratosphere (GOZCARDS) project, we have created monthly zonally averaged datasets of stratospheric composition on a common latitude/pressure grid, using satellite-based limb viewing instruments launched as early as 1979 (for ozone data in particular) and now continuing with instruments launched about a decade ago. The creation of this Earth System Data Record stays close to the data values themselves. Therefore, spatial or temporal gaps are typically not filled in; various methods can be used to try to produce continuous fits to time series, but we viewed this as being outside the scope of this data record creation. The GOZCARDS products arise from several high quality satellite datasets, namely from Stratospheric Aerosol and Gas Experiment instruments (SAGE I and SAGE II), the Halogen Occultation Experiment (HALOE) which flew aboard the Upper Atmosphere Research Satellite (UARS), the UARS Microwave Limb Sounder (MLS), the

Atmospheric Chemistry Experiment Fourier Transform Spectrometer (ACE-FTS) on SCISAT, and the Aura MLS experiment. Table 1 provides characteristics of the original datasets; validation papers from the instrument teams and other related studies give a certain degree of confidence in these datasets. However, the existence of validation references does not imply that there are no caveats or issues with a particular measurement suite. In this project, we have strived to optimize data screening and mitigate some undesirable features, such as the impact of outlier values or the effects of clouds or aerosols. All source datasets still have shortcomings or imperfections, but we have refrained from arbitrarily removing specific monthly means.

Based on original profiles from the various instruments, GOZCARDS "source" monthly zonal mean values were derived. After data screening, monthly average profiles were created by vertical interpolation onto the GOZCARDS pressure levels, followed by binning and averaging into monthly sets. In order to accomodate the lower vertical resolution of some limb viewers, such as UARS MLS, the GOZCARDS pressure grid was chosen as

$$p(i) = 1000 \times 10^{-\frac{i}{6}} \text{(hPa)} \quad (1)$$

with i varying from 0 to a product-dependent top; this grid width corresponds to ~2.7 km. The high resolution SAGE O_3 profiles were smoothed vertically onto this grid (see Sect. 5). Given the sampling of solar occultation instruments, which typically provide 15 sunrise (SR) and 15 sunset (SS) profiles every day (versus the emission-based sampling from MLS), we used latitude bins of width 10° (18 bins from 80°S-90°S to 80°N-90°N) to construct the monthly zonal means.

After the production of GOZCARDS source data files on the above grid, merged (combined) products were created. This involves the calculation of average biases between monthly zonal means from different source data during periods of overlap, followed by an adjustment (using calculated average offsets) of the time series. Non-zero biases always exist between datasets from different instruments for various reasons, such as systematic errors arising from Level 1 (radiances) or Level 2 (retrievals), different vertical resolutions, or sampling effects. A useful reference regarding the sampling effects, which can arise spatially (within a latitude bin) or temporally (within a month), is the recent work by Toohey et al. (2013). They studied sampling biases from a large suite of satellite-based stratospheric profiling instruments, based on simulations using fully-sampled model abundance averages versus averaged sampled results from sub-orbital track locations. The magnitude of such sampling errors is typically inversely

related to the number of available profiles routinely sampled, so that larger sampling errors arise from occultation than from emission measurements, which often sample thousands of profiles per day. Toohey et al. (2013) found that sampling-related biases can reach 10-15% in some regions/periods, notably at high latitudes when larger atmospheric variability exists. Sofieva et al. (2014) have also discussed sampling uncertainty issues for satellite ozone datasets.

We have observed very good correlations between GOZCARDS ozone and other long-term ozone datasets, such as the Stratospheric Water vapor and OzOne Satellite Homogenized (SWOOSH) dataset (Davis et al., in preparation) and homogenized Solar Backscatter Ultraviolet (SBUV) data; these analyses (along with related work on H_2O) will be discussed elsewhere. Results from GOZCARDS and other data relating to midlatitude ozone trends have appeared (e.g., Nair et al., 2013). Dissemination of trend results arising from analyses of GOZCARDS and other ozone profile data is planned as part of the SI^2N initiative, which stands for Stratospheric Processes And their Role in Climate (SPARC), International Ozone Commission (IOC), Integrated Global Atmospheric Chemistry Observations (IGACO-O_3), and the Network for the Detection of Atmospheric Composition Change (NDACC). Recent results on such ozone profile trend comparisons can be found in Tummon et al. (2014) and Harris et al. (2014).

This paper starts with a discussion of general data screening issues (Sect. 2), and then describes the GOZCARDS data production approach and methodology, followed by some atmospheric results for HCl (Sect. 3), H_2O (Sect. 4), and O_3 (Sect. 5). We provide specific diagnostics that indicate generally good correlations and small relative drifts between the main datasets being used to create the longer-term GOZCARDS merged time series. Section 6 briefly mentions the availability of a few other GOZCARDS products, namely N_2O, HNO_3, and temperatures derived from MERRA fields. The version of GOZCARDS described here is referred to as ESDR version 1.01 or ev1.01. ~~Each product's public GOZCARDS data record has an associated digital object identifier (DOI) along with a relevant dataset reference.~~

2 GOZCARDS source data and data screening

Data provenance information regarding the various measurements used as inputs for GOZCARDS is provided in Appendix A (Sect. A.1).

2.1 GOZCARDS data screening and binning

The screening of profiles for GOZCARDS has largely followed guidelines recommended by the various instrument teams and/or relevant publications; such screening procedures are rarely all described in one convenient location, so we review this briefly here for the various data sets. Data screening can reduce the total number of good profiles below our chosen threshold for flagging zonal monthly means; unless otherwise noted, we only provide monthly means constructed from 15 or more values in a given latitude/pressure bin.

~~For HALOE, cloud contamination may add retrieval artifacts and HALOE profiles were screened for clouds, following procedures described in Hervig and McHugh (1999); values at and below the cloud level (found in the netCDF files) were excluded. Also, HALOE profiles that may occasionally contain artifacts associated with either a faulty trip angle or constant lockdown angle registration were screened out, per recommendations from the HALOE data processing team (see http://haloe.gats-inc.com/user_docs/index.php for details).~~

~~For UARS MLS, we used screening recommendations documented by Livesey et al. (2003). In particular, MLS data points whose retrieved precisions are flagged with a negative sign were discarded, ensuring only a negligible contribution to the retrieval from a priori. Also, our data filtering followed the recommendations regarding the ''MMAF_STAT'' flag for operational status (we only used values of 'G', 't', or 'T' for this flag) and the product-specific ''QUALITY'' flag (for which we only used values equal to 4, thus eliminating bad radiance fits).~~

~~For Aura MLS data screening, the procedures are generally as follows: we only use profiles with even values of the Status field, Quality values larger than documented thresholds (indicating good radiance fits), Convergence values smaller than documented thresholds (indicating good convergence), and positive (unflagged) values of the estimated precisions. Species-specific validation papers give data screening recommendations, with appropriate flag values for Quality and Convergence; see Livesey et al. (2013) for such references and v3.3 data screening updates.~~

~~For ACE-FTS, a list of profiles with data issues is provided by the ACE-FTS team (see https://databace.scisat.ca/validation/data_issues_table.php) and these have been removed from our database.~~ However, we also found it necessary to remove occasional large outlier values that could significantly impact monthly zonal means by adding a bias (or "noise") to the time series if

such screening were not performed; such outliers are not otherwise routinely removed from the original ACE-FTS profiles. Our outlier screening procedure removed values outside 2.5 times the standard deviation, as measured from the median values in each latitude/pressure bin, for each year of data. This was deemed close to optimum by comparing the results to Aura MLS time series (which usually are not impacted by such outliers), as well as to independent zonal means (using 5° latitude bins) provided by the ACE-FTS instrument team. Up to 5% of the profile values in each bin in any given month were typically discarded as a result of this procedure, but the maximum percentage of discarded values can be close to 10% for a few months of ACE-FTS version 2.2 data, depending on year and species. Moreover, because of poor ACE-FTS sampling in the tropics, the threshold value for minimum number of (good) ACE-FTS profiles determining a monthly zonal average was allowed to be as low as 10 for mid- to high latitudes, and as low as 6 for low latitudes (bins centered from 25°S to 25°N). Our zonal mean datasets for ACE-FTS would become too sparse in some years if such lower threshold values were not used; ~~such data sparseness can introduce limitations (larger error bars) in the determination of trends, for example, although we also performed comparisons versus Aura MLS monthly means to provide some level of confidence in the results. In addition, ACE-FTS single profile values were discarded if the associated error was larger than the mixing ratio or smaller than 10^{-4} times the mixing ratio, following recommendations from the ACE-FTS team.~~

The binning of profiles occurs after the screened values are averaged (in each latitude/pressure bin). Negative monthly means have been flagged (set to -999.0) in the GOZCARDS files; while a negative mixing ratio that is smaller (in absolute value) than its associated standard error (or a few times this standard error) can in theory be meaningful, we deem that occasional small negative monthly means are unlikely to be very useful, scientifically.

The organization of profiles on a common pressure grid is straightforward when pressure values are present in the original files, as is the case for most data used here. Also, the vertical resolutions are similar for most of the instruments used for GOZCARDS (see more details in each species-specific section). The UARS MLS, HALOE, and Aura MLS native pressure grids are either the same as or a superset of the GOZCARDS pressure grid, so these datasets were readily sampled for the construction of the GOZCARDS monthly means. For ACE-FTS profiles, pressures are provided along with the fixed altitude grid, and we used linear interpolation versus

log(pressure) to convert profiles to the GOZCARDS grid. More details are provided in the O_3 section for SAGE I and SAGE II, for which density versus altitude is the native representation.

3 GOZCARDS HCl

3.1 GOZCARDS HCl source data records

We used HCl datasets from HALOE, ACE-FTS and Aura MLS to generate the monthly zonal mean source products for GOZCARDS HCl.

For the screening of HALOE HCl profiles, in addition to the procedures mentioned in Sect. 2, a first-order aerosol screening was applied: all HCl values at and below a level where the 5.26 μm aerosol extinction exceeds 10^{-3} km^{-1} were excluded.

For Aura MLS, the ongoing standard HCl product is retrieved using band 14 rather than band 13, which was used to measure HCl for the first 1.5 years after launch, but started deteriorating rapidly after Feb. 2006. Validation and error characterization for the Aura MLS HCl product (version 2.2) were provided by Froidevaux et al. (2008). The MLS version 3.3/3.4 HCl data used here (see Livesey et al., 2013) compare quite well with v2.2, with average biases within 5% in general. A high bias exists in MLS HCl versus aircraft data at 147 hPa at low latitudes (Froidevaux et al., 2008). Such regions with large uncertainties and biases are avoided (flagged) for the production of the GOZCARDS merged HCl dataset.

The use of the GOZCARDS source files for Aura MLS HCl, like the use of original Level 2 MLS HCl files, is not recommended for obtaining realistic trends in the upper stratosphere (at pressures < 10 hPa), even if monthly mean MLS HCl in this region displays reasonable values in comparison to other satellite-based measurements. Aura MLS switched to a backup band (band 14) to retrieve the daily HCl measurements after band 13 (originally targeted specifically for HCl) showed signs of rapid degradation in early 2006; as the remaining lifetime for band 13 is expected to be very short (days as opposed to weeks), this band has only been turned on for a few days since February 2006. However, for pressures ≥ 10 hPa, the long-term (band 14) HCl data now being routinely produced is deemed to be robust (because of the broader emission line

in this region, in comparison to the measurement bandwidth). These considerations have implications for how we treat MLS HCl upper stratospheric data in terms of the merging process.

Past validation studies have compared MLS HCl (v2.2), ACE-FTS (v2.2) and HALOE (v19) datasets using coincident pairs of profiles; such work was described by Froidevaux et al. (2008) for MLS HCl validation and by Mahieu et al. (2008) for ACE-FTS HCl validation. HALOE HCl values were found to be biased low by ~10-15% relative to both MLS and ACE-FTS, especially in the upper stratosphere; this low bias versus other (balloon- and space-based) measurements had been noted in past HALOE validation studies (Russell et al., 1996). Also, HALOE (v19) and ACE-FTS (v2.2) HCl data tend to lose sensitivity and reliability for pressures less than ~0.4 hPa.

3.2 GOZCARDS HCl merged data records

HCl is a good candidate for merging the main satellite data that have provided this measurement since 1991. Indeed, one can benefit from the strengths of all datasets in the lower stratosphere, but rely on HALOE and ACE-FTS for upper stratospheric trends, because of the Aura MLS HCl trend detection issue mentioned above. Aura MLS HCl time series were not included in the merging at pressures less than 10 hPa, so after November 2005, the GOZCARDS HCl upper stratospheric trends are dictated only by changes in ACE-FTS abundances. However, in order to derive the systematic offsets needed to adjust the time series from these three instruments in a continuous way (in the pressure dimension), we used the absolute Aura MLS HCl measurements at all pressure levels in 2004 and 2005, during the overlap period between the three instruments.

Figure 1 illustrates the merging process for HCl at 32 hPa for the 45°S latitude bin (which covers 40°S to 50°S). Given that there exists very little overlap between the three sets of measurements in the same months in 2004 and 2005, especially in the tropics, a simple 3-way averaging of the datasets is not practical and would lead to significant data gaps. Our methodology is equivalent to averaging all three datasets during this period (if one had full coverage from all datasets), but we use Aura MLS as a transfer dataset. This was done by first averaging ACE-FTS and Aura MLS data, where the datasets overlap, and then including the third dataset (HALOE) into the merging process with the intermediate (temporary) merged data. Although HALOE HCl is believed to be biased too low, modifying the HALOE values to somehow match ACE-FTS or Aura MLS values or a combination of these two datasets was

deemed to be too subjective. The combined weight of the other two datasets leads to a merged HCl dataset that is generally further away from HALOE than it is from either ACE-FTS or Aura MLS. The top left panel in Fig. 1 shows monthly zonal average GOZCARDS source data for HALOE, ACE-FTS, and Aura MLS during the overlap period, from Aug. 2004 (when Aura MLS data started) through Nov. 2005 (when HALOE data ended). The top right panel illustrates the result of step 1 in the merging procedure, with the temporary merged data values (orange) resulting from the adjustment of ACE-FTS and Aura MLS values to the mean reference indicated by the black dashed line; this reference is simply the average (over the overlap period) of these two datasets, formed from the average of the points which overlap during the same months, meaning whenever ACE-FTS obtained monthly data (since Aura MLS HCl means exist every month). The middle left panel shows the result of step 2, namely the merged values (brown) that arise from merging HALOE values with the temporary merged values (orange) from step 1. In this second averaging step, we weigh the intermediate merged values by 2/3 and HALOE values by 1/3 (leading to a mean reference illustrated by the dashed black line), in order for this process to be equivalent to averaging all three datasets, each with a weight of 1/3. The middle right panel shows the source data along with the final merged values during the overlap period. A simple mathematical description of the above procedure is provided in Appendix A. The bottom panel shows the same datasets but for 1991 through 2012, after the calculated additive offsets are applied to the whole source series, thus debiasing the datasets; these adjusted time series are then merged (averaged) together wherever overlap exists. We tested this procedure by using one or the other of the two occultation datasets as the initial one in step 1, and results were not found to differ appreciably. This methodology is used as well for the same three datasets for the H_2O merging; we have also checked our procedures and results by using two independent calculations from different institutions. We also found that the use of multiplicative adjustments generally produces very similar results as additive offsets. Some issues were found on occasion with multiplicative offsets, when combining very low mixing ratios, but additive offsets can also have drawbacks if the merged values end up being slightly negative, notably as a result of changes that modify the already low HCl values during Antarctic polar winter. This occurs on occasion as additive offsets tend to be weighted more heavily by the larger mixing ratios found during non-winter seasons; as a result, we decided not to offset the

lower stratospheric HCl source datasets in the polar winter seasons at high latitudes for any of the years (for interannual consistency). Procedural details regarding the merging of HCl data are summarized in the Supplementary material.

In Fig. 2, we display the offsets that were applied to the three HCl source datasets as a result of the merging process in each latitude/pressure bin; a positive value means that a dataset is biased low and needs to be increased (on average) by the offset value. These offsets show that in general, ACE-FTS and Aura MLS HCl values were adjusted down by 0.1-0.2 ppbv (a decrease of about 2-10%), while HALOE HCl was adjusted upward by 0.2-0.4 ppbv. Offset values tend to be fairly constant with latitude and the sum of the offsets equals zero. The generally homogeneous behaviour versus latitude is a good sign, as large discontinuities would signal potential issues in the merging (e.g., arising from large variability or lack of sufficient statistics). Figure S1 provides more detailed examples of some of these (upper and lower stratospheric) offsets versus latitude, including standard errors based on the variability in the offsets from month to month during the overlap period. Error bars in the offsets provide an indication of the results' robustness. Another indication of first-order compatibility between datasets is provided by a comparison of annual cycles. Figure S2 provides average annual cycle amplitudes obtained from simple regression model fits to HALOE, ACE-FTS, and Aura MLS series over their respective periods. While there are a few regions where noise or spikes exist (mainly for ACE-FTS), large annual amplitudes in the polar regions occur in all the time series; this arises from HCl decreases in polar winter, followed by springtime increases.

A more detailed analysis of interannual variability and trend consistency is provided from results in Fig. 3, which shows an example of ACE-FTS and Aura MLS time series. ~~We note that no v2.2 ACE-FTS data (for any species) are used after September 2010, because of a data processing problem; a fully updated version of ACE-FTS data was not available when the GOZCARDS data records were constructed.~~ We have used coincident points from these time series to compare the deseasonalized anomalies (middle panel in Fig. 3) from both instrument series; correlation coefficient values (R values) are also computed. In the Fig. 3 example, very good correlations are obtained and no significant trend difference between the anomalies (bottom panel) is found for ACE-FTS and Aura MLS HCl. A global view for all latitude/pressure bins of these correlations and drifts is provided in Fig. 4, where the top panel gives R values for

deseasonalized anomalies, and the bottom panel gives the ratio of the difference trends over the error in these trends. The results in Fig. 4 confirm that there are significant trend differences between the upper stratospheric HCl time series from ACE-FTS and that of Aura MLS (as a reminder, we did not use Aura MLS HCl for pressures less than 10 hPa). Fig. 4 also shows very low correlation coefficients from the deseasonalized HCl series in the uppermost stratosphere, because Aura MLS HCl exhibits unrealistically flat temporal behavior, whereas ACE-FTS HCl varies more. In the lower stratosphere, there is generally good agreement between the ACE-FTS and Aura MLS HCl time series, with R values typically larger than 0.7 and difference trend to error ratios smaller than 1.5. The few low R values for 100 hPa at low latitudes likely reflect more infrequent ACE-FTS sampling and some (possibly related) outlier data screening issues.

Figure S3 illustrates GOZCARDS merged 46 hPa HCl variations versus time; there is clearly a much more complete global view (with no monthly gaps) after the launch of Aura MLS. Gaps at low latitudes in 1991 and 1992 are caused by post-Pinatubo aerosol-related issues in the HALOE record, and gaps in later years arise from the decrease in coverage from UARS. In the upper stratosphere, there are more gaps compared to 10 hPa and below, as a result of the much poorer tropical coverage from ACE-FTS and the elimination of MLS data in this region.

An indication of systematic errors in the merged values can be obtained by providing estimates of the range of available monthly mean source data. We have made such a calculation, although these error values are not part of the public GOZCARDS data files. For each bin, we computed the ranges of monthly means above and below the merged values that include 95% of the available source data monthly means. These error bars are not usually symmetric about the merged values, especially if one dataset is biased significantly in relation to merged values. We did not have enough datasets here to consider a more statistical approach (such as actual standard deviations among source datasets). Figure 5 shows the result of such a systematic error calculation at 46 hPa for the 35°S latitude bin. The lower shaded region range gives the lower bound, determined by HALOE data, and the upper limit of the grey shading originates from ACE-FTS data. Figure 6 shows contour plots of these estimated systematic errors in HCl for all latitudes and pressures. These are fairly conservative error bars; however, even the source data averages at the 95% boundaries have their own systematic errors (rarely smaller than 5%), so our estimates do not really encompass all error sources. Error bars representing a range within which

95% of the source data values reside (see Figs. 5 and 6) can be a useful guide for data users or model comparisons; users can readily calculate such ranges (or we can provide these values).

~~Other quantities are provided in the netCDF GOZCARDS files, which are composed of one set of individual yearly files for all source datasets, and one set of yearly files for the merged products. The main data quantities are monthly averages, plus standard deviations and standard errors for these means. The GOZCARDS source files also provide the number of days sampled each month as well as minimum and maximum values for the source datasets. Other information includes average solar zenith angles and local solar times for individual sources. Note that for the species discussed here, sunset and sunrise occultation values in the same latitude bin during a given month are averaged together. Finally, formulae for monthly standard deviations of the merged data are given in Appendix A, where sample time series of the standard deviations and standard errors (not systematic errors) for both source and merged HCl data are also shown.~~

3.3 GOZCARDS HCl sample results and discussion

Stratospheric HCl is important because it is the main reservoir of gaseous chlorine and it can be used to follow the chlorine budget evolution over the past decades. This includes a significant increase before the mid-1990s as a result of anthropogenic chlorofluorocarbon (CFC) production, followed by a slower decrease as a result of the Montreal Protocol and subsequent international agreements to limit surface emissions that were correctly predicted to be harmful to the ozone layer (Molina and Rowland, 1974; Farman et al., 1985).

In Fig. 7, we provide an overview of the HCl evolution since 1991, based on GOZCARDS average merged HCl for 3 different latitude regions at 4 pressure levels, from the upper stratosphere to the lower stratosphere. In the upper stratosphere (at 0.7 hPa shown here), the rapid early rise in HCl was followed by a period of stabilization (1997-2000) and subsequent decreases. ~~The GOZCARDS HCl time series for pressures less than 10 hPa stop in September 2010 because after this, v2.2 ACE-FTS data were halted, due to technical retrieval issues with that data version.~~ Rates of decrease for stratospheric HCl and total chlorine have been documented based on such satellite-based upper stratospheric abundances, which tend to follow tropospheric source gas trends with a time delay of order 6 years, with some uncertainties in the modeling of this time delay and related age of air issues (Waugh et al., 2001; Engel et al., 2002;

Froidevaux et al., 2006). As summarized in WMO (2011), the average rate of decrease in stratospheric HCl has typically been measured at -0.6 %/yr to -0.9 %/yr, in reasonable agreement with estimated rates of change in surface total chlorine; see also the HCl upper stratospheric results provided by Anderson et al. (2000) for HALOE, Froidevaux et al. (2006) for the one and a half year Aura MLS data record (from the initially used primary band), and Jones et al. (2009) and Brown et al. (2011) for a combination of HALOE and ACE-FTS datasets. The WMO (2011) summary of trends also includes results from column HCl data at various NDACC Fourier transform infrared (FTIR) measurement sites; see Kohlhepp et al. (2012) for a comprehensive discussion of ground-based results, showing some scatter as a function of latitude. Figure 7 demonstrates that a global-scale decline in mid- to lower stratospheric HCl is visible since about 1997. We also notice that at 68 hPa in the tropics, the long-term rate of change appears to be near-zero or slightly positive. In addition, there are shorter-term periods in recent years when an average increasing "trend" would be inferred rather than a decrease, in particular, see the northern hemisphere data from 2005 through 2012 at 32 hPa.

To quantify the rates of change further, we created deseasonalized GOZCARDS merged monthly zonal mean HCl data for the different latitudes, and we show in Fig. 8 the linear rate of change that results from simple fits through such series (averaged into 20°-wide latitude bins). The long-term trends (1997 through 2013 for the lower stratosphere, and 1997 through 2010 for the upper stratosphere) are generally negative and between about -0.5%/yr in the upper stratosphere and -1%/yr in the lower stratosphere, depending on latitude. Some separation between northern and southern hemisphere results is observed in the lower stratosphere, with smaller trends in the northern hemisphere. Also, the scatter increases for 68 to 100 hPa and some positive (or essentially zero) trends occur at low latitudes in this region; however, we have less confidence in the results at 100 hPa, given the larger scatter and error bars in that region (and the smaller abundances). ~~Results at more polar latitudes (not shown here) tend to follow the adjacent midlatitude bin results, but with more scatter (and larger error bars), especially for shorter time periods. To explore these rates of change in the lower stratosphere in more detail, Fig. 9 shows the same type of analysis as Fig. 8 for three other time periods and for pressures of 10 hPa or more: (a) for a decade of data from 2003 through 2012, (b) for a shorter 6-yr period from 2006 through 2011, and (c) for the most recent 6-yr period from 2008 through 2013. For the results in~~

~~(a), a decadal decrease is still observed for the southern hemisphere bins and some of the tropics in the upper region, but increases can be detected in the northern hemisphere and at the higher pressures in the tropics. In (b), we see an accentuation of this hemispheric asymmetry in the short-term rates of change, with large positive changes in the northern hemisphere, and values on both negative and positive sides between 1 and 3 %/yr in many cases; during this past decade, this 6-yr period (2006-2011) is near the temporal peak of this asymmetric lower stratospheric behavior. In the most recent 6-yr period, however (see (c)), the rates of change have decreased for all the latitude bins shown, with all results from 10 to 68 hPa under 0.5 to 1%/yr (absolute value).~~ Without assigning an exact linear "trend" from these simple analyses, we illustrate here that there is considerable variability in lower stratospheric HCl short-term behavior, especially after 2005. Such lower stratospheric changes in HCl have been captured in column HCl FTIR data, as demonstrated by Mahieu et al. (2013, 2014). In the latter reference, it is shown that total column (FTIR) results and GOZCARDS lower stratospheric HCl trends agree quite well; also, these authors' analyses imply that a relative slowdown in the northern hemispheric circulation is responsible for these observed recent changes in the lower stratosphere. However, we note from Fig. 7, that such changes in lower stratospheric HCl appear to be fairly short-term in nature, with an apparent reversal in behavior occurring at both northern and southern midlatitudes since 2011 (e.g., at 32 hPa). The lower stratospheric changes are distinct from the upper stratospheric long-term decrease, which we expect to continue, as long as the Montreal Protocol agreements are fulfilled worldwide and total surface chlorine emissions keep decreasing.

The rate of change analyses above were repeated and shown in Fig. 10 for sliding time periods centered on different years (e.g., a 6-yr average for 2004 means an average from 2001 through 2006) in the upper and lower stratosphere for various latitude bins (covering 50°S to 50°N in 10° steps). As observed in Fig. 10(a), the sliding 6-yr results indicate that there has been an acceleration in the rate of decrease of upper stratospheric HCl between 2000 and 2004, followed by a flatter period until 2010 (this being the last year of GOZCARDS data available for the upper stratosphere, with a 6-yr period centered at the start of 2008). The rate of upper stratospheric HCl change reached a maximum close to -1%/yr, and has retreated to values near -0.5%/yr in more recent years. This is roughly in agreement with time-shifted curves showing the rates of change for surface total chlorine based on National Oceanic and Atmospheric

Administration (NOAA) surface data (Montzka et al., 1999), as shown in Fig. 10 (upper panels, green and purple curves) with the Earth System Research Laboratory Global Monitoring Division (website) data, time shifted by 6 or 7 years to approximately account for transport delays into the upper stratosphere. The tropospheric source gases for chlorine have also shown a reduction in the rate of decrease during the 2nd half of the past decade, as discussed by Montzka et al. (1999) and summarized more recently in WMO (2011). As discussed in the latter report, this arises from a combination of factors, including the initial rapid decrease in methyl chloroform (which now plays a much smaller role), slower rates of decrease from the sum of CFCs in more recent years, and increases in hydrochlorofluorocarbons (HCFCs), along with small contributions from very short-lived species, all of which requires continued monitoring. In Fig. 10, the lower stratospheric response is summarized (panel (b)) by considering the rates of change in partial column density between 68 hPa and 10 hPa. The lower stratospheric rates of change show more variability with latitude than in the upper stratosphere for short (6-yr) time periods, and a hemispheric asymmetry exists, peaking in 2009, when positive tendencies are seen in the northern hemisphere, as opposed to decreases in the south. ~~Figure 10 (panels (c) and (d)) also displays the sensitivity to the time period chosen, as we average the different latitudinal results (from the left panels) and add 8-yr sliding periods to this analysis of HCl changes.~~ The near-global results are not too dependent on whether 6-yr or 8-yr periods are used, but longer periods tend to smooth out the rates of change; interannual changes, including those arising from the quasi-biennial oscillation (QBO), will affect short-term results, especially in the lower stratosphere. It is worth noting (Fig. 10) that the patterns in the upper and lower stratosphere are qualitatively similar, and that rates of change in surface emissions will impact both regions, but carefully disentangling this from changes in the dynamics and in other constituents (e.g., CH_4) that can affect the partitioning of chlorine species will require more analyses and modeling.

4 GOZCARDS H_2O

4.1 GOZCARDS H_2O source data records

We used water vapor datasets from HALOE, UARS MLS, ACE-FTS, and Aura MLS to generate the monthly zonal mean source products for GOZCARDS H_2O.

494 In addition to the data screening procedures mentioned in Sect. 2, screening of HALOE H_2O
495 data for high aerosol extinction values was performed, in a way very similar to the method used
496 for the creation of merged H_2O for the Stratospheric Water vapor and OzOne Satellite
497 Homogenized (SWOOSH) dataset (Sean Davis, personal communication, 2013). This method
498 (see Fig. S4) screens out anomalous HALOE H_2O values that occurred mainly in 1991-1992,
499 when the aerosol extinction near 22 hPa exceeded $5x10^{-4}$ km^{-1}; for pressure levels at and below
500 22 hPa, we have excluded the corresponding H_2O values. Also, for upper mesospheric HALOE
501 data used here, care should be taken during high latitude summer months, as no screening was
502 applied for the effect of polar mesospheric clouds (PMCs). High biases (by tens of percent) in
503 H_2O above ~70 km have been shown to occur as a result of PMCs in the HALOE field of view
504 (McHugh et al., 2003). Indeed, monthly mean values larger than 8-10 ppmv are observed in
505 GOZCARDS H_2O merged data and in HALOE source data for pressures less than ~0.03 hPa. A
506 more recent HALOE data version (version 20), or the version labeled VPMC based on the above
507 reference, could be used to largely correct such PMC-related effects, although this was not
508 implemented for GOZCARDS H_2O. The Aura MLS and ACE-FTS measurements, obtained at
509 longer wavelengths than those from HALOE, do not yield such large H_2O values; a rough
510 threshold value of 8.5 ppmv could also be used (by GOZCARDS data users) to flag the pre-2005
511 merged dataset.

512 UARS MLS stratospheric H_2O for GOZCARDS was obtained from V6 (or V600) H_2O data.
513 This data version is identical to the original prototype (named V0104) from Pumphrey (1999),
514 who noted that UARS MLS H_2O often exhibits drier values (by 5-10%) than HALOE H_2O (see
515 also Pumphrey et al., 2000). The resulting GOZCARDS H_2O monthly zonal means span the
516 period from Sep. 1991 through April, 1993. We note that a significant fraction of UARS MLS
517 tropical data values at 100 hPa are flagged bad (as a result of diminishing sensitivity).

518 Summarizing briefly past validation results, SPARC WAVAS (2000) analyses pointed out the
519 existence of a small low bias in HALOE stratospheric data versus most other measurements
520 (from satellites or other means), except for UARS MLS. Lambert et al. (2007) showed
521 agreement within 5-10% between Aura MLS version 2.2 stratospheric H_2O and other satellite
522 data, including ACE-FTS H_2O (see also Carleer et al., 2008), as well as for comparisons between
523 Aura MLS and balloon data; Aura MLS H_2O values are slightly larger than HALOE H_2O in the

stratosphere, with differences increasing to 10-15% in the mesosphere. Changes from MLS v2.2 to v3.3 led to an increase of 0.2-0.3 ppmv in stratospheric values (Livesey et al., 2013). Past disagreements between aircraft water vapor measurements have made those datasets somewhat difficult to use as absolute validation of satellite-derived H_2O in the upper troposphere and lower stratosphere (UTLS) (Read et al., 2007, Weinstock et al., 2009). An intercomparison of measurements under controlled chamber conditions has helped to better constrain this issue (Fahey et al., 2014). Very good agreement exists between Aura MLS UTLS H_2O and measurements from Cryogenic Frost point Hygrometers (CFH), as discussed by Read et al. (2007) and Voemel et al. (2007) for MLS v2.2 data. At the lowest level (147 hPa) used here for merged H_2O, the latter study showed a dry bias (by ~10%) in the MLS v2.2 data versus CFH. Recent comparisons by Hurst et al. (2014) of MLS v3.3 H_2O data versus Boulder CFH time series show excellent overall agreement, and no significant trend differences between coincident profile sets. There is therefore support for systematic uncertainties as low as 5% for lower stratospheric MLS data. Aura MLS stratospheric H_2O v3.3 values are slightly larger (by up to ~5%) than the multi-instrument average from a number of satellite datasets, as discussed in SPARC Data Initiative comparisons by Hegglin et al. (2013). No large disagreements in interannual variations were noted by these authors for the GOZCARDS datasets ($p < 150$ hPa). From the mid-stratosphere to the upper mesosphere, excellent agreement between ground-based data from the Water Vapor Millimiter-wave Spectrometer (WVMS) and H_2O profiles from Aura MLS and ACE-FTS has been demonstrated by Nedoluha et al. (2007, 2009, 2011).

4.2 GOZCARDS H_2O merged data records

The merging process for H_2O is nearly identical to the method used for HCl. The main difference is an additional step that merges UARS MLS data with the already combined datasets from HALOE, ACE-FTS, and Aura MLS, by simply adjusting UARS MLS values to the average of the previously merged series during the early (1991-1993) overlap period; see Fig. S5 for an illustration at 22 hPa for the 5°N latitude bin. Typically, this requires an upward adjustment of the UARS MLS H_2O data, as these values are biased low versus most other datasets; nevertheless, the fairly short but global record from UARS MLS helps to fill the time series. After considering the channel drift issues for SAGE II H_2O (and following past advice from the

553 SAGE II team itself), we decided to use caution and did not include that dataset for GOZCARDS
554 merging, as some trend results could be affected to an unknown extent. Also, there is probably
555 some remaining retrieval contamination from volcanic aerosol effects for some time after the
556 volcanic eruptions of El Chichon (1982) and Mt. Pinatubo (1991), as well as after several smaller
557 eruptions; see Bauman et al. (2003) for a review of stratospheric aerosol climatology (1984-
558 1999) and Thomason et al. (2008) for the SAGE II stratospheric aerosol dataset.
559 Minor procedural merging details or issues for H_2O are included in the Supplement. Also,
560 data users should be aware of effects from unequal latitudinal sampling when no MLS data exist,
561 for regions where large latitudinal variations occur, as for H_2O at 147 hPa (the largest pressure
562 value). Indeed, global or latitudinal averages can be significantly biased in certain months and
563 month-to-month variability for such averages increases. This is because of poor sampling of the
564 full latitudinal variability, prior to Aug. 2004; after this, regular sampling exists from MLS every
565 month. Such variations in sampling can become an issue for temporal analyses of latitudinal or
566 global averages, unless additional fits or interpolations to mitigate such effects are undertaken.
567 In Fig. 11, we display the average offsets that were applied to the four H_2O source datasets;
568 these offsets follow previously known relative data biases (mentioned earlier). For example, low
569 biases in UARS MLS H_2O, especially in the mesosphere, were discussed by Pumphrey (1999)
570 and the UARS MLS offsets (see Fig. 11) correct that dataset upward. The application of offsets
571 derived for HALOE and UARS MLS raises the H_2O time series from these instruments, whereas
572 negative offsets lower the H_2O source data from ACE-FTS and Aura MLS. As we found for
573 HCl, the offset values generally display small variations versus latitude and are therefore fairly
574 stable systematic adjustments to the time series. Figure S6 displays the amplitudes of the fitted
575 annual cycles for HALOE, ACE-FTS, and Aura MLS. As for HCl, similar patterns emerge for
576 these datasets. Wintertime descent into the polar vortex regions is responsible for large annual
577 cycles at high latitudes, especially in the mesosphere; also, the seasonal impact of dehydration in
578 the lower stratospheric Antarctic region causes a large annual cycle in Aura MLS high southern
579 latitude data. Figure 12 provides some statistical information, as done for HCl in Sect. 3.2,
580 regarding the correlations and trend differences between ACE-FTS and Aura MLS. There are a
581 few regions with noisier relationships. While slow increases in H_2O are generally observed by
582 both instruments in the stratosphere and mesosphere, the tropical region near 0.1 hPa shows a

583 slight decreasing trend for the ACE-FTS points, thus leading to larger discrepancies; it is not
584 clear what the source of these discrepancies is. While the tropical ACE-FTS data are generally
585 sampled with a significantly lower temporal frequency, the same applies for all pressure levels;
586 however, a few outlier points can have a much larger impact when sampling is poorer. There are
587 also a few other spots, such as near 65°S and 65°N and near 5 hPa with a poor trend value for the
588 difference series, in comparison to the errors; this may be caused by a combination of poorer
589 sampling by ACE-FTS and higher atmospheric variability, which can lead to more scatter. At the
590 highest latitudes in the lower stratosphere, the observed slope differences are more within error
591 bars, but the larger variability means that a longer record is needed to determine if two time
592 series really trend differently. The main point here is to show the dataset characteristics and to
593 point out where the agreement is better or worse than typical. The merged dataset tends to be
594 much closer to Aura MLS in terms of trends because there are usually many more months of
595 Aura MLS data than ACE-FTS data, including the fact that the ACE-FTS time series (data
596 version 2.2) used here was halted for data after late 2010 due to technical retrieval issues.
597 Therefore, the overall impact of ACE-FTS data on the merged H_2O series is fairly small.
598 Figure S7 provides a visual representation of the merged GOZCARDS H_2O fields at
599 3 hPa and 68 hPa, respectively. Well-known features are displayed in these plots, given the good
600 global coverage in the post-2004 period in particular. In the upper stratosphere, descent at high
601 latitudes during the winter months leads to larger H_2O values, and low latitude QBO features are
602 also observed. In the lower stratosphere, one observes dehydration evidence at high southern
603 latitudes in the winter months, as well as a low latitude seasonal "tape recorder" signal; this
604 phenomenon is driven by tropopause temperatures and has been measured in satellite data since
605 the early 1990s (Mote et al., 1996; Pumphrey, 1999). A vertical cross-section of this lower
606 stratospheric tropical (20°S to 20°N) tape recorder in GOZCARDS merged H_2O for 1991-2013
607 is shown in Fig. 13; periods of positive anomalies alternate with negative anomalies, including
608 the post-2000 lows, as well as the most recent decreases in 2012-2013 (see also next section).
609 As we discussed for HCl, we have estimated systematic errors for the merged H_2O product.
610 This is illustrated by the contour plots in Fig. 14; these ranges encompass at least 95% of the
611 monthly mean source data values from HALOE, UARS MLS, ACE-FTS, and Aura MLS above
612 or below the merged series. These errors typically span 5 to 15% of the mean between 100 and

0.1 hPa; errors larger than 30% exist in the tropical upper troposphere (147 hPa), and similarly, large values in the upper mesosphere arise from the low bias in UARS MLS H_2O.

4.3 GOZCARDS H_2O sample results and discussion

Stratospheric H_2O variations have garnered attention in the past two decades, because of the radiative impacts of water vapor in the UTLS and the connection to climate change, as well as the stratospheric chemical significance of H_2O oxidation products. H_2O can influence changes in stratospheric and mesospheric ozone via the HO_x catalytic cycles. H_2O in the UTLS has a significant radiative impact (e.g., Forster and Shine, 2002) and has the potential to influence surface temperature changes in ways that could mitigate surface warming (Solomon et al., 2010) if H_2O exhibits a significant drop, as was observed right after 2000. A decrease of about 1 ppmv was also observed in *in situ* data (Fujiwara et al., 2010; Hurst et al., 2011; Kunz et al., 2013). Randel et al. (2004, 2006) correlated this post-2000 decrease with a decline in tropical cold point temperatures. An increasing trend in stratospheric H_2O since the 1950s (see Rosenlof et al., 2001) will have a surface warming tendency. We expect to see continued studies of the influence of cold point temperatures on stratospheric H_2O and the possible connections to changes in sea surface temperatures (see Rosenlof and Reid, 2008; Garfinkel et al., 2013). Efforts to better understand past changes in H_2O, and their causes and expected impacts, include the references above, and (among others) Dvortsov and Solomon (2001), Shindell (2001), Nedoluha et al. (2003), Urban et al. (2007), Dhomse et al. (2008), Scherer et al. (2008), Read et al. (2008), Tian et al. (2009), Schoeberl et al. (2012), Fueglistaler (2012), Fueglistaler et al. (2013), and the recent review of the tropical tropopause layer by Randel and Jensen (2013). The reconciliation of long-term trends in tropopause temperatures with changes in lower stratospheric water vapor is a task worthy of continued study, using additional datasets as well as model studies.

Individual water vapor datasets have been used here to produce a merged record now spanning more than two decades. Linear trend estimates can be quite sensitive to the starting and ending points of the time series, even for 22 years of data, and simple linear trends do not best describe the variations in stratospheric H_2O over the past two decades. We do not attempt here to characterize trends or to imply that recent tendencies will carry into the next decade or two. Rather, as variability is also of interest to climate modelers, we provide information below

regarding observed decadal-type (longer-term) variability in stratospheric water vapor. Figure 15 illustrates monthly, annual, and longer-term changes in stratospheric water vapor, based on the global GOZCARDS merged H_2O series; this shows the well-known H_2O minimum in the lower tropical stratosphere as well as an increasing vertical gradient in the upper stratosphere (as a result of methane oxidation). As we know from past studies (e.g., Randel et al., 2004), medium- to long-term changes in H_2O are large-scale in nature. However, lower stratospheric H_2O variations are more accentuated at low latitudes, in comparison to near-global (60°S-60°N) results. It has long been known (e.g., from the *in situ* balloon-borne measurements of Kley et al., 1979) that the hygropause is typically located a few km higher than the thermal tropopause; this is consistent with the tape recorder and Brewer-Dobson circulation concepts. We observe low water vapor mixing ratios at 68 hPa in the tropics, in comparison to 100 hPa values (near the tropopause). According to the 22-year GOZCARDS data record, annually-averaged H_2O values in the tropics (20°S-20°N) have varied between about 3.2 and 4.2 ppmv at 68 hPa. The rapid drop between 2000 and 2001 is observed at 100 and 68 hPa, with some dilution of this effect at higher altitudes. There is a clear difference in long-term behavior between the upper stratosphere, where changes in methane should have the clearest influence, and the lower stratosphere, especially in a narrow vertical region above the tropopause, where cold point temperatures and dynamical changes have a significant impact. To first-order, the last few years show ~10% larger values in the upper stratosphere than in the early 1990s, while the opposite holds in the lowest stratospheric region, where a decrease of order 10% is observed over the same period. The long-term upper stratospheric increase carries into the mesosphere (see below). Figure 15 also shows that month-to-month and seasonal variations (thin lines) are usually somewhat larger than the long-term changes in the lower stratosphere, most notably at 100 hPa.

In order to provide longer-term variability diagnostics for water vapor, we show in Fig. 16 the minimum to maximum spread in annual averages (tropics and mid-latitudes) from Fig. 15. These variability diagnostics are provided for the 22-yr period ~~(1992 to 2013) and also separated into the two 11-yr periods (thin and dashed lines); as expected, the 22-yr variability is always larger than the variability in either of the two decadal period subsets~~. We also note that the tropical variability is largest just above the tropopause (here this means at the 68 hPa GOZCARDS

level), where it reaches 20-28% (or 0.8 to 1 ppmv) depending on the time period. Such variability diagnostics should be useful for comparisons to various chemistry climate models.

The longer-term variability in water vapor increases above the stratopause and reaches close to 30% in the uppermost mesosphere, as seen in Fig. 17(a); this plot shows the monthly and annual near-global (60°S-60°N) H_2O variations at 0.01 hPa. Large seasonal changes in this region are driven by vertical advection associated with the mesospheric circulation, with each hemisphere's summertime peaks contributing to the maxima (two per year) in these near-global averages; such seasonal variations were compared to model results by Chandra et al. (1997), based on the first few years of HALOE H_2O data. The strong upper mesospheric variability in annual-mean H_2O is known from previous studies of ground-based and satellite H_2O data (Chandra et al., 1997; Nedoluha et al., 2009; Remsberg, 2010), and this region is where the solar (Lyman α) influence on H_2O is strongest. Figure 17(b) displays the near-global variations in annual upper mesospheric H_2O from 0.1 to 0.01 hPa. We clearly see increased variability in the uppermost mesosphere, and decreases in the mixing ratios as a result of H_2O photodissociation.

5 GOZCARDS ozone

A number of discussions relating to signs of ozone recovery have been presented before (Newchurch et al., 2003; Wohltmann et al., 2007; Yang et al., 2008; Jones et al., 2009; Hassler et al., 2011; Salby et al., 2011, 2012; Ziemke and Chandra, 2012; Gebhardt et al., 2013; Kuttipurath et al., 2013; Kirgis et al., 2013; Nair et al., 2013, 2014; Shepherd et al., 2014, Frith et al., 2014). While there are some indications of small increases in O_3 in the past 10-15 years, further confirmation of an increase in global O_3 and its correlation with column increases, is needed, in order to more clearly distinguish between long-term forcings, notably from the 11-yr solar cycle, slow changes in halogen source gases, temperature changes, and shorter-term variability. Continuing, good long-term ozone datasets are clearly needed for such studies.

5.1 GOZCARDS ozone source data records

We used ozone datasets from SAGE I, SAGE II, HALOE, UARS MLS, ACE-FTS, and Aura MLS to generate the monthly zonal mean source products for GOZCARDS. Due to time constraints, we did not use the newer SAGE II version 7 ozone (see Damadeo et al., 2013) as part of the GOZCARDS merged dataset. Our studies indicate that there are systematic

differences of a few percent between SAGE II V6.2 and V7 O_3 on their native coordinates (number density versus altitude). However, these 2 versions will exhibit different trends, mainly in the upper stratosphere, after the data are converted to mixing ratios on pressure surfaces (as shown later). These differences result mainly from different temperature trends between MERRA and analyses from the National Centers for Environmental Prediction (NCEP), which are used by the SAGE II V7 and V6.2 retrievals, respectively; the main differences between MERRA and NCEP temperatures occur in the upper stratosphere for time periods before 1989 and after mid-2000. ~~After June 2000, SAGE II V6.2 O_3 at upper stratospheric pressures ($\leq$ 3hPa) is not included in our merged data (see discussions in Sect. 5.2). In addition to the general data screening methods (Sect. 2), HALOE O_3 was screened for aerosols based on recommendations from Bhatt et al. (1999). Specifically, the O_3 profiles were screened for instances when either the 5.26 μm aerosol extinction exceeded 10^{-3} km^{-1} or a local aerosol extinction minimum was present near the tropopause; all O_3 values at or below the identified levels were flagged bad.~~

5.1.1 Treatment of SAGE ozone profiles

Both SAGE I and SAGE II used solar occultations during satellite sunrise and sunset to measure vertical profiles of ozone, along with other composition data and aerosol extinction (McCormick et al., 1989; Cunnold et al., 1989). It takes about 1 month for SAGE I and II to provide near global coverage (about 80°N to 80°S), with some dependence on season. The SAGE I measurements started in February 1979 and stopped in November 1981, while SAGE II provided data between October 1984 and August 2005. In the middle of July 2000, SAGE II had a problem in its azimuth gimbal system. Although this was corrected by November 2000, the instrument operation was switched to a 50% duty cycle, with either sunrise or sunset occultations occurring in monthly alternating periods, until the end of the mission.

It has been known that there were altitude registration errors in SAGE I (V5.9) data (Veiga et al., 1995; Wang et al., 1996). To correct this problem, an empirical altitude correction method based on Wang et al. (1996) had been applied to SAGE I (V5.9) data; these corrected SAGE I V5.9 profiles, which had been evaluated in previous trend studies (e.g. SPARC Report, 1998; WMO, 2003), were used to create the GOZCARDS SAGE I product (denoted as version V5.9_rev). We did not use reprocessed version 6.1 SAGE I data (L. W. Thomason, personal

communication) because the altitude registration problems had not been completely fixed and new altitude correction criteria should be derived and validated.

Ozone data screening details for the original SAGE I and SAGE II datasets are provided in the Supplementary material. The number density profiles were converted to mixing ratios on pressure levels by using NCEP temperature and pressure data provided with each profile. Derived ozone profiles were then interpolated to fixed pressure levels on the following grid:

$$p(i) = 1000 \times 10^{-\frac{i}{30}} (\text{hPa}) \quad i = 0, 1, 2, ... \tag{2}$$

Ozone values at each of the 5 levels centered on every GOZCARDS pressure level were then averaged (weighted by pressure) to derive mixing ratios at each GOZCARDS pressure level. By doing this, the SAGE profiles were smoothed to a vertical resolution comparable to that of the other satellite instruments used in this GOZCARDS work. Monthly zonal means were then computed for the SAGE ozone datasets on the GOZCARDS-compatible grid.

5.1.2 Comparisons of ozone zonal means

O_3 differences between SAGE II and other satellites are shown in Fig. S8. Zonal mean differences between SAGE II and HALOE are generally within 5% for 1.5 to 68 hPa at mid-latitudes, and for 1.5 to 46 hPa in the tropics. The relative biases are larger outside those ranges and increase to ~10% near the tropopause and also near 1 hPa. SAGE II data show better agreement with UARS and Aura MLS in the upper stratosphere and lower mesosphere, within 5% up to 0.68 hPa and for latitudes outside the polar regions. Aura MLS O_3 compares better with SAGE II data than does UARS MLS in the tropics for pressures larger than 68 hPa; the high bias in UARS MLS O_3 at 100 hPa has been discussed previously (Livesey et al., 2003). There are no months that include both SAGE II and ACE-FTS data in the northern hemisphere tropics (see the gap in Fig. S8, bottom right panel), largely due to the poorer coverage from ACE-FTS in the tropics. ACE-FTS O_3 shows the largest positive bias (greater than 10%) with respect to SAGE II, for pressures less than 1.5 hPa. The high bias in upper stratospheric ACE-FTS ozone has been mentioned in past validation work using ACE-FTS data (e.g., Froidevaux et al., 2008; Dupuy et al., 2008). The biases shown here are also consistent with recent O_3 intercomparison studies from a comprehensive array of satellite instruments by Tegtmeier et al. (2013). It has been known for some time that the HALOE and SAGE II ozone datasets, which govern the main variations of the

GOZCARDS merged ozone values before 2005, agree quite well (within 5%) in absolute value, and also in terms of temporal trends (Nazaryan et al., 2005), and versus ozonesondes (mostly above ~20 km or ~50 hPa). Larger percentage differences occur in the lowest region of the stratosphere at low latitudes, and especially in the upper troposphere, where HALOE values become significantly smaller than SAGE II data, which are already biased low (by ~50%) versus sondes (Wang et al., 2002); see also Morris et al. (2002), as well as results of SAGE II and HALOE comparisons versus solar occultation UV-Visible spectrometer measurements from long duration balloons (Borchi et al., 2005). We should note here that in this GOZCARDS merging work, we have largely avoided the upper tropospheric region.

Zonal mean differences between SAGE II and Aura MLS show some latitudinal structure between 1 and 3 hPa, with larger (5-10%) biases in the southern hemisphere, especially for 0 to 30°S (see Fig. S8). There are no such features between SAGE II and HALOE or UARS MLS. We found that this results from anomalous NCEP temperatures after 2000, which affect SAGE II data converted from number density/altitude to GOZCARDS VMR/pressure coordinates. Figure 18 shows an example of the ozone series from SAGE II and other satellite data for 10°S to 20°S from 1 to 6.8 hPa. At 1 hPa, SAGE II ozone drifts and is elevated after mid-2000, when compared to HALOE. Similar features are found down to pressures near 3 hPa. These anomalous values can be attributed to abnormal NCEP temperature trends compared to MERRA and HALOE during the same time period (for detailed views, see Figs. S9 and S10). Issues relating to anomalous upper stratospheric NCEP temperature trends were noted by McLinden et al. (2009). Because such NCEP-related artifacts are confirmed by both MERRA and HALOE, we decided not to include in the merging process any SAGE II O_3 values after June 30, 2000 for pressures equal to or less than 3.2 hPa. SAGE II ozone is not significantly affected by the conversion to mixing ratio/pressure coordinates at 4.6 and 6.8 hPa (Fig. 18).

5.2 GOZCARDS ozone merged data records

5.2.1 Methodology for GOZCARDS merged ozone

Ozone measurements from SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS and ACE-FTS, as described in Sect. 5.1, were used to establish a near-continuous monthly zonal mean record

from late 1979 through 2012 for the GOZCARDS merged O_3 product (ESDR version 1.01). ~~The monthly means from each instrument were produced after applying the screening described in Sect. 5.1.~~ The SAGE II dataset was used as a reference standard, since it has the longest period of measurements and has been extensively validated. A GOZCARDS ozone merged data record is constructed by combining these measurements after removing systematic biases with respect to SAGE II. This is done by applying additive offsets to all other instrument series, as determined from average differences between monthly zonal means and SAGE II during overlap time periods. The merged data are then derived by averaging all available adjusted datasets. Because there are gaps in overlap between SAGE II and ACE-FTS monthly mean data in some latitudes (Fig. S7), and as SAGE II ozone VMRs obtained from the vertical grid transformation were affected by anomalous NCEP temperatures after mid-2000 for pressures smaller than or equal to 3.2 hPa, a two-step approach is used to generate the merged product. First, SAGE II data are used as reference for pressures larger than 3.2 hPa to adjust HALOE, UARS MLS and Aura MLS based on overlapping months between 1991 and Nov. 2005; see the method overview schematic in Fig. 19. For $p \leq 3.2$ hPa, SAGE II O_3 is still used as a reference through June 2000, and HALOE and UARS MLS data are adjusted accordingly. This eliminates the effect of anomalous NCEP temperatures on SAGE II ozone and leads to more accurate offsets based on HALOE values, after they have been adjusted to SAGE II. The adjusted HALOE data (denoted as HALOE* in Fig. 19) are then used as a reference to derive estimated offsets for Aura MLS O_3, using the overlap period with HALOE from Aug. 2004 to Nov. 2005. In step 2, a new reference value is derived by averaging all available data from SAGE II, HALOE*, UARS MLS* and Aura MLS*. This new reference value is then used to adjust the ACE-FTS ozone values based on all overlapping months between March 2004 and Nov. 2005. By including Aura MLS in the dataset created in step 1, we obtain more complete spatial and temporal coverage than possible with SAGE II and HALOE, and ensure that there are overlapping months between this combined dataset and ACE-FTS source data. At the end of step 2, the final merged ozone is derived by averaging the temporary merged dataset from step 1 with the adjusted ACE-FTS data values.

5.2.2 Further considerations regarding GOZCARDS merged ozone data

~~Diurnal changes in ozone can affect measurement comparisons and could impact data merging. Measurements and models of the ozone diurnal variation from the lower stratosphere to the~~

mesosphere have been discussed previously (Ricaud et al., 1996; Haefele et al., 2008; Huang et al., 2010). Sakazaki et al. (2013) presented diurnal changes measured by the Superconducting Submillimeter-Wave Limb-Emission Sounder (SMILES), and Parrish et al. (2014) have analyzed ground-based microwave ozone profile variations versus local time in conjunction with satellite datasets. These studies indicate that ozone diurnal variations range from a few percent in the lower stratosphere to more than 10% in the upper stratosphere and lower mesosphere. SAGE II and other occultation instruments observe ozone at local sunrise or sunset, and the retrieved values are generally closer to nighttime values in the upper stratosphere and mesosphere. To characterize systematic differences between satellite data, coincident profiles with small differences in space and time are most often used; an example of mean differences and standard deviations between SAGE II and Aura MLS using both coincident profile and zonal mean methods is provided in Fig. S11. SAGE II and coincident Aura MLS nighttime O_3 values agree within ~5% between 0.46 and 100 hPa, except in the tropical lower stratosphere where comparisons are noisier due to weak O_3 signals and strong dynamical variability. Differences between zonal mean SAGE II and Aura MLS data are very close to the differences from averaged coincident values, except for pressures less than 2 hPa, where differences between the methods increase from a few % to ~10% at 0.3 hPa, consistent with what one expects from the diurnal cycle. Biases between SAGE II and Aura MLS based on coincident profiles versus zonal means could be different by more than 5% in the upper stratosphere and above. Zonal mean differences are likely to be less representative of "true" differences between the two instruments. By combining SAGE II with Aura MLS data adjusted by zonal mean biases, we provide a series adjusted to the average of sunrise and sunset, as measured by SAGE II. However, if Aura MLS data were adjusted by biases obtained using the coincident method, an upper stratospheric offset of several percent and artificial trends due to such a diurnal cycle effect could be introduced.

Even in the absence of diurnal variations, measurements from occultation sensors can yield larger sampling errors than those from more densely-sampled emission measurements (Toohey et al., 2013). The use of long-term data with consistent sampling should be an advantage for trend detection. Avoiding SAGE II data after mid-2000 also mitigates potential artifacts arising from different SAGE II sunrise/sunset sampling patterns versus time. The HALOE sampling remained fairly balanced between SR and SS events over its mission duration, although there

~~were also more data gaps in the later years. Similarly, the Aura MLS ozone data generated here are averaged from local times roughly in the middle of the day and the middle of the night, with repeatable and stable patterns over the years; ACE-FTS sampling patterns are also quite stable.~~

Figure 20 displays the average ozone offsets obtained from the calculated biases versus SAGE II data. The effect of a high bias in upper stratospheric and lower mesospheric ozone from ACE-FTS relative to other datasets is made evident by the need to apply a negative offset as large as 25% to the ACE-FTS series. Most of the stratospheric offsets applied to the other instrument datasets are in the 5-10% range; a lowering of O_3 from UARS MLS, HALOE, and Aura MLS in the lower mesosphere is generally required to match the SAGE II values. Sampling differences and data sparseness may be responsible for larger offsets at the highest latitudes. In these regions, more caution is required with the merged data, which is also less amenable to long-term analyses because of data gaps and larger variability (especially prior to 2004).

As shown in the Supplement (Fig. S12), we observe strong similarities in the ozone annual cycle amplitude patterns from SAGE II, HALOE, ACE-FTS, and Aura MLS over their respective measurement periods (e.g., peaks at midlatitudes near 10 hPa and 1.5 hPa). The middle stratospheric peaks are a result of the annual cycle in oxygen photolysis, whereas temperature variations drive the annual cycle in the upper stratosphere (Perliski et al., 1989). This sort of comparison provides some (first-order) reassurance regarding the consistency of the various datasets. For further details, Fig. 21 provides diagnostics similar to those presented for HCl and H_2O, namely the correlation coefficients and significance ratios for the slopes of the deseasonalized anomaly time series from SAGE II versus HALOE as well as from ACE-FTS versus Aura MLS (for 1992 through 1999, and 2005 through 2009, respectively). These diagnostic results for ACE-FTS and Aura MLS are of a quality that is comparable to the HALOE/SAGE II results; poorer fits occur mostly at high latitudes and in the upper stratosphere. Poorer correlations at upper altitude appear largely tied to a decrease in the amount of valid data in this region (especially at high latitudes), coupled with a relatively small variability. For regions with poorer agreement between ACE-FTS and Aura MLS, we often see small variability in the series from Aura MLS but larger changes (scatter) in the ACE-FTS series. Larger differences in trends between SAGE II and HALOE were noted by Nazaryan et al. (2005) at low latitudes near 50 km; this is also indicated by our simple linear fits (not shown here) to the

GOZCARDS source datasets from these two instruments and the existence of poorer agreements in Fig. 21 (2nd panel from top) for the slope of the differenced (anomaly) series in that region. The existence of good correlations in interannual ozone variations between a large number of satellite measurements was discussed by Tegtmeier at el. (2013). Regarding temporal drifts, Nair et al. (2012) have shown that small drifts (mostly within about ±0.5%/yr for the 20-35 km region) exist between most of the datasets from six ozone lidar sites and coincident HALOE, SAGE II, and Aura MLS measurements; similar results were obtained at two of these sites by Kirgis et al. (2013). Other recent or ongoing studies (in particular, the comprehensive study by Hubert et al., 2015) corroborate the very good stability of the longer-term ozone datasets used for GOZCARDS, which relies most heavily on data from SAGE II and Aura MLS. Jones et al. (2009) also studied the consistency of various satellite datasets (including SAGE II and HALOE) for 1979-2008; they concluded that small relative drifts existed between their averaged series and individual instrument series, although they did find that larger inter-instrument drifts exist in the extra-tropical upper stratosphere. While we feel justified in the use of the longer-term time series and generally robust datasets chosen for GOZCARDS O_3, data users should still note the existence of a few regions with poorer correlations or trend agreement (and, therefore, larger uncertainties) between different satellite ozone datasets, as indicated in Fig. 21. Long-term merged datasets from GOZCARDS and other sources should undergo continued scrutiny from the community, as done recently for trends by Tummon et al. (2014) and Harris et al. (2014). Sample cross-sectional views of two slices through the GOZCARDS merged O_3 field are provided in the Supplement (Fig. S13). Figure 22 shows estimated systematic errors from our calculation of the 95% ranges for the monthly mean source data used here, both above and below the merged values. In this case, as SAGE II is used as a reference dataset, the applied offsets (Fig. 20) correlate quite well with this plot depicting the ranges about SAGE II values. Minimum error bars can be slightly lower than 5% for the middle stratosphere at low latitudes, where ozone values are largest. This view of systematic error bars is generally consistent with results by Tegtmeier et al. (2013), who used a standard deviation measure, based on the larger set of satellite datasets analyzed for the SPARC Data Initiative. They also found that the regions with lowest errors (scatter) are in the middle stratosphere at low to mid-latitudes, where most monthly mean satellite data fit within ±5% of the multi-instrument mean.

5.3 GOZCARDS ozone sample results and discussion

Nair et al. (2013) used regression analyses to compare profile trend results from GOZCARDS merged O_3 at northern midlatitudes versus a combined O_3 dataset from lidar and coincident satellite data at the Observatoire de Haute Provence (OHP), France. They showed that good consistency exists for the decreasing ozone time period, from the early 1980s to 1997, and for the upper stratospheric increase since 1997, but some differences exist in the lower stratosphere during this second time period, when the GOZCARDS results show a near-zero trend in comparison to small positive trends from the combined (and more localized) dataset. The above results for the declining time period agree broadly with earlier work (for the 1979-1997 period) by Jones et al. (2009), who averaged various satellite ozone datasets and produced trend estimates; however, these authors obtained only a small positive, but statistically insignificant, linear trend for the post-1997 phase, most likely because of too short a time series at that time. Gebhardt et al. (2013) analyzed ozone profile trends from SCIAMACHY on ENVISAT, and compared this to trends from Aura MLS, Optical Spectrograph and InfraRed Imager System (OSIRIS) on the Odin satellite, and sondes; their results include the detection of localized ozone increases in the mid-stratosphere at low latitudes; see also Bourassa et al. (2014), who analyzed merged SAGE II and (OSIRIS) observations for 1984-2013, as well as results from Kyrölä et al. (2013) on combined SAGE II and Global Ozone Monitoring by Occultation of Stars (GOMOS) records for 1984-2012, and Eckert et al. (2014), who investigated ENVISAT MIPAS trends for 2002-2012. The shortness of data records since 1997, coupled with relative variability and potential drifts between various measurements may explain some differences in recent trend results, notably for the post-1997 period. More comprehensive analyses from the SI^2N initiative have focused on an intercomparison of profile changes from a variety of datasets, including GOZCARDS and other merged records (Tummon et al., 2014; Harris et al., 2014).

Here, we investigate ozone column results based on the global GOZCARDS dataset, given the work by Ziemke and Chandra (2012), hereafter generally referenced to as ZC12; these authors analyzed total column and stratospheric column data from satellite measurements, and their analyses yielded a rather strong near-global (60°S-60°N) average ozone increase since 1998. Their stratospheric column measurements depend on the convective-cloud differential (CCD) method, which uses Total Ozone Mapping Spectrometer (TOMS) and Ozone Monitoring

Instrument (OMI) column ozone data over convective clouds near the tropopause; below the clouds, little sensitivity exists, so the method can lead to stratospheric column estimates, especially in the tropics, where good cloud-related data of this kind exist. For midlatitudes, their methodology has focused on ozone data over the Pacific, along with a few assumptions relating to cloud heights and longitudinal invariance, in order to try to represent zonal mean stratospheric columns (see also Ziemke et al., 2005). In Fig. 23, we show the near-global total and stratospheric column values from ZC12 (J. Ziemke, private communication, 2013), along with (unscaled) GOZCARDS column densities down to three pressure levels (68, 100, and 215 hPa) for mid-latitudes (30°S-60°S and 30°N-60°N), low latitudes (30°S-30°N), and a near-global range (60°S-60°N). In an absolute sense, the GOZCARDS near-global columns above 215 hPa are larger than the ZC12 stratospheric columns, but quite close to the total column amounts from ZC12. Not too surprisingly, the GOZCARDS column values above 100 hPa are slightly lower than the stratospheric columns from ZC12, as the latter columns (estimated down to cloud tops) will capture more of the lower stratosphere in the extra-tropics. Most of the near-global decrease comes from the midlatitudes, as more of the lower stratospheric column resides in these regions, in an absolute sense. Randel and Thompson (2011) obtained small decreases (-2 to -4% per decade) in tropical lower stratospheric ozone for 1985-2009, from a combination of SAGE II ozone and sonde data from the Southern Hemisphere Additional Ozonesondes (SHADOZ) network. A fairly strong increase is observed from 2008 to 2010 at northern midlatitudes (orange curves); Steinbrecht et al. (2011) attributed large ozone enhancements in that region in 2010 to a coupling between the quasi-biennial and Arctic oscillations (and the North Atlantic oscillation). Long-term halogen source gas reductions that have occurred since the mid-1990s should only lead to an ozone increase of a few DU since 1997 (Steinbrecht et al., 2011).

In Fig. 24, we compare the changes in 60°S-60°N ZC12 column ozone data to the GOZCARDS column amounts above 68 hPa for that region; note that GOZCARDS values do not provide for a continuous long-term time series down to pressures of 100 hPa (or more) in the SAGE I years (1979-1981). To eliminate biases between stratospheric columns as calculated using the CCD methodology and the GOZCARDS fixed bottom pressure approach, we reference all stratospheric columns to the 1980 total column value. These column series include the SAGE I data record and are linearly interpolated between 1981 and 1984, when no GOZCARDS source

966 datasets exist. We observe that the relative changes in GOZCARDS columns follow the ZC12
967 curves within a few DU in the downward phase until about 1992, but the 1992-1997 decrease in
968 total columns does not compare very well. Some of this discrepancy may be because total
969 columns capture a stronger decrease from levels below 68 hPa, not fully represented in
970 GOZCARDS columns; there are also gaps in ZC12 stratospheric columns in 1993-1996 and
971 other years. Focusing on the late period for GOZCARDS data from Aura MLS and ACE-FTS,
972 we also show the GOZCARDS columns above 68 hPa, referenced to 2007 instead of 1980.
973 There is a good match in the variations between GOZCARDS and ZC12 columns during 2005-
974 2010, in agreement with the fact that very good correlations were obtained by ZC12 between
975 Aura MLS column variations and stratospheric column data from the CCD technique. The ZC12
976 values for stratospheric and total columns are in good agreement, although the stratospheric
977 values have gaps when not enough data were present for near-global estimates. Also, the large
978 increase in ZC12 data from 1997 to 1998 is not matched very well by GOZCARDS column data.
979 We removed the solar cycle from the deseasonalized anomalies, as was done in the ZC12
980 study, namely via a regression fit and subtraction of that component from the time series. In the
981 resulting plots (see Fig. 25), a 3-year smoothing is also applied (as done by ZC12). We focus in
982 these plots on the total time period from 1979 onward, and therefore, on the GOZCARDS
983 columns above 68 hPa. While agreement exists at the few DU level between the ZC12 relative
984 changes and the GOZCARDS columns, the apparent ‘recovery’ in the ZC12 datasets is quite
985 large (~3%) and is not matched by the changes in GOZCARDS columns. The latter columns
986 show an increase of less than 0.8% between 2001 and 2011, with some decrease (by ~ 0.5%)
987 from 2011 to 2013. We note that the recent analyses by Sheperd et al. (2014), who used a
988 chemistry-climate model constrained by observed meteorology to investigate potential causes of
989 long-term total column ozone variations, show a partial return, in 2010, towards the 1980 ozone
990 levels (for 60°S-60°N), but not nearly as much as implied by ZC12, neither in the model, nor in
991 the observations. It is possible that the discrepancies lie in the various datasets and their merging;
992 for example, it would be worthwhile to check if homogenized SBUV column O_3 data show
993 results that are substantially different from those of ZC12. Alternatively, the discrepancies could
994 mainly reflect differences between the coverage or meaning of the different ozone columns used
995 here (because of different methodologies, grids and/or sampling to properly determine a near-

global result). Although most column discrepancies are not that large as a percent of the total column values, a better consensus regarding the recovery of near-global ozone columns (and profile values discussed in other recent references) will be desirable in the future.

6 Other GOZCARDS data records

We now briefly mention the three other datasets that were part of the delivery of GOZCARDS records for public dissemination in 2013, namely N_2O, HNO_3, and temperature. For N_2O and HNO_3, the somewhat simpler merging procedure consisted of averaging the source datasets from ACE-FTS and Aura MLS over the overlap time period (Aug. 2004 through Sep. 2010) to obtain the additive offsets for each of the two individual records. We then simply used the correspondingly-adjusted and averaged series to create the merged results; this procedure is the same as we described for the first step in the HCl (or H_2O) merging process.

6.1 N_2O

This data set starts in August 2004, when the Aura MLS data record began; the only dataset after Sep. 2010 is the Aura MLS N_2O (version 3.3) data record, because we no longer had ACE-FTS version 2.2 data after that time due to ACE-FTS data processing issues mentioned earlier. Because of degradation in the main target MLS N_2O band (near 640 GHz) after the first few months of 2013, the N_2O standard MLS product will be reprocessed for the whole Aura MLS period using an alternate measurement band and an updated software version. As discontinuities in the version 3.3 MLS data are introduced after mid-2013, when the standard N_2O product was replaced with results from the 190 GHz band, there are currently no GOZCARDS N_2O zonal mean data after 2012 based on the original (640 GHz) MLS N_2O measurement band.

Validation results for the first few years of Aura MLS and ACE-FTS N_2O data were provided by Lambert et al. (2007) and Strong et al. (2008), respectively. Livesey et al. (2013) provided a minor update regarding the v3.3 Aura MLS N_2O data used here, which show typically small differences (within ±5%) in comparison to v2.2 data. The references mentioned above showed that excellent agreement (mostly within 5%) exists between the stratospheric ACE-FTS and Aura MLS N_2O profiles. Plots showing the average offsets applied to both MLS and ACE-FTS N_2O series as a function of latitude and pressure are provided in Fig. S14. These plots are in

agreement (in magnitude and in sign) with the above-referenced studies; the two datasets yield typical offsets (one half of the average differences) of less than 5%. Also, very good temporal agreement between these two time series (for 2004-2010) is illustrated by the quality of the N_2O diagnostic information displayed in Fig. S15 (computed as for other MLS and ACE-FTS comparisons discussed in this work). This generally shows very highly correlated fields, with insignificant drifts between the two separate time series of deseasonalized N_2O data; the poorest correlations are obtained near 100 hPa in the tropics.

Figure 26 shows sample contour plots for the N_2O merged field (2004-2012); as seen from the bottom panel (100 hPa), wintertime descent brings low N_2O values down at high latitudes (inside the polar vortices). N_2O is a conserved tracer in the lower stratosphere and its variations near the tropopause have implications regarding age of air. Variations in upper stratospheric N_2O are clearly affected by seasonal and dynamical effects; this is evident from the striking semi-annual, annual and QBO-related patterns displayed in Fig. 26 for the 6.8 hPa level (top panel).

6.2 HNO_3

As for N_2O, we merged the HNO_3 data from ACE-FTS (version 2.2) and Aura MLS (version 3.3) from Aug. 2004 onward, and included only the adjusted MLS dataset after Sep. 2010. The average offsets applied to MLS and ACE-FTS time series as a function of latitude and pressure for HNO_3 are provided in Fig. S16. The typical offsets (one half of the average differences) for HNO_3 are less than ~10% (and less than 0.5 ppbv). Despite somewhat larger percent absolute differences than for N_2O between Aura MLS and ACE-FTS HNO_3, there is very good agreement as a function of time between these two datasets in the stratosphere. This is illustrated by the quality of the HNO_3 diagnostic information provided in Fig. S17; the poorest correlations are obtained at or below the tropical tropopause.

Comparisons of v3.3 Aura MLS and v2.2 ACE-FTS nitric acid profiles have shown good agreement (see also Livesey et al., 2013), as the MLS HNO_3 v3.3 values are now generally larger than in v2.2, for which validation results were provided by Santee et al. (2007). Wolff et al. (2008) also compared MLS (v2.2) and ACE-FTS (v2.2) coincident profiles, and obtained similar results; in addition, they demonstrated that very good agreement exists between the HNO_3 profiles from ACE-FTS and coincident profiles from MIPAS on Envisat. Also, comparisons

between Aura MLS HNO_3 (v3.3) profiles and wintertime HNO_3 profiles retrieved by a Ground-based Millimeter-wave Spectrometer (GBMS) in Thule, Greenland, during the first 3 months of 2010, 2011, and 2012 have shown good agreement, mostly within 10-15% (Fiorucci et al., 2013).

Figure 27 (top two panels) displays the HNO_3 fields at 46 hPa from the UARS MLS period (1991-1997) as well as from the 2004-2013 period, for which a merged GOZCARDS product was produced, based on Aura MLS and ACE-FTS source datasets. Also shown (bottom two panels) are time series for 45°N and 32 hPa from both these periods; the bottom right panel includes the source and merged time series. We have performed additional investigations (not shown here) which lead us to believe that small upward adjustments to the UARS MLS HNO_3 values (by about 10%) are needed to better cross-correlate these datasets across the two distinct time periods; such relative biases are within the expected systematic errors. This is based on a consideration of ground-based Fourier Transform infrared column HNO_3 data covering the full time period, as well as past GBMS HNO_3 profile retrievals. Also, Aura MLS and ACE-FTS HNO_3 data match ground-based and other correlative data quite well, and typically better than the (intrinsically poorer quality) UARS MLS HNO_3 data. However, obtaining an optimum global set of adjustments for the UARS MLS nitric acid field will be limited by the number of sites with such ground-based data (as well as by the different vertical resolutions for these datasets versus MLS). More collaborative work regarding such analyses is needed in order to find the optimum adjustments to help tie together these two time periods for this species. Although we did not deliver the UARS MLS HNO_3 source data files for GOZCARDS, we could provide these monthly zonal mean series upon request, keeping the above caveats in mind.

6.3 Temperature

Finally, in terms of the initial set of delivered GOZCARDS products, and for the convenience of stratospheric composition data users, we have used temperatures (T) from the Modern-Era Retrospective Analysis for Research and Applications (MERRA) to produce a monthly mean GOZCARDS temperature data set from 1979 onward. MERRA is a NASA Goddard reanalysis (Rienecker et al., 2011) for the satellite era using Goddard Earth Observing System Data Assimilation System version 5 (GEOS-5); T is from the DAS 3d analyzed state MAI6NVANA, version 5.2 files (such as MERRA300.prod.assim.inst6_3d_ana_Nv.20110227.hdf). Data from

four daily MERRA files (for 00, 06, 12, and 18 hr UT) were averaged to provide daily mean temperature fields (appropriate for a mean time of 09 hr). Vertical interpolation was performed onto the GOZCARDS pressure grid, which, for temperature, covers 30 pressures levels from 1000 hPa to 0.0147 hPa. Averaged values were stored for the 10° GOZCARDS latitude bins, and daily results were binned to create the GOZCARDS monthly temperature data set (version 1.0).

7 Summary and conclusions

We have reviewed the MEaSUREs GOZCARDS project's production of merged data records of stratospheric composition using carefully screened satellite data, starting in 1979 with SAGE I O_3 and continuing with Aura MLS and ACE-FTS data. The source datasets have a high degree of maturity, and we have reinforced our confidence in their usefulness through investigations of various diagnostics (offsets, annual cycle amplitudes, temporal correlations and trend differences of deseasonalized series). We have focused here on the relatively long-term data records for HCl, H_2O, and O_3. These records are publicly available as GOZCARDS ESDR version 1.01 and can be referenced using DOI numbers (Froidevaux et al., 2013b, Anderson et al., 2013, and Wang et al., 2013, for the above species, respectively). The other GOZCARDS data records mentioned here also have dataset references, namely Schwartz et al. (2013) for the MERRA-based temperature records, and Froidevaux et al. (2013c, 2013d) for N_2O and HNO_3, respectively. Table 2 provides a summary of the monthly mean datasets produced for GOZCARDS. Yearly netCDF files were delivered to the GES DISC for public access (see http://mirador.gsfc.nasa.gov, where a README document is also available, see Froidevaux et al., 2013a). Temperature records based on MERRA are also included on the GOZCARDS grid (see Sect. 1). The merging methodology follows from the determination of mean biases (for each pressure level and 10° latitude bin) between satellite instrument datasets, based on the overlap periods between the various series. Each species is treated separately: for ozone, SAGE II data are the chosen reference, whereas for other species, the approach is equivalent to an average of the datasets during the periods of overlap. The merged data files contain the average offset values that were applied to each source data time series, along with standard deviations and standard errors. The GOZCARDS README document provides more details about the GOZCARDS data file quantities, including local time and solar zenith angle information, and a list of days with available data for each month. We have also presented here a compilation of

systematic error estimates about the merged values. While it is difficult to identify error sources specifically, we find that typical estimated systematic errors (ranging from ~5% to 15% for composition data) are consistent with the magnitude of observed relative biases.

The GOZCARDS HCl merged record in the upper stratosphere enables us to track long-term changes in this reservoir for stratospheric chlorine, and by implication, total stratospheric chlorine. The long-term increase in HCl prior to the late 1990s, and the subsequent gentler decrease in the 21st century, are delayed manifestations of changes in the sum of the surface source gas abundances as a result of regulations from the Montreal Protocol (and its amendments). From 1997 to 2010, the average rate of change in upper stratospheric HCl (50°S to 50°N) was about -0.4 to -0.7%/yr. There are smaller rates of decrease and a flattening or slight turn-around after 2003. In the lower stratosphere, where Aura MLS data weigh in heavily, recent short-term variations have shown a flattening out and, in particular for northern midlatitudes and at 50-70 hPa for the deep tropics, a significant reversal and increasing trend, compared to the decrease from the late 1990s to about 2004. Mahieu et al. (2014) have discussed the reversal in total column HCl trends for 2007-2012 (for northern, not southern midlatitudes), based on ground-based FTIR series at various sites; they also showed that column trends agree with those from lower stratospheric GOZCARDS abundances for the appropriate latitude bands. However, lower stratospheric HCl tendencies appear to be reversing in recent years (2011-2014), with decreases at northern midlatitudes (see Fig. 7 for 32 hPa) and some increasing tendencies at southern midlatitudes (Fig. 7). Continued data will be needed to track such short-term changes in HCl, but we expect to see long-term global HCl decreases in both upper and lower stratosphere, as long as the Montreal Protocol and its amendments are adhered to; also, surface chlorine shows smaller rates of decrease in recent years, for reasons that are largely understood (Sect. 3).

For water vapor, we have merged monthly mean datasets for 1991-2013 from the same satellite instruments as for HCl using the same basic methodology, except for the addition of 1991-1993 UARS MLS data, and the inclusion of Aura MLS H_2O data for all pressures. Mostly at the uppermost (mesospheric) altitudes, large variations that are anti-correlated with solar flux are clearly observed over the past two 11-yr solar cycles, as discussed previously by others, using shorter data records. Net long-term trends in lower stratospheric water vapor are quite small if one considers the past 22 years, but there has been considerable interannual change, as

1142 mentioned also in past work using satellite data and Boulder sonde records. Notably, the steep
1143 drop (by 0.5-0.8 ppmv depending on latitude and pressure from 46 to 100 hPa) from 2000 to
1144 2001 is clearly visible in the GOZCARDS record. While the trends have been generally positive
1145 in the decade following 2001 (see Sect. 3), the 68 and 100 hPa levels show equally steep
1146 decreases again from 2011 to 2013 (from ~ 0.5 ppmv for 60°S to 60°N averages to ~ 0.8 ppmv in
1147 the 20°S-20°N bin at 68 hPa); see also Urban et al. (2014). Long-term stratospheric trends may
1148 be observable most readily in the upper stratosphere. In the past 22 years, long-term global H_2O
1149 increases of order 10% are observed in the upper stratosphere and lower mesosphere, whereas a
1150 decrease of nearly 10% has occurred in the lower stratosphere (near 70-100 hPa). However, there
1151 is no regular monotonic change on decadal timescales, especially in the tropical lower
1152 stratosphere, where fairly sharp decreases followed by steadier increases may be a recurrent
1153 pattern. This remains to be better understood, with ongoing global datasets; Fueglistaler (2012)
1154 recently discussed the possibility of stepwise changes in water vapor. We have displayed the
1155 seasonal and decadal variability in stratospheric and mesospheric H_2O based on the GOZCARDS
1156 records for the past 22 years. As one might expect from the well-documented temperature
1157 influence on the tropical lower stratosphere, the H_2O variability based on maximum minus
1158 minimum yearly averages, is largest in the tropics and just above the tropopause. The elucidation
1159 of lower stratospheric water vapor changes over multiple decades is complicated by the
1160 significant low frequency variability in this region, and the occurrence of sudden changes; the
1161 addition of a few years of data can significantly modify trend results. More accurate studies of
1162 seasonal to decadal water vapor variability will be enabled by continuing merged H_2O data from
1163 the lower stratosphere to the upper mesosphere. A reduction in model spread for stratospheric
1164 H_2O is likely easier to achieve than tighter model results for water vapor (and ice water content)
1165 in the upper troposphere (for this region, see the data/model comparisons by Jiang et al., 2012).
1166 We should continue to improve model comparison results for H_2O (see, for example, Gettelman
1167 et al., 2010), although some studies should still focus on non-zonal aspects such as the Asian
1168 monsoon or Western Pacific regions, rather than zonal means like GOZCARDS.

1169 For ozone, we have used measurements from SAGE I, SAGE II, HALOE, UARS MLS, Aura
1170 MLS and ACE-FTS to produce a merged data record from late 1979 onward, after adjusting the
1171 monthly zonal mean series to SAGE II averages. Some complications arose because of the

conversion of the original SAGE II profiles from a density/altitude grid to the GOZCARDS mixing ratio/pressure grid. In particular, we observed temperature-related temporal drifts in the converted SAGE II series, as a result of the NCEP temperature data used in this conversion, mostly in the upper stratosphere after June 2000 (see also McLinden et al., 2009). To circumvent this issue, we used HALOE upper stratospheric O_3 ($p < 4$ hPa) as a reference for July 2000 to November 2005, after applying offsets to the HALOE series. Aura MLS and ACE-FTS provide the continuing data past 2004. The resulting GOZCARDS merged ozone profiles for northern midlatitudes have recently been used in regression analyses (Nair et al., 2013) to reveal significant decreases in the whole stratosphere during 1984-1996. Nair et al. (2014) extended this work to a broader range of latitudes; they found significant increasing trends in upper stratospheric GOZCARDS ozone since 1997, but no significant positive trends in the lower stratosphere. Studies of GOZCARDS (and other) O_3 profile trends have been recently discussed as part of the SI^2N analyses (Tummon et al., 2014; Harris et al., 2014), among other efforts.

Here, we looked into the consistency of column ozone data between stratospheric GOZCARDS results and the study by Ziemke and Chandra (2012), who noted, using a simple regression model, that a fairly rapid change ("recovery'') in near-global ozone columns from TOMS and OMI data could be inferred since the mid-1990s. We show here that, unlike the 2 to 3% net increase in near-global column ozone reported by ZC12 after the late 1990s, the similarly analyzed GOZCARDS record does not show such a strong reversal (or an upturn of more than 0.5-1% since that period). Reasons for these differences could include data coverage or merging-related issues in either dataset, as well as differences in column sensitivities, as column ozone data down to certain pressures (as used in GOZCARDS analyses) cannot be exactly equivalent to stratospheric or total column estimates from ZC12; for example, changes that occur at the lowest altitudes may not be that well captured (for all latitudes) by GOZCARDS-derived columns. Further studies regarding the consistency of various column ozone results and a recovery tendency are warranted; a recent global total ozone study (Sheperd et al., 2014) also points to less of a return towards the 1980 levels than implied by ZC12.

We also briefly described the creation of N_2O and HNO_3 GOZCARDS data records, based on Aura MLS and ACE-FTS monthly mean time series. The agreement between these two instruments' datasets for these species was shown to be generally very good. For HNO_3, UARS

MLS HNO_3 source datasets in the GOZCARDS format are available from the authors. However, a small upward adjustment (of order 10%) to the UARS MLS values is most likely needed based on our preliminary work comparing these time series to HNO_3 column results from FTIR measurements. More detailed work should help determine if adjustments can indeed be made in a more generalized way to the global UARS MLS HNO_3 dataset; lacking this, one should ensure that the error bars reflect the likely biases that can affect the continuity between satellite HNO_3 datasets before and after 2000, given the multi-year gap in satellite coverage for this species.

There is a Supplement related to this article.

Acknowledgements. We dedicate this work to the memory of Professor Derek Cunnold (Georgia Institute of Technology) who was a member of the original GOZCARDS proposal. The GOZCARDS data generation could not have been possible without the past work from instrument teams for SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS, and ACE-FTS, and related data usage documentation. At JPL, we thank Joe Waters for his leadership role in making MLS instruments and datasets possible and Bill Read for his key role; thanks to Vince Perun for MERRA-related work, and to Brian Knosp and Robert Thurstans for database and computer management assistance. Work at JPL was performed under contract with the National Aeronautics and Space Administration. We also thank Kaley Walker and Ashley Jones for comments regarding ACE-FTS data, Gloria Manney and William Daffer for help in making the original ACE-FTS data profiles available, and Joe Zawodny and Larry Thomason for their contributions and comments regarding SAGE data. We acknowledge the work of the GMAO team responsible for MERRA data used to generate the GOZCARDS temperatures, specifically, Steven Pawson and Jianjun Jin for discussions and cross-checks regarding temperature data. We acknowledge Jerry Ziemke for the ozone column data (from Ziemke and Chandra, 2012), and Sean Davis for discussions on data usage and screening, and the creation of long-term series. For early HNO_3-related work connecting ground-based data to MLS datasets, we thank Giovanni Muscari and Irene Fiorucci. We are thankful for the NOAA Earth System Research Laboratory (ESRL) Global Monitoring Division (GMD) website information and data on total surface chlorine. We obtained solar flux data for the Ottawa/Penticton sites from the NOAA National Geophysical Data Center (NGDC) website (www.ngdc.noaa.gov), for which we also acknowledge the National Research Council of Canada.

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Appendix A

A.1. GOZCARDS data provenance

The general origin of the datasets is summarized here. Data coverage from limb sounders (including the instruments used here) is displayed nicely in the work by Toohey et al. (2013).

SAGE I

SAGE I was launched February 18, 1979, aboard the Applications Explorer Mission-B (AEM-B) satellite. SAGE I was a sun photometer using solar occultation (Chu and McCormick, 1979), and it collected a global database for nearly three years on stratospheric aerosol, O_3, and NO_2. For more information, the reader is referred to http://sage.nasa.gov/SAGE1.

SAGE II

SAGE II was launched aboard the Earth Radiation Budget Satellite (ERBS) in October 1984 and its data gathering period ended in August 2005. During each sunrise and sunset, SAGE II measured stratospheric aerosols, O_3, NO_2, and H_2O via solar occultation. This long dataset has proven very valuable in determining past ozone trends. For more information on and data access to the (V6.2) dataset used for GOZCARDS, the reader is referred to http://sage.nasa.gov/SAGE2.

HALOE

Since its launch on September 12, 1991 from the Space Shuttle Discovery until November 2005, UARS HALOE collected profiles of atmospheric composition and temperature. HALOE (Russell et al., 1993) used solar occultation to measure vertical profiles of O_3, HCl, HF, CH_4, H_2O, NO, NO_2, temperature, aerosol extinction, and aerosol composition and size distribution. More information and access to the HALOE data can be obtained from http://haloe.gats-inc.com and http://disc.sci.gsfc.nasa.gov/UARS/data-holdings/HALOE. For GOZCARDS purposes, we have used Version 19 HALOE netCDF data files available at http://haloe.gats-inc.com.

UARS MLS

This instrument observed the Earth's limb in microwave emission using three radiometers, at frequencies near 63, 183 and 205 GHz (Waters, 1993; Barath et al., 1993), providing unique

daily global information on stratospheric ClO, along with other profiles, including O_3, H_2O, HNO_3, temperature, and cloud ice water content. The stratospheric H_2O data ceased on April 15, 1993, after the failure of the 183 GHz radiometer. After March 15, 1994, measurements became increasingly sparse in order to conserve the life of the MLS antenna scan mechanism and UARS power. Data exist until July 28, 1999, although for GOZCARDS, only data through mid-June 1997 are used, as data sparseness and degradation of the 63 GHz radiometer led to less 'trend-quality' data after this. Sampling patterns follow the alternating yaw cycles imposed on MLS by the precessing UARS orbit; MLS measurements were obtained continuously for all latitudes between 34°S and 34°N, with higher latitudes covered in either the northern or southern hemisphere with a roughly 36-day cycle. Livesey et al. (2003) provide more information on the UARS MLS instrument, retrievals, and results. For data access, the reader is directed to the relevant Goddard Earth Sciences and Information Services Center (GES DISC) data holdings at http://disc.sci.gsfc.nasa.gov/UARS/data-holdings/MLS. L3AT data files were used as the basis for the production of the GOZCARDS UARS MLS monthly source datasets.

ACE-FTS

ACE-FTS is the primary instrument onboard the SCISAT satellite, launched on August 12, 2003. It is a high spectral resolution (0.02 cm^{-1}) Michelson interferometer operating from 2.2 to 13.3 μm (750-4400 cm^{-1}); see Bernath et al. (2005) for an overview of the ACE mission. The instrument can simultaneously measure temperature and many trace gases (including all the species mentioned here for GOZCARDS), thin clouds, and aerosols, using the solar occultation technique. ACE-FTS data version 2.2, along with the version 2.2 update for ozone, were used here for GOZCARDS. For access to the public ACE-FTS datasets, with a routine measurement start date of March 2004, the reader is directed to http://www.ace.uwaterloo.ca.

Aura MLS

MLS is one of four instruments on NASA's Aura satellite, launched on July 15th 2004. Aura MLS is a greatly enhanced version of the UARS MLS experiment, providing better spatial coverage, vertical resolution, and vertical range, along with more continuous data over its lifetime (and with ongoing measurements at the time of writing). The instrument includes

radiometers at 118, 190, 240, and 640 GHz, and a 2.5 THz module (Waters et al., 2006). Aura MLS provides measurements of many chemical species, cloud ice, temperature and geopotential height. Continuous measurements have been obtained since August 2004, with the exception of OH, for which sparser measurements exist since August 2010, in order to preserve the life of the THz module. For more information and access to the Aura MLS datasets, the reader is referred to http://disc.sci.gsfc.nasa.gov/Aura/data-holdings/MLS. For GOZCARDS, we use the currently recommended Aura MLS data versions (version 2.2/2.3 for ozone and 3.3/3.4 for other species).

A.2. Calculation details for the iterative merging procedure

Given three time series, the merging procedure that we use first combines two out of the three time series, $y_1(i)$ and $y_2(i)$ (where index i represents time for each monthly mean value in a given latitude/pressure bin). We first obtain the temporary merged series $m_1(i)$ via:

$$m_1(i) = (1/2)\,(y_1(i) + y_2(i)) \quad (1)$$

with the average offsets for $y_1(i)$ and $y_2(i)$ being $(1/(2\,n_{12}))\,\Sigma\,(y_1(i) - y_2(i))$ and -1 times this value, respectively; n_{12} is the number of overlapping data points between the two time series. Then, we merge together the time series $m_1(i)$ and $y_3(i)$, keeping the weightings equal for all 3 time series (1/3 for each), so that we calculate the new merged time series $m(i)$ via:

$$m(i) = w_m\,m_1(i) + w_3\,y_3(i) = (1/3)\,(y_1(i) + y_2(i) + y_3(i)) \quad (2)$$

which will hold if the weights are $w_m = 2/3$ and $w_3 = 1/3$ (given equation (1) for $m_1(i)$). The average reference value (to which the adjustments of $m_1(i)$ and $y_3(i)$ in the 2nd step are made) is given by $(1/n_m)\,\Sigma\,(\,(2/3)\,m_1(i) + (1/3)\,y_3(i))$, where n_m represents the number of (overlapping pairs of) data values used in step 2. For the HCl and H_2O data merging procedure, we always use the Aura MLS time series as one of the first two series involved in the initial merging step, for example as $y_1(i)$, in order to maximize the overlap between the first two series and obtain more robust offset values. Then, we use the 3rd time series; the order used for HALOE and ACE-FTS (i.e., whether we use HALOE or ACE-FTS for y_2 or y_3) makes very little difference.

Calculation of the standard deviation for the merged data values

The average and standard deviation (square root of variance) for each y_k value (i.e. for each monthly zonal mean in a particular lat/p bin) are calculated from equations (3) and (4) below:

$$\bar{y}_k = \frac{1}{n_{yk}} \sum_j y_{kj} \quad (3)$$

and, for the variance,

$$\sigma_{yk}^2 = \frac{1}{n_{yk} - 1} \sum_j (y_{kj} - \bar{y}_k)^2 \quad (4)$$

where index "*j*" corresponds to individual data values within a month, index *k* represents a given instrument (data source), and *n* is the total number of data values for a given bin and source (instrument) time series point in time (or month). Each value $\overline{y_k}$ above is a monthly average (although we also use instead the simpler notation y_k), with standard deviation about the mean σ_{yk}. Now, given the merged series *u(i)* (where index *i* runs over a large number of months), the standard deviation of each merged data point (for a given month) can be obtained by considering the original datasets y_{kj} that were used to construct ***u*.** Specifically, we have the variance for the merged dataset

$$\sigma_u^2 = \frac{1}{n_u - 1} \sum_j (u_j - u_{ref})^2 \quad (5)$$

where u_{ref} is the merged value (which is not necessarily chosen to be the average value $\bar{u}$) and the u_j values represent the union of adjusted data values that make up the merged product, with the index j for this combined dataset covering all values (up to the total n_u) obtained from the original source values y_{kj}. In practice, we do not keep track of the individual data values that went into making the averages for the series y_k that are being merged, and we need to obtain σ_u based solely on the values $\bar{y}_k$, σ_{yk}, and the original number of points for each dataset y_k, namely n_{yk}. If we consider all the original values, we have a combined dataset with n_u points, such that $n_u = \sum_k n_{yk}$. Now, expanding equation (5), we get

$$(n_u - 1)\, \sigma_u^2 = \sum_j (u_j^2 + u_{ref}^2 - 2\, u_{ref}\, u_j) \quad (6)$$

or

$$(n_u - 1)\ \sigma_u^2 = \sum_j u_j^2 + n_u\, u_{ref}^2 - 2\, u_{ref} \sum_j u_j \qquad (7)$$

Expanding (4) for each individual dataset y_k, we get

$$(n_{yk} - 1)\ \sigma_{yk}^2 = \sum_j y_{kj}^2 + \bar{y}_k{}^2 - 2\overline{y_k} \sum_j y_{kj} \qquad (8)$$

which leads to

$$\sum_j u_j^2 = \sum_{k,j} y_{kj}^2 = \sum_k (n_{yk} - 1)\, \sigma_{yk}^2 \ + \sum_k n_{yk}\ \overline{y_k}^2, \qquad (9)$$

so that extracting the variance from equation (7) now leads to

$$s_u^2 = \frac{1}{(n_u - 1)} \left(\sum_k (n_{yk} - 1) s_{yk}^2 + \sum_k n_{yk}\ \overline{y_k}^2 + n_u\, u_{ref}^2 - 2 u_{ref} \sum_k n_{yk}\, \overline{y_k} \right) \qquad (10)$$

The adjusted time series are obtained from the original series y_k as Y_k, and we can write Equation (4) in the same manner for the Y_k data values, namely

$$\sigma_{Yk}^2 = \frac{1}{n_{yk} - 1} \sum_j (Y_{kj} - \overline{Y_k})^2 \qquad (11)$$

with $\sigma_{Yk} = \sigma_{yk}$ as the adjustments (offsets) are performed in an additive manner; if these adjustments were performed using multiplicative factors, those factors would also have to be considered in a multiplicative way to get the new σ_{Yk} values. We can thus write (10) for the adjusted datasets as:

$$s_u^2 = \frac{1}{(n_u - 1)} \left(\sum_k (n_{yk} - 1) s_{yk}^2 + \sum_k n_{yk}\ \overline{Y_k}^2 + n_u\, U_{ref}^2 - 2 U_{ref} \sum_k n_{yk}\, \overline{Y_k} \right) \qquad (12)$$

Equation (12) for the standard deviation of the merged dataset simplifies if the original datasets are adjusted to exactly the same reference value ref ($\overline{Y_k} = ref$) and the merged value U_{ref} is also equal to that value, as the sum of the last 3 terms in Eq. (10) (with Y_k replacing y_k) then reduces to $n_u\, ref^2 + n_u\, ref^2 - 2\, n_u\, ref^2$, which is zero. In this case, one obtains

$$s_u^2 = \frac{1}{(n_u - 1)} \left(\sum_k (n_{yk} - 1) s_{yk}^2 \right) \qquad (13)$$

However, in general, one should use equation (12) for the standard deviation of the merged dataset, given the adjusted datasets $\overline{Y_k}$ and the merged (or reference) value U_{ref}. Also, we often use a merged value equal to the average of the original data (over a given overlap period), so that

$$U_{ref} = \frac{1}{n_y} \sum_k \overline{y_k} \qquad (14)$$

where n_y is the total number of datasets (y_k), as opposed to having the merged value place more weight on the larger datasets (e.g., for emission-type measurements versus occultation-type), in which case one would consider using $U_{ref} = \frac{1}{n_u} \sum_k n_{yk} \overline{y_k}$. For ozone, we use a particular dataset (SAGE II ozone) as the primary reference, but equation (12) can be used to obtain the standard deviation for the merged dataset (about the SAGE II reference) in that case also. While it is useful to have the formalism above for obtaining the merged dataset standard deviation σ_u, we often find significant differences between the standard deviations of various datasets, so that this effect will have the greatest influence on the results, as opposed to the impact of the last 3 terms in the summation (in (12)). Finally, it is easy to test equation (12) (and we have done so) by using synthetic series and calculating the standard deviation of the combined set. In reality, the standard deviations of the time series monthly mean values are typically larger for MLS than for ACE-FTS, mainly because of the more complete sampling of variability from the daily global measurements acquired by MLS. Sample plots for standard deviations and standard errors in the case of HCl are shown in Fig. A1. As expected, merged standard deviations follow the standard deviations from HALOE HCl before Aug. 2004 and those from MLS HCl after this time. However, the merged standard errors for the MLS time period follow the smaller MLS standard errors, because these values vary inversely with the square root of the number of values sampled, and are therefore made smaller by the significantly larger daily and monthly MLS sampling rate and coverage.

A.3. Procedural merging details for GOZCARDS HCl, H_2O, and O_3

A.3.1. HCl

- The vertical data range for valid HCl merged values is between 0.46 hPa and 147 hPa (inclusive), as a result of data sparseness or data quality issues outside these ranges.
- At 147 hPa, no merged HCl values exist for latitude bins from 35°S to 35°N inclusive, because of unrealistically large Aura MLS HCl values in this region; also, there is not enough data at this level to provide a meaningful product from HALOE and ACE-FTS data alone.
- Because of occasional small negative merged values during southern hemisphere polar winter, we did not apply HCl data offsets in the lower stratosphere for the 65°S through 85°S bins from June through September and for pressures larger than or equal to 15 hPa. For vertical continuity purposes, we applied this method to all lower stratospheric pressure levels, although the small negative merged values only occurred in a small fraction of cases and the impact on the merged values is not large. Seasonal variations in other bins are milder and did not lead to such an Antarctic winter issue; also, this issue did not affect other species.
- As Aura MLS and ACE-FTS data exist in the 85°N and 85°S bins, but there are no HALOE measurements, we could not use our standard merging procedure there. We simply extended the offsets from the adjacent bins (at 75°N and 75°S) to these two bins to obtain a merged record after 2004 that exhibits continuity versus latitude.
- At 100 hPa, we used HCl offsets from the 5°S bin for the 5°N bin, as there was insufficient data from the three combined datasets in the latter bin to calculate meaningful offsets and merge the datasets. This procedure seems reasonable, given that the time series in these two adjacent tropical latitude bins (during years outside the 2004/2005 overlap period) look continuous and stable enough to justify identical adjustments in both bins and to avoid a data gap in the merged series at 5°N, although this does imply somewhat larger error bars at 5°N.

A.3.2. H_2O

- The vertical data range for valid H_2O merged values is between 0.01 hPa and 147 hPa (inclusive). Some H_2O data exist at 147 hPa for low latitudes, but more careful work would be needed to extend the merged data globally in such a region.

- Users should keep in mind the PMC-related caveats mentioned in Sect. 4 for summer at high latitudes in the upper mesosphere, prior to the end of the HALOE dataset (Nov. 2005).
- As for HCl, we could not use our standard merging procedure at the two most poleward latitude bins; we simply extended the offsets from the adjacent bins (at 75°N and 75°S) to these polar bins to obtain a merged record after 2004 that exhibits continuity versus latitude.
- Also as for HCl, at 100 hPa, we used H_2O offsets from the 5°S bin for the 5°N bin, as there was insufficient data from the combined datasets in the latter bin to calculate meaningful offsets and merge the datasets. This procedure avoids a data gap in the merged series at 5°N.

A.3.3. O_3

Screening of SAGE O_3 data

For SAGE I O_3, the main uncertainty is aerosol interference, especially below 15 to 20 km. All SAGE I values below (in altitude) where the aerosol extinction at 1.0 μm reaches a value larger than 1.0×10^{-3} km^{-1} are removed from the analysis (L. W. Thomason, personal communication).

For SAGE II ozone, the screening steps are based on Wang et al (2002) as follows:

- We removed the entire ozone profile when any reported error bar value exceeded 10% between 30 and 50 km, in order to filter out outliers affected by "short events" (Wang et al., 2002), which mainly occurred between mid-1993 and mid-1994, when SAGE II had a battery problem. In order to preserve power, sunset measurements were started later than normal while sunrise measurements were ended earlier. These "short events" had fewer extraterrestrial solar irradiance measurements for calibration and normalization.
- We used aerosol extinctions and extinction ratios to remove data affected by clouds, and aerosols from the June 1991 Mt. Pinatubo eruption. O_3 data were removed when the aerosol extinction at 0.525 μm exceeded 6×10^{-3} km^{-1}, thus removing data affected by this eruption for months and even years, in the lower stratosphere. For cases with extinctions less than 6×10^{-3} km^{-1} but greater than 1×10^{-3} km^{-1}, and extinction ratios (0.525 μm /1.02 μm) ≤ 1.4, the corresponding data were removed for additional filtering. Although more stringent criteria could be used to remove a few more outliers, this would also remove many more "good" ozone data that are not affected by aerosol/cloud. Fortunately, any artifacts from these few unfiltered data values are greatly reduced after binning the data into monthly zonal means.

- We removed anomalously low O_3 values resulting from very small SAGE II transmittances; O_3 error values in these cases were capped at 300% by the algorithm. Such low O_3 values (sometimes low by 2-3 orders of magnitude) generally occur close to the tropopause and in the troposphere, and can be identified by using this 300% error flag (Wang et al., 2002).
- It was found that SAGE II ozone data could be affected during high sun-orbit beta angle conditions (Wang et al., 1996). SAGE II profiles immediately following fully sunlit orbits with absolute values of beta greater than 40° are eliminated from monthly zonal means.

Other merging details for O_3

- SAGE I monthly mean source data are used for the merged dataset in the tropical bins (25°S to 25°N) from 1 through 68 hPa only and, at higher latitudes, from 1 through 100 hPa only.
- The vertical range for valid O_3 merged values is between 0.2 hPa and 215 hPa (inclusive), with the lower altitude bound varying with latitude. The merged product at 147 and 215 hPa has valid data only for the 35° to 85° latitude bins. Indeed, we limited merged data mostly to stratospheric values (larger than ~ 0.1 ppmv); the upper troposphere is more of a challenge for such a merging activity, given smaller abundances, more challenging measurements, and a larger impact from different instrument resolutions. The upper range limit was chosen to enable studies of the upper stratosphere and lower mesosphere, even if this is a region where diurnal ozone change occurs; arguments we have presented (se main text) suggest that the GOZCARDS merged ozone time series variations should not be subject to a large impact from diurnal variations, although high altitude regions should still be treated with caution.
- We omitted the use of UARS MLS at 100 hPa for low latitudes (from 25°S to 25°N), as these monthly values are biased quite high and also exhibit too large a seasonal cycle amplitude, in comparison to HALOE and SAGE II data; this appears to relate to a UARS MLS artifact.
- Since there is no (monthly) overlap between SAGE II and HALOE versus UARS MLS or Aura MLS in the 85°N and 85°S latitude bins, the same offsets as for 75°N and 75°S (respectively) are applied to the datasets at these two extreme latitude bins, in order to minimize latitudinal discontinuities in the merged data record.

- Because of discontinuities that appeared in merged O_3 at high latitudes above the stratopause, particularly in the 75°S bin, we flagged merged values for 75° and 85° (N and S) as bad, for pressures less than 1 hPa. This issue could be the result of a few bad data points or not enough data overlap. To minimize artifacts, we left the resolution of this issue for future investigations; also, the reduced amount of occultation data at these high latitudes makes the usefulness of a merged product with poorly sampled seasonal changes somewhat marginal (for certain years at least, the number of monthly values drops significantly at high latitudes).

Table 1. Characteristics of instrument datasets used to create GOZCARDS ESDRs (version ev1.01).

Instrument and Data Versions	Platform	Type of measurement	Time period (GOZCARDS source files)	Vertical Resolution (km)	Retrieved quantity and stratospheric vertical grid spacing
SAGE I V5.9_rev O_3	AEM-2	Solar occultation VIS/UV and near-IR	Feb. 1979 - Nov. 1981	1	Density on altitude grid 1 km spacing
SAGE II V6.2 O3	ERBS	Solar occultation VIS/UV and near-IR	Oct. 1984 - Aug. 2005	0.5 - 1	Density on altitude grid 0.5 km spacing
HALOE V19	UARS	Solar occultation mid-IR	Oct. 1991 - Nov. 2005	2.5	Volume Mixing Ratio on pressure grid with 30 levels per decade (LPD) change in p
MLS V5 O_3 V6 H_2O	UARS	Limb emission microwave / sub-mm	Oct. 1991 - June 1997 (May 1993 end for strat. H_2O)	**H_2O** 3 - 4 (strat.) 5 - 12 (mes.)	Volume Mixing Ratio on pressure grid with 6 LPD in stratosphere
				O_3 3.5 - 5 (strat.) 5 - 8 (mes.)	6 LPD in stratosphere
ACE-FTS V2.2 (V2.2 update for O_3)	SCISAT	Solar occultation mid-IR	Mar. 2004 through Sep. 2010 (2009 for O_3)	3 - 4	Volume Mixing Ratio on 1 km grid spacing (height and p provided)
MLS V3.3 V2.2 O_3	Aura	Limb emission microwave / sub-mm	Aug. 2004 through 2012	**HCl** 3 - 5	Volume Mixing Ratio on pressure grid with 6 LPD
				H_2O 3 - 4 (p > 0.1 hPa) 5 - 9 (0.1-0.01 hPa)	12 LPD
				O_3 3	6 LPD

2027 **Table 2.** Products and instrument source data making up the available GOZCARDS data records.

Merged Products and pressure range	**Source Datasets** (and years used)
HCl 147 – 0.5 hPa	HALOE (1991-2005), ACE-FTS (2004-2010), Aura MLS (2004 onward)
H_2O 147 – 0.01 hPa	HALOE (1991-2005), UARS MLS (1991-1993), ACE-FTS (2004-2010), Aura MLS (2004 onward)
O_3 215 – 0.2 hPa	SAGE I (1979-1981), SAGE II (1984-2005), HALOE (1991-2005), UARS MLS (1991-1997), ACE-FTS (2004-2009), Aura MLS (2004 onward)
HNO_3 215 – 1 hPa	ACE-FTS (2004-2010), Aura MLS (2004 onward)
N_2O 100 – 0.5 hPa	ACE-FTS (2004-2010), Aura MLS (2004 onward) \
Temperature 1000 – 0.015 hPa	GMAO MERRA (1979 onward)

2028

2029
2030

Fig. 1. Merging procedure illustration for HCl. Top left panel shows the HCl monthly mean source data during the overlap period (Aug. 2004 - Nov. 2005) for HALOE, ACE-FTS, and Aura MLS. Top right panel illustrates step 1 in the merging procedure, with the temporary merged data values (orange) resulting from the adjustment of ACE-FTS and Aura MLS values to the reference indicated by the black dashed line. Middle left panel shows the result of step 2, namely the merged values arising from merging HALOE data with the temporary merged values from step 1; the black dashed line is the new average reference value, obtained from a 2/3 and 1/3 weighting of the dashed orange and dashed blue line values, respectively (see text). Middle right panel shows all the source data and the final merged values during the overlap period. Bottom panel shows the source and merged time series from 1991 through 2012 after the calculated additive offsets are applied to the whole source datasets, which are then merged (averaged) together wherever overlap between instruments exists.

Fig. 2. Offsets applied to the HCl source datasets (top panels for HALOE, middle panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Nov. 2005) used here to compute the average offsets.

Fig. 3. Example of HCl time series analyses for 50°N-60°N and 32 hPa. (a) HCl monthly mean source data from ACE-FTS and Aura MLS; the MLS dots are filled when there is time overlap with ACE-FTS, and open if no such overlap exists. Simple linear fits are shown as colored lines for ACE-FTS and for Aura MLS (orange line for all red dots and red line for filled red dots only). Correlation coefficient values (R values) for the two time series are provided in the title. (b) Deseasonalized anomalies for both ACE-FTS and Aura MLS, with corresponding linear fits (and R values). (c) Difference of deseasonalized anomalies (ACE-FTS minus Aura MLS), with linear fit.

Fig. 4. Latitude/pressure contours of time series diagnostics obtained from analyses illustrated in Fig. 3 for HCl from Aura MLS and ACE-FTS. Top panel: Correlation coefficient for the deseasonalized time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the error in this slope.

Fig. 5. Illustration of GOZCARDS HCl monthly averages with systematic error estimates (shown as grey shading) at 46 hPa for 30°S-40°S; see text for the meaning of this shaded region. The source data from HALOE, Aura MLS, and ACE-FTS are shown in different colors (see legend), along with the merged values.

Fig. 6. Systematic error estimates for GOZCARDS HCl. One error (left panels) is relevant for values lower than (below) the merged values, and one (right panels) for values larger than the merged values; the top panels give the error estimates in ppbv, and the bottom panel errors are expressed as percent of the average merged values over the relevant time periods (see text). These error bars provide a range within which 95% of the source data values lie.

Fig. 7. Time series of the GOZCARDS monthly-averaged merged HCl abundance for 3 different latitude bin averages (see color legend in panel (a)) for (a) 0.7 hPa, (b) 10 hPa, (c) 32 hPa, and (d) 68 hPa.

Fig. 8. The average rate of change (percent per year) for HCl as a function of pressure for different latitude bin averages (see legend) for time periods corresponding to the appropriate GOZCARDS HCl values (see text) in the upper stratosphere (Jan. 1997 - Sep. 2010) and lower stratosphere (Jan. 1997 - Dec. 2012). Deseasonalized monthly data were used to obtain a long-term trend for these time periods; two-sigma error bars are shown.

Fig. 9. Same as Fig. 8, but for the lower stratosphere only and using different time periods to illustrate shorter-term changes in this region. Average rates of change in HCl are given for (a) 2003 through 2012, a decade exhibiting significant differences between northern and southern hemispheric change, (b) the 6-year period 2006 through 2011, when the largest changes occurred, and (c) the most recent 6-year period 2008 through 2013, when a significant decrease in the variability took place.

Fig. 10. Rates of change for GOZCARDS HCl (connected open circles) are given as a function of latitude in 10° latitude bins for sliding 6-year periods centered on Jan. 1 of each year (e.g., the 1998 point is an average for data from 1995 through 2000, and the 2011 point is for data from 2008 through 2013). (a) is for changes in upper stratospheric HCl at 0.7 hPa and (b) is for the change in the integrated HCl column between 68 hPa and 10 hPa. The two additional curves in (a) represent the rates of change in the estimated surface total chlorine from NOAA data (green is for a 6-year time shift, and purple for a 7-year time shift, to account for transport time to the upper stratosphere); see text for more details. Panels (c) and (d) are similar to (a) and (b), respectively, in terms of the pressure levels used, but the rates of change are averaged over all latitude bins covering 50°S to 50°N for two sets of sliding time periods (black is for 6-year periods, red is for 8-year periods). As in (a), surface total chlorine variations are also displayed in panel (c). Error bars indicate twice the standard errors in the means.

Fig. 11. Offsets applied to the H_2O source datasets as a function of latitude and pressure, similar to Fig. 2 for HCl.

Fig. 12. Latitude/pressure contours of time series diagnostics for H_2O from Aura MLS and ACE-FTS; this is similar to Fig. 4 for HCl.

Fig. 13. A depiction of the "tape recorder" evolution for tropical water vapor abundances from 147 to 10 hPa for October 1991 through December 2013. This plot was produced from GOZCARDS merged H_2O time series anomalies (differences from the long-term means) for the average of the 4 tropical bins covering 20°S to 20°N.

Fig. 14. Systematic error estimates for GOZCARDS H_2O (similar to Fig. 6 for HCl).

Fig. 15. Variations in stratospheric water vapor from the GOZCARDS H_2O merged data records (1992 through 2013) averaged from (a) 60°S to 60°N and (b) 20°S to 20°N. Monthly average values and annual averages are shown by thin and thick lines (connecting similarly-colored dots), respectively, for the pressure levels indicated in the plot legend.

Fig. 16. Stratospheric water vapor variability on decadal timescales for 1992 through 2013 (thick lines) for tropical (20°S-20°N in black) and mid-latitude (20°N-60°N in red and 20°S-60°S in blue) zonal means, based on the GOZCARDS merged H_2O data record. The variability is expressed here as the difference between maximum and minimum annual average abundances, from 100 to 1 hPa, in ppmv (left panel) and percent (right panel). The 22-year period is broken up into two 11-year periods to illustrate how the variability changes from the 1st period (dashed lines) to the 2nd period (thin solid lines).

Fig. 17. (a) Variations in upper mesospheric (0.01 hPa) water vapor mixing ratios averaged from 60°S to 60°N for Oct. 1991 through Dec. 2013, based on the GOZCARDS merged H_2O data records. Monthly average values and annual averages are shown by connected brown dots and connected black dots, respectively. (b) GOZCARDS merged H_2O annual averages (connected filled symbols) from 60°S to 60°N for 1992 through 2013 at pressure levels between 0.1 and 0.01 hPa. A time series of annually-averaged Lyman α solar flux values (open circles), scaled to arbitrary units, is also displayed (see text).

Fig. 18. Time series of monthly zonal mean O_3 for 10°S - 20°S between 1 hPa and 6.8 hPa (with pressure values given by ″pre″) from SAGE I, SAGE II, HALOE, UARS MLS, Aura MLS, and ACE-FTS, all color-coded following the legend in top left panel.

Fig. 19. Schematic diagram describing the creation of the merged GOZCARDS monthly zonal mean ozone data record from various satellite datasets. Instruments represented in red inside the boxes are used as a reference. Instruments whose measurements have already been adjusted to a reference are indicated with a "*" superscript. AMLS refers to Aura MLS and UMLS to UARS MLS. See text for more details.

Fig. 20. Offsets applied to the O_3 source datasets, similar to Fig. 2 for HCl.

Fig. 21. Latitude/pressure contours of time series diagnostics for O_3 from Aura MLS and ACE-FTS; this is similar to Fig. 4 for HCl. The correlation coefficients (R values) and slope trend diagnostics are provided for HALOE versus SAGE II in the top two panels (for 1993-1999 as the trend issue for converted SAGE II data occurs after mid-2000 and to avoid Pinatubo-related data gaps before 1993) and for ACE-FTS versus Aura MLS in the bottom two panels (for 2005-2009).

Fig. 22. Systematic error estimates for GOZCARDS O_3 (similar to Fig. 6 for HCl).

Fig. 23. Column ozone values (DU) from ZC12 (in red) for 60°S-60°N and from GOZCARDS averages in different latitude bins (see legend). The connected dots are for GOZCARDS column ozone densities above 68 hPa from 1979 through 2013. The lines with no symbols and the connected open dots are also for GOZCARDS columns, but for values above 100 hPa and above 215 hPa, respectively, for 1985 through 2013; there are no blue open dots because of the lack of GOZCARDS merged ozone data in the tropics for pressures larger than 100 hPa, during 1985-2013.

Fig. 24. Near-global (60°S to 60°N) results for average column ozone (total and stratospheric, from *Ziemke and Chandra,* 2012) compared to GOZCARDS O_3 columns above 68 hPa. Stratospheric columns are offset in order to more easily compare relative variations versus time; the black dots and red crosses are referenced to the 1980 total column values, while the cyan curves are referenced to 2007 to better illustrate the fits in the later years.

Fig. 25. Change in column ozone after removal of solar cycle signal and with 3-year smoothing applied; (a) gives relative changes in DU, and (b) shows percent changes relative to the average GOZCARDS stratospheric columns in 1979. Black symbols are from the GOZCARDS column values above 68 hPa, averaged over 60°S-60°N. Red symbols, for comparison, are from the ZC12 data over the same latitude range (after removal of the fitted solar cycle signal).

Fig. 26. Time evolution (Aug. 2004 through 2012) versus latitude of GOZCARDS merged N_2O (ppbv) at (a) 6.8 hPa and (b) 100 hPa.

Fig. 27. Sample results display the time evolution of satellite-retrieved HNO_3 (ppbv) for two different periods, 1992-1997 in (a) and (c) versus 2004-2013 in (b) and (d). Panels (a) and (b) are contour plots at 46 hPa from UARS MLS global data and the merged GOZCARDS global data after 2004, respectively; (c) and (d) show time series at 32 hPa and for the 40°N-50°N latitude bin, with (a) from UARS MLS data, and (d) from ACE-FTS, Aura MLS, and the merged combination (between the two source data sets).

Fig. A1. Illustration of the standard deviations (in (a)) and standard errors (in (b)) for monthly mean GOZCARDS HCl (source and merged records) at 46 hPa for 30°S-40°S. Source data from HALOE, Aura MLS, and ACE-FTS are given by the filled colored dots (see legend); each standard deviation is simply obtained from the range of values measured during the month. The large open brown circles give standard deviations for the merged HCl product; this Appendix provides the formulae to calculate these quantities.

Fig. S1: Illustration of the latitudinal dependence of the HCl offsets for HALOE, ACE-FTS, and Aura MLS at two pressure levels (top panel for 0.46 hPa, bottom panel for 46 hPa). Error bars represent twice the standard error in the derived offsets (based on variability during the overlapping period). Larger standard error values indicate that there were either fewer points of overlap or larger offset variability (standard deviations); we found that both of these factors contribute.

Fig. S2: Latitude/pressure contours of the fitted mean annual amplitudes (ppbv) from HCl time series for HALOE, ACE-FTS, and Aura MLS, based on their respective measurement periods (see text).

Fig. S3: Time evolution (Oct. 1991 through 2013) versus latitude of GOZCARDS merged HCl (ppbv) at 46 hPa.

Fig. S4: HALOE sunrise measurements of H_2O versus the 3.46 μm extinction coefficient for 1992, 1993, and 1999 at 22 hPa. The green vertical line represents the aerosol extinction value ($5x10^{-4}$ km^{-1}) used to screen anomalous HALOE H_2O values. It is apparent that anomalously low H_2O values occurred in 1992 when the 3.46 μm aerosol extinction exceeded about $5x10^{-4}$ km^{-1}. These artifacts were confined to 1991 and 1992; for these years, and for pressure levels at and below 22 hPa, the corresponding H_2O data values were excluded. This screening method eliminates about 10% of the global (lower stratospheric) measurements in 1992.

Fig. S5: Merging procedure illustration for H_2O at 5°N and 22hPa. This is similar to Fig. 2 (for HCl), but an additional step is illustrated for the end of this procedure, whereby stratospheric H_2O data from UARS MLS are adjusted to the early portion of the merged time series that was obtained after the 2nd step; this adds more coverage (more brown dots in the bottom panel for 1991-1993).

Fig. S6: Latitude/pressure contours of the fitted mean annual amplitudes (ppmv) from H_2O time series for HALOE, ACE-FTS, and Aura MLS, based on their respective measurement periods.

Fig. S7: Time evolution (Oct. 1991 through 2013) versus latitude of GOZCARDS merged H_2O (ppmv) at 3.2 hPa (top panel) and 68 hPa (bottom panel).

Fig. S8: Monthly zonal mean ozone differences (%) between SAGE II and (a) HALOE, (b) UARS MLS (UMLS for short), (c) Aura MLS (AMLS for short), and (d) ACE-FTS during their respective overlap periods. Differences are expressed (in percent) as 100 x [(SAGE II - Other) / (Other)]. Shaded areas indicate negative values.

Fig. S9: Monthly zonal mean temperature differences between NCEP (used by SAGE II) and HALOE temperatures relative to MERRA for 10°S - 20°S between 1 and 6.8 hPa, per color-coding indicated in bottom left panel; ‘‘pre‘‘ represents the pressure value. From 1 to 2.1 hPa, differences between NCEP and MERRA are generally within ± 4K before mid-2000. After that time, NCEP temperatures show a sharp increase and are systematically higher than MERRA values by 5 to 10K. However, this divergence and trend are not seen in HALOE temperatures. NCEP temperatures between 3.2 and 6.8 hPa are smaller than MERRA after mid-2000; negative trends (versus MERRA) also occur in the HALOE data at these levels.

Fig. S10: Relative trends (K/decade) in zonal mean temperature differences for NCEP – MERRA and HALOE – MERRA (color-coded as in Fig. S9) in the upper stratosphere. NCEP temperatures show positive trends versus MERRA of ~2-5 K/decade between 2.1 and 1 hPa for all latitudes. However, HALOE temperatures show no significant trends versus MERRA, except at 1.5 hPa in the southern hemisphere. For pressures between 3.2 and 6.8 hPa, the temperature analyses are not conclusive; although NCEP values show negative trends of ~2-3 K/decade versus MERRA, they agree with HALOE.

Fig. S11: Mean differences and standard deviations (horizontal bars) between SAGE II and Aura MLS ozone in three different latitude bins: 20°S to 60°S (left panel), 20°S to 20°N (middle panel), and 20°N to 60°N (right panel). Results based on monthly zonal mean and coincident profiles (see text for coincidence criteria) during overlap periods are shown in red and blue, respectively. To choose collocated profiles, coincidence criteria of ±1° in latitude and ±8° in longitude were used; the time difference criterion was chosen as 12 hours, but only nighttime measurements from Aura MLS were used.

Fig. S12: Latitude/pressure contours of the fitted mean annual amplitudes (ppmv) from O_3 time series for SAGE II, HALOE, ACE-FTS, and Aura MLS, based on their respective measurement periods.

Fig. S13: Illustration of the time evolution of the GOZCARDS merged O_3 data field versus latitude at 68 hPa (top panel) and versus pressure for the 40°N-50°N latitude bin (bottom panel).

Fig. S14: Offsets applied to the N_2O source datasets (top panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Sep. 2010) used here to compute the average offsets.

Fig. S15: Latitude/pressure contours of time series diagnostics derived from Aura MLS and ACE-FTS N_2O data comparisons (and obtained from analyses similar to those illustrated in Fig. 6 for HCl). Top panel: Correlation coefficient for the deseasonalized time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the error in this slope.

Fig. S16: Offsets applied to the HNO_3 source datasets (top panels for ACE-FTS, bottom panels for Aura MLS) as a function of latitude and pressure. The left column gives offsets in ppbv and the right column provides offsets as a percent of the zonal average merged mixing ratios during the overlap period (Aug. 2004 – Sep. 2010) used here to compute the average offsets.

Fig. S17: Latitude/pressure contours of time series diagnostics derived from Aura MLS and ACE-FTS HNO_3 data comparisons (and obtained from analyses similar to those illustrated in Fig. 6 for HCl). Top panel: Correlation coefficient for the deseasonalized time series. Bottom panel: Ratio of the slope of the difference between deseasonalized series over the error in this slope.