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**Characteristics of Monsoon Inversions over Arabian Sea observed by Satellite Sounder and
Reanalysis data sets**

Sanjeev Dwivedi¹, M. S. Narayanan¹, M. Venkat Ratnam^{2*} and D. Narayana Rao¹

¹Department of Physics, SRM University, Kattankulathur, Chennai - 603 203, India.

²National Atmospheric Research Laboratory (NARL), Gadanki, Tirupati- 517 502, India.

* vrtnam@narl.gov.in ; Phone: +91-8585-272123; Fax: +91-8585-272018

19 **Abstract**

20 Monsoon inversions (MIs) over Arabian Sea (AS) are an important characteristic associated
21 with the monsoon activity over Indian region during summer monsoon season. In the present study,
22 we have used five years (2009 - 2013) data of temperature and water vapor profiles obtained from
23 satellite sounder instrument, Infrared Atmospheric Sounding Interferometer (IASI) onboard MetOp
24 satellite, besides ERA - Interim data, to study their characteristics. The lower atmospheric data over
25 the AS have been examined first to identify the areas where monsoon inversions are predominant
26 and occur with higher strength. Based on this information, a detailed study has been made to
27 investigate their characteristics separately in eastern AS (EAS) and western AS (WAS) to examine
28 their contrasting features. The initiation and dissipation times of MI, their percentage occurrence,
29 strength etc., has been examined using the huge data base. The relation with monsoon activity
30 (rainfall) over Indian region during normal and poor monsoon years is also studied. WAS ΔT values
31 are ~ 2 K less than those over the EAS, ΔT being temperature difference between 950 and 850 hPa.
32 A much larger contrast between WAS and EAS in ΔT is noticed in ERA-Interim dataset Vis a Vis
33 those observed by satellites. The possibility of detecting MI from another parameter, Refractivity N,
34 obtained directly from another satellite constellation of GPS RO (COSMIC), has also been
35 examined. MI detected from IASI and Atmospheric InfraRed Sounder (AIRS) sounder onboard
36 NOAA satellite have been compared to see how far the two data sets can be combined to study the
37 MI characteristics. We suggest MI could also be included as one of the semi-permanent features of
38 southwest monsoon along with the presently accepted six parameters.

39

40 *Keywords:* Monsoon inversion, Arabian sea, lower atmospheric temperature, satellite sounders,
41 IASI, ERA

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44 **1. Introduction**

45 The Monsoon Inversion (MI) is one of the criteria providing a stability condition over the
46 western Arabian Sea (AS), extending sometimes through to the west coast of India. The MI controls
47 the mid tropospheric moisture content during the different phases of the monsoon. This shallow layer
48 of low level inversion will act as a barrier in uplifting of the moisture, and could act like a wave –
49 guide for transport of water vapour to the mainland. The fluctuation of the rainfall over the west
50 coast of India is more closely related to changes in monsoon circulation over the AS (Das, 2002).
51 The AS is located at the north head of the Indian Ocean. During the monsoon season, Indian rainfall
52 is fully dependent on the physical processes occurring over AS like SST, Somali Low Level Jet and
53 near by it Arabia desert is there which is putting more effect on MI. Thus, MI has been known to be
54 intimately associated with the activity of the Indian southwest monsoon and have a close link with
55 active and break spells (Narayanan and Rao, 2004).

56 MIs were first detected in 1964 during International Indian Ocean Expedition (IIOE) from
57 ship radiosonde data by Colon (1964) and Ramage (1966). Subsequently from satellite derived
58 temperature and humidity data, this feature was detected by Narayanan and Rao (1981). They
59 detected MI despite the coarse vertical resolution (~ 2 km) of the TIROS – N satellite temperature
60 sounding instruments (Thomas,1980) of 1970 – 80's compared to the vertical extent (about 1 to 1.5
61 km) of the phenomena itself. They used a simple differencing technique by finding the difference,
62 ΔT , of sea skin temperature and 1000 to 850 hPa mean layer temperature (MLT) from the satellite
63 sounding data. By adopting this differencing procedure, they assumed that most of the systematic
64 errors/limitations of retrieval methods and vertical resolution of satellite soundings may be getting
65 significantly minimized. Furthermore, the spatial and temporal nature of MIs is quite large compared
66 to normal boundary layer inversions observed over land and other oceans.

67 Using data of about 150 ship radiosonde and aircraft dropsonde profiles and concurrent
68 TIROS – N satellite sounder data of MONsoon EXperiment (MONEX) conducted in 1979, they

69 showed that regions with $\Delta T \leq 2$ K in satellite derived atmospheric temperatures are associated with
70 AS MI. Study of these MIs over the western AS was one of the three major objectives of MONEX /
71 FGGE -1979 (WMO, 1976). These are seen to be much stronger (temperature departures from
72 normal profiles in some cases being as high as ~ 6 K in the lower 1 - 2 km height region) in contrast
73 to the inversions observed over land or associated with trade wind inversions ($\sim 1 - 2$ K).

74 MIs are characterized by both a vertical temperature increase in the altitude region from 0.5
75 km (in some cases even from surface) to ~ 2 km and with a sharp fall in relative humidity (RH)
76 above this altitude region. Some of the observed features of MIs reported from the limited
77 observations to date (Colon, 1964; Ramage, 1966; Narayanan and Rao, 1981; 1989) are: (i) strength
78 decreases and base increases as one moves from the west to east AS, (ii) oscillation of its lateral
79 boundary from west to east with the activity of monsoon and (iii) associated oscillation of mid
80 tropospheric water vapor content from east to west, i.e. in the opposite sense to the boundary of
81 temperature inversion. The two primary causes proposed (Colon, 1964) for formation/maintenance
82 of monsoon inversion are: (a) hot air advection from Arabia (~ 700 hPa) riding over cool maritime air
83 (at levels below ~ 800 hPa) from south Indian Ocean and (b) subsidence over western AS associated
84 with monsoon convection over main land. This large scale subsidence had played a major role in the
85 maintenance of MI during the prolonged weak monsoon of 2002 (Narayanan et al., 2004).

86 However, not much attention was paid to the study of MI due to paucity of freely available
87 data over this region. The spatial density of TIROS – N satellite data available to the global, research
88 community in 1979 was just a single temperature – humidity profile a day in a latitude – longitude
89 grid box of $2.5^\circ \times 2.5^\circ$ (Kidder et al., 1995). Narayanan and Rao (1981) had to adopt temporally a
90 pentad and spatially a $5^\circ \times 5^\circ$ average to detect statistically significant results from the meager data
91 available then. Since 2008, the density of temperature and humidity profiles from polar orbiting
92 satellites is nearly two orders of magnitude higher (about one vertical profile every 50×50 km, twice
93 each day and from two satellites) besides with a much better vertical and spectral resolution. Thus, it

94 has become possible now to study MI phenomena in greater detail. However, no in-situ data after the
95 1979 experiment are available in this region.

96 In the present study, we have used the high resolution and better accuracy temperature and
97 humidity profiles data obtained from Infrared Atmospheric Sounding Interferometer (IASI) onboard
98 MetOp satellite. These data have higher vertical resolution, i.e., ~ 400 m below 700 hPa, which is
99 much better than those of TIROS – N of MONEX 1979 period. Further, ERA-Interim data have been
100 used to compare the MI features seen in them with those from the satellite data. For explaining the
101 relative contribution of subsidence and convection on MI, where only wind observations are
102 required, ERA-interim reanalysis data have been used. The temperature - humidity profile data are
103 also available from NOAA – Atmospheric InfraRed Sounder (AIRS) instrument since 2002, all of
104 which have also been analysed in the same way as the IASI data. However, we have not presented
105 those results here, because of some inconsistencies (i.e. sometimes ERA – interim data shows MI but
106 AIRS has different features like no MI present, profile to profile match between AIRS and ERA-
107 interim datasets are not seen i.e. inversion type changes or level of inversion changes) observed
108 between the IASI and AIRS data in studying the MI features. Thus, we have confined the present
109 study to data only from one instrument, viz., IASI, which had been reported to be performing better
110 (Smith et al., 2015). This is expected to also ensure that the results of temporal and spatial gradients
111 of ΔT presented here (featuring MI) will be mutually consistent – even if the absolute values of
112 temperature/humidity may be having some errors. We have, however, included one section
113 describing the discrepancies between the results of these two instruments for studying the MI
114 features. We have also shown to a limited extent the potential of the GPS RO measured ‘refractivity’
115 profiles in delineating inversion regions. For this we have also used the MONEX in-situ temperature
116 – humidity profiles of 1979.

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119 **2. Data**

120 As mentioned earlier, data from a variety of instruments have been used in this study – viz
121 from IASI satellite instrument, ERA-Interim reanalysis data and, in-situ dropsondes/ radiosondes
122 data obtained during MONEX – 1979. Limited AIRS sounder data and GPS RO data have also been
123 presented for comparison purposes. A short description of each of these data are given in the
124 following sub-sections and also summarized in Table 1.

125 **2.1. IASI observations**

126 The IASI instrument (Clerbaux et al., 2007; 2009) measures the profiles of temperature
127 profiles in the troposphere and lower stratosphere with a high accuracy (~1K root mean square) at a
128 vertical resolution of 1 km in the lower troposphere), as well as humidity profiles in the troposphere
129 (10–15% accuracy with a 1–2 km vertical resolution) primarily for numerical weather prediction
130 (Schlüssel et al., 2005). IASI is a thermal infrared nadir-looking Fourier transform spectrometer
131 which measures the Earth’s surface and the atmospheric radiation over a spectral range of 645–2760
132 cm^{-1} with a 0.5 cm^{-1} spectral resolution. The IASI field of view is a matrix of $2^\circ \times 2^\circ$ circular pixels,
133 each with a diameter footprint of 12 km at nadir. It measures on an average at each location on the
134 Earth’s surface twice a day (at 09:30 and 21:30 hr local time), every 50 km at nadir, with an
135 excellent horizontal coverage due to its polar orbit and its capability to scan across track over a swath
136 width of 2200 km. More details about retrieval and validation are presented in Kwon et al. (2012).
137 The support products, which we have used, are available at 100 pressure levels at 50 x 50 km
138 horizontal grid spacing, but we restrict the data from surface to 600 hPa only.

139 **2.2. Dropsonde / Radiosonde measurements MONEX (1979)**

140 For the in-situ ground truth comparisons over AS between the longitudes 55° - 75° E we also
141 make use of the aircraft dropsondes and ship radiosonde observations obtained during MONEX
142 1979. MONEX was conducted during May - July 1979 and there were 416 radiosondes and 412
143 dropsondes measurements over AS. It may be noted that after the MONEX campaign in 1979, no

144 campaign has been organized to get in-situ data over western or central AS. During the Indian
145 ARMEX programme (2002), however, some in-situ data were available but only in the far eastern
146 AS (east of 70°) near the coast of India. Table 2 summarizes the comparison of in-situ observations
147 with satellite data of 1979 by Narayanan and Rao (1981). This information on ΔT criterion has been
148 used as the basis in the present study.

149 **2.3. ERA-Interim data**

150 The European Centre for Medium Range Weather Forecasts (ECMWF)-Interim is one of
151 most advanced in operational use for diagnosing the global atmosphere with an accuracy that is less
152 than what is theoretically possible (Simmons and Hollingsworth, 2002; Simmons et al., 2007). The
153 selected variables are specific humidity along with the temperature on different pressure levels. The
154 atmospheric data are available at $0.125^\circ \times 0.125^\circ$ latitude and longitude grids on 37 pressure levels
155 from 1000 to 1 hPa; however, we have used data of 14 pressure levels from 1000 to 600 hPa for the
156 period of 2009 to 2013 for the present study. Vertical as well as horizontal strength of MI have been
157 examined from these data sets and compared with satellite observations.

158 **2.4. AIRS observations**

159 AIRS onboard the Earth Observing System (EOS) - Aqua satellite of NASA was launched in
160 2002. This is also a polar orbiting satellite which crosses the equatorial latitudes at 13:30 hr LT and
161 01:30 hr LT for the ascending and descending pass, respectively. The orbit period is 98.99 min, and
162 the orbit is sun synchronous with consecutive orbits separated by 2760 km at the equator. AIRS has a
163 field of view of 1.1° and provides a nominal spatial resolution of 13.5 km for IR channels and
164 approximately 2.3 km for visible/near-IR channels. AIRS data together with data from the Advanced
165 Microwave Sounder Unit (AMSU) (Lambrigtsen, 2003) are used in the present study. We make use
166 of AIRS support data which have higher vertical resolution with 100 levels between 1100 and 0.016
167 hPa. For the present study we restrict data only from surface to 600 hPa which have vertical

168 resolution of 30-20 hPa. Though these data are available since 2003, we make use data from 2009
169 only so as to compare with other data sets.

170 **2.5.COSMIC GPS RO**

171 GPS RO technique is also a remote sounding satellite technique, and it uses the radio signals
172 received onboard a low Earth orbiting satellite from atmospheric limb sounding. The GPS RO
173 measurements have a vertical resolution ranging from 400 m to 1.4 km, which is much better than
174 that of any other satellite data (Kursinski et al., 1997). COSMIC has vertical resolution of ~ 100 m in
175 the lower troposphere for temperature. The COSMIC GPS RO was successfully launched in mid-
176 April 2006 (Anthes et al., 2008). Since 17 July 2006, COSMIC GPS RO provides accurate and high
177 vertical resolution profiles of atmospheric parameters that are almost uniformly distributed over the
178 globe. COSMIC provides a direct estimate of refractivity (from measurement of bending angle by
179 GPS technique) at very high vertical resolution, but have poor repetitivity.

180 **3. Methodology and analysis procedure**

181 As mentioned earlier, MI was first observed by Colon (1964) and Ramage (1966) over the
182 AS from ship upsonde profiles. They reported that MI lies between 900 and 800 hPa with strong
183 intensity over western AS (WAS) and weakens as its base rises and comes to eastern AS (EAS).
184 Following this study, Narayanan and Rao (1981) have shown MI's presence using the temperature
185 difference (ΔT) between the TIROS-N derived sea skin temperature and atmospheric layer mean
186 temperature (between 1000 hPa and 850 hPa).

187 Note that lapse rate (dT / dz) of atmosphere at the tropospheric altitudes is a negative
188 quantity. However, in this study (and also of Narayanan and Rao, 1981), we have considered ΔT as
189 temperature difference between a lower level (higher temperature) and a higher level (lower
190 temperature), so is normally a positive quantity of value $\sim + 6$ to $+ 7$ K. For inversion regions, it is
191 negative or a small positive quantity (i.e. less than $+ 2$ K).

192 After considering several limitations in the satellite data of that time, Narayanan and Rao
193 (1981) finally considered MI when the difference ΔT , between surface and layer mean temperature
194 (of 1000 to 850 hPa), is 2 K or less, which otherwise was greater than 3 K. Since then, several
195 improvements in the satellite instruments, retrieval techniques and data products have come up in
196 these three decades.

197 Extensive in-situ observations of AS MI features were obtained during FGGE-MONEX 1979
198 experiment. Fig. 1a shows a typical example of MI observed in T (temperature) and RH (Relative
199 Humidity) data obtained on 27 June 1979 at 0656 GMT at 20°N, 62°E from radiosonde. In this
200 example MI starts from surface and temperature departure is as high as ~ 10 K from a normal lapse
201 rate profile at 900 hPa. The vertical extent of inversion varies from 0.5 km to even more than 1 km.
202 It is to be noted that AS MI are much stronger and long lasting i.e. less diurnal variation than normal
203 boundary layer and persist for many days compared to those over land regions.

204 A detailed analysis is made in this study by considering several thousands of profiles
205 obtained from different satellite observations now available over AS for redefining MI. Since the
206 MIs occur at low levels, first we tried with the earlier adopted criteria of Narayanan and Rao (1981)
207 i.e., by taking difference between sea surface (skin) temperature and 925 hPa level (mean pressure
208 level of 1000 - 850 hPa MLT of TIROS-N data of the 1980 time frame) temperature and found those
209 to be noisy for detecting MI. To avoid the surface emissivity effects in the retrieval at / near surface
210 (from the sounder instrument), we adopted the lower level in the present study as 950 hPa instead of
211 sea surface / skin temperature. It was considered not appropriate to use SST/skin temperature
212 (though may be of higher accuracy) from a different source (viz imager onboard the same satellite)
213 for estimating ΔT . It was felt that this will not give the advantage of the differencing procedure
214 employed earlier to detect inversion (Narayanan and Rao, 1981). This level criterion (950 – 850 hPa)
215 was arrived at after a detailed examination of ΔT at a few more level intervals (viz 1000 – 900 hPa,
216 1000 – 850 hPa, etc).

217 Thus, we have used:

$$218 \quad \Delta T = T(950 \text{ hPa}) - T(850 \text{ hPa}) \quad (1)$$

219 to delineate MI. However, the actual levels used were 958 hPa and 852 hPa at which the support data
220 are available from the NOAA website.

221 While considering the normal atmospheric lapse rate of + 6 to +7 K / km (average of 340
222 non-inversion cases obtained during MONEX, figure not shown), it is expected to observe a ΔT of +
223 6 to +7 K between 950 and 850 hPa (~ 1 km height difference). Note that Narayanan and Rao (1981)
224 have identified inversion (non-inversion) region as $\Delta T \leq + 2 \text{ K}$ ($\Delta T > + 2 \text{ K}$) in TIROS – N satellite
225 data for a height range difference of ~ 0.75 km. For the present study (for 1 km height difference) the
226 same would translate to $\Delta T \sim + 2.7 \text{ K}$ for inversion delineation. However, to be on the safe side and
227 to provide margin of error, we have still considered $\Delta T \leq + 2 \text{ K}$ as criterion of inversion region. The
228 interval 2.0 K to 2.7 K may still be a grey region which could be interpreted as inversion region on
229 some occasions. The criterion of $\Delta T \geq + 4 \text{ K}$ as non – inversion regions has been adopted. In the
230 example shown in Fig. 1a, ΔT is (minus) - 1.3 K (note however, that the actual inversion value is ~ -
231 5 K between surface and 900 hPa).

232 In general, a sudden drop in the water vapor just above the inversion is observed (e.g. RH
233 drop of ~ 70% in Fig 1a). Since all the data sources mentioned in section 2 provide water vapor
234 information, we also have examined the changes happening in water vapor near/above the inversion
235 altitude. In general, inversion is identified in the temperature (water vapor) where it increases
236 (decreases sharply) instead of decreasing (decreasing gradually) with altitude. For obtaining detailed
237 characteristics of MIs over the Arabian sea, we have selected three $3^\circ \times 3^\circ$ grid boxes centered at
238 latitude 18.5° N , and located at longitudes 60° E as WAS, 64° E as CAS (central AS), 71° E as EAS
239 (as shown in Fig.3).

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242 **3.1. Quality checks for the profiles and volume of data**

243 Each temperature profile from the satellite data was interpolated from surface to 500 hPa (26
244 levels of support data) at 0.25 km intervals for our preliminary analysis. We have used the quality
245 flag 0 and 1 from the given data set which are corresponding to best and good. There were many
246 erroneous profiles which could be observed even from a cursory examination of the data. The
247 temperatures at a few / more levels were far wide of the normal profile. To account for these types of
248 profiles, we applied a quality check to filter out spurious data. All profiles of July and August
249 months of 2009 (poor monsoon year) and 2011 (normal monsoon year) were sorted out in 3 x 3
250 boxes of WAS and EAS. For each month the mean and standard deviation were obtained for each
251 interpolated levels separately. Those profiles for which the data at any one level was lying beyond + /
252 - 2 sigma of the mean, were not considered for further analysis. From this procedure we saw that
253 nearly 25 – 30 % of profiles were getting filtered out.

254 Using these quality checked profiles, the procedure for selecting the right levels for
255 calculating ΔT was established. Thereafter, for all the other monsoon days of the five years, we have
256 computed ΔT for individual profiles by an automated procedure (without resorting to examining each
257 profile). They were grouped and their ΔT values averaged in $1^\circ \times 1^\circ$ bins over the whole AS region.
258 Diurnal variation of ΔT was examined for a few months of data. Once we made sure that this is not
259 discernible, the day and night data of a calendar day were merged in $1^\circ \times 1^\circ$ boxes.

260 For further analysis, the average ΔT values for the day (24 hr period) at $1^\circ \times 1^\circ$ grids have
261 been used. Due to averaging of ΔT of all the profiles in $1^\circ \times 1^\circ$ box and morning and evening passes
262 (~ 6 to 8 values of ΔT in 24 hours), the strength of MI may be getting somewhat reduced (as MI
263 occur at slightly different levels within a vertical range of 25 - 50 hPa, for different profiles in the
264 same $1^\circ \times 1^\circ$ box). For some studies (e.g. for Fig 2, 4, 5, etc), we have used only a limited data from
265 this total data set. The total number of profiles considered for the five years amount to nearly half a

266 million, each for AIRS and IASI – considering that nearly 30 % profiles did not pass through our
267 quality check.

268 **4. Results and Discussions**

269 **4. 1. Monsoon Inversions observed in satellite and ERA-Interim datasets**

270 Fig. 1a and 1b show MI observed on 27 June 1979 at 0730 GMT at 20°N, 60°E through
271 MONEX radiosonde and ERA – Interim data, respectively. The detailed comparison study between
272 TIROS – N satellite data of 1979 and concurrent in-situ MONEX radiosonde profiles for 1979
273 southwest monsoon carried out by Narayanan and Rao (1981) is summarized in Table 2. This was
274 the only occasion (1979) when in-situ data were available over AS to compare with satellite
275 soundings. Thus, comparison of current satellite observations is being done in this study with ERA-
276 Interim data. In this case, ERA – Interim data also catches the inversion but with a less rise in
277 temperature (~ 3 - 4 K) and decrease in RH (~ 60%). To show how the present day satellites reveal
278 MI, typical profiles of temperature and RH obtained from collocated IASI and ERA-Interim on 30
279 July 2009, 0530 GMT are plotted in Fig. 1c, and 1d, respectively. A clear MI in the satellite profile
280 and ERA-Interim can be noticed though with somewhat varying strengths and base of inversion
281 height. However, the top height of inversion is consistent. These are the first reported results of MI
282 features seen directly from the satellite observations over the AS which were shown earlier by
283 Narayanan and Rao (1981) in an indirect way by using ΔT indices. In general, in the individual
284 satellite profiles, we are able to see the MI strengths ranging from ~ + 2 to - 6 K (-8.8 K being the
285 actual temperature difference between 930 hPa and 850 hPa in Fig. 1c). These MI lie mostly below
286 850 hPa level, but in rare occasions we could see them even up to 700 hPa over the EAS – but of
287 much weaker strength. The strength of MI is also seen to be decreasing from WAS to EAS which
288 will be discussed in detail in later sections.

289 Thus, in Fig. 1, we have seen examples of MI comparison between radiosonde and ERA
290 interim (1979) and between IASI and ERA-Interim (2009). There are some minor inconsistencies by

291 way of inversion heights in individual profiles of the three data sets. However, our objective here is
292 to examine the large scale characteristics of MI by considering average ΔT computed from individual
293 profiles in $1^\circ \times 1^\circ$ grids.

294 **4. 2. Contrasting behavior of MI between WAS and EAS**

295 As observed from Fig. 1, MI can lie between surface and ~ 2 km during Indian Summer
296 Monsoon (ISM) season (JJAS). Careful examination of time evolution of ΔT over the western
297 Arabian sea reveals that the MI start forming around first half of May and dissipate around late
298 September. Fig. 2 shows the evolution of the MI during two contrasting years (2009 a poor monsoon
299 year and 2011 a normal monsoon year). During the peak monsoon season of July – August, the
300 difference in ΔT between the two years are prominently noticed. Also MI is more frequently
301 observed with higher strength during the peak monsoon months of July and August. To investigate
302 further their contrasting features in WAS and EAS, data only of July and August from 2009 to 2013
303 are presented.

304 In Fig. 3 we have summarized the three important characteristics of MI viz their base altitude,
305 strength (as revealed by ΔT) and percentage occurrence during the complete season. For brevity, the
306 results of only July and August months, averaged for all the five years 2009 – 2013 are shown in the
307 figures. Fig. 3a and 3b show the spatial variation of base altitude of MI during July and August,
308 respectively. The contrasting feature of base altitude of occurrence of MI is seen mainly north of 15°
309 N from the selected three grid boxes. It increases from WAS (below 1 km) to EAS (above 1.5 km)
310 through CAS (1.0 -1.5 km).

311 As mentioned earlier, from very limited observations previous studies (Colon, 1964; Ramage,
312 1966; Narayanan and Rao, 2004) had suggested that strength and frequency of occurrence of the MI
313 will be more over WAS than over EAS. To investigate this contrasting behavior of MI in detail from
314 satellite soundings, we examined the spatial variations of ΔT . Fig. 3c (July) and 3d (August) shows
315 the strength of MI increasing from EAS to the WAS and is prevalent mainly north of 15° N latitude

316 extending from 15°N to 25° N latitude and 55° E to 68° E longitude. The strength of MI can be
317 noticed as $\sim + 2$ K near Arabia coast and as we approach Indian coast, the normal environmental
318 lapse rate condition of $+ 6$ to $+ 7$ K/km are encountered. From these figures a clear contrast in ΔT a
319 difference of around 2 K in the southeast quadrant of AS between July and August is also noticed. In
320 general, the AS is covered with lapse rate of $+ 4$ K/km, which is the condition for taking the
321 atmosphere towards stability during the August month. The region of Somali low level jet is the
322 location of permanent region of MI during the month of July. In the spatial distribution of monsoon
323 low level jet shown by Roja Raman et al. (2011) reveals that the center of the core is seen around
324 13°N and 60°E and exists strong shear between 850 hpa and 700 hpa. Strong surface winds of south-
325 west monsoon produce an Ekman transport perpendicular to the wind flow with strong upwelling in
326 the region which in turn brings the cool water from the deeper layers to surface. Simon et al. (2007)
327 showed that WAS region is the region of Somali upwelling, and also since the low level jet and
328 surface wind are of the order of ~ 20 m/s, they produce sufficient cooling and the air above this
329 region is still warmer when compared to the upwelling area, producing strong inversion.

330 Fig. 3e and 3f shows the spatial variation of percentage occurrence (PO) of MI during July
331 and August months. PO is calculated corresponding to $\Delta T \leq + 2$ K criteria. In general, it is observed
332 that WAS show more number of MI cases (50 to 70%) compared to EAS (10 to 20%). ERA-Interim
333 data show only 30 to 50% cases of MI over WAS which will be dealt in detail in the following sub-
334 sections. The maximum PO during the four months of monsoon over the WAS are 40 % (June), 60
335 % (July), 50 % (August) and 30 % (September) (figure not shown). The areal extent of the maximum
336 PO is seen during July. During September, very small area of Northern AS is covered with ~ 50 %.
337 No inversion is seen in the EAS box during the June and September periods. Despite its low strength
338 (ΔT) PO show maximum occurrence of 60% in July. Since the PO and strength of MI over the CAS
339 is in between the features of EAS and WAS, for further discussions pertain, only WAS and EAS
340 boxes.

341 The PO of ΔT value in different ranges observed in IASI for the five monsoon seasons is
342 shown in Fig. 4. ΔT values range from -2 to + 6 K (0 to + 7 K) in WAS (EAS) with peak occurring
343 around + 1 to + 2 K (+3 to +4 K). There are only a few values of ΔT less than + 2 K in EAS. Similar
344 analysis is also made using ERA-Interim data and is shown in bottom panels of Fig. 4. ERA-Interim
345 data shows the contrast between WAS and EAS more clearly. In case of q at 700 hPa a difference of
346 about 2 g/kg can be noticed, with EAS having higher humidity values than WAS in IASI. However,
347 ERA-Interim data does not show this distinction.

348 To further examine the contrasting behavior between EAS and WAS, time series of ΔT and
349 water vapour at 700 hPa is considered for different years. Daily mean variations of ΔT and specific
350 humidity, q , at 700 hPa in WAS and EAS during the monsoon season of the year 2012 observed by
351 IASI is shown in Fig. 5. Note that we have included results of all the days irrespective of whether MI
352 is present or not. Three point average smoothed curves are shown in the respective panels. In
353 general, it can be seen that WAS ΔT (q at 700 hPa) values are $\sim + 2$ K (1 - 2 g/kg) less than those
354 over EAS for the season as a whole (Fig. 5a and 5b). During all the years (2009 - 2013) of the
355 present study, IASI reveals (figure not shown) this feature. Similar analysis has been carried out
356 using ERA-Interim reanalysis data and is shown in Fig. 5c and 5d. A clear contrast between WAS
357 and EAS in ΔT can be noticed in ERA-Interim data. A mean difference of ~ 2 K (~ 1 g/kg) can be
358 noticed in ΔT (q at 700 hPa) between WAS and EAS, EAS values being lower. A cyclic behavior in
359 ΔT variations with a period of ~ 20 -25 days in case of ERA-Interim is noticed but not observed in
360 the satellite measurements. There exists no significant diurnal variation in ΔT (figure not shown).
361 This was verified before averaging ΔT of all profiles (day and night) in the $1^\circ \times 1^\circ$ grids. Due to
362 inversion and stability, moisture is getting trapped at lower levels over WAS compared to EAS as
363 indicated in Fig. 5b and 5d observed from IASI and ERA-Interim, respectively.

364

365

366 **4.3. Relation between MI over AS and monsoon activity**

367 Past investigations (e.g. Gadgil and Joseph, 2003) showed that the mesoscale monsoon
368 features largely vary with the activity of the monsoon. In general during the active phase of the ISM,
369 typically there will be more precipitation over central India (18°-28°N and 65° to 88°E). Similar
370 variations in precipitation during the monsoon season can also be expected on regional scales.
371 Gadgil and Joseph (2003), Kripalani et al. (2004), Rajeevan et al. (2006) have considered the daily
372 rainfall time series over central India during monsoon months along with the climate normal to
373 delineate 'active' and 'break' periods over the Indian region. On the basis of this data, Rajeevan and
374 Bhate (2009) have defined active and break phases over central India by considering the days
375 exceeding the climate mean with +1 (-1) standardized anomaly as active (break) periods provided it
376 should persist at least for 3 days (triad).

377 Fig. 6 shows the latitude - longitude cross section of ΔT and q at 700 hPa for active (14 - 17
378 July 2009) and break (30 July - 11 Aug. 2009) spells for the monsoon season of 2009 observed using
379 IASI and ERA-Interim data. Irrespective of the data source, ΔT and associated q at 700 hPa reveal
380 that a large part of WAS is covered with MI ($\Delta T \leq +2$ K and less moisture values) up to west of ~
381 68° E during the break spell as seen in Fig. 6a and 6e. In the north AS, MI reach as close as Gujarat
382 coast during break spells (especially in ERA-Interim data), but are restricted to WAS during active
383 spells. During the active spell, the inversion regions from ΔT maps are patchy west of 65° E in Fig.
384 6c. Also strengths of ΔT in WAS are more as observed by ERA-Interim than by IASI during break
385 spells. ERA-Interim shows (Fig. 6e and 6g) more smoothed results and there is less change in area
386 extent in this case. Specific humidity q at 700 hPa shows clear result that during the break spell AS
387 has less moisture and more during the active spell. One can notice the feature of inversion from the
388 figure where water vapor is being trapped in the lower portion resulting in less moisture over WAS
389 and more over the EAS. Thus, the q values also give a good indication of the inversion feature.

390

391 **4.4. MI during normal and poor monsoon years**

392 It is well known that strong MI suppresses the vertical development of clouds; rain cannot
393 occur in such situations (Sathiyamoorthy et al., 2013). Using ARMEX-I (2002) data, Bhat (2006)
394 could notice strong and persistent inversions in the atmosphere over the AS and west coast of India.
395 This data proved very valuable as July 2002 rainfall was the lowest in the recorded history and the
396 data collected over the AS and on the west coast helped in understanding the conditions that
397 prevailed over the eastern AS during one of the worst monsoon years. The relation between MI and
398 central India rainfall is further investigated by separating the MI observed during normal (2010 -
399 2013) and poor monsoon (2009) years. Time variations of ΔT observed over WAS during two
400 contrasting years of 2009 and 2011 obtained from IASI measurements and ERA-Interim data are
401 shown in Fig. 7. It can be seen that good monsoon year 2011 has higher ΔT than poor monsoon year
402 2009 (Fig. 7a), and is the same for q i.e. higher value for the year 2011 (Fig. 7b). ΔT is observed to
403 be lower by about 2 K during the season as a whole in the poor monsoon year when compared to the
404 good monsoon year, suggesting the possibility of a variation of this parameter between normal and
405 poor monsoon years. This aspect is clear from the right panels where difference between 2011 and
406 2009 observed in ΔT (Fig. 7c) and q at 700 hPa (Fig. 7d) are shown. From this figure we can infer
407 that the year 2009 has less value of ΔT and less value for q suggesting stronger MI during poor
408 monsoon year. Note that during most of the time, the temperature in 2011 is higher (the difference
409 between 2011 and 2009 showing positive values) and less temperature lapse rate means more stable
410 layered atmosphere. In 2011, WAS temperature show higher values revealing less MI over AS when
411 compared to 2009. The decreasing trend in ΔT is discernible in difference plots for some particular
412 epochs. In general, ERA-Interim also show these features (Fig. 7e and 7f), but only to a moderate
413 extent. It may be noted that these inferences are based on the results of only one poor monsoon year
414 (2009).

415

416 **4.5. Inter-comparison of MI features with IASI, AIRS and ERA**

417 Inter-comparison of the gross features of PO of MI (with $\Delta T \leq 2$ K) in WAS and EAS
418 estimated for the five years of monsoon season by IASI, AIRS and ERA-Interim data are shown in
419 Fig. 8. In general, when we consider ΔT as a parameter to detect MI, clear contrasting feature
420 between WAS and EAS with higher PO in WAS can be noticed in all the data sources mentioned
421 above. PO in the IASI measurements ranges from 23% to 54%. Among these data sets, ERA-Interim
422 shows huge difference in the percentage occurrences between WAS and EAS, to the extent that not
423 even a single MI is seen in EAS in any year. Since the vertical resolution of the IASI temperature
424 profiles is better than AIRS, higher PO of MI in the WAS is noticed throughout when compared to
425 AIRS, except in the case of 2012. However, ERA- Interim being a combination of model and
426 observations, it is not able to pick up the MI in the EAS where the strength of inversion is also
427 weak. The artifact of the model appears to be smoothening the MI features of IASI when it is
428 assimilated in the ERA – Interim.

429 Coming to the satellite observations, during five years, IASI shows higher PO of MI than
430 AIRS except for 2012 for WAS. A distinct contrast between WAS and EAS with higher PO in the
431 former region can be noticed. When we consider EAS as a place to detect MI, AIRS observed always
432 higher PO than IASI and almost nothing is noticed in ERA-Interim. Thus, we may infer that IASI is
433 performing better than AIRS for detecting MI (as ERA is in better agreement with IASI rather than
434 with AIRS). Note that large inter-annual variability in MI is observed and this is expected to reflect
435 in the monsoonal activity over Indian region. It can also be seen that there is a steady decrease of PO
436 of MI as observed by IASI from 2009 to 2013. No such feature is observed in AIRS – which shows
437 more random behavior over the different years.

438 We have made the scatter plot of ΔT observed by IASI and AIRS over WAS and EAS (figure
439 not shown). The scatter does not suggest that these two data sets can be combined to study the small
440 changes of ΔT in their intra-seasonal and inter-annual variations. This and the other differences

441 related to q at 700 hPa constrained us not to combine the AIRS data with IASI data in the present
442 study.

443 **4.6. Monsoon Inversion derived from other parameters**

444 Narayanan and Rao (1989) had also considered equivalent potential temperature (θ_e)
445 differences to study MI. θ_e incorporates the effect of both temperature and humidity. However, the
446 dynamic range of $\Delta\theta_e$ is no better than that of ΔT . Recall that the troposphere is statically stable on
447 average, with a potential temperature gradient of 3.3 K/km (Wallace et. al., 2006). We make use of
448 another index here viz atmospheric refractivity (N) for identifying MI. Similar to θ_e , Refractivity
449 (N), is another atmospheric parameter which is a function of temperature and water vapor. It was
450 shown that better information on boundary layer can be obtained from refractivity profiles than
451 virtual potential temperature though both has temperature and water vapor information (Basha and
452 Ratnam, 2009). Refractivity, N has a higher dynamic range and vertical variation as compared to
453 temperature (~ 15 N units vis a vis 2 K). More advantage of using N for delineating MI will be
454 available, provided, it is measured directly, for example, using GPS Radio Occultation technique,
455 instead of computing it from temperature and water vapor obtained from the sounders or from
456 radiosonde. However, the spatio-temporal density of direct N observations is too sparse to get
457 meaningful statistics over equatorial regions.

458 We have computed refractivity N , from temperature and water vapor data of IASI (and
459 MONEX radiosonde data), given by the expression:

$$460 \quad N = 77.6 \left(\frac{P}{T} \right) + 3.73 \times 10^5 \left(\frac{e}{T^2} \right) \quad (2)$$

461 Where P is pressure, T temperature and e water vapor pressure.

462 Similar to ΔT we have defined an index ' ΔN ' as:

$$463 \quad \Delta N = N(950 \text{ hPa}) - N(850 \text{ hPa}) \quad (3)$$

464 Profile of N computed from the temperature and humidity profiles of dropsonde (Fig. 9a) of
465 MONEX time is shown in Fig. 9b. A drastic decrease in N (by 129 N units between 950 and 850
466 hPa) can be noticed near MI altitudes in this example. Thus, N can also be taken as a potential
467 parameter to delineate inversion and for studying spatial and temporal variations of MI.

468 In order to see the relation between ΔT and ΔN , we have estimated ΔN using all the MONEX
469 profiles obtained over AS. These include both inversion and non-inversion cases. There were 32
470 (346) profiles with inversion (non- inversion). Note that $\Delta T \leq + 2$ K and $\Delta T > + 4$ K are only
471 considered for obtaining above statistics and there exists 34 profiles in the transition zone (+ 2 to + 3
472 K). Scatter plot between ΔT and ΔN for all 411 in-situ profiles of MONEX over AS is shown in Fig.
473 9c. Correlation coefficients between the two parameters are found to be 0.56 with 15.7 as standard
474 deviation. Note that $\Delta T \leq + 2$ K (inversion region) corresponds to $\Delta N > 50$ N units which is shown
475 as blue line in Fig. 9c. We can infer that if ΔN is less than 50 N units it corresponds to non-inversion
476 region (ΔN more than 50 may be inversion or otherwise). ΔN is thus a supportive parameter to ΔT in
477 identifying inversion / non inversion. Because of its larger dynamic range, details of inversion have
478 been identified in the ΔT and ΔN maps (figure not shown).

479 It is well known that COSMIC satellites are able to provide N profiles directly. The spatial
480 and temporal sampling of COSMIC at any particular region are, however, very meager. The
481 comparison map of ΔN from IASI and ΔN from COSMIC combined for a long break spell from 30
482 July to 11 August 2009 has been studied. This long period accumulation of data was necessary to
483 have sufficient data points from COSMIC to cover the entire AS. One can see ΔN values above 50 N
484 units (inversion region) covering the entire Arabian sea corresponding to ΔT values being below 2 K
485 (shown by IASI, figure not shown). Over the AS region ΔN observed for all the five years of our
486 study was combined to produce the frequency distribution of ΔN over Western AS (5 – 25 °N, 56 –
487 65 °E, excluding land) and Eastern AS (5 – 25 °N, 66 – 75 °E, excluding land) and is shown in Fig
488 10. Over WAS, 712 cases and over EAS 547 cases are showing $\Delta N > 50$ N units (which may be

489 supportive to inversion). A difference of about 10 N units can be noticed, with WAS having higher
490 ΔN values.

491 **5. Summary and Conclusions**

492 Low level MI characteristics, which usually occur below 700 hPa over the AS during
493 southwest monsoon months, have been identified directly from operational satellite temperature
494 retrievals. For the first time we have shown here cases of direct and unambiguous delineation of MI
495 from the satellite temperature and water vapor retrieval observations. We have used five years (2009-
496 2013) data of two different satellite sounder instruments (mainly from IASI and for inter comparison
497 AIRS) along with ERA-Interim reanalysis data to delineate the characteristics of MI over AS. Their
498 percentage occurrence, base height and strength have been studied. For supporting our findings, we
499 also compare with the campaign of MONEX 1979 in-situ measurements over AS. The main findings
500 obtained from the observational study are summarized in the following:

- 501 1. Percentage occurrences of MI over WAS (up to $\sim 65^\circ\text{E}$) is $\sim 60 - 70\%$ and are always higher
502 and stronger than over EAS. WAS ΔT values are $\sim 2\text{ K}$ less than those over EAS.
- 503 2. MI is stronger during poor monsoon year (2009) and occurs on more occasions in WAS
504 during break spells. Whether this is true or not for all poor monsoon years need to be checked
505 with more years of data.
- 506 3. ERA-Interim is also able to provide these features but is restricted to some parts of AS with
507 more smoothed variability.
- 508 4. Inter-comparison of IASI and AIRS profiles from the view of study of inversion suggests the
509 differences do not warrant a mix of these two data sets for this study.
- 510 5. The refractivity data has only a supporting role to identify monsoon inversion regions.

511 Thus, MI seems to be a semi-permanent feature of Indian summer monsoon. It is suggested to
512 include this feature also in future monsoon diagnostic and forecast studies.

513

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520 (<http://www.nsof.class.noaa.gov>/for IASI) for IASI through ftp. We also thank IMD for providing
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597 **Figure captions:**

598 **Figure 1.** Typical examples showing MI in T and RH on (a) 27 June 1979 at 0730 GMT at 20°N,
599 60°E obtained from radiosonde from MONEX experiment, (b) same as (a) but at 0600 GMT from
600 ERA, (c) 30 July 2009 at 0514 GMT at 22°N, 68°E by IASI, (d) 30 July 2009 by ERA-Interim at
601 same location but at 0600 GMT. Note that scale for RH is shown in the top axis of (a) and (b).

602 **Figure 2.** Time series of ΔT for starting and ending of MI from April to October 2009 (black) and
603 2011 (blue). Green vertical lines are showing starting (01 May 2009) and ending (07 October 2009)
604 time for MI.

605 **Figure 3.** Base altitude occurrence of MI during (a) July, (b) August, ΔT (Strength) of MI (c) July,
606 (d) August, and Percentage occurrence of MI days (e) July, (f) August, averaged during 2009-2013
607 observed by IASI. (We are selecting WAS, CAS and EAS from this figure).

608 **Figure 4.** Percentage occurrence of (a) ΔT and (b) q at 700 hPa observed in WAS and EAS during
609 monsoon season of the years 2009-2013 for various ranges of ΔT and q at 700 hPa by IASI. (c) and
610 (d) same as (a) and (b) but obtained from ERA-Interim data.

611 **Figure 5.** Time series of (a) ΔT and (b) q at 700 hPa observed over WAS and EAS grid boxes
612 during the monsoon season of the year 2012 by IASI, (c) and (d) same as (a) and (b) but obtained
613 using ERA – Interim data. 3-point smoothed curves are shown.

614 **Figure 6.** MI observed in (a) ΔT and (b) q at 700 hPa during break spells (30 July – 11 August 2009)
615 of the year 2009 by IASI, (c) and (d) same as (a) and (b) but observed during active spells (14-17
616 July 2009). (e) and (f) and (g) and (h), same as (a) and (b) and (c) and (d) but observed by ERA-
617 Interim, respectively.

618 **Figure 7.** Time variations of (a) ΔT and (b) q at 700 hPa observed over WAS during two contrasting
619 years of 2009 and 2011 by using IASI measurements. Difference between 2011 and 2009 observed
620 in (c) ΔT and (d) q at 700 hPa. (e) to (h) same as (a) to (d) but observed by using ERA-Interim data
621 products.

622 **Figure 8.** Percentage occurrence of MI observed with (a) $\Delta T \leq 2K$ using IASI, AIRS and ERA-
623 Interim data during monsoon seasons of 2009-2013 over WAS and EAS.

624 **Figure 9.** Typical examples showing MI in temperature and RH on (a) 27 June 1979 at 0656 GMT at
625 $20^{\circ}N$, $62^{\circ}E$ obtained from dropsondes from MONEX experiment, (b) N profile (c) Scatter plot of
626 ΔT and ΔN .

627 **Figure 10.** Frequency of ΔN observed in Western AS and Eastern AS during monsoon season of the
628 years 2009-2013 for various ranges of ΔN by COSMIC. Western AS is showing higher
629 values means inversion is there.

630

631 **Table captions:**

632 **Table 1:** Data details for accuracy/error and availability.

633 **Table 2:** Comparison of aircraft profiles with satellite data.

634 **Table 1:** Data details for accuracy/error and availability.

	IASI	AIRS	COSMIC GPS - RO	ERA-Interim	MONEX 1979 In-situ data
Launch of satellite	MetOp – A launched in October 2006, 8461 spectral Channels	Aqua launched in May 2002, 2378 spectral channels	GPS – RO microsatellite receiver launched in April 2006	---	May – August 1979
Data availability from	August 2008	2003	April 2006	1979	May – August 1979
Data used in the present study	June – September 2009 - 2013	June – September 2009 – 2013	June – September 2009 - 2013	June – September 2009 - 2013	May – August 1979
Accuracy in Temperature	~ 1 K(RMS) at a vertical resolution of 1 Km(Clerbaux et al., 2007; 2009)	~ 1 K at a vertical resolution of 1 Km(Susskind et al., 2003)	Generally ~ 100m in the lower troposphere (not for T)	0.5 – 1.0 K at a vertical resolution of 0.8 – 1.0 km	± 1 °C in 4 vertical levels resolution(WMO report)
Accuracy in Humidity	~10 – 15 % accuracy with a 1 – 2 Km vertical resolution(Clerbaux et al., 2007; 2009)(Schlüssel et al., 2005)	~15 % accuracy with a 2 Km vertical layer resolution(Susskind et al., 2003)	---	~7.0 – 20 % at a vertical resolution of 0.8 – 1.0 km	± 30 % at a vertical resolution of 4 levels.
Accuracy in Refractivity	---	---	400 m to 1.4 km (Kursinski et al., 1997),		
Horizontal resolution	15 Km	25 Km	2000 soundings per day	1.5° x 1.5° (~ 80 km)	500 km
Pressure levels	1100- 0.0161 hPa - 100	1100 – 0.0161 hPa – 100	70% of occultations penetrate below 1 km (Anthes et al., 2008)	1013 – 1 hPa 37	1000 – 294 Different -2
Local equator crossing time	0930 LT descending node	1330 LT ascending node	-----	----	---
Swath	2200 km	1650 Km	-----		

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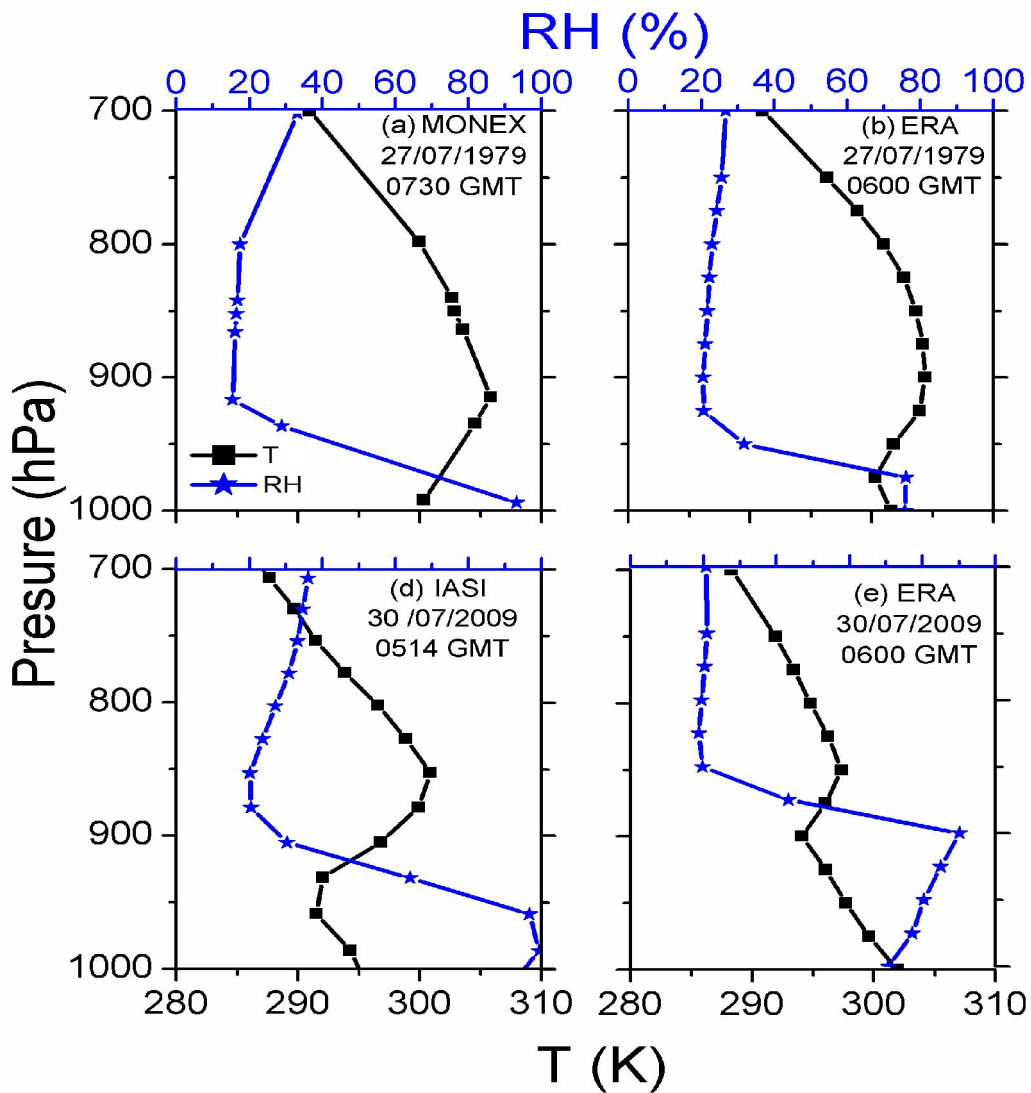
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637 **Table 2:** Comparison of aircraft profiles with satellite data.

	Aircraft profiles	Near simultaneous satellite data	
		$\Delta T \leq 2^{\circ}\text{C}$	$\Delta T \geq 3^{\circ}\text{C}$
No. Of profiles with well – marked inversion below 850 mbar	30	23	7 (for four of them $\Delta T = 3^{\circ}\text{C}$)
No. Of profiles without well – marked inversion	129	0	129

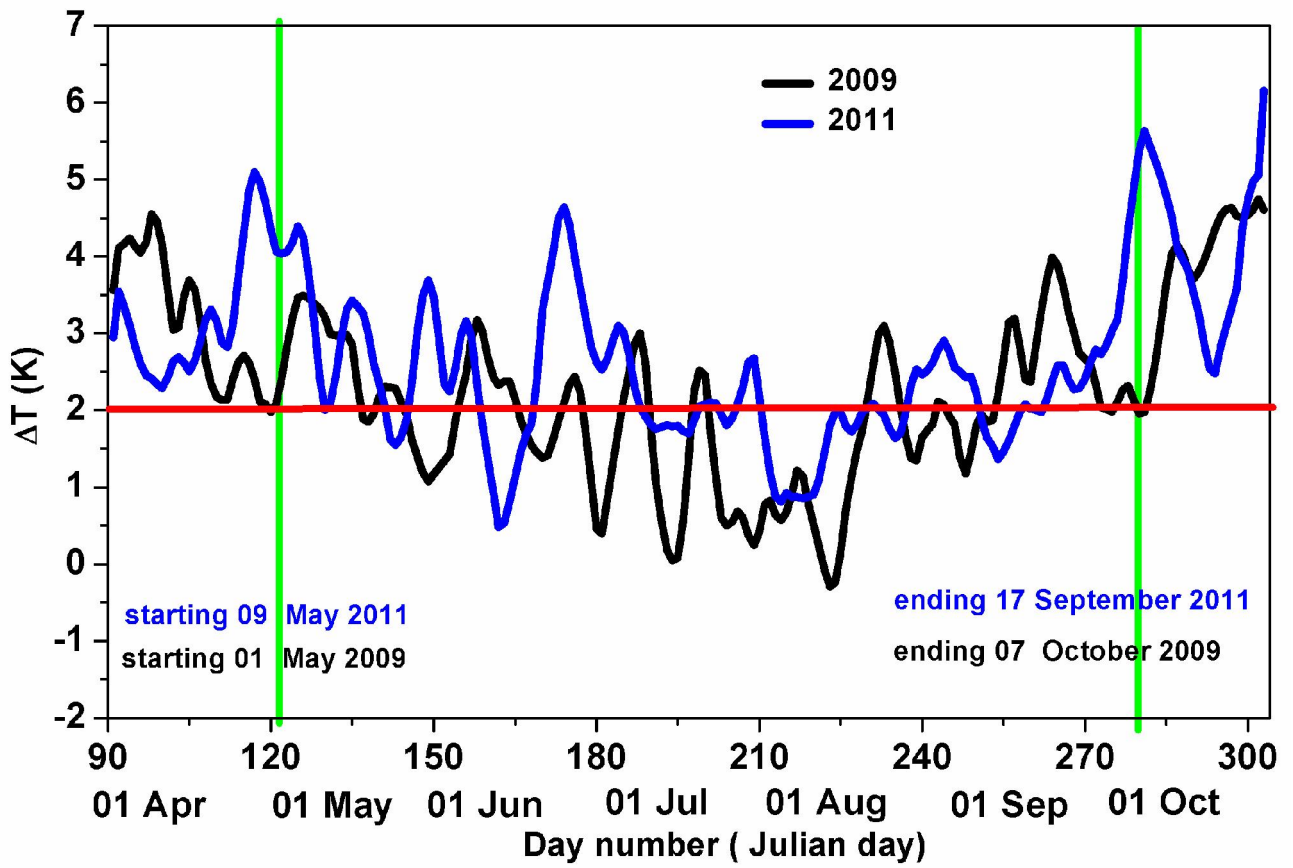
638 (Regenerated from Narayanan et al., 1981)

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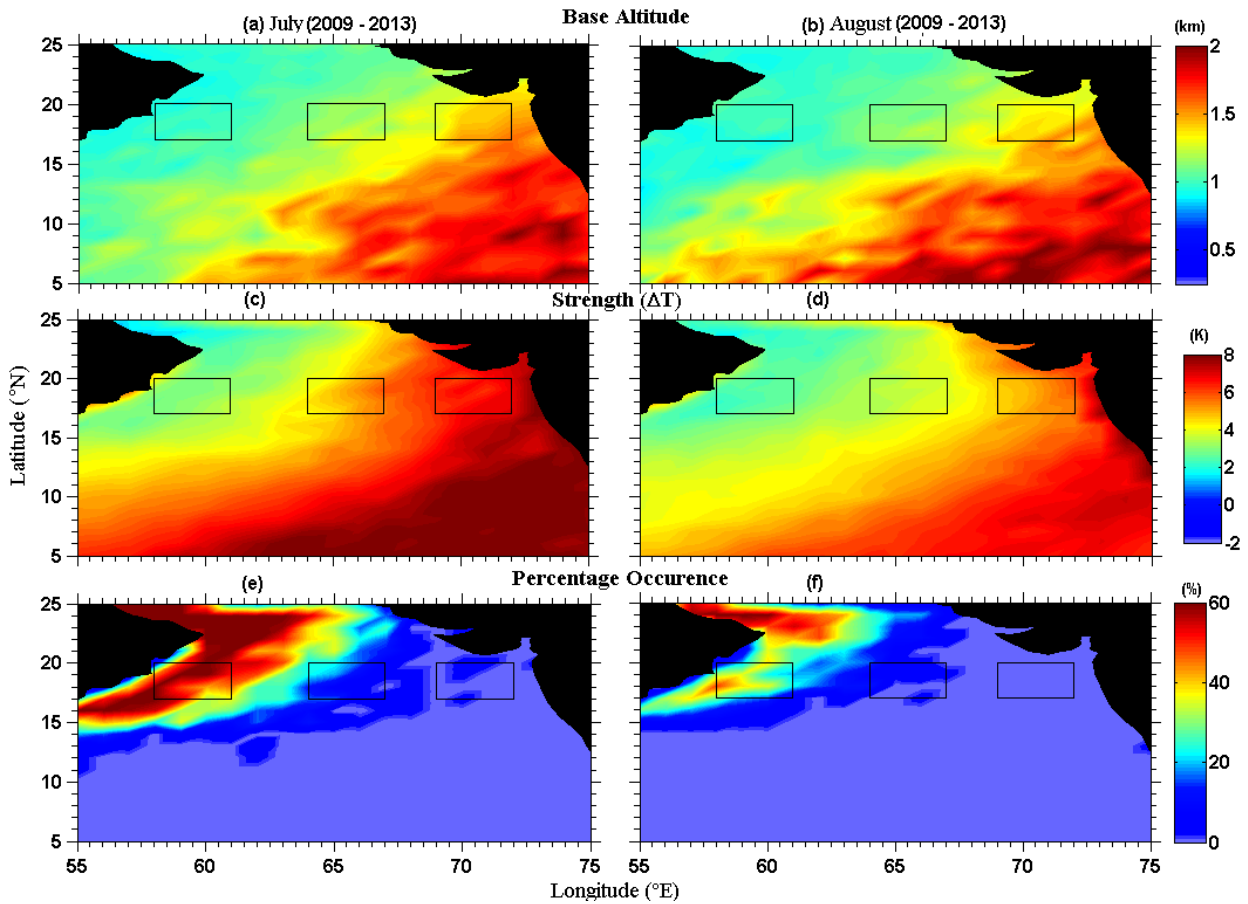
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646

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 648 2011 (blue). Green vertical lines are showing starting (01 May 2009) and ending (07 October 2009)
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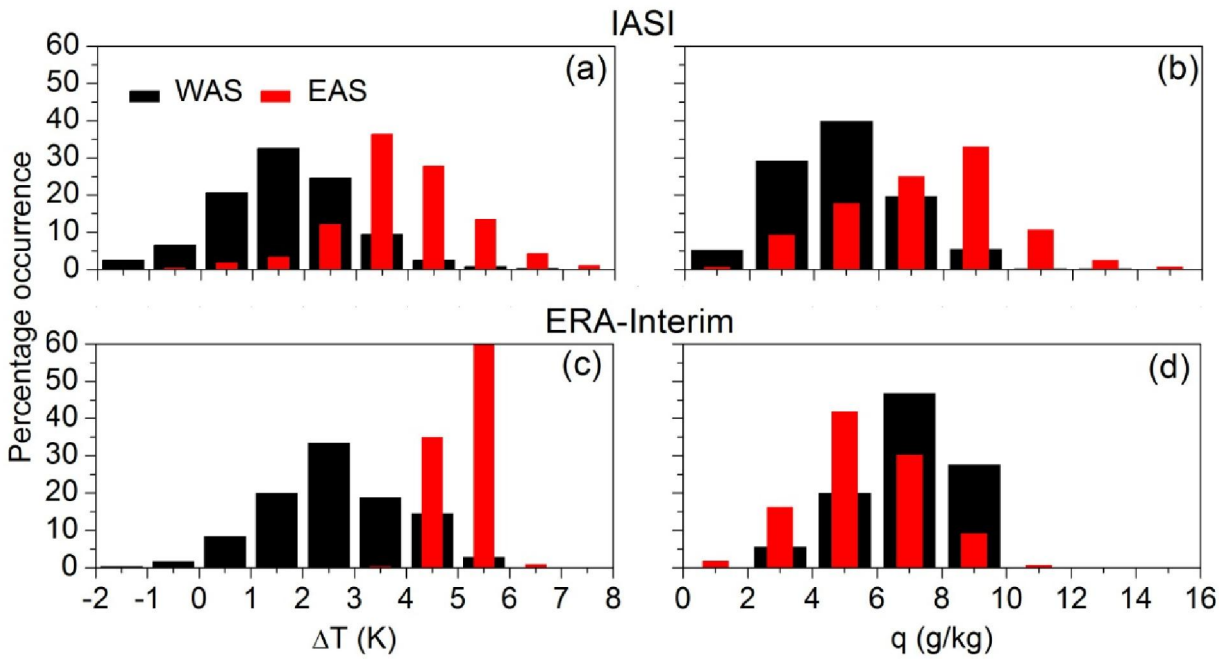
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666

667 **Figure 4.** Percentage occurrence of (a) ΔT and (b) q at 700 hPa observed in WAS and EAS during
 668 monsoon season of the years 2009-2013 for various ranges of ΔT and q at 700 hPa by IASI. (c) and
 669 (d) same as (a) and (b) but obtained from ERA-Interim data.

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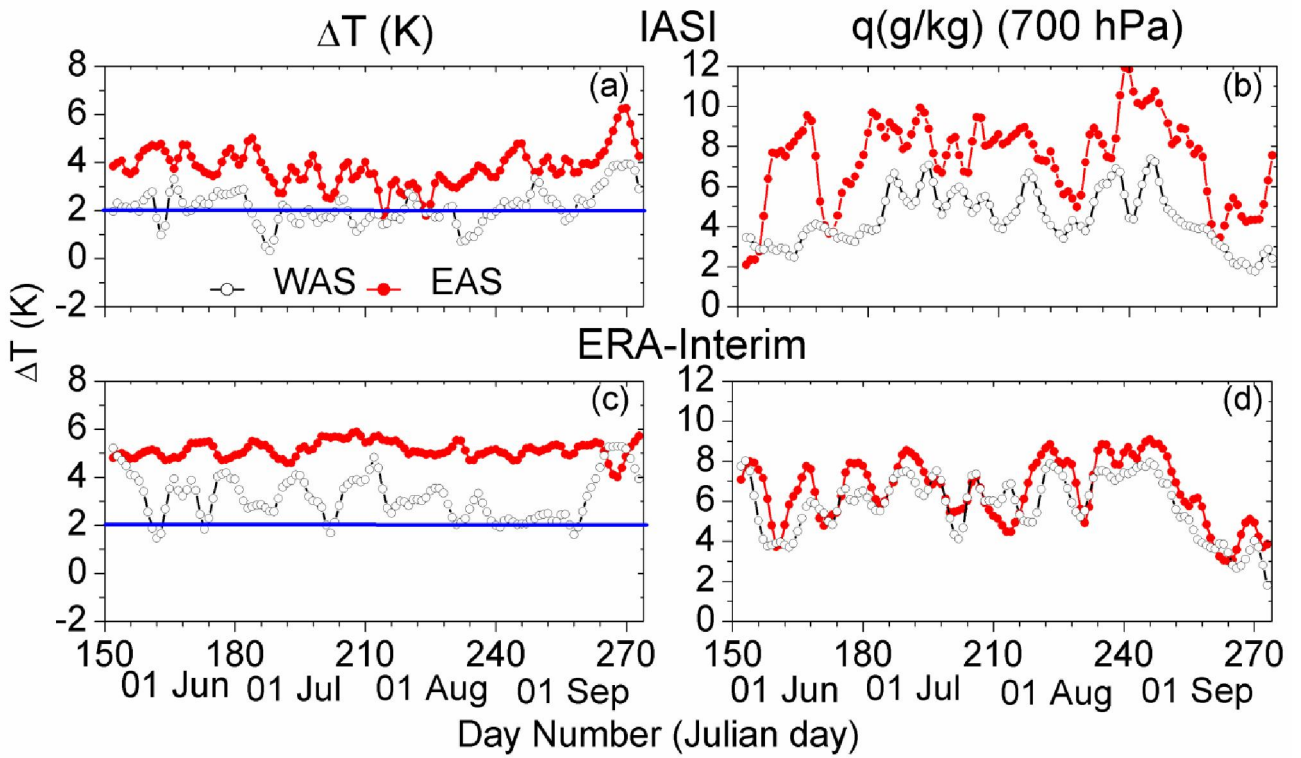
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681 **Figure 5.** Time series of (a) ΔT and (b) q at 700 hPa observed over WAS and EAS grid boxes
 682 during the monsoon season of the year 2012 by IASI, (c) and (d) same as (a) and (b) but obtained
 683 using ERA – Interim data. 3-point smoothed curves are shown.

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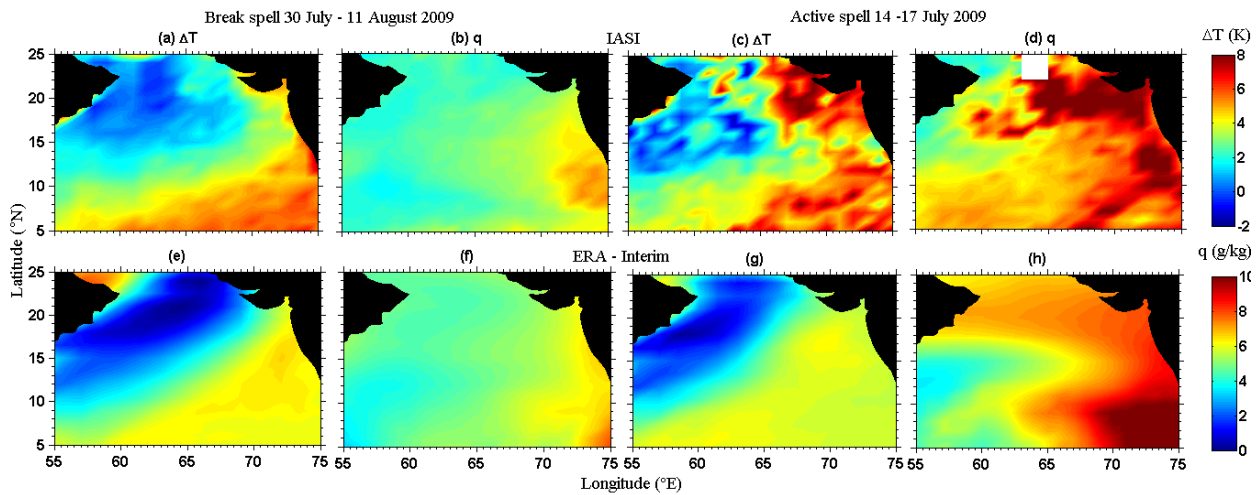
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 691 **Figure 6.** MI observed in (a) ΔT and (b) q at 700 hPa during break spells (30 July – 11 August
 692 2009) of the year 2009 by IASI, (c) and (d) same as (a) and (b) but observed during active spells
 693 (14-17 July 2009). (e) and (f) and (g) and (h), same as (a) and (b) and (c) and (d) but observed by
 694 ERA-Interim, respectively.

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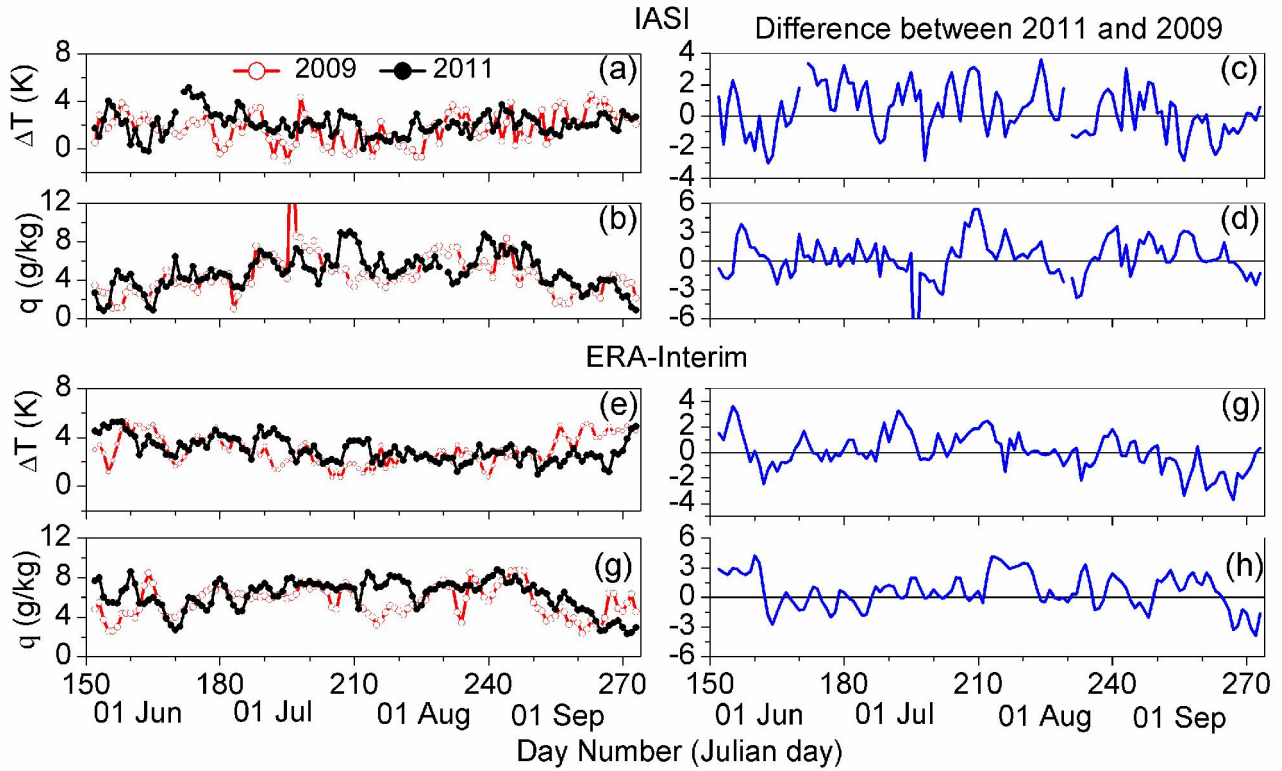
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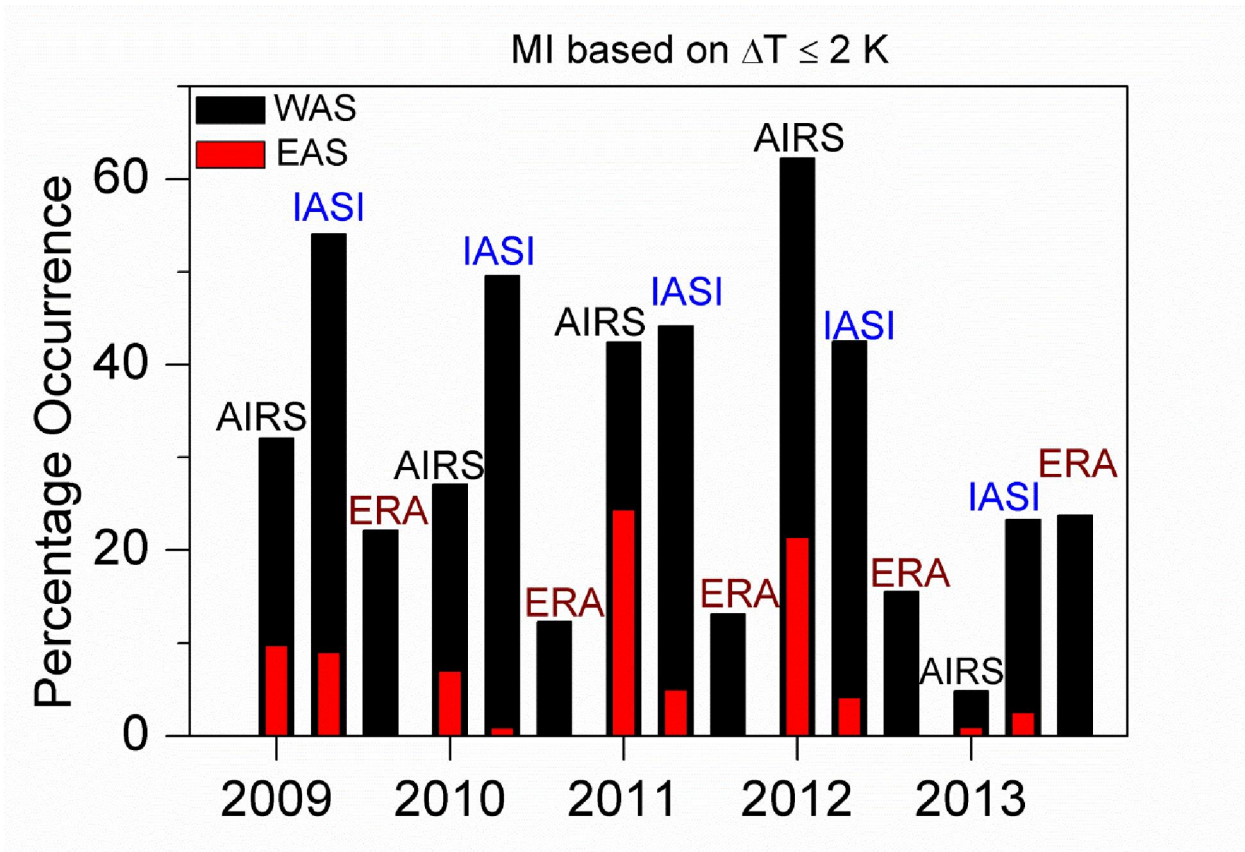
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Figure 7. Time variations of (a) ΔT and (b) q at 700 hPa observed over WAS during two contrasting years of 2009 and 2011 by using IASI measurements. Difference between 2011 and 2009 observed in (c) ΔT and (d) q at 700 hPa. (e) to (h) same as (a) to (d) but observed by using ERA-Interim data.



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Figure 8. Percentage occurrence of MI observed with (a) $\Delta T \leq 2K$ using IASI, AIRS and ERA-

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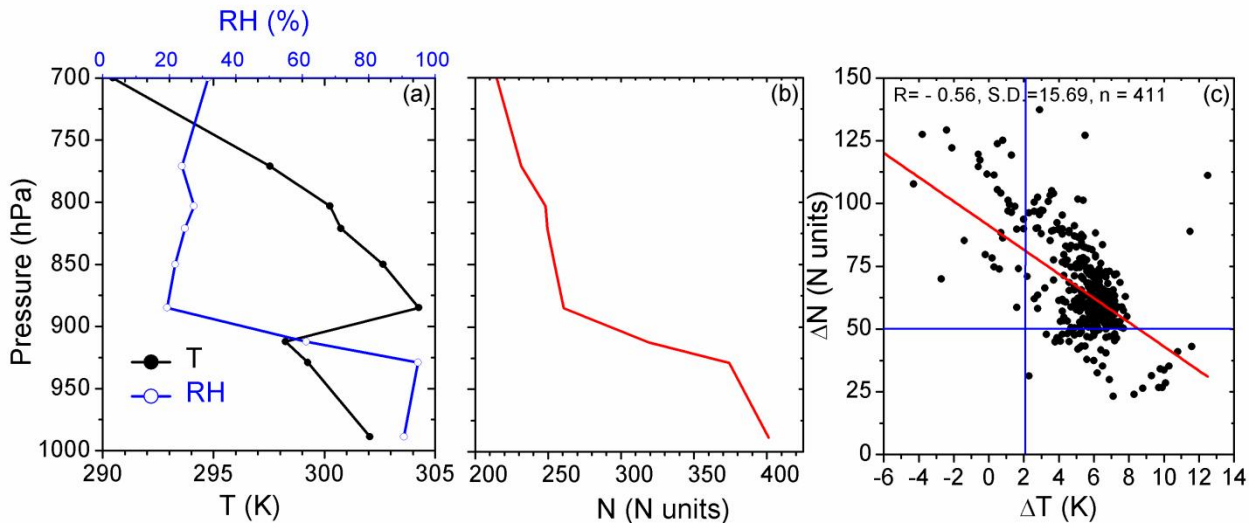
Interim data during monsoon seasons of 2009-2013 over WAS and EAS.

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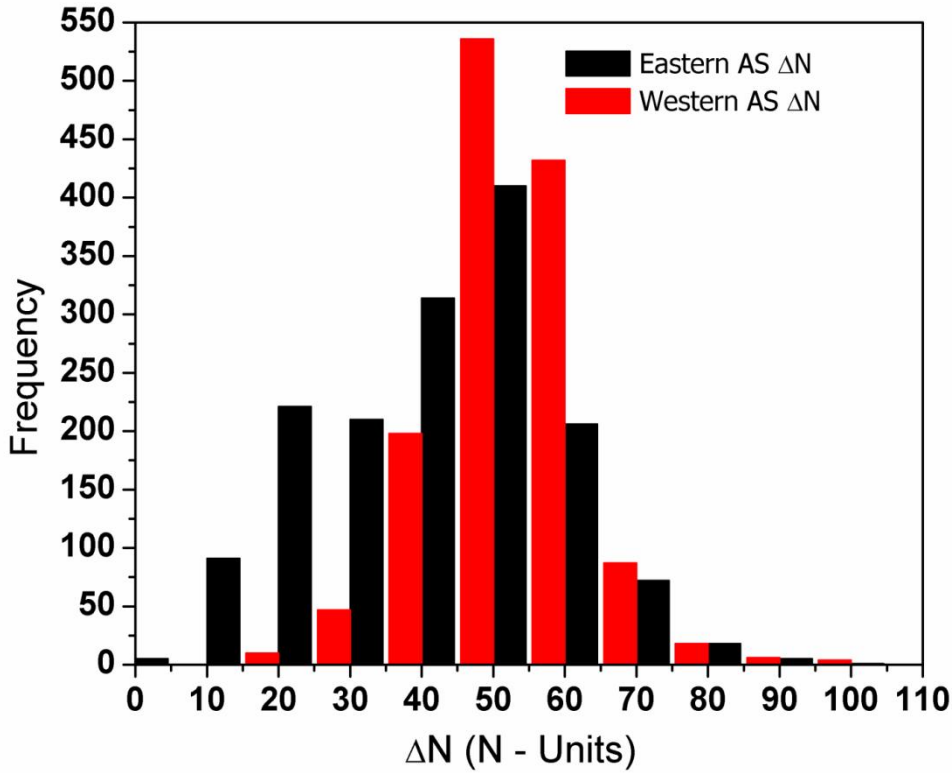
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Figure 9. Typical examples showing MI in temperature and RH on (a) 27 June 1979 at 0656 GMT at 20°N, 62°E obtained from dropsondes from MONEX experiment, (b) N profile (c) Scatter plot of ΔT and ΔN .

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Figure 10. Frequency of ΔN observed in Western AS and Eastern AS during monsoon season of the years 2009-2013 for various ranges of ΔN by COSMIC. Western AS is showing higher values means inversion is there.