1 Change in turbopause altitude at 52° and 70°N

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20 Abstract

21 The turbopause is the demarcation between atmospheric mixing by turbulence (below) and 22 molecular diffusion (above). When studying concentrations of trace species in the 23 atmosphere, and particularly long-term change, it may be important to understand processes 24 present, together with their temporal evolution that may be responsible for redistribution of 25 atmospheric constituents. The general region of transition between turbulent and molecular mixing coincides with the base of the ionosphere, the lower region in which molecular 26 27 oxygen is dissociated, and, at high latitude in summer, the coldest part of the whole 28 atmosphere.

This study updates previous reports of turbopause altitude, extending the time series by half a decade, and thus shedding new light on the nature of change over solar-cycle timescales. Assuming there is no trend in temperature, at 70°N there is evidence for a summer trend of ~1.6 km/decade, but for winter and at 52°N there is no significant evidence for change at all. If the temperature at 90 km is estimated using meteor trail data, it is possible to estimate a cooling rate, which, if applied to the turbopause altitude estimation, fails to alter the trend significantly irrespective of season.

The observed increase in turbopause height supports a hypothesis of corresponding negative trends in atomic oxygen density, [O]. This supports independent studies of atomic oxygen density, [O], using mid-latitude timeseries dating from 1975, which show negative trends since 2002.

41 Introduction

42 The upper mesosphere and lower thermosphere (UMLT) regime of the atmosphere exhibits a 43 number of features, the underlying physics of which are interlinked and, relative to processes 44 at other altitudes, little understood. At high latitude, the summer mesopause, around 85km is the coldest region in the entire atmosphere. The UMLT is, inter alia, characterized by the 45 46 base of the ionosphere, dissociation of molecular species (for example oxygen) by sunlight, and, the focus in this study, the transition from turbulent mixing to distribution of constituents 47 by molecular diffusion. The altitude at which transition turbulence-dominated mixing gives 48 49 way to molecular diffusion is known as the turbopause, and typically occurs around 100 km, 50 but displaying a seasonal variation, being lower in summer (e.g. ~95 km) and higher in winter 51 (e.g. ~110km) (Danilov et al., 1979). Many processes in the UMLT are superimposed and 52 linked. One example is where the mesopause temperature structure determines the altitude 53 dependence of breaking of upwardly propagating gravity waves (e.g. McIntyre, 1991) and 54 thus generation of turbulence. Indeed, the concept of a "wave turbopause" was proposed by 55 Offermann et al. (2007) and compared with the method used forthwith by Hall et al. (2008). 56 Prevailing winds filter or even inhibit propagation of gravity waves generated in the lower atmosphere and the static stability (or lack of it) of the atmosphere dictates the vertical 57 58 distribution of gravity wave saturation and breaking. The generation of turbulence and its 59 height distribution vary with season and similarly affect the turbopause altitude (e.g. Hall et 60 al., 1997). Turbulence is somewhat distributed through the high latitude winter mesosphere, 61 whereas in summer the gravity waves "save their energy" more until reaching the "steep beach" (a visualization attributable to M. E. McIntyre - private communication) of the 62 summer mesopause near 85km. Vertical transport by turbulent mixing and horizontal 63 64 transport by winds redistribute constituents such as atomic oxygen, hydroxyl and ozone.

Thus, long-term change in trace constituents cannot be fully explained in isolation fromstudies of corresponding change in temperature and neutral dynamics.

67 One means of locating the turbopause is to measure the concentration of particular species as 68 a function of height and noting where the constituents exhibit scale heights that depend on 69 their respective molecular weights, e.g. Danilov et al., (1979). Detection of turbulence and 70 estimation of its intensity is non-trivial because direct measurement by radar depends on turbulent structures being "visible" due to small discontinuities in refractive index, e.g. 71 72 Schlegel et al. (1978) and Briggs (1980). At 100km, this implies some degree of ionisation 73 and even in situ detectors often depend on ionisation as a tracer (e.g. Thrane et al. 1987). A 74 common means of quantifying turbulent intensity is the estimation of turbulent energy 75 dissipation rate, ε . In the classical visualization of turbulence in two dimensions, large 76 vortices generated by, for example breaking gravity waves or wind shears form progressively 77 smaller vortices (eddies) until inertia is insufficient to overcome viscous drag in the fluid. Viscosity then "removes" kinetic energy and transforms it to heat. This "cascade" from large-78 79 scale vortices to the smallest scale eddies capable of being supported by the fluid, and 80 subsequent dissipation of energy, was proposed by Kolmogorov (1941) but more accessibly 81 described by Batchelor (1953) and (e.g.) Kundu (1990). At the same time, a minimum rate of 82 energy dissipation by viscosity is supported by the atmosphere (defined subsequently). The 83 altitude at which these two energy dissipation rates are equal is also a definition of the 84 turbopause and corresponds to the condition where the Reynolds number, the ratio between 85 inertial and viscous forces, is unity.

The early work to estimate turbulent energy dissipation rates using medium frequency (MF) radar by Schlegel et al. (1978) and Briggs (1980) was adopted by Hall et al. (1998a). The reader is referred to these earlier publications for a full explanation, but in essence, velocity fluctuations relative to the background wind give rise to fading with time of echoes from

90 structures in electron density drifting through the radar beam. While the drift is determined by cross-correlation of signals from spaced receiver antennas, autocorrelation yields fading 91 92 times which may be interpreted as velocity fluctuations (the derivation of which is given in 93 the following section). The squares of the velocity perturbations can be equated to turbulent 94 kinetic energy and then, when divided by a characteristic timescale become energy 95 dissipation rates. Energy is conserved in the cascade to progressively smaller and more numerous eddies such that the energy dissipation rate is representative of the ultimate 96 97 conversion of kinetic energy to heat by viscosity. Hall et al. (1998b and 2008) subsequently 98 applied the turbulent intensity estimation to identification of the turbulent. The latter study, 99 which offers a detailed explanation of the analysis, compares methods and definitions and 100 represents the starting point for this study. In addition, Hocking (1983 and 1996) and 101 Vandepeer and Hocking (1993) offer a critique on assumptions and pitfalls pertaining to 102 observation of turbulence using radars. For the radars to obtain echoes from the UMLT, a 103 certain degree of ionization must be present and daylight conditions yield better results than 104 night-time, and similarly results are affected by solar cycle variation. However, there is a 105 trade-off: too little ionization prevents good echoes while too much gives rise to the problem 106 of group delay of the radar wave in the ionospheric D-region. Space weather effects that are 107 capable of creating significant ionization in the upper mesosphere are infrequent, and aurora 108 normally occur on occasional evenings at high latitude, and then only for a few hours 109 duration at the most. Of the substantial dataset used in this study, however, only a small 110 percentage of echo profiles are expected to be affected by auroral precipitation that would 111 cause problematic degrees of ionisation below the turbopause. While it must be accepted that group delay at the radar frequencies used for the observations reported here cannot be 112 113 dismissed, the MF-radar method is the only one that has been available for virtually uninterrupted measurement of turbulence in the UMLT region over the past decades. 114

Full descriptions of the radar systems providing the underlying data used here are to be found in Hall (2001) and Manson and Meek (1991) and the salient features of the radars, relevant for this study are given in Table 1.

118

119 Analysis methodology

120 The characteristic fading time of the signal, τ_c , is used to define an indication of the upper 121 limit for turbulent energy dissipation present in the atmosphere, ε' , as explained above. First, 122 velocity fluctuations, v' relative to the background wind are identified as:

123
$$\mathbf{v}' = \frac{\lambda \sqrt{\ln 2}}{4\pi \tau_c} \tag{1}$$

124 where λ is the radar wavelength. This relationship has been presented and discussed by 125 Briggs (1980) and Vandepeer and Hocking (1993). In turn v⁻² can be considered to represent 126 the turbulent kinetic energy of the air such that the rate of dissipation of this energy is 127 obtained by dividing by a characteristic timescale. If the Brunt-Väisälä period T_B (= $2\pi/\omega_B$ 128 where ω_B is the Brunt-Väisälä frequency in rad s⁻¹) can be a characteristic timescale, then it 129 has been proposed that:

130
$$\varepsilon' = 0.8 \mathrm{v}^{\prime 2} / T_B \tag{2}$$

131 the factor 0.8 being related to an assumption of a total velocity fluctuation (see Weinstock,

132 1978). Alternatively, this can be expressed as:

133
$$\varepsilon' = 0.8 \mathrm{v}^{\prime 2} \omega_{\scriptscriptstyle B} / 2\pi \tag{3}$$

134 wherein the Brunt-Väisälä frequency is given by

135
$$\omega_{B} = \sqrt{\left(\frac{dT}{dz} + \frac{g}{c_{p}}\right)\frac{g}{T}}$$
(4)

136 where *T* is the neutral temperature, *z* is altitude, *g* is the acceleration due to gravity and c_p is 137 the specific heat of the air at constant pressure. Due to viscosity, there is a minimum energy 138 dissipation rate, ε_{\min} , present in the atmosphere, given by

139
$$\varepsilon_{\min} = \omega_B^2 v / \beta$$
 (5)

140 where v is the kinematic viscosity. The factor β , known as the mixing or flux coefficient 141 (Oakey, 1982; Fukao et al., 1994; Pardyjac et al., 2002), is related to the flux Richarson 142 Number R_f ($\beta = R_f/(1-R_f)$). R_f is in turn related to the commonly used gradient Richardson number, Ri by the ratio of the momentum to thermal turbulent diffusivities, or turbulent 143 144 Prandtl number (e.g. Kundu 1990). Fukao et al. (1994) proposed 0.3 as a value for β . The 145 relationships are fully described by Hall et al. (2008). To use the MF radar system employed 146 here to estimate turbulence is not well suited to estimating *Ri* due to the height resolution of 147 3km; moreover more detailed temperature information would be required to arrive at R_{f} .

Anywhere in the atmosphere, energy dissipation is by the sum of the available processes. In this study, therefore, the turbulent energy dissipation rate can be considered the total rate minus that corresponding to viscosity:

151
$$\varepsilon = \varepsilon' - \varepsilon_{\min}$$
 (6)

152 Importantly, the kinematic viscosity is given by the dynamic viscosity, μ , divided by the 153 density, ρ :

$$154 \quad v = \mu/\rho \tag{7}$$

155 Thus, since density is inversely proportional to temperature, kinematic viscosity is 156 (approximately) linearly dependent on temperature; ω_B^2 is inversely proportional to 157 temperature and therefore ε_{min} is approximately independent of temperature. On the other 158 hand, ε' is proportional to ω_B and therefore inversely proportional to the square root of 159 temperature. 160 If we are able to estimate the energy dissipation rates described above, then the turbopause 161 may be identified as the altitude at which $\varepsilon = \varepsilon_{\min}$. This corresponds to equality of inertial and 162 viscous effects and hence the condition where Reynolds number, *Re*, is unity as explained 163 earlier.

164 To implement the above methodology, temperature data are required. Since observational temperature profiles cannot be obtained reliably, NRLMSISE-00 empirical model (Picone et 165 166 al., 2002) profiles are, of necessity, used in the derivation of turbulent intensity from MF-167 radar data. The reasons for this are discussed in detail in the following section. While a 168 temperature profile covering the UMLT region is not readily available by ground-based 169 observations from Tromsø, meteor-trail echo fading times measured by the Nippon/Norway 170 Tromsø Meteor Radar (NTMR) can be used to yield neutral temperatures at 90 km altitude. Any trend in temperature can usefully be obtained (the absolute values of the temperatures 171 172 being superfluous since they are only available for one height). The method is exactly the 173 same as used by Hall et al. (2012) to determine 90 km temperatures over Svalbard (78°N) using a radar identical to NTMR. Hall et al. (2005) investigate the unsuitability of meteor 174 175 radar data for temperature determination above ~95km and below ~85 km. In summary: 176 ionization trails from meteors are observed using a radar operating at a frequency less than 177 the plasma frequency of the electron density in the trail (this is the so-called "underdense" 178 condition). It is then possible to derive ambipolar diffusion coefficients D from the radar echo 179 decay times, τ_{meteor} (as distinct from the corresponding fading time for the medium-frequency 180 radars) according to:

181
$$\tau_{meteor} = \frac{\lambda^2}{16\pi^2 D}$$
(8)

182 wherein λ is the radar wavelength. Thereafter the temperature *T* may be derived using the 183 relation:

9 of 29

184
$$T = \sqrt{\frac{P \cdot D}{6.39 \times 10^{-2} K_0}}$$
(9)

185 where P is the neutral pressure and K_0 is the zero field mobility of the ions in the trail (here we assume $K_0 = 2.4 \times 10^{-4} \text{ m}^{-2} \text{ s}^{-1} \text{ V}^{-1}$ (McKinley, 1961; Chilson et al., 1996; Cervera and 186 187 Reid, 2000; Holdsworth et al., 2006). The pressure, P, was obtained from NRLMSISE-00 for 188 consistency with the turbulence calculations. In the derivations by Dyrland et al. (2010) and 189 Hall et al. (2012), for example, temperatures were then normalized to independent 190 measurements by the MLS (Microwave Limb Sounder) on board the EOS (Earth Observing 191 System) Aura spacecraft launched in 2004. The MLS measurements were chosen because the 192 diurnal coverage was constant for all measurements and it was therefore simpler to estimate 193 values that were representative of daily means, than other sources such as SABER. In this 194 way, the influence of any systematic deficiencies in NRLMSISE-00 (e.g. due to the age of the 195 model) were minimized.

196

197 Results and implications for changing neutral air temperature

198 Following the method described above and by Hall et al. (1998b and 2008), the turbopause 199 position is determined as shown in Fig. 1. The time and height resolutions of the MF radars 200 used for the investigation are 5 minutes and 3 km respectively, and daily means of turbulent 201 energy dissipation rate profiles are used to determine corresponding turbopause altitudes. The 202 Figure shows the evolution since 1999, 70°N, 19°E (Tromsø) in the upper panel and 52°N, 203 107°W (Saskatoon) in the lower panel. Results are, of course specific to these geographical 204 locations and it must be stressed that they are in no way zonally representative (hereafter, though, "70°N" and "52°N" may be used to refer to the two locations for convenience). Data 205 are available from 1 January 1999 to 25 June 2014 for Saskatoon but thereafter, technical 206 207 problems affected data quality. Data are shown from 1 January 1999 to 25 October 2015 for 208 Tromsø. The cyan background corresponding to the period 16 February 1999 to 16 October 209 2000 in the 70°N (Tromsø) panel indicates data available but using different experiment 210 parameters and thus 70°N data prior to 17 October 2000 are excluded from this analysis. A 211 30-day running mean is shown by the thick lines with the shading either side indicating the 212 standard deviation. The seasonal variation is clear to see, and for illustrative purposes, trend 213 lines have been fitted to June and December values together with hyperbolae showing the 214 95% confidence limits in the linear fits (Working and Hotelling, 1929); the seasonal 215 dependence of the trends is addressed in more detail subsequently. The months of June and 216 December are chosen simply because these correspond to the solstices and thus to avoid any 217 a priori conception of when one could anticipate the maxima and minima to be. It is evident 218 that, apart from the seasonal variation, the mid-latitude turbopause changes little over the 219 period 1999-2014, whereas at high latitude there is more change for the summer state over 220 the period 2001-2015 (the summers of 1999 and 2000 being excluded from the fitting due to 221 changes in experiment parameters for the Tromsø radar). To investigate the seasonal 222 dependence of the change further, the monthly values for 70°N and 52°N are shown in Fig. 2. 223 Since 2001, the high latitude turbopause has increased in height during late spring and midsummer but otherwise remained constant. Since individual months are selected the possibility 224 225 of "end-point" biases are not an issue in the trend-line fitting as would be the case if 226 analyzing entire datasets with non-integer numbers of years. Even so, certain years may be 227 apparently anomalous, for example the summer of 2003. In this study, the philosophy is to 228 look for any significant change in the atmosphere over the observational period. If anomalous 229 years are caused by, for example, changes in gravity-wave production (perhaps due to an increasing frequency of storm in the troposphere) and filtering in the underlying atmosphere, 230 231 these too should be considered part of climate change. The trend (or overall change) over the observation period is indeed sensitive to exclusion of certain years. Although not illustrated 232

233 here, this was tested briefly: selecting data from only 2004 onwards indicates no significant 234 change for summer, but a slightly increased negative winter change (to -1.7 ± 0.2 K decade 235 ¹); excluding only 2003 (the visually anomalous year) fails to alter the summer and winter 236 values significantly at all. The above findings represent an update of those by Hall et al. 237 (1998b and 2008), adding more years to the time series and therefore now covering a little 238 over one solar cycle (the latter half of cycle 23 and first half of 24). As for the preceding 239 papers and for consistency the neutral atmosphere parameters (temperature and density) 240 required have been obtained from the NRLMSISE-00 empirical model (Picone et al., 2002) 241 and have been assumed not to exhibit any trend over the observation period. In other terms, 242 one-year seasonal climatology temperature models at one-day resolution for 70°N and 52°N 243 and altitude range appropriate for the respective radars are therefore used for all years for 244 consistency with earlier results and for consistency between the two latitudes studied here. 245 Satellite-based temperature determinations are, of course available, including, for example SABER (Sounding of the Atmosphere by Broadband Emission Radiometry) on board 246 247 TIMED (Thermosphere Ionosphere Mesosphere Energetics and Dynamics) which was 248 launched in 2001. The temporal sampling by such instruments makes the estimation of (for 249 example) daily means somewhat complicated. Moreover, the measurements are not 250 necessarily representative for the field of field of view of the radar because the geographical 251 coverage of remote sensing data needs to be sufficiently large to obtain the required annual 252 coverage, since the sampling region can vary with season (depending on the satellite). Choice 253 of the somewhat dated NRLMSISE-00 model at least allows the geographical location to be 254 specified and furthermore ensures a degree of consistency between the two sets of radar observations and also earlier analyses. The only ground-based temperature observations both 255 256 available and suitable are at 70°N and 90 km altitude as described earlier and used 257 subsequently.

258 Next, we have attempted to investigate the effects of changing temperature. In a very simplistic approach, hypothetical altitude-invariant trends are imposed on the NRLMSISE-00 259 profiles. In other words, the same hypothetical trend is applied to all heights (for want of 260 better information) in the NRLMSISE-00 profile to generate evolving (cooling or warming) 261 temperature time-series. The suggested trends vary from -20Kdecade⁻¹ to +20Kdecade⁻¹, thus 262 263 well encompassing any realistically conceivable temperature change (c.f. Blum and Fricke, 264 2008; Danilov, 1997, Lübken, 1999). The result of applying hypothetical temperature trends to the time-invariant turbopause heights shown earlier is demonstrated in Figure 3. Given the 265 seasonal differences identified earlier, four combinations are shown: summer (average of 266 267 May, June and July) and winter (average of November, December and January) for each 268 geographic location. Realistic temperature trends can be considered within the range ± 6 Kdecade⁻¹ such that the only significant response of turbopause height to temperature trend is 269 270 for 70°N in summer. In addition, the figure includes estimated trends obtained from observations, which shall be explained forthwith. The salient point arising from the Figure is 271 272 that no realistic temperature trend (at least given the simple model employed here) has the 273 capability of reversing the corresponding trend in turbopause height.

274 In a recent study, Holmen et al. (2015) have built on the method of Hall et al. (2012) to determine 90 km temperatures over NTMR, as has been described in the previous section. 275 276 This new work presents more sophisticated approaches for normalisation to independent 277 measurements and investigating the dependence of derived temperatures on solar flux. 278 Having removed seasonal and solar cycle variations in order to facilitate trend-line fitting (as 279 opposed to isolating a hypothetical anthropogenic-driven variation), Holmen et al. (2015) arrive at a temperature trend of -3.6 ± 1.1 Kdecade⁻¹ determined over the time interval 2004-280 281 2014 inclusive. This can be considered statistically significant (viz. significantly non-zero at the 5% level) since the uncertainty ($2\sigma = 2.2 \text{ Kdecade}^{-1}$) is less than the trend itself (e.g. Tiau et al., 1990).

284 Estimation of changes in temperature corresponding to the period for determination of the turbopause were only viable for 70°N, these being -0.8 ± 2.9 Kdecade⁻¹ for summer and -8.1285 \pm 2.5 Kdecade⁻¹ for winter, and these results are indicated in Fig. 3. Again using the simple 286 287 idea of superimposing a gradual temperature change (the same for all heights) on the temperature model used for the turbulence determination thus fails to alter the change in 288 289 turbopause height significantly, for the ~decade of observations. Although direct temperature 290 measurements are not available for the 52°N site, Offermann et al, (2010) report cooling rates of ~2.3 K decade⁻¹ for 51°N, 7°E, and She et al. (2015) ~2.8 K decade⁻¹ for 42°N, 112°W. As 291 292 for 70°N, these results do not alter the conclusions inferred from Fig. 3.

293

294 **Discussion**

295 The aim of this study has been to update earlier reports (viz. Hall et al. 2008) of turbopause 296 altitude and change determined for two geographic locations: 70°N, 19°E (Tromsø) and 297 52°N, 107°W (Saskatoon). An effort has been made to demonstrate that conceivable 298 temperature trends are unable to alter the overall results, viz. that there is evidence of 299 increasing turbopause altitude at 70°N, 19°E in summer, but otherwise no significant change 300 during the period 2001 to 2014. Assimilating results from in situ experiments spanning the 301 time interval 1966-1992, Pokhunkov et al. (2009) present estimates of turbopause height 302 trends for several geographical locations, but during a period prior to that of our observations. 303 For high latitude the turbopause is reported to have fallen by ~2-4 km between 1968 and 1989 – the opposite sign of our finding for 2001-2014. More recently, further evidence has 304 305 been presented for a long-term descent of the turbopause, at least at mid-latitude (Oliver et al., 2014 and references therein). The rationale for this is that the atomic oxygen density [O] 306

307 has been observed to increase during the time interval 1975-2014 at a rate of approximately

308 1% year⁻¹. The associated change in turbopause height may be estimated thus:

$$309 \qquad H = RT / mg \tag{10}$$

where *H* is scale height, *R* is the universal gas constant (=8.314 J mol⁻¹ K⁻¹), *m* is the mean molecular mass (kg mol⁻¹) and *g* is the acceleration due to gravity. At 120 km altitude, *g* is taken to be 9.5 ms⁻². For air and atomic oxygen, m = 29 and 16 respectively. For a typical temperature of 200K, the two corresponding scale heights are therefore $H_{air} = 6.04$ km and $H_{oxygen} = 10.94$ km. If the change (fall) in turbopause height is denoted by Δh_{turb} , then Oliver et al. (2014) indicate that the factor by which [O] would increase is given by:

316
$$\exp(\Delta h_{turb}/H_{air})/\exp(\Delta h_{turb}/H_{oxygen})$$
 (11)

Note that Oliver et al. (2014) state that '[O] ... would increase by the amount', but, since Eq. 317 318 (11) is dimensionless, the reader should be aware this is a factor, not an absolute quantity. At 319 first, there would appear to be a fundamental difference between the findings derived from 320 [O] at a mid-latitude station and those for ε from a high-latitude station, and indeed the 321 paradox could be explained by either the respective methods and/or geographic locations. Usefully, in this context, Shinbori et al. (2014) and Kozubek et al. (2015) investigate such 322 323 geographical diversity. However if one examines the period from 2002 onwards 324 (corresponding to the high-latitude dataset, but only about one guarter of that from the mid-325 latitude station), a decrease in [O] corresponds with an increase in Δh_{turb} . If, then, Δh_{turb} . for the measured summer temperature change at high latitude (viz. 0.16 km year⁻¹ from Fig. 3) is 326 327 inserted in Eq. (11) together with the suggested scale heights for air and atomic oxygen, one obtains a corresponding decrease in [O] of 16% decade⁻¹, e.g. over the period 2002-2015. The 328 corresponding time interval is not analysed per se by Oliver et al (2014) but a visual 329 330 inspection suggests a decrease of the order of 20%; the decrease itself is incontrovertible and 331 therefore in qualitative agreement with our high-latitude result.

332 It is somewhat unfortunate that it is difficult to locate simultaneous and approximately co-333 located measurements by different methods. The turbopause height-change derived by Oliver 334 et al. (2014) are by measurements of [O] and at mid-latitude; those by Pokhunkov et al. 335 (2009), also by examining constituent scale-heights, include determinations for Heiss Island (80°N, 58°E) but this rocket sounding programme was terminated prior to the start of our 336 337 observation series (Danilov et al., 1979). It should be noted, however that the results of seasonal variability presented by Danilov et al. (1979) agree well with those described here 338 339 giving credence to the method and to the validity of the comparisons above.

340 Finally, the change in turbopause altitude during the last decade or more should be placed in 341 the context of other observations. The terrestrial climate is primarily driven by solar forcing, 342 but several solar cycles of data would be required to evaluate the effects of long-term change 343 in space weather conditions on turbulence in the upper atmosphere. A number of case-studies 344 have been reported, however that indicate how space weather events affect the middle atmosphere (Jackman et al., 2005; Krivolutsky et al., 2006). One recurring mechanism is 345 346 forced change in stratospheric chemistry (in particular, destruction and production of ozone 347 and hydroxyl); the associated perturbations in temperature structure adjust the static stability 348 of the atmosphere through which gravity waves propagate before reaching the mesosphere. In 349 addition, greenhouse gases causing global warming in the troposphere act as refrigerants in 350 the middle atmosphere and so changing the static stability and therefore the degree to which 351 gravity waves shed turbulence en route to the UMLT. Not a subject of this study, it is 352 hypothesised that changes in the troposphere and oceans give rise to a higher frequency of violent weather; this in turn could be expected to increase the overall gravity wave activity 353 354 originating in the lower atmosphere but propagating through the middle atmosphere. Sudden 355 stratospheric warmings (SSWs) also affect (by definition) the vertical temperature structure 356 and thus gravity wave propagation (e.g. de Wit et al., 2015; Cullens et al., 2015). Apart from 357 direct enhancements of stratospheric temperatures, SSWs have been demonstrated to affect 358 planetary wave activity even extending into the opposite hemisphere (e.g. Stray et al., 2015). 359 If such effects were capable of, for example, triggering the springtime breakdown of the polar 360 vortex, associated horizontal transport of stratospheric ozone contributes to determination of the tropopause altitude (e.g. Hall, 2013) and again, gravity wave propagation. Overall change 361 362 in the stratosphere is proposed as the origin of the observed strengthening of the Brewer Dobson circulation during the last 35 years at least (Fu et al., 2015). Closer to the 70°N, 19°E 363 (Tromsø) observations, Hoffmann et al. (2011) report increases in gravity wave activity at 364 55°N, 13°E during summer, including at 88km. Although not co-located, the increasing 365 366 gravity wave flux, with waves breaking at the summer high latitude mesopause would 367 similarly increase turbulence intensity and support the change reported here. Further 368 references to long-term change in the middle and upper atmosphere in general can be found 369 in Cnossen et al. (2015). Background winds and superimposed tides thus affecting gravity 370 wave propagation and filtering in the atmosphere underlying the UMLT also vary from 371 location to location at high latitude and the two studies by Manson et al. (2011a and 2011b) 372 study this zonal difference and compare with a current model. Although for approximately 373 10° further north than the Tromsø radar site, these studies give valuable background 374 information, on not only the wind field, but also on tidal amplitude perturbation due to 375 deposition of gravity waves' horizontal momentum.

376

377 Conclusion

Updated temporal evolutions of the turbopause altitude have been presented for two locations: 70°N, 19°E (Tromsø) and 52°N, 107°W (Saskatoon), the time interval now spanning 1999 to 2015. These turbopause altitude estimates are derived from estimates of turbulent energy dissipation rate obtained from medium-frequency radars. The method entails 382 a knowledge of neutral temperature that had earlier (Hall et al., 2008) been assumed to be 383 constant with time. Here the response of the change in turbopause heights over the period of the study to temperature trends - both hypothetical and observed - is examined. No 384 385 temperature trend scenario was capable of altering the observed turbopause characteristics 386 significantly; at 70°N, 19°E an increase in turbopause height is evident during the 1999-2015 387 period for summer months, whereas for winter at 70°N, 19°E and all seasons at 52°N, 107°W the turbopause height has not changed significantly. In evaluating these results, however, 388 389 there are a number of caveats that must be remembered. Firstly, the radar system does not 390 perform well with an aurorally disturbed D-region – the study, on the other hand incorporates 391 well over 100,000 hours of data for each radar site, and auroral conditions are occasional and 392 of the order of a few hours each week at most. Secondly, an influence of the semi-empirical 393 model used to provide both density and Brunt-Väisälä frequencies cannot be disregarded. It 394 should also be stressed that a change is being reported for the observational periods of 395 approximately 15 years (i.e. just over one solar cycle) and parameterized by fitting linear 396 trend-lines to the data; this is distinct from asserting long-term trends in which solar and 397 anthropogenic effects can be discriminated.

398 At first, this conclusion would appear to contradict the recent report by Oliver et al. (2014) 399 and Pokhunkov et al. (2009), however, closer inspection shows that if one considers the time 400 interval 2002-2012 in isolation, there is a qualitative agreement. In fact, we note that Oliver et 401 al. (2014) deduce a turbopause change based on changing atomic oxygen concentration and 402 so we are similarly able to deduce a change in atomic oxygen concentration based on the 403 change in turbopause height obtained from direct estimation of turbulence intensity. Given an average (i.e. not differentiating between seasons) temperature change of -3.4 ± 0.5 K decade⁻¹ 404 405 for 70°N, 19°E (Tromsø), the change in turbopause height in summer over the same time interval is 1.6 ± 0.3 km decade⁻¹ suggesting a decrease in atomic oxygen concentration of 16%. 406

407 The primary result of this study is to demonstrate the increasing altitude of the summer 408 turbopause at 70°N, 19°E and the apparently unvarying altitude in winter and at 52°N, 409 107°W during the time interval 1999-2014. Independent studies using a radically different method demonstrate how to infer a corresponding decrease in atomic oxygen concentration, 410 411 as a spin-off result. Finally, the question as to the exact mechanism causing the evolution of turbulence in the lower thermosphere at, in particular 70°N, 19°E, remains unanswered, and 412 413 furthermore, dynamics at this particular geographic location may be pathological. The 414 solution perhaps lies in seasonally dependent gravity wave filtering in the underlying 415 atmosphere being affected by climatic tropospheric warming and/or middle atmosphere 416 cooling; hitherto, however, this remains a hypothesis.

417

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419 The authors thank the referees of this paper.

420

423 Table 1. Salient radar parameters

Parameter	Tromsø	Saskatoon
Geographic coordinates	69.58°N, 19.22°E	52.21°N, 107.11°E
Operating frequency	2.78 MHz	2.22 MHz
Pulse length	20 µs	20 µs
Pulse repetition frequency	100 Hz	60 Hz
Power (peak)	50 kW	25 kW
Antenna beamwidth	17° at -3dB	17° at -6dB
Altitude resolution	3 km	3 km
Time resolution (post-analysis	5 min	5 min

424

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Figure 1. Turbopause altitude as determined by the definition and method described in this paper. The thick solid line shows the 30-day running mean and the shading behind it the corresponding standard deviations. The straight lines show the fits to summer and winter portions of the curve. Upper panel: 70°N (Tromsø); lower panel: 52°N (Saskatoon). The cyan background in the 70°N panel indicates data available but unused here due to different experiment parameters



Figure 2. Trends for the period as a function of month. Upper panel: 70°N (Tromsø); lower
panel: 52°N (Saskatoon).



Figure 3. Response of turbopause trend line to different upper-mesosphere/lowerthermosphere temperature trends. Hypothetical trends range from an unrealistic cooling of 20K/decade to a similarly unrealistic warming. Top-left: 70°N summer (average of May, June and July); top-right: 70°N winter (average of November, December and January); bottomleft: 52°N summer; bottom-right: 52°N winter. Observed values for 70°N are also identified on the upper panels (dashed vertical lines) together with uncertainties (shading).