

1 **The contribution of oceanic halocarbons to marine and**
2 **free troposphere air over the tropical West Pacific**

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12
13 ***Response to Editor***

14
15 *Editor:*

16 *Comments to the Author:*

17 *Dear authors,*

18 *as discussed in our separate email discussion and according to your question in the answer*
19 *to my latest review, you wish to resubmit another revised manuscript. I am calling this a*
20 *minor revision, since I need to review it again.*

21 *kind regards*

22 *Michael*

23
24 **Authors:**

25 We thank the editor that he agreed with our suggestions in the reply to the previous editor
26 report. Below you find a list with relevant changes and the revised manuscript including the
27 improved method to compute free troposphere VSLs mixing ratios.

28

29 **Relevant changes in the manuscript**

30

- 31 • Adjusted signs in Eq. 3 – 6 and Table 3 to simplify and visualise the mass balance
- 32 equations (source / loss processes) and to be consistent with the current review of
- 33 Fuhlbrügge et al. 2015. This didn't affect the results.
- 34 • Added dependencies of variables in Eq. 11 and 12
- 35 • Replaced Eq.13 with the new method.
- 36 • Removed box approach and Figure 8.
- 37 • Updated results in 5.3.1 – 5.3.3, 6. according to the new method (Eq. 13).
- 38 • Updated former Figure 10, now 9 without box approach.
- 39 • Updated former Figure 11, now 10 using new method (Eq. 13).

40

41

42 **References**

43 Fuhlbrügge, S., Quack, B., Atlas, E., Fiehn, A., Hepach, H., and Krüger, K.: Meteorological
44 constraints on oceanic halocarbons above the Peruvian Upwelling, Atmos. Chem. Phys.
45 Discuss., 15, 20597–20628, 10.5194/acpd-15-20597-2015, 2015.

The contribution of oceanic halocarbons to marine and free troposphere air over the tropical West Pacific

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Abstract

Emissions of halogenated very short lived substances (VSLs) from the oceans contribute to the atmospheric halogen budget and affect tropospheric and stratospheric ozone. Here, we investigate the contribution of natural oceanic VSLs emissions to the Marine Atmospheric Boundary Layer (MABL) and their transport into the Free Troposphere (FT) over the tropical West Pacific. The study concentrates on bromoform, dibromomethane and methyl iodide measured on ship and air craft during the SHIVA (Stratospheric Ozone: Halogen Impacts in a Varying Atmosphere) campaign in the South China and Sulu Seas in November 2011. Elevated oceanic concentrations for bromoform, dibromomethane and methyl iodide of on average 19.9, 5.0 and 3.8 pmol L⁻¹ in particular close to Singapore and at the coast of Borneo with high corresponding oceanic emissions of 1486, 405 and 433 pmol m⁻² h⁻¹, respectively, characterize this tropical region as a strong source of these compounds. Atmospheric mixing ratios in the MABL were unexpectedly relatively low with 2.08, 1.17 and 0.39 ppt for bromoform, dibromomethane and methyl iodide. We use meteorological and chemical ship and aircraft observations, FLEXPART trajectory calculations and source-loss estimates to identify the oceanic VSLs contribution to the MABL and to the FT. Our results show that the well-ventilated MABL and intense convection led to the low atmospheric mixing ratios in the MABL despite the high oceanic emissions. Up to 45 % of the accumulated bromoform in the FT above the region originates from the local South China Sea area, while dibromomethane is largely advected from distant source regions and the local ocean only contributes 20 %. The accumulated methyl iodide in the FT is higher than can be explained with local

34 contributions. Possible reasons, uncertainties and consequences of our observations and
35 model estimates are discussed.

36

37 **1. Introduction**

38 Halogens play an important role for atmospheric chemical processes. Chlorine, bromine and
39 iodine radicals destroy ozone in the stratosphere (e.g. Solomon, 1999) and also affect
40 tropospheric chemistry (e.g. Saiz-Lopez and von Glasow, 2012). Halogens are released
41 following the photochemical breakdown of organic anthropogenic and natural trace gases. A
42 large number of very short lived brominated and iodinated organic substances, originating
43 from macro algae, seaweed, phytoplankton and other marine biota, are emitted from tropical
44 oceans and coastal regions to the atmosphere (Gschwend et al., 1985; Carpenter and Liss,
45 2000; Quack and Wallace, 2003; Quack et al., 2007; Liu et al., 2013). In particular, marine
46 emissions of bromoform (CHBr_3), dibromomethane (CH_2Br_2) and methyl iodide (CH_3I) are
47 major contributors of bromine and iodine to the atmosphere (Montzka and Reimann, 2011).
48 Annually averaged mean tropical lifetimes of these halogenated very short-lived substances
49 (VSLs) in the boundary layer are 15 (range: 13 – 17) days for CHBr_3 , 94 (84 – 114) days for
50 CH_2Br_2 and 4 (3.8 – 4.3) days for CH_3I . The mean tropospheric lifetimes of these compounds
51 at 10 km height are 17 (16 – 18) days, 150 (144 – 155) days, respectively 3.5 (3.4 – 3.6) days
52 (Carpenter et al., 2014). Climate change could strongly affect marine biota and thereby
53 halogen sources and the oceanic emission strength (Hughes et al., 2012; Leedham et al.,
54 2013; Hepach et al., 2014).

55 Aircraft measurements from Dix et al. (2013) suggest that the halogen-driven ozone loss in
56 the Free Troposphere (FT) is currently underestimated. In particular, elevated amounts of the
57 iodine oxide free radical (IO) in the FT over the Central Pacific indicate that iodine may have
58 a larger effect on the FT ozone budget than currently estimated by chemical models.
59 Coinciding with this study, Tegtmeier et al. (2013) projected a higher CH_3I delivery to the
60 Upper Troposphere / Lower Stratosphere (UTLS) over the tropical West Pacific than
61 previously reported, using an observation based emission climatology by Ziska et al. (2013).
62 Recent studies reported significant contributions of bromine and iodine to the total rate of
63 tropospheric and stratospheric ozone loss (e.g. von Glasow et al., 2004; Yang et al.,
64 2005; Saiz-Lopez et al., 2014; Yang et al., 2014; Hossaini et al., 2015). Deep tropical
65 convective events (Aschmann et al., 2011; Tegtmeier et al., 2013; Carpenter et al., 2014) as
66 well as tropical cyclones, i.e. typhoons (Tegtmeier et al., 2012), are projected to transport
67 VSLs rapidly from the ocean surface to the upper tropical tropopause layer. Despite the

68 importance of halogens on tropospheric and stratospheric ozone chemistry, halogen sources
69 and transport ways are still not fully understood. While the tropical West Pacific comprises
70 strong VLS source regions (Krüger and Quack, 2013), only low mean atmospheric mixing
71 ratios were observed during ship campaigns in 1994 and 2009 (Yokouchi et al., 1997; Quack
72 and Suess, 1999) and in 2010 (Quack et al., 2011; Brinckmann et al., 2012). None of these
73 previous studies investigated the contribution of oceanic VLS emissions to the marine
74 atmospheric boundary layer (MABL) and to the FT in this hot spot region with large oceanic
75 sources and strong convective activity.

76 The SHIVA ('Stratospheric Ozone: Halogen Impacts in a Varying Atmosphere') ship, aircraft
77 and ground-based campaign during November and December 2011 in the Southern South
78 China and Sulu Seas investigated oceanic emission strengths of marine VLS, as well as their
79 atmospheric transport and chemical transformation from the ocean surface to the upper
80 troposphere. For more details about the SHIVA campaign see the ACP special issue
81 (http://www.atmos-chem-phys.net/special_issue306.html).

82 In this study, we present campaign data from the research vessel (R/V) SONNE and the
83 research aircraft (R/A) FALCON. We identify the contribution of oceanic emissions to the
84 MABL and their exchange into the FT applying in-situ observations, trajectory calculations
85 and source-loss estimates. The results are crucial for a better process understanding and for
86 chemical transport model validation (Hossaini et al., 2013; Aschmann and Sinnhuber, 2013).
87 An overview of the data and the methods used in this study is given in Chapter 2. Chapter 3
88 provides results from the meteorological observations along the cruise. Chapter 4 compares
89 atmospheric VLS measurements derived on R/V SONNE and R/A FALCON. The
90 contribution of the oceanic emissions to the MABL and FT air is investigated and discussed
91 in Chapter 5. Finally, a summary of the results is given in Chapter 6.

92

93 **2. Data and Methods**

94 **2.1. Ship and aircraft campaigns**

95 The R/V SONNE cruise started on November 15, 2011 in Singapore and ended on November
96 29, 2011 in Manila, Philippines (Figure 1). The ship crossed the southwestern South China
97 Sea towards the northwestern coast of Borneo from November 16 – 19, 2011. From
98 November 19 - 23, 2011 the ship headed northeast along the northern coast of Borneo
99 towards the Sulu Sea. Two diurnal stations took place on November 18, 2011 at 2.4° N /
100 110.6° E and on November 22, 2011 at 6.0° N / 114.8° E. Two meetings between ship and
101 aircraft were carried out on November 19 and 21, 2011, where R/A FALCON passed R/V

102 SONNE within a distance of about 100 m several times to simultaneously measure the same
103 air masses. On November 24, 2011 the ship entered the Sulu Sea, and after 4 days transect,
104 R/V SONNE reached the Philippine coast.

105 16 measurement flights were carried out with R/A FALCON between November 16 and
106 December 11, 2011 as part of the SHIVA campaign to investigate halogenated VSLs from
107 the surface up to 13 km altitude over the South China and Sulu Seas. Observations were
108 performed between 1° N and 8° N, as well as 100° E and 122° E, from Miri, Borneo
109 (Malaysia) as the aircraft base. A detailed description of the VSLs measurements and flight
110 tracks can be found in Sala et al. (2014).

111

112 **2.2. Meteorological observations during SHIVA**

113 **2.2.1. Measurements on board R/V SONNE**

114 Meteorological parameters (temperature, air pressure, humidity and wind) were recorded at
115 20 m height every second. A 10 minute running mean of this data is used for this study. An
116 optical disdrometer ('ODM-470') measured the amount and intensity of precipitation during
117 the cruise at 15 m height every minute (see Supplementary Material for further details). To
118 obtain atmospheric profiles of air temperature, relative humidity and wind from the surface to
119 the stratosphere 67 GRAW DFM-09 and 6 GRAW DFM-97 radiosondes were launched
120 every 6 hours at standard UTC times (0, 6, 12, 18) from the working deck of R/V SONNE at
121 about 2 m above sea level. At the 24 h stations, the launch frequency was increased to 2 – 3
122 hours to analyse short term diel variations of the atmospheric boundary layer. The radiosonde
123 data was integrated in near real time into the Global Telecommunication System (GTS) to
124 improve meteorological reanalyses such as ERA-Interim (Dee et al., 2011), which is used as
125 input data for the trajectory calculations (Section 5).

126

127 **2.2.2. Marine atmospheric boundary layer**

128 The MABL is the atmospheric surface layer above the ocean in which trace gas emissions are
129 mixed vertically by convection and turbulence on a short time scale of about an hour (Stull,
130 1988;Seibert et al., 2000). The upper boundary of the MABL is either indicated by a stable
131 layer e.g. a temperature inversion or by a significant reduction in air moisture. Determination
132 of the MABL height can be achieved by theoretical approaches, e.g. using critical Bulk
133 Richardson number (Troen and Mahrt, 1986;Vogelezang and Holtslag, 1996;Sorensen, 1998)
134 or by practical approaches summarized in Seibert et al. (2000). An increase with height of the
135 virtual potential temperature, the temperature an air parcel would acquire if adiabatically

136 brought to standard surface pressure with regard to the humidity of the air parcel, identifies
137 the base of the stable layer, which is typically found between 100 m and 3 km altitude (Stull,
138 1988). In this study, we use the height of the base of the stable layer increased by half of the
139 stable layer depth as the definition for the MABL height. The height of the MABL is
140 determined from the atmospheric profiles measured by radiosondes launched on board the
141 ship, as described in detail by Fuhlbrügge et al. (2013).

142

143 **2.3. VLSL measurements and flux calculation**

144 VLSL in marine surface air and sea water were sampled synchronously on R/V SONNE
145 along the cruise track. From these data the oceanic emissions of the compounds during the
146 SHIVA campaign were calculated (Section 2.3.3). Additionally, VLSL were measured in the
147 MABL and the FT by R/A FALCON (Sala et al., 2014; Tegtmeier et al., 2013).

148

149 **2.3.1. Atmospheric samples**

150 Air samples were taken 3 hourly along the cruise track, and 1 – 2 hourly during the 24 hour
151 stations on R/V SONNE resulting in a total of 195 samples during the cruise. The air was
152 pressurized to 2 atm in pre-cleaned stainless steel canisters with a metal bellows pump. The
153 samples were analyzed within 6 months after the cruise at the Rosenstiel School for Marine
154 and Atmospheric Sciences (RSMAS, Miami, Florida) according to Schauflier et al. (1999)
155 with an instrumental precision of ~5 %. Further details of the analysis are described in
156 Montzka et al. (2003) and Fuhlbrügge et al. (2013). On R/A FALCON ambient air was
157 analysed in situ by a GhOST-MS (Gas Chromatograph for the Observation of Stratospheric
158 Tracers – coupled with a Mass Spectrometer) by the Goethe University of Frankfurt (GUF).
159 Additionally 700 ml glass flasks were filled with ambient air to a pressure of 2.5 bar with the
160 R/A FALCON whole air sampler (WASP) and analysed within 48 hours by a ground-based
161 gas chromatography – mass spectrometry (GC/MS) instrument (Agilent 6973) of the
162 University of East Anglia (Worton et al., 2008). During the flights GhOST measurements
163 were conducted approximately every 5 minutes with a sampling time of 1 minute, while
164 WASP samples were taken every 3 – 15 minutes with a sampling time of 2 minutes. Further
165 details on the instrumental precision and intercalibration on R/A FALCON are given in Sala
166 et al. (2014). Given that the ground-based GC/MS investigated only brominated compounds,
167 CH₃I data is not available from WASP. Measurements from R/V SONNE and R/A FALCON
168 were both calibrated with NOAA standards.

169

170 **2.3.2. Water samples**

171 VLSL sea water samples were taken 3 hourly from the moon pool of R/V SONNE at a depth
172 of 5 m from a continuously working water pump. Measurements were interrupted between
173 November 16, 00 UTC to November 17, 2011 12 UTC due to permission issues in the
174 southwest South China Sea. The water samples were analysed on board with a purge and trap
175 system, attached to a gas chromatograph with mass spectrometric detection in single-ion
176 mode and a precision of 10 % determined from duplicates. The method is described in detail
177 by Hepach et al. (2014).

178

179 **2.3.3. Sea – air flux**

180 The sea – air flux (F) of CHBr_3 , CH_2Br_2 and CH_3I is calculated with k_w the concentration
181 gradient, and Δc the concentration gradient between the water and atmospheric equilibrium
182 concentrations (Eq. 1). For the determination of k_w , the wind speed-based parameterization
183 of Nightingale et al. (2000) was used and a Schmidt number (Sc) correction to the carbon
184 dioxide derived transfer coefficient k_{CO_2} after Quack and Wallace (2003) was applied for the
185 three gases (Eq. 2).

186

$$F = k_w \cdot \Delta c \quad (\text{Eq. 1})$$

187

$$k_w = k_{CO_2} \cdot \frac{Sc^{-\frac{1}{2}}}{600} \quad (\text{Eq. 2})$$

188

189 Details on measuring the air – sea concentration gradient are further described in Hepach et
190 al. (2014) and references therein.

191

192 **2.4. Oceanic VLSL contribution to the MABL and FT**

193 **2.4.1. Trajectory calculations**

194 The air mass transport from the surface to the FT was calculated with the Lagrangian Particle
195 Dispersion Model FLEXPART from the Norwegian Institute for Air Research in the
196 Department of Atmospheric and Climate Research (Stohl et al., 2005). The model has been
197 extensively evaluated in earlier studies (Stohl et al., 1998;Stohl and Trickl, 1999) and
198 includes parameterizations for turbulence in the atmospheric boundary layer and the FT as
199 well as moist convection (Stohl and Thomson, 1999;Forster et al., 2007). Meteorological

200 input fields are retrieved from the ECMWF (European Centre for Medium-Range Weather
 201 Forecasts) assimilation reanalysis product ERA-Interim (Dee et al., 2011) with a horizontal
 202 resolution of $1^\circ \times 1^\circ$ and 60 vertical model levels. The ship-based 6 hourly radiosonde
 203 measurements were assimilated into the ERA-Interim data (Section 2.2.1) and provide air
 204 temperature, horizontal and vertical wind, boundary layer height, specific humidity, as well
 205 as convective and large scale precipitation. For the trajectory analysis, 80 release points were
 206 defined along the cruise track. Time and position of these release events are synchronized
 207 with the water and air samples (Section 2.3). At each event, 10,000 trajectories were launched
 208 from the ocean surface within a time frame of ± 30 minutes and an area of $\sim 400 \text{ m}^2$.

209

210 **2.4.2. VSLS source-loss estimate in the MABL**

211 The time scales of air mass transport derived from FLEXPART together with the oceanic
 212 emissions and chemical losses of the VSLS are used for a mass balance source-loss estimate
 213 over the South China and Sulu Seas. For each release event, a box given by the in-situ height
 214 of the MABL and by the horizontal area of the trajectory releases ($\sim 400 \text{ m}^2$ centred on the
 215 measurement location) is defined. The MABL source-loss estimate is based on the
 216 assumption of a constant VSL S mixing ratio (given by the atmospheric measurements), a
 217 constant sea – air flux, the chemical loss rate, and a VSLS homogeneous distribution with the
 218 box during each release.

219 The Oceanic Delivery (OD) is given as the contribution of VSLS sea – air flux (in mol per
 220 day) to the total amount of the VSLS in the box (in mol) in percentage per day. The loss of
 221 MABL air to the FT caused by vertical transport, denoted here Convective Loss (COL), is
 222 calculated from the mean residence time of the FLEXPART trajectories in the observed
 223 MABL during each release and is given as a negative number in percentage per day. COL
 224 equals the loss of VSLS from the MABL to the FT. The Chemical Loss (CL), in the form of
 225 reaction with OH and photolysis, is estimated in percentage per day (negative quantity) and is
 226 based on the tropical MABL lifetime estimates of 15 days for CHBr_3 , 94 days for CH_2Br_2 and
 227 4 days for CH_3I (Carpenter et al., 2014).

228 Relating the delivery of VSLS from the ocean to the MABL (OD) and the loss of MABL air
 229 containing VSLS to the FT (COL) results in an Oceanic Delivery Ratio (ODR) (Eq. 3):

230

$$\left| \begin{aligned} ODR = \frac{OD [\%d^{-1}]}{-COL [\%d^{-1}]} = \frac{\text{Sea-Air flux contribution} [\%d^{-1}]}{\text{Loss of MABL air to the FT} [\%d^{-1}]} \end{aligned} \right. \quad (\text{Eq. 3})$$

231

232 Similarly, the Chemical Loss in the MABL (CL) related to the MABL VSLs loss into the FT
 233 (COL) leads to a Chemical Loss Ratio (CLR) (Eq. 4):

234

$$CLR = \frac{CL [\%d^{-1}]}{-COL [\%d^{-1}]} = \frac{\text{Loss through chemistry} [\%d^{-1}]}{\text{Loss of MABL air to the FT} [\%d^{-1}]} \quad (\text{Eq. 4})$$

235

236 The oceanic delivery, chemical loss and loss to the FT must be balanced by advective
 237 transport of air masses in and out of the box. We define the change of the VSLs through
 238 advective transport as Advective Delivery (AD) in percentage per day (Eq. 5). Additionally,
 239 we define the ratio of change in VSLs caused by advection (AD) to the loss of VSLs out of
 240 the MABL to the FT as Advective Delivery Ratio (ADR) in Eq. 6:

241

$$AD = -COL - \mp CL - OD \quad (\text{Eq. 5})$$

$$ADR = \frac{AD [\%d^{-1}]}{-COL [\%d^{-1}]} = 1 - \mp CLR - ODR \quad (\text{Eq. 6})$$

242

243 Note that for the VSLs within the MABL box, COL and CL are loss processes ~~and given as~~
 244 ~~negative numbers~~ while OD and AD (besides very few exceptions for the latter) are source
 245 processes ~~and given as positive numbers~~. In order to derive the ratios, we divided CL, OD
 246 and AD by COL ~~and therefore end up with negative ratios for the loss process and positive~~
 247 ~~ratios for the source processes~~.

248 In a final step, we relate the source-loss ratios (ODR, CLR and ADR) to the MABL VSLs
 249 volume mixing ratio VMR_{MABL} in the box (Eq. 7 – 9), in order to estimate VSLs newly
 250 supplied from oceanic delivery (VMR_{ODR}), lost by chemical processes (VMR_{CLR}) and
 251 supplied by advective transport (VMR_{ADR}).

252

$$VMR_{ODR} = ODR \cdot VMR_{MABL} \quad (\text{Eq. 7})$$

$$VMR_{CLR} = CLR \cdot VMR_{MABL} \quad (\text{Eq. 8})$$

$$VMR_{ADR} = ADR \cdot VMR_{MABL} \quad (\text{Eq. 9})$$

253

254 2.4.3. Oceanic and MABL VSLs contribution to the FT

255 We use a simplified approach to calculate the mean contribution of boundary layer air masses
 256 observed from various oceanic regions in the South China Sea on the ship, and the oceanic
 257 compounds therein, to the FT above the South China and Sulu Seas. The contribution is

258 determined as a function of time and altitude based on the distribution of the trajectories
259 released at each measurement location along the ship track. According to R/A FALCON
260 observations and our trajectory calculations we assume a well-mixed FT ~~within 5° S – 20° N,~~
261 ~~100° E – 125° E~~. Observations on R/V SONNE, on the other hand, are characterized by large
262 variability and are considered to be representative for the area along the cruise track where
263 the VSLS were measured in the water and atmosphere. We constrain our calculations to this
264 area and define 80 vertical columns along the cruise track. Each column extends horizontally
265 over the area given by the starting points of the trajectories (20 m x 20 m centred on the
266 measurement location) and vertically from the sea surface up to the highest point of R/A
267 FALCON observations around 13 km altitude. For each of the 80 columns along the cruise
268 track, 10,000 trajectories were launched and assigned an identical MABL air parcel
269 containing air with the VSLS mixing ratios observed on R/V SONNE during the time of the
270 trajectory release. The volume of the air parcel is given by the in-situ height of the MABL
271 and the horizontal extent of the release box (20 m x 20 m) divided by 10,000 trajectories. The
272 transport of the MABL air parcels is specified by the trajectories, assuming that no mixing
273 occurs between the parcels during the transport. Chemical loss of the VSLS in each air parcel
274 is taken into account through chemical degradation according to their specific tropospheric
275 lifetimes. The VSLS mixing ratios in the FT from the aircraft measurements are considered
276 representative for the whole South China Sea area. Thus we average over the volume and
277 mixing ratios of all trajectories within the South China Sea area, independent of their exact
278 horizontal location. Due to the decreasing density of air in the atmosphere with height, the
279 volume of the MABL air parcels expands along the trajectories with increasing altitude. The
280 expanding MABL air parcels take up an increasing fraction of air within the FT column,
281 which is taken into account in our calculations using density profiles from our radiosonde
282 measurements.

283 We calculate the contribution of oceanic compounds to the FT ~~for 25 layers of 500-m height~~
284 ~~intervals between 0.5 km and up to~~ 13 km altitude, ~~the upper height of R/A FALCON~~
285 ~~observations~~, within the column above the measurement location. For each layer, the ratio
286 r_{MABL} of the volume of the MABL air parcels with the VSLS mixing ratio VMR_{MABL} to the
287 whole air volume of the layer is calculated. The ratio of advected FT air with a mixing ratio
288 VMR_{AFT} to the whole air volume of the layer is r_{AFT} , respectively, with $r_{MABL} + r_{AFT} = 1$. In
289 our simulation, the FT air with a mixing ratio VMR_{FT} observed by R/A FALCON at a
290 specific height is composed of the MABL air parcels and of the advected FT air parcels (Eq.
291 10):

292

$$r_{MABL} \cdot VMR_{MABL} + r_{AFT} \cdot VMR_{AFT} = (r_{MABL} + r_{AFT}) \cdot VMR_{FT} \quad (\text{Eq. 10})$$

293

294 The relative contribution C_{MABL} of VSLS observed in the MABL to the VSLS observed in the
 295 FT at height z and time t is computed in altitude steps of 500 m (Eq. 11):

296

$$C_{MABL}(z, t) [\%] = 100 \cdot (r_{MABL}(z, t) \cdot VMR_{MABL}(CL(t))) / VMR_{FT}(z) \quad (\text{Eq. 11})$$

297

298 The oceanic contribution C_{ODR} of the South China Sea emissions to the VSLS in the FT is
 299 computed after Eq. 12:

300

$$C_{ODR}(z, t) [\%] = 100 \cdot (r_{MABL}(z, t) \cdot VMR_{ODR}(CL(t))) / VMR_{FT}(z) \quad (\text{Eq. 12})$$

301

302 The simplified approach also allows deriving mean VSLS mixing ratios accumulated in the
 303 FT from both MABL VSLS and oceanic emissions. The FT VSLS mixing ratios are
 304 simulated for each of the 80 columns by initiating a new trajectory release event using same
 305 meteorological conditions and VSLS MABL observations, ~~when the former MABL air has~~
 306 ~~been transported into the FT, according to the specific residence time in the MABL. The~~
 307 ~~initial FT background mixing ratios are 0 ppt for each VSLS.~~ The accumulated mean mixing
 308 ratio of a compound at a specific height z is then ~~computed iterated~~ after Eq. 13:

309

~~$$VMR_{MFT} = r_{MABL_z} \cdot VMR_{MABL_z} + r_{MABL_z} \cdot VMR_{MABL_z} + \dots + r_{MABL_t} \cdot VMR_{MABL_t} \quad (\text{Eq. 13})$$~~

310

$$VMR_{FT}(z) = \sum_{i=1}^n \left(\begin{array}{c} r_{MABL}(z, t(i)) \cdot VMR_{MABL}(CL(t(i))) \\ + \\ (1 - r_{MABL}(z, t(i))) \cdot VMR_{FT}(z, t(i-1)) \cdot CL(COL) \end{array} \right) \quad (\text{Eq. 13})$$

311

312 Here, n is the number of runs for each of the 80 profiles, according to the residence time of
 313 the trajectories in the MABL and the total runtime of the trajectories. For example a
 314 trajectory residence time of 7 hours in the MABL in combination with a total trajectory
 315 runtime of 16 days leads to $n = (24 / 7) * 16 = 54$. r_{MABL} gives the volume ratio of MABL air
 316 parcels at height z and time t to the total volume of a specific height layer. VMR_{MABL} gives the
 317 compounds mixing ratio in the air parcels including chemical degradation (CL) since the air
 318 was observed in the MABL. Since we use a mean tropospheric lifetime, CL and thus also

319 VMR_{MABL}(CL) are independent from the height z . The initial FT background mixing ratios
320 (VMR_{FT}($z, t(i=0)$)) are set to 0 ppt for each VSLS _{i} , followed by n times iteration of VMR_{FT}(i -
321 1). The difference between VMR_{FT}($i = n-1$) and VMR_{FT}($i = n$) during the last two steps of the
322 iteration is less than 1 % for each compound within the 16 days runtime.
323 ~~Here, VMR_{FT} is the modelled accumulated FT mixing ratio, r_{MABL_i} is the ratio of MABL air~~
324 ~~parcels in 20 m x 20 m x 500 m layers between 0.5 km and 13 km altitude to the total volume~~
325 ~~of each layer, VMR_{MABL_i} is the mixing ratio in the MABL air parcels including chemical~~
326 ~~degradation since release from the MABL, and i is the number of initiated runs per release. A~~
327 ~~steady state for the compounds is reached, when variations in their mixing ratios vary less~~
328 ~~than 1 % between two initiated runs. For CHBr₃ the steady state is reached after 11.0 ± 2.1 d~~
329 ~~(mean $\pm \sigma$), 11.8 ± 2.4 d for CH₂Br₂ and 8.0 ± 1.4 d for CH₃I. The modelled overall mean FT~~
330 ~~mixing ratio in the South China Sea is derived as the mean from the 80 individually~~
331 ~~calculated FT mixing ratios determined along the cruise. The oceanic contribution to the FT~~
332 ~~compounds is calculated with VMR_{ODR} from Eq. 7 inserted as VMR_{MABL} in Eq. 13.~~

333

334 **3. Meteorological conditions in the MABL and the FT**

335 **3.1. Meteorology along the ship cruise**

336 Moderate to fresh trade winds were dominating the South China and Sulu Seas during the
337 cruise (Figure 1a-b), indicated by the overall mean wind direction of northeast ($50^\circ - 60^\circ$)
338 and a mean wind speed of 5.5 ± 2.9 ms⁻¹. The wind observations reveal two different air mass
339 origins. Between November 15 and 19, 2011 a gentle mean wind speed of 3.7 ± 1.8 ms⁻¹ with
340 a *northern* wind direction was observed, influenced by a weak low pressure system (not
341 shown here) over the central South China Sea moving southwest and passing the ship
342 position on 17.11.2011. During November 20 – 29, 2011 the wind direction changed to
343 *northeast* and the mean wind speed increased to moderate 6.4 ± 3.0 ms⁻¹. A comparison
344 between 6 hourly ERA-Interim wind and a 6 hourly averaged mean of the observed wind on
345 R/V SONNE reveals an underestimation of the wind speed by ERA-Interim along the cruise
346 track by 1.6 ± 1.4 ms⁻¹ on average (not shown here). The mean deviation of the wind
347 direction between reanalysis and observation is 2 ± 37 degree. Reanalysis and observed wind
348 speeds correlate with $R = 0.76$ and the wind directions with $R = 0.86$, reflecting a good
349 overall agreement between ship observation and ERA-Interim winds. With an observed mean
350 surface air temperature (SAT) of 28.2 ± 0.8 °C and a mean SST of 29.1 ± 0.5 °C the SAT is
351 on average 1.0 ± 0.7 °C below the SST, which benefits convection of surface air (Figure 2).

352 Indeed enhanced convective activity and pronounced precipitation events have been observed
353 during the cruise (S-Figure 1). Figure 3a shows the time series of the relative humidity
354 measured by the radiosondes launched on R/V SONNE from the surface up to the mean
355 height of the cold point tropopause at 17 km. Elevated humidity is found on average up to
356 about 6 km, which implies a distinct transport of water vapour to the mid troposphere during
357 the cruise by deep convection or advection of humid air from a nearby convective cell.

358

359 **3.2. Marine atmospheric boundary layer**

360 Higher SSTs than SATs (Figure 2) cause unstable atmospheric conditions (negative values)
361 between the surface and about 50 – 100 m height (Figure 3b). Surface air is heated by warmer
362 surface waters and is enriched with humidity both benefiting moist convection. The stability
363 of the atmosphere increases above 420 ± 120 m and indicates the upper limit of the MABL at
364 this altitude range derived from radiosonde data (Figure 3b). The MABL height given by
365 ERA-Interim along the cruise track is with 560 ± 130 m systematically higher (not shown),
366 but still within the upper range of the MABL height derived from the radiosonde
367 measurements. The unstable conditions of the MABL and the increase of the atmospheric
368 stability above the MABL reflect the characteristics of a convective, well-ventilated tropical
369 boundary layer. In contrast to cold oceanic upwelling regions with a stable and isolated
370 MABL (Fuhlbrügge et al., 2013; Fuhlbrügge et al., 2015), the vertical gradient of the relative
371 humidity measured by the radiosondes (Section 3.1) and the height of the MABL do not
372 coincide. This is caused by increased mixing through and above the MABL by turbulence
373 and convection, which leads to the convective, well-ventilated MABL.

374

375 **4. Atmospheric VSLs over the South China and Sulu Seas**

376 **4.1. Atmospheric surface observations on R/V SONNE**

377 Overall, the three VSLs show a similar pattern of atmospheric mixing ratios along the cruise
378 track with lower atmospheric surface abundances before 21 November 2011 and higher
379 concentrations afterwards, which can be attributed to a change in air mass origin (Figure 1).
380 A decrease from 3.4 to 1.2 ppt of CHBr_3 occurs at the beginning of the cruise (Figure 4a)
381 when the ship left Singapore and the coast of the Malaysian Peninsula. On November 16 –
382 19, 2011, when the ship passed the southern South China Sea, lower mixing ratios (\pm
383 standard deviation 1σ) of 1.2 ± 0.3 ppt prevail and also the lowest mixing ratios for CHBr_3
384 during the whole cruise of 0.8 ppt are observed. At the coast of Borneo and the Philippines,
385 the average mixing ratio of CHBr_3 increases to 2.3 ± 1.4 ppt. The overall mean CHBr_3

386 mixing ratio during the cruise is 2.1 ± 1.4 ppt (Table 1) and therefore higher than earlier
387 reported CHBr_3 observations of 1.2 ppt in January – March 1994 (Yokouchi et al., 1997), 1.1
388 ppt in September 1994 (Quack and Suess, 1999) and 1.5 ppt in June – July 2009 (Nadzir et
389 al., 2014) further offshore in the South China Sea. The higher atmospheric mixing ratios
390 during the R/V SONNE cruise in November 2011 in contrast to the lower mixing ratios in
391 these previous studies may point to stronger local sources, strong seasonal or interannual
392 variations, or even to long-term changes. CH_2Br_2 shows a mean mixing ratio of 1.2 ± 0.2 ppt
393 (Table 1). Yokouchi et al. (1997) observed a lower mean atmospheric mixing ratio of 0.8 ppt
394 and Nadzir et al. (2014) of 1.0 ppt in the South China Sea. An increase of the CH_2Br_2 mixing
395 ratios from 1.0 ± 0.1 ppt to 1.3 ± 0.2 ppt is observed after November 21, 2011 coinciding
396 with an increase of the CH_3I concentrations from primarily 0.3 ± 0.0 ppt to 0.4 ± 0.1 ppt
397 (Figure 4a). The highest mixing ratio of CH_3I was detected in the southwestern Sulu Sea on
398 November 25, 2011 with 0.8 ppt. The overall mean atmospheric mixing ratio for CH_3I , of 0.4
399 ± 0.1 ppt (Table 1) is lower than the mean of 0.6 ppt observed by Yokouchi et al. (1997).

400 The concentration ratio of CH_2Br_2 and CHBr_3 (Figure 4b) has been used as an indicator of
401 relative distance to the oceanic source, where a ratio of 0.1 was observed crossing strong
402 coastal source regions (Yokouchi et al., 2005; Carpenter et al., 2003). The ten times elevated
403 CHBr_3 has a much shorter lifetime, thus degrading more rapidly than CH_2Br_2 , which
404 increases the ratio during transport. Overall, the mean concentration ratio of CH_2Br_2 and
405 CHBr_3 is 0.6 ± 0.2 , which suggests that predominantly older air masses are advected over the
406 South China Sea.

407

408 **4.2. Oceanic surface concentrations and emissions from R/V SONNE**

409 VSLs in the surface sea water along the cruise track show highly variable distributions
410 (Figure 4c and Table 1). Oceanic CHBr_3 surface concentrations range from 2.8 – 136.9 pmol
411 L^{-1} with a mean of 19.9 pmol L^{-1} during the cruise, while CH_2Br_2 concentrations range from
412 2.4 – 21.8 pmol L^{-1} with a mean of 5.0 pmol L^{-1} . CHBr_3 and CH_2Br_2 have similar distribution
413 patterns in the sampling region with near shore areas showing typically elevated
414 concentrations. CH_3I concentrations range from 0.6 – 18.8 pmol L^{-1} with a mean of 3.8 pmol
415 L^{-1} and show a different distribution along the ship track which might be ascribed to
416 additional photochemical production of CH_3I in the surface waters (e.g. Manley and Dastoor,
417 1988; Manley and de la Cuesta, 1997; Richter and Wallace, 2004).

418 High levels of all VSLs are found in waters close to the Malaysian Peninsula, especially in
419 the Singapore Strait on November 16, 2011, possibly showing an anthropogenic influence on

420 the VSLs concentrations. VSLs concentrations decrease rapidly when the cruise track leads
421 to open ocean waters. Along the west coast (November 19 - 23, 2011) and northeast coast of
422 Borneo (November 25, 2011), bromocarbon concentrations are elevated, and especially
423 CHBr_3 concentrations increase in waters with lower salinities, indicating an influence by
424 river run off. Elevated CHBr_3 concentrations are often found close to coasts with riverine
425 inputs caused by natural sources and industrial and municipal effluents (see Quack and
426 Wallace, 2003; Fuhlbrügge et al., 2013 and references therein).

427 Oceanic emissions were calculated from synchronized measurements of sea water
428 concentrations and atmospheric mixing ratios, sea surface temperatures and wind speeds,
429 measured on the ship (Section 2.3.3). The overall VSLs distribution along the ship track is
430 opposite for the oceanic and atmospheric measurements (Figure 4a-d). While the sea water
431 concentrations of VSLs generally decrease towards the Sulu Sea, the atmospheric mixing
432 ratios increase, leading to a generally lower concentration gradient of the compounds
433 between sea water and air in the Sulu Sea (not shown here).

434 Coinciding low VSLs atmospheric background concentrations, high SSTs, elevated oceanic
435 VSLs concentrations and high wind speeds, lead to high emissions of VSLs for the South
436 China and Sulu Seas (Figure 4d) of $1486 \pm 1718 \text{ pmol m}^{-2} \text{ h}^{-1}$ for CHBr_3 , $405 \pm 349 \text{ pmol m}^{-2} \text{ h}^{-1}$
437 for CH_2Br_2 and $433 \pm 482 \text{ pmol m}^{-2} \text{ h}^{-1}$ for CH_3I . In particular, CHBr_3 fluxes are very
438 high and thus confirm elevated coastal fluxes from previous campaigns in tropical source
439 regions (Quack et al., 2007). They often exceed $2000 \text{ pmol m}^{-2} \text{ hr}^{-1}$ in the coastal areas and
440 are sometimes higher than $6000 \text{ pmol m}^{-2} \text{ hr}^{-1}$, as in the Singapore Strait on November 15,
441 2011 and on November 22, 2011 at the northwest coast of Borneo, which was also an area of
442 strong convection (Figures 1, 4b).

443

444 **4.3. VSLs intercomparison: R/A FALCON and R/V SONNE**

445 The two profiles of the bromocarbon mixing ratios and the profile for CH_3I from the surface
446 to 13 km altitude as observed on R/A FALCON with the GhOST and WASP instruments
447 (Sala et al., 2014; Tegtmeier et al., 2013) are shown in Figure 5. Mean CHBr_3 mixing ratios
448 are 1.43 ppt (GhOST) and 1.90 ppt (WASP) in the MABL (0 – 450 m, determined from
449 meteorological aircraft observations similarly as for the radiosondes, Section 2.2.2) and 0.56
450 ppt (GhOST) and 1.17 ppt (WASP) in the FT (0.45 km – 13 km, Table 1). The GhOST
451 mixing ratios in the MABL are lower than those observed on R/V SONNE (2.08 ppt). A very
452 good agreement of the measurements is given for the longer lived CH_2Br_2 with 1.17 ppt (R/V
453 SONNE), 1.19 ppt (GhOST) and 1.15 ppt (WASP). CH_3I mixing ratios measured by GhOST

454 are 0.59 ± 0.30 ppt within the MABL of 450 m height, which is about 0.2 ppt higher than the
455 values from R/V SONNE. Above the MABL, the average mixing ratio of CH₃I decreases to
456 0.26 ± 0.11 ppt (Figure 5).

457 CHBr₃ and CH₂Br₂ concentrations in the MABL correlate with $R = 0.83$ for all instruments
458 (Figure 6). CHBr₃ and CH₃I concentrations correlate with $R = 0.55$ and CH₂Br₂ and CH₃I
459 with $R = 0.66$; all correlations are significant at 99 %. Even higher correlations are found if
460 only measurements on R/V SONNE are taken into account with $R = 0.92$ for CHBr₃ and
461 CH₂Br₂, $R = 0.64$ for CHBr₃ and CH₃I, and $R = 0.77$ for CH₂Br₂ and CH₃I.

462 Comparison of R/A FALCON and R/V SONNE data are obtained from their meetings on
463 November, 19 and 21, 2011 (Table 2), when aircraft and ship passed each other within 100 m
464 distance several times, measuring the same air masses. During both meetings, deviations
465 between the GhOST and WASP instruments on the aircraft are larger for the bromocarbons
466 than the deviation between the WASP and the ship measurements. According to Sala et al.
467 (2014) the agreement between the GhOST and WASP instruments are within the expected
468 uncertainty range of both instruments which is then assumed to be also valid for the ship
469 measurements (this study). The good agreement between WASP and ship data might be
470 caused by the same sampling and analysis method, both using stainless steel canisters and
471 subsequent analysis with GC/MS, while GhOST measures in-situ with a different resolution.
472 Since GhOST and WASP measurements together cover a larger spatial area and higher
473 temporal resolution, a mean of both measurements is used in the following for computations
474 in the free troposphere. For CH₃I significantly higher mixing ratios were measured during the
475 meetings between ship and aircraft (Table 2). Whether this offset is systematic for the
476 different methods, needs further investigation.

477

478

479 **5. Air mass and VSLs transport from the surface to the free troposphere**

480 **5.1. Timescales and intensity of vertical transport**

481 Forward trajectories computed with FLEXPART starting at sea level along the cruise track
482 yield an average MABL residence time of 7.8 ± 3.5 hr before the trajectories enter the FT
483 (Figure 7), reflecting a relatively fast exchange due to the convective well ventilated MABL
484 (Figure 3). The trajectories generally show a strong contribution of surface air masses to the
485 FT, despite some exceptions during November 18 – 22, 2011 (Figure 7). Most intense and
486 rapid transport of MABL air masses up to 13 km height occurs on November 17 and 23,
487 2011. ~~To estimate the loss of air masses out of the South China Sea area between 5° S—20°~~

~~N and 100° E—125° E (Section 2.4.3) we determine the loss of trajectories out of this area after their release (Figure 8). After 4 days, 88 % of all trajectories released along the cruise track are still within this defined area of the South China Sea, and 31 % after 10 days.~~

5.2. Contribution of oceanic emissions to VSLs in the MABL

From the sea – air fluxes (Section 4.2) and the residence times of the surface trajectories in the MABL (Section 5.1), the Oceanic Delivery (OD) and the CONvective Loss (COL) were computed (Table 3) using the method described in Section 2.4.2.

Based on the OD and the COL, the Oceanic Delivery Ratio (ODR) is calculated in order to characterize the relative contribution of the local oceanic emissions compared to the loss of MABL air into the FT (Table 3, Figure 8). The average ODR during the cruise is 0.45 ± 0.55 for CHBr_3 , which means that the loss from the MABL to the FT is balanced to 45 % by oceanic emissions along the cruise track. The ODR for CH_2Br_2 is 0.20 ± 0.21 and for CH_3I 0.74 ± 1.05 , respectively, suggesting that the major amount of CH_3I originates from nearby sources. Similarly to the ODR the CL is related to the COL to derive the Chemical Loss Ratio (CLR) for the VSLs, which is 0.03 ± 0.01 for CHBr_3 , 0.01 ± 0.00 for CH_2Br_2 and 0.09 ± 0.04 for CH_3I . When compared to the other source and loss processes, the chemical loss appears negligible for all three gases. The ratio of the Advective Delivery (ADR) is 0.58 ± 0.55 for CHBr_3 , 0.80 ± 0.21 for CH_2Br_2 and 0.35 ± 1.02 , implying that most of the observed CH_2Br_2 (80 %) in the MABL is advected from other source regions. Applying the ODR to the observed mixing ratios in the MABL gives an estimate of the VSLs originating from the local oceanic emissions (VMR_{ODR} , Table 4). The local ocean emits a concentration that equates to 0.89 ± 1.12 ppt CHBr_3 , 0.25 ± 0.26 ppt CH_2Br_2 and 0.28 ± 0.40 ppt CH_3I in the MABL. The average transport from the MABL to the FT ($\text{Flux}_{\text{MABL-FT}}$), computed from the MABL concentrations and the trajectory residence time in the MABL, is 4240 ± 1889 $\text{pmol m}^{-2} \text{ hr}^{-1}$ for CHBr_3 , 2419 ± 929 $\text{pmol m}^{-2} \text{ hr}^{-1}$ for CH_2Br_2 and 865 ± 373 $\text{pmol m}^{-2} \text{ hr}^{-1}$ for CH_3I . Calculations with the ERA-Interim MABL height, which is on average 140 m higher than the one derived from the radiosondes, leads to similar estimates (S-Table 1).

Since the wind is a driving factor for oceanic emissions and advection of VSLs, changes in wind speed are assumed to affect atmospheric VSLs mixing ratios in the MABL during this cruise. Significant correlations are found between wind speed and the observed mixing ratios of all three VSLs in the MABL with correlation coefficients of $R = 0.55$ (CHBr_3), $R = 0.57$ (CH_2Br_2) and $R = 0.56$ (CH_3I), respectively. Mixing ratios that originate from oceanic

emissions (VMR_{ODR}) correlate significantly to the wind speed with $R = 0.52$, $R = 0.72$ and $R = 0.62$, respectively. On the opposite, VMR_{ADR} , which is calculated as the residual from VMR_{ODR} , is negatively correlated to the wind speed with $R = -0.21$, $R = -0.32$ and $R = -0.53$. The correlations reveal that the contribution of oceanic emissions to MABL VSLs increase for higher wind speeds, while the advective contribution decreases.

527

5.3. Oceanic contribution to the FT

5.3.1. Identification of VSLs MABL air in the FT

With a simplified approach (method description in Section 2.4.3) we are able to estimate the contribution of MABL air and regional marine sources observed on R/V SONNE to the FT. Individual MABL air masses during the cruise show ~~the strongest~~ a strong contribution to the FT air up to 10 % after 10 days ~~within the lower 3 km of the atmosphere during the first 6 days after release~~ (Figure 9a-f). The MABL air is rapidly transported into high altitudes before it is dispersed in the FT column. ~~Once the MABL air is spread in the FT column, a decrease of the contribution with time is due to the transport of trajectories out of the predefined South China and Sulu Seas area and their chemical loss (Figure 8).~~

The average contribution of VSLs concentrations in the MABL air to the FT concentrations (ϵ_{MABL}) is generally highest for CHBr_3 , with ~~about 5–10~~ about 10 – 15 % between 4 – 10 km and up to 25 % above 10 km height ~~% above 3 km height within 10–3 – 10 days after release in the MABL~~ (coloured contours in Figure 9a), followed by CH_2Br_2 (~~2–5~~ 5 - 10 % below 10 km and up to 16 % above 10 km height, Figure 9b). The lowest contribution is found for ~~the short-lived CH_3I and only with 0–3~~ up to 4 % within 3 days ~~for the short-lived CH_3I~~ (Figure 9c). In general the contribution below 5 km height decreases after about 10 days, when most of the MABL air is transported into higher altitudes. For CH_3I , ~~The the~~ chemical degradation ~~of CH_3I~~ , according to its short tropospheric lifetime of 3.5 days, leads to a rapid decrease of the contribution ~~already 3~~ 2 days after release.

To identify the contribution of the oceanic emissions to the FT VSLs during the cruise, the VMR_{ODR} of each compound is used as the initial mixing ratio in the MABL air mass. For CHBr_3 and CH_2Br_2 the local emissions contribute only up to ~~6–11~~ % and ~~2–4~~ % to the FT concentrations (ϵ_{ODR} , Figure 9d-e) compared to the ~~25–10~~ %, respectively ~~5–16~~ %, of ϵ_{MABL} . In contrast, the contribution of the local oceanic emissions of CH_3I (Figure 9f) is almost similar to the contribution of the observed MABL concentrations (Figure 9c).

554

5.3.2. Accumulated VSLs in the free troposphere

By simulating a steady transport of MABL air masses into the FT, mean accumulated VSLs mixing ratios in the FT along and during the cruise were computed (Figure 10) as described in 2.4.3. The simulated FT mixing ratios of CHBr_3 and CH_2Br_2 from the observed MABL (VMR_{MABL}) decrease on average from 1.72.1 and 1.1-2 ppt at 0.5 km height at the surface to 0.6 and 0.6-0.8 ppt at 7-3 km height. The simulated CHBr_3 mixing ratios are constant up to 8 km height followed by an increase up to 0.9 ppt at 13 km height. Simulated CH_2Br_2 mixing ratios increase between 3 and 11.5 km height up to 1.1 ppt and remain constant above 11.5 km height. ~~and increase again above 8 km up to 0.8 (CHBr_3) and 0.8 ppt (CH_2Br_2).~~ Simulated CH_3I shows a decrease from 0.20-4 ppt at 0.5 km the surface to 0.05-08 ppt at 3 km. Above this altitude, the simulated mixing ratios of CH_3I are almost constant with 0.05 ppt before they slowly increase above 9 km height to 0.10 ppt.

To estimate the accumulated FT mixing ratios solely from oceanic emissions, the VMR_{ODR} is used as the initial MABL mixing ratio (Figure 10). The simulated FT mixing ratios using either VMR_{MABL} or VMR_{ODR} as input reveal a similar vertical pattern, since both simulations are based on the same meteorology and trajectories. While FT mixing ratios based on VMR_{MABL} and VMR_{ODR} are similar for CH_3I and differ by about 0.03 ppt (due to the large oceanic contribution to the MABL mixing ratios), FT mixing ratios from VMR_{ODR} are on average ~0.5-4 ppt lower for CHBr_3 and ~0.6-8 ppt CH_2Br_2 than from VMR_{MABL} . Comparing the simulated VMR_{MABL} FT mixing ratios with the observed FT mixing ratios from R/A FALCON reveals generally stronger vertical variations for the ~~simulations-observations~~ in contrast to the ~~observations-simulations~~. Still, CHBr_3 observations ~~is~~ are well reflected in the ~~overestimated in the~~ VMR_{MABL} simulation between 0-2 and 4-11 km altitude, ~~as well as above 12 km.~~ Also the ~~S~~ simulated CH_2Br_2 in the FT based on VMR_{MABL} ~~underestimates~~ reflects the ~~observed-observations~~ of R/A FALCON ~~mixing ratios between~~ very well from the surface up to 9-6 and 10-km height. ~~In particular, the maximum between 7 and 9 km height is not reflected in the simulations. However, observations of both bromocarbons are within 1σ of the FT simulations with MABL air.~~ The CH_3I simulations show a distinct underestimation of the observed FT mixing ratios. Adjusting the R/A FALCON values by the identified offset to R/V SONNE (Section 4.3 and Table 2) reveals a better agreement between observed and simulated FT mixing ratios (Figure 10).

586

5.3.3. Discussion

587

588 Oceanic emissions of CHBr_3 from the South China Sea contribute on average 45 % to the
589 simulated FT mixing ratios (Figure 11). Simulated FT mixing ratios from MABL
590 observations and observed FT mixing ratios agree quite well up to ~~10~~ 11 km height. Above
591 this altitude the simulated FT mixing ratios, in contrast to the observations increase, which is
592 caused by an overestimation of the convective activity ~~outflow~~ in our method. However,
593 despite using this simple approach we are able to simulate mean FT mixing ratios up to 11
594 km height above the South China Sea in good agreement with observations. Thus, we assume
595 that the observed MABL mixing ratios of CHBr_3 and CH_2Br_2 are representative for the South
596 China and Sulu Seas. ~~In case of CH_2Br_2 the underestimation of simulated FT mixing ratios
597 between 6 and 10 km using observed MABL mixing ratios may be due to higher background
598 concentrations of CH_2Br_2 in the mid to upper troposphere given the long atmospheric lifetime
599 of 150 days at these levels.~~

600 On average, 45 % and 20 % of CHBr_3 and CH_2Br_2 abundances, ~~respectively~~, in the MABL,
601 observed ~~only~~ by the ship, originate from local oceanic emissions along the ship track. Thus,
602 advection from stronger source regions, possibly along the coast (for CHBr_3) and from the
603 West Pacific (CH_2Br_2), are necessary to explain the **observed MABL** mixing ratios.

604 In contrast to CHBr_3 and CH_2Br_2 , the simulated mixing ratios of CH_3I in the FT are strongly
605 underestimated no matter whether observed MABL mixing ratios or oceanic emissions are
606 used. The offset between the simulated and observed FT CH_3I could be caused by additional
607 strong sources of CH_3I in the South China Sea area. Furthermore, modelling or measurement
608 uncertainties may add to this offset.

609 The simulations use constant atmospheric lifetimes for each compound and neglect **lifetime**
610 variations with altitude which could impact the simulated abundances. However, the altitude
611 variations of the CH_3I lifetime in the MABL and FT are around 0.5 days (Carpenter et al.,
612 2014) and thus impacts on the simulated abundances are quite small. Therefore, it seems
613 unlikely that the lifetime estimate causes a large underestimation of the FT CH_3I . Additional
614 uncertainties may arise from cloud induced effects on photolysis rates (Tie et al., 2003) and
615 OH levels (e.g. Tie et al., 2003; Rex et al., 2014) impacting the VLSL lifetimes. Deficiencies
616 in the meteorological input fields and the FLEXPART model, in particular in the boundary
617 layer and in the convection parameterizations would affect **the transport of** all compounds
618 and their contribution to the FT concentrations in a similar way and thus seems to be unlikely
619 as well. Ship and aircraft measurements revealed a possible instrumental offset for CH_3I
620 (Section 4.3). When we adjust the observations for the offset of CH_3I between R/V SONNE
621 and R/A FALCON the simulated and observed FT mixing ratios match better. Thus, an

622 instrumental offset causing, at least partially, the calculated discrepancy for CH₃I appears
623 likely (Section 2.3.1).

624 Another explanation for the elevated CH₃I in the FT is advection of fresh air with elevated
625 CH₃I mixing ratios in the FT from e.g. South East Asia or the Philippines. These areas are
626 known to comprise strong sources for atmospheric CH₃I from e.g. rice plantations (Redeker
627 et al., 2003; Lee-Taylor and Redeker, 2005). In combination with convective activity over
628 land, which is common in this area (Hendon and Woodberry, 1993), the high observed FT
629 mixing ratios of CH₃I could be explained, despite the low oceanic contribution during the
630 cruise. The low observed MABL mixing ratios of CH₃I on R/V SONNE may thus also not be
631 representative for the area. Yokouchi et al. (1997) observed higher atmospheric CH₃I mixing
632 ratios in the South China Sea.

633 Finally, the method of our simplified approach includes uncertainties as well. Different
634 parameterizations for the transfer coefficient k_w such as Liss and Merlivat (1986), which is at
635 the lower end of reported parameterizations, and Wanninkhof and McGillis (1999), which is
636 at the higher end, are discussed in Lennartz et al. (2015). Both lead to a reduction of the
637 oceanic contribution to the atmospheric mixing ratios at the observed average moderate wind
638 speeds ($\sim 6 \text{ ms}^{-1}$) when applied to our data. Nonetheless, the general conclusion that local
639 oceanic sources of CHBr₃ and CH₃I significantly contribute to MABL mixing ratios remains
640 for the cruise. In times of possible higher wind speeds ($>10 \text{ ms}^{-1}$), which are likely for this
641 region, the flux variations between the different parameterizations as well as the oceanic
642 contribution to atmospheric abundances, would increase. Further uncertainties may result
643 from the approach to simulate FT mixing ratios from MABL observations, which does not
644 include e.g. turbulent mixing or entrainment of different air masses containing different
645 VSLs mixing ratios. Since observational studies quantifying the oceanic contribution to
646 atmospheric abundances of VSLs are quite rare, it is difficult to evaluate our findings at the
647 moment and more studies for different oceanic regimes should be carried out to validate our
648 results.

649

650 **6. Summary**

651 The contribution of oceanic VSLs emissions to marine atmospheric boundary layer (MABL)
652 and free troposphere (FT) air during the SHIVA campaign in November 2011 in the South
653 China and Sulu Seas was investigated in this study. Meteorological parameters were
654 measured near the ocean surface and in the troposphere by regular radiosonde launches on
655 R/V SONNE during the cruise. Oceanic VSLs emissions were determined from simultaneous

656 atmospheric observations and sea surface water concentrations. The transport from the
 657 surface through the MABL into the FT was computed with the trajectory model FLEXPART.
 658 The ship campaign was dominated by north-easterly winds with a characteristic moderate
 659 mean wind speed of 5.5 ms^{-1} . The radiosonde launches revealed a convective, well-ventilated,
 660 weakly developed MABL with an average height of $420 \pm 120 \text{ m}$ during the cruise. 800,000
 661 forward trajectories, launched from the ocean surface along the cruise track, show a rapid
 662 exchange of MABL air with the FT within 7.8 hrs. The observations on R/V SONNE reveal
 663 high mean ocean surface concentrations and emissions for CHBr_3 ($19.94 \text{ pmol L}^{-1}$ and 1486
 664 $\text{ pmol m}^{-2} \text{ hr}^{-1}$), CH_2Br_2 (4.99 pmol L^{-1} and $405 \text{ pmol m}^{-2} \text{ hr}^{-1}$) and CH_3I (3.82 pmol L^{-1} and
 665 $433 \text{ pmol m}^{-2} \text{ hr}^{-1}$) in comparison to other oceanic source regions. Atmospheric mixing ratios
 666 in the MABL, on the other hand, are relatively low with mean values of 2.08 ppt CHBr_3 , 1.17
 667 ppt CH_2Br_2 , and 0.39 ppt CH_3I . The contribution of the oceanic VSLS emissions to their
 668 MABL concentrations was evaluated by simple source-loss estimates, resulting in an Oceanic
 669 Delivery Ratio (ODR). The ODR for CHBr_3 is 0.45, revealing that up to 45 % of CHBr_3
 670 mixing ratios in the MABL above the marginal seas originated, on average, from local
 671 oceanic sources, while 74 % of CH_3I and only 20 % of CH_2Br_2 originates from the local
 672 ocean. This indicates that the long-lived CH_2Br_2 is largely advected in the MABL.
 673 We extend our analysis to the FT using VSLS observations from R/A FALCON above the
 674 South China Sea. A single MABL air release contributes up to ~~28-25~~ % (CHBr_3), ~~12-16~~ %
 675 (CH_2Br_2) and ~~5-4~~ % (CH_3I) to the FT mixing ratio (Section 5.3.1). The mean contributions of
 676 the local oceanic VSLS to the FT within this MABL air release are up to ~~13-11~~ % (CHBr_3),
 677 ~~3-4~~ % (CH_2Br_2) and 4 % (CH_3I). In order to estimate if the accumulated contributions from
 678 the single MABL air releases are sufficient to explain the accumulated VSLS mixing ratios
 679 observed in the FT, a steady transport of observed MABL air masses and oceanic emissions
 680 into the FT above the South China Sea was simulated. The simulations for CHBr_3 based on
 681 the volume mixing ratios in the MABL (VMR_{MABL}) reflect the observed mixing ratios in the
 682 FT up to 11 km height, while the simulations based on the local oceanic emissions (VMR_{ODR})
 683 explained about ~~60-45~~ %. In the MABL, the local oceanic emissions along the cruise track
 684 also explain half of the CHBr_3 . Thus, we conclude that the observed mixing ratios of CHBr_3
 685 in the MABL are influenced by stronger, possibly coastal sources, in the region.
 686 CH_2Br_2 in the FT, simulated from observed MABL mixing ratios, shows a good agreement
 687 between observations and simulations up to 9 km height. ~~However, slightly lower simulated~~
 688 ~~mixing ratios in the mid to upper troposphere compared to the observations indicate an~~
 689 ~~accumulation of CH_2Br_2 at these levels due to its longer atmospheric lifetime compared to~~

690 | ~~CHBr₃~~-CH₃I in the FT is underestimated in the simulations, using both the observed MABL
691 mixing ratios and the oceanic emissions. Even addressing an unresolved offset between the
692 ship and aircraft data leads to an underestimation of CH₃I in the FT, which points to
693 additional CH₃I sources, e.g. rice plantations in the region together with pronounced
694 convection.

695 Our investigations show how oceanic emissions of VSLS in a strong oceanic source region
696 contribute to the observed atmospheric mixing ratios in the MABL. Furthermore, the
697 contributions of these atmospheric mixing ratios, and the local oceanic VSLS, to the observed
698 VSLS in the FT above this source region are derived. The results reveal strong links between
699 oceanic emissions, atmospheric mixing ratios, MABL conditions and prevailing convective
700 activity in the troposphere. The methods should be applied to other oceanic regions to derive
701 a better process understanding of the contributions of air-sea gas exchange on atmospheric
702 abundances. For the detection of future climate change effects on ocean surface trace gas
703 emissions and their influence on atmospheric chemistry and composition it is important to
704 study the complex interplay between oceanic sources and emissions, meteorology,
705 atmospheric mixing ratios, and transport to the upper atmosphere.

706

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722

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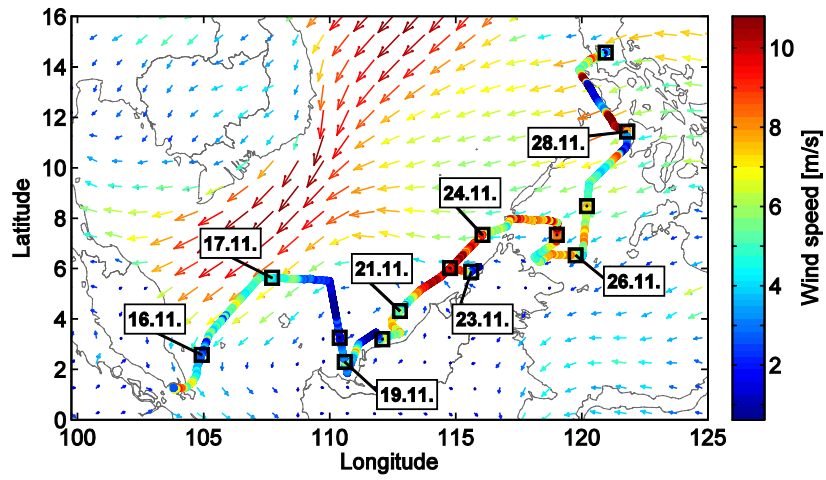
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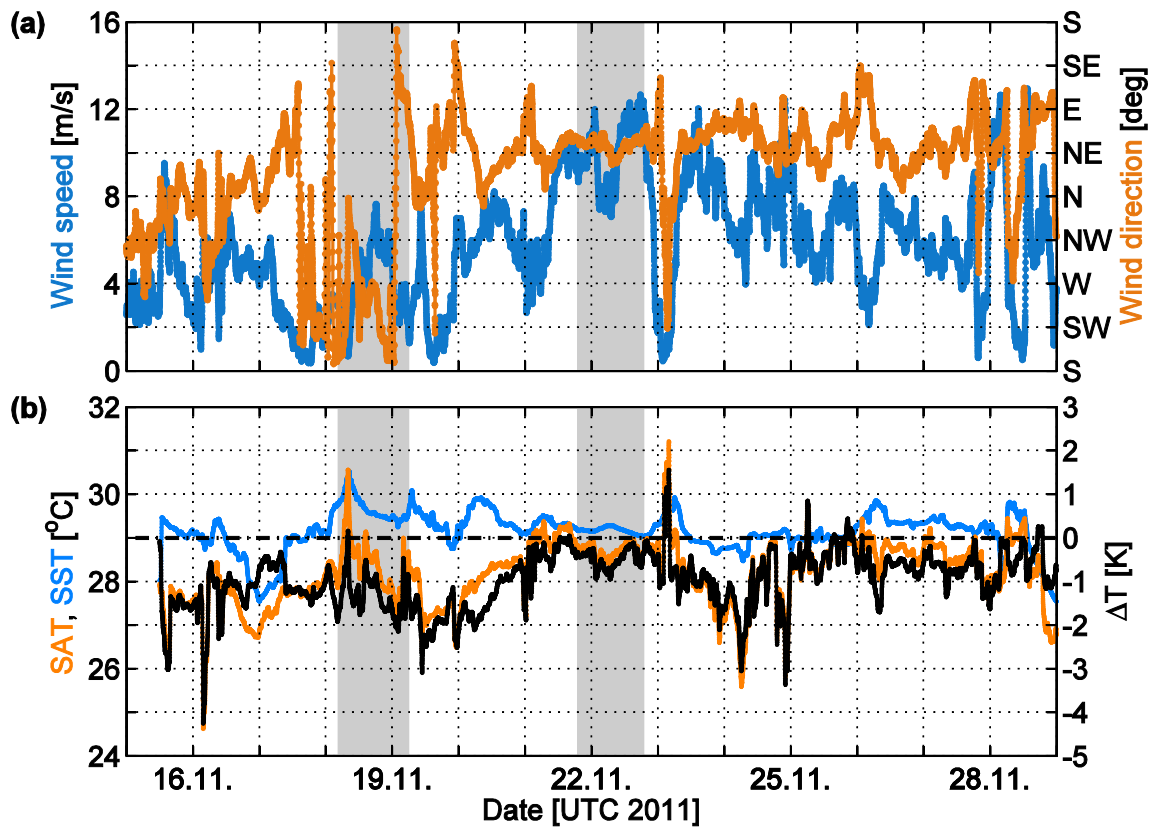
899 **Figures**

900



901 Figure 1: ERA-Interim mean wind field November 15 – 30, 2011 (arrows) and 10 minute
902 running mean of wind speed observed on R/V SONNE as the cruise track. The black squares
903 show the ships position at 00 UTC each day.

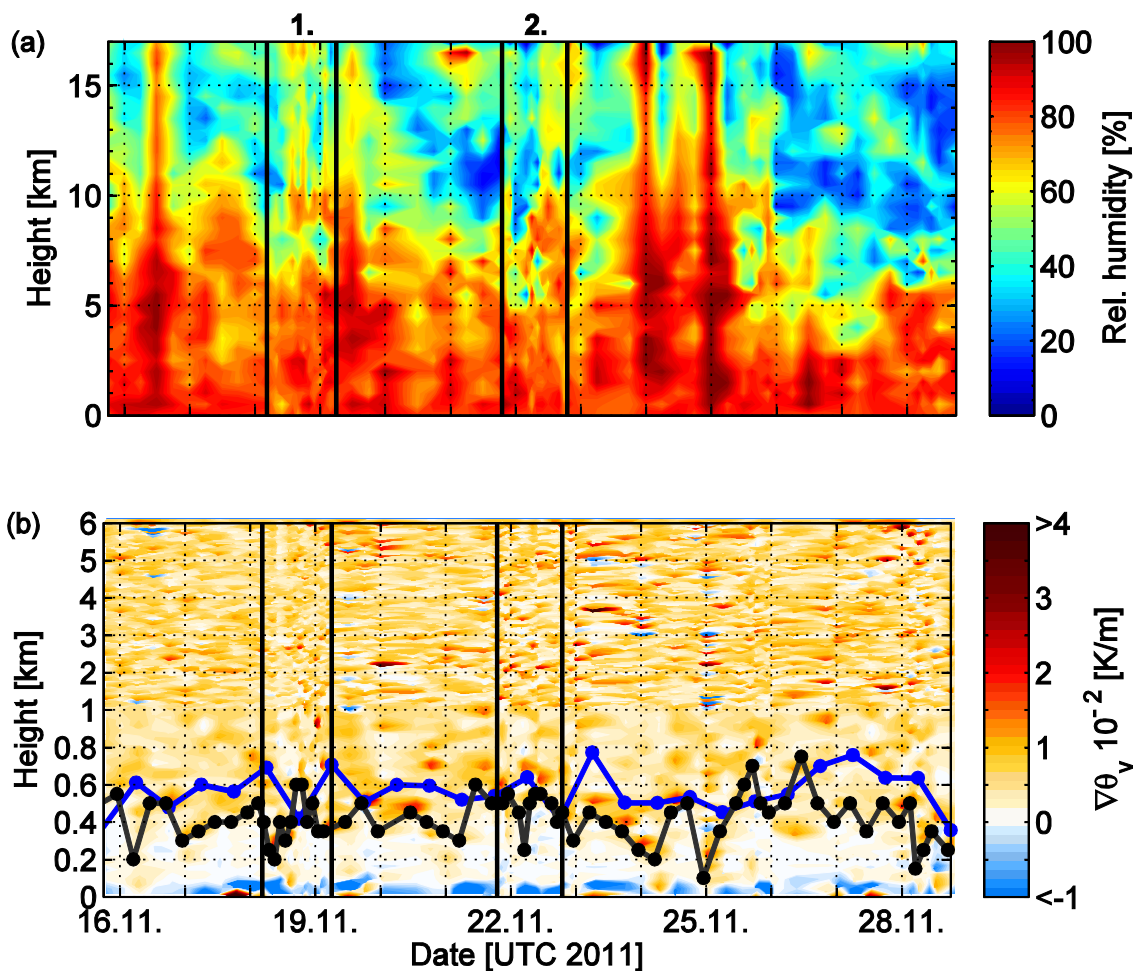
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906 Figure 2 a-b: Time series of (a) wind speed (blue) and wind direction (orange) and (b) surface
 907 air temperature (SAT, orange) and sea surface temperature (SST, blue) on the left scale, as
 908 observed on R/V SONNE. The temperature difference of SAT and SST (ΔT) is given on the
 909 right scale in (b). The ΔT of 0 K is drawn given by a dashed line. The shaded areas (grey) in
 910 the background show the 24 h stations. The data are averaged by a 10-minute running mean.

911

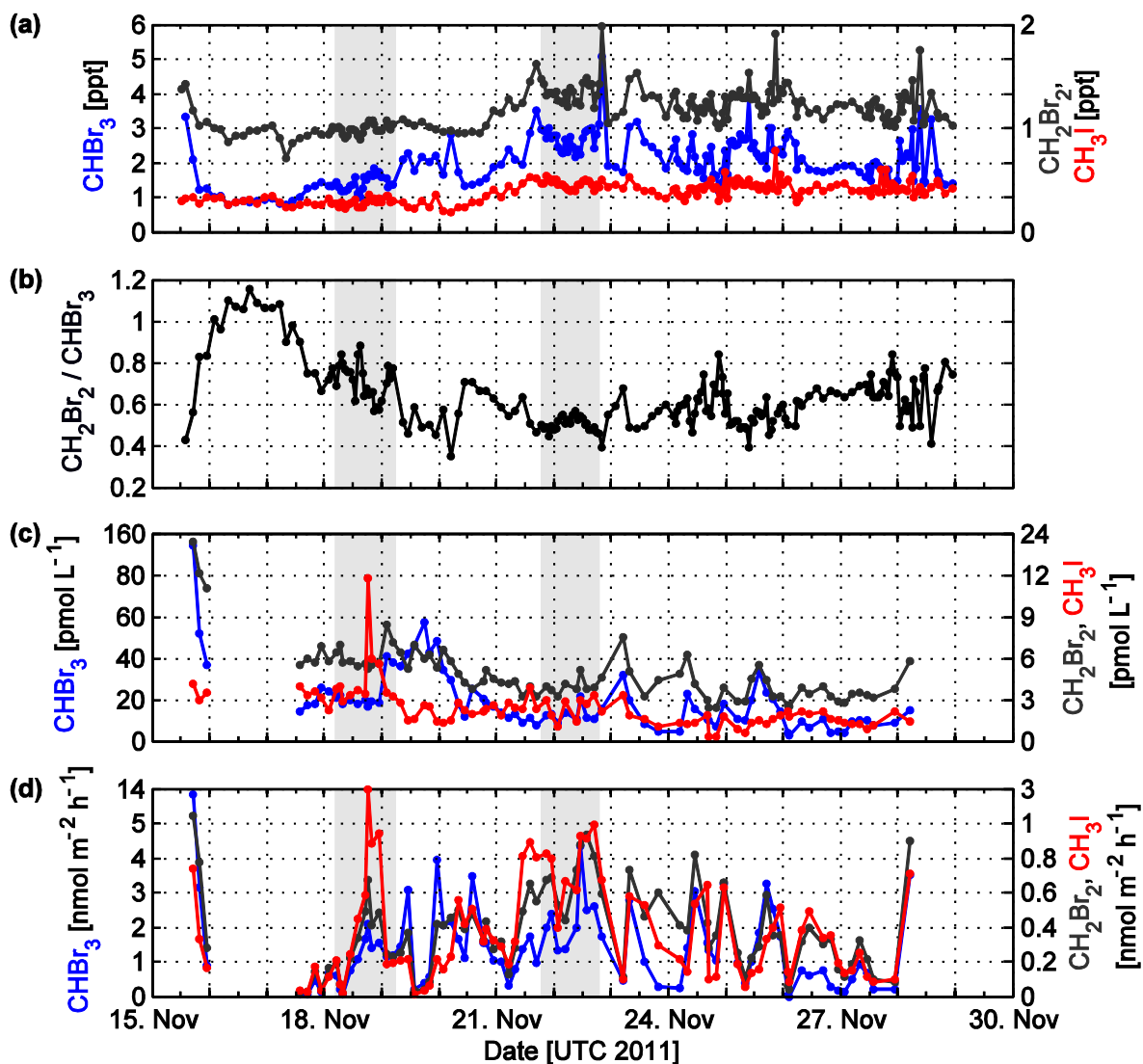


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914 | Figure 3: (a) Relative humidity **from**by radiosondes up to 17 km height, the mean cold point
 915 | tropopause level. The dashed lines and the two numbers above the figure indicate the two 24
 916 | h stations. (b) Virtual potential temperature gradient as indicator for atmospheric stability
 917 | (red for stable, white for neutral and blue for unstable) with MABL height **derived**
 918 | from radiosondes (black curve) and from ERA-Interim (blue curve). The y--axis is non-linear. The
 919 | lower 1 km is enlarged to display the stability around the MABL height. The vertical lines
 920 | and the two numbers above the figures indicate the two 24 h stations.

921



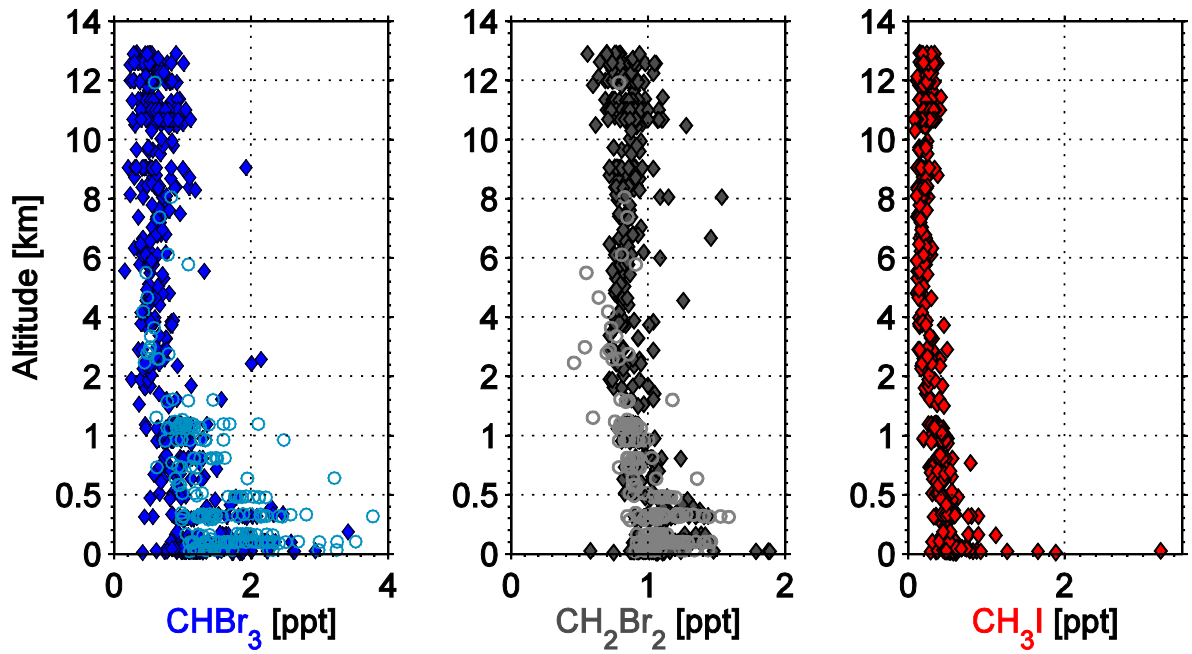
922

923 Figure 4 a-d: (a) Atmospheric mixing ratios of CHBr_3 (blue), CH_2Br_2 (dark grey) and CH_3I
 924 (red) measured on R/V SONNE. (b) Concentration ratio of CH_2Br_2 and CHBr_3 on R/V
 925 SONNE. (c) Water concentrations of CH_3I , CHBr_3 and CH_2Br_2 measured on R/V SONNE.
 926 (d) Calculated emissions of CHBr_3 , CH_2Br_2 and CH_3I from atmospheric and water samples
 927 measured on R/V SONNE. The two shaded areas (light grey) in the background show the 24
 928 h stations. Y-axis for (c) and (d) are non-linear.

929 Table 1: Mean \pm standard deviation and range of atmospheric mixing ratios observed on R/V
 930 SONNE (195 data points) and R/A FALCON (GhOST-MS with 513 and WASP GC/MS with
 931 202 data points) in the MABL and the FT, water concentrations observed on R/V SONNE
 932 and the computed sea – air fluxes. MABL and FT mixing ratios on R/A FALCON are
 933 adopted from Sala et al. (2014) and Tegtmeier et al. (2013). The R/A FALCON MABL
 934 height was analysed to be 450 m (Sala et al., 2014).

			CHBr ₃	CH ₂ Br ₂	CH ₃ I	
Atmosph. mixing ratios [ppt]	R/V SONNE		2.08 \pm 1.36 [0.79 – 5.07]	1.17 \pm 0.19 [0.71 - 1.98]	0.39 \pm 0.09 [0.19 – 0.78]	
	R/A FALCON	GhOST	MABL	1.43 \pm 0.53 [0.42 – 3.42]	1.19 \pm 0.21 [0.58 – 1.89]	0.59 \pm 0.30 [0.29 - 3.23]
		GhOST	FT	0.56 \pm 0.17 [0.16 – 2.15]	0.87 \pm 0.12 [0.56 – 1.54]	0.26 \pm 0.11 [0.08 – 0.80]
	WASP	WASP	MABL	1.90 \pm 0.55 [0.99 – 3.78]	1.15 \pm 0.14 [0.85 – 1.59]	/
		WASP	FT	1.17 \pm 0.50 [0.43 – 3.22]	0.88 \pm 0.14 [0.46 – 1.36]	/
Water concentrations [pmol L ⁻¹]			19.94 \pm 17.90 [2.80 – 136.91]	4.99 \pm 2.59 [2.43 – 21.82]	3.82 \pm 2.43 [0.55 – 18.84]	
Sea – air flux [pmol m ⁻² h ⁻¹]			1486 \pm 1718 [-8 – 13149]	405 \pm 349 [16 - 2210]	433 \pm 482 [13 - 2980]	

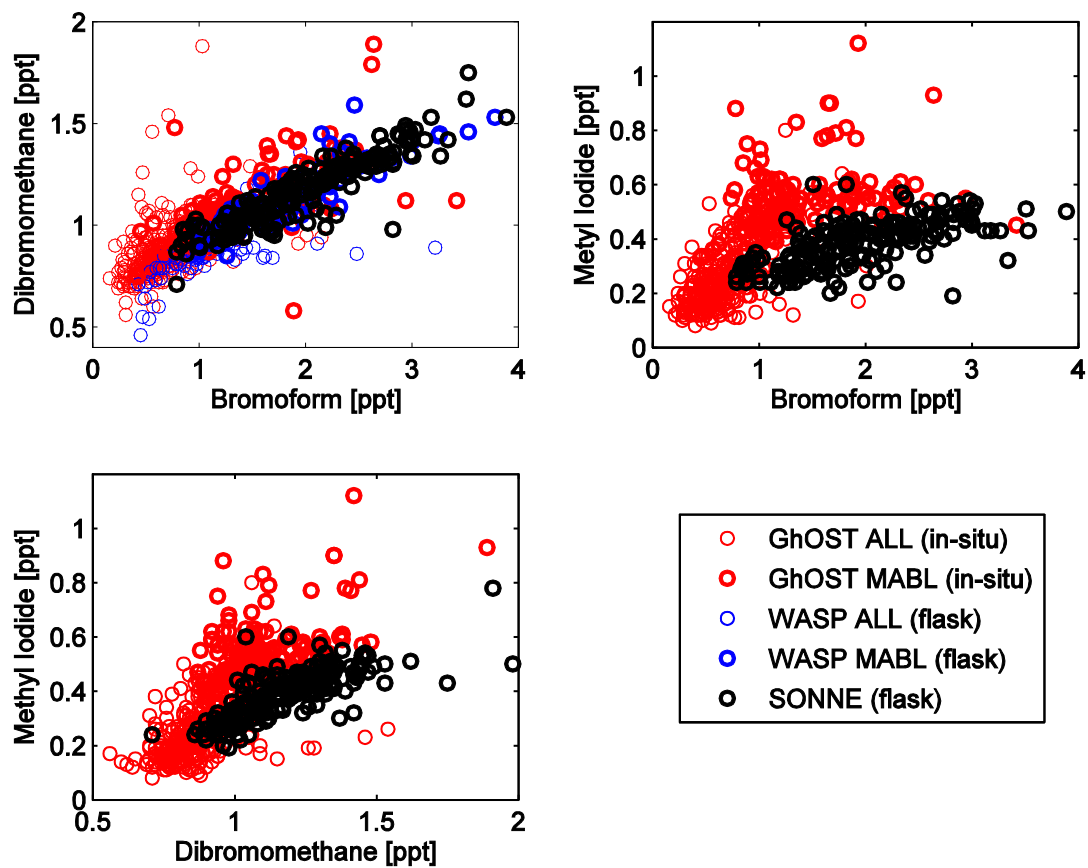
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937 Figure 5: Vertical distribution CHBr₃ (blue), CH₂Br₂ (grey) and CH₃I (red) mixing ratios
 938 measured in-situ by GhOST (diamonds) and with flasks by WASP (circles) on R/A
 939 FALCON. CH₃I was only measured in-situ by GhOST. The lower 2 km are non-linear
 940 displayed.

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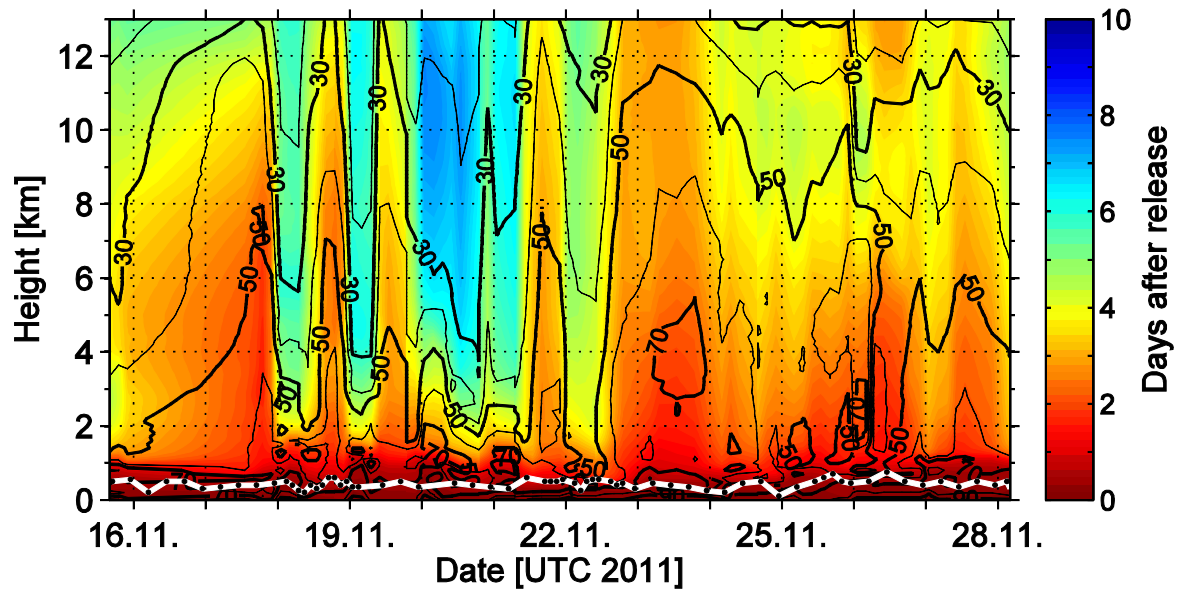
943 Figure 6: Correlation of bromoform and CH_2Br_2 (upper left), bromoform and CH_3I (upper
 944 right), and CH_2Br_2 and CH_3I (lower left) from GhOST and WASP for all heights (ALL) and
 945 only within the MABL (MABL) and from R/V SONNE.

946 Table 2: Mean atmospheric mixing ratios of CHBr_3 , CH_2Br_2 and CH_3I observed on R/V
 947 SONNE and R/A FALCON during two case studies on 19.11.2011 at 3.2° N and 112.5° E
 948 and on 21.11.2011 at 4.6° N and 113.0° E. During the two meetings two (one) measurements
 949 have been taken by R/V SONNE, 20 (5) measurements on R/A FALCON by GhOST and 17
 950 (21) by WASP.

		CHBr_3 [ppt]	CH_2Br_2 [ppt]	CH_3I [ppt]
November 19, 2011	R/V SONNE	1.37	0.99	0.29
	R/A FALCON: GhOST / WASP	1.02 / 1.37	0.94 / 1.03	0.45 / -
November 21, 2011	R/V SONNE	2.05	1.08	0.28
	R/A FALCON: GhOST / WASP	1.63 / 2.00	1.31 / 1.08	0.82 / -

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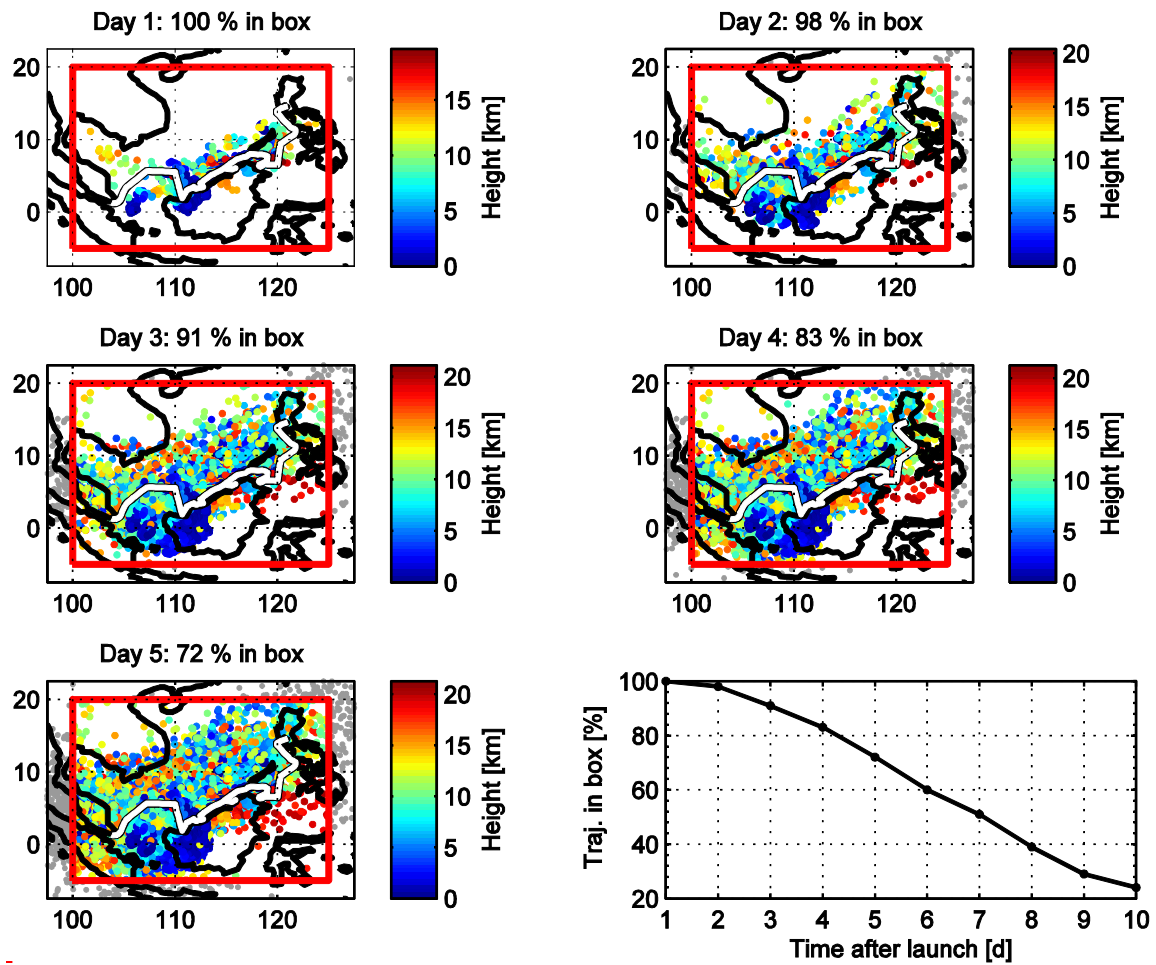
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954 Figure 7: Forward trajectory runs along the cruise track with FLEXPART using ERA-Interim
 955 data. The black contour lines show the mean amount of trajectories (in %) reaching this
 956 height within the specific time (colour shading). The white line indicates the radiosonde
 957 MABL height.

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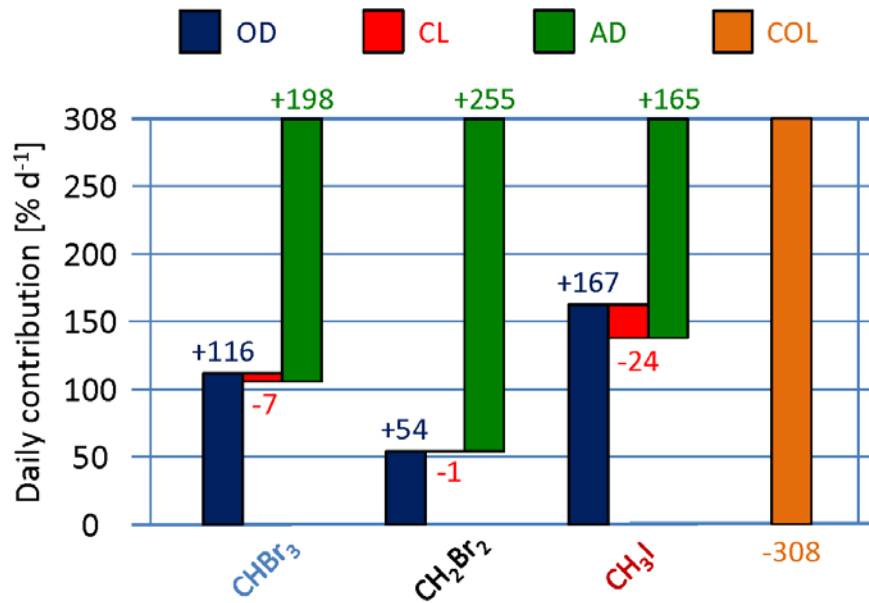
Figure 8: Horizontal distribution, altitude and amount of trajectories over time during the cruise. The red box represents the South China and Sulu Seas area. The lower right plot shows the amount of trajectories that remain in the box with time from all trajectory releases.

964 Table 3: Mean \pm standard deviation of Oceanic Delivery (OD), CONvective Loss (COL),
 965 Chemical Loss (CL), Advective Delivery (AD), Oceanic Delivery Ratio (ODR), Chemical
 966 Loss Ratio (CLR), and Advective Delivery Ratio (ADR) for CHBr_3 , CH_2Br_2 and CH_3I .

	OD [% d ⁻¹]	COL [% d ⁻¹]	CL [% d ⁻¹]	AD [% d ⁻¹]	ODR	CLR	ADR
CHBr₃	116.4	-307.6		198.2	0.45	-0.03	0.58
	\pm	\pm	-7.1	\pm	\pm	\pm	\pm
	163.6	124.3		199.7	0.55	0.01	0.55
CH₂Br₂	54.2	-307.6		254.6	0.20	-0.00	0.80
	\pm	\pm	-1.2	\pm	\pm	\pm	\pm
	66.7	124.3		131.9	0.21	0.00	0.21
CH₃I	166.5	-307.6		165.2	0.74	-0.09	0.35
	\pm	\pm	-24.0	\pm	\pm	\pm	\pm
	185.8	124.3		242.3	1.05	0.04	1.02

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970 Figure 8: Average budgets of the Oceanic Delivery (OD, blue), Chemical Loss (CL, red),
 971 Advective Delivery (AD, green) and Convective Loss (COL, orange) of CHBr₃, CH₂Br₂ and
 972 CH₃I in the Marine Atmospheric Boundary Layer (MABL).

973

974 | Table 4: Mean \pm standard deviation of observed ~~Volume-Mixing-Ratios~~ mixing ratios in the
 975 MABL on R/V SONNE (VMR_{MABL}) versus the amount of VMR originating from oceanic
 976 emissions (VMR_{ODR}), chemically degraded according to the specific lifetime (VMR_{CLR}),
 977 originating from advection (VMR_{ADR}) and the Flux from the MABL into the FT ($\text{Flux}_{\text{MABL-FT}}$)
 978 for CHBr_3 , CH_2Br_2 and CH_3I .

	VMR_{MABL} [ppt]	VMR_{ODR} [ppt]	VMR_{CLR} [ppt]	VMR_{ADR} [ppt]	$\text{Flux}_{\text{MABL-FT}}$ [$\text{pmol m}^{-2} \text{hr}^{-1}$]
CHBr_3	2.08	0.89	-0.06	1.18	4240
	\pm	\pm	\pm	\pm	\pm
	1.36	1.12	0.04	1.20	1889
CH_2Br_2	1.17	0.25	-0.01	0.92	2419
	\pm	\pm	\pm	\pm	\pm
	0.19	0.26	0.00	0.27	929
CH_3I	0.39	0.28	-0.04	0.13	865
	\pm	\pm	\pm	\pm	\pm
	0.09	0.40	0.02	0.37	373

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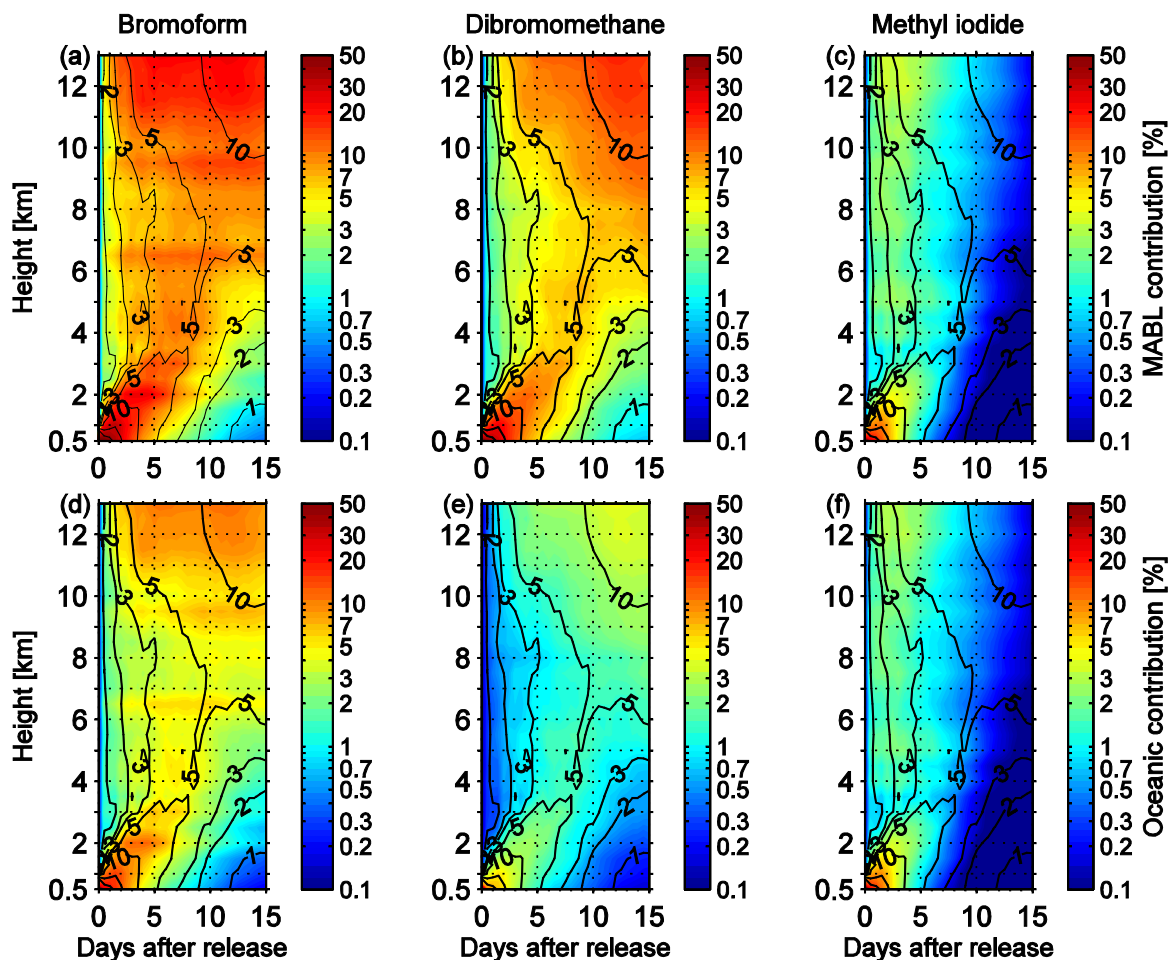
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981 Table 5: Correlation coefficients between wind speed and VSLS MABL mixing ratios
 982 (VMR_{MABL}), the Oceanic Delivery (OD), the CONvective Loss to the FT (COL), the
 983 Advective Delivery (AD), computed as the residual of OD, and the mixing ratios originating
 984 from the OD (VMR_{ODR}) and from the AD (VMR_{ADR}) of each compound. Bold numbers are
 985 significant at the 95 % level (p-value).

<i>Wind speed</i>	CHBr_3	CH_2Br_2	CH_3I
VMR_{MABL}	0.55	0.57	0.56
OD	0.31	0.48	0.52
COL	-0.33		
AD	-0.46	0.56	-0.57
VMR_{ODR}	0.52	0.72	0.62
VMR_{ADR}	-0.17	-0.31	-0.49

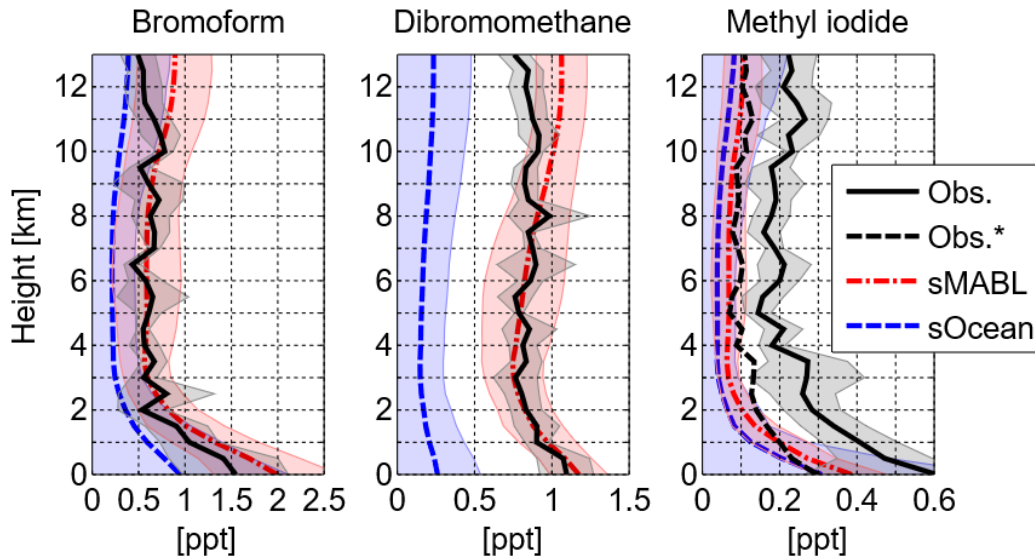
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989 Figure 9: Mean MABL air contribution (a – c) and oceanic contribution (d – f) to observed
 990 FT mixing ratios observed by R/A FALCON for three VSLs. The black contour lines show
 991 the mean portion of MABL air masses in the FT [%], the colours show the oceanic
 992 contribution to the observed compounds in the FT at specific height and day after release [%]
 993 including chemical degradation, ~~the loss out of the South China Sea area with time~~ and the
 994 vertical density driven extension of MABL air masses. The scale of the coloured contour is
 995 logarithmic.



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997 Figure 10: Mean FT mixing ratios (solid lines) and 1 standard deviation (shaded areas) from
 998 in-situ and flask observations on R/A FALCON (Obs., black) versus simulated mean FT
 999 mixing ratios from MABL air (sMABL, red) and oceanic emissions (sOcean, blue) observed by R/V SONNE. R/A FALCON in-situ observations have been adjusted for
 1000 CH₃I (Obs.*, dashed black) according to measurement deviations during the meetings of
 1001 R/V SONNE and R/A FALCON (Table 2; Section 4.3).