



Consistent response
of Indian summer
monsoon to Middle
East dust

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This discussion paper is/has been under review for the journal Atmospheric Chemistry and Physics (ACP). Please refer to the corresponding final paper in ACP if available.

Consistent response of Indian summer monsoon to Middle East dust in observations and simulations

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Received: 31 March 2015 – Accepted: 23 May 2015 – Published: 11 June 2015

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Published by Copernicus Publications on behalf of the European Geosciences Union.

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Abstract

The response of the Indian summer monsoon (ISM) circulation and precipitation to Middle East dust aerosols on sub-seasonal timescales is studied using observations and the Weather Research and Forecasting model with chemistry (WRF-Chem). Satellite data shows that the ISM rainfall in coastal southwest India, central and northern India, and Pakistan are closely associated with Middle East dust aerosols. The physical mechanism behind this dust–ISM rainfall connection is examined through ensemble simulations with and without dust emission. Each ensemble includes 16 members with various physical and chemical schemes to consider the model uncertainties in parameterizing shortwave radiation, the planetary boundary layer, and aerosol chemical mixing rules. Experiments show that dust aerosols increase rainfall by about 0.44 mm day^{-1} ($\sim 10\%$) in coastal southwest India, central and northern India, and northern Pakistan, a pattern consistent with the observed relationship. The ensemble mean rainfall response over India shows much stronger spatial correlation with the observed rainfall response than any of the ensemble members. The largest modeling uncertainties are from the boundary layer schemes, followed by shortwave radiation schemes. In WRF-Chem, the dust AOD over the Middle East shows the strongest correlation with the ISM rainfall response when dust AOD leads rainfall response by about 11 days. Further analyses show that the increased ISM rainfall is related to the enhanced southwesterly flow and moisture transport from the Arabian Sea to the Indian subcontinent, which are associated with the development of an anomalous low pressure system over the Arabian Sea, the southern Arabian Peninsula, and the Iranian Plateau due to dust-induced heating in the lower troposphere (800–500 hPa). This study demonstrates a thermodynamic mechanism that links remote desert dust emission in the Middle East to the ISM circulation and precipitation variability on sub-seasonal timescales, which may have implications for ISM rainfall forecasts.

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Our study uses observations and model experiments to understand the discrepancies among the above studies. We have three research questions. First, in what areas is rainfall sensitive to Middle East dust aerosols? The AOD–ISM rainfall relationship based on observations can provide a baseline for model evaluations. Secondly, how are the observed AOD–ISM interactions represented in the Weather Research and Forecasting model (Skamarock et al., 2005) with chemistry (WRF-Chem) (Grell et al., 2005) and how do the modeling uncertainties affect our conclusions? In V2014, 19 ensemble simulations were conducted during a short period (10 days). In S2015, three ensemble simulations were created by perturbing the boundary conditions. In our study, 16 pairs of ensemble simulations are conducted using a perturbed physics and chemistry ensemble (PPCE) method during the boreal summer 2008, a period with strong dust emission. We believe that by using PPCE members we can better capture the uncertainties in the monsoon response to dust because the AOD–ISM rainfall hypothesis is based on both chemical properties (e.g. aerosol chemical mixing rules) of dust and their impact on atmospheric physical processes (e.g. radiation and circulations) (McFiggans et al., 2006). The dust-induced impact is then examined by the ensemble mean differences. Finally, is the 13-day maximum cross-correlation found in observations in J2014 captured by WRF-Chem? This question is critical because if the AOD–ISM rainfall hypothesis is true, AOD must lead the ISM rainfall in the model.

2 Model and experiment

2.1 Model

WRF-Chem simultaneously simulates the evolution of trace gases and aerosols and their interactions with meteorological fields. It incorporates the second-generation Regional Acid Deposition Model (RADM2) gas-phase chemical mechanism (Stockwell et al., 1997) and the Modal Aerosol Dynamics Model for Europe (MADE) primary aerosol scheme (Schell et al., 2001) coupling the Secondary Organic Aerosol Model

2.2 Emissions

2.2.1 Dust emission

The Goddard Chemistry Aerosol Radiation and Transport (GOCART) dust emission scheme (Ginoux et al., 2001) coupled with the MADE-SORGAM aerosol scheme is used to simulate dust emission. In GOCART, dust emission is calculated based on wind speed and an erodibility map (Prospero et al., 2002; Zender et al., 2003b) as

$$G = CSs_p u_{10m}^2 (u_{10m} - u_t) \quad (1)$$

where G is dust emission flux ($\mu\text{g m}^{-2} \text{s}^{-1}$), C is a dimensionless empirical proportionality constant, S is soil erodibility of potential dust source regions (Fig. 1b), s_p is a fraction of mass of each size bin of dust emission, u_{10m}^2 is wind speed at 10 m above the surface and u_t is the wind speed threshold under which dust emission does not occur. GOCART represents dust emission in 5 bins, with averaged radii of 0.73, 1.40, 2.40, 4.50, and 8.00 μm . The computed dust emission are divided into the accumulation and coarse modal modes for MADE-SORGAM to match the representation of particle size distribution in the aerosol scheme.

2.2.2 Other emissions

The anthropogenic emissions for WRF-Chem come from the mixture of the Reanalysis of the Tropospheric chemical composition emissions inventory (<https://verc.enes.org/>), Emission Database for Global Atmospheric Research (<http://www.mnp.nl/edgar/introduction>), and GOCART. These anthropogenic emissions include 20 gases and 3 aerosol species, including BC, OC, and sulfate. Background emissions are from GOCART dataset, including nitrate, H_2O_2 hydroxyl radical, and DMS. Biogenic emissions are calculated online using the Model of Emissions of Gases and Aerosols from Nature (Guenther et al., 2012). Biomass burning emissions are obtained from

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the Global Fire Emissions Database version 2 with 8-day temporal resolution (<http://www.globalfiredata.org/index.html>).

2.3 Experiment design

In this study, WRF-Chem 3.5 is configured over the Middle East and the ISM region (−9.68° S–49.60° N, 20.41–109.87° E, see Fig. 1a) with 160 × 120 grid points centered over the AS (23° E, 65° N), at a 54 km horizontal resolution and 30 vertical layers up to 50 hPa. This domain covers major dust source regions represented by erodibility in the model, as shown in Fig. 1b. A few key areas are selected to examine the relationship between Middle East dust and the ISM rainfall. The DST box is chosen to include the heavy dust loading areas, i.e. the AS, the AP, the IP, Afghanistan, and Turkmenistan. The WHI box and CNI boxes are chosen to represent the rainfall over the whole India and central and northern Indian, respectively.

All simulations cover a 104-day period from 20 May 2008 to 31 August 2008 without nudging. The first 12 days for each integration are discarded as “spin up” to reduce the impact of initial conditions, and the analysis focuses on the monsoon season from 1 June 2008 to 31 August 2008. Meteorological fields from the European Centre for Medium-Range Weather Forecasts (ECMWF) Interim Reanalysis (ERA-I) global re-analysis data are prescribed as lateral and lower boundary conditions (e.g. SST) and initial conditions.

WRF-Chem provides multiple options for physical and chemical parameterizations. For physical parameterizations, we used the Noah land surface model (Chen et al., 2001), the RRTMG SW and longwave (LW) radiation scheme, Yonsei University (YSU) planetary boundary layer (PBL) scheme (Hong et al., 2006), and Lin’s second-moment microphysical scheme (Lin et al., 1983). An updated version of the Grell-Devenyi (Grell and Devenyi, 2002) cumulus parameterization scheme is used, which includes feedback from the parameterized convection to the atmospheric RRTMG radiation scheme and Fast-J photolysis scheme. For the chemical parameterizations, we used RADM2 gas-phase chemistry, MOSAIC-SORGAM aerosol chemistry with aqueous reactions,

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and Fast-J photolysis scheme. Table 1 summarizes the model schemes used in this study.

To understand how the model uncertainties and errors affect our results, additional alternative physical and chemical parameterization schemes are used to create ensemble simulations. V2014 and S2015 created ensemble members by perturbing initial and boundary conditions, respectively. In this study, two groups of simulations were designed based on the presence and absence of dust emission. The reference group considers all aerosol forcing (including mineral dust, sea salt, biomass burning, biogenic emission, and anthropogenic emission; ALLF) and the perturbed group is identical to the first group but without dust emission (NDST). Within each group, 16 ensemble members were created using the PPCE method, because we are more interested in how differences in the aerosol chemical mixing rules, aerosol diffusion in the atmospheric boundary layer, and radiation schemes may affect the simulations of the ISM rainfall and variability.

Four different aerosol chemical mixing rules are used to calculate the aerosol optical properties: volume approximation, Maxwell-Garnett approximation, exact volume, and exact Maxwell-Garnett schemes (Fast et al., 2006; Barnard et al., 2010). The volume approximation assumption calculates refractive indices based on the volume average of each aerosol species. The Maxwell-Garnett method assumes a random distribution of black carbon in spherical particles. Both of the volume and Maxwell-Garnett schemes call the full Mie calculation only at the first time step (Ghan et al., 2001). However, the exact volume and exact Maxwell-Garnett schemes call the full Mie calculation at each time step.

Two SW radiation schemes-RRTMG and Goddard, and two PBL schemes-YSU and Bougeault-Lacarrère (BouLac) are employed. The sub-grid cloud parameterization in RRTMG can simulate interactions between aerosol radiative forcing and sub-grid clouds. YSU and BouLac represent two types of PBL schemes-turbulent kinetic energy and first-order closure schemes (Shin and Hong, 2011). Totally, there are 32 ensemble members, which are comprised of 2 PBL \times 2 SW \times 4 composition \times 2 dust

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options as detailed in Table 2. Note that the LW radiation induced uncertainties are not considered in this study, because only the RRTMG LW radiation scheme is coupled with the aerosol scheme.

3 Datasets

3.1 AOD

The Moderate-resolution Imaging Spectroradiometer (MODIS) instrument aboard the National Aeronautics and Space Administration (NASA) Terra and Aqua platforms is uniquely designed to observe and monitor atmospheric trace gases, clouds, and tropospheric aerosols. MODIS provides two kinds of AOD data, “dark target” and “deep-blue” daily and monthly AOD at the wavelength of 550 nm with resolution of $1^\circ \times 1^\circ$ from 1 March 2000 to the present (<http://modis-atmos.gsfc.nasa.gov/>). The “dark target” algorithm provides best results over the oceans and on land with low surface albedo; the “deep-blue” algorithm retrieves AOD over regions with bright surfaces, such as desert (Hsu et al., 2004). In our study, both kinds of AOD are combined to expand the spatial coverage by simply assigning the “deep-blue” AOD to the grid if the “dark target” AOD is missed. In this way, the global AOD is available over both oceans and land. Because of a broad swath of 2330 km, MODIS images the entire earth every 1–2 days.

The Multi-angle Imaging Spectroradiometer (MISR) instrument provides detailed aerosol properties at the global scale. MISR onboard Terra, NASA’s first Earth Observing System spacecraft, is designed to improve our understanding of the regional and global impacts of different types of atmospheric particles and clouds on climate (Diner et al., 1998). With nine cameras, MISR views Earth in nine different directions, and each piece of Earth’s surface below is successively imaged by all nine cameras, in each of four wavelengths (blue, green, red, and near-infrared). This specific feature of MISR can help estimate aerosol particle size and composition with unprecedented accuracy. Based on the particle size information, the aerosol’s effects on climate caused

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by natural sources and human activities can be isolated. The swath for MISR is only 360 km, which gives MISR a longer global span time of 9 days.

The ECMWF Monitoring Atmospheric Composition and Climate (MACC) ($0.5^\circ \times 0.5^\circ$) reanalysis data provide partial AOD of various aerosol types (Benedetti et al., 2009).

MACC assimilates MODIS dark target production collection 5, which does not include AOD retrievals over the bright surface using the “deep-blue” algorithm.

3.2 Precipitation

Four precipitation datasets are used. The Tropical Rainfall Measuring Mission (TRMM) ($0.25^\circ \times 0.25^\circ$) monthly 3B43 (version 7) and daily 3B42 (version 7) precipitation datasets are employed (Huffman et al., 2007). These two products combine multiple independent satellite precipitation estimates and Global Precipitation Climatology Centre rain gauge analysis. The Global Precipitation Climatology Project (GPCP) ($1^\circ \times 1^\circ$) monthly and daily (version 2) rainfall datasets are produced by combining multiple satellite retrieved precipitation and surface precipitation gauge analyses (Huffman et al., 2001). The Climate Prediction Center Merged Analysis of Precipitation (CMAP) ($2.5^\circ \times 2.5^\circ$) monthly and pentad rainfall datasets are produced based on five kinds of satellite retrievals (Xie and Arkin, 1997). The National Oceanic and Atmospheric Administration PRECipitation REConstruction over land (PREC/L) ($0.5^\circ \times 0.5^\circ$) monthly rainfall datasets are constructed by precipitation gauge observations (Chen et al., 2002).

3.3 Reanalyses

ERA-I ($0.5^\circ \times 0.5^\circ$) (Rienecker et al., 2011) is adopted to provide WRF-Chem with lateral and lower boundary conditions, as well as initial conditions. The Modern Era-Retrospective Analysis for Research and Applications (MERRA) ($1/2^\circ \times 2/3^\circ$) (Dee et al., 2011) is also used for comparison with ERA-I and model evaluation in terms of geopotential height (GPH) and circulations.

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4 Evaluation of ISM and AOD simulated by WRF-Chem

Figure 2a–c illustrate the spatial patterns of the observed and modeled rainfall averaged for June–July–August (JJA) 2008. During the ISM season, TRMM and GPCP observed heavy rainfall in coastal southwest India (CSWI) and CNI. The WRF-Chem ensemble mean rainfall in 16 ALLF members (Fig. 2c) shows a spatial pattern quite consistent with that of TRMM and GPCP. Note that the model overestimates rainfall in CSWI and underestimates rainfall in CNI. Similar rainfall difference between model simulations and observations were also found in both RCM (e.g. Solmon et al., 2015) and GCM (e.g. Levine and Turner, 2012) studies. The underestimated rainfall in CNI can be partly attributed to the lack of representation of agricultural irrigation in the model. Intensive irrigation activities occurring during JJA over the IGP can increase local evapotranspiration, and thus increase rainfall (Douglas et al., 2009; Guimberteau et al., 2012). Figures 2d–2f show the GPH and circulation at 850 hPa from reanalysis and WRF-Chem. The ISM system is featured by strong cross-equator southerly winds in the tropical Indian Ocean and southwesterly winds in the lower troposphere in the AS and the Bay of Bengal. Another ISM feature is the deeper low-pressure centered over north India and the IP. In general, the model can capture the ISM features quite well.

A reliable representation of the spatial distribution of dust concentration in the model is essential for examining dust impacts on the ISM rainfall. The modeled AOD at 300, 400, 600, and 999 nm are converted to 550 nm using the Ångström exponent and evaluated using satellite retrieved AOD, which is usually at 550 nm. By tuning the empirical proportionality in Eq. (1) against satellite AOD, the spatial pattern of modelled AOD is quite consistent with multiple satellite retrievals and aerosol reanalysis results. Figure 3 demonstrates that WRF-Chem captures the observed high dust loading in the AP, the Thar Desert, and the IP. However, the model underestimates AOD over the northern AS and overestimates AOD in the southern AP in comparison to the other four datasets. Note that MISR, MODIS Aqua, and MACC show much higher AOD over

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the AS than the southern AP, whereas modeled AOD shows the opposite. This discrepancy between the modeled and satellite data could be attributed to two potential contributors. First, the assumption regarding dust mass distribution on dust particle size in the MADE-SORGAM aerosol scheme is not suitable for Middle East dust aerosols.

In MADE-SORGAM, only 7% of the mass of total dust emission is assigned to the accumulation mode, whereas the other 93% percent goes into the coarse mode; consequently, most dust emission is deposited quickly in the dust source regions in the AP and only very little is transported long-distance to the AS. However, in the Mineral Dust Entrainment and Deposition model (Zender et al., 2003a), 17% of the mass of dust emission is assigned to the accumulation mode. Note also that the underestimation of fine particles reduces dust-induced atmospheric heating because fine particles absorb 3 to 5 times the solar radiation absorbed by coarse particles (Mahowald et al., 2014). Secondly, the model does not adequately represent the impact of relative humidity on AOD calculation. Increased relative humidity can lead to higher AOD because more water vapor can be taken up by dust particles, an effect known as aerosol humidification (Myhre et al., 2007).

5 Observed relationship between dust and ISM rainfall

The observed relationship between AOD and ISM rainfall is studied using regression analysis. Figure 4 shows the spatial patterns of AOD regressed on the area-averaged ISM rainfall in WHI using their JJA monthly anomalies during 2000 and 2013. Figure 4a shows the regressed AOD using MISR and NOAA observations. Positive anomalies of ISM rainfall in WHI are associated with heavy aerosol loading over the northern AS, the southern AP, and the IP. This spatial pattern of regressed AOD persists or becomes stronger in other AOD and rainfall datasets, as shown in Fig. 4b–d. Over northeastern India, dust is negatively correlated with rainfall because local dust is removed through wet deposition. This spatial pattern is consistent with the modeled atmospheric heating pattern induced by dust in Fig. 12a, which will be discussed later.

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To compare this observed dust-ISM rainfall relationship with model simulations, Fig. 5 shows the precipitation regressed on the area-averaged AOD in the DST region. In general, the various datasets show a consistent spatial pattern. The positive response of rainfall to AOD is primarily located in the IGP, the central India, and CSWI, while a weak negative response is seen in southeast India. This observed north-positive and southeast-negative correlation patterns differ from the results of V2014 (Fig. 4b in V2014) and is almost opposite to those of S2015 (Fig. 5b in S2015), but it is very similar to our observation-based analysis in J2014 (Fig. 2c in J2014). See Table 3 for a summary.

The regression analyses of AOD–ISM rainfall relationship based on observations provide a baseline for evaluating the model results.

6 Modeled ISM rainfall response to dust

The ISM rainfall response to Middle East dust is represented by the ensemble mean differences in rainfall from 16 ALLF and 16 NDST simulations. Figure 6 shows the spatial pattern of rainfall response averaged during JJA. In general, rainfall increases over most of India with an average magnitude of 0.44 mm day^{-1} ($\sim 10\%$) in WHI. The largest rainfall increases occur in CNI, northern Pakistan, and CWSI. Note that the spatial pattern of rainfall increase closely follows the topography along CSWI and the Tibetan Plateau, indicating that increased rainfall can be mainly related to large-scale circulation changes. Conversely, decreased rainfall is simulated in southeastern and northeastern India. The spatial pattern of the modeled rainfall response is generally consistent with the observed AOD–ISM rainfall relationship in Fig. 5 and Fig. 2c in J2014 (Table 3).

Figure 7 shows the centered spatial correlations between the regressed rainfall pattern in Fig. 5c and rainfall responses in each of the 16 ensemble pairs (ALLF-NDST) as well as their ensemble mean. It can be seen that 15 out of 16 members show positive spatial correlations between the modeled rainfall response and regressed rainfall in

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observations, with a magnitude of 0.1 to 0.5, which indicates that most of the members can capture the observed spatial patterns of dust-induced rainfall changes. Interestingly, the rainfall response from the ensemble mean shows a spatial correlation of about 0.6, much larger than any of the ensemble members. This indicates that the ensemble mean may cancel out and reduce model errors raised from various parameterization schemes.

Figure 7 also illustrates that ensemble members 2, 6, 10, and 14 show very small correlation coefficients. These schemes share the same SW radiation and PBL schemes—Goddard and YSU. The low correlation coefficients are attributed mainly to the BouLac PBL scheme, because ensemble members 4, 8, 12, and 16 show high correlation coefficients and share the same SW radiation schemes but have a different PBL scheme—BouLac from members 2, 6, 10, and 14. Furthermore, the ensemble members using RRTMG SW radiation schemes generally show higher correlation coefficients than those using the Goddard SW radiation scheme, while those using the different aerosol chemical mixing rules show very little differences in correlation coefficients.

Figure 8 shows the scatter plot of the area-averaged rainfall in WHI and CNI in each ALLF ensemble member (y axis) against NDST members (x axis). In WHI, all members are located above or on the 1:1 line except the number 2, indicating that dust-induced rainfall increases are quite robust. Along with the number 2, which shows decreased rainfall, the numbers 4, 6, 8, 10, and 16 show very weak rainfall increase. These simulations share the same Goddard SW scheme. The situation is similar in CNI, but with more members showing decreased rainfall than in WHI. Therefore, we conclude that the largest effects on our simulation results are from the PBL scheme, followed by the SW radiation scheme, and the aerosol chemical mixing rule has a very weak effect.

Figure 9 shows the time series of WRF-Chem simulated daily rainfall in each ensemble member and the ensemble mean of rainfall response and AOD in ALLF simulations. In general, the model can capture the temporal variation of rainfall in July and August, with two peaks in early July and the first half of August in both the model runs and observations. The model has a notable low bias in June, which is larger in CNI than in

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positive). The negative SW radiative forcing is very strong over water bodies, such as the Red Sea, the AS, the Persian Gulf, and the Caspian Sea, with a magnitude around 30 W m^{-2} . However, Fig. 11a shows a quite weak or even positive SW radiative forcing over regions with bright surfaces, such as in eastern North Africa, the AP, the IP, and the Taklimakan Desert. The contrasting spatial patterns of SW radiative forcing over water and land are due to the high surface albedo contrast between them. Due to the high albedo of dust layers, when dust plumes reach the ocean, they reflect and scatter more SW radiation back to space than water bodies. Note that the small and sparsely distributed positive SW radiative forcing at the TOA is due to changes in clouds, which can be caused by dust-induced changes in atmosphere dynamics and cloud microphysics. In the atmosphere, dust aerosols can absorb SW radiation and heat the atmosphere. Figure 11b shows positive SW radiative forcing over the entire domain. This positive radiative forcing can attain a magnitude of 20 to 25 W m^{-2} in Iraq, the southern AP, the northern AS, and Pakistan. The land-ocean contrast in radiative forcing observed at the TOA is not seen within the atmosphere, because the spatial distribution of dust-induced atmospheric heating is mainly determined by dust concentration rather than surface properties. However, some studies have proposed that dust aerosols can absorb SW radiation more effectively over bright surfaces than over dark surfaces due to multiple reflections of SW radiation between overlying dust layers and bright land surfaces (Lau et al., 2006; Kuhlmann and Quaas, 2010). At the surface, dust aerosols block SW radiation from reaching the surface through scattering and absorption, which results in a surface cooling effect as shown in Fig. 11c. The maximum magnitude of the negative SW radiative forcing is about 30 W m^{-2} in the Red Sea, the southern AP, the northern AS, the Persian Gulf, and Pakistan.

For LW, at the TOA Fig. 11d shows that dust causes positive radiative forcing between 1 to 5 W m^{-2} . The dust layer can absorb surface-emitted LW radiation, and then reemit it back to the surface. Because the dust layer is cooler than the surface due to its higher altitude, it emits less LW radiation to space than the surface on dust-free days, resulting in positive radiative forcing at the TOA. In the atmosphere, due to the

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blocking effect of dust, less LW radiation is absorbed by the atmosphere above the dust layer, which results in an atmospheric cooling effect. Figure 11e illustrates that the LW cooling effect of dust can reach magnitudes of 10 W m^{-2} in dust source regions and decreases to 5 W m^{-2} in most other regions except for the northern Indian Ocean and the Bay of Bengal, where positive forcing is observed. At the surface, the blocking effect of dust results in a positive forcing, i.e. warming effect, of about 10 W m^{-2} over the AP and Indian subcontinent as shown in Fig. 11f. Note that the LW radiative forcing in the atmosphere and at the surface is stronger over land than over ocean. There are several reasons for this. First, the LW radiation interacts more efficiently with large particles than small particles. Large particles are mainly located in dust source regions and very few can be transported over long-distance to the ocean. Secondly, there is less water vapor between the surface and dust layer over land than over ocean. This strengthens the interactions between surface-emitted LW radiation and overlying dust aerosols. Third, the hotter land surface can emit more LW radiation than the cooler ocean surface.

It is obvious that the LW and SW radiative effects of dust have opposite signs, and the SW forcing has a much greater magnitude than LW forcing (Fig. 11a–f). Therefore, the dust net radiative forcing is dominated by the SW forcing (Fig. 11g–i). The area-averaged radiative forcing of dust over the whole domain is summarized in Table 4. By simple comparison of values in Table 4, we can conclude that a quarter to one-third of the SW radiative forcing is counterbalanced by LW radiative effects, which is consistent with previous studies (Huang et al., 2009).

7.2 Dust impact on circulation

Dust aerosols can change large-scale circulations through their surface cooling effects and atmospheric warming effects. The resultant circulation change depends on the net effect of the two.

Figure 12a illustrates the spatial pattern of the ensemble mean thickness differences in ALLF and NDST simulations between 800 and 500 hPa averaged for JJA 2008. Ac-

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cording to the hypsometric equation, the thickness between two isobaric surfaces is proportional to the mean temperature of the layer. Figure 12a illustrates the increased thickness over the AS and southeast AP with a magnitude of 25 to 30 m, equal to 2.5 to 3 K. The spatial pattern of the thickness differences generally follows the AOD spatial pattern (Fig. 3e). Figure 12b shows the vertical profiles of area-averaged atmospheric heating sources at all-sky conditions in DST. LW radiative forcing and sensible heating contribute to the atmospheric cooling from the surface to about 600 hPa. The SW radiative forcing is the only source of atmospheric warming effect, which is strongest near 950 hPa and diminishes to zero near 400 hPa. The latent heat shows little changes. Net atmospheric heating, which is the sum of LW, SW, sensible heat, and latent heat, demonstrates the atmospheric heating effect of dust from the near surface to 400 hPa except for an anomalous cooling effect at 900 hPa, which is caused by cloud effects.

Due to dust-induced atmospheric heating in the lower troposphere, a low-pressure system at 850 hPa can be observed (shown as contours in Fig. 13a) over the AS, Shamal wind regions, the IP, and the Caspian Sea and nearby regions. Associated with the low-pressure system is a convergence region centered over the north AS and western India at 850 hPa, illustrated by arrows in Fig. 13a. In this convergence area, the strengthened southwesterly winds transport more water vapor from the AS north-eastward to the Indian subcontinent. When moist airflows meet the mountains in CSWI, CNI, Tibetan Plateau, and north Pakistan, they are lifted, converged and cooled, which forms the orographic rainfall (Fig. 6). Additionally, the strengthened northwesterly winds over the AP can result in more dust emission over the AP and transport them from the AP to the AS, thus forming a positive feedback. Figure 13b shows that dust can also modulate the atmospheric circulation in the upper troposphere, e.g. 500 hPa. There are two dust-induced convergence regions at 500 hPa: the IP and Iraq, CNI and North China. The role of dust-induced upper troposphere heating and circulation changes in modulating ISM is not clear and needs to be addressed in future studies. Dai et al. (2013) showed that the south-north thermal contrast in the mid-upper troposphere is more important for the Asian monsoon than that in the lower troposphere.

7.3 Dust impact on moisture budget

By disturbing large-scale circulations, dust can also modulate the moisture budget of ISM. Contours in Fig. 14 show the spatial distribution of the ensemble mean of precipitable water differences in the entire atmospheric column in ALLF and NDST experiments during JJA 2008. Increased PW is simulated over the entire Indian subcontinent with maximum increases of 2 mm in CSWI and IGP. The increased PW is attributed to the strengthened moisture transport by southwesterly winds over the AS and southeasterly winds over the IGP, as indicated by arrows in Fig. 14. A minor PW increase occurs in east India around 20°N, which causes decreased rainfall in this area as shown in Fig. 6. Note that the increased PW is less than that of the increased rainfall in CSWI and central India. This is because rainfall process usually involves moisture convergence, which brings in moisture from surrounding regions (Trenberth, 1999).

8 Discussion and conclusions

Frequent dust storms develop in the boreal summer due to the strong “Shamal” winds in the AP and IP. After a long-distance transport, these dust storms can reach the AS and interact with ISM. Using observational datasets, we found a positive correlation between the ISM rainfall and the remote Middle East dust aerosols. To disclose the physical mechanism responsible for this correlation, a regional meteorological model with online chemistry, WRF-Chem, is used to examine the radiative effects of Middle East dust on ISM. The primary conclusions are drawn below.

WRF-Chem is capable of simulating the major ISM features and heavy dust loading in the Middle East and the AS during the boreal summer. The model can capture the ISM circulations quite well, e.g. the cross-equator circulation and the southwesterly winds over the AS. It also reproduces precipitation patterns quite similar to observations, with heavy precipitation located in CSWI and IGP. The low bias of precipitation in central India is partly attributed to the lack of representation of agriculture irrigation. In

addition, by tuning the empirical proportionality constant in the GOCART dust emission scheme against satellite observations, the model can capture the main spatiotemporal features of dust aerosols in the Middle East and its surrounding regions.

Satellite retrievals show that AOD in DST is positively correlated with the ISM rainfall in CSWI, CNI, and Pakistan. This correlation is examined here by using WRF-Chem. Two groups of experiments with the presence and absence of dust emission are designed to isolate dust impacts on ISM rainfall. Ensemble model results based on PPCE show that mineral dust increases the ISM rainfall by 0.44 mm day^{-1} (about 10%) in WHI. The spatial pattern of the modeled rainfall increase is highly consistent with the observational results shown in J2014 and this study. However, V2014 and S2015 show different patterns. Using a GCM, V2014 demonstrated significantly decreased rainfall or very weakly increased rainfall in central India. The differences in rainfall response between their studies and ours may be attributed to different atmospheric heating patterns caused by dust (Fig. 12 in this study vs. Fig. S4a in V2014), which may result from different representations of dust emission, dust diffusion in the PBL, or model resolution. An accurate representation of elevated orography of the Tibetan Plateau is critical in simulating ISM due to the large elevation gradient (Bollasina et al., 2011). Using an RCM, S2015 showed results that were opposite to those of J2014 and this study in terms of rainfall response. In S2015, the decreased rainfall in CNI is attributed to high GPH anomalies over CNI, Pakistan, the central and north AS. Our further analyses show that the modeled rainfall response from the ensemble mean shows a much stronger spatial correlation with the observed rainfall response than any of the ensemble members, indicating that ensemble mean can cancel out and reduce model errors from various parameterizations. By comparing the performance of individual ensemble member in Figs. 7 and 8, we conclude that the largest effects on our simulation results are from the PBL schemes, followed by the SW radiation schemes, and the aerosol chemical mixing rules have a relatively weak effect.

Cross-correlation analysis of modeled daily dust AOD in DST and the ISM rainfall response in WHI and CNI shows a maximum cross-correlation when dust AOD leads

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rainfall response by 11 days (Fig. 10). This finding is very similar to the 13 days found in observational datasets by J2014. Note that the model experiments here have separated the rainfall response due to dust for the correlation analysis, while observations include various forcings and responses. Considering that a typical dust event usually lasts 3 to 5 days, and the time for water vapor to be transported from the equator to CN1 is about a week assuming a wind speed of 10 m s^{-1} , the total time is about 10 to 12 days. Therefore, the timescale of 11 days is reasonable. This cross-correlation provides further evidence to support the causal relationship between Middle East dust and ISM rainfall.

Dust-induced ISM rainfall increase can be explained by the dynamic impacts of dust radiative forcing on water vapor transport from the AS to the ISM region. By absorbing solar radiation in the atmosphere (Fig. 11h), dust heats the lower troposphere (800–500 hPa) by about 2 to 2.5 K over the AS, the south AP, and the IP (Fig. 12). The dust-induced atmospheric heating causes a low-pressure system at 850 hPa over the AS and surrounding regions, which is associated with a convergence anomaly over the AS and north India (Fig. 13). The southwestern branch of the convergence anomaly transports more water vapor from the AS to the Indian subcontinent, resulting in more precipitable water in the atmospheric column (Fig. 14). This strengthened southwesterly wind due to dust-induced heating is responsible for the dust-ISM rainfall correlation observed in both satellite data and model simulations. Furthermore, the northwestern branch of the convergence anomaly over the AP can create more dust emission and transport these dust particles from the AP to the AS, building a positive feedback.

Two issues warrant further discussion. First, SST is prescribed during the three-month simulation period. The SST in the AS has been shown to play an important role in modulating the ISM rainfall (e.g. Levine and Turner, 2012). The surface cooling effect of dust can decrease the SST, which may influence the ISM rainfall response to Middle East dust. Secondly, although the aerosol indirect effect is turned on in this study, experiments with a relatively coarse horizontal resolution of 54 km cannot resolve convective clouds (typically 1–5 km wide), which in turn fails to consider the indirect aerosol

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Table 1. Configuration options of WRF-Chem used in this study.

Atmospheric Process		Model Option
Physics	Longwave radiation	RRTMG
	Shortwave radiation	RRTMG-Goddard
	Surface layer	Monin-Obukhov
	Land surface	Noah
	Boundary layer	YSU/BouLac
	Cumulus clouds	Grell-Freitas
	Cloud microphysics	Lin et al. (1983)
Chemistry	Gas-phase chemistry	RADM2
	Aerosol chemistry	MOSAIC-SORGAM with aqueous reactions
	Photolysis	Fast-J
Emission	Dust emission	GOCART
	Sea-salt emission	MADE/SORGAM
	Anthropogenic emission	Reanalysis of the Tropospheric and Emission Database for Global Atmospheric Research
	Biogenic emission	The Model of Emissions of Gases and Aerosols from Nature version 2
	Fire emission	MODIS

Table 2. Various schemes employed to create the ensemble members.

Scheme	Option	Option	Description
Shortwave radiation	sw4	Goddard	1. 11 spectral bands (7 UV, 1 PAR, 3 infrared) 2. A two-stream adding algorithm 3. No interacts with sub-grid clouds
	sw2	RRTMG	1. 14 spectral bands (3 UV, 2 PAR, 9 infrared) 2. A two-stream algorithm for multiple scattering 3. Represents sub-grid cloud variability by McICA with maximum-random cloud overlap.
Planetary boundary layer	pb1	YSU	1. First-order closure 2. <i>K</i> profile and non-local mixing 3. Explicit treatment of entrainment
	pb8	BouLac	1. One-and-a-half order closure 2. Prognostic turbulent kinetic energy equation
Aerosol chemical mixing rules	op1	Volume	1. Internal-mixing of aerosol composition 2. Full Mie calculations are called only at the first time step
	op2	Maxwell	1. Randomly distributes black carbon within a particle 2. Full Mie calculations are called only at the first time step
	op3	Exact volume	Same as volume, but use full Mie calculation at each time step
	op4	Exact Maxwell	Same as Maxwell, but use full Mie calculation at each time step

UV, PAR, and McICA stand for ultraviolet, photosynthetically active radiation, and the Monte Carlo Independent Column Approximation, respectively.

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Table 3. Rainfall response in various regions of India to Middle East dust in this study and other references.

Reference		Pakistan	CNI	CSWI	Southeast India
This study	Satellite	+	+	+	-
	Model	+	+	+	-
V2014		Not shown	No change	+	+
J2014		+	+	+	-
S2015		-	-	+	+

“+” and “-” mean increased and decreased rainfall, respectively.

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Table 4. Area-averaged direct radiative forcing of dust simulated by WRF-Chem over the entire model domain for JJA 2008. The acronyms have the same meaning as in Fig. 11. Units: W m^{-2} .

	SW	LW	Net
TOA	-3.32	1.07	-2.25
ATM	6.44	-1.64	4.80
BOT	9.76	2.71	-7.05

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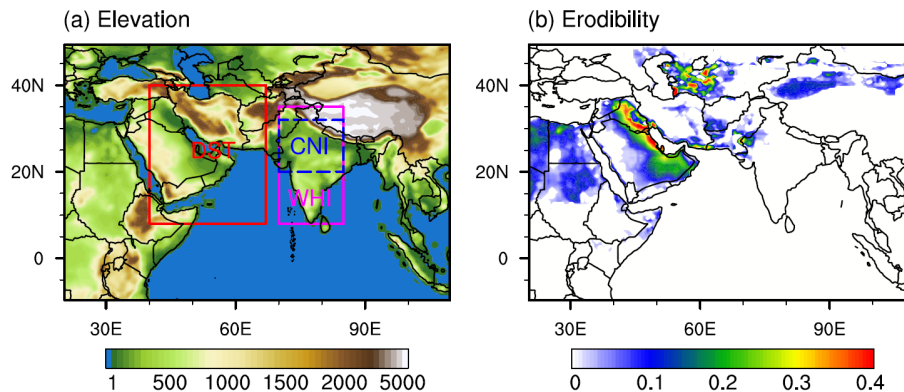


Figure 1. (a) Elevation map in WRF-Chem domain (unit: m). “DST”, “WHI”, and “CNI” represent the remote dust region (8–40° N, 40–67° E), the entire India (8–35° N, 70–85° E), and the central and northern India (20–32° N, 70–85° E), respectively. (b) Soil erodibility (unitless) map used in WRF-Chem dust emission scheme.

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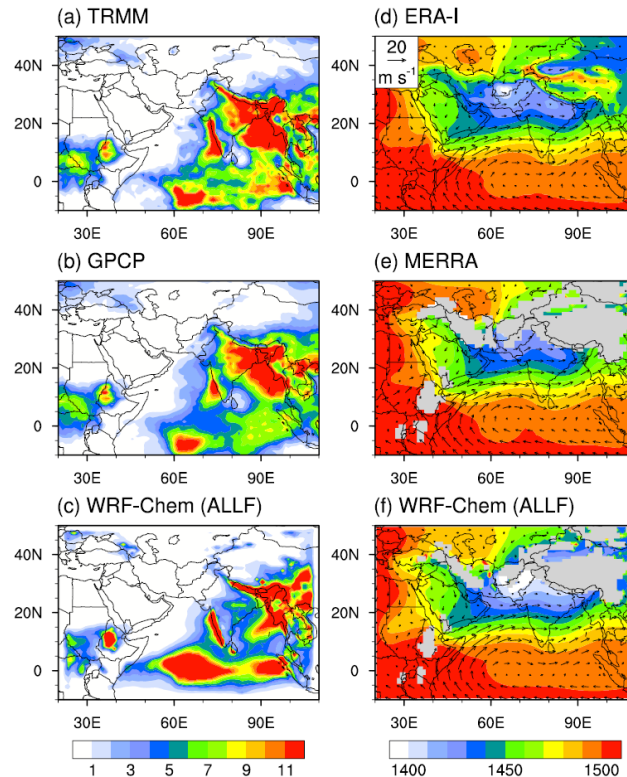


Figure 2. (Left) Precipitation (mm day^{-1}) from (a) TRMM, (b) GPCP, (c) WRF-Chem (ALLF). (Right) 850 hPa GPH (shading; unit: m) and wind (arrows; units: m s^{-1}) from (d) ERA-I, (e) MERRA, (f) WRF-Chem (ALLF). All variables are average for JJA 2008. In Panels (e) and (f), topography is masked out in grey.

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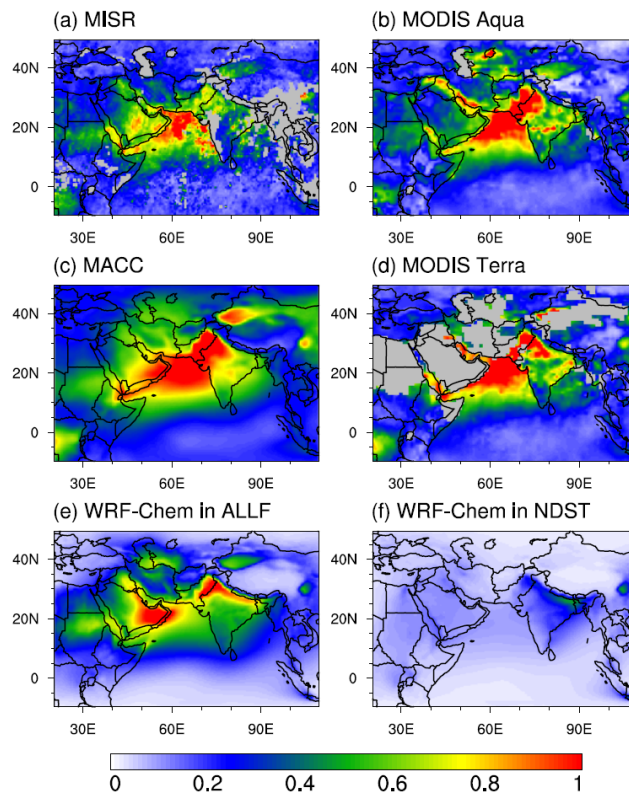


Figure 3. Spatial patterns of AOD (unitless) from **(a)** MISR (558 nm), **(b)** MODIS Aqua (550 nm), **(c)** MACC (550 nm), **(d)** MODIS Terra (550 nm), **(e)** WRF-Chem (ALLF; 550 nm), and **(f)** WRF-Chem (NDST; 550 nm) averaged for JJA 2008. Missing values are masked in grey in Panels **(a)**, **(b)** and **(d)**.

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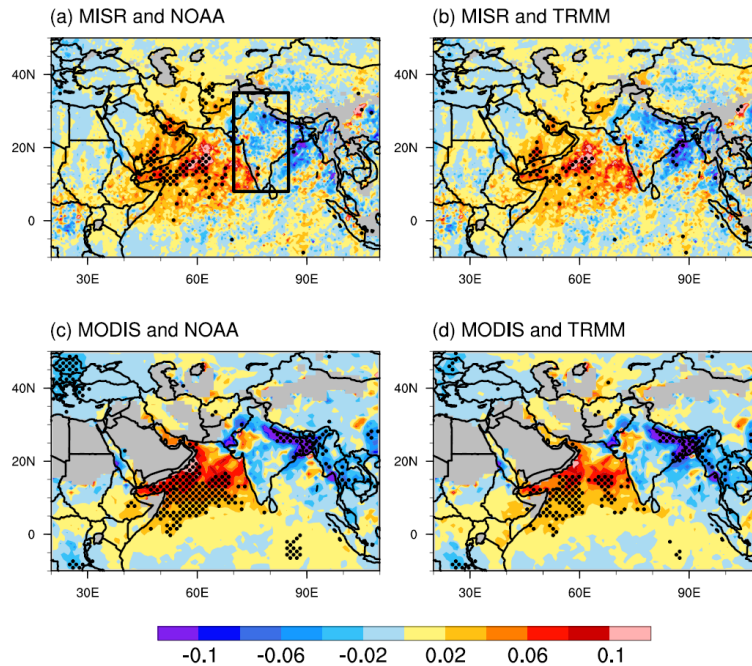


Figure 4. Spatial patterns of AOD regressed on area-averaged ISM rainfall in WHI (box in **a**) using JJA monthly anomalies from 2000 to 2013. Two different AOD datasets (MISR and MODIS Terra) and two different rainfall datasets (NOAA and TRMM) are used, so there are four different results. The black dots represent grid points that are 95% confident based on the t test. Missing values are masked out in grey.

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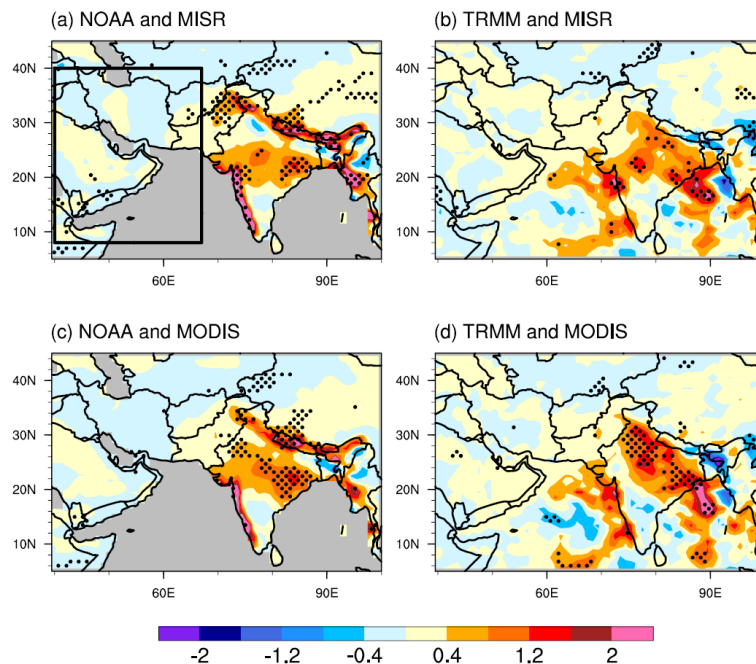


Figure 5. Same as Fig. 4, but for the spatial patterns of precipitation (mm day⁻¹) regressed on area-averaged AOD in the AS, the south AP, and the IP (box in a).

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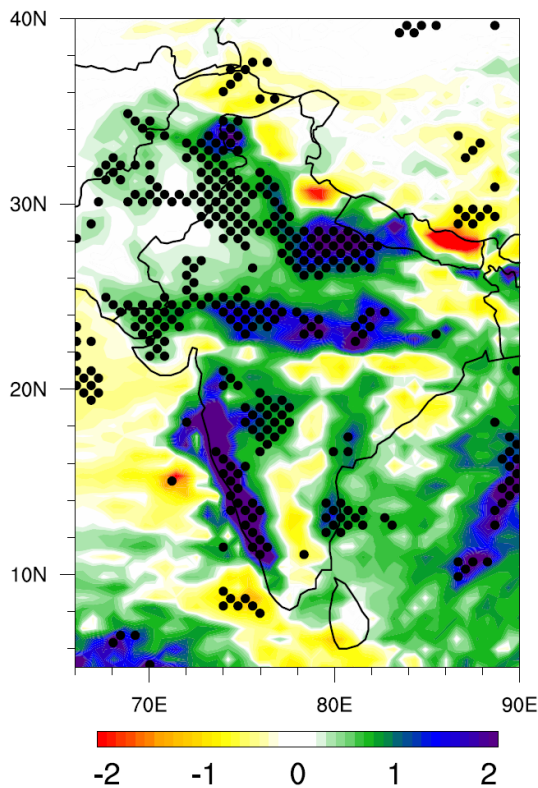


Figure 6. Spatial pattern of WRF-Chem ensemble mean differences in rainfall (mm day^{-1}) between AFFL and NDST experiments during JJA 2008. Areas that are confident at the 90 % level based on a one-sided Student's t test are dotted.

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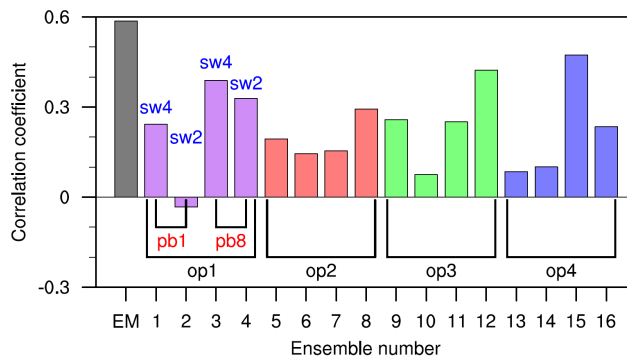


Figure 7. The centered spatial correlation coefficients between the regressed rainfall change pattern (Fig. 5c) and the modeled rainfall response (Fig. 6) from the ensemble members (marked by numbers from 1 to 16) and their ensemble mean (marked by “EM”). “EM” stands for the ensemble mean. The region for calculating the spatial correlation is WHI. Using other figures in Fig. 5 for the evaluations gets similar results. The uncentered spatial correlation shows the similar results. See Table 2 for the donations.

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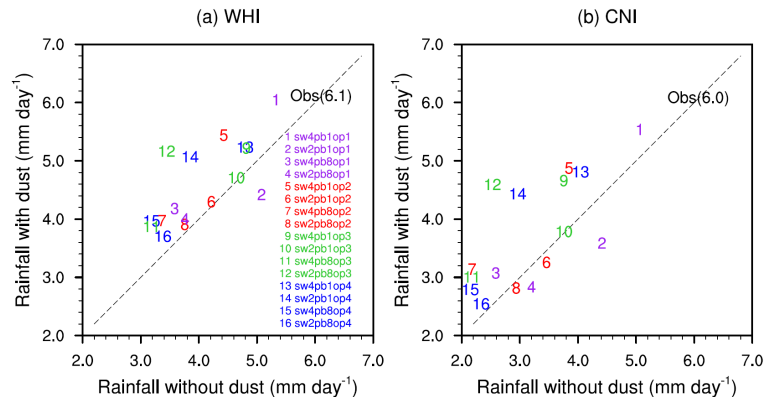


Figure 8. Scatter plot of area-averaged rainfall (mm day^{-1}) in WHI and CNI in 16 pairs of ALLF and NDST ensemble simulations and observations. “Obs” means averaged observational rainfall from TRMM, GPCP, and CMAP, with area-averaged mean values of 6.1 and 6.0 in parentheses in WHI and CNI, respectively. The ensemble numbers are the same as in Fig. 7. See Table 2 for the denotations.

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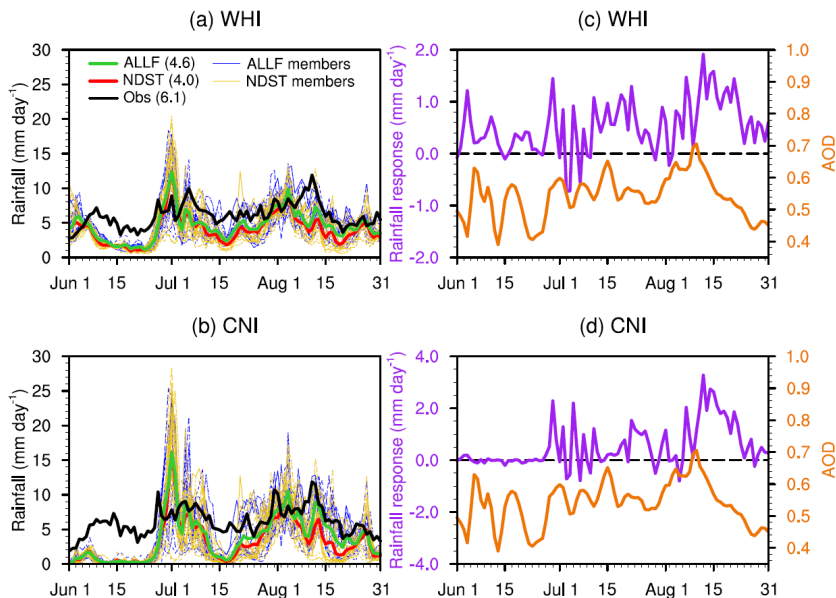


Figure 9. (Left) Time series of rainfall (mm day^{-1}) in 32 ensemble members and ensemble means of ALLF and NDST experiments and (Right) ensemble mean rainfall responses (mm day^{-1}) in WHI and CNI and AOD in DST (from ALLF). The numbers in parentheses are time-averaged rainfall.

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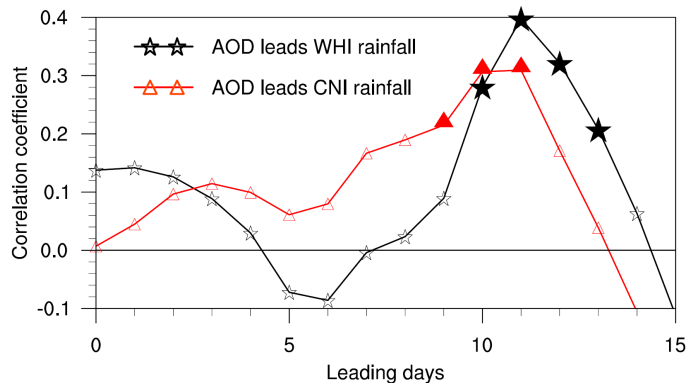


Figure 10. Cross-correlation coefficients between WRF-Chem simulated rainfall responses in WHI and CNI and dust AOD (ALLF minus NDST) in remote DST region. All correlations are calculated based upon daily anomalies obtained by subtracting the 21-day running mean from the daily data. The filled markers represent that the correlation coefficients are 95 % confident based on the t test. The unfilled markers are not significant.

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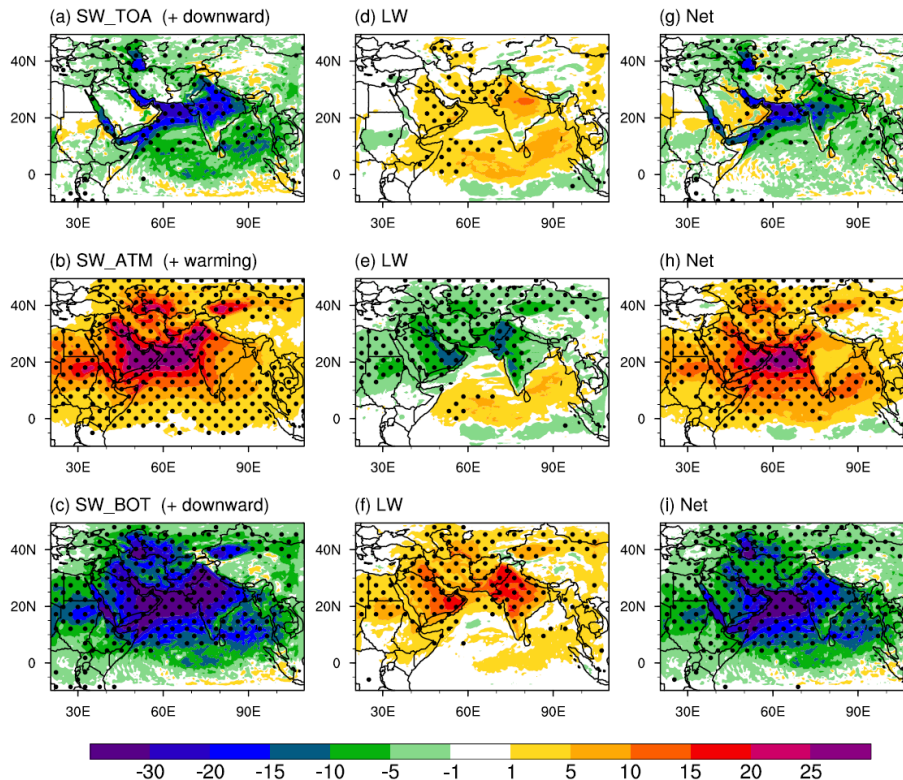


Figure 11. Spatial patterns of dust direct radiative forcing (W m^{-2}) at the top of the atmosphere (TOA), in the atmosphere (ATM), and at the surface (BOT) averaged during JJA 2008. Calculated from the ensemble mean differences between ALLF and NDST experiments of WRF-Chem. Downward radiation is defined as positive at the TOA and the surface; therefore, positive (negative) value means absorb/warming (irradiate/cooling) effects in the atmosphere. Net radiative forcing is the sum of SW and LW radiative forcing. The dotted areas mean that radiative forcing is 95 % confident based on one-sided Student's t test.

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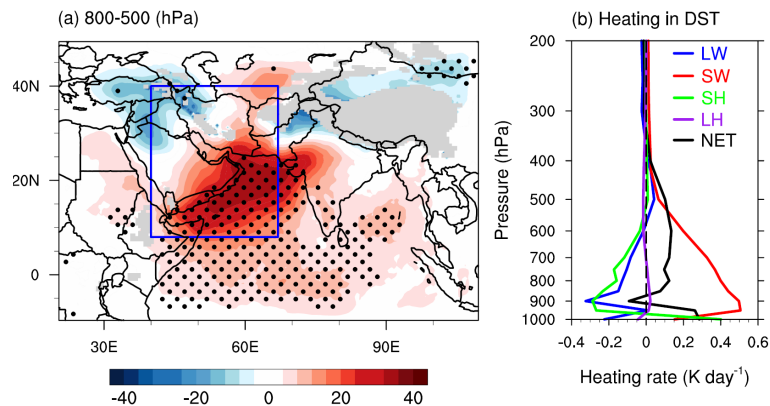


Figure 12. (a) Spatial pattern of WRF-Chem ensemble mean differences in thickness between 800 and 500 hPa pressure layers between the ALLF and NDST experiments (unit: m). (b) Vertical profiles of the WRF-Chem ensemble mean differences in atmospheric heating rate area-averaged in DST (box in a) between the ALLF and NDST experiments (units: K day^{-1}). The dotted areas in (a) mean that radiative forcing is 95 % confident based on one-sided Student's t test. All heating rates are for all-sky conditions. “SH”, “LH”, and “NET” stand for sensible heating, latent heating, and net heating rate. Net heating is the sum of SW, LW, SH, and LH.

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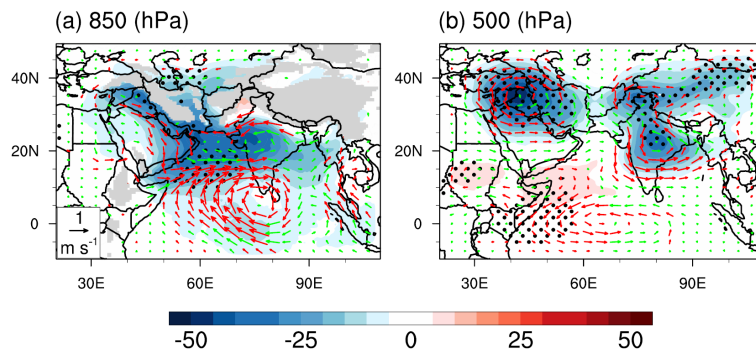


Figure 13. WRF-Chem ensemble mean differences in GPH (shading; unit: m) and winds (arrows; units: m s^{-1}) at **(a)** 850 hPa and **(b)** 500 hPa between ALLF and NDST experiments. The dotted areas are 95% confident based on one-sided Student's *t* test. The red arrows show wind differences that are 95% confident based on one-sided Student's *t* test, and green arrows are other wind differences (not confident).

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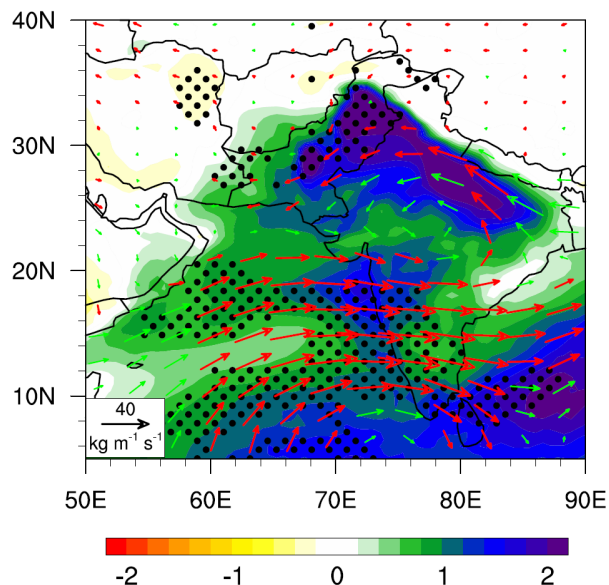


Figure 14. Same as Fig. 6, but for precipitable water (shading; unit: mm) and water vapor flux (arrows; units: $\text{kg m}^{-1} \text{s}^{-1}$) both integrated within the entire atmospheric column. Black dots represent precipitable water differences that are 95 % confident based on a one-sided Student's t test. The red arrows represent wind differences that are 95 % confident, and the green arrows represent other wind differences (not confident).

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