

# Diurnal, seasonal and long-term variations of global formaldehyde columns inferred from combined OMI and GOME-2 observations

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## 1 Abstract

2 We present the new version (v14) of the BIRA-IASB algorithm for the retrieval of  
3 formaldehyde (H<sub>2</sub>CO) columns from spaceborne UV-Visible sensors. Applied to OMI  
4 measurements from Aura and to GOME-2 measurements from MetOp-A and B, this  
5 algorithm is used to produce global distributions of H<sub>2</sub>CO representative of mid-morning and  
6 early afternoon conditions. Its main features include (1) a new iterative DOAS scheme  
7 involving three fitting intervals to better account for the O<sub>2</sub>-O<sub>2</sub> absorption, (2) the use of  
8 earthshine radiances averaged in the equatorial Pacific as reference spectra, (3) a destriping  
9 correction and background normalisation resolved in the [alongacross](#)-swath position. For the  
10 air mass factor calculation, *a priori* vertical profiles calculated by the IMAGES chemistry  
11 transport model at 09:30 am and 13:30 pm are used. Although the resulting GOME-2 and  
12 OMI H<sub>2</sub>CO vertical columns are found to be highly correlated, some systematic differences  
13 are observed. Afternoon columns are generally larger than morning ones, especially in mid-  
14 latitude regions. In contrast, over tropical rainforests, morning H<sub>2</sub>CO columns significantly  
15 exceed those observed in the afternoon. These differences are discussed in terms of the H<sub>2</sub>CO  
16 column variation between mid-morning and early afternoon, using ground-based MAX-  
17 DOAS measurements available from seven stations in Europe, China and Africa. Validation  
18 results confirm the capacity of the combined satellite measurements to resolve diurnal  
19 variations in H<sub>2</sub>CO columns. Furthermore, vertical profiles derived from MAX-DOAS  
20 measurements in the Beijing area and in Bujumbura are used for a more detailed validation  
21 exercise. In both regions, we find an agreement better than 15% when MAX-DOAS profiles

1 are used as *a priori* for the satellite retrievals. Finally regional trends in H<sub>2</sub>CO columns are  
2 estimated for the 2004-2014 period using SCIAMACHY and GOME-2 data for morning  
3 conditions, and OMI for early afternoon conditions. Consistent features are observed such as  
4 an increase of the columns in India and Central-East China, and a decrease in Eastern US and  
5 Europe. We find that the higher horizontal resolution of OMI combined to a better sampling  
6 and a more favourable illumination at mid-day allow for more significant trend estimates,  
7 especially over Europe and North America. Importantly, in some parts of the Amazonian  
8 forest, we observe with both time series a significant downward trend in H<sub>2</sub>CO columns,  
9 spatially correlated with areas affected by deforestation.

## 10 **1 Introduction**

11 Atmospheric formaldehyde (H<sub>2</sub>CO) is an intermediate product common to the degradation of  
12 many volatile organic compounds. While the global formaldehyde background is due to  
13 methane oxidation, emissions of non-methane volatile organic compounds (NMVOCs) from  
14 biogenic, biomass burning and anthropogenic continental sources result in important and  
15 localised production of H<sub>2</sub>CO. The global sink of H<sub>2</sub>CO is due to photolysis and oxidation by  
16 OH, resulting in a photochemical lifetime of only a few hours. Elevated concentrations of  
17 H<sub>2</sub>CO can therefore be related to the emission of reactive NMVOCs. Monitoring the spatial  
18 and temporal variability of NMVOC emissions is essential for a better understanding of the  
19 processes that not only control the production and the evolution of tropospheric ozone, key  
20 actor in air quality and climate change, but also of the hydroxyl radical OH and secondary  
21 organic aerosols. For these reasons, H<sub>2</sub>CO satellite observations have been increasingly used  
22 in combination with tropospheric chemistry transport models to constrain NMVOC emissions  
23 [in top-down inversion approaches](#) (*i.a.* Palmer et al., 2006; Fu et al., 2007; Millet et al., 2008;  
24 Stavrakou et al., 2009a, 2009b; 2014; Curci et al., 2010; Barkley et al., 2013; Fortems-  
25 Cheiney et al., 2012; Marais et al., 2012, Zhu et al., 2014).

26 For more than 15 years, mid-morning formaldehyde tropospheric columns have been  
27 retrieved from the successive nadir scanning spectrometers GOME on ERS-2 (1996-2003)  
28 (Chance et al., 2000; Palmer et al., 2001; De Smedt et al., 2008), SCIAMACHY on  
29 ENVISAT (2002-2011) (Wittrock et al., 2006; De Smedt et al., 2008) and GOME-2 on  
30 MetOp-A & B (2006- & 2012-) (De Smedt et al., 2012; Hewson et al., 2013). Since 2004,  
31 complementary early afternoon H<sub>2</sub>CO columns can also be retrieved from the imaging  
32 [spectrometer OMI on Aura](#) (Kurosu, 2008; Millet et al., 2008; González Abad et al., 2015a).

1 | [and since 2011 from OMPS on SUOMI-NPP](#) -(Li et al., 2015; González Abad et al., 2015b).

2 In addition to formaldehyde, glyoxal – another short-lived NMVOC – has also successfully  
3 been retrieved from SCIAMACHY (Wittrock et al., 2006), GOME-2 (Vrekoussis et al., 2010;  
4 Lerot et al., 2010) and OMI (Alvarado et al., 2014; Chan Miller et al., 2014).

5 From 2017 onward, the morning observations will be continued with a third GOME-2  
6 instrument to be launched on MetOp-C (Callies et al., 2000), while the afternoon observations  
7 will be extended with the TROPOMI instrument (Veefkind et al., 2012), to be launched in  
8 2016 as part of the Copernicus Sentinel 5 Precursor (S-5P) mission, and later with the  
9 Sentinel-5 mission to be operated on the MetOp Second Generation platform (Ingmann et al.,  
10 2012). Also at the 2020 horizon, the Sentinel-4 instrument on the geostationary Meteosat  
11 Third Generation (MTG) platform will allow for hourly observations of H<sub>2</sub>CO over Europe,  
12 while TEMPO (NASA) and GEMS (KARI) will provide geostationary measurements for  
13 North America and Asia respectively. To realize the full potential of these missions, it is  
14 crucial to develop high quality and consistent retrieval algorithms applicable to the different  
15 satellite sensors, taking into account their differences in horizontal resolution and sampling.  
16 Likewise, it is essential to understand the diurnal variations of the sources and sinks of  
17 formaldehyde, in order to exploit the synergy between the morning and afternoon satellite  
18 observations. To our knowledge, so far, the use of combined morning and afternoon H<sub>2</sub>CO  
19 satellite observations has only been reported over Amazonia, using SCIAMACHY and OMI  
20 measurements (Barkley et al., 2011; 2013).

21 Ground-based measurements are essential to quantitatively assess the seasonal and diurnal  
22 variations of the tropospheric H<sub>2</sub>CO columns. However, up to now, very few validation  
23 studies have been reported for satellite H<sub>2</sub>CO observations (Wittrock et al., 2006; Vigouroux  
24 et al., 2009) because of the general lack of suitable ground-based measurements, in particular  
25 for tropical regions where H<sub>2</sub>CO columns are among the highest worldwide (Stavrakou et al.,  
26 2009b; Marais et al., 2012; Barkley et al., 2013). Also, very little attention has been paid to  
27 the diurnal variations of the H<sub>2</sub>CO columns and to their local dependencies, which are a  
28 complex blend of local NMVOC emission variations, H<sub>2</sub>CO production and loss via oxidation  
29 and photolysis depending on local chemical regimes and season. To this regard, the latest  
30 generation of MAX-DOAS instruments and retrieval algorithms offer new perspectives for  
31 the validation of tropospheric trace gas concentrations and aerosol optical densities (Clémer et  
32 al., 2010; Pinardi et al., 2013; Vlemmix et al., 2014; Wang et al., 2014).

1 | This study focuses on [tropospheric](#) formaldehyde retrievals from OMI, using an algorithm  
2 | historically developed within the TEMIS (Tropospheric Emission Monitoring Internet  
3 | Service) framework and applied to morning observations from the GOME, SCIAMACHY  
4 | and GOME-2 sensors (<http://h2co.aeronomy.be>). We present several adaptations that have  
5 | been implemented to handle observations from the OMI imaging spectrometer, as well as  
6 | number of more general improvements to the algorithm, giving rise to a new version of the  
7 | BIRA-IASB H<sub>2</sub>CO retrieval algorithm (version 14). This version has been applied to the  
8 | complete time series of OMI measurements, as well as to the GOME-2 measurements from  
9 | MetOp-A and B platforms. For the first time, differences between morning and afternoon  
10 | H<sub>2</sub>CO columns are estimated at the global scale and discussed in terms of H<sub>2</sub>CO diurnal  
11 | variations, horizontal resolution effects and retrieval uncertainties. Moreover, ground-based  
12 | measurements at seven stations, covering mid-latitude and tropical locations, are used to  
13 | validate the observed H<sub>2</sub>CO columns and their diurnal changes as derived from the combined  
14 | satellite data sets.

15 | The paper is structured as follows: section 2 introduces the main characteristics of the OMI  
16 | and GOME-2 instruments. Section 3 describes the new version (v14) of the H<sub>2</sub>CO retrieval  
17 | algorithm. The H<sub>2</sub>CO tropospheric columns obtained from GOME-2 and OMI measurements  
18 | are presented and compared in section 4, and the main results of our validation studies are  
19 | outlined in section 5. Finally, the long-term variations of the H<sub>2</sub>CO columns over the last  
20 | decade are discussed in section 6.

## 21 | **2 Satellite instruments**

### 22 | **2.1 OMI on Aura**

23 | The Aura satellite was launched in July 2004, in a sun-synchronous polar orbit crossing the  
24 | equator [around](#) 13:30 LT (in ascending mode). It is the third major component of the NASA  
25 | Earth Observing System (EOS) following Terra (launched 1999) and Aqua (launched 2002).  
26 | OMI (Ozone Monitoring Instrument) is a nadir viewing imaging spectrometer that measures  
27 | the solar radiation backscattered by the Earth's atmosphere and surface over the wavelength  
28 | range from 270 to 500 nm with a spectral resolution of about 0.5 nm (Levelt et al., 2006). The  
29 | light entering the telescope is depolarised using a scrambler and then split into two channels: a  
30 | UV channel (wavelength range 270 - 380 nm) and a VIS channel (wavelength range 350 - 500  
31 | nm). The 114° viewing angle of the telescope corresponds to a 2600 km wide swath on the  
32 | earth surface, which enables [nearly](#) daily global coverage. In the nominal global operation

1 mode, the OMI ground pixel size varies from 13x24 km<sup>2</sup> at nadir to 28x150 km<sup>2</sup> at the edges  
2 of the swath. For this work, we have used the OMI Level 1B UV Global Radiances Data  
3 Product (OML1BRUG - Version 003) provided on the NASA website  
4 ([http://disc.sci.gsfc.nasa.gov/Aura/data-holdings/OMI/oml1brug\\_v003.shtml](http://disc.sci.gsfc.nasa.gov/Aura/data-holdings/OMI/oml1brug_v003.shtml)).

## 5 **2.2 GOME-2 on MetOp-A & B**

6 The MetOp-A and MetOp-B satellites were respectively launched in October 2006 and  
7 September 2012, in sun-synchronous polar orbits with equator crossing times of 09:30 and  
8 09:00 LT (in descending node). Both satellites carry the same type of GOME-2 (Global  
9 Ozone Monitoring Experiment) instruments. Hereinafter, we will refer to them by GOME-2A  
10 and GOME-2B (or G2A and G2B). GOME-2 (Callies et al., 2000; Munro et al., 2006) is an  
11 improved version of the GOME instrument which flew on the ERS-2 satellite. It is a nadir  
12 viewing scanning spectrometer with four main optical channels, covering the spectral range  
13 between 240 and 790 nm with a spectral resolution between 0.26 nm and 0.51 nm.  
14 Additionally, two polarisation components are measured with polarisation measurements  
15 devices (PMDs) at 30 broad-band channels covering the full spectral range. A direct sun  
16 spectrum is also measured via a diffuser plate once per day. The default swath width of the  
17 GOME-2 scan is 1920 km allowing for global Earth coverage within 1.5–3 days at the  
18 equator. The nominal ground pixel size is 80×40 km<sup>2</sup>. For this work, we have used the  
19 EUMETSAT GOME-2A and GOME-2B level 1B data version 5.3.0 from the beginning of  
20 their time series up to mid-June 2014, and version 6.0.0 afterwards.

## 21 **3 Formaldehyde retrievals**

22 We use a DOAS algorithm (Differential Optical Absorption Spectroscopy), including three  
23 main steps, further detailed in Sect. 3.1, 3.2 and 3.3: (1) the fit of absorption cross-section  
24 databases to the log-ratio of measured Earth reflectance to retrieve H<sub>2</sub>CO slant columns ( $N_s$ ),  
25 (2) a background normalisation procedure to eliminate remaining unphysical dependencies,  
26 and (3) the calculation of tropospheric air mass ( $M$ ) factors using radiative transfer  
27 calculations and modelled *a priori* profiles. The tropospheric H<sub>2</sub>CO vertical column ( $N_v$ ) is  
28 related to intermediate quantities by the equation:

$$29 \quad N_v = \frac{\Delta N_s}{M} + N_{v,0,CTM}$$

1 where  $\Delta N_s$  is the background-corrected slant column density, and  $N_{v,0,CTM}$  is the model  
2 background column in the reference sector. More detailed equations can be found in our  
3 previous publications (e.g. De Smedt et al., 2011 and 2014).

4 ~~Since 10 years, the TEMIS website provides~~ Level 2 formaldehyde products developed at  
5 ~~BIRA-IASB are provided via the TEMIS website, for all European nadir viewing UV-VIS~~  
6 ~~spectrometers~~. The algorithms used to generate these products ~~were~~ are designed to be as  
7 consistent as possible, in order to optimise the overall coherency of the resulting time series.  
8 Over the years, scientific developments are leading to step-by-step improvements in the  
9 quality of the data products. This entails regular reprocessing of the datasets. The retrieval  
10 settings presented here for the GOME-2 and OMI measurements are based on the BIRA  
11 algorithm developed for GOME-2 (De Smedt et al., 2012) but include a number of  
12 adaptations allowing for the efficient processing of imaging instruments, and additional  
13 improvements as further detailed in this section. The main differences compared to version 12  
14 are (1) the use of daily radiances as DOAS reference spectra, (2) the inclusion of O<sub>2</sub>-O<sub>2</sub> in the  
15 BrO and H<sub>2</sub>CO retrieval intervals and (3) the pre-fit of O<sub>2</sub>-O<sub>2</sub> in a dedicated retrieval interval.  
16 For the reference sector correction, which so far was resolved in latitude and time, an  
17 additional dimension (viewing zenith angle or detector row) is now introduced as destriping  
18 procedure. An updated version of the IMAGES model *a priori* profiles is ~~sampled~~ used at the  
19 respective satellite overpass times. Quality flags have been defined, in order to better select  
20 the valid satellite observations. Finally, the format of the level 2 data files has been changed  
21 to HDF-5. As for the previous algorithm versions, level 2 and level 3 data products of version  
22 14 are openly available on TEMIS (<http://h2co.aeronomie.be>). Comparisons between the  
23 BIRA-IASB OMI H<sub>2</sub>CO product and the OMI operational product from SAO can be found in  
24 González Abad et al. (2015b), while comparisons with the NASA OMPS PCA product can be  
25 found in Li et al. (2015).

### 26 **3.1 Slant Columns**

27 The formaldehyde slant columns are retrieved in the interval 328.5-346 nm. Retrieval settings  
28 are summarized in ~~Table 1~~ Table 1. We use the QDOAS software developed at BIRA-IASB  
29 for the DOAS retrieval of trace gases from many common satellite and ground-based  
30 instruments. QDOAS is distributed under the GNU GPL license version 2.0 (Danckaert, Fayt  
31 and Van Roozendael, 2014, <http://uv-vis.aeronomie.be/software/QDOAS>).

1 In QDOAS, the wavelength registration of the reference spectrum is fine-tuned by means of a  
2 calibration procedure making use of the solar Fraunhofer lines. To this end, a highly accurate  
3 reference solar atlas (Chance and Kurucz, 2010) is degraded at the resolution of the  
4 instrument, through convolution by the instrumental slit function. The absorption cross  
5 sections of the different trace gases are also convolved with the instrumental slit function. In  
6 the case of GOME-2, the slit function shape is fitted during the calibration procedure, in order  
7 to take into account its changes with time (Dikty and Richter, 2011; De Smedt et al., 2012). In  
8 the case of OMI, pre-flight measured slit functions (Dobber et al., 2006) are used, and the  
9 calibration is performed for each binned spectrum of the detector array (60 rows). Except for  
10 the reported row anomaly that has dynamically evolved over the years  
11 (<http://www.knmi.nl/omi/research/product/rowanomaly-background.php>), the performance of  
12 the OMI instrument has proven to be very stable in time (Dobber et al., 2008). In contrast to  
13 previous work, we now moved to the systematic use of daily radiance spectra averaged in the  
14 equatorial Pacific (15°S-15°N, 180°E-240°E) as reference for the DOAS retrieval. Different  
15 reference spectra are selected daily for each OMI row and their wavelength registration is  
16 optimised as described above. This serves as a first correction for the OMI stripe effect.  
17 Consequently, all retrieved slant columns are differential columns relative to the mean  
18 reference spectra.

19 A 5<sup>th</sup> order polynomial is used to fit the low-frequency variations of the spectra, as well as a  
20 linear offset term. The H<sub>2</sub>CO Meller and Moortgat (2000) laboratory measurements are fitted  
21 to the differential absorption features. The absorption cross-sections of O<sub>3</sub> at 228K and 243K,  
22 NO<sub>2</sub> and BrO are included. To take into account the Ring effect, two cross-sections are used  
23 (Vountas et al., 1998). They have been calculated in an ozone-containing atmosphere for low  
24 and high SZA using LIDORT RRS (Spurr et al., 2008). Two additional terms (called O<sub>3</sub>L and  
25 O<sub>3</sub>O<sub>3</sub> in [Table 1](#)~~Table 1~~), resulting from the Taylor expansion of the O<sub>3</sub> absorption as a  
26 function of the wavelength, are included in order to better cope with strong O<sub>3</sub> absorption  
27 effects (Puķīte et al; 2010; De Smedt et al., 2012). Introduced in the previous version 12, a  
28 second (larger) retrieval interval is used to pre-fit the BrO slant columns (Fig.1). The benefit  
29 of this procedure is to decorrelate the H<sub>2</sub>CO and BrO absorption features, resulting in a  
30 reduction of the noise on the BrO and H<sub>2</sub>CO slant columns. However the O<sub>2</sub>-O<sub>2</sub> (O<sub>4</sub>)  
31 absorption had not been taken into account in version 12, neither in the larger nor in the  
32 shorter interval. While the weak and smooth O<sub>4</sub> signature appears to be well fitted by the 5<sup>th</sup>  
33 order polynomial used in the shorter interval (Hewson et al., 2013), not including the stronger

1 O<sub>4</sub> term in the larger interval is clearly a shortcoming. Nevertheless, this solution has often  
 2 been selected in past studies, because experience has shown that including O<sub>4</sub> tends to  
 3 destabilize the BrO fit (Kurosu, 2008; Begoin et al., 2010; Theys et al., 2011) hence leading to  
 4 increased noise on the H<sub>2</sub>CO slant columns retrieved in the shorter interval. In version 14, we  
 5 propose the addition of a third fitting interval (339-364 nm) covering the entire O<sub>4</sub> absorption  
 6 band around 360 nm (Fig.1). We use the recently published O<sub>4</sub> absorption cross-sections by  
 7 Thalman and Volkamer (2013), and our H<sub>2</sub>CO retrieval scheme therefore now includes three  
 8 fitting intervals: w1 for the pre-fit of O<sub>4</sub>, w2 for the pre-fit of BrO, and w3 for the fit of the  
 9 H<sub>2</sub>CO slant columns, in which the O<sub>4</sub> and BrO slant columns are fixed to values determined in  
 10 the other intervals.

11 It should be noted that for the GOME and SCIAMACHY H<sub>2</sub>CO retrievals, only one retrieval  
 12 interval is used (328.5-346 nm) because the quality of the recorded spectra has been found to  
 13 be insufficient in the 360 nm region (De Smedt et al., 2008). O<sub>4</sub> absorption effects are  
 14 therefore significantly reduced.

15 Table 1 : Summary of retrieval settings used in version 14 of the BIRA H<sub>2</sub>CO retrieval  
 16 algorithm, applicable to GOME-2 and OMI measurements.

<b>Settings for the DOAS equation parameters</b>		
Calibration	Accurate solar atlas (Chance and Kurucz, 2010)	
Slit function	OMI: 1 slit function per binned spectrum as a function of wavelength (Dobber et al., 2006)	GOME-2: asymmetric Gaussian slit function fitted during calibration (De Smedt et al., 2012)
Reference spectrum	Daily average of radiances selected in the equatorial Pacific	
Polynomial	5 <sup>th</sup> order	
Intensity offset	Linear offset	
Hot pixels treatment	Iterative spike removal algorithm (Richter et al., 2011, De Smedt et al., 2014)	
<b>Absorption cross-section datasets</b>		
H <sub>2</sub> CO	Meller and Moortgat (2000)	
O <sub>3</sub>	Brion et al. (1998); Daumont et al. (1992); Malicet et al. (1995)	
BrO	Fleischmann et al. (2004)	
NO <sub>2</sub>	Vandaele et al. (2002)	
O <sub>2</sub> -O <sub>2</sub> (O <sub>4</sub> )	Thalman and Volkamer (2013)	
Ring effect	2 Ring cross-sections calculated in an ozone containing atmosphere for low and high SZA, using LIDORT RRS (Spurr et al., 2008).	
Non-linear O <sub>3</sub> absorption effect	2 pseudo-cross sections from the Taylor expansion of the wavelength and the O <sub>3</sub> optical depth (Puķīte et al., 2010).	
<b>Fitting interval w1: O<sub>4</sub></b>	<b>339-364 nm</b>	
Included cross-sections	O <sub>4</sub> (293K), O <sub>3</sub> (228K), BrO (223K), H <sub>2</sub> CO (298K), NO <sub>2</sub> (220K), Ring1, Ring2	
<b>Fitting interval w2: BrO</b>	<b>328.5-359 nm</b>	



Included cross-sections	BrO (223K), H <sub>2</sub> CO (298K), O <sub>3</sub> (228K and 243K), NO <sub>2</sub> (220K), O <sub>4</sub> (293K, not fitted), Ring1, Ring2, O <sub>3</sub> L, O <sub>3</sub> O <sub>3</sub>
<b>Fitting interval w3: H<sub>2</sub>CO</b>	<b>328.5-346 nm</b>
Included cross-sections	H <sub>2</sub> CO (298K), O <sub>3</sub> (228K and 243K), BrO (223K, not fitted), NO <sub>2</sub> (220K), O <sub>4</sub> (293K, not fitted), Ring1, Ring2, O <sub>3</sub> L, O <sub>3</sub> O <sub>3</sub>

1 Figure 2 illustrates the correlation effects occurring between the O<sub>4</sub>, BrO and H<sub>2</sub>CO  
2 absorptions, for one OMI orbit on July, 1<sup>st</sup> 2005. The upper panel shows the O<sub>4</sub> differential  
3 slant columns retrieved in w1 (339-364 nm, light blue) or in w2 (328.5-359 nm, dark green)  
4 as a function of the latitude. The improved quality of the O<sub>4</sub> slant columns in w1 is observed,  
5 as expected from Fig.1. The second and third panels show the differences observed in the BrO  
6 and H<sub>2</sub>CO slant columns depending on whether O<sub>4</sub> is included in the fits or not (v14-v12), as  
7 a function of the O<sub>4</sub> slant columns. Those differences are shown for two cases: O<sub>4</sub> slant  
8 columns from w1 (light blue, v14) or from w2 (dark green). The pre-fit of O<sub>4</sub> in w1 does  
9 reduce the noise on the O<sub>4</sub>, BrO and H<sub>2</sub>CO slant columns, in the three intervals. The  
10 correlation between the 3 molecules is not reduced by the introduction of this third interval  
11 (the slope of the differences remains the same), but the slant columns are limited to more  
12 realistic values for each molecule, allowing for an effective reduction of the noise. It is  
13 interesting to note that the observed slopes of the differences in BrO and H<sub>2</sub>CO columns are  
14 exactly the same for GOME-2 and OMI retrievals and for different periods of the years (not  
15 shown), pointing to a fundamental spectral effect rather than an instrumental feature.

16 From Fig.2, one can conclude that the net effect of including O<sub>4</sub> in the fits is a positive  
17 correlation of the H<sub>2</sub>CO slant columns with the O<sub>4</sub> differential columns, with H<sub>2</sub>CO column  
18 deviations of less than 10<sup>16</sup> molec.cm<sup>-2</sup> but sometimes reaching 2x10<sup>16</sup> molec.cm<sup>-2</sup> for  
19 extreme O<sub>4</sub> values. Figure 3 presents O<sub>4</sub> differential slant columns retrieved in w1 from OMI  
20 measurements in February and August 2007, giving an idea of the spatial distribution of the  
21 expected differences.

22 Over the Pacific Ocean, the O<sub>4</sub> differential slant columns are generally positive and increase  
23 with the solar zenith angle. Figure 4 presents the averaged zonal variation of the H<sub>2</sub>CO, BrO  
24 and O<sub>4</sub> differential slant columns for January, April, July and October 2007, as a function of  
25 the latitude (a) when O<sub>4</sub> is not included in the fit (v12), and (b) when O<sub>4</sub> is pre-fitted in w1  
26 and included in the fit (v14). Dashed lines show the slant columns before the background  
27 normalisation procedure. The decrease of the H<sub>2</sub>CO slant columns with latitude is greatly  
28 reduced when O<sub>4</sub> is considered, with a minimum impact on the standard deviations of the  
29 columns, owing to the pre-fit of O<sub>4</sub>. However, a positive artificial dependency remains, that is

1 related to ozone absorption interferences and still needs to be corrected (see Sect. 3.2). The O<sub>4</sub>  
2 effect is exactly the opposite for BrO, for which a decrease of the columns is observed for  
3 large O<sub>4</sub> slant columns (about -20%). This can have a significant impact for BrO studies in  
4 polar regions (Salawitch et al., 2010; Choi et al., 2012), but it is beyond the scope of this  
5 paper.

6 Over the continents, where H<sub>2</sub>CO column enhancements are expected, the differential O<sub>4</sub> slant  
7 columns are almost always negative (*i.e.* lower than over oceans), due to the combined effect  
8 of higher altitude and lower surface reflectivity. We observe therefore a decrease of the H<sub>2</sub>CO  
9 columns over continental emission regions, by 0 down to -25%. This reduction is the same for  
10 GOME-2 and OMI, *i.e.* it has no impact on the observed diurnal variations. As an example,  
11 Fig.5 presents the time series of GOME-2 and OMI H<sub>2</sub>CO columns over India and Equatorial  
12 Africa, representative for mid-latitudes and tropical emission regions. Results without (v12)  
13 and with (v14) O<sub>4</sub> included are plotted, as well as the differences between OMI and GOME-2  
14 for the 2 versions. A reduction of the columns is observed when including O<sub>4</sub> in the fits, but  
15 the OMI-GOME-2 differences are equivalent for v12 and v14.

### 16 **3.2 Across-track and zonal reference sector correction**

17 [The use of daily radiance spectra as reference for the DOAS retrievals result in differential](#)  
18 [slant columns, close to zero in the equatorial Pacific. However, Residual-latitude-dependent](#)  
19 biases in the H<sub>2</sub>CO columns, due to unresolved spectral interferences, remain a limiting factor  
20 for the DOAS retrieval of weak absorbers such as H<sub>2</sub>CO. Furthermore, in the case of a 2D-  
21 detector array such as OMI, across-track striping arises, due to imperfect calibration and  
22 different dead/hot pixel masks for the detectors. Such instrumental effects also affect scanning  
23 spectrometers like GOME-2, but since these instruments have one single detector, these errors  
24 do not appear as stripes, but rather as constant offsets (Boersma et al., 2011). These different  
25 retrieval artefacts can be compensated to a certain extent, using normalisation approaches  
26 such as the reference sector correction (Kurosu, 2008; De Smedt et al., ~~2012~~2008). As for the  
27 DOAS reference radiance selection, the reference sector is chosen in the Pacific Ocean, where  
28 the only significant source of H<sub>2</sub>CO is the CH<sub>4</sub> oxidation. [The H<sub>2</sub>CO background is replaced](#)  
29 [by model simulations in the same region.](#) The reference sector correction is also meant to  
30 handle possible time-dependent instrumental degradation effects, for example the evolution of  
31 the OMI stripe artefacts, or the GOME-2 signal degradation (De Smedt et al., 2012). Note that  
32 our analysis show that the most efficient method to reduce across-track stripes in OMI H<sub>2</sub>CO

1 retrievals is to use row-dependent mean radiances as reference spectrum in the DOAS  
2 equation (see Sect. 3.1).

3 We apply a 2-step normalisation of the H<sub>2</sub>CO slant columns. In a first step, a row-dependent  
4 median H<sub>2</sub>CO value is determined in the equatorial Pacific (15°S-15°N, 180°E-240°E) and  
5 subtracted from all the columns (in the case of GOME-2, we use a viewing angle-dependent  
6 correction). The aim is to reduce possible remaining offsets between rows, resulting from the  
7 different detectors. In addition to the destriping procedure, those OMI rows presenting a level  
8 of noise and fitting residuals significantly higher than the average of the other rows for a  
9 particular day, are assigned a bad quality flag, and not further used in our applications. This  
10 criterion removes a few rows from the analysis in 2005, and more importantly, all the rows  
11 affected by the row anomaly which started in June 2007, and further developed over the years  
12 (<http://www.knmi.nl/omi/research/product/rowanomaly-background.php>). As illustrated by  
13 the first line of Fig.6, the affected rows can be identified using the fitting residuals. This  
14 filtering procedure is less systematic than the use of flags provided in the level 1 files, and is  
15 aimed to keep as many observations as possible in the analysis, which is of fundamental  
16 importance to mitigate the noise on formaldehyde observations. In a second step, the  
17 latitudinal dependency of the offset-corrected H<sub>2</sub>CO slant columns is modelled by a  
18 polynomial in the entire reference sector (90°S-90°N, 180°E-240°E). These two corrections  
19 are sequentially subtracted from the global slant columns and replaced by the latitudinal  
20 dependency of the modelled columns in the same region. The result of the across-track and  
21 zonal reference sector correction is illustrated in Fig.4 as a function of the latitude and in the  
22 second line of Fig.6, showing daily maps of normalised OMI H<sub>2</sub>CO slant columns in 2005  
23 and 2014. The loss of coverage due to the OMI row anomaly and to our filtering scheme is  
24 clearly visible when comparing the 2005 and 2014 maps.

### 25 **3.3 Air Mass Factors**

26 In the troposphere, scattering by air molecules, clouds and aerosols lead to complex altitude-  
27 dependent air mass factors. Multiple scattering calculations are required for the determination  
28 of the air mass factors, and the vertical distribution of the absorber has to be assumed *a priori*.  
29 In the case of optically thin absorbers, the formulation of Palmer et al. (2001) is used and has  
30 been described for the TEMIS H<sub>2</sub>CO retrievals in De Smedt et al. (2012). It decouples the  
31 vertical sensitivity of the measurements (the scattering weighting functions, derived with  
32 radiative transfer model calculations) from the vertical profile shape of the species of interest

1 (vertical shape factors, taken from an atmospheric chemistry transport model or from some  
2 other prior knowledge of the vertical distribution of the absorber). Details on these two  
3 calculation steps are given below. The decoupling of the AMF calculation allows addressing  
4 separately the radiative transfer effects, including clouds, and the atmospheric composition of  
5 optically thin absorbers like H<sub>2</sub>CO. Vertical columns might be improved for particular  
6 locations by using more accurate *a priori* profiles, for example based on input from regional  
7 models, ground-based or aircraft measurements. Furthermore, using shape factors from an  
8 atmospheric chemistry model ensures consistency for subsequent evaluation of the model  
9 with the retrieved vertical columns (Barkley et al., 2013). For these reasons, the averaging  
10 kernels and the *a priori* profiles are provided in the level 2 data files for each individual  
11 measurement. The third line of Fig.6 presents global daily maps of AMF in April 2005 and  
12 2014. These maps have been filtered for effective cloud fractions larger than 0.4. Clouds and  
13 *a priori* profile shapes are further discussed below and in the validation Section (5.2).

### 14 **3.3.1 Scattering weighting functions**

15 Scattering weighting functions calculated at 340 nm, using the LIDORT v3.3 radiative  
16 transfer model (Spurr, 2008) are tabulated according to their dependencies with solar, viewing  
17 and relative azimuth angles, surface altitude and surface reflectivity. We use the surface  
18 reflectivity data base derived from OMI by Kleipool et al. (2008), in both GOME-2 and OMI  
19 H<sub>2</sub>CO retrievals. Radiative cloud effects are corrected using the independent pixel  
20 approximation (IPA, Martin et al., 2002) and the respective cloud products of the instruments  
21 provided on the TEMIS website, namely the GOME-2 O<sub>2</sub> A-band Frescov6 product (Wang et  
22 al., 2008) and the OMI O<sub>4</sub> cloud product (Stammes et al., 2008). While the cloud fractions are  
23 in general good agreement between OMI and GOME-2, we observe larger discrepancies for  
24 the cloud altitudes. ~~While~~ The differences between OMI and GOME-2 cloud-free AMFs  
25 range from 0 to -10% (observation geometry effects), but the differences between IPA cloud-  
26 corrected AMFs can reach -20% where and when the cloud lies in the lower troposphere. This  
27 implies an uncertainty of about 10% in the final product, since it is not clear whether those  
28 cloud-related differences reflect real differences in cloud properties or differences in the cloud  
29 retrieval algorithms. This stresses the need for a multi-instrument homogenised cloud  
30 product, for example based on the O<sub>4</sub> absorption band that can be measured by all sensors. No  
31 explicit correction is applied for aerosols but the cloud correction scheme accounts for a large  
32 part of their scattering effect (Boersma et al., 2011). The uncertainty related to aerosol effects

1 [is estimated to be lower than 15% in average \(Leitao et al., 2008; Castellanos et al., 2015;](#)  
2 [Theys et al., 2015\).](#)

### 3 **3.3.2 Vertical shape factors: IMAGES-v2**

4 The *a priori* profile shapes are extracted from daily simulations performed with the IMAGES  
5 model, at 09:30 am for GOME-2 and 13:30 pm for OMI. The IMAGESv2 CTM calculates the  
6 global distributions of 90 long-lived and 41 short-lived trace gases at a resolution of 2°  
7 (latitude) x 2.5° (longitude) and on 40 vertical levels from the surface to the lower  
8 stratosphere. The current model version is thoroughly described in Stavrakou et al. (2013).  
9 The model time step is set to 4 hours. Diurnal changes in the photolysis and kinetic rates,  
10 meteorological fields, and the emissions are taken into account through correction factors  
11 calculated from a simulation with a 20-min time step (Stavrakou et al., 2009a) and applied to  
12 model runs using longer time steps. Simulations have been performed for all years between  
13 2005 and 2013, spun up by a period of four months.

14 Anthropogenic VOC emissions are obtained from the RETRO 2000 global database (Schultz  
15 et al., 2008) and are kept constant throughout the years. Over Asia, RETRO is overwritten by  
16 the REASv2 inventory (Kurokawa et al., 2013) until 2008, whereas 2008 values are used for  
17 more recent years. Anthropogenic VOC emissions are equal to 147 and 150 Tg VOC in 2005  
18 and 2006, respectively, and to 156 Tg VOC for the following years. Emissions of isoprene  
19 from vegetation are taken from the MEGAN-MOHYCAN-v2 inventory (Stavrakou et al.,  
20 2014). The global annual fluxes range between 323 Tg and 363 Tg isoprene, the lowest and  
21 the highest values corresponding to 2008 and to 2010, respectively. Open biomass burning  
22 emissions are taken from the GFEDv3 inventory (van der Werf et al., 2010) until 2011,  
23 whereas a climatological mean based on 1997-2011 GFEDv3 emissions is used for 2012 and  
24 2013. Global annual fire emissions range between 70 Tg VOC (in 2009) and 105 Tg VOC (in  
25 2010).

26 The photochemical production of H<sub>2</sub>CO is estimated at ca. 1600 Tg annually. [The main](#)  
27 [formaldehyde sinks are the oxidation by OH \(Sander et al., 2011\) which leads to CO](#)  
28 [production and conversion of OH to HO<sub>2</sub>, and two photolysis reactions which produce CO](#)  
29 [and HO<sub>2</sub> radicals. Based on IMAGESv2 model calculations, photolysis is by far the dominant](#)  
30 [removal process, estimated at 71% of the global sink, whereas the OH sink is less efficient](#)  
31 [\(26%\). Photolysis of H<sub>2</sub>CO and oxidation by OH account for about 71% and 26% of the](#)

1 | ~~global sink, respectively, while d~~Dry and wet deposition account for the remainder (<3%).

2 The global photochemical H<sub>2</sub>CO lifetime is estimated at 4.5 hours.

### 3 **3.4 Quality criteria of the H<sub>2</sub>CO vertical columns**

4 The H<sub>2</sub>CO level-2 files of version 14 are provided in HDF-5 format. They include all the  
5 intermediate quantities of the H<sub>2</sub>CO retrieval from slant columns to vertical columns, air mass  
6 factors (cloud free or including the IPA cloud correction) and averaging kernels. A detailed  
7 error budget is also provided. Error contributions resulting from each step of the retrieval to  
8 the final vertical column error are provided separately, including their random and systematic  
9 parts (for details see De Smedt et al., 2014). This allows for an estimation of the total error on  
10 the column averages. A quality flag is also provided for each observation. The following  
11 criteria are considered for assigning a bad (<0) quality flag:

- 12 • per orbit: fit residuals larger than three times the averaged fit residual
- 13 • per day: unsuccessful across-track and zonal reference sector correction (see Sect. 3.2)
- 14 • per orbit: corrected slant columns lower than the mean corrected column minus three  
15 times the column standard deviations
- 16 • per pixel: effective cloud fractions larger than 0.4
- 17 • per pixel: snow or ice flag in the cloud product
- 18 • per pixel: solar zenith angles larger than 70°
- 19 • per pixel: individual vertical column errors larger than three times the column

## 20 **4 GOME-2 and OMI H<sub>2</sub>CO vertical columns**

### 21 **4.1 Background values and precision**

22 Figure 7 presents time series of monthly averaged GOME-2A&B and OMI H<sub>2</sub>CO vertical  
23 columns in the reference sector (first panel) together with the retrieval residuals, standard  
24 deviations and number of observations (second to fourth panels). In the first panel,  
25 uncorrected columns are displayed with dotted lines, while background-corrected columns are  
26 represented with plain lines. From this figure, it can be concluded that even the uncorrected  
27 OMI H<sub>2</sub>CO columns are remarkably stable in time (yet across-track and zonal corrections are  
28 needed). In the case of GOME-2A retrievals however, the use of radiances as DOAS  
29 reference spectra does not completely compensate for a weakening of the signal over the

1 years, although the loss of signal is reduced by a factor 2 compared to the use of solar  
2 irradiances, as it was the case in version 12 (De Smedt et al., 2012 ).

3 Regarding the fit residuals and the standard deviations of the monthly averaged H<sub>2</sub>CO  
4 columns, the OMI retrievals also show a very good stability in time, the random errors having  
5 increased by less than 4% in 10 years, while the GOME-2A random errors have increased at a  
6 rate of about 10% per year due to degradation effects. Accordingly, the OMI H<sub>2</sub>CO individual  
7 pixel precision is about  $7 \times 10^{15}$  molec.cm<sup>-2</sup>, while the GOME-2A precision degraded from  
8  $5 \times 10^{15}$  molec.cm<sup>-2</sup> to  $8 \times 10^{15}$  molec.cm<sup>-2</sup>. The GOME-2B performances are found to be  
9 equivalent to those of GOME-2A in its early lifetime; however a degradation rate similar to  
10 the one of GOME-2A is expected.

11 As shown in the fourth panel, the number of OMI observations passing our quality criteria has  
12 decreased by 40% since the beginning of the mission due to the row anomaly issue. The  
13 resulting change of sampling has an impact on the calculated monthly means that should be  
14 taken into account in trend analyses (see Sect. 6).

## 15 **4.2 Continental emission regions**

16 Figure 8 presents multi-year maps of H<sub>2</sub>CO vertical columns retrieved from GOME-2A and  
17 OMI between 2007 and 2013, and their averaged absolute differences over the same period.  
18 For these maps (and further below), the quality flags as defined in Sect. 3.4 have been used  
19 to filter out invalid satellite observations. We observe consistent H<sub>2</sub>CO distributions, with  
20 highest columns over tropical regions, India, China, Southeast China and Southeastern US.  
21 We also observe noticeable differences, OMI showing more elevated columns at mid-  
22 latitudes and over regions with moderate H<sub>2</sub>CO concentrations, but lower columns in the  
23 Tropics, where the H<sub>2</sub>CO columns are highest. To better visualize the regional structures,  
24 Figure 9 presents H<sub>2</sub>CO vertical columns derived from both sensors over Europe (from April  
25 to September) and Asia (from March to November), averaged between 2007 and 2013.

26 Different effects can be identified:

27 (1) The better horizontal resolution of OMI allows for a better identification of strong  
28 hotspots over localised H<sub>2</sub>CO sources, as for example over cities like Mexico, Pretoria,  
29 Hong-Kong/Guangdong, Beijing, Cairo, Tehran and Mumbai. This resolution effect can  
30 also be identified along coastal areas (Cape-Town/South Africa, the Algerian Coast, the

1 Turkish coasts, Kerala/India or California/US) and along mountain chains (North of  
2 India, Pyreneens, Alps).

3 | (2) The increase in H<sub>2</sub>CO columns between 9h30 and 13h30 –appears to be largest over  
4 South Europe, South Australia, Northeastern China or Central Siberia, while an inverse  
5 diurnal variation, of equivalent magnitude, is observed over the Equatorial forests of  
6 Amazon, Africa and Indonesia. The same effect has been reported for glyoxal columns,  
7 another important indicator of NMVOC emissions (Alvarado et al., 2014). The observed  
8 diurnal variations will be further discussed in Sect. 5, using ground-based observations.

9 (3) Differences in retrieval sensitivity appear when comparing GOME-2 and OMI. We note  
10 that over oceans, poleward of 30°, OMI columns are found to be higher than GOME-2.  
11 These differences cannot be explained only by the diurnal variation of the CH<sub>4</sub> oxidation,  
12 but they also result from a relative lack of sensitivity of GOME-2 to lower tropospheric  
13 H<sub>2</sub>CO in comparison to OMI. This is due to the morning overpass time of GOME-2,  
14 which leads to larger SZA over mid- and high- latitudes (therefore to lower sensitivity to  
15 the lowest atmospheric layers) and to lower H<sub>2</sub>CO concentrations (therefore to lower  
16 absorption).

17 | [Table 2](#) provides a list of regions where the highest annual H<sub>2</sub>CO columns are  
18 observed. The annual means are provided with the standard deviations of the monthly  
19 averaged columns, as an indication of the amplitude of the seasonal variations. The same  
20 regions stand out both in the GOME-2 and OMI observations, with equivalent seasonal  
21 variations. However, as shown on the maps, the GOME-2 observations are maximal in the  
22 Tropics (Africa and South America), while the OMI observations maximize over megacities  
23 like Hong-Kong/Guandong or Delhi. As can be deduced from the seasonal deviations  
24 | provided in [Table 2](#), the largest seasonal variations are found in Rondônia/Brazil and  
25 Tianjin/China, for both instruments. Interestingly, the OMI H<sub>2</sub>CO columns exhibit large  
26 seasonal variations along the coasts of India, which are less pronounced in GOME-2  
27 observations (Goa, Kerala, Orissa). This could be explained by resolution effects or by  
28 diurnal variations in the H<sub>2</sub>CO production and loss processes during the monsoon season. In  
29 Sect. 6, the complete time series of monthly averaged H<sub>2</sub>CO vertical columns in China, India,  
30 Africa, South America, North America and Europe are presented.



1 Table 2: GOME-2A and OMI largest annual mean H<sub>2</sub>CO columns between 2007 and 2013  
 2 (10<sup>15</sup> molec.cm<sup>-2</sup>)

Continent	Country	Province/State	GOME-2A	seas.dev.	OMI	seas.dev.
Africa	Congo		9.37	1.71	7.74	1.27
Africa	Ghana		8.97	2.79	8.47	2.6
Africa	Sierra Leone		10.26	3.76	9.21	3.95
Africa	Togo		8.96	2.77	8.49	2.62
Asia	Bangladesh		9.76	1.48	10.5	2.3
Asia	Cambodia		8.63	3.39	8.74	3.71
Asia	China	Anhui	7.96	2.85	9.21	3.18
Asia	China	Guangdong	8.59	1.44	9.82	1.34
Asia	China	Guangxi	8.6	1.88	9.04	1.63
Asia	China	Hong Kong	9.05	2.25	11.86	3.53
Asia	China	Jiangsu	7.81	3.24	9.39	3.64
Asia	China	Tianjin	7.8	4.75	9.86	3.84
Asia	India	Bihar	9.67	1.44	10.5	2.01
Asia	India	Dadra and Nagar Haveli	7.93	2.32	9.5	3.92
Asia	India	Daman & Diu	8.35	2.5	9.27	4.31
Asia	India	Delhi	10.2	2.85	12.2	2.91
Asia	India	Goa	7.73	3.08	9.41	4.61
Asia	India	Haryana	9.24	2	10.95	2.16
Asia	India	Kerala	7.26	2.76	9.34	4.10
Asia	India	Orissa	9.06	2.38	9.92	3.01
Asia	India	Punjab	9.09	1.97	10.97	2.26
Asia	India	Tripura	8.76	1.68	9.79	2.47
Asia	India	Uttar Pradesh	8.79	1.58	10.2	1.69
Asia	India	West Bengal	9.85	1.61	10.79	2.34
Asia	Singapore		9.14	2.79	8.9	3.62
Asia	Thailand		8.5	3.17	8.79	3.58
South America	Brazil	Mato Grosso	11.19	4.23	8.72	3.82
South America	Brazil	Rondônia	11.04	5.29	8.8	4.71
South America	Brazil	Tocantins	10.55	2.66	8.62	2.57
South America	Bolivia		8.91	2.96	6.19	2.46

3

## 1 5 Validation using MAX-DOAS measurements

2 We validate the observed satellite column variations using ground-based measurements  
 3 operated by BIRA-IASB at seven stations: Cabauw, Brussels, the Haute-Provence  
 4 Observatory, Beijing, Xianghe, Bujumbura and Reunion Island. Details of the ground-based  
 5 measurements, and related publications, are summarized in [Table 3](#)~~Table–3~~. In Reunion  
 6 Island, measurements have been performed with an FTIR instrument (Vigouroux et al., 2009),  
 7 while in all other stations, multi-axis DOAS (MAX-DOAS) instruments have been used.  
 8 Although installed at the end of 2013, the Bujumbura instrument could only deliver 6 months  
 9 of data due to a failure of the UV channel in April 2014.

10 The MAX-DOAS measurement technique has been developed to retrieve tropospheric trace  
 11 gas total columns and profiles. The most recent generation of MAX-DOAS instruments allow  
 12 for measurement of aerosols and a number of tropospheric pollutants, such as NO<sub>2</sub>, H<sub>2</sub>CO,  
 13 SO<sub>2</sub>, HONO, O<sub>4</sub> and CHOCHO (see e.g. Heckel et al., 2005; Clémer et al., 2010; Irie et al.,  
 14 2011; Ma et al., 2013, Hendrick et al., 2014; Vlemmix et al., 2014; Wang et al., 2014). H<sub>2</sub>CO  
 15 slant columns have been retrieved from the six MAX-DOAS instruments using consolidated  
 16 settings published in Pinardi et al. (2013). While the scientific-grade instruments installed in  
 17 China and in Bujumbura allow for vertical profile retrievals by optimal estimation (Clémer et  
 18 al., 2010, Hendrick et al., 2014), only vertical columns could be retrieved from the less  
 19 sensitive instruments operated in the other stations. We therefore use the ground-based total  
 20 columns to validate the seasonal and diurnal variations in all stations, while a more  
 21 quantitative validation making use of the profile information is performed in Beijing/Xianghe  
 22 and Bujumbura.

23 Table 3: Summary of ground-based measurements available at BIRA-IASB

Station/Country (lat, long)	Instrument	Period	Retrieved quantity	Reference
Cabauw/The Netherlands (52° N, 5° E)	MAX-DOAS	18/06/2009-21/07/2009	VCD	Pinardi et al., 2013
Brussels/Belgium (50.78° N, 4.35° E)	Mini-MAX- DOAS	01/05/2011-23/04/2012	VCD	Gielen et al., 2014
OHP/France (43.94°N, 5.71°E)	MAX-DOAS	26/06/2007-20/03/2013	VCD	Valks et al., 2011
Beijing/China (39.98°N, 116.38°E)	MAX-DOAS	03/07/2008-17/04/2009	VCD + Profile	Vlemmix et al., 2014
Xianghe/China (39.75° N, 116.96° E)	MAX-DOAS	07/03/2010-26/12/2013	VCD + Profile	Vlemmix et al., 2014
Bujumbura/Burundi (3°S, 29°E)	MAX-DOAS	25/11/2013-22/04/2014	VCD + Profile	-
Reunion Island/France (20.9° S, 55.5° E)	FTIR	01/08/2004-25/10/2004 21/05/2007-15/10/2007	VCD	Vigouroux et al., 2009

## 5.1 Seasonal and diurnal variations

Figure 10 presents the mean diurnal variations of the H<sub>2</sub>CO columns, averaged by season over the complete period of measurements, as observed from the ground (in black) and from space (in red). For this figure, measurements in Cabauw and Brussels have been combined, as well as those in Beijing and Xianghe. We therefore consider five validation sites, representative for different emission sources and illumination conditions. The satellite observations are averaged within 100 km around each station, and filtered as described in Sect. 3.4. Although larger than the typical length of air masses sampled by a MAX-DOAS spectrometer, which is less than a few tens of kilometres (Gomez et al., 2014), this radius allows including enough satellite pixels to ensure significant analysis. MAX-DOAS observations are filtered using similar thresholds on total errors (retrieval errors larger than three times the columns) and solar zenith angles (70°). Cloudy observations are excluded using the multiple-scattering cloud filter described in Gielen et al. (2014) since such sky conditions can potentially degrade the quality of the MAX-DOAS retrievals. The error bars shown in Fig.10 include random and systematic error contributions. The random errors are taken as the standard deviations of the daily averaged observations divided by the square root of the number of days considered in the seasonal means. Winter time satellite columns in Europe and China are not shown, because of their reduced quality resulting from a combination of low H<sub>2</sub>CO values and lower satellite sensitivity close to the surface. Furthermore, ~~Table 4~~ [Table 4](#) provides quantitative differences between the afternoon and morning observations for the five locations, in each season. Both for satellite and ground-based measurements, the correlations between 09h30 and 13h30 columns have been considered to estimate the error on the differences, using a standard uncertainty propagation formulation.

The first conclusion that can be drawn from Fig.10, is the general underestimation of the satellite columns by 0 up to 50% compared to the MAX-DOAS observations especially during the seasons of maximum concentration. In comparison, the agreement is found to be better with tropical background FTIR measurements (Vigouroux et al., 2009). The second conclusion is the qualitative agreement between the satellite and ground-based measurements regarding the H<sub>2</sub>CO column variations. The seasonal variations are in a very good agreement, with a maximum in summertime in mid-latitude regions, especially in the South of France and in China. Regarding diurnal variations, the differences between the OMI and GOME-2

1 columns should reflect the change in H<sub>2</sub>CO between 9h30 and 13h30. However, uncertainties  
 2 on the satellites H<sub>2</sub>CO columns are large (De Smedt et al., 2012), and so are the errors on their  
 3 differences, as reflected in [Table 4](#). Nevertheless, the sign of the differences between  
 4 GOME-2 and OMI agrees well with both MAX-DOAS and FTIR measurements. Early  
 5 afternoon values are almost always equal to or larger than mid-morning values, except in  
 6 Bujumbura where morning columns are larger. The amplitude of the diurnal variation inferred  
 7 from ground-based and satellite data is also in relatively good agreement, as can be seen by  
 8 comparing the slopes between 09h30 and 13h30, and from [Table 4](#). This is less true  
 9 for summertime columns in Brussels/Cabauw, or for spring time columns in OHP and in  
 10 Beijing/Xianghe, where the differences OMI-GOME-2 are larger than the corresponding  
 11 MAX-DOAS variations, possibly pointing to an underestimation of the GOME-2 retrievals in  
 12 those regions/periods.

13 Table 4: Mean diurnal variations of the H<sub>2</sub>CO columns as observed with ground-based and  
 14 satellite instruments. Values are given for 5 regions where BIRA-IASB operates ground-  
 15 based measurements: Cabauw and Brussels, OHP, Beijing and Xianghe, Bujumbura and  
 16 Reunion Island. Details of the ground-based measurements are summarized in [Table 3](#)

Station	Instrument	Absolute 13h30-9h30 difference (10 <sup>14</sup> molec.cm <sup>-2</sup> )			
		DJF	MAM	JJA	SON
Cabauw, Brussels	OMI-GOME-2	-	14±24	25±27	21±37
	MAX-DOAS	-	12±25	08±23	20±21
OHP	OMI-GOME-2	-	30±27	34±34	29±34
	MAX-DOAS	00±08	09±10	35±24	18±14
Beijing, Xianghe	OMI-GOME-2	-	34±31	39±67	25±32
	MAX-DOAS	-12±08	-08±18	44±40	18±15
Bujumbura	OMI-GOME-2	-22±26	-28±23	-24±41	-22±40
	MAX-DOAS	-49±28	-50±28	-	-
Reunion Island	OMI-GOME-2	14±22	13±27	16±18	14±30
	FTIR	-	13±20	07±06	06±09

17

## 18 5.2 Vertical columns in China and in Burundi

19 Monthly averaged columns retrieved from GOME-2, OMI and MAX-DOAS measurements  
 20 are shown in Fig.11 and 12, respectively for the stations of Beijing/Xianghe (averaged over  
 21 2008-2013) and Bujumbura (averaged over 2013-2014). As the instrument in Bujumbura was  
 22 installed at the end of 2013, we use the more recent H<sub>2</sub>CO columns from GOME-2B, but the

1 GOME-2A columns are shown up to the end of 2013 (light green). The monthly means have  
2 been calculated using days in common between respectively GOME-2 and morning MAX-  
3 DOAS data (8h-11h), and OMI and afternoon MAX-DOAS data (12h-15h). Again, the  
4 satellite observations are averaged within 100 km around each station, and filtered as  
5 described in Sect. 3.4. In China, the resulting number of correlative days amounts to 711 for  
6 GOME-2 and 807 for OMI. In Burundi, we obtain respectively ~~78-58~~ and ~~908~~ days over the 5  
7 months of measurements. [Quantitative comparisons between GOME-2, OMI and MAX-](#)  
8 [DOAS columns are provided in Table 5.](#)

9 Three sets of monthly averages are ~~shown-used~~ for the satellites, in order to evaluate  
10 uncertainties related to the AMF calculation: (1) the vertical columns calculated using the  
11 IMAGES *a priori* profile shapes and no cloud correction (plain dots), (2) same but applying  
12 the IPA to correct for cloud radiative effects (empty dots), and (3) the vertical columns  
13 calculated using as *a priori* the H<sub>2</sub>CO vertical profile shapes retrieved from the MAX-DOAS  
14 measurements and the IPA cloud correction (diamonds). In Fig.11, the lower panels display  
15 scatter plots between monthly averaged observations for the satellite retrievals versions (2)  
16 and (3). In Beijing/Xianghe, the correlation coefficients are similarly high between the  
17 different AMF versions, and slightly better for OMI than GOME-2 (about 0.8 for GOME-2  
18 and ~~0.85-9~~ for OMI). However, the slopes ~~and offsets~~ of the regression lines between the  
19 satellite and ground-based observations show larger dependency on the AMF calculation  
20 settings. We find slopes of respectively ~~0.78~~, 0.6 and 0.9 for GOME-2 and ~~0.86~~, ~~0.67~~, ~~1-1~~ for  
21 OMI. In Bujumbura, the number of ground-based measurements is unfortunately not large  
22 enough to draw conclusions about the correlations, ~~and we rather provide the mean~~  
23 ~~differences between the satellite and MAX-DOAS columns.~~ The ~~slopes-mean differences~~ are  
24 found to be respectively ~~0.7, 0.7, and 1.14-2.8,-2.6~~ and ~~1.1 x10<sup>15</sup> molec.cm<sup>-2</sup>~~ for GOME-2  
25 and ~~0.6, 0.7, and 0.9-3.3,-2.6~~ and ~~-0.9 x10<sup>15</sup> molec.cm<sup>-2</sup>~~ for OMI. Results in China and in  
26 Burundi suggest that the cloud correction has little systematic influence, and therefore a  
27 limited effect on the monthly averaged columns, as indicated by the ~~mean~~ difference  
28 ~~variations~~ between ~~slopes~~-(1) and (2) (~~about-less than~~ 10%). On the contrary, the *a priori*  
29 vertical profile has a larger systematic effect on the vertical columns, as indicated by the  
30 different ~~resultsee~~ between ~~slopes~~-(2) and (3). Both for GOME-2 and OMI, switching from  
31 modelled to measured profile shapes increase the H<sub>2</sub>CO columns by 20 to 50%, bringing the  
32 satellites and ground-based observations to a satisfactory agreement within 15%. [We note that](#)  
33 [the OMI columns in Beijing/Xianghe present a positive offset of about 3x10<sup>15</sup> molec.cm<sup>-2</sup>.](#)

1 The large effect of the *a priori* profile shape is explained by the vertical sensitivity of the  
 2 satellite measurements, decreasing strongly in the lowest atmospheric layers, and by the shape  
 3 of the H<sub>2</sub>CO vertical distribution, peaking near the surface. This is illustrated in Fig.13, where  
 4 the H<sub>2</sub>CO profile shapes modelled by IMAGES and those retrieved from the MAX-DOAS  
 5 measurements, are plotted next to the satellite vertical sensitivity for June 2010 in Xianghe. It  
 6 must be noted that the retrieved MAX-DOAS profiles also have their own uncertainties  
 7 (Vlemmix et al., 2014), however using them to re-calculate the satellite AMFs allows to  
 8 remove from the comparison the error associated to the *a priori* profile shapes (Eskes and  
 9 Boersma, 2003). Indeed, only the shape of the *a priori* profiles impact the satellite AMFs, not  
 10 their total columns (Palmer et al., 2001). The satellite averaging kernels (AKs) are much  
 11 closer in shape to the FTIR AKs than to the MAX-DOAS retrievals, which may explain the  
 12 better agreement of the columns (Vigouroux et al., 2009).

13 Table 5: Results of the comparisons between GOME2, OMI and MAX-DOAS columns,  
 14 shown in Figure 11 for Beijing/Xianghe and in Figure 12 for Bujumbura. Three satellite VCs  
 15 are used: IMAGES a.p. profile/no cloud correction, IMAGES a.p. profile/IPA cloud  
 16 correction, and MAX-DOAS a.p. profile/IPA cloud correction. Mean differences (satellite-  
 17 MAX-DOAS) are given in both stations. In Beijing/Xianghe, the slopes and offsets of a linear  
 18 regression between MAX-DOAS and satellite columns are provided.

	<u>GOME-2</u>			<u>OMI</u>		
	<u>IMAGES</u> <u>ap profiles</u> <u>No cloud</u> <u>correction</u>	<u>IMAGES</u> <u>ap profiles</u> <u>Cloud</u> <u>correction</u>	<u>MAXDOAS</u> <u>ap profiles</u> <u>Cloud</u> <u>correction</u>	<u>IMAGES</u> <u>ap profiles</u> <u>No cloud</u> <u>correction</u>	<u>IMAGES ap</u> <u>profiles</u> <u>Cloud</u> <u>correction</u>	<u>MAXDOAS</u> <u>ap profiles</u> <u>Cloud</u> <u>correction</u>
<u>Beijing/Xianghe</u>						
<u># common days</u>	<u>711</u>			<u>807</u>		
<u>Mean Difference</u> <u>10<sup>15</sup> molec.cm<sup>-2</sup></u> <u>(%)</u>	<u>-4.4</u> <u>(-33%)</u>	<u>-5.4</u> <u>(-41%)</u>	<u>-1.4</u> <u>(-11%)</u>	<u>-2.8</u> <u>(-19%)</u>	<u>-3.4</u> <u>(-24%)</u>	<u>2.1</u> <u>(15%)</u>
<u>Correlation</u> <u>coefficient</u>	<u>0.82</u>	<u>0.85</u>	<u>0.84</u>	<u>0.88</u>	<u>0.87</u>	<u>0.88</u>
<u>Slope</u>	<u>0.80</u>	<u>0.64</u>	<u>0.91</u>	<u>0.61</u>	<u>0.62</u>	<u>0.97</u>
<u>Offset</u> <u>10<sup>15</sup> molec.cm<sup>-2</sup></u>	<u>-1.6</u>	<u>0.6</u>	<u>0.2</u>	<u>3.0</u>	<u>2.3</u>	<u>3.2</u>
<u>Bujumbura</u>						
<u># common days</u>	<u>58</u>			<u>90</u>		

<a href="#">Mean Difference</a>						
<a href="#">10<sup>15</sup> molec.cm<sup>-2</sup></a>	<a href="#">-2.8</a>	<a href="#">-2.6</a>	<a href="#">1.1</a>	<a href="#">-3.3</a>	<a href="#">-2.6</a>	<a href="#">-0.9</a>
<a href="#">(%)</a>	<a href="#">(-29%)</a>	<a href="#">(-27%)</a>	<a href="#">(11%)</a>	<a href="#">(-39%)</a>	<a href="#">(-31%)</a>	<a href="#">(-9%)</a>

1  
2  
3 The effect of the rather coarse resolution of the global CTM on the modelled profiles (here  
4 2°x2.5°) needs to be further investigated, [particularly for anthropogenic sources](#), as well as  
5 possible other effects of vertical transport and chemical processes. For example, a recent  
6 analysis of in situ measurements in Beijing (Chen et al. 2014) indicated that primary sources  
7 of H<sub>2</sub>CO are responsible for as much as about 32% of the total H<sub>2</sub>CO source in the area,  
8 suggesting a strong underestimation of this direct source in current inventories. In the  
9 standard IMAGES simulation using the REASv2 inventory, this primary source contributes to  
10 less than 1% of the total H<sub>2</sub>CO source during the summer. Adjusting the primary source in the  
11 model to match the direct fraction estimated by Chen et al. (2014) results in the IMAGES  
12 profile shapes also shown on Fig.13 (IMAGES prim1), found to agree much better with  
13 MAX-DOAS profiles. It is not clear whether this large apparent contribution of direct  
14 emissions is real or reflects e.g. a fast chemical production from highly reactive VOCs  
15 currently not well represented in inventories and models. Whatever the reasons might be for  
16 the improved model profiles in comparison with MAX-DOAS data, generalization to other  
17 areas would appear premature. It is worth noting that the IMAGES model was found to very  
18 well reproduce observed vertical profiles of formaldehyde measured during aircraft  
19 campaigns over the United States and over oceans (Stavrakou et al., 2009c). More studies and  
20 measurements are required to refine our understanding of factors governing the formaldehyde  
21 vertical profiles.

## 22 **6 Long-term variations**

23 [Formaldehyde columns are mainly formed by oxidation of NMVOCs from biogenic, biomass](#)  
24 [burning and anthropogenic sources. Column inter-annual variabilities are mainly driven by](#)  
25 [fire events and temperature changes \(Millet et al., 2008; Barkley et al., 2009; Stavrakou et al.,](#)  
26 [2014\). However, over industrialised regions, changes in anthropogenic emissions have also](#)  
27 [been identified as drivers of observed H<sub>2</sub>CO column trends \(De Smedt et al., 2010; Zhu et al.,](#)  
28 [2014; Khokhar et al., 2015; Mahajan et al., 2015; Stroud et al., 2015\).](#)

1 The TEMIS time series of morning H<sub>2</sub>CO columns, consisting of GOME, SCIAMACHY and  
2 GOME-2A & B data, spans over 15 years of observations, while the afternoon time series  
3 derived from OMI covers 10 years of observations. Meaningful trend studies are therefore  
4 possible. We used a linear model with a seasonal component to fit the monthly averaged  
5 columns, separately for the morning and afternoon time series. The error and statistical  
6 significance of the inferred trends are also estimated, as described in De Smedt et al. (2010). It  
7 is understood that the accuracy of such a trend analysis is limited in the case of H<sub>2</sub>CO  
8 because, like for the amplitude of the diurnal variation (between 1 and 4 x 10<sup>15</sup> molec.cm<sup>-2</sup>),  
9 the amplitude of the trends to be detected (about 1-2 x 10<sup>15</sup> molec.cm<sup>-2</sup> in 10 years) is one  
10 order of magnitude smaller than the H<sub>2</sub>CO columns, ~~and~~ their errors and their seasonal  
11 variations. However, statistically significant trends are detected in several regions  
12 independently with both datasets, which gives confidence in our H<sub>2</sub>CO column long-term  
13 variation estimates.

14 Figures 14 to 16 present the time series of monthly averaged H<sub>2</sub>CO vertical columns in India,  
15 China, India, South America, Africa, ~~South America~~, North America and Europe. On those  
16 figures, the mid-morning time series (in green) combine SCIAMACHY (2003-2011),  
17 GOME-2A (2007-2013) and B (from 2013) measurements, while early afternoon columns (in  
18 red) are derived from OMI measurements. The good agreement between the different  
19 morning observations can be noted (see also De Smedt et al., 2012), as well as the previously  
20 described differences and similarities between afternoon and morning time series. In  
21 particular, the higher morning columns over tropical forest are also observed in the  
22 SCIAMACHY time series. The results of our trend analysis are displayed whenever they  
23 have been found statistically significant, i.e. if the absolute value of the calculated trend is  
24 larger than twice the associated error. Furthermore, Fig.17 presents a global map of the most  
25 significant trends found in the OMI H<sub>2</sub>CO columns between November 2004 and August  
26 2014. The analysis has been performed at country level on the global scale, at province or  
27 state level in the largest countries, and in a radius of 100 km around the main urban areas.

28 For these calculations, “sampling-corrected” OMI columns are used. Indeed, as shown in  
29 Fig.6, the OMI daily spatial sampling has been reduced over the years because of the growing  
30 row anomaly. This has an impact on the H<sub>2</sub>CO monthly averaged columns, which tend to  
31 decrease in time if this effect is not taken into account. This decrease could be explained by  
32 the fact that a large number of central rows (rows 27-44 since January 2009), which have the  
33 finest spatial resolution, are affected by the anomaly and need to be filtered out. For this



1 reason, we have calculated special OMI monthly averages, selecting only the rows that were  
2 still valid at the end of 2013. The net effect is a slight decrease of the columns at the  
3 beginning of the time series, almost negligible when looking at the absolute values, but  
4 significant when considering trends.

5 In India and China (Fig.14), we observe increasing H<sub>2</sub>CO columns. The spatial distribution of  
6 the observed increases, and their values, is similar in the SCIAMACHY-GOME-2 and OMI  
7 time series. No trend is observed in the largest Chinese cities like Beijing and the Pearl River  
8 Delta. However, large positive trends are detected in the surrounding provinces, located in the  
9 centre of the country (e.g. Henan and Hunan provinces). This spatial shift of the Chinese  
10 trends outside the largest cities was already reported in De Smedt et al. (2010), using GOME  
11 and SCIAMACHY observations between 1997 and 2009. This might be related to a larger  
12 influence of newer, cleaner technologies in these mega-cities compared to the Chinese  
13 heartland (Zhang et al., 2009). In India, we observe positive trends almost everywhere, with  
14 an (absolute and relative) amplitude increasing from South to North of the country. Values  
15 obtained from morning and afternoon observations are found to be consistent, and similar to  
16 the ones reported for the period 1997-2009.

17 | In Eastern US and Western Europe ([Figure 16](#)~~Fig-15~~), we observe decreasing trends, more  
18 clearly detected in the OMI retrievals. The morning H<sub>2</sub>CO observations have larger errors and  
19 the wintertime observations cannot be used. Negative trends have previously been reported,  
20 but with less statistical significance (De Smedt et al., 2010). Here, the correlation between the  
21 SCIAMACHY/GOME-2 and the OMI observations is found to be good even in these  
22 retrieval-sensitive regions, giving some new confidence in the morning time series over North  
23 Europe and North America. The expected negative trends due to emission controls are better  
24 captured by the early afternoon observations of OMI, allowing the detection of significant  
25 decreases in Germany and France. In South European countries like Spain, negative trends  
26 can be detected in both time series, which show remarkable correlation, with for example very  
27 low values in spring 2008, that could be related to power plant emission reductions as  
28 reported for NO<sub>2</sub> by Curier et al. (2014).

29 | Over the African continent (right panels of Fig.[Figure 15](#)~~46~~), almost no change is detected,  
30 with the exception of Madagascar, and its capital Antananarivo, where a large positive trend  
31 is found in the OMI time series. The origin of this trend is not understood for the moment. No  
32 similar increase has been reported in satellite NO<sub>2</sub> observations over this region (see for

1 | example Hilboll et al., 2013). In South America (left panels of [Figure 15](#)~~Fig-16~~), we observe  
2 | very significant negative trends over the Brazilian state of Rondônia, of  $-2 \times 10^{14}$  molec.cm<sup>-2</sup>  
3 | yr<sup>-1</sup>, in both morning and afternoon time series. Such negative trends are also present, with  
4 | lower amplitudes, in the surrounding Brazilian States covered by the Amazon forest. The  
5 | largest H<sub>2</sub>CO columns worldwide are observed in those regions, with very large variations  
6 | between the dry and the wet season (see [Table 2](#)~~Table-2~~). Figure 18 shows the H<sub>2</sub>CO columns  
7 | in Rondônia between 2003 and 2014, the MODIS (on Terra and Aqua satellites) Collection 5  
8 | Active Fire Product (<ftp://fuoco.geog.umd.edu/modis/C5/cmg>, Giglio, 2013), and the yearly  
9 | deforestation rates reported by the Brazilian INPE (<http://www.obt.inpe.br/prodes/index.php>),  
10 | in selected Amazonian States. Deforestation in Rondônia amounted to 23% of its surface area  
11 | between 1988 and 2013. This is the highest surface ratio among all Brazilian states. The years  
12 | showing the highest deforestation rates are 1995 and 2004. In Rondônia, a strong decrease of  
13 | the deforestation rate has been observed between 2005 and 2010, and a slight increase is  
14 | again observed since 2011. As illustrated by the middle panel of Fig.18, we find high  
15 | correlation coefficients between the SCIAMACHY/GOME-2 and OMI H<sub>2</sub>CO columns and  
16 | the MODIS fire product, of respectively 0.8 and 0.9 (see also Barkley et al., 2013). We have  
17 | also compared the H<sub>2</sub>CO vertical columns with GISS surface temperature anomalies  
18 | (Gridded Monthly Maps of Temperature Anomaly Data, <http://data.giss.nasa.gov/gistemp/>;  
19 | Hansen et al., 2010), but no correlation was found in this area. Vegetation burning related to  
20 | deforestation appears to have strongly decreased in Rondônia, while it is not yet the case in  
21 | the surrounding areas. It should be noted that the strong enhancement of natural fire  
22 | emissions during very dry years (such as 2005 and 2010), somehow reduce the observed  
23 | downward trend in the fires and H<sub>2</sub>CO columns, and therefore the correlation with the  
24 | reported deforestation rates.

25 | Besides these direct effects of biomass burning activity changes on the H<sub>2</sub>CO columns, more  
26 | studies are needed in order to assess the impact of deforestation and land use changes, and  
27 | possibly related meteorological changes, on biogenic NMVOC emissions (Stavrakou et al.,  
28 | 2014). It is worth noting that BIRA-IASB is currently installing a FTIR instrument in Porto  
29 | Velho, in the Rondônia state. This will bring new information concerning both the diurnal  
30 | cycle of H<sub>2</sub>CO columns and its chemistry, and more generally, on the carbon cycle molecules  
31 | (CO<sub>2</sub>, CH<sub>4</sub>, CO and NMVOCs).

## 1 **7 Conclusions**

2 This paper presents a new version of the BIRA-IASB formaldehyde retrieval algorithm that  
3 has been applied to the complete time series of OMI and GOME-2 measurements and  
4 delivered for public use on the TEMIS web site. Our focus is the continuity and the  
5 consistency of the H<sub>2</sub>CO dataset, as well as a good characterisation of the satellite  
6 observations. The spectral fits have been improved by means of a better treatment of the  
7 interference between O<sub>4</sub>, BrO and H<sub>2</sub>CO differential structures, resulting in H<sub>2</sub>CO columns of  
8 higher accuracy and precision. Daily remote radiance spectra are used as DOAS reference,  
9 and a destriping procedure is included in the background sector correction, reducing the  
10 impact of the OMI row anomaly, but also the GOME-2 across-track variability. Daily  
11 morning and afternoon *a priori* profiles are provided by a state-of-the-art version of the  
12 IMAGES global CTM.

13 The GOME-2 and OMI H<sub>2</sub>CO datasets agree very well qualitatively, both in terms of long-  
14 term variations and seasonal variations. Vertical columns also agree reasonably well although  
15 systematic differences are observed depending on the geographical location. The morning  
16 H<sub>2</sub>CO observations are higher than the afternoon observations over tropical rainforests of  
17 Amazon basin, Africa, and Indonesia. The OMI observations are larger than the GOME-2  
18 columns over urban areas (pointing to a horizontal resolution effect) and more generally over  
19 all mid-latitude regions (pointing to a combination of actual diurnal variation effects and  
20 differences in retrieval sensitivities between morning and afternoon observations).

21 A detailed validation study has been performed using correlative ground-based MAX-DOAS  
22 measurements in Belgium/Netherlands, South of France, North Eastern China, Burundi and  
23 FTIR measurements in Reunion Island. We show that the differences observed between the  
24 GOME-2 and OMI H<sub>2</sub>CO columns are mainly consistent with the diurnal variations observed  
25 from the ground, within the error bars of the satellite and ground-based observations. In  
26 Beijing/Xianghe and Bujumbura, MAX-DOAS vertical profiles have been used to re-calculate  
27 the satellite air mass factors, allowing to eliminate from the comparison the error coming from  
28 the *a priori* profiles. By doing so, the satellite and MAX-DOAS columns are found to agree to  
29 within 15% or better.

30 To conclude, while the precision is driven by the signal to noise ratio of the recorded spectra,  
31 the accuracy is limited by our current knowledge of the external parameters needed for the  
32 retrieval, mainly the *a priori* profile shapes and their diurnal variation and the cloud and

1 aerosol properties. To fully exploit the potential of satellite data, scientific studies relying on  
2 tropospheric H<sub>2</sub>CO observations require consistently retrieved long-term time series, provided  
3 with well characterised errors and averaging kernels. In the framework of prototype algorithm  
4 developments for the future TROPOMI instrument to be flown on the ESA Copernicus  
5 Sentinel-5 Precursor mission, we are currently investigating the impact of using global CTM  
6 profiles on a finer horizontal resolution. Verification and validation studies are ongoing with  
7 the aim to further improve the retrieval algorithms. Furthermore, in the context of the EU  
8 QA4ECV project ([www.qa4ecv.eu/](http://www.qa4ecv.eu/)), a H<sub>2</sub>CO Climate Data Record (CDR) using all the  
9 satellite instruments based on a jointly optimised European algorithm is currently under  
10 development.

## 11 **8 Acknowledgements**

12 The H<sub>2</sub>CO data products from GOME-2 were generated at BIRA using level-1 data developed  
13 by EUMETSAT. Level-2 and level-3 H<sub>2</sub>CO scientific products from GOME-2 have been  
14 jointly supported by Belgian PRODEX (A3C and TRACE-S5P), ESA (PROMOTE) and EU  
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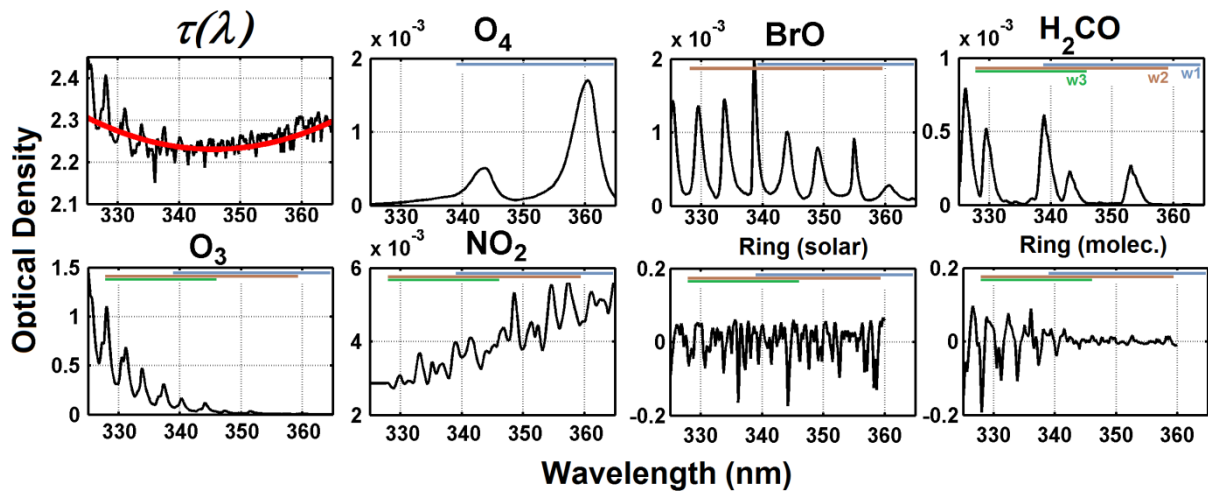


Figure 1: Typical log-ratio of the measured Earth reflectance and fitted low-order polynomial (black and red lines in the first panel), and optical densities of  $O_2-O_2$  ( $O_4$ ), BrO,  $H_2CO$ ,  $O_3$ ,  $NO_2$  and Ring effect (solar lines and molecular) in the near UV. The slant columns have been taken as  $0.4 \times 10^{42} \text{ molec}^2 \cdot \text{cm}^{-5}$  for  $O_4$ ,  $10^{14} \text{ molec} \cdot \text{cm}^{-2}$  for BrO,  $10^{16} \text{ molec} \cdot \text{cm}^{-2}$  for  $H_2CO$ ,  $10^{20} \text{ molec} \cdot \text{cm}^{-2}$  for  $O_3$  and  $1 \times 10^{16} \text{ molec} \cdot \text{cm}^{-2}$  for  $NO_2$ . The different wavelength intervals used in the retrieval are indicated as w1, w2 and w3.



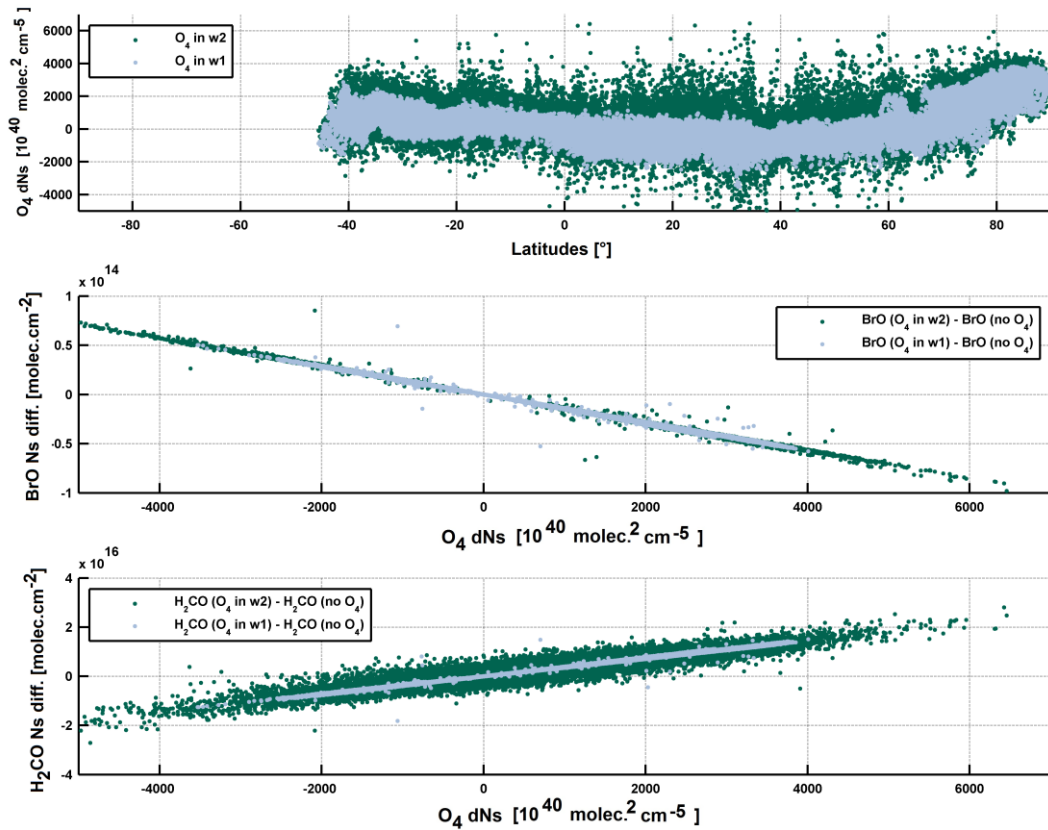


Figure 2: Interdependencies between  $O_4$ , BrO and  $H_2CO$  slant columns shown for one OMI orbit of 01/07/2005 (UT 0548). The first panel shows the differential  $O_4$  slant columns retrieved in w1 (339-364 nm, light blue) and in w2 (328.5-359 nm, dark green) as a function of the latitude. The second and third panels show the differences obtained in BrO and  $H_2CO$  slant columns as a function of the  $O_4$  slant columns, when including or not  $O_4$  in the fits. Differences are shown for two cases:  $O_4$  slant columns retrieved in w1 (light blue) or in w2 (dark green). The limits of the wavelength intervals are indicated in Fig. 1.

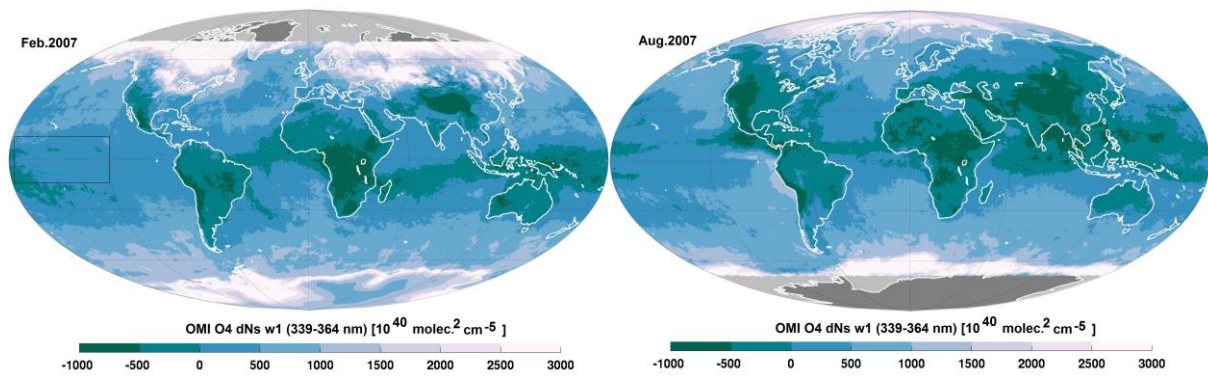


Figure 3: Monthly averaged O<sub>2</sub>-O<sub>2</sub> (O<sub>4</sub>) differential slant columns retrieved in the interval 339-364 nm. The slant columns are differential since radiance spectra over the Equatorial Pacific region (delimited by the black box) are used as reference spectra.

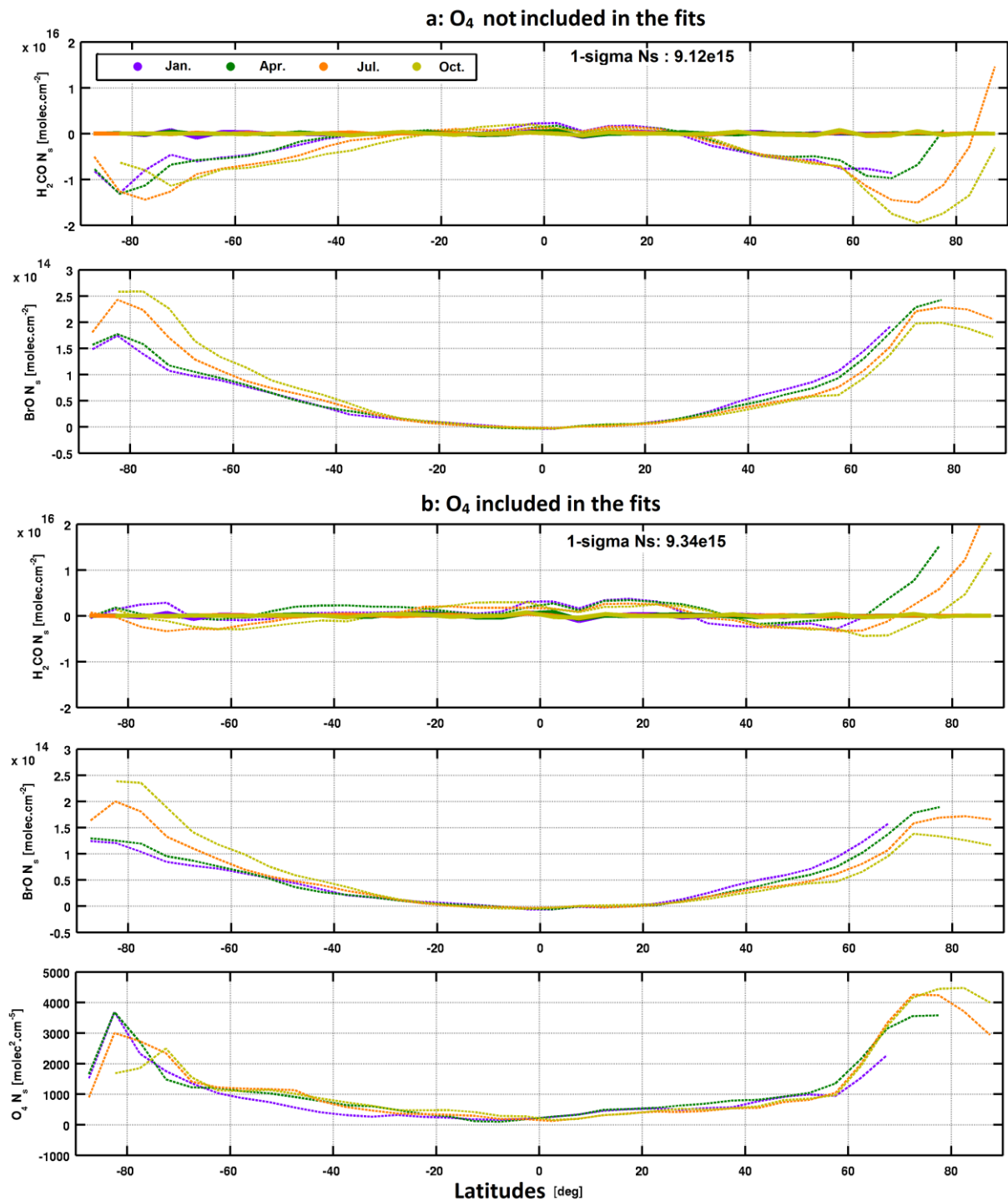


Figure 4: Averaged zonal variation of the H<sub>2</sub>CO, BrO and O<sub>4</sub> differential slant columns for 4 months of OMI observations in 2005 (January, April, July, October), over the Pacific Ocean, as a function of the latitude (a) when O<sub>4</sub> is not included in the fit (v12), (b) when O<sub>4</sub> is included and pre-fitted in w1 (v14). Dashed lines show the slant columns before the background normalisation procedure, while the plain lines show the normalized H<sub>2</sub>CO slant columns.

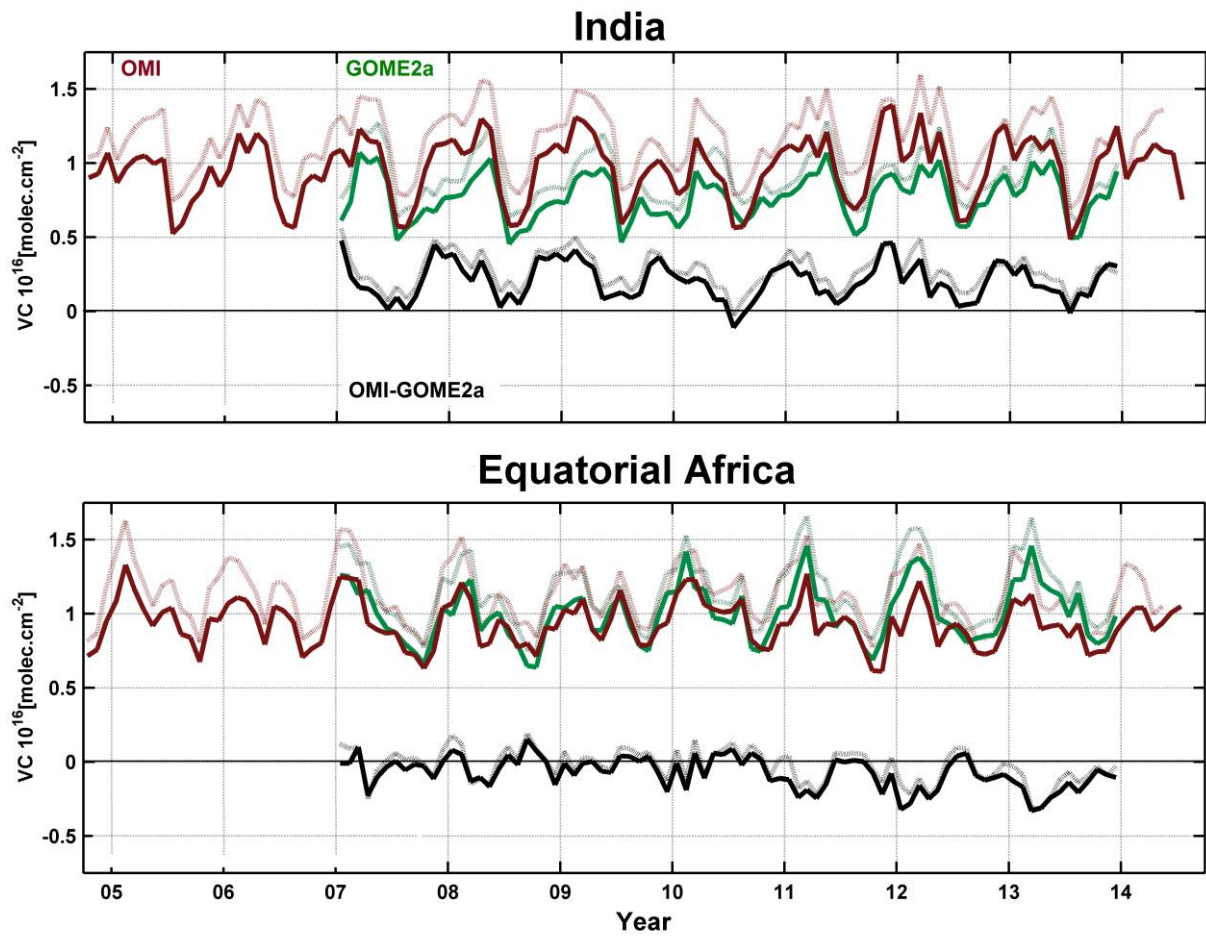


Figure 5: Impact of O<sub>4</sub> on the absolute OMI and GOME-2A H<sub>2</sub>CO columns (respectively in red and green) and on their differences (in black). Dashed lines show results when O<sub>4</sub> is not included in the fits, while solid lines show v14 results.

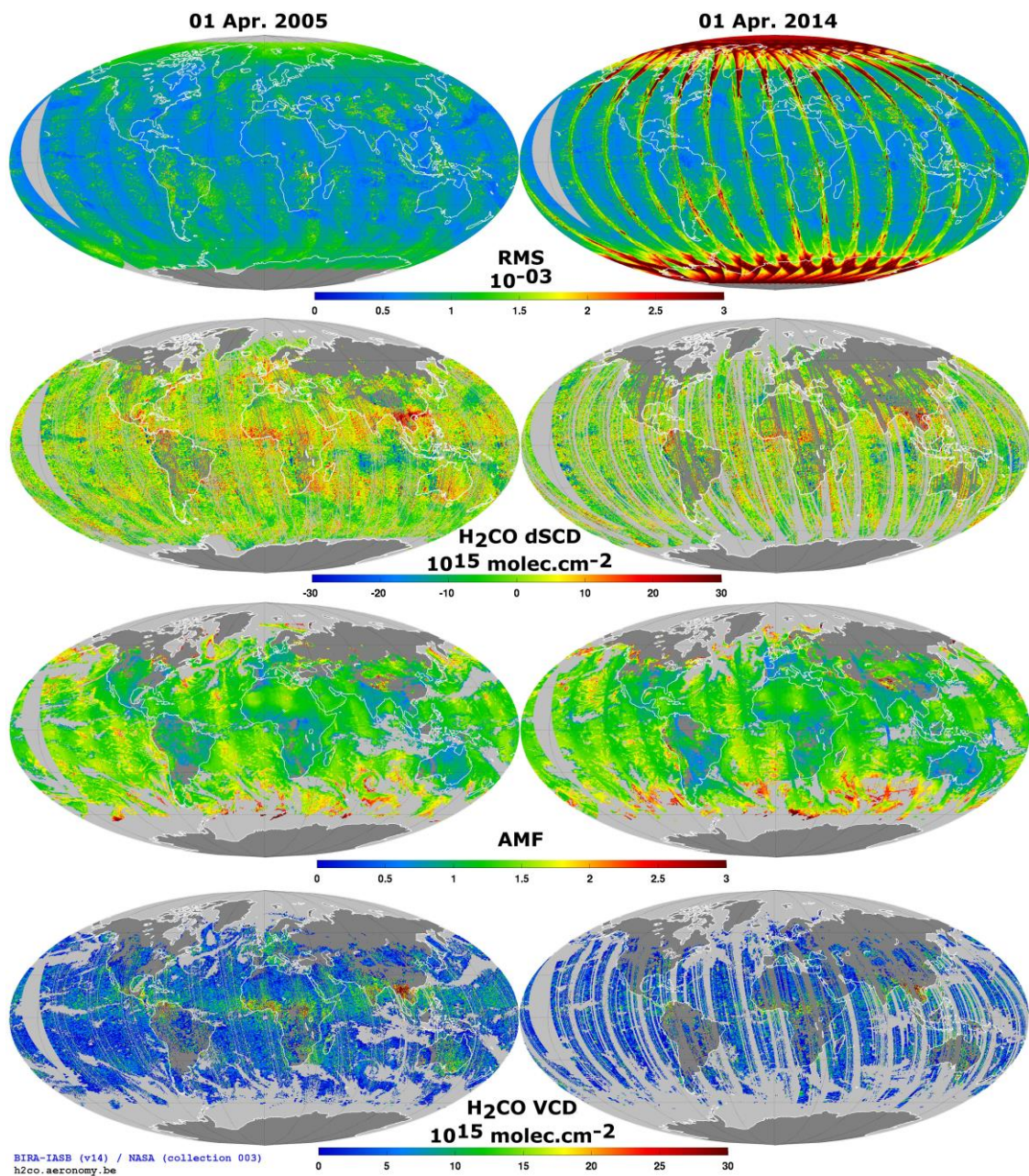


Figure 6: Intermediate retrieval quantities of the H<sub>2</sub>CO retrieval algorithm illustrated with OMI on 01 April 2005 (first column), and 01 April 2014 (second column). The first line shows the fit residuals, while lines 2 to 4 show respectively the H<sub>2</sub>CO reference sector corrected slant columns, the air mass factors and the vertical columns.

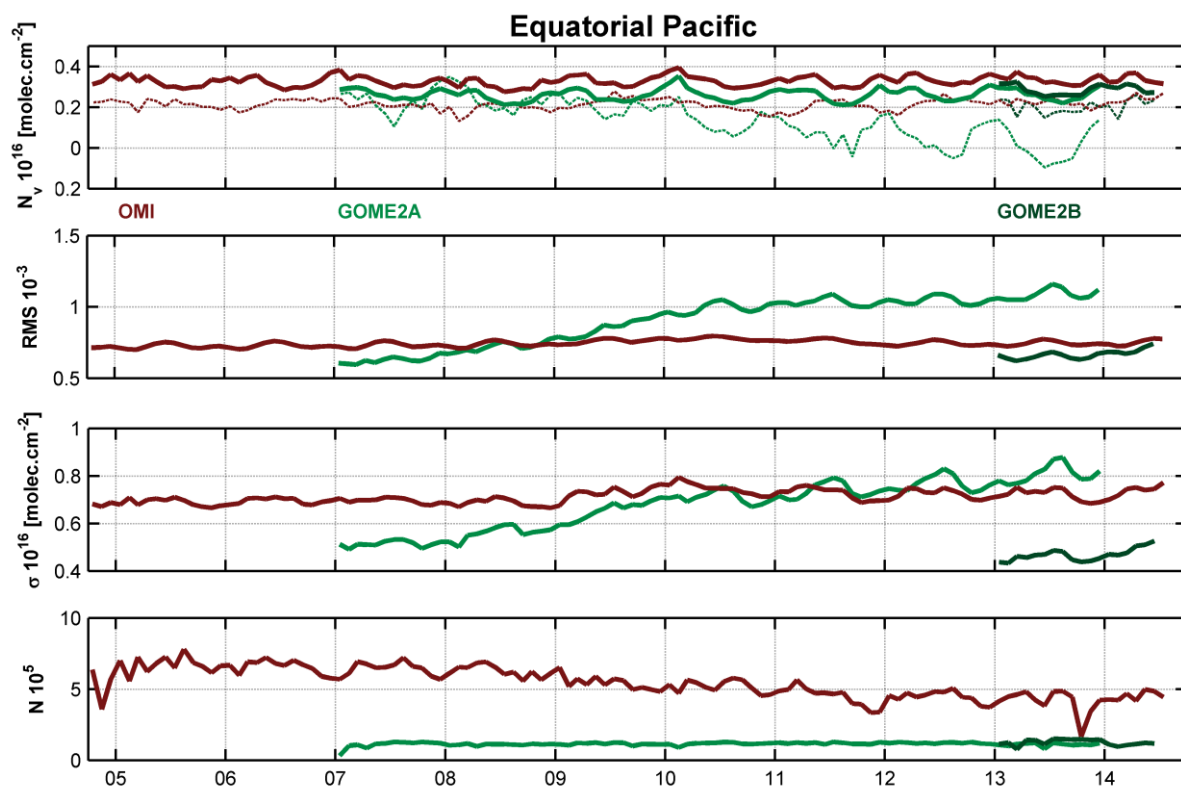


Figure 7: Time series of H<sub>2</sub>CO retrieval statistics for OMI and GOME-2 on METOP A and B in the remote equatorial Pacific region ( $[-15\ 15]^{\circ}$  lat.,  $[180\ 240]^{\circ}$  long.). First panel presents the H<sub>2</sub>CO vertical columns, before and after background correction (dotted and plain lines). The second, third and fourth panels present respectively the fit residuals, the column standard deviation and the number of valid observations. See text for details on the DOAS retrieval settings.

**H<sub>2</sub>CO VC [10<sup>15</sup> molec.cm<sup>-2</sup>]**

**2007-2013**

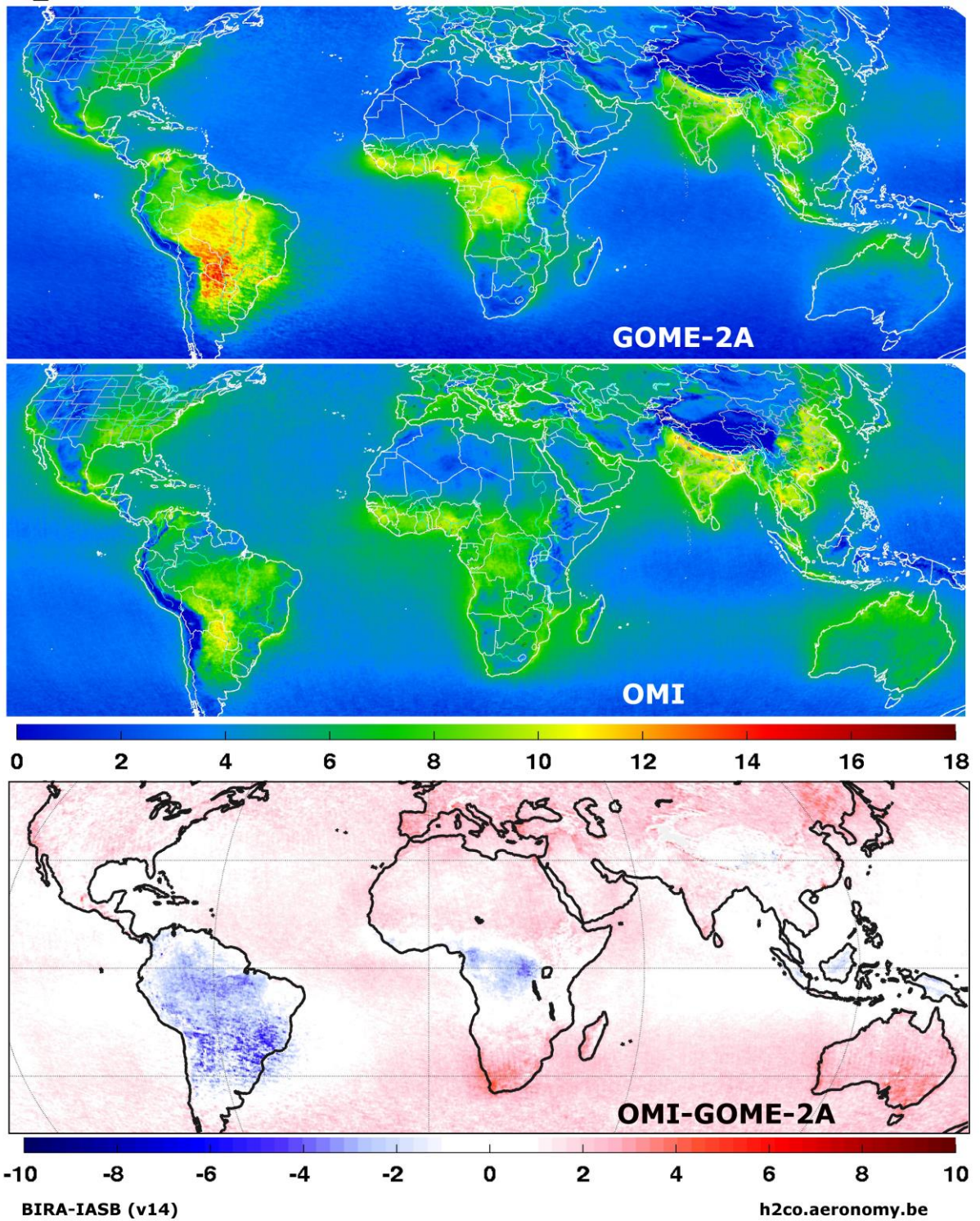


Figure 8: H<sub>2</sub>CO vertical columns retrieved from GOME-2/MetOp-A (first panel, EUMETSAT level 1 data), OMI/Aura (second panel, NASA level 1 data), and their absolute differences (third panel) between 2007 and 2013.

H<sub>2</sub>CO VC [10<sup>15</sup> molec.cm<sup>-2</sup>]

2007-2013

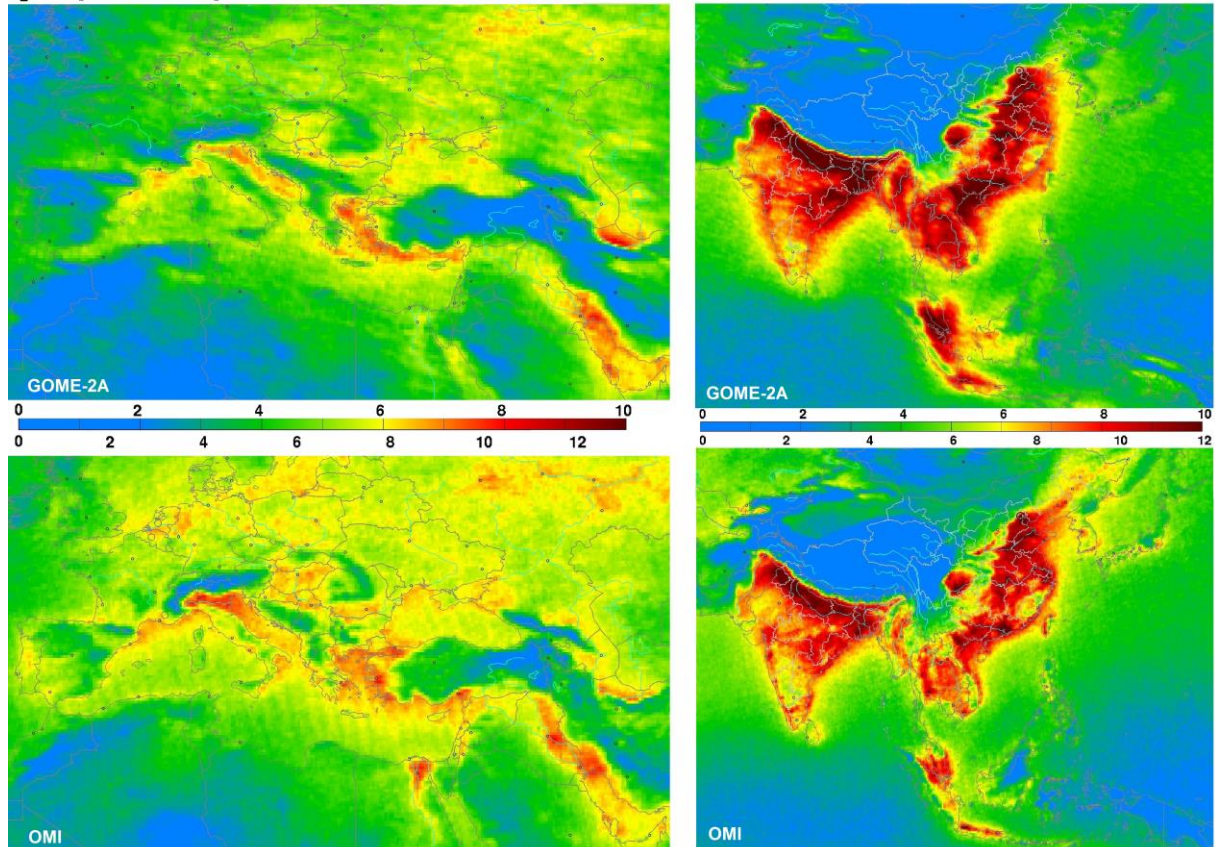


Figure 9: H<sub>2</sub>CO vertical columns retrieved from GOME-2/MetOp-A (first row) and OMI/Aura (second row) over Europe (from April to September) and Asia (from February to November), between 2007 and 2013. Note that different colour scales have been used for GOME-2 and OMI.



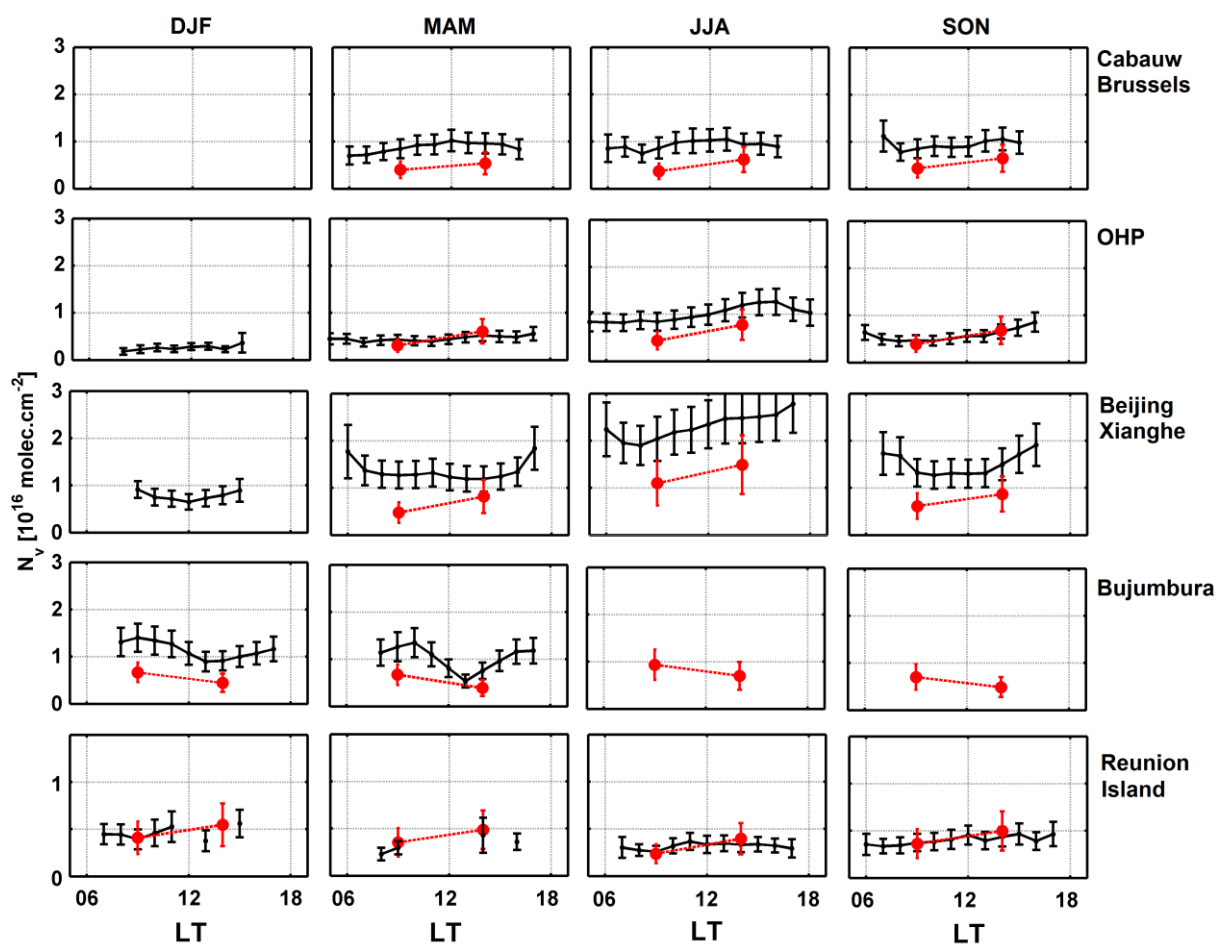
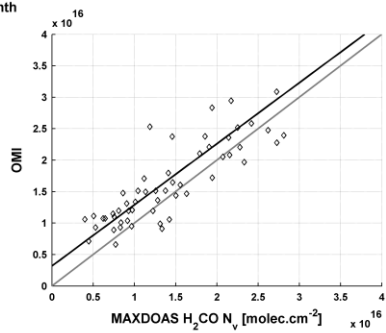
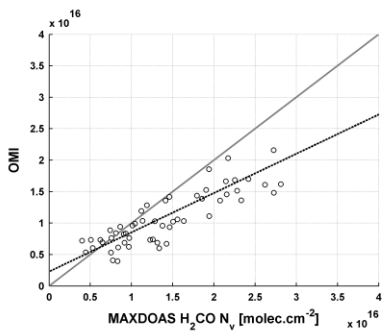
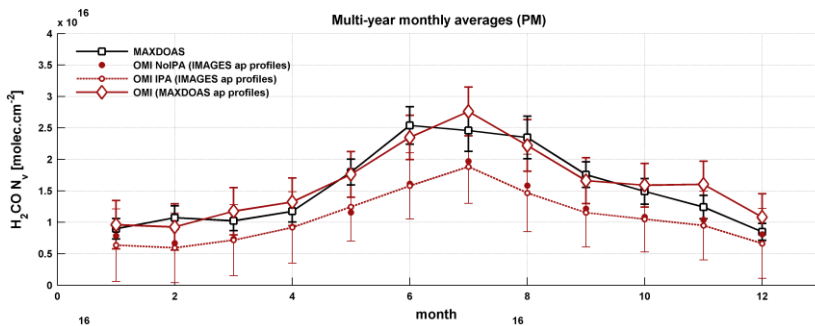
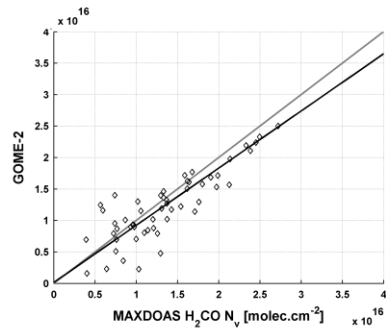
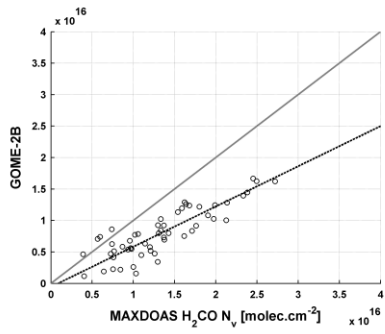
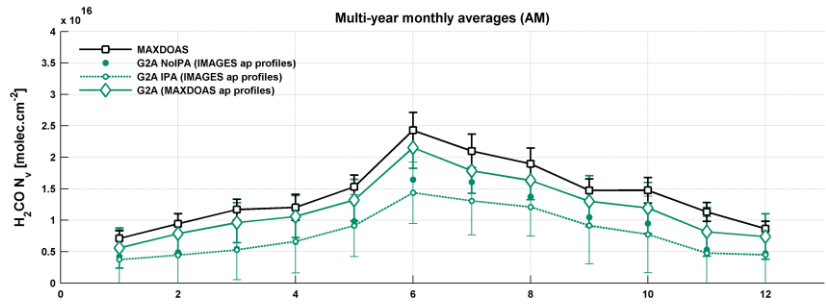


Figure 10: Mean seasonal and diurnal variations of the  $\text{H}_2\text{CO}$  columns as observed by MAX-DOAS or FTIR instruments (in black) and by GOME-2 and OMI (in red). Values are shown for 5 regions where BIRA-IASB operates ground-based measurements: Cabauw and Brussels (Northern Europe), OHP (Southern France), Beijing and Xianghe (North Eastern China), Bujumbura (Central Africa) and Reunion Island (Southern Africa). Satellite measurements have been averaged within 100km around each location, and filtered using quality criteria as defined in Sect. 3.4. Details of the ground-based measurements are summarized in Table 3.



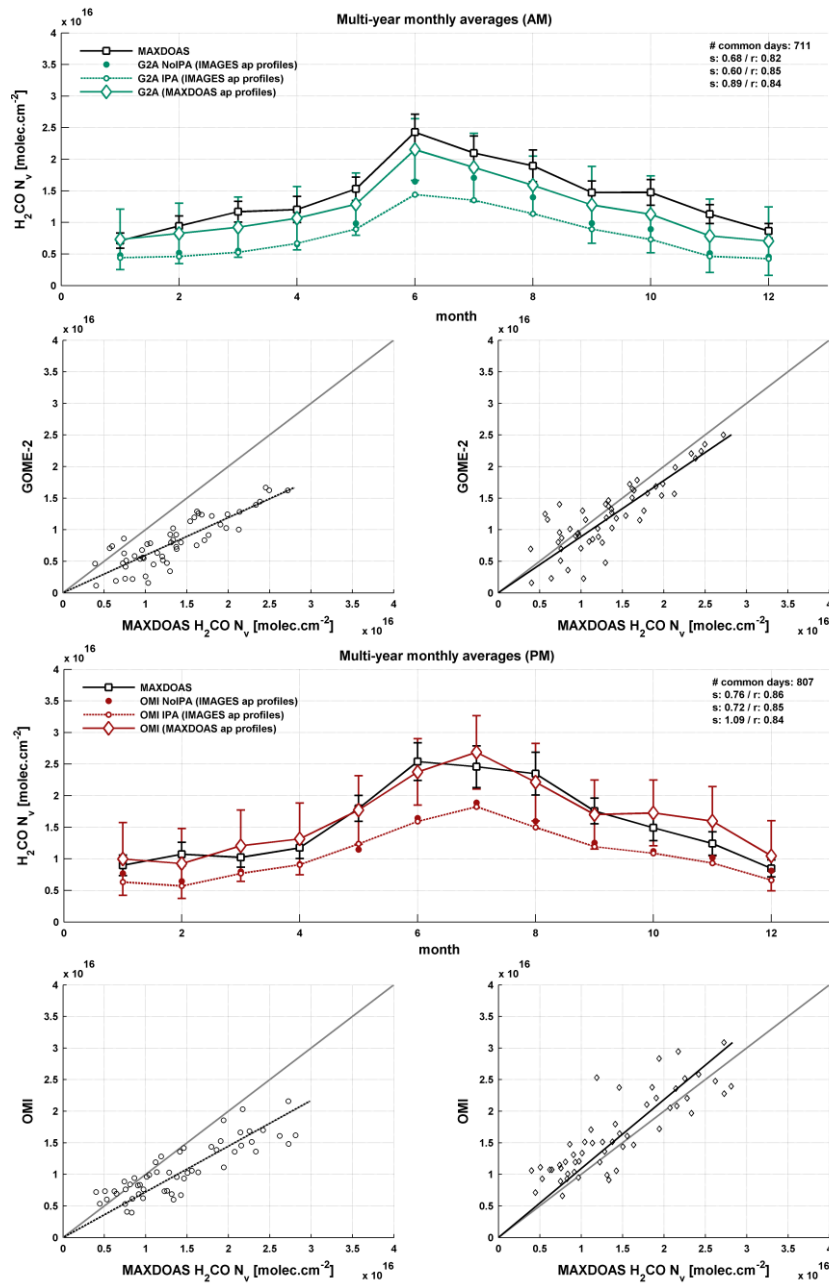


Figure 11: Validation of GOME-2 and OMI retrievals in Beijing and Xianghe, using MAX-DOAS retrievals (represented by black squares). Upper panel: mid-morning observations (GOME-2 and MAX-DOAS averaged over 8-11h). Lower panel: early afternoon observations (OMI and MAX-DOAS averaged over 12-15h). Observations have been averaged per month, over the period 2008-2013, selecting correlative days between GOME-2/OMI and the MAX-DOAS instrument. Satellite measurements have been averaged within 100km around each location, and filtered using quality criteria as defined in Sect. 3.4. Three satellite VCs are presented: IMAGES a.p. profile/no cloud correction, IMAGES a.p. profile/IPA cloud correction, and MAX-DOAS a.p. profile/IPA cloud correction. Correlation plots are shown for the two latter cases, respectively on the left and right panels.

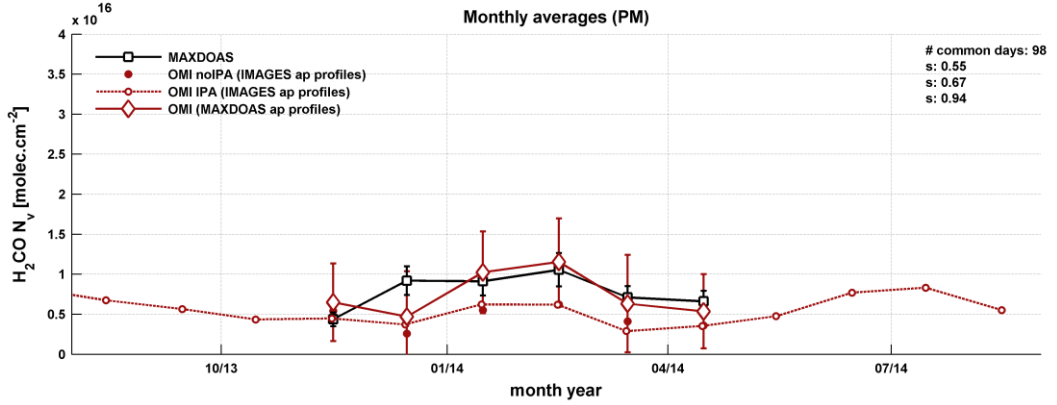
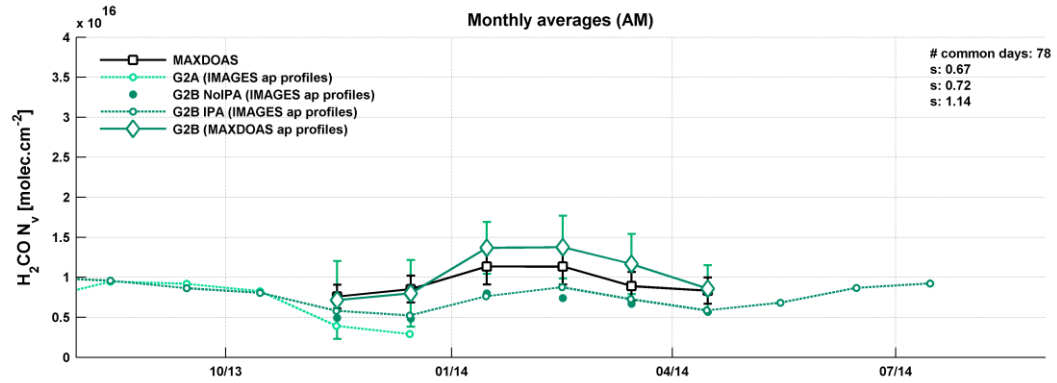
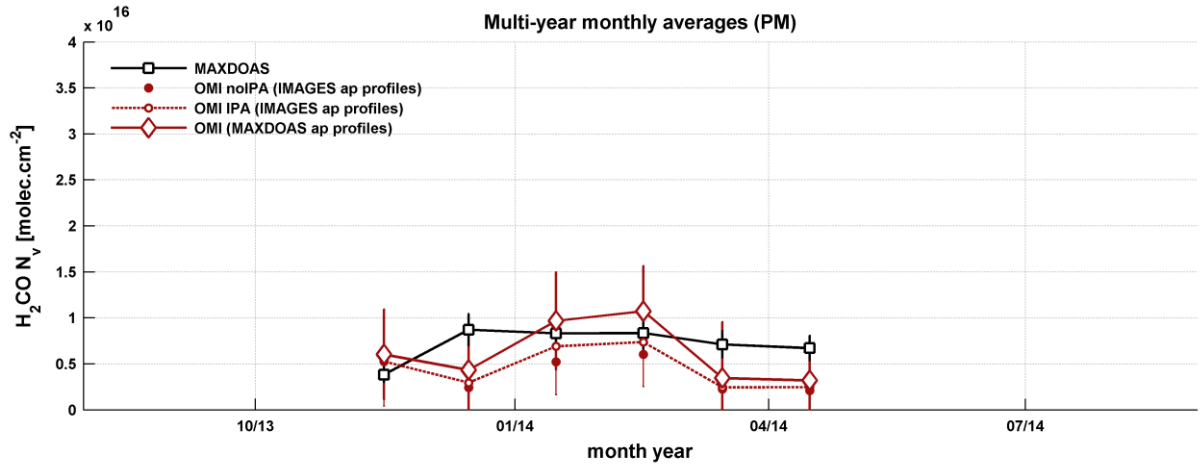
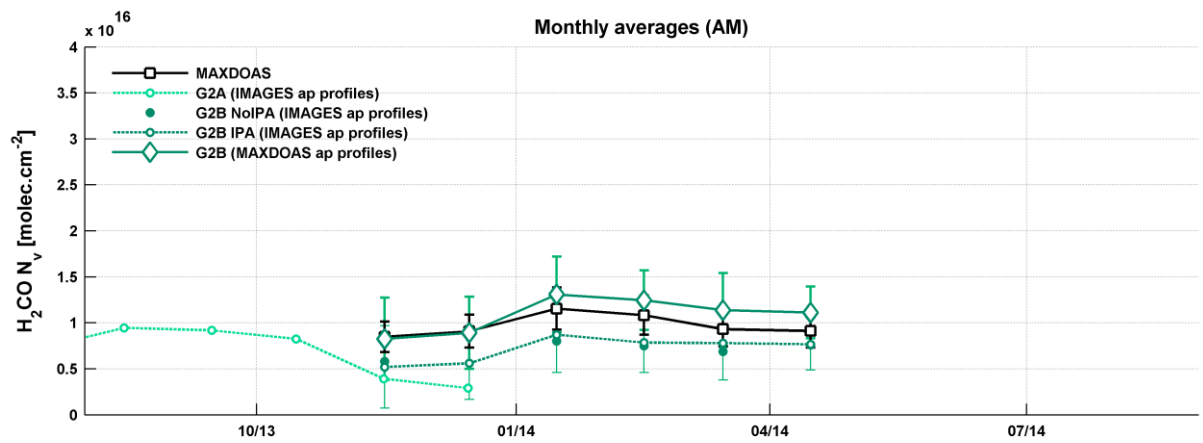


Figure 12: Validation of GOME-2 and OMI retrievals in Bujumbura, using MAX-DOAS retrievals (represented by black squares). Upper panel: mid-morning observations (GOME-2 A&B and MAX-DOAS averaged over 8-11h). Lower panel: early afternoon observations (OMI and MAX-DOAS averaged over 12-15h). Observations have been averaged per month, over the period 2013-2014, selecting correlative days between GOME-2/OMI and the MAX-DOAS instrument. Satellite measurements have been averaged within 100km around each location, and filtered using quality criteria as defined in Sect. 3.4. Three GOME-2B and OMI VCs are presented: IMAGES a.p. profile/no cloud correction, IMAGES a.p. profile/IPA cloud correction, and MAX-DOAS a.p. profile/IPA cloud correction.

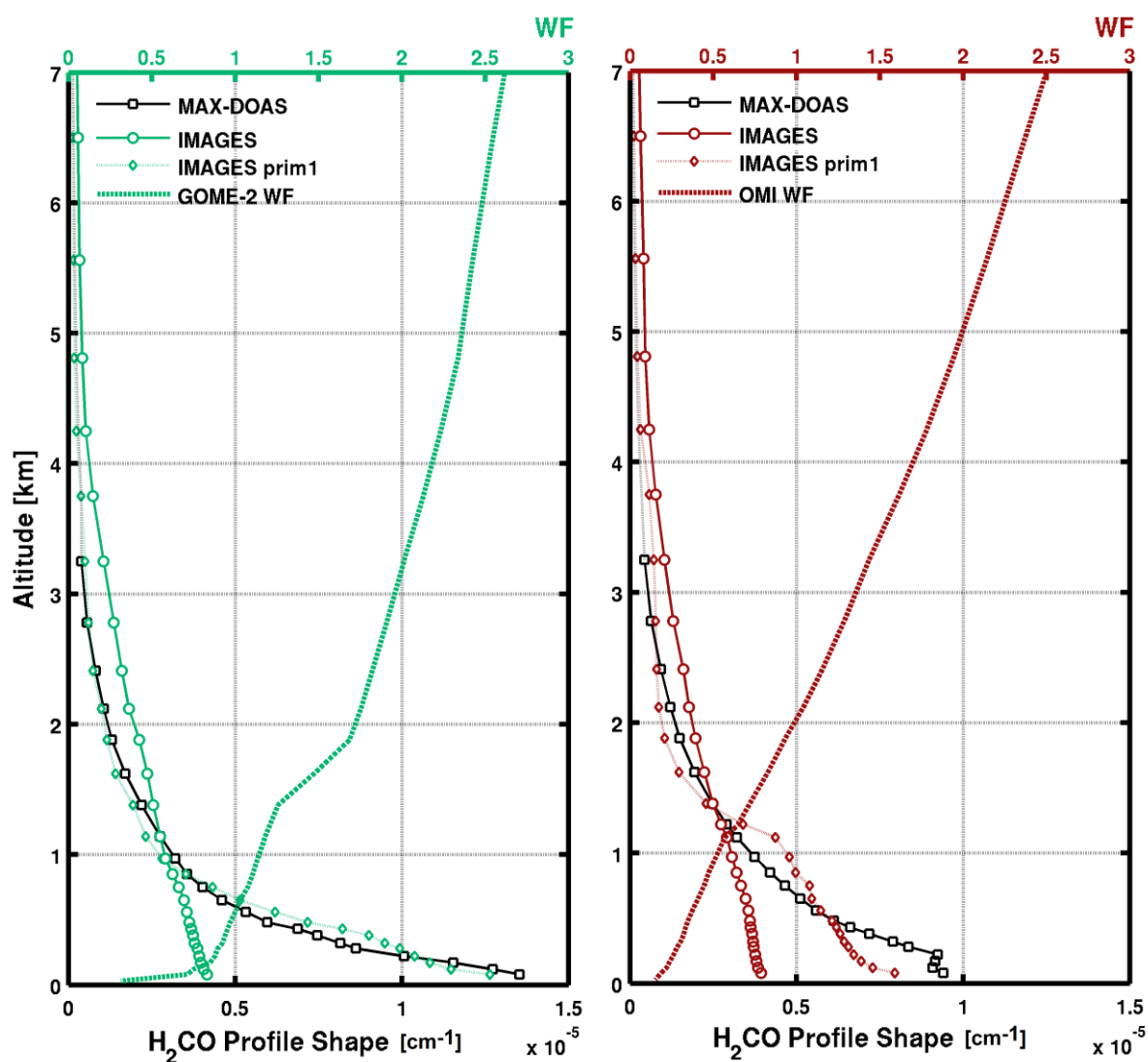


Figure 13: IMAGES CTM and MAX-DOAS retrieved  $H_2CO$  profile shapes (concentration profiles divided by their total columns) averaged in June 2010 in Xianghe. IMAGES prim1

includes an additional primary source, as reported by Chen et al. (2014). The corresponding mean satellite scattering weighting function are also displayed.

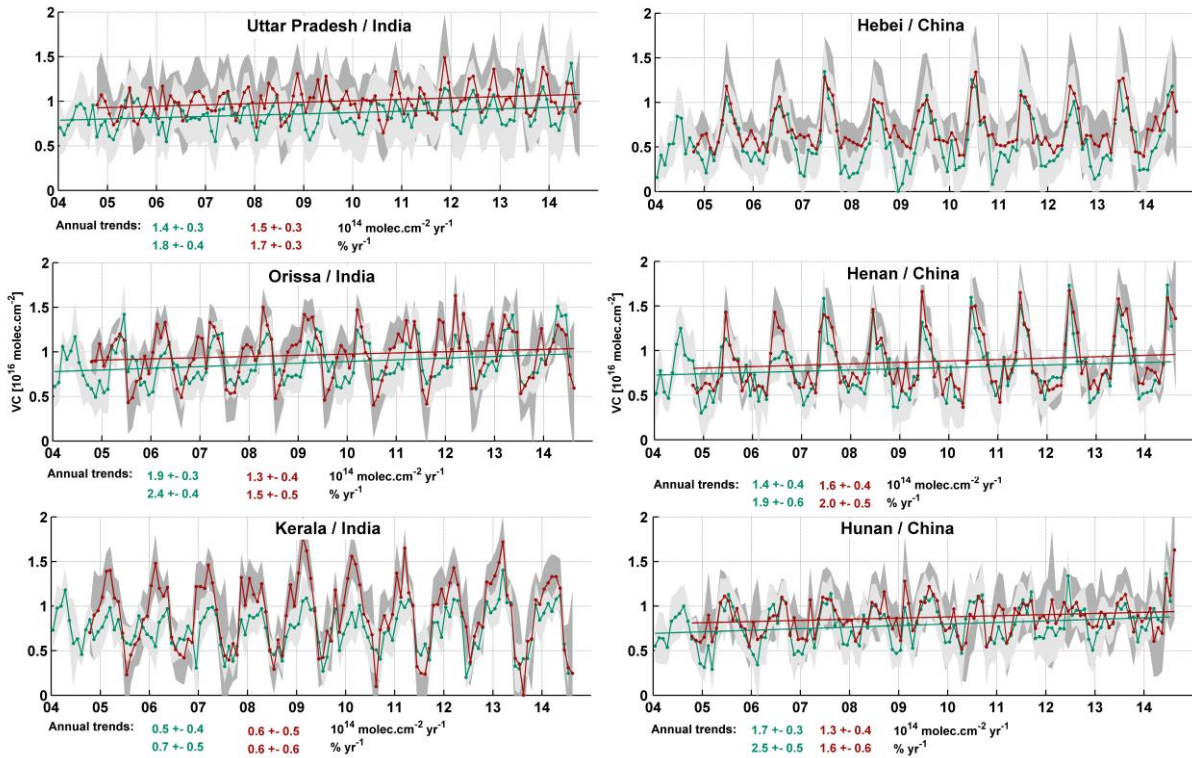


Figure 14: Monthly averaged H<sub>2</sub>CO vertical columns as observed from satellite instruments in India and China. Mid-morning columns (in green) consist in **SCIAMACHY (2003-2011)** and **GOME-2A (2007-2013)** and **B (2013- )** measurements, while early afternoon columns (in red) are derived from OMI measurements. If statistically significant, results of the trend analysis are displayed (De Smedt et al., 2010).

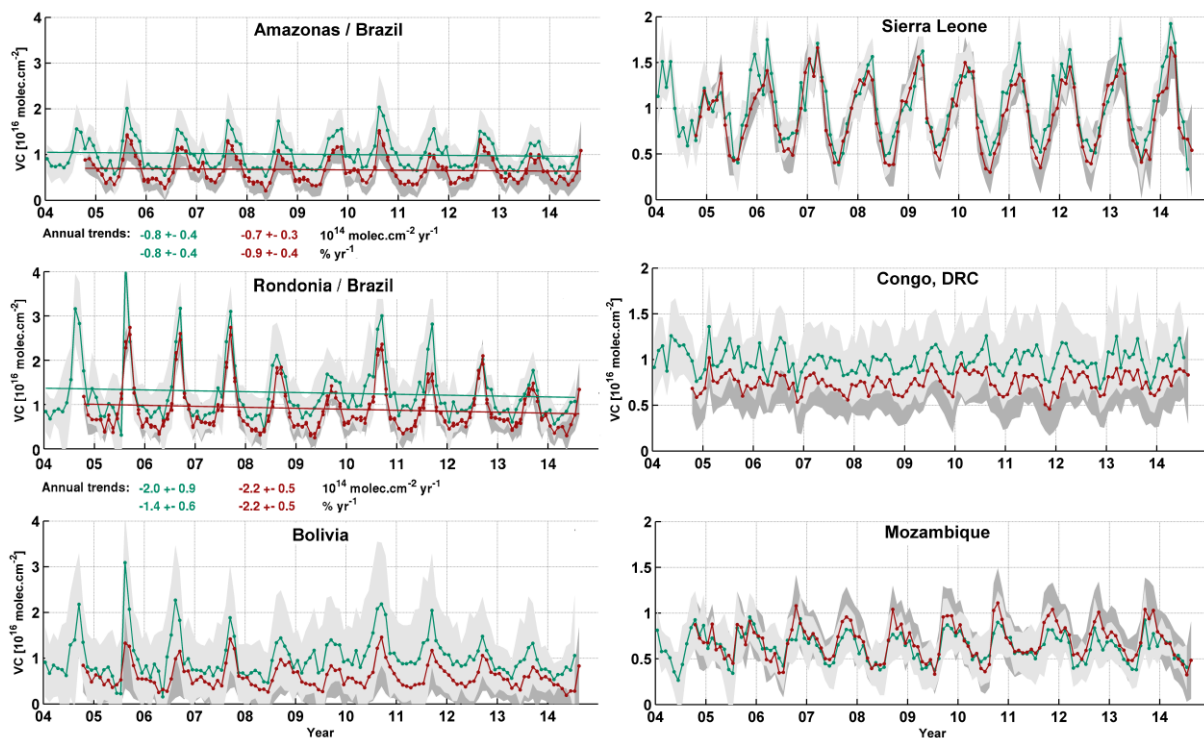


Figure 15: Monthly averaged H<sub>2</sub>CO vertical columns as observed from satellite instruments in South America and Africa. Mid-morning columns (in green) consist in **SCIAMACHY (2003-2011)** and **GOME-2A (2007-2013)** and **B (2013- )** measurements, while early afternoon columns (in red) are derived from OMI measurements. If statistically significant, results of the trend analysis are displayed (De Smedt et al., 2010).

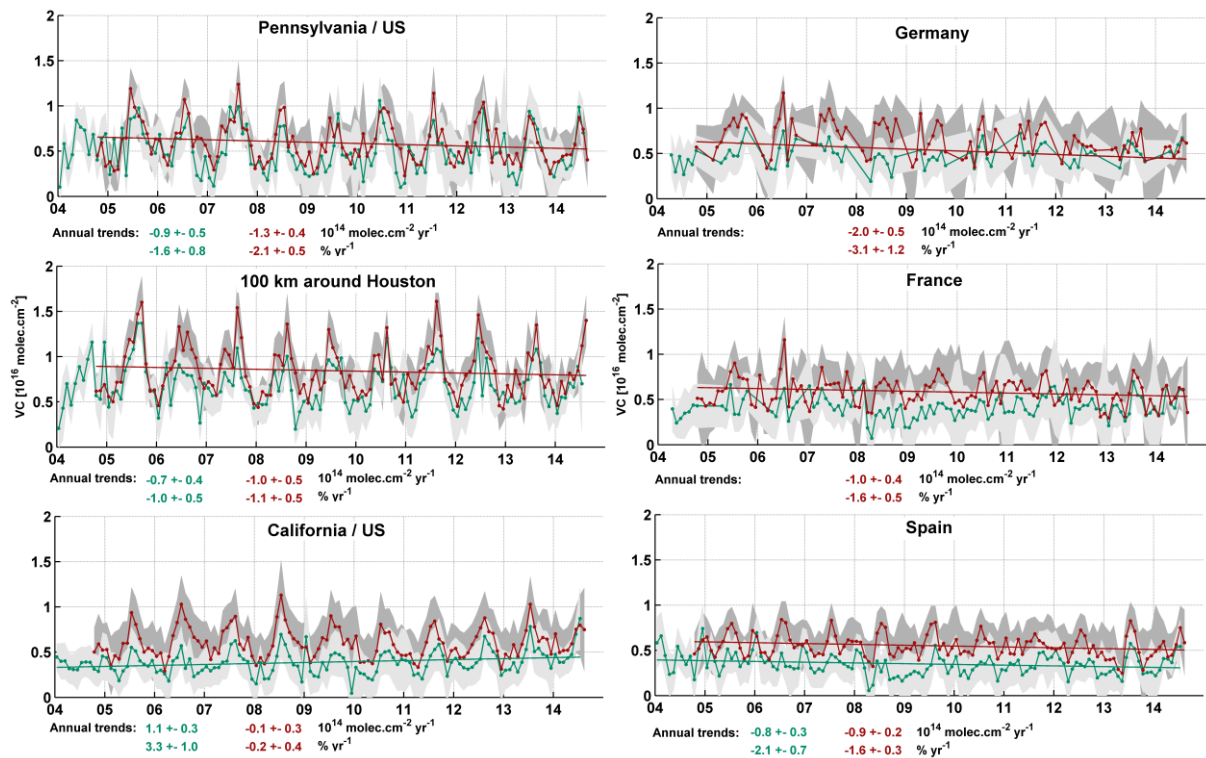


Figure 16: Monthly averaged H<sub>2</sub>CO vertical columns as observed from satellite instruments in US and Europe. Mid-morning columns (in green) consist in **SCIAMACHY (2003-2011) and GOME-2A (2007-2013) and B (2013- ) measurements**, while early afternoon columns (in red) are derived from OMI measurements. If statistically significant, results of the trend analysis are displayed (De Smedt et al., 2010).



H<sub>2</sub>CO Annual Trend [ $10^{14}$  molec.cm<sup>-2</sup>.yr<sup>-1</sup>]: 2004-2014

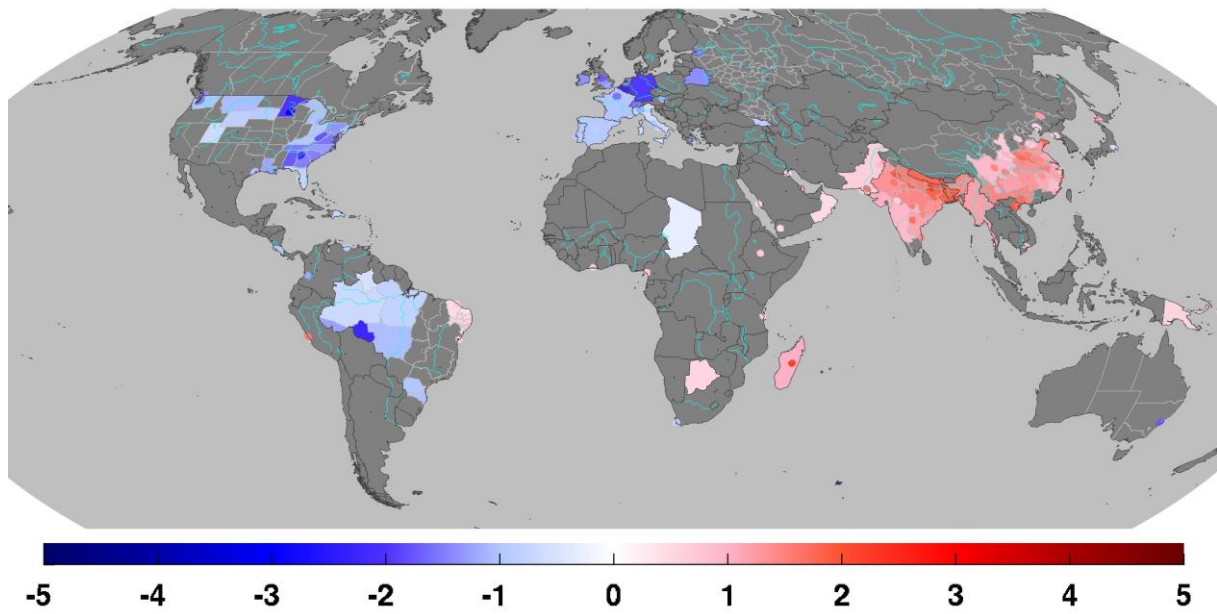


Figure 17: Annual absolute trends observed in the OMI H<sub>2</sub>CO columns between Nov.2004 and Aug.2014. The change in the OMI spatial sampling over the years has been taken into account in this trend analysis.

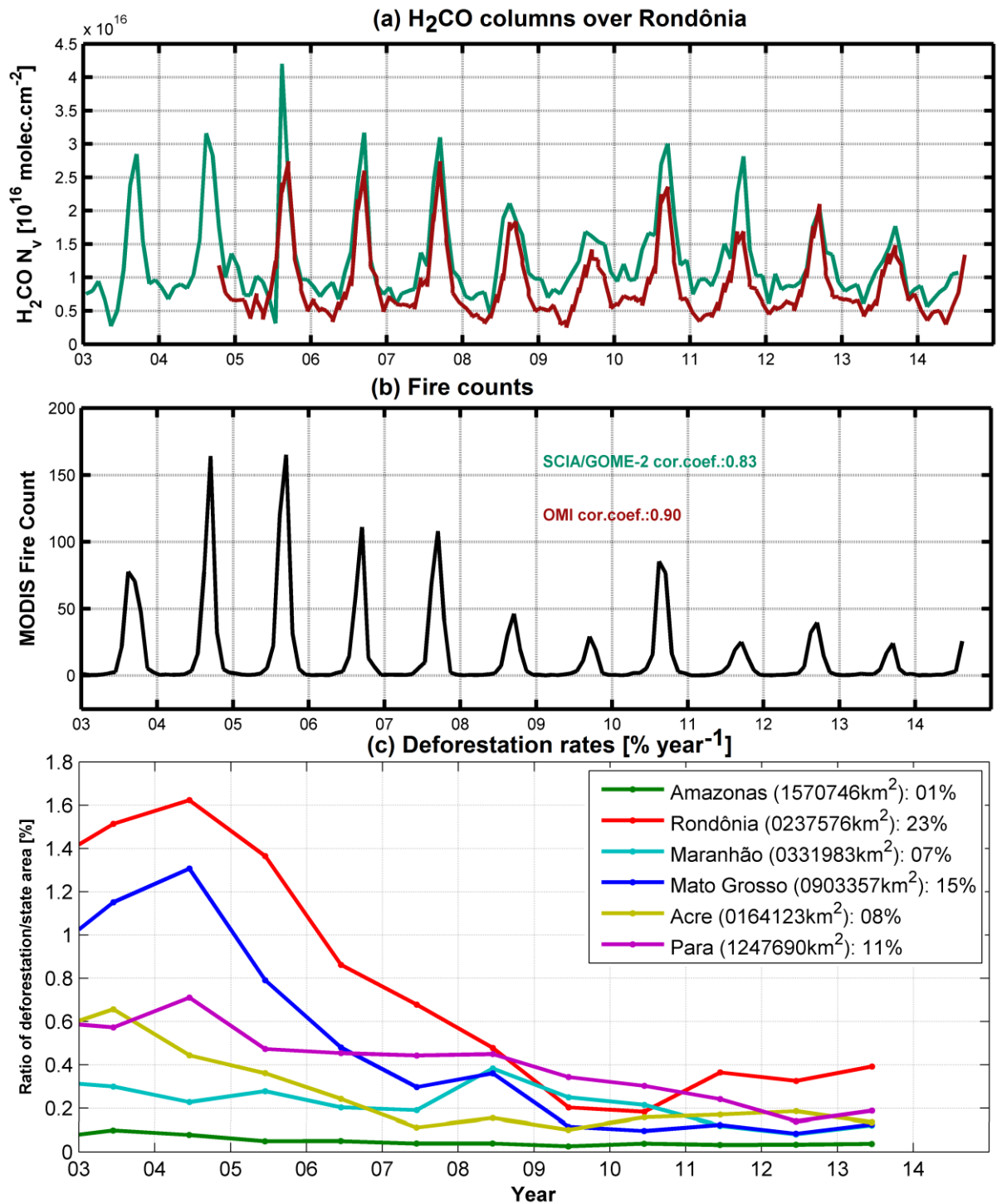


Figure 18: (a) Monthly averaged H<sub>2</sub>CO vertical columns from SCIAMACHY/GOME2 (in green) and from OMI (in red) (first panel) and (b) MODIS fire count (second panel) over the Rondônia Brazilian State. Inset values are the correlation coefficient between the satellite H<sub>2</sub>CO columns and the fire counts. (c) Reported yearly deforestation rates in selected Brazilian States, relative to their respective surfaces (third panel, source: Brazil INPE). Inset values are the total surface of the State, and the total rate of deforestation between 1988 and 2013.