

[Interactive
Comment](#)

Interactive comment on “Modeling dust as component minerals in the Community Atmosphere Model: development of framework and impact on radiative forcing” by R. A. Scanza et al.

R. A. Scanza et al.

ras486@cornell.edu

Received and published: 17 November 2014

Review comments are in black, while responses to the reviewer are in red. When text from the manuscript is quoted, new text is in bold face.

We thank both reviewers for their careful reading of the text, and find that their remarks improve the clarity and usefulness of the manuscript. We revise the text as suggested by the reviewers.

[Full Screen / Esc](#)

[Printer-friendly Version](#)

[Interactive Discussion](#)

[Discussion Paper](#)

Referree 1: Review on the paper entitled: “Modeling dust as component minerals in the Community Atmosphere Model: development of framework and impact on radiative forcing”, by R. A. Scanza et al., for publication in Atmos. Chem. Phys.

General comments:

The study is a first attempt at modeling global mineralogy of dust. It is focused on building the framework required to carry mineral tracers and on synching them with radiative codes. This paper presents important model developments and a complete discussion on what are the impacts on the simulation of radiative forcing. Simulations performed with CAM4 and CAM5 model versions are presented and discussed to point out the improvement and the limits of the model developments. The paper is well-written and I recommend its publication in Atmos. Chem. Phys. Before its publication, I only have minor comments which are reported below. The main comments are related to the simulated vs observed mineral dynamic range, the optical properties used in this study, and the optimizations presented for the different simulations.

(In case the authors have any comments or questions, they can contact me at benoit.laurent@lisa.u-pec.fr).

We thank the reviewer for the very helpful comments, and revise the text to clarify the points addressed by the reviewer.

Specific comments: p.17760: A table gathering optical properties or at least a discussion on the values used would be useful to complete table 3. We modified the supplement to include a spreadsheet with the spectral range of the complex refractive indices for the eight minerals as well as the ‘Zender’ dust blend with and without hematite. We added the following text on p.17760 line 16 (online) or line 237: “The wavelength dependent complex refractive indices for all eight minerals along with the “rest of dust” blend (“Zender,” Table 3) with (Mahowald et al., 2006) and without hematite (this study) are provided in the supplementary material (S2).”

[Interactive
Comment](#)

[Full Screen / Esc](#)

[Printer-friendly Version](#)

[Interactive Discussion](#)

[Discussion Paper](#)

We added the following text in the supplementary material: “Supplementary material S2: Complex refractive indices for minerals. Illite, Kaolinite, and Montmorillonite are from Egan and Hilgeman, 1979 in SW and from Querry, 1987 in LW (<http://www.atm.ox.ac.uk/project/RI/mineral.html>.) Hematite is from <http://www.atm.ox.ac.uk/project/RI/hematite.html> cited as A. H. M. J. Triaud, private communication (2005). Quartz is the average of O-ray and E-ray data from HITRAN96. Calcite and Gypsum are calculated from data presented in Long, L.L., M. R. Querry, R. J. Bell, and R. W. Alexander (1993), Optical properties of calcite and gypsum in crystalline and powdered form in the infrared and far-infrared, *Infrared Physics* 34(2), 191–201 using code provided by, Zender, C. S., personal communication, 2013). Feldspar refractive indices are for SW only from Egan and Hilgeman, 1979.”

p.17762: The authors made a great effort to perform several simulations and to present them as clearly as possible. Nevertheless, CAM4-t and CAM5-t simulations result of several optimizations from the initial versions that make difficult the understanding of the impact of each optimization on the results. This could be explained more in detail. Here we are discussing the changes from the default CAM model, for CAM model users, which are explained in Albani et al., 2014, and we build from these changes. We modified the text to make this more clear on p.17762, line 23(online) We added the following text: “The tuning parameterizations are described in detail in Albani et al., 2014, and were used for both tuned and mineralogy runs in CAM4 and CAM5. The only changes from the default release for CAM we tested explicitly was the particle size distribution at emission (Kok (2011)). “

In table 4, CAM4-m and CAM4-mH seem to correspond, for instance, to the same simulations (FSDBAM, Kok (2011), and table 3) if the prescription of Hematite is not indicated. Please add a column in the table or add a comment in the caption. We modified the caption for Table 4 to clarify the difference between the mineralogy cases with (CAM4-m, CAM5-m) and without (CAM4-mH, CAM5-mH) hematite prescribed from the coarse soil mode on p.17793 (online)

[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)[Discussion Paper](#)

“Table 4: Description of the model simulation used in this study. All cases are eight year simulation with the last six years used for analysis, All cases are run at 1.9x2.5 resolution. FSDBAM indicates CAM4 physics, bulk aerosols, active atmosphere, land and sea ice components, data ocean, slab glacier and GEOS5 meteorology. FC5 indicates CAM5 physics, modal aerosols, stand alone atmosphere with land and sea ice components, data ocean, slab glacier, and GEOS5 meteorology. Default, tuned and tuned plus mineralogy cases are listed in the upper portion of the table and the lower portion of the table designates the simulations part of the sensitivity study section. “The suffix “-m” refers to the prescription of hematite from both fine and coarse soil fractions while the suffix “-mH” refers to hematite prescribed solely from the fine soil fraction.”

p.17766, 17767 and 17768: The authors point out that the models overestimate the dynamic range in mineral mass fraction close to source areas (Fig. 4), and do not capture the dynamic range far away from the dust sources, as it is for instance the case for K/I surface concentrations (Fig. 5) and deposition (Fig. 7 and 8). Kaolinite and Illite being clay minerals, they seem to be distributed in the same aerosol bins at the emission. If these two minerals are transported and deposited in the same way in the model, the simulated K/I ratio should thus stay constant from their emission in desert source areas and during their atmospheric transport and deposition. Does the K/I ratio, generally encountered in the main desert soil from which dust is emitted, show a dynamic range as large as the one observed in the dust surface concentration and dust deposition? If not, this may also partly explain the model difficulty to reproduce the K/I variability observed far away from emission areas. The point of view of the authors on this would be appreciated in their answer and/or added in the paper. This is a good question and to answer this, we added the following text on p.17769 line 1:

“Interestingly, mineral ratios in most of the main desert soils exhibit range of variability within the range of the observations of variability in mineral concentrations. This suggests that in theory, the soil maps we are using could capture the observed ranges in mineral ratios. For example, the variability of the mineral ratio K/I in N. Africa is be-

[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)[Discussion Paper](#)

tween about 0.2 to 5. Since there were more observations in this region accounted for in the mineral maps from Claquin et al., 1999, along with N. Africa accounting for up to 80% of global dust emission, this heterogeneity is promising. However, due to the coarse resolution of the model, the mineral ratios in the simulations do not capture observations of mineral ratios in dust deposition or concentrations near the surface. In addition, the variability over desert regions in Australia is low (between 1 and 2), while in China, nearly all grid boxes of soil mineralogy K/I are around 0.5 which suggests that the assumed soil mineral variabilities are not adequate in these regions. While in the model we include kaolinite and illite with the same assumed size distribution, in reality, kaolinite tends to be in a slightly larger size fraction than illite ($0.5\text{--}4\mu\text{m}$ and $0.1\text{--}1\mu\text{m}$, respectively) (Glaccum and Prospero, 1980). So in the model these values will tend to stay constant as the model advects them downwind, while in reality these should be more fractionation occurring with transport. It is unclear how more resolution of the size fractions of the minerals in the soils would improve our simulations.”

p.17771: A discussion on the imaginary and real parts mentioning the values used should be added. We modified the text to make this more clear (bold language is new) on p.17771 line 12(online) “The real part (scattering) and the imaginary part (absorbing) of the refractive index at 533 nm is larger for tuned dust than for each of the minerals except for the real part in montmorillonite and for hematite (dust($\lambda=533$ nm): 1.515 -i0.00236, illite($\lambda=533$ nm): 1.415 -i0.00103, kaolinite($\lambda=533$ nm): 1.493 -i9.954e-5, montmorillonite($\lambda=533$ nm): 1.529 -i0.00185, hematite($\lambda=533$ nm): 2.967 -i0.7997, rest of dust blend($\lambda=533$ nm): 1.51 -i0.00105).”

p.17777 l.25-p.17778 l.1: The interpretation of the figures is a little bit tricky. Please rephrase to make it more understandable. The figure reference was incorrect. We modify the text to fix this on p.17778 line 1 (online): “(Figure S2d, Figure 13d, Figure 3d, Table 8a)”

p.17778: Uncertainties in mineral optical properties are mentioned. The presentation of the values used could be interesting. This is a good point, we add in the follow-

[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)[Discussion Paper](#)

ing bold text on p.17778, line 20 (online) “We do not have enough data to estimate the uncertainties in the mineral optical properties, although it is clear that the refractive indices for a given mineral can vary due to imperfections or inclusions which may reflect the geographic location of minerals. For example, chemical composition can vary between two samples collected at a single location, and have different refractive indices. (Egan and Hilgeman, 1979). Additionally, two samples of the same mineral from different geographic locations can also have different refractive indices, Egan and Hilgeman, 1979).”

Technical corrections:

p.17753 I.2: Please add (CESM) after “Community Earth System Model” Text is modified on p.17753, line 2 (online) “This study addresses the direct radiative forcing (DRF) of natural mineral aerosols in the Community Earth System Model (CESM).”

p.17755 I.5: Saltation process is mentioned but not the sandblasting process. The text is modified on p. 17755, line 5 (online) “Measurements and theory show that dust aerosols (0.1-50 μm) are primarily emitted through saltation, the bouncing motion of sand-sized ($\sim 100\text{-}200 \mu\text{m}$) particles that disaggregate and emit dust aerosols via sandblasting from the saltating particles (Gillette et al., 1974; Shao et al., 1993; Kok et al., 2012).”

P.17755 Eq.1: It is the normalized mass distribution which is presented. If M_{ij} refers to a mass fraction, a mass term is missing in this equation. We add the following text to clarify this on p.17756, line 2 (online) “The mass fraction in Equation 1 is 0.87 for particle diameters $D=0.1\text{-}10 \mu\text{m}$ with the remaining fraction 0.13 centered around 19 μm . We assume this fraction is insignificant for long range transport (Zender et al., 2003). “

p.17756 I.3: The sum of the mass fractions is equal to 99.9% and not to 100%. Because we only include significant digits out to 2 decimal points, there is the possibility of round off error, which is not a significant source of uncertainty compared to the many

[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)[Discussion Paper](#)

other assumptions in the model.

Table 1: The sum is not equal to 100% for the Zg, Zo and ST soil types. Please check. The sums do add to 100 if another significant digit is added.

p.17756 l.3 and p.17763 l.6: Please use the same writing for the mass fractions. The sum of the mass fractions is not equal to 100%. The text was modified on p.17763, line 6 (online) “For the first sensitivity study, the tuning parameterizations for dust in both CAM4 and CAM5 are kept constant except the new size distribution was replaced with the size distribution in the release version of the model with mass fractions of 3.8, 11.1, 17.2 and 67.8% for bins 1-4 (CAM4-trs) and mass fractions of 3.2 and 96.8% for modes 1 and 3 (CAM5-trs).”

p.17763 l.12: Please write “Cwiertny et al. (2008)” instead of “(Cwiertny et al., 2008)”. The text was modified on p.17763, line 2 (online): “Cwiertny et al., (2008) finds. . .”

p.17766 l.3: Please add “in aerosols” after “The spatial distribution of minerals”. The text was modified on p.17766, line 3 (online): “The spatial distribution of minerals in aerosols in CAM4 and CAM5”

Table 5: Replace “Isreal” with “Israel” in the Location column. This is corrected.

p.17769 l.18: Add a space between “533” and “nm”. The text was modified to include the space between 533 and nm on p17769, line 18 (online) “. . .are simulated for each model at 533 nm and compared..”

p.17771 l.15: imaginary and real parts instead of imaginary and real portions The text was modified on p.17771, line 15 (online) “Hematite has much larger imaginary and real parts “

Please also note the supplement to this comment:

<http://www.atmos-chem-phys-discuss.net/14/C9302/2014/acpd-14-C9302-2014-supplement.pdf>

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

Discussion Paper



Interactive comment on Atmos. Chem. Phys. Discuss., 14, 17749, 2014.

ACPD

14, C9302–C9309, 2014

Interactive
Comment

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

Discussion Paper

C9309

