

1 **Climatology of the free tropospheric humidity: Extension** 2 **into the SEVIRI era, evaluation and exemplary analysis**

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14 **Abstract**

15 A new free tropospheric humidity (FTH) data record is presented. It is based on observations
16 from the Meteosat Visible and Infrared Imager (MVIRI) onboard Meteosat-2 – Meteosat-5,
17 and Meteosat-7 and from the Spinning Enhanced Visible and Infrared Imager (SEVIRI)
18 onboard Meteosat-8 and Meteosat-9 at the water absorption band near 6.3 μm . The dataset is
19 available under clear-sky and low-level cloud conditions. With the extension to SEVIRI
20 observations the data record covers the period 1983-2009 with a spatial resolution of $0.625^\circ \times$
21 0.625° and a temporal resolution of 3 hours.

22 The FTH is the mean relative humidity (RH) in a broad layer in the free troposphere. The
23 relation between the observed brightness temperature (BT) and the FTH is well established.
24 Previous retrievals are refined by taking into account the relative humidity Jacobians in the
25 training process of the statistical retrieval. The temporal coverage is extended into the SEVIRI
26 period, the homogenization of the BT record is improved and the full archive is reprocessed
27 using updated regression coefficients.

1 The FTH estimated from the Meteosat observations is compared to the FTH computed from
2 the RH profiles of the Analysed RadioSoundings Archive (ARSA). An average relative bias
3 of -3.2% and a relative root mean square difference (RMSD) of 16.8% are observed. This
4 relative RMSD agrees with the outcome of an analysis of the total uncertainty of the FTH
5 product. The decadal stability of the FTH data record is $0.5\pm 0.45\%$ per decade.

6 As exemplary applications the inter-annual standard deviation, the differences on decadal
7 scales, and the linear trend in the FTH data record and in the frequency of occurrence of
8 $FTH < 10\%$ (FTHp10) are analyzed per season. Inter-annual standard deviation maxima as
9 well as maxima in absolute decadal differences are featured in gradient areas between dry and
10 wet regions as well as in areas where FTH reaches minima and FTHp10 reaches maxima. An
11 analysis of the FTH linear trends and of the associated uncertainty estimates is achieved to
12 identify possible problems with the data record. Positive trends in FTHp10 are featured in
13 gradient areas between wet and dry regions, in regions where the FTH is minimum, in regions
14 where the FTHp10 is maximum, and in regions where differences between FTHp10 averaged
15 over the 2000s and over the 1990s are negative. However, these positive trends in FTHp10
16 are associated with maximum standard deviation and they are thus hardly significant. This
17 analysis and inter-comparisons with other humidity data records are part of the Global Energy
18 and Water Cycle Experiment (GEWEX) Water Vapor Assessment (G-VAP).

19

20 **1 Introduction**

21 The importance of relative humidity (RH) in the free troposphere originates from the non-
22 linear interaction between RH and longwave radiation. In order to realistically assess the
23 impact of RH in the free troposphere on longwave radiation, the full probability distribution
24 of RH needs to be considered. However, the outgoing longwave radiation (OLR) is much
25 more sensitive to perturbations at the dry end than at the moist end of the distribution
26 (Spencer and Braswell, 1997; Roca et al., 2011). The magnitude of the differential impact of a
27 given change at the dry end of the RH distribution compared to at the moist end of it can
28 exceed a factor of 3 (Roca et al., 2011).

29 Most of the knowledge on the impact of climate change on RH in the free troposphere arises
30 from theoretical and from global average considerations (Pierrehumbert et al., 2007;
31 Sherwood et al., 2010b; Shi and Bates, 2011; Roca et al., 2011 and references therein). On a
32 global scale, the assumption of a constant upper tropospheric RH under climate change

1 conditions is supported by the work of Soden and Held (2000) and Soden et al. (2002) among
2 others. However, the water vapor pressure is expected to be below equilibrium in most of the
3 atmosphere. This is true in particular for the dry regions in the free troposphere. On the zonal
4 mean scale, a common pattern emerges for a doubling of CO₂: subtropical dry regions shift
5 poleward (Sherwood et al., 2010a), dry minima become drier (Hurley and Galewsky, 2010)
6 and the width of these dry regions increases. Such impacts on the frequency of occurrence of
7 the dry end of the FTH range (below 10%, noted FTHp10 in the following) have been
8 simulated with the Institut Pierre Simon Laplace climate model based on the analysis of the
9 FTH distribution under a climate change scenario (Roca et al., 2011). On a regional scale, it is
10 challenging to assess whether any of these large-scale features are reproduced and what is the
11 expected change in their pattern if they are not reproduced. Sherwood et al. (2010a) reviewed
12 the processes that determine the RH distribution in the intertropical region. A strong
13 connection between large-scale dynamics and water vapor was found and the role of eddies
14 (e.g. mesoscale convection, circulation transients) in establishing these links highlighted the
15 broad range of scales implied in the humidity distribution. The authors further indicated the
16 need for the available theory to be better constrained.

17 These modifications of the spatial distribution of FTH correspond to the expected climate
18 responses either at equilibrium after a CO₂ doubling or at the end of the 21st century. It is
19 unclear how these modifications have been at play over the last 25-30 years and if they can
20 already be identified in the observational record of the last 25-30 years. Satellites observing
21 the humidity in the free troposphere, and in particular geostationary platforms, appear to be
22 very well suited to contribute to constraining the available theory by providing observations at
23 kilometer- and hour-scale resolutions over a 30 year-long period. One of the aims of the
24 present paper is to provide a well-qualified data record to explore such possibility at the
25 temporal and spatial scales of the Meteosat observations.

26 Infrared imagers and infrared sounders, which channels are centered at the water vapor
27 absorption band around 6.3 μm , as well as microwave sounders, which channels are centered
28 at 183.31 GHz, allow for the sounding of the free troposphere. Soden and Bretherton (1993)
29 developed the commonly applied relationship between observed brightness temperature (BT)
30 and free tropospheric humidity (FTH) for infrared observations of cloud-free scenes. This
31 proportionality is relative to a vertical average of the RH in the free troposphere and relies on
32 the assumptions of random strong line theory and constant lapse rate. The regression

1 coefficients depend on the averaging kernel and the performance of the regression can be
2 improved by considering the so-called scaled reference pressure. Roca et al. (2003)
3 statistically determined the regression coefficients and used reanalyses to determine the scaled
4 reference pressure, whereas Schmetz et al. (1995) used reanalyses to determine regression
5 coefficients on a pixel basis. Both retrievals were originally designed for Meteorological
6 Satellite (Meteosat) Visible and Infrared Imager (MVIRI) observations. Buehler and John
7 (2005) adapted the Soden and Bretherton (1996) method to intercalibrated Advanced
8 Microwave Sounding Unit-B (AMSU-B) observations (John et al., 2013). The AMSU-B-
9 based FTH data record (Buehler et al., 2008) is affected by orbital drift and it contains valid
10 observations under all sky conditions except when intense scattering occurs, such as during
11 precipitation events. The method described in Soden and Bretherton (1996) has recently been
12 applied to three of the six channels of the Sondeur Atmosphérique du Profil d'Humidité
13 Intertropicale par Radiometrie onboard the Megha-Tropiques satellite (Brogniez et al., 2014)
14 launched in October 2011. Shi and Bates (2011) spent significant effort on recalibrating and
15 inter-calibrating the water vapor observations at $6.7 \mu\text{m}$ of the High-resolution Infrared
16 Radiation Sounder (HIRS). A subsequent application of the Soden and Bretherton (1996)
17 method yields to a FTH record that is longer than 30 years (Shi et al., 2013). The time series
18 is affected by orbital drift and as it is the case for the MVIRI observations, the FTH is
19 retrieved under clear-sky and low-level cloud conditions. Jackson and Bates (2001) assessed
20 FTH algorithms that are applicable to HIRS observations at $6.7 \mu\text{m}$. They concluded that the
21 averaging kernel has a significant effect on the FTH retrieval and that the results are improved
22 in the tropics when considering the scaled reference pressure in the retrieval (see Soden and
23 Bretherton, 1996).

24 Brogniez et al. (2009) assessed the quality of the MVIRI-based FTH data record for the
25 period 1984-2005. It is a precursor of the FTH data record presented here, and it was
26 compared to FTH values which have been computed from RH profiles of a previous version
27 of the Analysed RadioSoundings Archive (ARSA). The mean difference between the MVIRI
28 FTH and the ARSA FTH over the period 1984-2005 is -1.2 \%RH and the standard deviation
29 of the difference is 1.7 \%RH , indicating the stability of the MVIRI archive over this period.
30 The value of the FTH product has been demonstrated e.g. in Brogniez et al. (2005) who
31 compared the FTH product against 14 climate models in the framework of the Atmospheric
32 Model Intercomparison Project phase 2. Buehler and John (2005) estimated the theoretical
33 uncertainty of the AMSU-B-based FTH retrieval scheme. They found a bias of 2 \%RH for

1 low values of FTH and 7 %RH for high values of FTH. John et al. (2011) further found a
2 systematic clear-sky bias of 9 %RH by sampling the all sky AMSU-B product with the HIRS
3 clear-sky mask and by comparing the clear-sky and the all sky AMSU-B FTH products. An
4 exemplary application of the AMSU-B FTH data record is given in Moradi et al. (2010). They
5 found that the spatial distribution of radiosonde types coincide with the difference between
6 AMSU-B data and data from the Integrated Global Radiosonde Archive (Durre et al., 2006).
7 The HIRS data record has been evaluated in radiance space (see Shi and Bates, 2011) and the
8 correlation of the BT with various climate indices has been analyzed (Shi et al., 2013),
9 highlighting the valuable contribution of this data record to the analysis of global
10 teleconnections.

11 The generation of FTH data records from Meteosat, AMSU-B and HIRS was part of a pilot
12 project within the World Meteorological Organization (WMO) Sustained, Coordinated
13 Processing of Environmental Satellite Data for Climate Monitoring (SCOPE-CM), which
14 aims at establishing a network of facilities ensuring continuous and sustained provision of
15 high-quality satellite products. Within new SCOPE-CM projects, the Meteosat, AMSU-B and
16 HIRS data records will be homogenized in format, metadata and documentation as much as
17 possible and the observations from all available geostationary imagers will be recalibrated and
18 intercalibrated. The impact of this Fundamental Climate Data Record (FCDR) on the
19 Meteosat-based FTH product will also be evaluated within the SCOPE-CM projects.

20 The GEWEX Data and Assessments Panel has initiated the GEWEX Water Vapor
21 Assessment (G-VAP). One element of G-VAP is the intercomparison of available long-term
22 FTH data records and the analysis and the comparison of temporal changes in these data
23 records. The present work supports the latter G-VAP activity by setting-up the technical
24 framework and by starting the analysis with the Meteosat-based FTH data record.

25 The objectives of this paper are 1) to describe the data and the algorithms used to produce the
26 extended long-term stable FTH data record from the Meteosat observations, 2) to characterize
27 the quality and the stability of this observational record by comparing it with independent
28 data, and 3) to confront the variability of FTH on various temporal and spatial scales with the
29 linear trends and their significance. The input data and the homogenization of Meteosat
30 observations are introduced in Sect. 2 and Sect. 3, respectively. The retrieval scheme, an
31 overview of the technical specifications and exemplary figures on the characteristics of the
32 FTH data record are presented in the next sections. The theoretically expected uncertainties

1 are discussed in Sect. 5 and the evaluation results are shown in Sect. 6. The applications
2 presented in Sect. 7 feature an analysis of the variability of the FTH and FTHp10, of the
3 FTHp10 differences on decadal scales, and of the linear trend in FTH and FTHp10.
4 Conclusions are finally drawn.

5 A series of statistical quantities and a list of abbreviations are given in the Appendices.

6

7 **2 Input data**

8 This section briefly describes the instruments and the radiance input data records used to
9 retrieve the FTH.

10 MVIRI is a three channel imaging radiometer flown on Meteosat-2, -3, -4, -5, -6, and
11 Meteosat-7, which belong to the first generation of Meteosat satellites. It continuously
12 observed the Earth from a geostationary orbit at 0° latitude / 0° longitude every 30 minutes
13 between 1982 and 2006. The spatial sampling distance of the observations is approximately 5
14 km at nadir and it increases with distance from the sub-satellite point.

15 The Spinning Enhanced Visible and Infrared Imager (SEVIRI) performs observations at 12
16 channels that cover the visible and thermal infrared spectral range. SEVIRI is on board the
17 geostationary satellites Meteosat-8, -9 and Meteosat-10, which are positioned at 0° latitude /
18 0° longitude in operational mode. The SEVIRI full disc observations are repeated every 15
19 minutes between 2004 and the present day. The spatial sampling distance is 3 km, increasing
20 with distance from the sub-satellite point (Schmetz et al., 2002).

21 The elaboration of the Meteosat clear-sky radiance (CSR) archive is described in Brogniez et
22 al. (2006) and it is briefly recalled in Sect. 3. The 6.3 μm BTs, as observed by Meteosat-2 to
23 Meteosat-5 as well as by Meteosat-7, were taken from the International Satellite Cloud
24 Climatology Project (ISCCP, Rossow and Schiffer, 1999) at the DX pixel resolution (ISCCP-
25 DX, see <http://isccp.giss.nasa.gov/products/products.html> for details). All observations are
26 adapted to the Meteosat-5 spectral response function. The CSR archive built at LMD is used
27 as input. It covers the period July 1983-June 2005. This CSR record has been extended using
28 ISCCP-DX data for the period July 2005-June 2006. From July 2006 onwards, the SEVIRI
29 observations are sampled to mimic the ISCCP-DX radiance data.

30 The FTH retrieval can reliably be applied under clear-sky and low-level cloud (cloud top
31 pressure larger than 700 hPa) conditions. The sampling is largely improved when

1 observations from the latter case are included. The scene selection is performed using the
2 cloud information (e.g. cloud cover and cloud top pressure) of ISCCP-DX for the period
3 1983-2009.

4

5 **3 Homogenization and extension to the SEVIRI era**

6 Prior to the FTH inversion, the CSR data record is adapted to the Meteosat-5 spectral
7 response function. Scatterplots of simulated BTs from Meteosat-5 and Meteosat-8 as well as
8 from Meteosat-5 and Meteosat-9 exhibit an excellent linear behavior with correlations >0.99
9 (not shown). A linear equation with slope a and with intercept b is thus used for adaptation.
10 The coefficients for Meteosat-5/Meteosat-8 are $a=1.0160$ and $b=-2.3498$. The coefficients for
11 Meteosat-5/Meteosat-9 are $a=1.0174$ and $b=-2.6033$.

12 The Meteosat time series is not fully homogeneous and exhibits breakpoints in the BT time
13 series mainly due to satellite changes and changes in calibration. Such breakpoints can be
14 removed using homogenization approaches. The homogenization applied here largely follows
15 the work by Picon et al. (2003).

16 The basic approach is 1) to use the European Centre for Medium-Range Weather Forecasts
17 (ECMWF) reanalyses ERA-Interim data (Dee et al., 2011) as input to the Radiative Transfer
18 for the TIROS Operational Vertical Sounder (RTTOV) 9.3 model, 2) to simulate Meteosat-5
19 BT and 3) to apply a linear regression to the observed BT for a month prior and after the
20 breakpoint (adapted from Picon et al., 2003):

21

$$\begin{aligned} BT_{\text{corrected}} &= \frac{a_{\text{before}}}{a_{\text{after}}} BT_{\text{original}} + b_{\text{before}} - b_{\text{after}} \frac{a_{\text{before}}}{a_{\text{after}}} \\ &= a' BT_{\text{original}} + b' \end{aligned} \quad (1)$$

22

23 Output from the regression is used to modify the satellite observed BT after the breakpoint in
24 a way that preserves the bias and the root mean square difference (RMSD) between observed
25 and simulated BT. The underlying assumption is the stability of ERA-Interim simulations
26 over the two-month period.

27 In order to perform the comparison, the following criteria have been applied:

1 • Only data at 6 UTC and 12 UTC have been considered.

2 • Simulations are performed in clear-sky condition only. This is further constrained by
3 considering the warmest 80% in simulated radiances only.

4 • The subdomain covers $\pm 45^\circ\text{N/S}$ and $\pm 45^\circ\text{E/W}$.

5 The homogenization coefficients a' and b' can be computed after a double application of the
6 linear regression and after substituting for the simulated BT.

7 This approach is applied to homogenize the change in calibration in January 2001 as well as
8 the Meteosat-7/8 and Meteosat-8/9 transitions using ERA-Interim data for the months of
9 December 2000 and January 2001, June and July 2006, and April and May 2007, respectively.
10 The following parameters have been computed and applied:

11 • From January 2001 onwards: $a'=0.98908$ and $b'=2.10135$,

12 • From July 2006 onwards: $a'=1.01510$ and $b'=1.00681$,

13 • From May 2007 onwards: $a'=0.974119$ and $b'=5.31705$.

14 After homogenization, the inter-calibration to HIRS (Breon et al., 2000) is applied to the CSR
15 data. The results are consistent with HIRS channel 12 observations onboard the NOAA12
16 (National Oceanic and Atmospheric Administration satellite) and the known bias of Meteosat-
17 5 is removed. This extended, homogenized, and inter-calibrated CSR data is used as input to
18 the FTH retrieval described in the next section.

19 Figure 1 shows the deseasonalized anomaly of the original and the updated BT as well as their
20 difference. The intensity of a breakpoint is the difference between the anomaly difference
21 (black values) prior and after the breakpoint. The breakpoints in January 2001, in July 2006
22 and in May 2007 have the following intensities: 0.5 K, -4.5 K, and 0.8 K, respectively. The
23 degree of homogeneity and stability has been largely improved.

24 Results from the Global Space-based Inter-Calibration System over the period May 2008-
25 December 2008 exhibit a difference in bias of slightly less than 0.5 K between Meteosat-8
26 and Meteosat-9 relative to the Infrared Atmospheric Sounding Interferometer. The observing
27 periods and the magnitude of the bias are different but the result confirms the presence of a
28 small bias between Meteosat-8 and Meteosat-9 in the early Meteosat-9 phase. Note that this
29 bias is significantly smaller in 2009 and later on.

1 As far as the first generation of Meteosat observations is concerned, the vicarious calibration
2 has been replaced by the calibration using the onboard black body in May 2000, and an
3 updated version has been implemented in January 2001
4 ([http://www.eumetsat.int/groups/ops/documents/document/pdf_ten_blackbody-
6 calibration.pdf](http://www.eumetsat.int/groups/ops/documents/document/pdf_ten_blackbody-
5 calibration.pdf)). In the meantime, the eclipses that occurred in 2000 and in 2001 affected the
7 overall performance (Köpken, 2001). Finally several gain changes have been applied, in
8 particular on January 9th 2001 ([http://www.eumetsat.int/groups/ops/documents/document/
10 pdf_rep_gains_hist_met7.pdf](http://www.eumetsat.int/groups/ops/documents/document/
9 pdf_rep_gains_hist_met7.pdf)).

11 **4 The free tropospheric relative humidity retrieval scheme**

12 **4.1 Rationale for the retrieval**

13 Assuming a random strong line theory and a constant lapse rate, Soden and Bretherton (1993)
14 showed that the observed BT is proportional to the logarithm of the mean RH over a deep
15 layer of the troposphere. Under these assumptions, the observed BT is mainly a function of
16 RH alone and not of temperature and specific humidity separately. The FTH is determined
17 from the following equation, which was analytically determined by Soden and Bretherton
18 (1996):

$$19 \ln\left(\frac{\langle RH \rangle p_0}{\cos \theta}\right) = a \times BT_{6.3\mu m} + b. \quad (2)$$

20 This equation links the clear-sky BT at 6.3 μm to the mean RH (defined with respect to water
21 only) of a broad layer of the troposphere. Eq. (2) also corrects for the effect of the satellite
22 viewing angle θ and it includes a scaling parameter p_0 , which is defined as the ratio between
23 the pressure at a temperature of 240 K and 300 hPa. The parameter p_0 represents the deviation
24 from a standard tropical profile where the 240 K isotherm is located at 300 hPa (see Soden
25 and Bretherton, 1993). In practice, this thermal parameter p_0 is computed using ERA-Interim
26 and ERA-40 (Uppala et al., 2005) temperature profiles collocated with Meteosat observations.
27 Level profiles from ERA-Interim (since January 2006) and ERA-40 (until December 2005)
28 models are preferred over the standard pressure levels. Indeed, Roca et al. (2009) proved that
29 the vertical resolution does not significantly affect the quality of the estimation of the FTH as
long as there is a given minimum number of pressure levels.

1 The fitting parameters (a and b) of the BT-to-⟨RH⟩ retrieval are determined once using a
2 representative dataset of thermodynamic profiles and by sampling the satellite field of view.
3 This training database is composed of temperature (T) and specific humidity (q) profiles
4 extracted from ERA-Interim restricted to clear-sky profiles. The database contains profiles
5 from the 1st day (4 time steps per day) of the January, April, July and October months for the
6 years 2001, 2006 and 2007 in order to have a significantly large set of profiles sampling the
7 various thermodynamical conditions encountered in the area limited to 30°N/S covered by
8 Meteosat. The clear-sky cases are defined using the ERA-Interim cloud fraction with a strict
9 value of 0 at all levels. In addition to the cloud screening, a quality check is performed for the
10 RH profiles (determined with respect to the water phase only) to remove the dry and
11 supersaturated cases where the RH reaches values below 1 %RH and above 100 %RH in the
12 free troposphere.

13 **4.2 Definition of FTH and selection of the vertical averaging operator**

14 There are several definitions of the vertical averaging operator ⟨•⟩ involved in the FTH
15 retrieval according to the interpretation of the observed radiation (Brogniez et al., 2009):

- 16 • Idealized Jacobian $\Delta BT/\Delta RH$, for which the weights are defined in temperature
17 coordinates (e.g. Soden and Bretherton, 1993; Soden and Bretherton, 1996),
- 18 • Local relative humidity Jacobian $J_{RH} = \partial BT / \partial RH$ (e.g. Roca et al., 2003; Brogniez,
19 2004; Brogniez et al., 2004),
- 20 • Transmission-derived weighting function $\partial \tau / \partial \ln(p)$ (e.g. Schmetz and Turpeinen,
21 1988; Stephens et al., 1996).

22 The selection of the most appropriate operator for the retrieval is based on a comparison
23 between the three regressions defined from the three definitions of FTH and using ERA-40
24 data. Figure 2 illustrates the results of this evaluation with scatter plots of the bias between the
25 weighted RH profiles (“observed”) and the computed FTH using simulated BTs (“retrieved”)
26 and the “observed” FTH. The statistics provided in Figure 2 clearly highlight the higher
27 quality of the fit obtained with J_{RH} , and thus the more precise definition of the FTH using J_{RH} .
28 Roca et al. (2009) analyzed the differences in spatial distribution of the peak height of the
29 three different averaging operators. They showed that the spatial distribution of the peak
30 heights resembles the spatial distribution of RH when using J_{RH} and the transmission-derived

1 weighting function. The latter exhibits a bimodal distribution that is not featured in J_{RH}
2 results.

3 The FTH is thus defined as the mean RH weighted by J_{RH} and normalized by the sum of
4 weights. The layer between 150 and 700 hPa is considered in the FTH computation:

$$5 \quad FTH(RH) = \frac{\sum_{p=700hPa}^{150hPa} RH(p) \times J_{RH}(p)}{\sum_{p=700hPa}^{150hPa} J_{RH}(p)}. \quad (3)$$

6 with $RH(p)$ defined between 0 and 100% with respect to the liquid phase of water.

7 Using this definition, the training of the algorithm yields to $a=-0.1248$ and $b=33.46$ (Roca et
8 al. (2009)).

9 **4.3 Practical considerations and examples**

10 The retrieval was applied to observations from Meteosat-2 to Meteosat-5 and from Meteosat-
11 7 to Meteosat-9 and provides FTH values within $\pm 45^\circ$ longitude and $\pm 45^\circ$ latitude. The FTH
12 is available at 3-hour resolution and as monthly averages (arithmetic averages over all valid
13 observations) on a regular latitude/longitude grid with a spatial resolution of $0.625^\circ \times 0.625^\circ$.
14 The data record covers the period from July 1983 to December 2009. For the reasons given in
15 Brogniez et al. (2009) the Meteosat-6 period (i.e. from March 1997 to May 1998) is not
16 covered.

17 In the following, relative units of FTH are given in % and the absolute units of FTH in %RH.

18 Figure 3 shows examples of instantaneous and monthly averaged products. Figure 4 illustrates
19 the FTH seasonal averages featuring strong FTH minima over northern and southern Africa
20 during boreal summer and strong FTH maxima in the Inter Tropical Convergence Zone
21 (ITCZ). The location and the extent of dry and wet areas and the corresponding minimum and
22 maximum FTH values strongly depend on the season. Figure 5 shows the time series of FTH
23 spatially averaged in the three regions shown in Figure 4. The three time series exhibit large
24 differences in amplitude and in shape of the annual cycle. Note that minor changes in the
25 definition of the regions have a noticeable impact on the time series, in particular on outliers.
26 Exemplary outliers are observed in March 1992 (in the South Atlantic) and in April 1990

1 (over northeast Africa) and they seem to be caused by deviations from the climatological
2 behavior of atmospheric dynamics on the regional scale.

3 As mentioned earlier, the identification of clear-sky and low-level clouds in the Meteosat
4 observations depends on ISCCP-DX data, and the quality of FTH depends on the quality of
5 the cloud classification. A strong degradation of the FTH quality can be expected when high-
6 level clouds are not correctly identified. Coastal areas exhibit reduced FTH data quality due to
7 problems with cloud detection before February 1997. The retrieval is also not reliable over
8 elevated terrain with surface pressures less than 700 hPa, since the observed signal might
9 contain contributions from the surface.

10 The FTH data record is referenced under digital object identifier (doi):
11 10.5676/EUM_SAF_CM/FTH_METEOSAT/V001 and is freely available at
12 <http://www.cmsaf.eu/wui>.

13

14 **5 Towards an uncertainty budget estimate**

15 This section briefly discusses the uncertainty budget estimate for the FTH product. Following
16 Chambon et al. (2012), the uncertainty budget is composed of three main uncertainty source
17 terms: 1) calibration uncertainty, 2) retrieval uncertainty and 3) sampling uncertainty. In line
18 with Roca et al. (2010) the calibration uncertainty is considered to be a systematic difference,
19 whereas the retrieval uncertainty depends on the details of the underlying algorithm. The
20 representativeness uncertainty depends on the temporal and spatial averaging and it vanishes
21 at the instantaneous pixel scale. This uncertainty depends on the number of independent
22 observations. In order to estimate this number, the correlation length in space and time for the
23 FTH has been estimated by analyzing variograms (see Roca et al., 2012 for details). Typical
24 spatial correlation lengths of 350 km and temporal correlation lengths of 7.5 hours have been
25 observed.

26 An upper bound calibration uncertainty of 1 K is considered (van de Berg et al., 1995). In this
27 case, the relative uncertainty on FTH is equivalent to the intercept of the retrieval (b in Eq.
28 (2)) and is between 10% and 15% (Roca et al., 2009). Results from the training of the FTH
29 retrieval allow the estimation of the retrieval uncertainty. Based on the tropical training a
30 RMSD of 2 %RH (8% when assuming an average FTH of 25%) and an average difference of
31 0.3 %RH were estimated. Assuming a daily average over a 2.5° grid box and a typical

1 standard deviation of 20% yields to a 10% relative sampling uncertainty. As a result, in this
2 idealized case, the total uncertainty in the estimation of the mean FTH is equally driven by the
3 calibration and by the sampling terms and to a lesser extent by the algorithm term. The
4 estimated total uncertainty is the square root of the sum of the three variances and in this case,
5 it is around 16-19% at one sigma (Roca et al., 2012).

6

7 **6 Evaluation**

8 **6.1 Data record for evaluation**

9 The ARSA version 2.7 is an archive of global radiosonde measurements of temperature, water
10 vapor and ozone profiles, which have been quality controlled and combined with auxiliary
11 observations. The ARSA archive has been developed and provided by the Atmospheric
12 Radiation Analysis group at the Laboratoire de Météorologie Dynamique (LMD, Paris,
13 France). In a first processing step, the radiosonde observations are quality controlled. For
14 example water vapor observations are considered only when they are available up to a
15 minimum pressure of 350 hPa, and the Thermodynamic Initial Guess Retrieval (also
16 developed at the LMD) climatological database is used to remove outliers. In a second step,
17 existing radiosonde measurements are combined with other reliable data sources. This step
18 depends on ERA-Interim data, which is also used for extrapolation into upper levels of the
19 atmosphere. Finally, the profiles are interpolated to 43 pressure levels from sea level pressure
20 to 0.0025 hPa. ARSA covers the period between 1979 and 2013 with few 10,000 observations
21 per month. More details can be found at
22 <http://ara.abct.lmd.polytechnique.fr/index.php?page=arsa>.

23 **6.2 Methodology**

24 The evaluation approach using ARSA follows the approach given in Brogniez et al. (2006,
25 2009). In order to simulate the Meteosat-5 observations, the RTTOV9.3 model (Matricardi et
26 al., 2004) has been applied to the ARSA profiles. The RTTOV model uses fast transmittance
27 algorithms based on accurate transmittances obtained from line-by-line computations
28 (GENLN2 for the 3-20 μm spectral range, Edwards, 1992). It thus depends on the
29 spectroscopic database (HITRAN-2000, Rothman et al., 2003). For the specific case of the 6.3
30 μm strong vibration-rotation absorption band by water vapor, the RTTOV model takes into

1 account the water vapor continuum (foreign-broadening and self-broadening, model CKD-
2 2.4, Clough et al., 1989) that has a non-negligible contribution in the water vapor band (e.g.
3 Stephens et al., 1996 and Soden et al., 2000). Based on the work of Brunel and Turner (2003)
4 referenced in the RTTOV v9 user guide, the bias uncertainty of RTTOV with respect to the
5 Meteosat water vapor channel is <0.1 K with a standard deviation of >0.3 K. RTTOV also
6 includes the K-matrix model that computes standard Jacobians ($=\partial BT/\partial X$, where X is any
7 atmospheric parameter of the model).

8 The following selection criteria are applied for the validation:

- 9 • Night time only,
- 10 • FTH $> 5\%$ to reduce potential surface contributions,
- 11 • Absolute differences in BTs < 3 K (i.e. approximately 3σ) and simulated and observed
12 BTs > 220 K to minimize cloud detection uncertainties of the ISCCP-DX algorithm.

13 A pair of ARSA and Meteosat observations is considered to be collocated when the temporal
14 distance is within 1.5 hours and the spatial distance is within 0.625° . Although ARSA
15 contains measurements from radiosondes launched on ships and at small islands, the
16 validation is dominated by observations over land.

17 To evaluate the quality of the FTH data record the relative and absolute bias, the
18 corresponding RMSD and the decadal stability are determined on a monthly basis and as
19 spatial averages. These uncertainty parameters are only considered when the number of valid
20 observations is larger than 10. More detail is given in Appendix A.

21 **6.3 Results**

22 The time series of monthly averages of absolute and relative biases of the FTH from Meteosat
23 and the FTH from ARSA, of their absolute and relative RMSD, and the number of valid
24 observations (N) are shown in Figure 6. Over the period July 1983 to December 2009 the
25 averaged relative bias, the averaged relative RMSD and the averaged N are -3.2% , 16.8% and
26 170, respectively. The averaged relative bias and 55% of the monthly relative biases between
27 the FTH from Meteosat and the FTH from ARSA are smaller than the Global Climate
28 Observing System (GCOS) requirement for the FTH (5%, verify GCOS-154). The relative
29 bias between the FTH from Meteosat and the FTH from ARSA exhibits strong temporal
30 fluctuations as well as a standard deviation of 4.5%. N strongly fluctuates in time and neither

1 small nor large values of N systematically coincide with maxima or minima in relative bias.
2 No correlation is found ($R = -0.01$) between N and the relative bias (see Eq. (A3) for a
3 definition of the Pearson's correlation coefficient R). Besides several local maxima and
4 minima in the time series of the relative bias between the FTH from Meteosat and the FTH
5 from ARSA the following features are noteworthy:

- 6 • Increase in relative bias between the summers 1988 and 1990,
- 7 • Maximum in relative bias in January 1996, with spurious biases in 1996.

8 These features most likely originate from changes in calibration procedures or
9 instrumentation as discussed in Sect. 3. The main difference between the comparisons
10 performed in the FTH space and in the BT space is that the features mentioned above appear
11 to be enhanced in the relative bias of FTH due to the exponential relation between CSR and
12 FTH and due to the normalization. Sharp summer minima in the relative bias with values
13 down to almost -15% are often featured. These minima are less obvious before 2001 and they
14 are most likely caused by the increased uncertainty in the FTH retrieval when the troposphere
15 is very dry. The relative RMSD between the FTH from Meteosat and the FTH from ARSA
16 exhibits weaker variations than the relative bias and a slight decrease between 1988 and 2006.
17 More than 66% of the monthly relative RMSD values are within 16-19%, which is the
18 estimated uncertainty of the FTH product.

19 The features mentioned above are also highlighted on the time series of the absolute bias and
20 RMSD (Figure 6, third panel) but they appear to be damped. The most obvious difference is
21 the lack of decrease in RMSD. The averaged bias and RMSD are stable and small over the
22 period 1983-2009 (-1.2 %RH and 5.0 %RH, respectively). The normalization step with
23 respect to the FTH likely causes the decrease in relative RMSD, which is an indication for a
24 general increase in FTH. Note that the normalization is done with respect to the FTH from
25 ARSA. Finally, the annual cycle of the absolute bias between the FTH from Meteosat and the
26 FTH from ARSA is less pronounced than the one of the relative bias because the
27 normalization has an amplifying effect (e.g. the months of July and August exhibit the
28 strongest minima and are characterized by larger dry regions than during the other months of
29 the year).

30 Scatter plots and histograms of the difference between the FTH from Meteosat and the FTH
31 from ARSA have been performed (not shown). A small bias between both FTH data records
32 is noteworthy, with larger ARSA-based FTH values than Meteosat-based FTH values. The

1 histogram of the differences peaks at -1 ± 1 %RH and is slightly skewed towards negative
2 values.

3 The stability of the data record is considered to be the slope of the linear regression between
4 the difference of the Meteosat FTH and the corresponding ARSA FTH (see Appendix A).
5 Based on the differences shown in Figure 6 (top panel) and after conversion from %/month
6 into %/decade, the decadal stability is found to be $0.5\pm 0.45\%$, which envelops the GCOS
7 requirement of 0.3% (GCOS-154). The uncertainty of the decadal stability is relatively large
8 and is obvious in Figure 6.

9

10 **7 Variability and trend analysis**

11 After introducing the frequency of occurrence of $FTH < 10\%$ (FTHp10), an analysis of the
12 standard deviation of FTH and FTHp10 on inter-annual scale is performed and the correlation
13 to the El Nino 3.4 and to the Quasi-Biennial Oscillation (QBO) indices is presented as well as
14 the differences between FTHp10 averaged over the 2000s and over the 1990s. This analysis
15 contributes to the discussion of linear trends and associated uncertainties in Sect. 7.3. Intra-
16 seasonal variability also contributes to the uncertainty but it is not considered here.

17 Throughout this section, full years are considered from January 1984 to December 2009.

18 **7.1 FTHp10**

19 Roca et al. (2011) introduced the frequency of occurrence of dry air as a marker of the
20 behavior of the dry part of the FTH distribution, namely the frequency of occurrence of
21 FTHp10. This climatological indicator corresponds to the radiatively sensitive range of FTH
22 and it is more resilient to the various assumptions in the retrieval (e.g. cloud clearing). A
23 strong contrast between minima and maxima reveals the spatial distribution of the moisture
24 field as seen in the FTHp10 seasonal climatology shown in Figure 7. The locations of
25 frequently dry areas coincide with locations of the dry seasonal averages in Figure 4. Figure 7
26 further highlights the dry area in the Southern tropical Atlantic Ocean where the air is very
27 dry more than 70% of the time, and it highlights the strong FTHp10 maxima in amplitude and
28 in spatial extent that occurs mostly during the boreal summer. The maximum FTHp10 is
29 found in the northeastern Mediterranean Sea in the boreal summer climatology.

1 7.2 Variability

2 An analysis of the FTH and the FTHp10 temporal and spatial variability is of general interest,
3 as outlined in the introduction. Also, the significance of FTH and FTHp10 trends is assessed
4 in the next section. In general, the coverage probability (or level of confidence) is a function
5 of the standard deviation and thus of the temporal variability. Therefore, the spatial
6 distribution of the standard deviation is analyzed on the inter-annual scale, the correlation to
7 El Nino 3.4 and QBO indices is discussed and the differences between FTHp10 averaged over
8 the 2000s and over the 1990s are analyzed.

9 In order to assess the interannual variability of FTH and FTHp10, Figure 8 shows the relative
10 standard deviation in FTH and in FTHp10 for each season. The relative standard deviation
11 maxima in FTH are located over the South Atlantic, the North Atlantic and central-east Africa
12 in DJF and over northeast Africa in JJA. Minima are associated with the ITCZ and the extra-
13 tropics. Note that the SON season clearly features the minimum averaged relative standard
14 deviation, mainly due to largely reduced maxima during this season. The positions of maxima
15 and minima in relative FTH standard deviation obviously coincide with the positions of the
16 FTH minima and maxima but also with gradient areas between dry and wet regions. These
17 results are in good agreement with the findings of Brogniez et al. (2009) and we recall here
18 the outstanding relevance of the variability in the dry free troposphere on the outgoing
19 longwave radiation (e.g. Udelhofen and Hartman, 1995; Sohn and Schmetz, 2004). Keeping
20 in mind that FTH minima are associated with FTHp10 maxima, similar results are found for
21 FTHp10 but only for the dry end of the distribution. Also, the strong maximum in relative
22 standard deviation of FTH featured in the northeast Africa region in JJA appears as a local
23 minimum in standard deviation of FTHp10, indicating a strong stability of the occurrence of
24 dry events in this region.

25 Following Roca et al. (2011) large-scale dynamics have a strong impact on the FTH
26 distribution and its variability. Brogniez et al. (2009) analyzed the FTH from MVIRI over
27 northeast Africa over the period 1983-2004 for the months of July/August and separated the
28 analysis into dry and wet years. The air masses of the driest years have been shown to mainly
29 originate from the tropics with a contribution from the extra-tropics, whereas the wet air
30 masses originate from tropical regions only. The variability of the extent of specific features
31 in FTHp10 over the northeast Africa in JJA and their position must also be analyzed to better

1 understand the overall dynamics and underlying distributions of FTH, their impact on OLR
2 and their changes over time.

3 The correlation coefficient R between seasonal averages of FTH, the FTHp10, and the El
4 Nino 3.4 index (available at <http://www.esrl.noaa.gov/psd/data/climateindices/list/>) has been
5 analyzed (not shown). Only DJF averages have been considered because El Nino events are
6 more intense during boreal winter. The coverage probability of the correlation has been
7 computed as described in Shi et al. (2013). The averaged correlation coefficient for positive
8 values only and for negative values only was found to be 0.14 and -0.11, respectively. The
9 number of grid cells with coverage probability of 95% or larger, relative to the total number
10 of grid cells (in the following referred to as “area fraction”) is 1.8% for FTH and 3.6% for
11 FTHp10. The FTH and the FTHp10 exhibit significant correlations over northeast Africa and
12 over parts of the Arabian Peninsula, with FTH (FTHp10 respectively) values being positively
13 (resp. negatively) correlated with the El Nino 3.4 index (0.48; resp. -0.45). The positive
14 correlation for FTH is consistent with results of Shi et al. (2013) who analyzed the correlation
15 between BT based on HIRS observations and the El Nino 3.4 index. Similarly, the correlation
16 between deseasonalized FTH and FTHp10 values and the QBO index was analyzed. The total
17 area fraction of significant correlation is slightly larger than for the correlation with the El
18 Nino 3.4 index but it is associated with averaged correlations about only 0.13. El Nino events
19 and the QBO thus have a minor contribution to the overall variability in FTH and FTHp10
20 over the considered area.

21 The FTH data record covers two full decades, namely the 1990s and the 2000s. The
22 difference between FTHp10 averaged over the 2000s and over the 1990s are shown per
23 season in Figure 9. Negative values occur when FTHp10 values in the 2000s are larger than in
24 the 1990s. Thus, the FTHp10 is generally larger in the 2000s than in 1990s. The maximum
25 area fraction of negative difference is 90% in DJF and the minimum area fraction of negative
26 difference is 71% in SON. The regions of minimum difference mainly coincide with gradient
27 areas between dry and wet regions and, to a smaller extent, with the dry regions themselves.
28 The largest connected area of positive difference is found over northeast Africa in SON and is
29 located at the west-southwest border of a regional FTHp10 maximum. This most likely
30 corresponds to an east-northeast retreat of the dry region between the 1990s and 2000s. These
31 differences were compared to the noise, that is, to the square root of the squared sum of the
32 internal decadal standard deviations (not shown). The maximum area fraction with an

1 absolute ratio between the differences and the noise that is larger than one is 9% in DJF. The
2 minimum area fraction with an absolute ratio between the differences and the noise that is
3 larger than one is 1% in SON. Areas of large absolute ratios are typically found between the
4 ITCZ and the neighboring dry areas. Even though most of the full area of interest exhibits an
5 increase in the frequency of dry events, it cannot be concluded that this tendency is
6 significant.

7

8 **7.3 Linear trend analysis**

9 **7.3.1 Methodology**

10 Two methods to analyze linear trends are tested: the “median of pair wise slopes regression”
11 method (named “Theil-Sen slope estimator”, Theil, 1950) and the linear regression method.
12 The “Theil-Sen slope estimator” method is more robust (i.e. less sensitive to outliers) than
13 linear regression and better suited to analyze linear trends in climatological data series. This
14 estimator takes into account the median of all pair wise slopes in the data. Approximately 600
15 pairs are needed to accurately estimate coverage probabilities (Wilcox, 2001) but the FTH
16 data record only has 312. A simple linear regression computation has thus also been
17 performed. Minor differences in absolute values and patterns of the different trends are
18 observed (not shown). The linear regression method is thus used in the following. The
19 estimated uncertainty of the trend is computed as described in Wilks (2011). Here, the
20 autocorrelation is neglected because seasonal averages are considered. Based on the estimated
21 uncertainty the coverage probability is estimated from a two-sided Student’s t-test.

22 To increase the accuracy of the trend analysis, 5x5 degrees averages are used instead of the
23 full resolution of the original product.

24

25 **7.3.2 Results**

26 Figures 10 and 12 show the linear trend in FTH and in FTHp10 as well as the associated
27 coverage probabilities per season over the period 1984-2009. Positive (negative respectively)
28 trends in FTH largely coincide with negative (positive resp.) trends in FTHp10, except over
29 southeast Europe in DJF where FTHp10 exhibits trends that are around 0%/year. Note that

1 FTHp10 only reflects the very dry events whereas the FTH has been averaged over the full
2 range of FTH values.

3 In general and for both the FTH and the FTHp10, the strongest trends are observed in areas
4 where there is a strong gradient between dry and wet regions and in dry areas. The dipole
5 structure of positive and negative trends in FTH and in FTHp10 over northeast Africa in DJF
6 and JJA is located at the borders of the associated dry regions. This can be an indication of a
7 displacement of (frequently) dry events in these areas. The observed trends in FTH and in
8 FTHp10 are hardly significant at the 95% confidence level. The minimum area fraction with
9 significant trends at this confidence level is 2.5% (in SON) and the maximum area fraction is
10 19.8% (in MAM), both found for the FTHp10. Largest connected areas of significant trends
11 are found in the extra-tropics. They coincide with generally large (resp. small) values of FTH
12 (resp. FTHp10) and small estimated uncertainties. They should be interpreted with care
13 because they are affected by potential oversimplifications of the retrieval scheme, which
14 might occur in the extra-tropical environment where the assumption of a constant lapse rate is
15 no longer valid. The estimated uncertainties are shown in Figures 11 and 13 for the FTH and
16 the FTHp10, respectively. Areas of large absolute trends frequently coincide with large
17 estimated uncertainties but also appear slightly shifted compared to the estimated uncertainty
18 maxima. An exception is the area of negative FTH trends, which is located over Brazil and
19 over the neighboring South Atlantic in SON. Strong similarities between the estimated
20 uncertainty and the interannual variability arise from comparing Figures 11 and 13 with
21 Figure 8. The interannual variability thus dominates the estimated uncertainty. Together with
22 the length of the record (26 years), this explains that significant trends in FTH and FTHp10
23 are hardly observed in this analysis.

24 Brogniez et al. (2009) reported that 1985 was among the driest years of the full FTH data
25 record. Since this specific year is at the beginning of the time series, it strongly impacts the
26 trend estimate. This impact has been assessed for the FTHp10 by removing the first two years
27 from the data record (not shown). After recomputing the trends, dry areas and gradient areas
28 between dry and wet areas exhibit larger positive trends and connected areas of such trends
29 cover larger areas. In particular, the dry area over northeast Africa in JJA and the dry area
30 over the South Atlantic in DJF exhibit significant positive trends at the 95% confidence level.
31 Interestingly, the areas of significant trends in the extra-tropics in MAM and over central
32 Africa in DJF almost disappear, which is associated with trends being close to zero. As stated

1 in Santer et al. (2011), these results confirm that trend estimates, their significance, and their
2 uncertainty depend on the period considered and in particular on the statistics of the data at
3 the start and at the end of the period.

4 Even though the estimated trends and the differences between FTHp10 averaged over the
5 2000s and over the 1990s are practically not significant and cover different periods, the
6 spatial patterns of increasing FTHp10 values generally coincide with negative differences.
7 Further analysis is needed to check if trends in dry areas, in particular for the FTHp10, are
8 reproduced in reconstructions of tropospheric RH from back trajectory models, in order to
9 check if such an increase can be related to a change in the large-scale dynamics of the last
10 saturations statistics. In this context, the dipole structures of FTH and FTHp10 trends over
11 northeast Africa in JJA and DJF are noteworthy. The associated feature in the estimated
12 uncertainty coincides with a similar feature in interannual variability. The back trajectory
13 analysis should be extended with an analysis of the position and extent of the dry areas,
14 specifically over northeast Africa.

15

16 **8 Conclusions**

17 Meteosat-2 to Meteosat-5 and Meteosat-7 to Meteosat-9 observations at $6.3\ \mu\text{m}$ are used to
18 retrieve information on humidity in the free troposphere. The inversion from BT to FTH is
19 reliable in the clear-sky case and in the presence of low-level clouds. Temperature data from
20 ERA-Interim is used to slightly improve the performance of the statistical retrieval scheme.
21 Thanks to a successful cooperation between a research institute and an operational service, the
22 FTH data record was extended into the SEVIRI era. The FTH data record is now released,
23 free of charge, by the European Organization for the Exploitation of Meteorological Satellites
24 (EUMETSAT) Satellite Application Facility on Climate Monitoring (CM SAF) at
25 <https://www.cmsaf.eu/wui> and is referenced under doi:
26 10.5676/EUM_SAF_CM/FTH_METEOSAT/V001. The FTH data record is available within
27 $\pm 45^\circ\text{N/S}$ and $\pm 45^\circ\text{E/W}$ with a spatial resolution of $0.625^\circ \times 0.625^\circ$, and it covers the period
28 1983-2009 with a temporal resolution of 3 hours. Monthly averages are also available.

29 Based on the comparison against FTH derived from the ARSA archive by using the RH
30 Jacobians for the integration of the RH profiles, the average relative bias of the FTH product
31 is -3.2% and its relative RMSD is 16.8%. The relative RMSD is in agreement with the
32 estimated uncertainty. The decadal stability is $0.5 \pm 0.45\%$. The relatively large uncertainty

1 estimate covers the GCOS requirement on humidity in the free troposphere of 0.3% per
2 decade. Due to the increase in bias between the summer 1988 and the summer 1990 and due
3 to a maximum bias in January 1996, with generally spurious biases in 1996, and even though
4 significant efforts have been dedicated to the homogenization of the Meteosat time series, the
5 quality of the FTH data record will benefit from a FCDR of the full Meteosat time series,
6 including the recovery of Meteosat-6 data in order to close data gaps in the time series.

7 The inter-annual relative standard deviation of FTH and FTHp10, the differences between
8 FTHp10 averaged over the 2000s and over the 1990s, and linear trends using seasonal
9 averages of FTH and FTHp10 have been analyzed. Maxima in inter-annual standard
10 deviations generally coincide with minima in FTH and maxima in FTHp10. Maxima in
11 absolute estimates of the trends in seasonal FTH and FTHp10 are associated with maxima in
12 standard deviation. As a result, the estimated trends are hardly significant. In the ITCZ region
13 where the results could be corrupted by the cloud-clearing method, the trends and their
14 uncertainties must be interpreted with caution. However, the maxima in trend estimate of
15 FTHp10 coincide with maximum absolute differences in FTHp10 averaged over the 2000s
16 and over the 1990s. The linear analysis performed in the dry free tropospheric subtropical
17 regions leads to results that are not significant but that are consistent with theoretical
18 considerations, in both the sign and the small magnitude of the change over the last ~25 years.
19 The combination of trend estimates, coverage probability and estimated uncertainty provides
20 valuable information to further analyze changes in the climate system. The analysis of the dry
21 end of the FTH distribution is very relevant not only because of its impact on OLR but also
22 because of the observed indication of small changes in value, in area, and in the associated
23 large variability.

24 This analysis will benefit from the availability of a Meteosat FCDR and a gap-free input data
25 record. It will also benefit from the extension of the temporal coverage to the most recent
26 times in order to promote a robust view on the decadal changes estimated here using linear
27 trends computations. Within the WMO SCOPE-CM, EUMETSAT leads the activity on the
28 “Inter-calibration of imager observations from time-series of geostationary satellites”. Among
29 others, this activity aims at the development and the provision of a Meteosat FCDR. This
30 FCDR will be used to improve the quality of the Meteosat-based FTH data record.

31 Initial comparisons to other available FTH records (e.g. based on HIRS and AMSU-B
32 observations) have already been carried out by CM SAF (<http://www.cmsaf.eu/docs>) and by

1 G-VAP (<http://www.gewex-vap.org>). sThe work presented here is part of the analysis of long-
2 term temporal changes within G-VAP. The extension of this analysis to other FTH data
3 records and the inter-comparison is in progress and is being performed by an international
4 team associated with G-VAP.

5

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15

16 **Appendix A: Statistical quantities**

17 In this Appendix, frequently used statistical quantities are defined:

- 18 • The relative systematic difference (or relative bias) between two estimations x_i and y_i of
19 the same variable is computed as:

$$\text{relative bias} = \frac{1}{N} \sum_{i=1}^N \frac{(y_i - x_i)}{x_i} \times 100 \quad (\text{A1})$$

20 The relative RMSD is defined as follows:

$$\text{relative RMSD} = \sqrt{\frac{1}{(N-1)} \sum_{i=1}^N \left(\frac{((y_i - \text{bias}) - x_i)}{x_i} \times 100 \right)^2} \quad (\text{A2})$$

21 In Eqs. (A1) and (A2) the sum is computed for all valid pairs within a given month, noted N .
22 Absolute bias and RMSD are computed by omitting the factor $100/x_i$ in Eqs. (A1) and (A2).
23 The term “absolute” is usually omitted.

- 24 • Decadal stability is computed by applying linear regression analysis to the results from
25 Eq. (A1). The slope of the regression is the temporal change of the relative bias per month in
26 %/month. When multiplying by 120, the decadal stability is in %/decade.

- 1 • The Pearson's correlation coefficient R between the variables x and y , each having N
2 elements, is defined as follows:

$$R = \frac{N \sum_{i=1}^N x_i y_i - \sum_{i=1}^N x_i \sum_{i=1}^N y_i}{\sqrt{N \sum_{i=1}^N x_i^2 - \left(\sum_{i=1}^N x_i \right)^2} \times \sqrt{N \sum_{i=1}^N y_i^2 - \left(\sum_{i=1}^N y_i \right)^2}} \quad (\text{A3})$$

- 3 Note that the N can differ from the N in Eqs. (A1) and (A2).

4

5 **Appendix B: Abbreviations**

- 6 Table C1 provides a list of the frequently used abbreviations.

7

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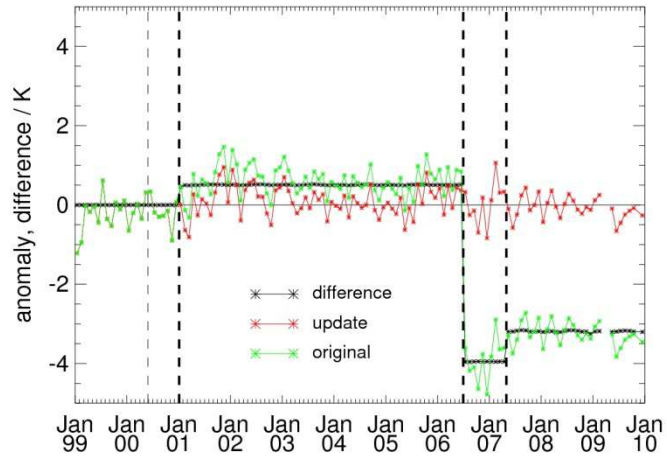
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1 Table C 1: List of abbreviations.

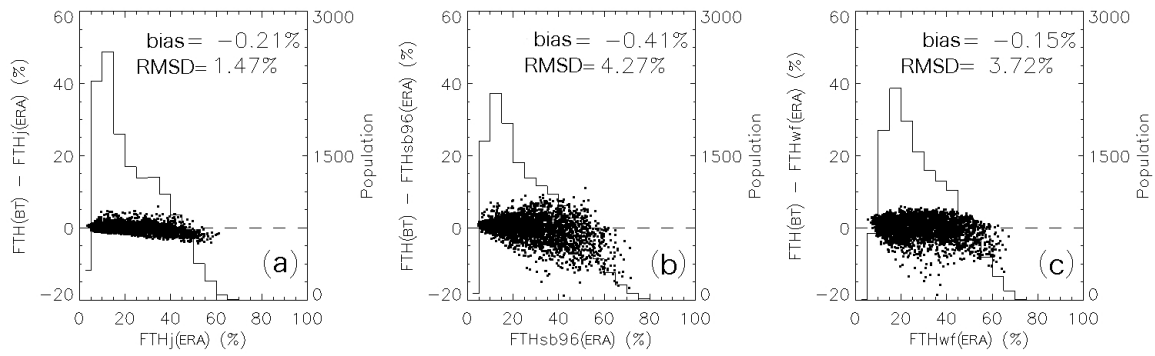
Abbreviation	Word/phrase
AMSU-B	Advanced Microwave Sounding Unit-B
ARSA	Analysed RadioSoundings Archive
BT	Brightness Temperature
CM SAF	Satellite Application Facility on Climate Monitoring
CSR	Clear Sky Radiance
DJF	December/January/February
ECMWF	European Centre for Medium-Range Forecasts
ERA-Interim, ERA-40	ECMWF Reanalysis
EUMETSAT	European Organisation for the Exploitation of Meteorological Satellites
FCDR	Fundamental Climate Data Record
FTH	Free Tropospheric Humidity
FTHp10	Frequency of occurrence of FTH<10%RH
GCOS	Global Climate Observing System
GEWEX	Global Energy and Water cycle EXperiment
G-VAP	GEWEX water VAPor assessment
HIRS	High-resolution Infrared Radiation Sounder
ISCCP-DX	International Satellite Cloud Climatology Project, DX type
ITCZ	Inter Tropical Convergence Zone
JJA	June/July/August
LMD	Laboratoire de Météorologie Dynamique
MAM	March/April/May
Meteosat	Meteorological Satellite
MVIRI	Meteosat Visible Infra-Red Imager
NetCDF	Network Common Data Format
NOAA	National Oceanic and Atmospheric Administration
OLR	Outgoing Longwave Radiation
QBO	Quasi-Biennial Oscillation
RH	Relative Humidity
RMSD	Root Mean Square Difference
RTTOV	Radiative Transfer for the TIROS Operational Vertical Sounder
SAF	Satellite Application Facility
SCOPE-CM	Sustained, Coordinated Processing of Environmental Satellite Data for Climate Monitoring
SEVIRI	Spinning Enhanced Visible and InfraRed Imager
SON	September/October/November
WMO	World Meteorological Organization



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2 Figure 1. Monthly deseasonalized clear sky brightness temperature anomaly for the original
 3 data (green) and for the updated homogenized data (red). The black line shows the difference
 4 between both anomalies. The thick dashed vertical lines represent the time when
 5 homogenization was applied and the thin dashed line represent the time when the black body
 6 calibration was implemented.

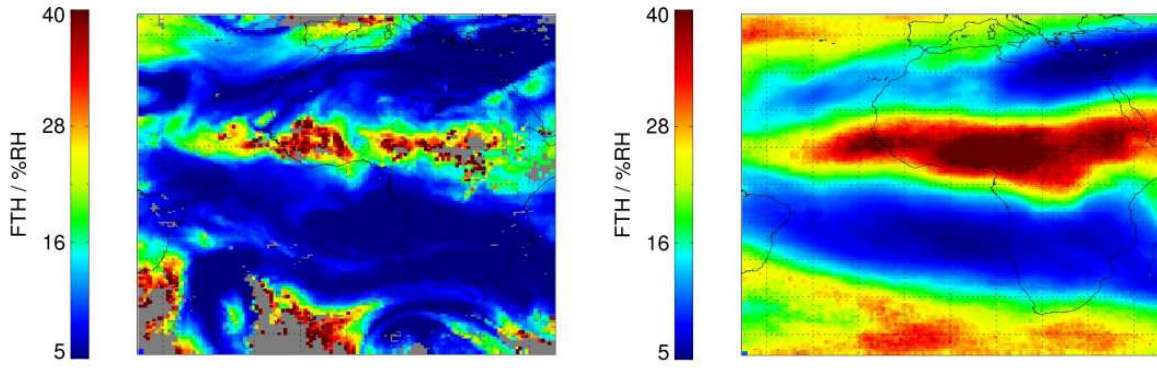
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2 Figure 2. Scatter plot of the bias between “retrieved” FTH from simulated BT and “observed”
 3 FTH using the local Jacobian (“FTHj”, left panel), the idealized Jacobian (“FTHsb96”,
 4 middle panel) and the transmission-derived weighting function (“FTHw”, right panel).
 5 Temperature and specific humidity from ERA-40 were used as input. The histogram gives the
 6 “observed” FTH population described on the right-hand side of the graphs. The average bias
 7 and the average RMSD are also given.

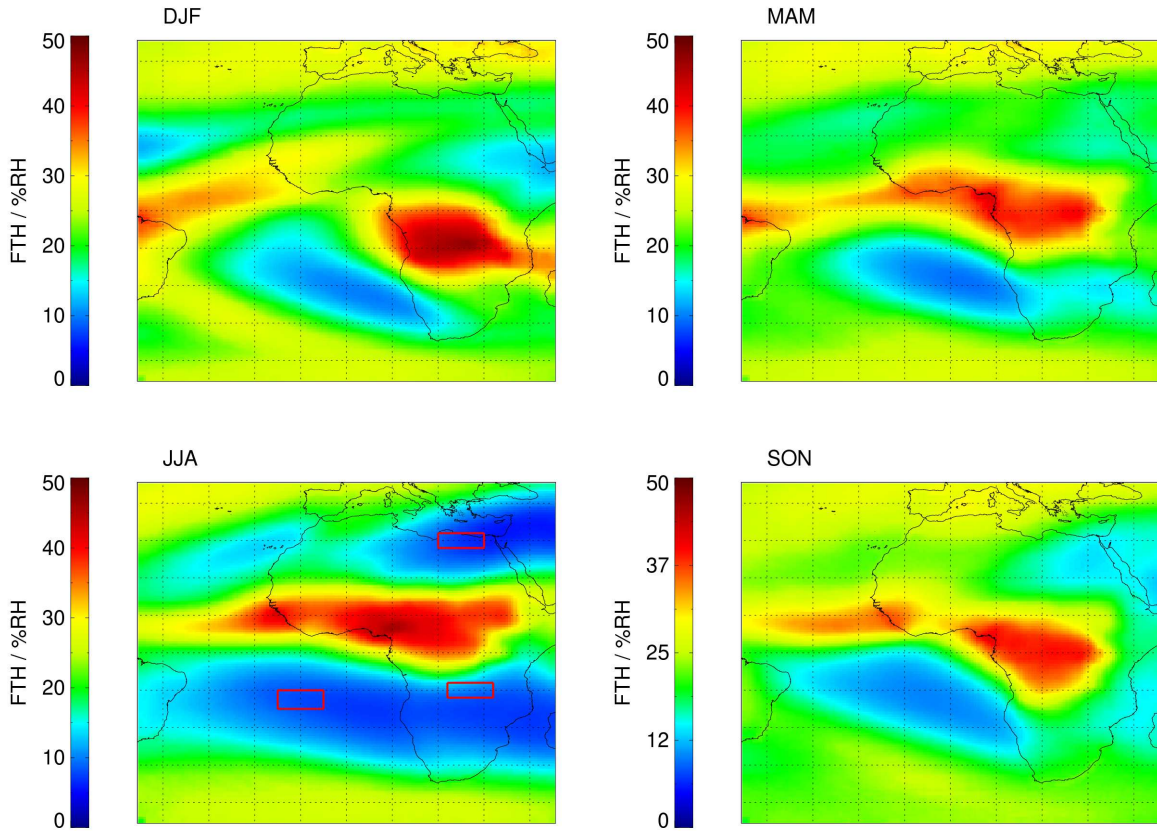
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2 Figure 3. Instantaneous FTH at 12 UTC on July 15th, 2009 (left panel) and monthly averaged
3 FTH for July 2009 (right panel). Undefined areas are in grey and they are usually associated
4 to cloud top pressures above 700 hPa.

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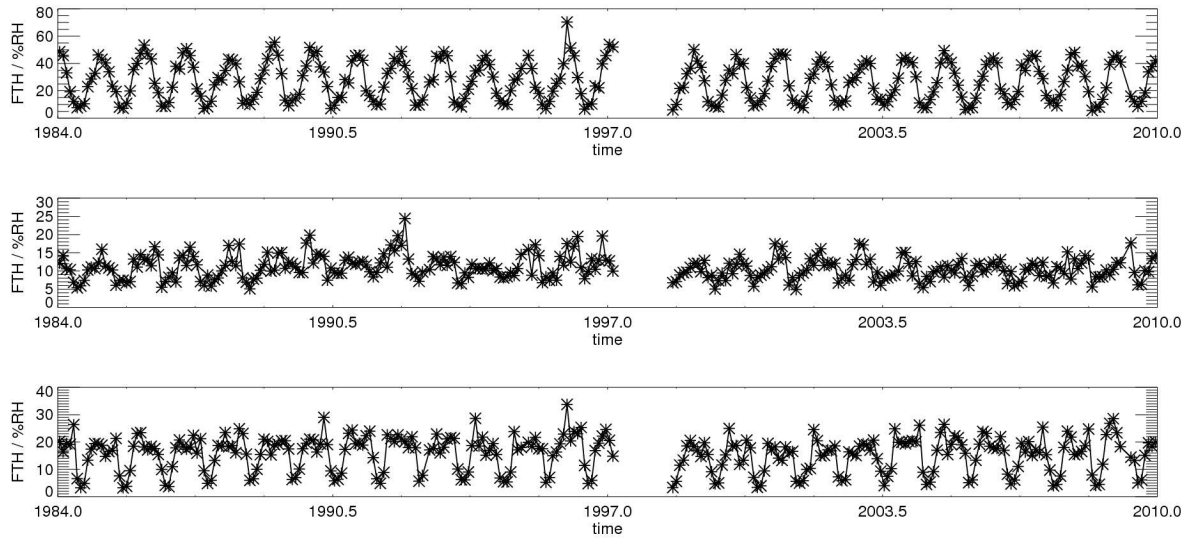


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3 Figure 4. FTH seasonal averages: December/January/February (DJF, top left),
 4 March/April/May (MAM, top right), June/July/August (JJA, bottom left) and
 5 September/October/November (SON, bottom right). The period considered is 1984-2009. The
 6 red boxes indicate regions for which the average time series is plotted in Figure 5.

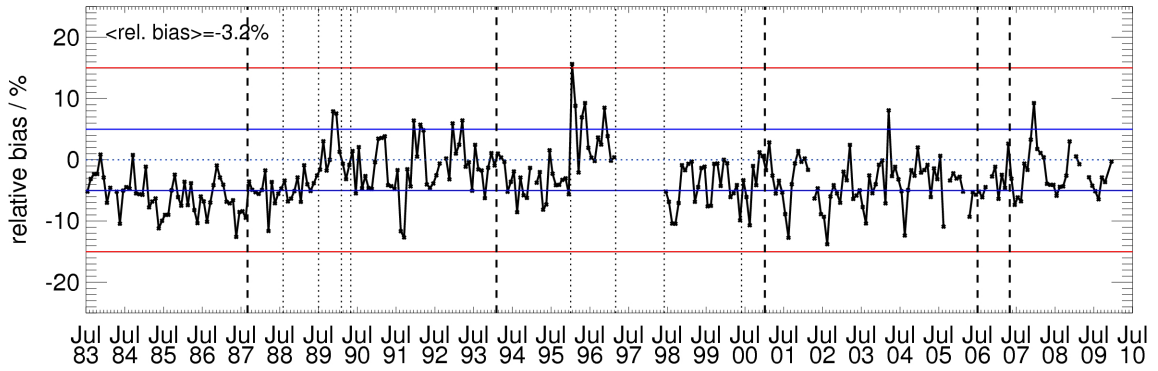
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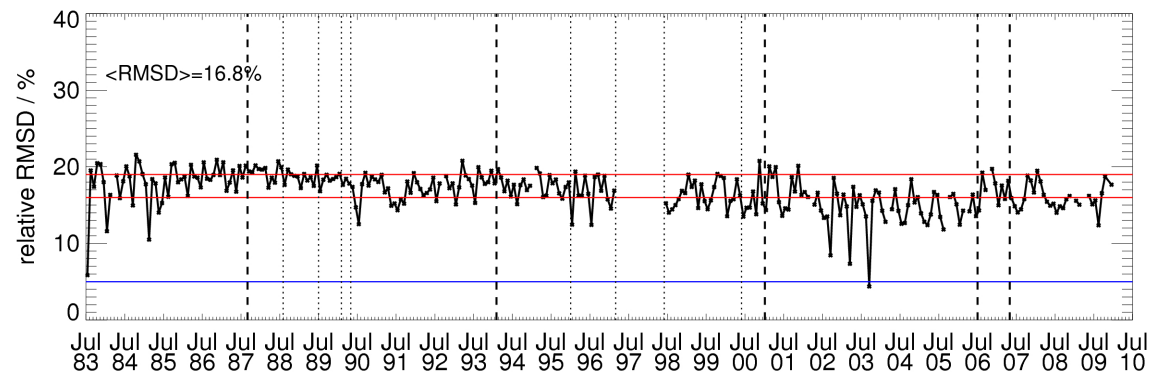
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2 Figure 5. Time series of regional FTH averaged over south central Africa (top), South
 3 Atlantic (middle) and northeast Africa (bottom). The regions are shown on the bottom left
 4 panel in Figure 4.

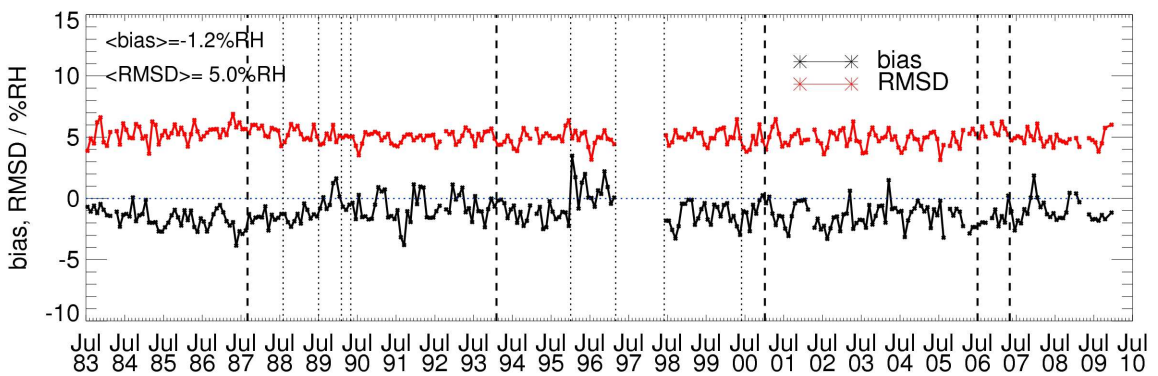
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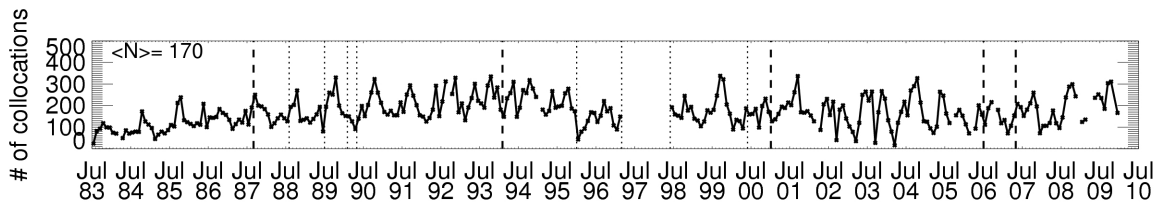
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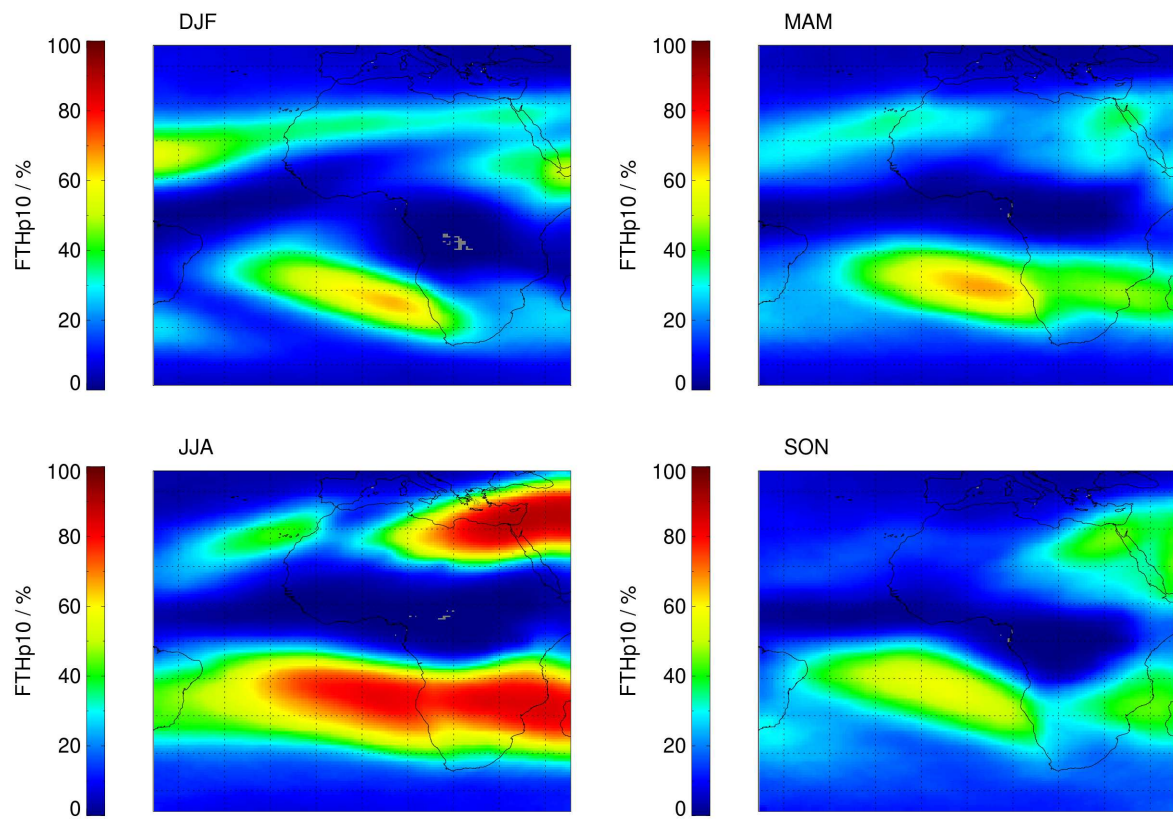
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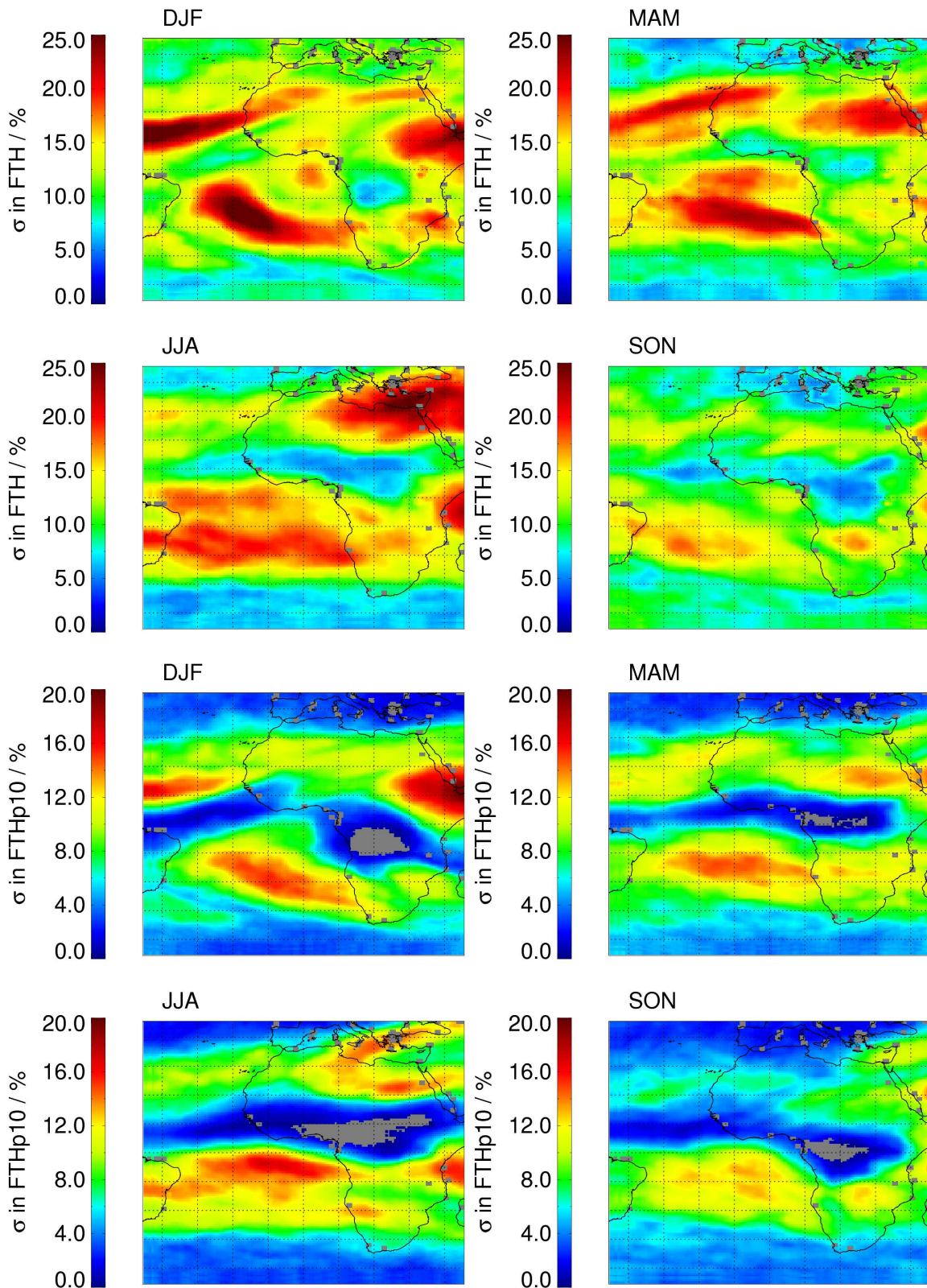
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5 Figure 6. Monthly mean of the relative bias between the FTH from METEOSAT and the FTH
 6 from ARSA (top panel), monthly mean of the relative RMSD between the METEOSAT- FTH
 7 and the ARSA-FTH (second panel), monthly mean of the absolute bias between the
 8 METEOSAT- FTH and the ARSA-FTH, monthly mean of the RMSD (third panel) and

1 number of valid observations (N, bottom panel). The thick dashed and the thin dotted vertical
2 lines indicate homogenization and major radiometric events. The averaged bias, the averaged
3 RMSD and the averaged N are also given. The colored lines in the first two panels highlight
4 the FTH requirements from Global Climate Observing System (GCOS-154), the error budget
5 estimate from Section 5 and a line at 15%, which is close to the peak values in maximum
6 relative bias.



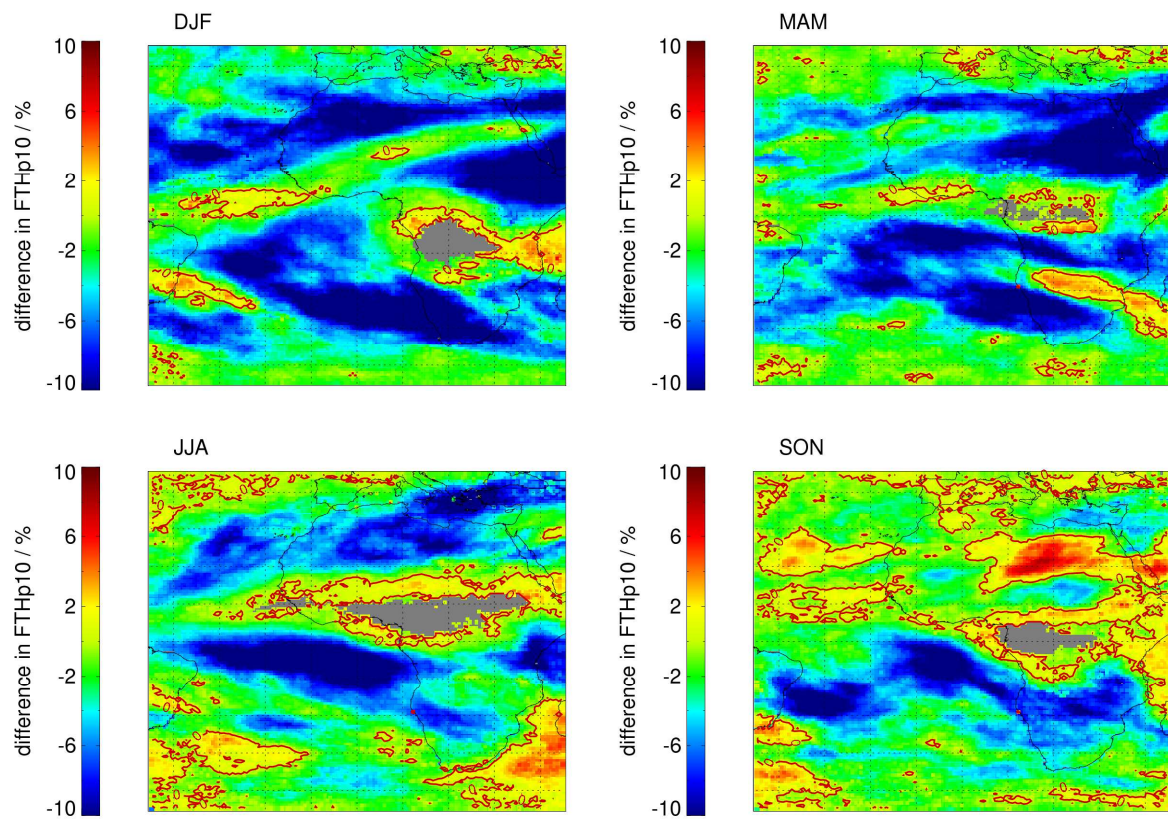
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- 2 Figure 7. Same as Figure 4 but for the frequency of occurrence of $FTH < 10\%$ (FTHp10).
- 3 Areas where only a small number of observations are valid are shown in grey.



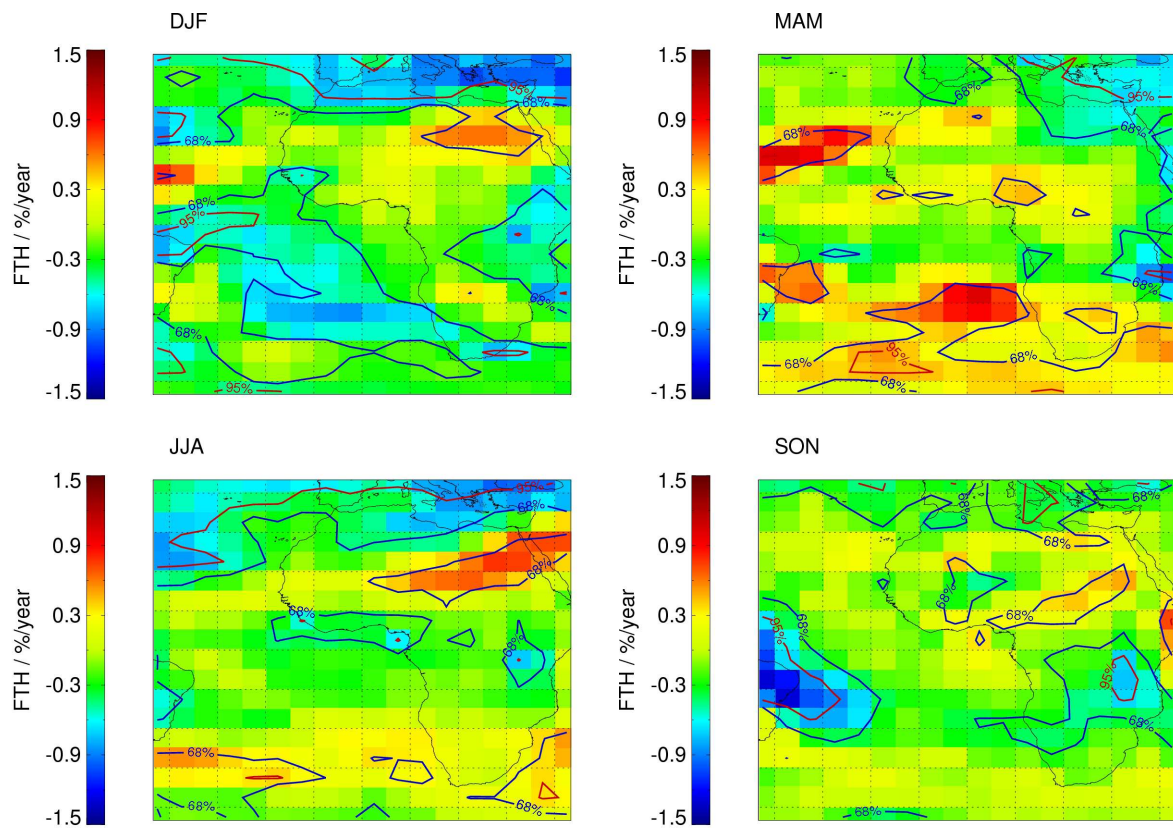
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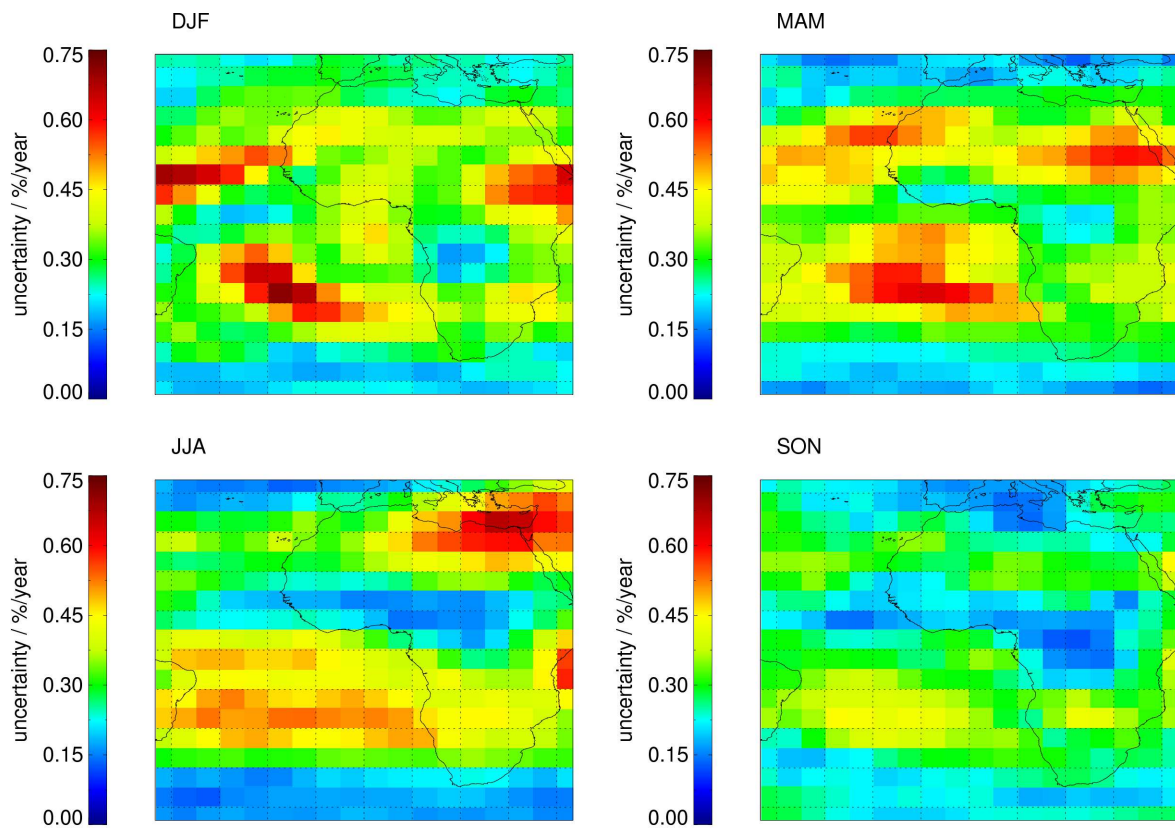
3 Figure 8. Relative standard deviation in FTH for each season (top four panels) and in FTHp10
 4 (bottom four panels) over the period 1984-2009. Areas where only a small number of
 5 observations are valid are shown in grey.



1
 2 Figure 9. Difference between decadal averages of FTHp10 in the period 1990-1999 and in the
 3 period 2000-2009. The difference was computed per season. Red contour lines indicate a 0%
 4 difference. Areas where only a small number of observations are valid are shown in grey.
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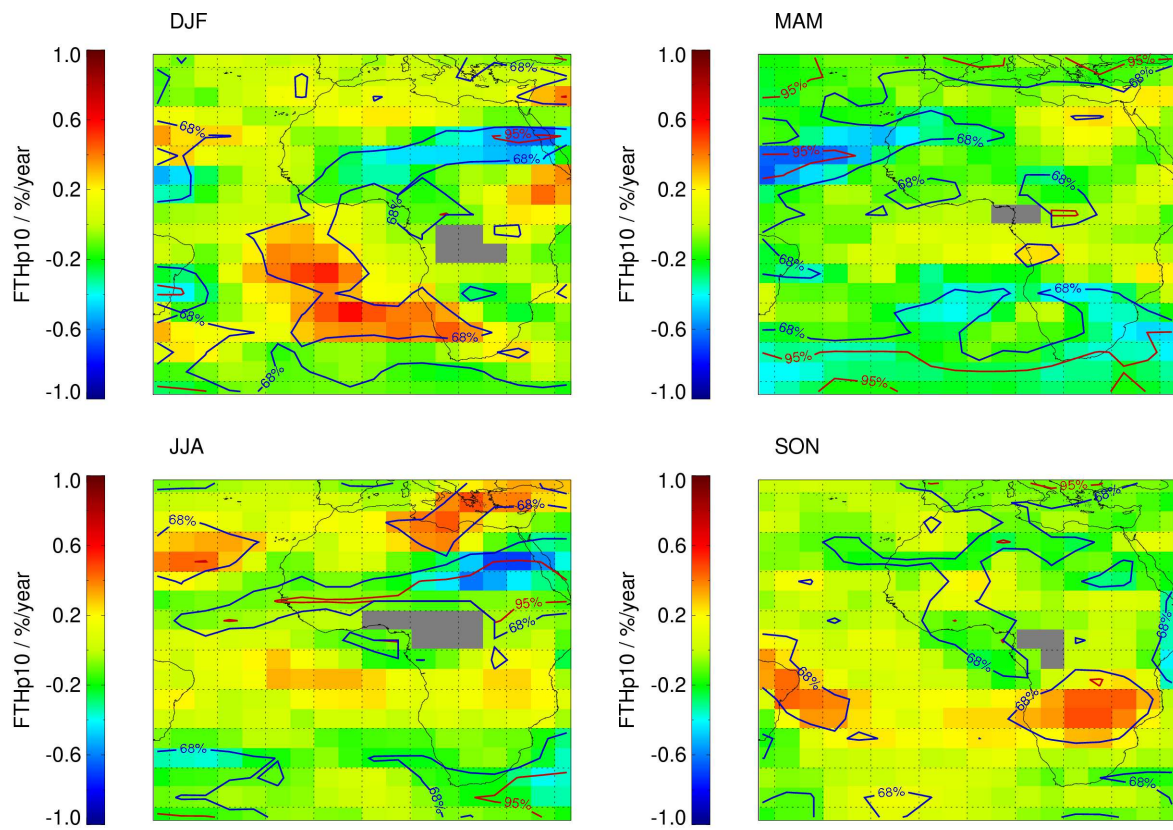


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 2 Figure 10. Same as Figure 4 but for the linear trend in relative FTH. The blue and red contour
 3 lines show coverage probabilities of 68% and 95%, respectively. Areas where only a small
 4 number of observations are valid are shown in grey.
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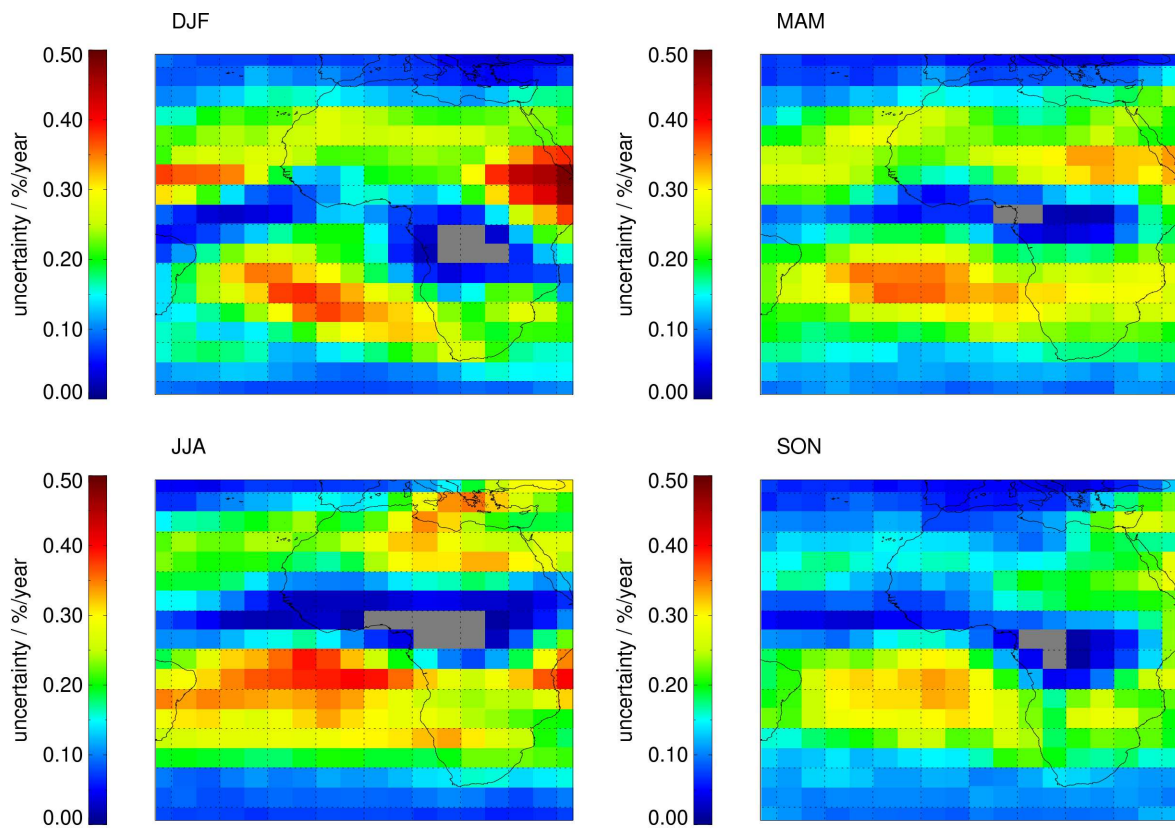
Figure 11. Same as Figure 10 but for the uncertainty of the linear trend in relative FTH.



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2 Figure 12. Same as Figure 10 but for FTHp10.

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1

2 Figure 13. Same as Figure 11 but for the uncertainty of the linear trend in FTHp10.

3