

Aerosol radiative effects in the ultraviolet, visible, and near-infrared spectral ranges using long-term aerosol data series over the Iberian Peninsula

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20 Abstract

21 A better understanding of the aerosol radiative properties is a crucial challenge for climate change
22 studies. This study aims to provide a complete characterization of aerosol radiative effects in
23 different spectral ranges within the shortwave (SW) solar spectrum. For this purpose, long-term
24 datasets of aerosol properties from six AERONET stations located in the Iberian Peninsula
25 (Southwestern Europe) are analyzed in term of climatological characterization and trends. Aerosol
26 information is used as input to the libRadtran model in order to determine the aerosol radiative
27 effect at the surface in the ultraviolet (ARE_{UV}), visible (ARE_{VIS}), near-infrared (ARE_{NIR}), and the
28 entire SW range (ARE_{SW}) under cloud-free conditions. Over the whole Iberian Peninsula, yearly
29 aerosol radiative effects in the different spectral ranges are: $-1.1 < ARE_{UV} < -0.7 \text{ W m}^{-2}$, $-5.7 <$
30 $ARE_{VIS} < -3.5 \text{ W m}^{-2}$, $-2.6 < ARE_{NIR} < -1.6 \text{ W m}^{-2}$, and $-8.8 < ARE_{SW} < -5.7 \text{ W m}^{-2}$. The four
31 variables showed positive statistically significant trends between 2004 and 2012, e.g., ARE_{SW}
32 increased $+3.6 \text{ W m}^{-2}$ per decade. This fact is linked to the decrease in the aerosol load, which
33 presents a trend of -0.04 unit of aerosol optical depth at 500 nm per decade, hence a reduction of
34 aerosol effect on solar radiation at the surface is seen. Monthly means of ARE show a seasonal
35 pattern with larger values in spring and summer. The aerosol forcing efficiency (AFE), ARE per
36 unit of aerosol optical depth, is also evaluated in the four spectral ranges. AFE exhibits a
37 dependence on single scattering albedo and a weaker one on Ångström exponent. AFE is larger (in
38 absolute value) for small and absorbing particles. The contributions of the UV, VIS, and NIR ranges
39 to the SW efficiency vary with the aerosol types. Aerosol size determines the fractions of
40 AFE_{VIS}/AFE_{SW} and AFE_{NIR}/AFE_{SW} . VIS range is the dominant region for all types, although non-
41 absorbing large particles cause a more equal contribution of VIS and NIR intervals. The
42 AFE_{UV}/AFE_{SW} ratio shows a higher contribution for absorbing fine particles.

43

44 1. Introduction

45 Atmospheric aerosol particles can absorb and scatter part of the total amount of solar radiation
46 entering the Earth's atmosphere. In fact, aerosols directly influence the Earth's energy budget and
47 act as cloud condensation nuclei modifying the cloud structure (e.g., Boucher et al., 2013). Aerosols
48 can either be produced by ejection into the atmosphere or by physical and chemical processes
49 within the atmosphere. Aerosol particles affect the radiative field by attenuating the direct
50 component thereby enhancing (or reducing under a highly absorbing aerosol) the diffuse one. They
51 also produce indirect effects by perturbing the Earth's atmospheric radiative balance by modulating
52 cloud albedo and fraction.

53 The aerosol radiative effect (ARE) is defined as the change in net radiation due to changes in
54 atmospheric aerosol properties and content. This is a key quantity in the determination of climate
55 change (e.g., Hansen et al., 1998). Most studies dealing with ARE have focused on discrete
56 wavelengths, whole shortwave (SW) solar radiation spectrum (e.g., Rajeev and Ramanathan, 2001;
57 García et al., 2008; di Sarra et al., 2008; Foyo-Moreno et al., 2014; Mateos et al., 2013a), longwave
58 (LW) radiation (e.g., Panicker et al., 2008; di Sarra et al., 2011; Antón et al., 2014), ultraviolet (UV)
59 interval (e.g., Hatzianastassiou et al., 2004; Kazadzis et al., 2009; Nikitidou et al., 2013), and visible
60 (VIS) range (e.g., Jayaraman et al., 1998; Horvath et al., 2002; Bush and Valero, 2003; Meloni et
61 al., 2003). With regards to surface SW radiative effect (ARE_{SW}), Di Biagio et al. (2010) obtained
62 the maximum radiative daily effects for different aerosol types in the central Mediterranean in the
63 period 2004-2007: -61 Wm^{-2} (desert dust aerosols), -26 Wm^{-2} (urban/industrial - biomass burning
64 aerosols) and -43 Wm^{-2} (mixed aerosols). All these negative figures point out a cooling of the
65 Earth's surface. Aerosol radiative effects for dust particles, for which the LW effect is relevant, in
66 the LW (ARE_{LW}) are expected to be smaller than in the SW and with positive sign (see, e.g., di
67 Sarra et al., 2011; Antón et al., 2014). Hence, this heating effect at the surface can partly offset the

68 cooling induced in the SW range. With respect to the ARE for the UV range (ARE_{UV}), Nikitidou et
69 al. (2013) analyzed the ARE in two different spectral regions in the UV range, 300-315 and 315-
70 360 nm. They found a stronger attenuation in the UV-B than in the UV-A.

71 The main goal of this study is to evaluate the ARE at the surface over the Iberian Peninsula, which
72 is a region of great interest because of its geographical position in Southwestern Europe, near the
73 African continent and the interface between the Atlantic Ocean and the Mediterranean Basin. Thus,
74 it is affected by frequent desert dust intrusions which modulate their aerosol climatology (Toledano
75 et al., 2007a, Bennouna et al., 2011; Pey et al., 2013; Valenzuela et al., 2012). In addition, this area
76 is also affected by a great variety of air masses loaded with different aerosol types: clean
77 continental, polluted plumes of central Europe, and marine aerosols. Hence, aerosol climatology at
78 six stations (Palencia, Barcelona, Cabo da Roca, Évora, Granada, and El Arenosillo) is also carried
79 out for different time periods between 2001 and 2012. Aerosol radiative effects as well as their
80 efficiency are calculated in four regions of the solar spectrum (ultraviolet, visible, near-infrared, and
81 shortwave) and the relative contribution of each range with respect to the whole solar spectrum is
82 analyzed as a function of the aerosol properties. Therefore, this study is intended to contribute to the
83 understanding of the aerosol impact on radiative budget over the Iberian Peninsula.

84 This article presents the following outline: detailed descriptions of the aerosol stations and the
85 database used are performed in Section 2; Section 3 includes the followed methodology; the results
86 obtained in the different analyses about the climatology of aerosol properties, aerosol radiative
87 effects, and aerosol forcing efficiencies are shown and discussed in Sections 4,5, and 6,
88 respectively. Finally, the main conclusions of this article are summarized in Section 7.

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90

91 **2. Columnar aerosol optical data**

92 The aerosol data are obtained from the Aerosol Robotic Network (AERONET) (Holben et al.,
93 1998). Six AERONET sites operating in the Iberian Peninsula are selected in this study: Palencia,
94 Barcelona, Évora, Cabo da Roca, Granada and El Arenosillo (see Table 1), all of them with a
95 minimum of 8 years of data sets of continuous observations. These sites present the largest records
96 of aerosol properties in the Iberian Peninsula in the AERONET network.

97 The standard instrument used in AERONET is the Cimel CE-318 radiometer. It performs direct sun
98 measurements at several wavelengths in the spectral range 340-1020 nm. Furthermore, the
99 instrument also measures sky radiance in the solar almucantar and principal plane configurations at
100 440, 670, 870 and 1020 nm wavelengths. A detailed description of this instrument was provided by
101 Holben et al. (1998). The direct sun observations are used to derive the spectral aerosol optical
102 depth (AOD) and the corresponding Ångström exponent. The sky radiances together with the AOD
103 are employed to retrieve a set of aerosol optical and microphysical properties via inversion methods
104 (Dubovik and King, 2000; Dubovik et al., 2006). These include particle size distribution, complex
105 refractive index, single scattering albedo (SSA), phase function, asymmetry parameter, fraction of
106 non-spherical particles, etc. (see
107 http://aeronet.gsfc.nasa.gov/new_web/Documents/Inversion_products_V2.pdf). Data are provided
108 in three database levels: 1.0 (raw data), 1.5 (cloud-screened) and 2.0 (cloud-screened and quality
109 assured).

110 The calibration of these instruments is performed following AERONET protocols by AERONET-
111 NASA, PHOTONS and RIMA networks every 12 months of operation approximately. The
112 estimated uncertainty is 0.01-0.02 for AOD (larger at shorter wavelengths) and ~5% for the sky
113 radiances (Holben et al., 1998). The SSA has an absolute uncertainty about 0.03-0.07 depending on
114 the aerosol load and type (Dubovik et al., 2000).

Level 2.0 aerosol optical depth data have been used in this work. However, it is well-known that when level 2.0 inversion data are used, the number of available observations of single scattering albedo (SSA) and asymmetry factor (g) is quite limited because these variables are only considered reliable when $AOD_{440nm} > 0.4$ ¹. Such AOD is mainly reached in the study region during Saharan dust or biomass burning events, therefore we would not have information on SSA and g for other conditions. To solve this issue, we have reduced the threshold of the level 2.0 inversion products. For this, we started with the level 1.5 data (for those quality-assured almucantar data that reached level 2.0) and applied the same criteria used by AERONET to elaborate the level 2.0 regarding the number of symmetrical angles, retrieval error and solar zenith angle (see http://aeronet.gsfc.nasa.gov/new_web/Documents/AERONETcriteria_final1_excerpt.pdf). However, a less restrictive threshold is applied to the AOD, which we restricted to cases with $AOD_{440nm} > 0.15$, instead of 0.4. This choice must be considered a compromise between the amount and the quality of the data. This kind of approach has been adopted by other authors using AERONET absorption data (e.g. Mallet et al., 2013). The threshold of 0.15 seems adequate analyzing the typical values of the AOD in the Iberian Peninsula (e.g., Bennouna et al., 2011; Obregón et al., 2012), because it can be considered a value to separate background aerosol conditions from episodic events with moderate or high aerosol loadings. The level 1.5-filtered data of SSA and g are daily averaged in order to have one value per day. In these conditions, the estimated uncertainty of the single scattering albedo is $\pm 0.05-0.07$ (Dubovik et al., 2000). Furthermore, for those days presenting level 2.0 data but also measurements in the 1.5-filtered level, we tested the uncertainty of our approach. We evaluated the difference in the SSA values of the level 1.5-filtered data with respect to the closest level 2.0 data. The mean relative differences in the

¹ Other inversion products, like the volume size distributions, are provided for all AOD levels.

137 SSA values between both methodologies are smaller than 1%, being in the same order that the
138 inversion uncertainty.

139 Lastly, when the AOD is low (<0.15 at 440 nm), there is no reliable information on the absorption
140 properties in the almucantar retrievals. Such low AOD is typical in our study region (e.g. almost
141 70% of observations at Palencia, Granada and Évora are below this threshold). If only cases with
142 $AOD_{440nm} > 0.15$ are considered in our study, the derived aerosol radiative effect would be
143 unrealistically large. To overcome this problem of representativeness, fixed values of SSA (0.90)
144 and g (0.75) have been used for the cases with $AOD < 0.15$ at 440 nm, considering typical values
145 for continental, desert, and maritime aerosols (e.g., Hess et al., 1998). In spite of the associated
146 uncertainties, our approximation (daily level 1.5-filtered values of these aerosol properties for AOD
147 > 0.15 together with a typical fixed value for low AOD cases) provides a good characterization of
148 the aerosol absorption of the particles present in the atmosphere. The data products and AERONET
149 database level are summarized in Table 2, where the estimated absolute uncertainties of AOD and
150 SSA are also provided.

151

152 **3. Methodology**

153 The ARE calculations are performed in the ultraviolet (ARE_{UV} , 280-400 nm), visible (ARE_{VIS} , 400-
154 700 nm), near-infrared (ARE_{NIR} , 700-2800 nm), and shortwave (ARE_{SW} , 280-2800 nm) intervals.
155 For this purpose, cloud-free simulations are carried out by means of a radiative transfer code.

156 The libRadtran model (Mayer and Kylling, 2005) has been shown to be a useful tool for obtaining
157 solar radiation data, presenting high accuracy (e.g., Román et al., 2014). Version 1.7 of the
158 libRadtran is used in this study with inputs of aerosol, total ozone column (TOC), precipitable water
159 vapor column (PWC), and surface albedo data. We performed simulations of ultraviolet (280-400

160 nm), visible (400-700 nm), near-infrared (700-2800 nm), and shortwave (280-2800 nm) radiation
161 during the periods indicated in Table 1. Total ozone column is provided by the Ozone Monitoring
162 Instrument (OMI) and Total Ozone Mapping Spectrometer (TOMS). Daily values of these
163 instruments are obtained from the Daily Level 3 Global Gridded products, which are downloaded
164 using the Giovanni application (<http://disc.sci.gsfc.nasa.gov/giovanni>). Level 2.0 AERONET PWC
165 data are used in the calculations. The uncertainty of this parameter is 10-15% (Holben et al., 1998).
166 In addition, retrievals of surface albedo at 440, 675, 870 and 1020 nm from the AERONET
167 algorithm are also used in this work. For land surface cover, this algorithm relies on the Lie–Ross
168 model (Lucht and Roujean, 2000), but considering the bidirectional reflectance distributions from
169 MODIS (Moody et al., 2005).

170 Aerosol properties obtained from AERONET measurements are also used as input to the libRadtran
171 model. Ångström coefficients, α and β , are utilized to compute a spectral aerosol optical depth in
172 the wavelengths of interest (Schuster et al., 2006). Ångström exponent α is obtained with the
173 measurements between 440 and 870 nm, while the turbidity β is obtained from the α value and
174 aerosol data at 1020 nm. Since the aerosol asymmetry factor, single scattering albedo, and surface
175 albedo are obtained at four wavelengths from AERONET in each measurement, three different
176 spectral regions are simulated with the libRadtran model. For computations in the UV range (280-
177 400 nm), the AERONET retrievals of aerosol asymmetry factor, aerosol single scattering albedo,
178 and surface albedo at 440 nm are used. The AERONET retrievals at 675 nm of the same variables
179 are used in the visible range (400-700 nm), while in the near-infrared region (700-2800 nm) we
180 used the average properties retrieved at 870 and 1020 nm. In each interval, these properties are
181 considered as wavelength independent. This choice to perform the radiative transfer simulations is
182 proven as adequate in the Appendix A. Other options in the model set-up are: extraterrestrial
183 irradiance values are taken from Gueymard (2004); profiles of temperature, air density, ozone and
184 other atmospheric gases are taken from the midlatitude summer/winter standard atmospheres; and

185 the radiative equation solver is the improved version of the discrete ordinate method of Stamnes et
 186 al. (2000) (DISORT2) calculated by 16-streams (e.g., de Miguel et al., 2011). After computing the
 187 solar irradiance in the different spectral intervals, the SW irradiance is evaluated by adding up the
 188 contributions of these three spectral regions.

189 In order to evaluate the aerosol radiative effect, the simulations under aerosol-free conditions are
 190 also computed with the same inputs as explained above, but with a fixed β value of 0.001.

191 The use of radiative transfer models fed with reliable experimental aerosol data to determine the
 192 ARE has been also employed in other studies (e.g., Barja and Antuña, 2011; Valenzuela et al., 2012;
 193 García et al., 2014).

194 Once the simulated radiometric values are obtained, ARE is derived for each interval (X represents
 195 UV, VIS, NIR, and SW) at the surface by:

196

$$197 \quad ARE_X = (X_{aer}^{\downarrow} - X_{aer}^{\uparrow}) - (X_{NOaer}^{\downarrow} - X_{NOaer}^{\uparrow}) \quad (1)$$

198

199 where X_{aer} and X_{NOaer} are the irradiances ($W m^{-2}$) for the X range under actual and aerosol-free
 200 conditions, respectively.

201 Daily values are obtained by the integration of the hourly data during the whole day (24 h)
 202 considering $ARE = 0 W m^{-2}$ for $SZA > 90^\circ$ (e.g., Bush and Valero, 2003; Valenzuela et al., 2012)
 203 and assuming cloud-free conditions along the day:

204

$$ARE_{daily} = \sum ARE_{hourly} \frac{dt}{24} \quad (2)$$

205

206 The aerosol forcing efficiency (AFE) is defined as the rate at which the radiative effect varies per
207 unit of AOD (e.g., Di Biagio et al., 2009; and the references therein). The linear relationship
208 between aerosol radiative effect and AOD is well known (see, e.g., Costa et al., 2004, 2006; Di
209 Biagio et al., 2009). Hence, in this study, ARE is obtained as the slope of linear fits in the ARE vs
210 AOD_{500nm} relationships. Therefore, AFE values are expressed in W m⁻² per AOD_{500nm}-unit (Wm⁻²τ⁻¹).
211

212 With respect to the temporal trends calculated in this study, the Sen's method (Sen, 1968) is applied
213 to evaluate the slope of a time series using the Mann-Kendall non parametric test to determine the
214 significance of these rates. The Sen's method is not greatly affected by outliers and can be
215 computed when there are gaps in the database (Collaud Coen et al., 2013). This is a common and
216 adequate method in temporal trend evaluation (e.g., Sánchez-Lorenzo et al., 2013). The trends
217 calculated in this study are obtained in the corresponding physical units per year. However, to unify
218 notation with previous studies dealing with the radiative effect trends of clouds and aerosols (e.g.,
219 Mateos et al., 2013b), the results are multiplied by 10 and expressed in physical units per decade. In
220 this way, the trends are also easier to read.

221

222 **4. Analysis of aerosol properties over the Iberian Peninsula**

223 A direct CIMEL retrieval (AOD at 440 nm) is selected to perform the climatological analysis
224 because the estimations of AOD_{500nm} (used in the ARE calculations) are obtained using α values.
225 Hence, we minimized the impact of other uncertainty sources in the AOD analysis. Besides, the
226 results for AOD_{440nm} and AOD_{500nm} do not differ excessively. In order to identify the differences in
227 the aerosol climatology over the six sites analyzed in this study, the monthly distribution of the

228 daily values of the AOD_{440nm} and α are evaluated using the database mentioned in Table 1. All the
229 available level 2.0 AERONET measurements are used in this section.

230 Figure 1 shows the climatology of the aerosol load by box whisker plots. Several conclusions can
231 be drawn from this figure. The highest values of the AOD occur in Barcelona, as can be expected
232 because it is a large city. With respect to the monthly average values (triangles in the figure), the
233 central stations in the Iberian Peninsula (Palencia and Évora) exhibit AOD_{440nm} below 0.2, while the
234 southern sites (Cabo da Roca, Granada, and El Arenosillo) show aerosol load over 0.2 during
235 summer months. The AOD_{440nm} seasonal distribution is seen, with maximum values in summer and
236 minimum ones in winter. However, the seasonality becomes more evident in the stations outside the
237 central area of the Iberian Peninsula. The large differences between median and average values for
238 some months evidence a large impact of high aerosol optical depth events on the monthly
239 climatology. In this line, the bimodality of the monthly AOD climatology (with two maximum
240 monthly means occurring in March and summer months) observed for the El Arenosillo site has
241 been already reported by previous studies (e.g., Bennouna et al., 2011), and directly attributed to
242 desert dust intrusions from the African continent.

243 To go further in the characterization, α allows for a better understanding of the particle size over
244 each site. Figure 2 shows the climatology of this variable over the six stations using also box
245 whisker plots. Analyzing the monthly average means, α values larger than one, indicative of the
246 predominance of fine particles, are dominant over Barcelona, Palencia, and Évora. The other three
247 stations (Cabo da Roca, Granada, and El Arenosillo) present monthly α averages over and below 1,
248 which means a larger variety of aerosol sizes over these stations. A seasonal dependence over
249 Granada site is seen, with winter months dominated by fine particles and summer months by larger
250 ones (see also Navas-Guzman et al., 2013). Values of α present a large variability during summer
251 which is indicative of the influence of different aerosol types including biomass burning events and

252 Saharan dust transport (e.g., Pérez-Ramírez, 2008). The monthly distribution of α is symmetric with
253 similar average and median values through the year for the six sites.

254 With the daily AOD and α values, it is possible to classify the origin of the aerosol particles.
255 Previous studies suggest different thresholds of AOD and α (e.g., Hess et al., 1998; Pace et al.,
256 2006; Toledano et al., 2007b). A simple classification, which can be used for the whole Iberian
257 Peninsula, of aerosol type is carried out in this study. The threshold between fine and large particles
258 is placed at $\alpha = 1$, while the situations with a high aerosol load are those with $\text{AOD}_{440\text{nm}} > 0.2$.
259 Therefore, aerosol particles can be classified in four types: maritime ($\text{AOD}_{440\text{nm}} < 0.2$ and $\alpha < 1$),
260 desert dust ($\text{AOD}_{440\text{nm}} > 0.2$ and $\alpha < 1$), continental clean ($\text{AOD}_{440\text{nm}} < 0.2$ and $\alpha > 1$), and
261 continental polluted ($\text{AOD}_{440\text{nm}} > 0.2$ and $\alpha > 1$). Note that the limit of $\text{AOD}_{440\text{nm}} < 0.2$ is arbitrary
262 and this value could be adjusted according to the sites, which likely produce a different distribution
263 in the pie diagrams. Actually, even close stations can present slight differences in the α -AOD
264 classification (see, e.g., Obregón et al. 2012). However it is not the aim of this work to provide an
265 extensive aerosol climatology, but rather to demonstrate the great variety of air masses over Iberia
266 which transport different aerosol types. Although other types, such as biomass burning or mixed
267 aerosols, are placed in the boundaries of these types, this simple classification can provide
268 information about the aerosol sources for the six sites. The classification used here is in line with
269 the previous studies. For instance, Toledano et al., (2007b) proposed for El Arenosillo site similar
270 thresholds (see their Table V), although they identified continental polluted aerosols with an
271 $\text{AOD}_{440\text{nm}}$ larger than 0.35 and $\alpha > 1.4$. Pace et al., (2006) proposed at Lampedusa island (Central
272 Mediterranean) a desert dust identification when $\text{AOD}_{440\text{nm}} \geq 0.15$ and $\alpha \leq 0.5$.

273 Figure 3 shows pie diagrams with the frequency of occurrence of the four aerosol types. The six
274 diagrams agree pointing at continental clean as the main type of aerosols over the Iberian Peninsula.
275 In Barcelona, there is also an important contribution of continental polluted, since Barcelona is a

276 large coastal city with relevant pollution levels from vehicular and ship traffic (e.g., Reche et al.,
277 2011). The influence of maritime aerosols is notable at El Arenosillo, Cabo da Roca, and Évora
278 sites (see also e.g., Bennouna et al., 2011; Obregón et al., 2012). Furthermore, desert dust events are
279 shown to be common in the Iberian Peninsula with a higher occurrence at Granada and El
280 Arenosillo sites (the two closest points to the African continent and hence to the Saharan desert)
281 (see also Toledano et al., 2007b; Guerrero-Rascado et al., 2009; Antón et al., 2012). For instance,
282 the minimum values of α obtained for Granada station during summer months are linked to the
283 higher likelihood of desert dust events (Valenzuela et al., 2012), being sometimes associated with
284 high aerosol loads (Córdoba-Jabonero et al., 2011). These results corroborate the findings obtained
285 by previous studies about desert dust events over the Iberian Peninsula (see, e.g., Lyamani et al.,
286 2005; Toledano et al., 2007b; Cachorro et al., 2008).

287 The temporal trend of aerosol load can be established over the last decade in the Iberian Peninsula.
288 The yearly values of AOD_{440nm} at the six sites are shown in Figure 4. The geographical distribution
289 of AOD through the Spanish geography is observed in the figure. Barcelona site presents yearly
290 values over ~0.2. Granada, El Arenosillo, and Cabo da Roca exhibit yearly means in the interval
291 between 0.15 and 0.22, while the means for Palencia and Évora sites are slightly lower in the range
292 0.12-0.18. Analyzing the six sites together, the year of 2010 presents one of the minimum values of
293 AOD_{440nm}, while the maximum averages seem to appear at the early 2000s. The different sampling
294 of AOD measurements in the six sites can produce discrepancies because different events are or are
295 not captured in each database. In addition, possible technical problems and meteorological
296 conditions (CIMEL aerosol data are recorded under cloud-free skies) cause a non-equally
297 distribution through the year. Overall, summer is the season with the largest contribution of data,
298 followed by spring, autumn, and winter. Looking at the years with a large sampling (>200 days in,
299 at least, four stations), 2005, 2007, and 2011, all the features mentioned above are corroborated for
300 these particular years. The minimum of 2010 occurs when two Southern sites (El Arenosillo and

301 Cabo da Roca) have not enough data to evaluate the yearly mean. Hence, we cannot ensure that the
302 apparent minimum of AOD recorded that year is linked to global-scale phenomena or to more local
303 conditions at the other sites. During 2010 a persistent negative phases of North Atlantic Oscillation
304 (NAO) and Quasi Biennial Oscillation (QBO) indexes was observed (e.g., Steinbrecht et al., 2011),
305 and the connection between air mass transport at global scale and particulate matter (at the surface)
306 is proved by Pey et al., (2013) in the Eastern Iberian Peninsula.

307 With respect to the temporal trend rates, the evolution of these yearly values seems to be weak. The
308 evaluation of the trend rates (see Section 3 for details) produces the more statistically significant
309 trend for the Barcelona site, where a decrease of the aerosol load of 0.09 AOD_{440nm}-unit per decade
310 is observed with a *p value* of 0.02. Évora and Palencia stations showed trend rates of -0.06 and -
311 0.04 AOD_{440nm}-unit per decade with *p values* of 0.06 and 0.10, respectively. The *p values* for the
312 other sites point out non-statistically significant trends. In spite of that, the sign of the temporal
313 trends is negative for all of them. Hence, a slight reduction of the aerosol load over the Iberian
314 Peninsula is observed since 2000. This result obtained in the Southeastern Europe is in line with the
315 long-term analysis of AOD series performed in Northern Germany and Switzerland by Ruckstuhl et
316 al. (2008). These authors highlight a strong decrease of aerosol load starting in 1985, while the
317 values are stabilized since about 2000.

318 The reasons behind the decrease in the aerosol load since the early 2000s are a mixed of
319 anthropogenic and natural sources. As was reported by Aas et al. (2013), the particulate matter
320 (PM) emissions in the Iberian Peninsula have decreased around 25% between 2000 and 2011.
321 Furthermore, observational PM data in different Spanish sites have also shown a decrease trend in
322 the 2000s (e.g., Barmpadimos et al., 2012; Cusack et al., 2012; Pey et al., 2013; Bennouna et al.,
323 2014; Mateos et al., 2014). This fact can be understood by the effect of the current economic crisis
324 and the implementation of new environmental laws to control the pollution (e.g., Querol et al.,

2014). In addition, recent studies have shown that natural aerosols have also decreased in the last decade. For instance, Gkikas et al. (2013) reported, using satellite AOD estimations, that strong and extreme desert dust episodes in the Mediterranean decreased in the period 2000-2007 over land surfaces. This trend is understood due to the low spring and summer frequencies in 2005 and 2007 and the high frequencies in 2000 and 2003. As it was shown by Pey et al. (2013), one possible reason behind this trend is the atypical trajectories followed by the air masses emerging from North Africa in summer since 2006. Hence, both columnar and surface aerosols have pointed out a decrease in the aerosol load over the Iberian Peninsula, which has increased solar radiation levels reaching the surface in the 2000s (Mateos et al., 2014).

334

335 **5. Inter-annual and intra-annual evolution of ARE**

From the daily data, the yearly ARE for each station and spectral range are evaluated to analyze their inter-annual changes (see Figure 5). In spite of the high variability of the yearly values with large standard deviations (see the vertical bars for Palencia station in the figure), the radiative effects of atmospheric aerosols have slightly declined over the last years. The patterns of ARE in the UV, VIS, NIR, and SW ranges are similar, since the inter-annual changes are simultaneously observed in the four spectral intervals. The significance levels of the temporal trends (Mann-Kendall nonparametric test at the 95% confidence interval) are evaluated. Évora and Palencia sites exhibit statistically significant trends in the periods 2005-2012 and 2003-2012, respectively. The trends for the aerosol effects for Palencia (Évora) are: $+4.9 (+3.2) \text{ Wm}^{-2}$ per decade in ARE_{SW} , $+3.3 (+2.1) \text{ Wm}^{-2}$ per decade in ARE_{VIS} , $+0.1 (+0.08) \text{ Wm}^{-2}$ per decade in ARE_{NIR} , and $+0.06 (+0.03) \text{ Wm}^{-2}$ per decade in ARE_{UV} . The *p values* of the trends for Évora site range between 0.009 and 0.019, while for Palencia site are between 0.02 and 0.03. The other four stations present positive trends in all the spectral ranges, but they are not statistically significant, at least, at the 95%

349 confidence interval. The decrease in the aerosol effects is in line with the previous fall in the AOD
350 values in the 2000s mentioned in Section 4. Furthermore, this slight reduction in the radiative
351 effects of the atmospheric aerosol over the Iberian Peninsula could partially contribute to the
352 increase in the levels of SW radiation at the surface (the brightening phenomenon) in this region
353 reported by Sanchez-Lorenzo et al. (2013) and Mateos et al. (2013b).

354 To establish the general behavior of the ARE over the whole Iberian Peninsula, the yearly values
355 using the six ground-based stations are evaluated. Only those years with, at least, simultaneous
356 measurements at three sites are considered in these averages, and consequently, the time period is
357 limited to 2004-2012. Figure 6 shows the evolution of the ARE and AOD at 500 nm for the entire
358 peninsula. The previously discussed reduction of aerosol load in the six individual datasets is again
359 corroborated with the mean data series. This decline produces a consequent decrease in the aerosol
360 radiative effect at the four spectral ranges. The temporal trends of these yearly values are evaluated,
361 and all the trends resulted statistically significant at the 95% significance level are shown in Figure
362 6. Overall, ARE_{SW} over the Iberian Peninsula increased 3.6 W m^{-2} per decade while the aerosol
363 reduced $0.04 \text{ AOD}_{500\text{nm}}$ -unit per decade. The yearly aerosol radiative effects over the entire
364 peninsula are in the ranges: $-1.1 < ARE_{UV} < -0.7 \text{ W m}^{-2}$, $-5.7 < ARE_{VIS} < -3.5 \text{ W m}^{-2}$, $-2.6 <$
365 $ARE_{NIR} < -1.6 \text{ W m}^{-2}$, and $-8.8 < ARE_{SW} < -5.7 \text{ W m}^{-2}$. The larger contribution of the visible
366 spectral region with respect to the whole solar spectrum was also noticed by Bush and Valero
367 (2003), and this is expected since the maximum of shortwave radiation is found in this interval. The
368 relationship between ARE and $AOD_{500\text{nm}}$ is analyzed more in detail in Section 6, when the aerosol
369 forcing efficiency is evaluated for each ground-based station.

370 In addition to the inter-annual changes, the intra-annual behavior is also analyzed. For this purpose,
371 the annual cycle (12 monthly means) is evaluated for the six stations (see Figure 7). A seasonal
372 pattern is seen in ARE_{UV} and ARE_{VIS} , and therefore, ARE_{SW} . However, ARE_{NIR} does not follow a

373 seasonal pattern, particularly at the Évora and Palencia stations given that ARE_{NIR} remains nearly
 374 constant. Small differences among the six stations are observed in the annual cycle during the cold
 375 seasons. The aerosol radiative effects are stronger during summer months. This can be related to the
 376 higher likelihood of desert dust or biomass burning events over the Iberian Peninsula in these
 377 months (e.g., Cachorro et al., 2008; Valenzuela et al., 2012), as was mentioned above. This is
 378 corroborated by the increase of the differences among the stations during the warm season, likely
 379 due to the variability in the impact of the desert dust episodes which strongly depend on the
 380 geographical location of each site. The higher occurrence of large aerosol loads during the warm
 381 seasons (see Figure 1), can explain the more negative ARE during summer and spring in Figure 7.
 382 For instance, the Barcelona station, with the largest values of AOD_{440nm} , is the bottom curve of each
 383 panel in Figure 7. Furthermore, the influence of mineral dust aerosol (with high aerosol optical
 384 depth) during these months also causes strong radiative effects, as was also reported by previous
 385 studies (e.g., Cachorro et al., 2008; Guerrero-Rascado et al., 2009; Antón et al., 2011; Román et al.,
 386 2013; García et al., 2014). In addition, the bimodality of the monthly AOD climatology mentioned
 387 in Section 4 has its impact on the radiative effects. The annual AOD cycle (see Figure 1, El
 388 Arenosillo site) causes the inverse monthly distribution of ARE with a first minimum in March.
 389 This effect is more clearly seen in ARE_{NIR} and ARE_{SW} .

390

391 **6. Aerosol radiative forcing efficiency in different spectral ranges**

392 The daily AFE values are calculated (following the methodology described in Section 3) in all the
 393 spectral ranges. AFE is a function of the aerosol optical properties, where both the aerosol particle
 394 size distribution and absorptive properties play a key role (e.g., Antón et al., 2011). As we assumed
 395 a fixed value of $SSA = 0.90$ in the simulations with $AOD_{440nm} < 0.15$ (see Table 2), the AFE is
 396 calculated only for those cases showing AOD_{440nm} larger than 0.15.

397 To identify the influence of SSA and α on AFE, this variable is calculated for several intervals of
 398 each aerosol property. Four categories of single scattering albedo at 675 nm are established in the
 399 calculation of the AFE: $1.0 \geq \text{SSA}_1 > 0.95$, $0.95 \geq \text{SSA}_2 > 0.90$, $0.90 \geq \text{SSA}_3 > 0.85$, and $0.85 \geq$
 400 $\text{SSA}_4 > 0.80$. Furthermore, aerosol size is classified in three intervals: $0 \leq \alpha_1 \leq 1$, $1 < \alpha_2 \leq 1.5$, and
 401 $1.5 < \alpha_3 \leq 2$. Note that two intervals in the range of α larger than 1 have been considered. One for
 402 median particles and another one for fine particles, because of the relevant importance of median
 403 size particle (continental or mixed aerosol aerosols types) over the Iberian Peninsula (see Figure 3).
 404 Although the general classification between fine and coarse particles requires a more refined
 405 classification (Schuster et al., 2006; Prats et al., 2011), the more general intervals selected in this
 406 study are adequate to perform a study of the aerosol sizes at the six stations together.

407 Figure 8 shows the AFE obtained for the UV (AFE_{UV}), VIS (AFE_{VIS}), NIR (AFE_{NIR}), and SW
 408 (AFE_{SW}) ranges for all these intervals. The threshold to evaluate the average in each sub-interval is
 409 fixed at 10 data points. From these figures it is seen that, the stronger the absorption by aerosols, the
 410 stronger their forcing efficiency. That is a decrease in the absolute values of the AFE for increasing
 411 SSA and for all particle size. In general, the groups of non-absorbing particles exhibit a good
 412 agreement among the six stations (see, for instance, AFE values in all the spectral ranges in the
 413 interval $1 < \alpha \leq 1.5$). Larger differences are obtained in the case of more absorbing aerosol particles.
 414 These can be understood because of the different types of aerosols presented over each site (see
 415 Section 4) and the different data numbers. The average AFE values over the whole Iberian
 416 Peninsula (considering the six stations together) are presented in Table 3 as a function of α and
 417 SSA, separately. The role played by the aerosol size on AFE values is different in the three sub-
 418 intervals of the shortwave radiation. AFE_{UV} and AFE_{VIS} are larger (in absolute value) for fine
 419 particles, while the opposite occurs in the case of AFE_{NIR} . As a result of these mixed effects, AFE_{SW}
 420 shows also a decrease in its values with increasing α , but this effect is weaker than for the visible
 421 and ultraviolet parts. SSA exhibits a more dominant role. As was observed before, the most

negative values are achieved for the most absorbing aerosols considered in this study (group 1 of SSA, see Table 3).

The average values of forcing efficiency obtained in this study (see Table 3) are in line with those found by other authors. Table 4 summarizes the results obtained by previous studies. It is difficult to assess some features in the comparison with previous reported AFE values, because of the different aerosol types, time periods and methods that are analyzed. Our study presents the evaluation of ARE with six long-term databases of aerosol properties. In spite of that, the values shown in Table 3 agree with those of Table 4, but the larger discrepancies are observed with the studies focusing on specific events. Our results match better with the results reported by, e.g., Zhou et al. (2005), Meloni et al. (2005), and Di Biagio et al. (2010). As was noticed by, e.g., Costa et al. (2004, 2006) and Di Biagio et al. (2010), AFE at the surface is larger (in absolute term) for aerosols characterized by smaller and absorbing particles. This result is corroborated by the findings shown in this study. Furthermore, as was pointed out by Di Biagio et al. (2010), the aerosol absorption is the dominant factor on AFE evaluated at the surface.

To evaluate the contribution of each spectral range with respect to the shortwave, the dependence of each AFE ratio (VIS to SW and NIR to SW) on SSA and α is shown in Figure 9. AFE_{VIS}/AFE_{SW} and AFE_{NIR}/AFE_{SW} ratios are shown in the figure since their contributions are the dominant. AFE_{UV}/AFE_{SW} ratio can be obtained as 100% minus the sum of the percentage of the two other ranges. As expected, non substantial differences are observed in the behavior of the six stations considered in this study. The NIR contribution becomes more decisive for large particles ($\alpha < 1$). It is expected that larger particles interact more with the longer wavelengths, while the smaller particles present more interaction with the shorter wavelengths. The presence of large particles with low SSA (high absorption) leads to a reduction of the AFE_{NIR}/AFE_{SW} ratio as well as an increase of the AFE_{VIS}/AFE_{SW} ratio. However, for non-absorbing (high SSA) large particles, the

446 AFE_{NIR}/AFE_{SW} ratio increases, and the contributions of the visible and infrared parts become more
447 similar (both around ~40-50%). The difference between AFE_{VIS}/AFE_{SW} and AFE_{NIR}/AFE_{SW}
448 increases for intermediate - fine particles. For these particles, the AFE_{VIS}/AFE_{SW} ratio does not
449 show a dependence on SSA. The smallest contribution of the NIR interval is around ~25% under
450 strong absorbing aerosols and fine particles, while AFE_{VIS}/AFE_{SW} is still over 60%. For this case,
451 the contribution of the ultraviolet range achieves a maximum of ~15%, being almost comparable
452 with the near infrared contribution. In summary, aerosol size determines the relevance of VIS-NIR
453 ranges, while SSA plays a key role, particularly, for large particles.

454

455 **7. Conclusions**

456 Six long-term datasets of aerosol properties over the Iberian Peninsula were analyzed and used as
457 input in a radiative transfer model to simulate ultraviolet, visible, near-infrared, and shortwave
458 radiation. The aerosol radiative effect (ARE) and aerosol forcing efficiency (AFE) were calculated.
459 The main conclusions are as follows:

460 1) The annual cycles of AOD and α values of atmospheric aerosols over the six analyzed stations
461 present high variability among them, emphasizing the inhomogeneity of the Iberian Peninsula,
462 mainly due to the different aerosol types over each station. The Barcelona site presents the largest
463 values of AOD, although Southern stations (Granada and El Arenosillo sites) frequently exhibit
464 daily values over 0.2 during summer months. The classification α -AOD has shown that continental
465 (mainly, clean) is the principal type of aerosol over the Iberian Peninsula. However, maritime
466 aerosols are also common in the Cabo da Roca, El Arenosillo and Évora sites. Desert dust events
467 are registered at the six sites, with the highest frequency at Granada and El Arenosillo, but the most
468 relevant feature is the South-North gradient of desert dust load which modulates the aerosol
469 climatology over the Iberian Peninsula.

2) The aerosol load over the Iberian Peninsula has shown a decrease trend between 2004 and 2012 (-0.04 per unit of AOD_{500nm} per decade, being statistically significant at the 95% of significance level). Yearly values of the AOD at 440 nm have also shown a statistically significant trend of -0.09 AOD_{440nm}-unit per decade at Barcelona site. The temporal trends for the rest of the stations are not statistically significant at the 95% significance level, but all of them are negative. Hence, a reduction of the aerosol column load over the Iberian Peninsula is observed in the last decade.

3) In the whole Iberian Peninsula, yearly ARE_{UV} ranges between -1.1 and -0.7 Wm⁻², ARE_{VIS} ranges between -5.7 and -3.6 Wm⁻², and ARE_{NIR} has values between -2.6 and -1.6 Wm⁻². As a result, ARE_{SW} is in the range between -8.8 and -5.7 Wm⁻². The temporal trends of ARE_{UV}, ARE_{VIS}, ARE_{NIR}, and ARE_{SW} exhibit positive statistically significant trends between 2004 and 2012. For instance, the trend rate for the ARE_{SW} is +3.6 Wm⁻² per decade (statistically significant at the 95% of significance level).

4) The intra-annual ARE cycle exhibits larger values during the spring and summer months when the likelihood of high aerosol loading over the Iberian Peninsula increases. In general, the annual AOD cycle is driven by the occurrence of Saharan dust events.

5) The AFE values at the six stations used in this study are in good agreement. Conditions of high α (small particles predominate) and low SSA (high absorption) lead to the largest negative AFE values. Overall, as an average for the Iberian Peninsula: AFE_{UV} = -6 Wm⁻² τ^{-1} , AFE_{VIS} = -34 Wm⁻² τ^{-1} , AFE_{NIR} = -19 Wm⁻² τ^{-1} , and AFE_{SW} = -59 Wm⁻² τ^{-1} .

6) The contribution of the ultraviolet, visible, and infrared to total shortwave aerosol forcing efficiency is governed by the aerosol type. In general, the visible part of the spectrum is the most dominant part. Non-absorbing large particles cause a more equal contribution of VIS and NIR intervals, while the UV range shows a higher contribution for absorbing fine particles.

493

494

495 **Appendix A**

496 The two choices in the performance of radiative transfer simulations from the libRadtran code
497 concerning aerosol properties are justified in this section.

498 First at all, as it is mentioned in the text, most of the data present $AOD_{440nm} < 0.15$ (~70% for
499 Palencia, Granada, and Évora sites). For these low values, $SSA = 0.9$ and $g = 0.75$ are selected by
500 the representativeness of the local aerosols in the six sites of study (e.g., Cachorro et al., 2010). To
501 analyze possible uncertainties emerging from this choice, the radiative net fluxes are also evaluated
502 for SSA and g values covering the most variety of aerosols observed in the Iberian Peninsula.
503 Hence, $SSA_1 = 0.8$, $SSA_2 = 1.0$, $g_1 = 0.65$, and $g_2 = 0.80$ are selected in this analysis. Four
504 possibilities or scenarios are simulated mixing the two values of the aerosol properties. The
505 radiation obtained in each scenario is compared with the assumed case of $SSA = 0.9$ and $g = 0.75$.
506 The two optical properties are also fixed as non-wavelength-dependent in this analysis. The
507 AOD_{440nm} used is 0.15, the worst scenario possible for these cases because the higher the AOD the
508 stronger the impact of aerosol properties. The simulations are performed for the four spectral
509 ranges. Table A1 shows the mean relative difference observed for the four scenarios and two
510 different SZAs (30° and 60°). The assumption considered in this study causes, in the worst possible
511 scenarios, errors in the ARE retrievals (obtained as the expanded errors from the radiative
512 uncertainty) $< 10\%$, $< 6\%$, $< 3\%$, and $< 5\%$ for the UV, VIS, NIR, and SW ranges, respectively. As
513 the cases with $AOD_{440nm} < 0.15$ are the large majority of the Iberian Peninsula, they should be
514 included in the study. The aerosol characterization, with respect to SSA or g , for these 'clean' cases
515 present large uncertainties itself, and no reliable information is extracted from them. Hence, the
516 results of this sensitivity study are adequate. As the SSA influences the diffuse radiation, the worst

517 results are obtained at large SZAs. The impact of g on the net fluxes is very weak. In conclusion,
 518 the choice of $SSA = 0.9$ and $g = 0.75$ in a clean scenario ($AOD_{440nm} < 0.15$) is proven as adequate
 519 because of two reasons: a) representativeness of the local aerosols which can be mixture of different
 520 types, and b) the low uncertainty produced in the simulations by SSA and g under these conditions.

521 The choice of fixed SSA and g values within each of the spectral ranges (UV, VIS, and NIR)
 522 represented by the CIMEL spectral measurements is also justified here. The aerosol models by
 523 Shettle (1989) included in the libRadtran code (see Mayer and Kylling, 2005) are used to evaluate
 524 the uncertainty of using this approximation. The continental clean aerosols (most common type in
 525 the Iberian Peninsula, see Figure 3), and continental polluted aerosols (also very common, which
 526 present an extreme case of absorption) are tested in this analysis. The simulations are performed for
 527 the expected spectral behavior of SSA and g following Shettle (1989) and the case of fixed
 528 properties in the UV (SSA and g values at 440 nm), VIS (SSA and g values at 675 nm), and NIR
 529 intervals (SSA and g average of values at 870 and 1020 nm). Figure A1 presents the evolution of
 530 the relative error (considering as reference the net flux with the expected spectral dependence of
 531 aerosol properties) for several AOD values between aerosol-free and $AOD_{550nm} = 0.6$. In the case of
 532 continental clean aerosols (Figure A1.a), the error of using our assumption is lower than 5% for all
 533 SZAs and spectral ranges. Therefore, as the large majority of aerosol particles are of this type, the
 534 methodology used and proposed in this study only introduces a relative error below 5% in the
 535 majority of the simulations. With respect to the continental polluted aerosols (Figure A1.b), the
 536 error increases achieving a maximum around 20% for the UV range and very turbid conditions. For
 537 large AOD conditions in the Iberian Peninsula (e.g., $AOD_{550nm} = 0.4$) but with low frequency of
 538 occurrence in contrast to $AOD_{440nm} < 0.15$, the error of the SW range is below 5%. However, the
 539 UV range is more sensitive to our method and the error is around 15% at $SZA = 60^\circ$. As it was
 540 mentioned above, the errors are larger for large SZAs because of the possible interaction between
 541 absorption and scattering processes resulting the diffuse radiation. The visible range is more

sensitive to the spectral variations than the NIR interval, which exhibits a maximum error around 11% at $\text{SZA} = 60^\circ$ and $\text{AOD}_{550\text{nm}} = 0.6$. The daily net radiative fluxes are also evaluated for the two aerosol types in order to quantify the uncertainty in the final simulated data used in this study. For Palencia site (and the corresponding SZA evolution), a daily value for the June 20th is simulated assuming $\text{TOC} = 300 \text{ DU}$ and $\text{PWC} = 1 \text{ cm}$. The results for the continental polluted case with $\text{AOD}_{440\text{nm}} = 0.4$ exhibit differences between the spectral and fixed-band aerosol properties of: 7.5%, 5.3%, 4.0%, and 4.8% for the UV, VIS, NIR, and SW intervals. The relative errors for the same intervals with continental clean (and same AOD value) are: 1.9%, 1.2%, 1.4%, and 1.4%, respectively. Therefore, the uncertainty due to fixed optical properties in each spectral range is dependent on the aerosol type but the error caused can be considered as acceptable. Since actual aerosols often present mixtures of different types, the uncertainty of using the theoretical spectral evolution for one type (given by an aerosol model) can also produce uncertainties which should be taken into account. Although other aerosol types are not tested in this analysis, a similar behavior can be expected. For instance, for the case of desert dust aerosols, Román et al. (2013) found a slight influence of spectral aerosol absorption properties on UV irradiance analyzing a strong Saharan intrusion over Granada site.

Therefore, the two assumptions performed in this study in the simulations are adequate for the evaluation of net fluxes and aerosol radiative effects. The uncertainties that can be introduced in the daily values are acceptable being around or smaller than 5% for the net SW radiation. This uncertainty is usually achieved in clear-sky modeling (e.g., Mateos et al., 2013a).

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580

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851 **Table 1.** Coordinates and time interval of the six AERONET sites used in this study.

Station	Latitude (°N)	Longitude (°E)	Altitude a.s.l. (m)	Time interval
Palencia	41.99	-4.52	750	2003-2012
Barcelona	41.39	2.12	125	2004-2012
Cabo da Roca	38.78	-9.50	140	2003-2011
Évora	38.57	-7.91	293	2005-2012
Granada	37.16	-3.61	680	2004-2012
El Arenosillo	37.11	-6.73	0	2000-2009

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856 **Table 2.** Summary of AERONET data used for ARE calculations: aerosol optical depth (AOD),
857 single scattering albedo (SSA), asymmetry factor (g), precipitable water vapor column (PWC).
858 Estimated absolute uncertainty of AOD and SSA is given according to Dubovik et al. (2002), and
859 PWC error from Holben et al. (1998).

	AERONET database	Estimated uncertainty
AOD	Level 2.0	± 0.01 -0.02
SSA, g (AOD ₄₄₀ >0.4)	Level 2.0	± 0.03 (in SSA)
SSA, g (0.15<AOD ₄₄₀ <0.4)	Level 1.5-filtered*	± 0.05 -0.07(in SSA)
SSA, g (AOD ₄₄₀ <0.15)	Fixed value	
PWC	Level 2.0	10-15%

860 *Filters applied are the same as in level 2.0 except for AOD₄₄₀ (see text).
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Table 3. AFE values and their standard error for the UV, VIS, NIR, and SW ranges for, separately,
four SSA and three α intervals over the Iberian Peninsula. Units are $\text{Wm}^{-2}\tau^{-1}$. SSA groups: $0.85 \geq$
SSA₁ > 0.80 (group 1), $0.90 \geq$ SSA₂ > 0.85 (group 2), $0.95 \geq$ SSA₃ > 0.90 (group 3), and $1.0 \geq$
SSA₄ > 0.95 (group 4); and α groups: $0 \leq \alpha_1 \leq 1$ (group 1), $1.0 \leq \alpha_2 \leq 1.5$ (group 2), and $1.5 < \alpha_3 \leq 2$
(group 3). The average values without any classification are also presented.

Variable	Group	AFE _{UV}	AFE _{VIS}	AFE _{NIR}	AFE _{SW}
α	1	-5.41 ± 0.06	-30.1 ± 0.3	-20.9 ± 0.2	-56.5 ± 0.5
	2	-6.60 ± 0.09	-38.3 ± 0.4	-19.1 ± 0.2	-64.0 ± 0.6
	3	-7.06 ± 0.10	-39.4 ± 0.4	-16.9 ± 0.2	-63.3 ± 0.7
SSA	1	-9.7 ± 0.2	-52.8 ± 0.8	-24.9 ± 0.5	-87.4 ± 1.4
	2	-8.19 ± 0.10	-44.6 ± 0.4	-21.2 ± 0.2	-74.0 ± 0.6
	3	-6.37 ± 0.05	-35.9 ± 0.2	-19.5 ± 0.2	-61.8 ± 0.3
	4	-4.59 ± 0.05	-26.6 ± 0.2	-18.1 ± 0.2	-49.3 ± 0.3
Average		-5.98 ± 0.05	-33.7 ± 0.2	-19.34 ± 0.11	-59.1 ± 0.3

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873 **Table 4.** Daily Forcing Efficiencies at the surface by previous studies. Legend: desert dust (DD),
874 continental-biomass burning (C-BB), and forest fires (FF).
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Reference	Aerosol Type	AFE _x	Value (Wm ⁻² τ ⁻¹)	Time period	Region	More info.
Díaz et al. (2007)	Mixed	AFE _{UV}	-3	July 2002	Spain	290-363 nm
Meloni et al. (2005)	DD Mixed	AFE _{VIS}	-28.4 -45.6	July 2002	Central Mediterranean	
Lyamani et al. (2006)	Mixed	AFE _{VIS}	-75.8	August 2003	Spain	2003 heat wave
Di Biagio et al. (2010)	DD C-BB Mixed	AFE _{SW}	-68.9 -59.0 -94.9	2004-2007	Central Mediterranean	At the equinox
Esteve et al. (2014)	Mixed	AFE _{SW}	-139.0	2003-2011	Spain	200 cloud-free days
Santos et al. (2008)	FF	AFE _{SW}	-113.0	2004-2005	Portugal	Absorbing aerosols
di Sarra et al. (2011)	DD	AFE _{SW}	-55	25-26/03/2010	Central Mediterranean	Strong event
García et al. (2014)	DD	AFE _{SW}	-59	2009-2012	Canary Islands	386 cloud-free days
Costa et al. (2006)	DD	AFE _{SW}	-116.9	7/04/2000	Korea	SSA = 0.76
Zhou et al. (2005)	DD	AFE _{SW}	-80/-48	Monthly aerosol climatology	North Africa and Arabian Peninsula	Depending on surface albedo
Saha et al. (2008)	C-BB Mixed	AFE _{SW}	-97.6 -81.5	2005-2006	French Mediterranean	0.7 < SSA < 0.8
Valenzuela et al. (2012)	DD	AFE _{SW}	-70	2005-2010	Spain	

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880 **Table A1.** Mean relative difference (RD) in the UV, VIS, NIR, and SW net fluxes if SSA = 0.90
881 and g = 0.75 are compared with different SSA and g scenarios for different SZA values.
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SZA	RD _{UV} (%)	RD _{VIS} (%)	RD _{NIR} (%)	RD _{SW} (%)
30	±3.4	±1.9	±0.9	±1.5
60	±4.9	±3.0	±1.5	±2.4

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885 Figure Captions

886 **Figure 1.** Annual cycle of daily values of AOD at 440 nm by box whisker plots. Triangles and
887 horizontal solid lines indicate the monthly average and median values, respectively.

888 **Figure 2.** Annual cycle of daily values of α ('alpha' in the figure) by box whisker plots. Triangles
889 and horizontal solid lines indicate the monthly average and median values, respectively.

890 **Figure 3.** Relative frequency of aerosol type occurrence: maritime (MA), desert dust (DD),
891 continental clean (CC), and continental polluted (CP).

892 **Figure 4.** Yearly values of AOD_{440nm} at the six sites: Barcelona (blue diamonds), Palencia (purple
893 triangles), Évora (red squares), Cabo da Roca (grey crosses), Granada (black stars), and El
894 Arenosillo (green circles). The text points out the statistically significant trend obtained.

895 **Figure 5.** Evolution of yearly ARE_{UV} (a), ARE_{VIS} (b), ARE_{NIR} (c), and ARE_{SW} (d) at the six sites:
896 Barcelona (blue diamonds), Palencia (purple triangles), Évora (red squares), Cabo da Roca (grey
897 crosses), Granada (black stars), and El Arenosillo (green circles). Vertical bars indicate the standard
898 deviation of each yearly value at Palencia station.

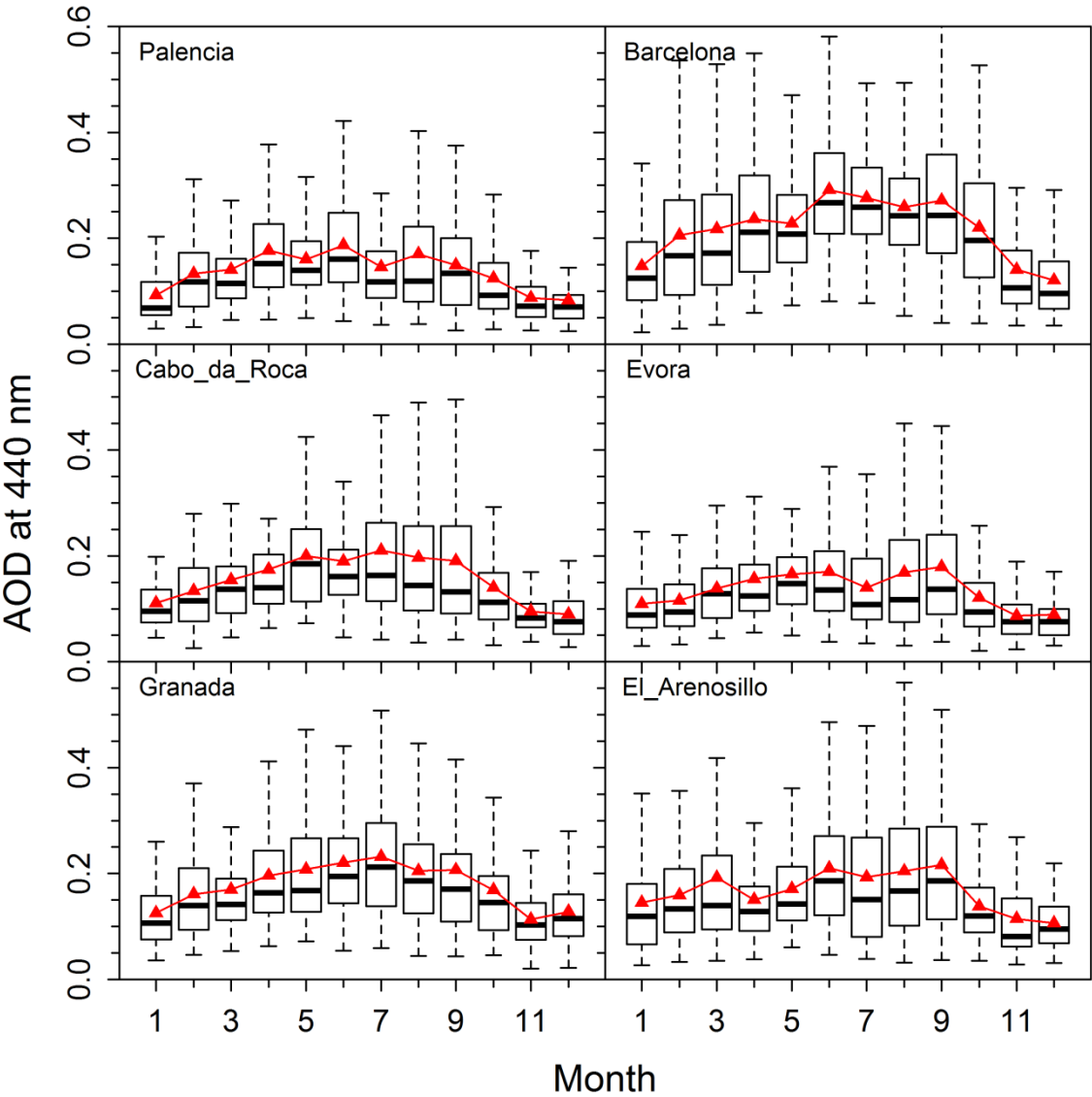
899 **Figure 6.** Evolution of annual ARE at the four spectral ranges (ARE_{UV} purple diamonds, ARE_{VIS}
900 red squares, ARE_{NIR} green triangles, and ARE_{SW} black circles) and AOD at 500 nm (blue stars)
901 averaging the data from the six Iberian ground-based sites (only years with at least three sites
902 considered). Dashed lines point out the linear trends (see text).

903 **Figure 7.** Annual cycle of ARE_{UV} (a), ARE_{VIS} (b), ARE_{NIR} (c), and ARE_{SW} (d) at the six sites:
904 Barcelona (blue diamonds), Palencia (purple triangles), Évora (red squares), Cabo da Roca (grey
905 crosses), Granada (black stars), and El Arenosillo (green circles). Vertical bars point out the
906 standard deviation of each monthly value at Évora station.

907 **Figure 8.** AFE_{UV} , AFE_{VIS} , AFE_{NIR} , and AFE_{SW} against four groups of aerosol single scattering
908 albedo and three intervals of α at the six sites: Barcelona (blue diamonds), Palencia (purple
909 triangles), Évora (red squares), Cabo da Roca (grey crosses), Granada (black stars), and El
910 Arenosillo (green circles).

911 **Figure 9.** Dependence of AFE_{VIS}/AFE_{SW} (a, c, e) and AFE_{NIR}/AFE_{SW} (b, d, f) ratios on SSA for
912 large (a, b), medium (c, d) and small (e, f) particles at the six sites: Barcelona (blue diamonds),
913 Palencia (purple triangles), Évora (red squares), Cabo da Roca (grey crosses), Granada (black stars),
914 and El Arenosillo (green circles).

915

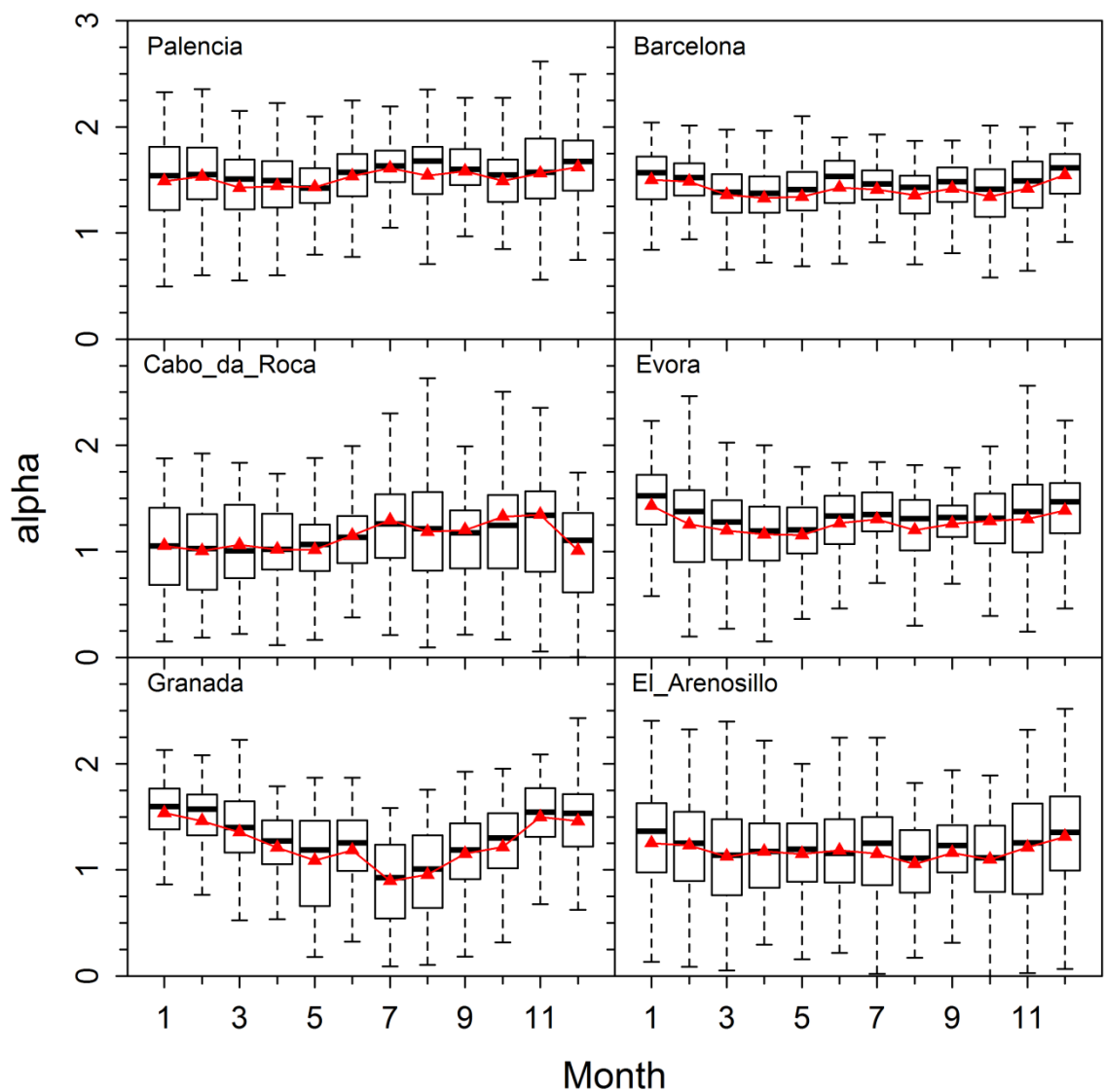


917

918 **Figure 1.** Annual cycle of daily values of AOD at 440 nm by box whisker plots. Triangles and
919 horizontal solid lines indicate the monthly average and median values, respectively.

920

921 Figure 2



922

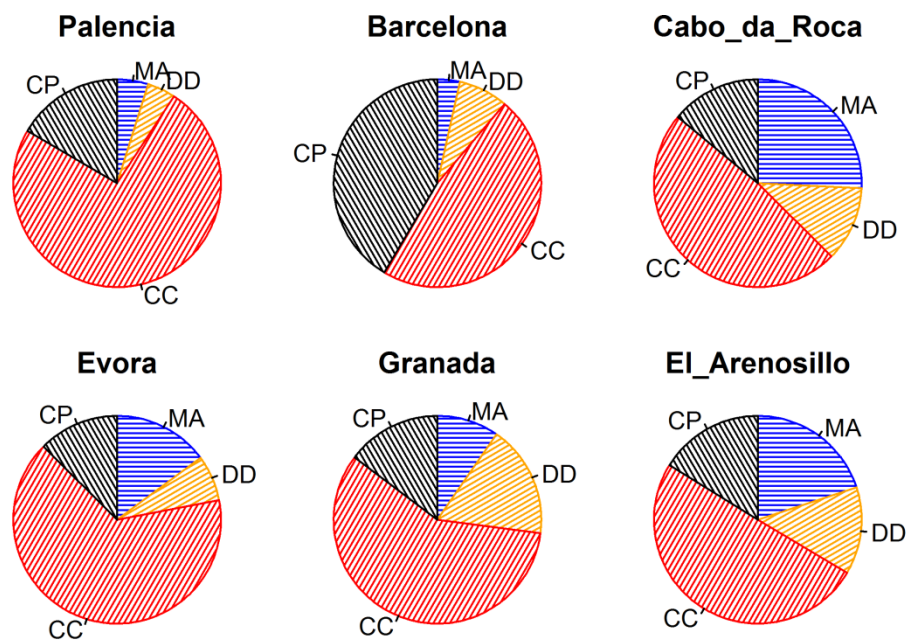
923 **Figure 2.** Annual cycle of daily values of α ('alpha' in the figure) by box whisker plots. Triangles
924 and horizontal solid lines indicate the monthly average and median values, respectively.

925

926

927 Figure 3

928



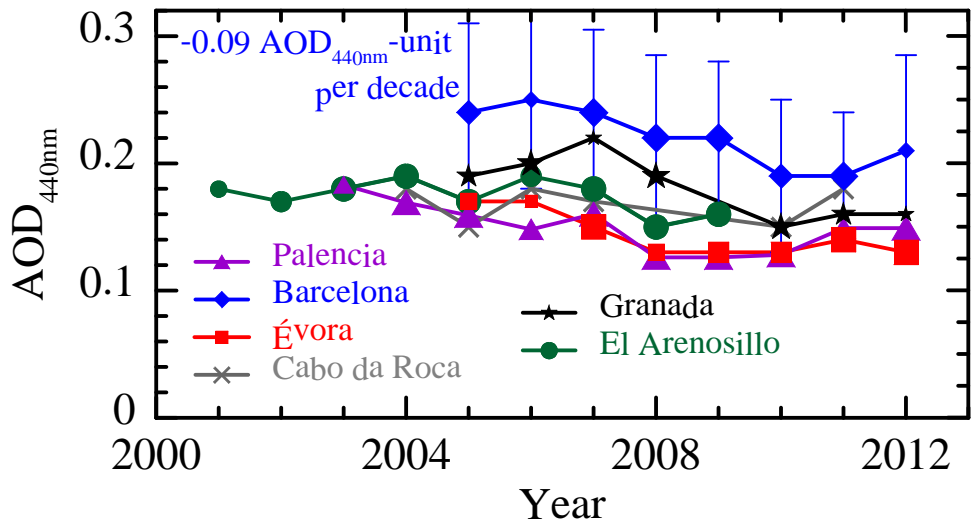
929

930 **Figure 3.** Relative frequency of aerosol type occurrence: maritime (MA), desert dust (DD),
931 continental clean (CC), and continental polluted (CP).

932

933 Figure 4

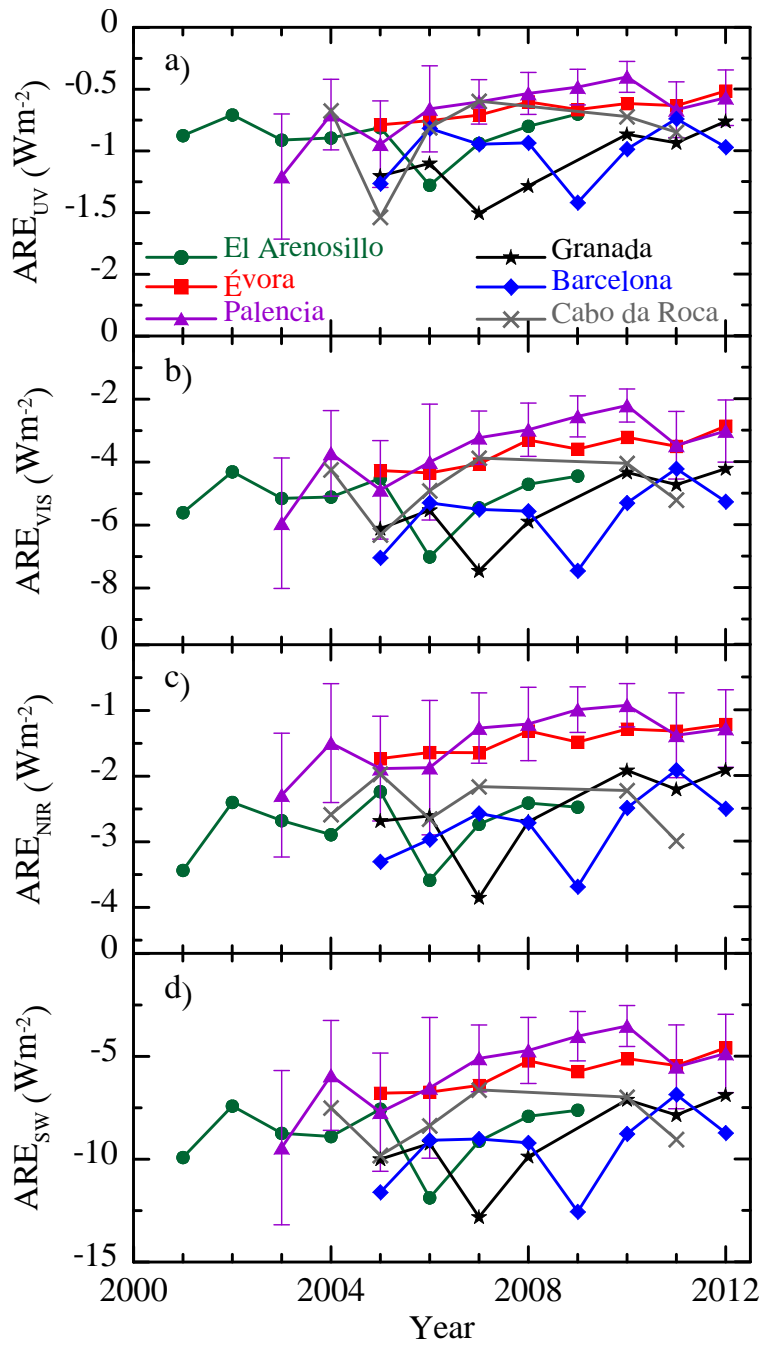
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935

936 **Figure 4.** Yearly values of AOD_{440nm} at the six sites: Barcelona (blue diamonds), Palencia (purple
937 triangles), Évora (red squares), Cabo da Roca (grey crosses), Granada (black stars), and El
938 Arenosillo (green circles). The text points out the statistically significant trend obtained. Vertical
939 bars indicate the standard deviation of each yearly value at Barcelona station. Larger the symbols,
940 large amount of data number that year.

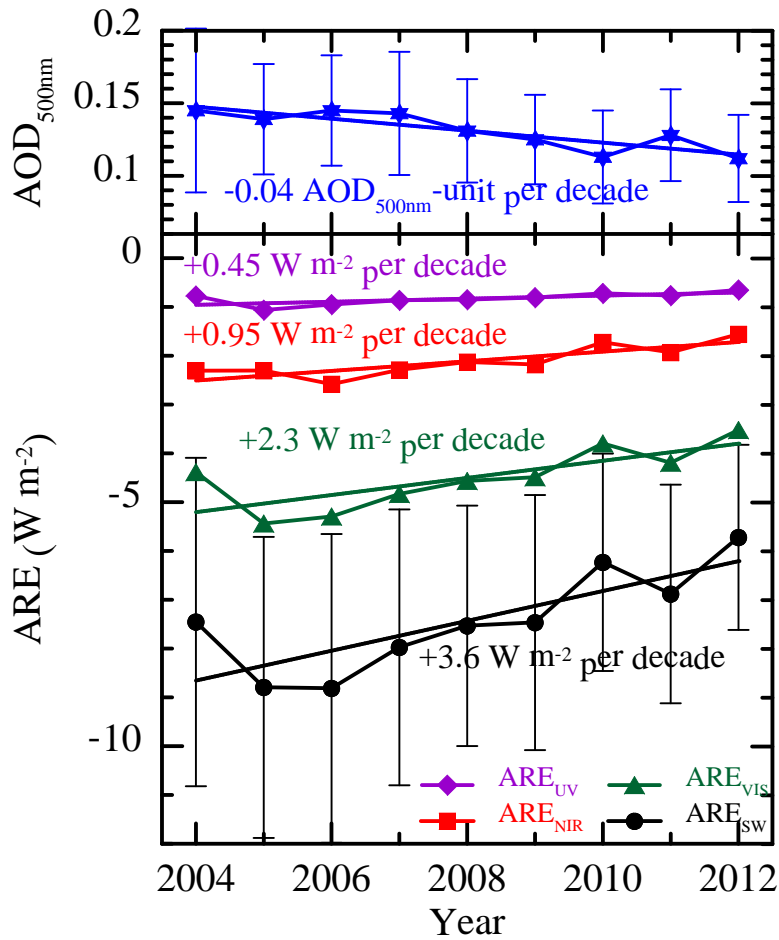
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943

944 **Figure 5.** Evolution of yearly ARE_{UV} (a), ARE_{VIS} (b), ARE_{NIR} (c), and ARE_{SW} (d) at the six sites:
945 Barcelona (blue diamonds), Palencia (purple triangles), Évora (red squares), Cabo da Roca (grey
946 crosses), Granada (black stars), and El Arenosillo (green circles). Vertical bars indicate the standard
947 deviation of each yearly value at Palencia station.

948 Figure 6
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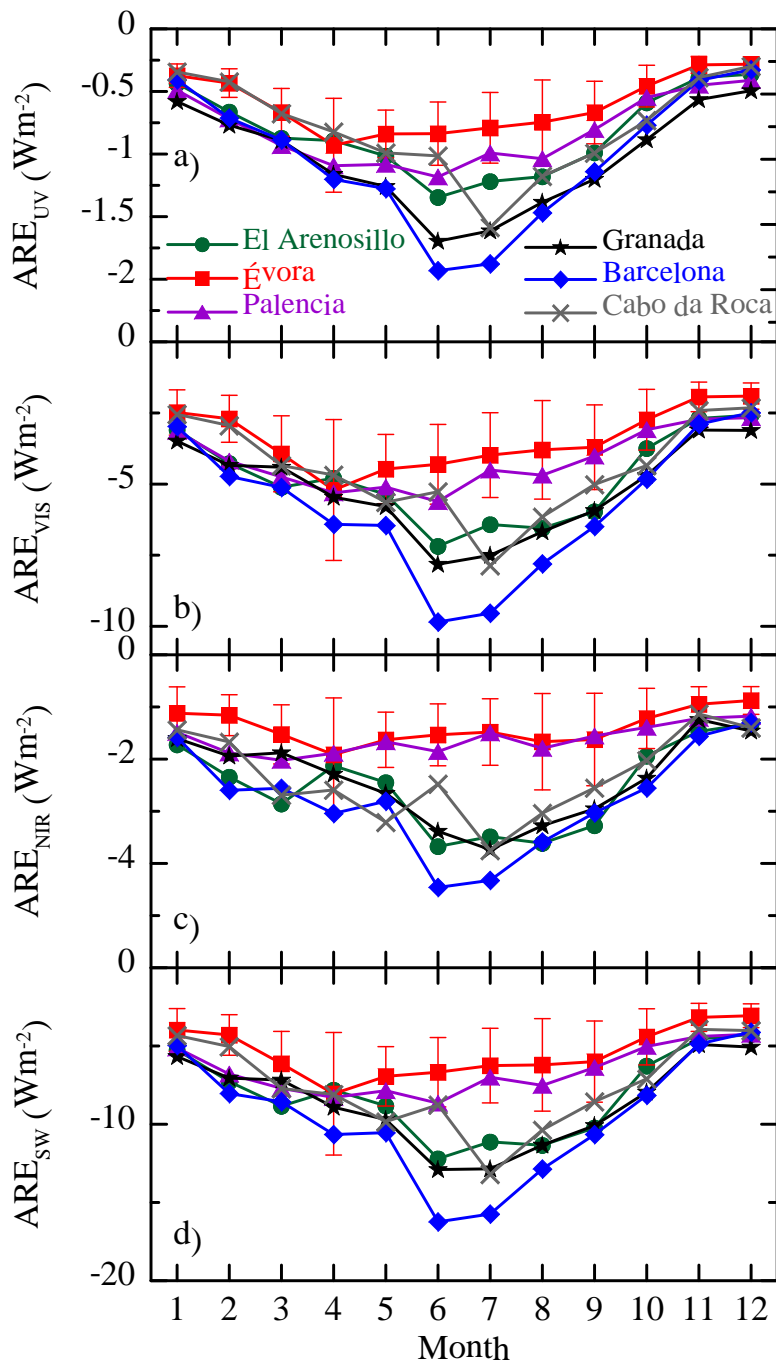


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952

953 **Figure 6.** Evolution of annual ARE at the four spectral ranges (ARE_{UV} purple diamonds, ARE_{VIS}
954 red squares, ARE_{NIR} green triangles, and ARE_{SW} black circles) and AOD at 500 nm (blue stars)
955 averaging the data from the six Iberian ground-based sites (only years with, at least, three sites are
956 considered). Dashed lines point out the linear trends (see text). Vertical bars indicate the standard
957 deviation.

958

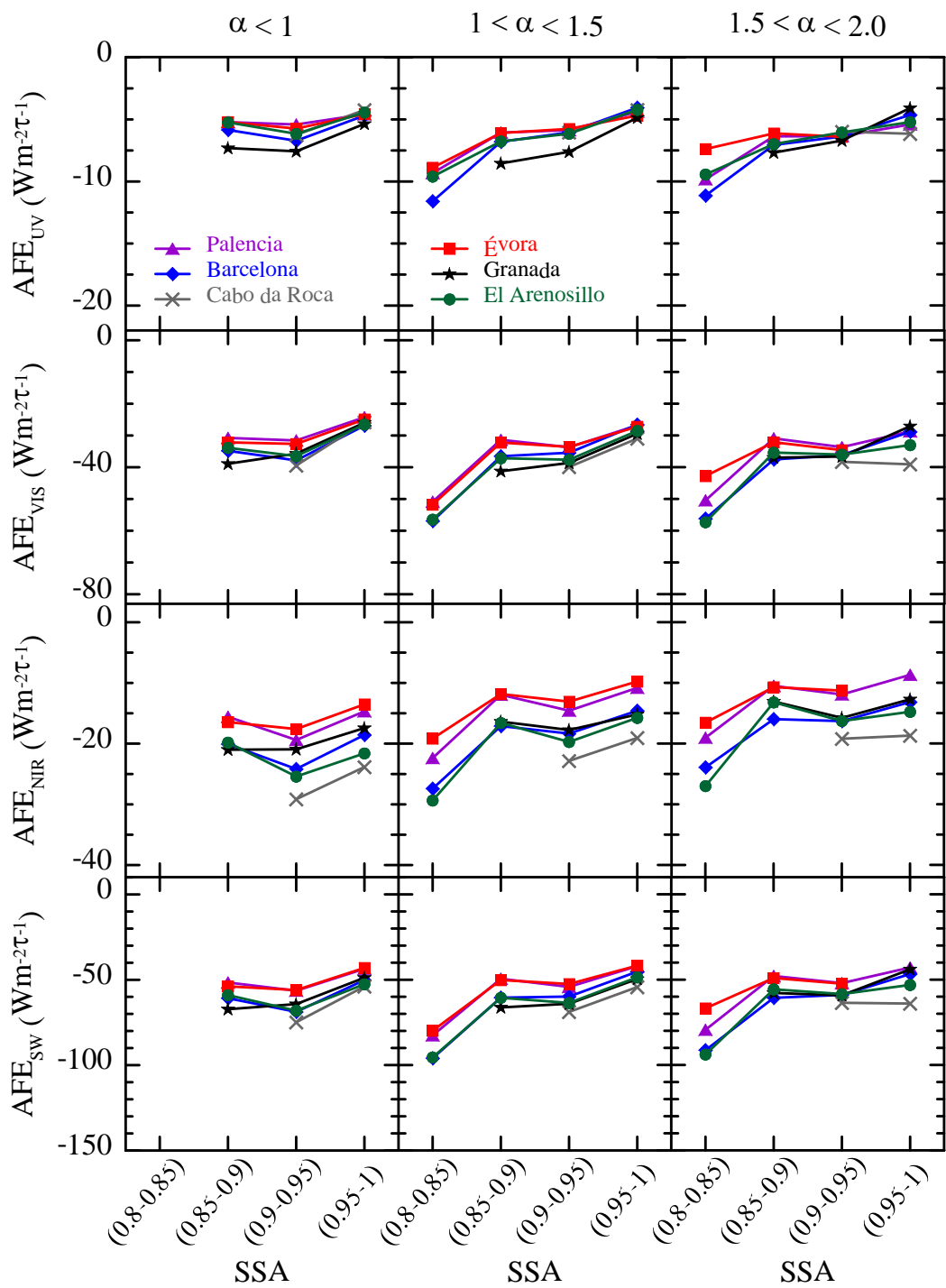
959 Figure 7
960



961

962 **Figure 7.** Annual cycle of ARE_{UV} (a), ARE_{VIS} (b), ARE_{NIR} (c), and ARE_{SW} (d) at the six sites:
963 Barcelona (blue diamonds), Palencia (purple triangles), Évora (red squares), Cabo da Roca (grey
964 crosses), Granada (black stars), and El Arenosillo (green circles). Vertical bars point out the
965 standard deviation of each monthly value at Évora station.

966 Figure 8

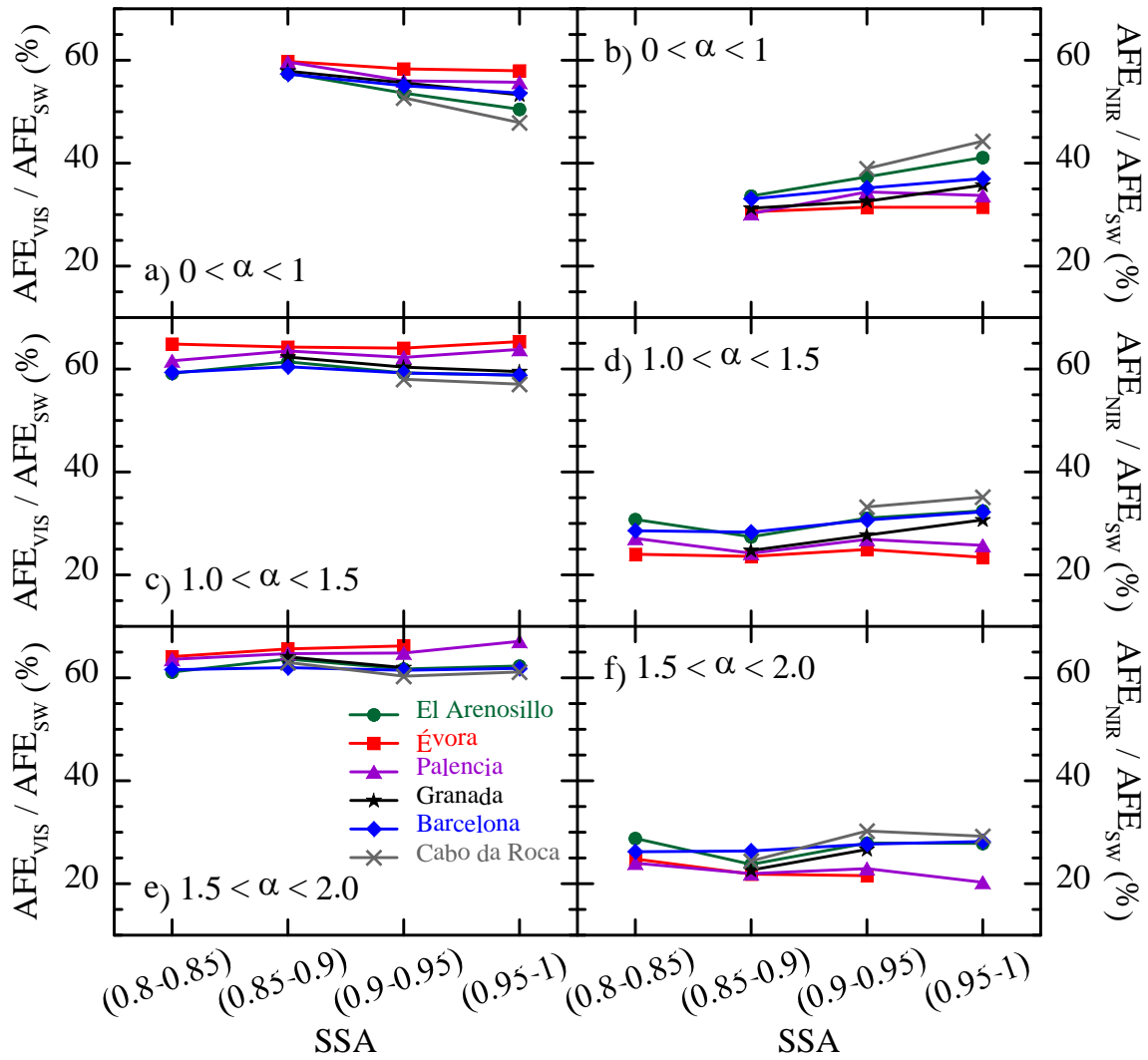


967

968 **Figure 8.** AFE_{UV} , AFE_{VIS} , AFE_{NIR} , and AFE_{SW} against four groups of aerosol single scattering
969 albedo and three intervals of α at the six sites: Barcelona (blue diamonds), Palencia (purple
970 triangles), Évora (red squares), Cabo da Roca (grey crosses), Granada (black stars), and El
971 Arenosillo (green circles).

972 Figure 9

973



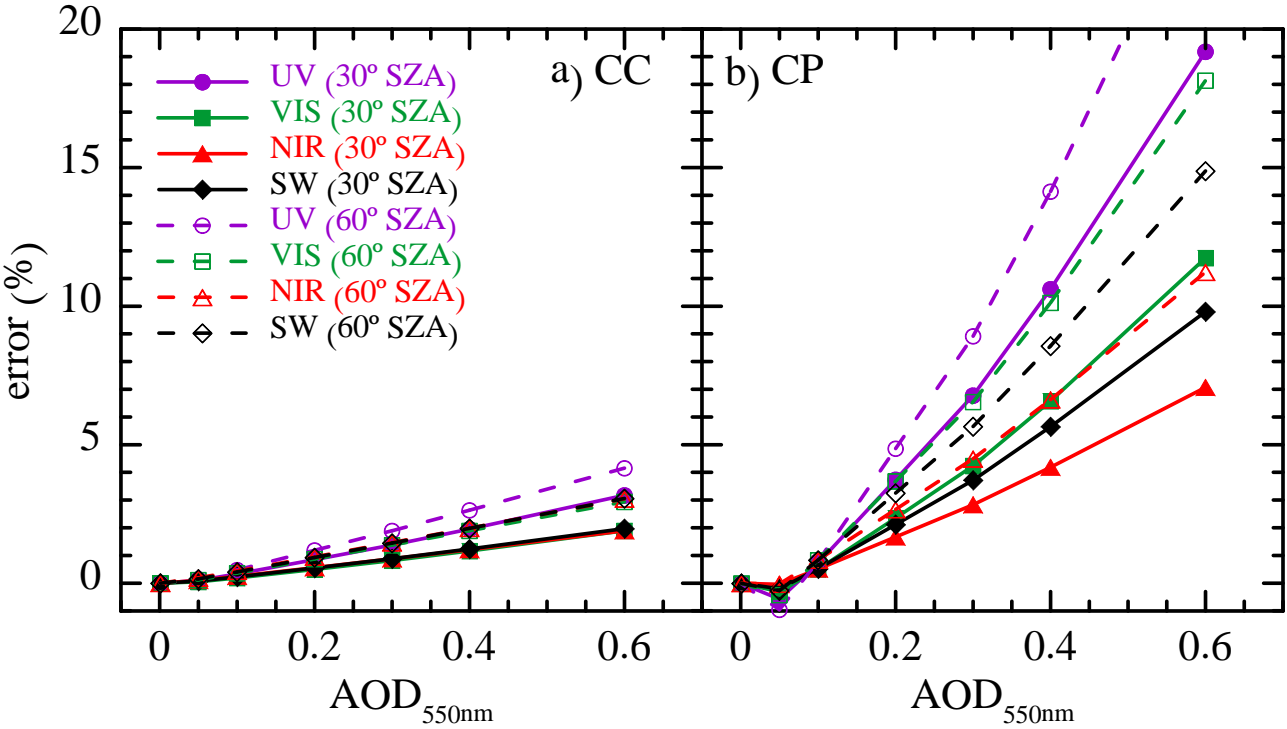
974

975 **Figure 9.** Dependence of AFE_{VIS}/AFE_{SW} (a, c, e) and AFE_{NIR}/AFE_{SW} (b, d, f) ratios on SSA for
976 large (a, b), medium (c, d) and small (e, f) particles at the six sites: Barcelona (blue diamonds),
977 Palencia (purple triangles), Évora (red squares), Cabo da Roca (grey crosses), Granada (black stars),
978 and El Arenosillo (green circles).

979

980 Figure A1

981



982

983 Figure A1. Evolution of the error committed when the optical properties are fixed in the different
984 spectral range for two SZAs, and continental clean (CC, a) and continental polluted (CP, b)
985 aerosols.

986