

**Sulphate aerosols
from Nabro volcano**

M. J. M. Penning de Vries
et al.

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Characterisation of a stratospheric sulphate plume from the Nabro volcano using a combination of passive satellite measurements in nadir and limb geometry

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Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Abstract

The eruption of the Nabro volcano (Eritrea), which started on 12 June 2011, caused the introduction of large quantities of SO₂ into the lower stratosphere. The subsequently formed sulphate aerosols could be detected for several months following the eruption. It is generally assumed that the formation of sulphate aerosols in the stratosphere takes about a month, but in plumes from explosive eruptions significant amounts of aerosols have been seen to form within a few hours.

We show that sulphate aerosols were present in the lower stratosphere within hours of the onset of the eruption of Nabro. Evidence comes from nadir UV Aerosol Index (UVAI) and SO₂ measurements by SCIAMACHY, GOME-2 and OMI, and limb aerosol measurements by SCIAMACHY. The sulphate plume displays negative UVAI in the western part of OMI's swath and positive UVAI in the eastern part – an effect that is due to the strong viewing angle dependence of UVAI and can only be caused by a high-altitude (> 11 km), non-absorbing (single-scattering albedo > 0.97) aerosol plume. For the retrieval of the aerosol profile from limb measurements, the horizontal dimensions and the position of the aerosol plume need to be taken into account, otherwise both extinction and layer height may be underestimated appreciably. By combining nadir SO₂ column density and UVAI with limb aerosol profiles, a stratospheric plume from Nabro could be tracked from 13 to 17 June, before the plumes from later, lower-altitude explosions started interfering with the signal. Our findings are in agreement with ground-based lidar and sun-photometer data from an MPLNET/AERONET station in Israel and with data from the satellite-borne CALIOP lidar.

1 Introduction

Volcanic eruptions are an important source of sulphate aerosols to the stratosphere. Once the aerosols reach the stratosphere, low deposition rates cause lifetimes on the order of months to years (e.g. Bluth et al., 1997, and references therein). The purely

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Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



(Rose et al., 2003; Guo et al., 2004; Bitar et al., 2010). Nabro was reported to emit large quantities of water vapour (Smithsonian Institution, 2011), and we show here that the first volcanic plume from Nabro, emitted on 12–13 June 2011, contained an amount of sulphate aerosols large enough to be visible from satellite within at most 36 h of the eruption. This is in agreement with a study by (Sawamura et al., 2012), who report the presence of an aerosol plume in the lower stratosphere at 17 km altitude using data from a ground-based lidar station in Sede Boker, Israel, approximately 30 h downwind of the first explosion of Nabro.

For the investigation of the first plume from the Nabro eruption we use the satellite-borne spectrometers SCIAMACHY, GOME-2, and OMI. The primary goal of these instruments is the measurement of various trace gases in the atmosphere, which is why spatial resolution is sacrificed in favour of higher spectral resolution – a dedicated aerosol instrument like MODIS has a 300x (for OMI) to 3000x (GOME-2) higher spatial resolution. With these large ground pixels, cloud contamination is a serious issue for aerosol retrieval from such spectrometers (Krijger et al., 2007). To avoid this problem, we make use of UV Aerosol Indices (UVAI), which can be determined in the presence of clouds (Torres et al., 1998; de Graaf et al., 2005; Penning de Vries and Wagner, 2011). Despite the fact that the quantitative application of UVAI is not straightforward, it can be used to distinguish between UV-absorbing aerosols and aerosols that do not or barely absorb UV radiation (Penning de Vries et al., 2009), especially for high-altitude aerosol plumes. This makes the UVAI a tool particularly suited for the study of volcanic plumes, which may contain either (absorbing) ash or (non-absorbing) sulphate aerosols, or both. It has been shown in the past that UVAI depend on solar and viewing geometry (de Graaf et al., 2005), but we show that the effect can be particularly dramatic for high-altitude plumes, and is even capable of changing the sign of UVAI caused by purely scattering aerosol layers from negative to positive.

Once aerosols reach the upper troposphere or lower stratosphere, e.g. in the course of a volcanic eruption, their profile can be studied using space-based measurements in limb geometry. The SCIAMACHY instrument (2002–2012) made consecutive scans

in limb and nadir geometry, allowing the study of trace gas profiles together with their co-located total column amount (Bovensmann et al., 1999). First successful attempts at aerosol retrieval from SCIAMACHY limb measurements have only recently been published (Ovigneur et al., 2011; Taha et al., 2011; Ernst et al., 2012). We present results from a new limb aerosol retrieval approach and demonstrate the limitations of conventional limb retrievals when faced with horizontally narrow plumes such as those emitted by volcanoes and wildfires.

2 Methods and data

2.1 Instrumentation

The SCanning Imaging Absorption spectroMeter for Atmospheric CHartography (SCIAMACHY) was launched on ENVISAT in March 2002, into a daytime-descending polar orbit located at approximately 800 km from the Earth's surface (Bovensmann et al., 1999; Wagner et al., 2008; Gottwald and Bovensmann, 2011). SCIAMACHY's swath is 960 km wide, and global coverage is achieved approximately every 6 days. The instrument measures backscattered radiance in the UV-NIR range (240–2380 nm). One feature that makes SCIAMACHY unique is the alternating measurements in the limb and nadir modes, which allow co-located measurements of the total column of atmospheric constituents (in nadir geometry) and their stratospheric profile (in limb geometry). The alternation of measurements in limb and nadir geometry is the reason for the division of the orbit into subsets of measurements called states. Whereas nadir ground pixels generally measure 30 km × 60 km, limb pixels are approximately 240 km wide and have a vertical resolution of about 3 km (Bovensmann et al., 1999; Gottwald and Bovensmann, 2011). Since April 2012, all communications with ENVISAT have been lost, which put an end to several successful and long-lasting missions on the platform – SCIAMACHY among them.

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



GOME-2 (the second Global Ozone Monitoring Experiment) on METOP-A is very similar to SCIAMACHY, although it measures in a smaller spectral range (240–790 nm) and lacks the possibility of limb measurements (Callies et al., 2000). The spatial resolution is comparable (40 km × 80 km), but because GOME-2's swath is twice as wide as SCIAMACHY's, and due to the fact that it measures in nadir geometry only, GOME-2 achieves global coverage in 1.5 days. Once per month, the instrument operates in the narrow-swath mode, reducing the pixel width (and swath width) by a factor of 8, approximately 40 km × 10 km: the same resolution as the polarisation measurement devices (PMDs). The main function of the PMDs is to provide input for the polarisation correction of measured radiances, but they can also be used for the detection of clouds (Grzegorski et al., 2009) or the determination of UVAI (de Graaf et al., 2013).

The OMI (Ozone Monitoring Instrument) on the AURA satellite has a much higher spatial resolution: away from the swath edges, the pixels measure 13 km × 24 km. OMI has been performing measurements in the UV-Visible wavelength range from its daytime ascending polar orbit since 2004 (Levelt et al., 2006). Its wide swath of 2600 km allowed daily global coverage until the first occurrence of the so-called row anomaly in June 2007, an instrumental problem that causes grievous radiance errors in up to half of OMI's ground pixels (van Hoek and Claas, 2010). Currently, neither SO₂ columns (from our retrieval), nor UVAI can be reliably determined from pixels affected by the row anomaly which we therefore removed from the data set prior to analysis (although different criteria were used to filter SO₂ and UVAI data sets). In the future, many more OMI pixels might be eligible for SO₂ retrieval – see, e.g. the study by Yan et al. (2012).

2.2 UV aerosol indices

To investigate if ash was present in the volcanic plume we use the UVAI (also known as “residue”). The positive part of UVAI, known as AAI (Absorbing Aerosol Index) is a well-known and much-used qualitative indicator of aerosols that absorb radiation in the UV range, such as mineral dust, biomass burning smoke, and volcanic ash (Herman et al., 1997; Torres et al., 1998; de Graaf et al., 2005). The AAI depends on plume altitude

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



and is very sensitive to elevated layers of absorbing aerosols, in particular if these are located over bright surfaces such as ice or clouds (see, e.g. Fromm et al., 2010). Its counterpart, the SCattering Index (SCI), was only recently introduced as an indicator for aerosols that do not or only weakly absorb UV radiation (Penning de Vries et al., 2009). A big advantage of the UVAI is that they can be meaningfully interpreted in the presence of clouds; however, clouds influence UVAI and must be taken into account for quantitative analysis of UVAI (Penning de Vries and Wagner, 2011). In contrast to the AAI, the SCI does not depend on plume altitude (Torres et al., 1998; de Graaf et al., 2005; Penning de Vries et al., 2009), but only on the size distribution and amount of aerosols (assuming purely scattering particles). In addition, viewing and solar geometry play an important role.

There are several different definitions of UVAI in existence; here we use KNMI's AAI product for SCIAMACHY, which is determined using 340 nm and 380 nm radiances (available at <http://www.temis.nl/>). For consistency, OMI UVAI data obtained from the same wavelengths were used (obtained from <http://mirador.gsfc.nasa.gov/>). GOME-2 UVAI are available from radiances measured using two of the instrument's PMD channels centred at 338 nm and 382 nm (de Graaf et al., 2013). Not only do the PMD UVAI have the advantage of an approximately 8x higher spatial resolution, the UVAI also appears less noisy than the science channel product, which probably has to do with the better co-location of the radiance measurements at the two wavelengths required for the UVAI calculation. The correlation between UVAI determined from GOME-2's science channels on the one hand, and the PMD UVAI on the other is excellent, although we found that an offset of 0.5 units had to be subtracted – possibly to account for differences in polarisation sensitivity and calibration.

2.3 Limb extinction profiles

For the determination of the extinction profile of the volcanic aerosol plume we applied a newly developed retrieval algorithm to SCIAMACHY limb intensity data measured at 750 nm. Because the intensity at a certain tangent height (TH) is not only influenced by

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



particles at that particular altitude, but also by particles in the layers above, an onion peeling method is applied. That means that the retrieval is applied to the intensity measured at each TH in sequence, starting at the top (TH = 31.5 km), taking into account the retrieval results from the higher layers. The intensity at TH = 34.5 km is used to normalize the data. For the presented study of a volcanic plume, the onion peeling method is iterated at least twice to ensure the effects of particles in lower layers (which influence the intensity at the THs above if the layer is optically dense) are properly accounted for. The retrieval at each TH is done by performing radiative transfer calculations in which the extinction is varied iteratively until the simulated intensity matches the measurement. The simulations are performed using the full spherical 3-D Monte Carlo model McArtim (Deutschmann et al., 2011) with an effective surface albedo from the matching nadir measurements and aerosol parameters representative of sulphate particles (single-scattering albedo 1.0, $g = 0.6$, corresponding to a size distribution with a mean diameter of 80 nm and width = 1.6) as input. The SCIAMACHY limb retrieval method will be described in more detail in a forthcoming publication.

The findings from the extinction retrieval are supported by results from the Colour Index method. Colour Indices (CIs) have been used in the past to estimate particle abundance in the stratosphere (Sarkissian et al., 1991; Enell et al., 1999; von Savigny et al., 2005). The ratio of the intensity (I) at two wavelengths (e.g. $\lambda_1 = 750$ nm and $\lambda_2 = 1090$ nm) at a certain tangent height θ is a measure of the particle loading of the atmosphere which is based on the difference between the effect of weakly wavelength-dependent scattering on particles and that of strongly wavelength-dependent molecular scattering. The CI R_c is thus defined as:

$$R_c(\theta) = \frac{I(\lambda_2, \theta)}{I(\lambda_1, \theta)} \quad (1)$$

Scenes containing aerosols and/or clouds show higher colour index values than those with background conditions.

2.4 SO₂ data

As one of the most abundant gases in a volcanic plume, SO₂ is often used as a tracer for volcanic emissions by a variety of spectroscopic remote sensing techniques. It can be very well measured from space using DOAS (Differential Optical Absorption Spectroscopy, see e.g. Platt and Stutz, 2008; Richter and Wagner, 2011) due to the strong characteristic absorption features in the UV spectral range, both from the ground (e.g., Galle et al., 2003; Bobrowski and Platt, 2007) and from space (e.g. Eisinger and Burrows, 1998; Khokhar et al., 2005). The DOAS method is usually only applied to weak absorbers with optical densities below 0.05, in which case the effective light path can be assumed to be wavelength-independent within the considered range of analysis. Standard DOAS evaluation schemes take advantage of the relatively high sensitivity to SO₂ within the wavelength range of 310–325 nm, but in the course of a major volcanic eruption, the SO₂ optical density can easily exceed 0.2 within that range. In such cases, the effective light path becomes dependent on wavelength. As a consequence, the SO₂ vertical column density is strongly underestimated by standard fit procedures in the UV range (Yang et al., 2007; Bobrowski et al., 2010; Hörmann et al., 2013). To account for the non-linear radiative transfer effects of the DOAS retrieval during the SO₂-rich eruption of Nabro, a combined, multi-wavelength fit-window approach was adopted for the retrieval of SO₂ by SCIAMACHY, GOME-2, and OMI as described in (Hörmann et al., 2013); satellite measurements clearly affected by strong SO₂ absorptions in the standard wavelength range of 312.1–324 nm (i.e., surpassing a threshold SO₂ slant column density of 1×10^{18} molec cm^{-2} , or about 37 DU) were re-analysed using an alternative wavelength range of 326.5–335.3 nm. As the volcanic SO₂ plume was located at high altitudes, vertical column densities were calculated by using a geometrical airmass factor.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



3 The June 2011 eruption of Nabro

The Nabro volcano had been presumed extinct before the sudden, explosive eruption that started in the late evening of 12 June 2011 (Smithsonian Institution, 2011). An ice cloud was seen to emerge from Nabro at 21:00 UTC by SEVIRI (Vernier et al., 2013) and an hour later by MODIS-AQUA (Fromm et al., 2013). Nabro's eruption, which was monitored by many different satellite instruments, was exceptional with regard to the total amount of emitted SO₂: 4–5 Tg were estimated to have been released within the first 15 days (Theys et al., 2012), the most active period of the eruption. This compares to approximately 20 Tg SO₂ from Pinatubo (Bluth et al., 1993), the largest eruption of the past 100 years. The total stratospheric aerosol optical depth, resulting from the conversion of volcanic SO₂ to sulphate aerosols, was the highest measured by the space-based limb-sounding OSIRIS instrument since its launch in 2001 (Bourassa et al., 2012). Whereas at least one – the first – of the numerous explosive eruptions was observed to inject gases and particles into the stratosphere (as presented here, but also in Vernier et al., 2013; Fromm et al., 2013; Clarisse et al., 2013; Fairlie et al., 2013), most of the emissions occurred at much lower altitudes. We here focus on the first plume from the eruption of Nabro because it is clearly separated from other plumes; the continuous volcanic emission of SO₂ during the following days complicates the identification of individual plumes. The first plume was initially detected by GOME-2 and IASI at 06:25 UTC on 13 June and contained approximately 0.5 Tg SO₂ (L. Clarisse and N. Theys, personal communication, 2013), thereby contributing roughly 10% to the total SO₂ mass.

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



4 Results

4.1 Nadir satellite data: SO₂ columns and UV Aerosol Indices

Figure 1 shows nadir data from four different satellite instruments (GOME-2, SCIAMACHY, OMI, and MODIS) taken on 13 June, shortly after the onset of the eruption. Assuming that Nabro started emitting SO₂ at 21:00 UTC on 12 June (Vernier et al., 2013), the plume's leading edge is 11 and 14 h old at the time of the GOME-2 and OMI overpasses, respectively. The SO₂ plume initially stretches from the volcano (pink triangle in the lower right of each figure) to the Northwest and fans out slightly, as shown in Fig. 1 for GOME-2, SCIAMACHY, and OMI (panels a–c). Panels d–f present the corresponding UVAI; please note that the colour scale of the UVAI was adapted to show non-absorbing aerosols (SCI, or UVAI < 0 in yellow-red tones) in addition to UV-absorbing particles (AAI, or UVAI > 0 in blue).

The spatial agreement between SO₂ and the negative UVAI plume is very good for OMI, but is less clear in the GOME-2 PMD and SCIAMACHY data. Only the later (western) GOME-2 orbit shows a negative UVAI signal coinciding with the SO₂ plume, and although UVAI values < 0 are seen over the full width of the SCIAMACHY swath, the eastern part is probably also affected by clouds (see below). For a more intuitive view of the situation, in panels g and h the true-colour images from the MODIS instruments on TERRA and AQUA, respectively, are given. The time difference between the two sets of instruments is very small; there are 10 min between SCIAMACHY and MODIS-TERRA overpasses, 75 and 25 min between the two GOME-2 orbits and the MODIS-TERRA overpass, and between OMI and MODIS-AQUA measurements lie 15 min. We assume that the plume has not moved or changed appreciably within such a short time frame.

Most of the particle extinction observed in the MODIS images from 13 June (Fig. 1g and h) is most probably due to the presence of ice particles that fall out rapidly; on the following day, the plume can no longer be seen in MODIS true-colour images (Fig. 2g and h) and neither is a cloud detected by the IR retrieval (not shown here, but see: http://ladsweb.nascom.nasa.gov/browse_images/granules.html). Evidence of the presence

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



of large (i.e. cloud) particles within the 13 June volcanic plume also comes from MODIS IR cloud retrieval (Fromm et al., 2013) and from SEVIRI measurements (Vernier et al., 2013). The possible effects of an ice cloud on the UVAI data measured on 13 June will be discussed in Sect. 5.1.

5 The SO₂ and UVAI data from GOME-2, SCIAMACHY and OMI measured on 14 June are shown in Fig. 2. Only a small part of the SO₂ plume was captured by GOME-2 (panel a), because the instrument measured in narrow-swath mode at the time. No sign of sulphate aerosols can be seen in the accompanying UVAI plot (d). SCIAMACHY caught the tail end of the volcanic plume (panels b and e): SO₂ columns of up to 30 DU
10 were measured and a slightly negative UVAI feature can be seen at the corresponding position in panel e. In the OMI SO₂ data displayed in panel c not only the first volcanic plume can be seen (over Northern Egypt and Israel), but also elevated SO₂ columns near the volcano; these were emitted later at lower altitudes and are ignored here. The northern SO₂ plume measured by OMI corresponds to an area with clearly negative
15 UVAI values. In panel f the volcanic plume is difficult to separate from the clouds directly to the north. More importantly, the figure does not correctly display all OMI measurements: parts of the two consecutive orbits overlap in the figure, in particular those parts where the sulphate plume is seen. Figure 3 shows both orbits separately, allowing us to focus on an interesting UVAI phenomenon: in the earlier orbit (36 772;
20 panels a, c, and e) the high-altitude volcanic plume is viewed from the East, causing it to exhibit negative UVAI values. During the following overpass (orbit 36 773; panels b, d, and f), only 100 min later, the viewing direction of the instrument is exactly opposite and positive UVAI are found for the same plume. This is an effect of the phase functions of scattering on air molecules and aerosol particles that can be well reproduced
25 with RTM simulations using McArtim (as detailed in the Supplement). The dependence on viewing geometry is most pronounced for aerosol (or cloud) layers at high altitude (> 11 km) and for line-of-sight zenith angles (at the sub-satellite point) larger than about 45°, corresponding to the swath edges of OMI. As the clouds directly north of the volcanic plume somewhat muddy the picture, we selected only those pixels that have SO₂

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



columns larger than 1.5 DU for panels c and d of Fig. 3. Despite the fact that not all selected plume pixels contain a pronounced aerosol signal, the viewing angle effect can be clearly seen. We performed RTM simulations using the aerosol parameters representative of volcanic sulphate particles: $g = 0.6$, single-scattering albedo of 1.0 (also used in the limb retrieval) and with an Ångström exponent of 1.5. We modelled the UVAI from the sulphate plume by assigning an AOT (at 380 nm) of 0.4 to those pixels with SO_2 column densities exceeding 30 DU, an AOT of 0.2 for pixels with SO_2 column densities between 15 and 30 DU, and an AOT of 0.1 to those with SO_2 columns smaller than 15 DU. The results of this exercise are shown in panels e and f of Fig. 3. Although the measurements in panels c and d are not in perfect agreement with the model results (which may be due to cloud contamination in orbit 36 722, or inaccuracies in the assumed aerosol parameters), the viewing angle effect is very well reproduced. This leads to two conclusions: (1) the volcanic plume must be at high altitude (> 11 km) and hence, (2) the aerosols must have a high single-scattering albedo (≥ 0.97), or they would not have shown negative UVAI at any viewing angle (see the Supplement for more evidence from RTM calculations).

4.2 SCIAMACHY Limb data: colour index and extinction profiles

Figure 4 shows the centres of SCIAMACHY limb pixels (tangent points) in the region around the Nabro volcano on 13 June (panel a) and 14 June 2011 (b). Each dot is colour-coded by the CI at a TH of 18.6 km, which is clearly above the tropopause (at 16.9 km, according to ECMWF). The only significant enhancements ($\text{CI} > 0.25$) in this region are found very close to the SO_2 plume (shown in grey tones) and indicate the presence of an elevated layer of particles.

The extinction profile retrieval was applied to the two SCIAMACHY states marked by pink boxes in Fig. 4. Whereas the four limb pixels within the selected state from 13 June all detect the aerosol plume ($\text{CI} > 0.25$), the tangent points of the first two (western) limb pixels on 14 June lie outside of the volcanic plume and can be regarded as background measurements. The extinction profiles for all selected pixels are found

**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



in Fig. 5. To avoid cluttered plots, the retrieved extinction is depicted as a point at a certain altitude, whereas in reality, the value represents the extinction for a 3 km high box centred at the respective TH. The profiles for the two scenes outside of the volcanic plume (pixels 1 and 2 in panel b, and marked as background 1 and 2 in panel a) are very similar, reaching a maximum of $1 \times 10^{-4} \text{ km}^{-1}$, corresponding to the stratospheric background aerosol layer (SPARC, 2006). All non-background pixels in panel a (13 June, orbit 48 555, state 20), on the other hand, show a pronounced peak from 19.5 km down to 13.5 km, with a maximum extinction of $4.4 \times 10^{-3} \text{ km}^{-1}$ at 13.5 km. In comparison, pixels 3 and 4 in panel b (14 June, orbit 48 570, state 19) show a smaller extinction maximum ($< 1.4 \times 10^{-3} \text{ km}^{-1}$) and a profile that is shifted to higher altitude, with a pronounced peak at 16.5 km.

The retrieval leading to the results shown in Fig. 5 was performed under the assumption of a homogeneous aerosol layer of 3 km thickness, spreading out infinitely in the horizontal plane through the tangent point (TP). This assumption is usually justified for stratospheric background aerosols, but not for the volcanic plume studied here, as illustrated in Fig. 6. Not only are the plume dimensions (indicated by the grey shaded nadir SO_2 data) finite; the plume is also not located symmetrically around the TPs (approximate positions marked by yellow numbers). As a result, both the retrieved extinction and the layer altitude are affected. The effects of observing an aerosol plume (or cloud) with finite dimensions can be intuitively understood by studying the lines-of-sight of the limb instrument at various THs (solid black lines in Fig. 7). The infinite aerosol layer assumed in the retrieval is depicted in green in panel a. If the retrieval algorithm is applied to a finite aerosol layer (panel b), the extinction is underestimated, because the retrieval “expects” the plume to have an infinite horizontal extent. If, in addition, the layer is positioned off-centre with respect to the TP (panel c), both the extinction and the altitude of the layer are underestimated: the aerosol layer is mainly located within the line-of-sight at 16.5 km, therefore the retrieval algorithm wrongly places the extinction at that altitude. The magnitude of the underestimation of aerosol layer altitude depends on the distance to the TP. Taking the limb lines-of-sight into account leads to the TH-

dependent sensitivity areas depicted in Fig. 6 for observations with TH = 19.5 km. If an aerosol plume at 18–21 km of finite width would be located only within the blue boxes, the retrieval algorithm would (correctly) assign the extinction to 19.5 km. If, on the other hand, the plume were confined to the green or even red regions, an altitude of 16.5 km or 13.5 km would be retrieved. Similar displacements are found for observations at other THs. Note that the retrieval algorithm can lead to a significant underestimation of aerosol layer altitude, but never to an overestimation.

In view of the above, the fact that the volcanic plume is located somewhat off-centre on 13 June (particularly for pixel 4) may be the explanation for the West-to-East (pixel 1–4) decrease in extinction observed on 13 June – although the decrease may also be physical. The exceptionally high extinction retrieved for 13 June (in comparison to extinctions determined for 14–17 June) almost certainly results from the ice cloud, mentioned in Sect. 4.1, of which the top was at an altitude of 19.5 km, according to SEVIRI data (Vernier et al., 2013). The ice cloud may be the reason that no well-defined aerosol layer can be seen in panel a of Fig. 5, but rather a profile of which the extinction increases towards the surface. In contrast, in panel b (14 June), the volcanic plume bottom is clearly seen in addition to the plume top. The reason that both the extinction and the layer altitude reach higher values in pixel 4 than in pixel 3 is probably that the TP of limb pixel 4 falls exactly within the volcanic plume, whereas the TP of pixel 3 lies to the North of the plume. However, the possibility that a smaller amount of aerosols in the field of view leads to a smaller extinction within pixel 3 cannot be excluded, either.

4.3 Transport and evolution of the volcanic plume

The first volcanic plume from Nabro could be distinctively seen in GOME-2, SCIAMACHY, and OMI data until 17 June, when overlap with the plumes from various later explosions no longer allowed a clear attribution of the emission time (Fig. 8). Although there is an apparent merging of SO₂ plumes on 15 June (best seen in panel a between 40–70° E) due to overlap in the horizontal plane, the difference in plume altitude and

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



A-band) algorithm (Wang et al., 2008). On 16 June, the stratospheric plume fills the width of both SCIAMACHY state 16 (orbit 48 598) and state 17 (orbit 48 597). In the former (panel c), extinction maxima of $7\text{--}9 \times 10^{-3} \text{ km}^{-1}$ are found at 16.5 km. The plume is located further from the TPs of pixels 1 and 4, which translates in a downward shift of the retrieved profiles relative to those of pixels 2 and 3. Another explanation could be that the plume detected in pixels 2 and 3 reaches a higher altitude, but this seems less plausible – particularly because the nadir SO_2 data support the first explanation (Fig. 8, panel e). In the latter (panel d), the plume is clearly off-centre, again causing an underestimation of plume altitude (as seen when comparing to pixels 2–3 in panel c) and extinction. The increased extinction found at lower altitude in pixel 4 may originate from the tropospheric plume, glimpsed in the South-eastern corner of the state. The extinction profiles from state 17, orbit 48 611, measured on 17 June (panel e), look very similar to those shown in panel c: in all pixels strongly enhanced aerosol extinction with maxima of $0.7\text{--}1.0 \times 10^{-3} \text{ km}^{-1}$ at 16.5 km are found; in pixels 3 and 4, where the highest nadir SO_2 columns are found, aerosol extinction exceeds $0.6 \times 10^{-3} \text{ km}^{-1}$ even at 19.5 km. The variability of retrieved aerosol extinction at the lowest altitude of 13.5 km may be due to the underlying clouds. FRESCO+ detected clouds at an altitude of about 4 km in the nadir pixels corresponding to limb pixels 1, 3 and 4 and clouds at 7–8 km in limb pixel 2. Interpreting the results from the aerosol retrieval from state 17, orbit 48 610 (panel f) is not straightforward, in particular because we can no longer be guided by the SO_2 measurements due to the overlap of the first plume from Nabro and later emissions. The stratospheric plume is still clearly present, as seen by the enhanced particle extinction at 16.5 km, but underlying clouds and (possibly) lower-lying volcanic aerosol layers cause the retrieval of high extinctions down to 13.5 km.

5 Discussion

We have presented evidence that the eruption of Nabro on 12/13 June 2011 caused the injection of sulphate aerosols into the stratosphere. Using a new SCIAMACHY limb

retrieval algorithm, the extinction profile could be determined for parts of the volcanic plume on every day from 13–17 June. The aerosol plume was simultaneously detected in nadir by UVAI from SCIAMACHY and (with a small time difference) from GOME-2 and OMI on 13 and 14 June. We discuss the presented results in the following sections.

5.1 Sulphate aerosols

The major part of stratospheric SO₂ from Nabro was probably converted to sulphate on the time scale of weeks (Bourassa et al., 2012), yet we showed that an amount large enough to be detected from satellites was formed within hours of the eruption. Using the UVAI, similar results were found for volcanic eruptions in the past, e.g. Hekla in 2000 (Rose et al., 2003), Soufrière Hills and Nyamuragira in 2006, Okmok and Kasatochi in 2008, and Sarychev Peak in 2009 (unpublished results). Rapid sulphate formation was also observed in the young eruption plume of Pinatubo in 1991 (Guo et al., 2004), indicating that in this sense, the Nabro eruption was not unique.

Although there were reports of ash fallout near the volcano (Smithsonian Institution, 2011), no evidence of ash particles within the volcanic plume is seen either in the true-colour images or in the UVAI plots. UVAI are very sensitive to absorbing particles, particularly if they are elevated, hence even a relatively small amount of ash would cause UVAI to increase to near zero, or even positive values. Positive UVAI values (blue tones) are seen in most of the satellite pixels not covered by clouds or the SO₂ plume and could, in principle, be caused by volcanic ash. But the presence of elevated dust, clearly seen over the Red Sea in panels g and h of Fig. 1, makes this a very uncertain attribution.

The effect of the ice cloud on UVAI is potentially substantial, but is difficult to quantify due to the complications regarding the modelling of ice clouds. Radiative transfer model (RTM) calculations have shown that cloud droplets cannot cause UVAI values much lower than -1 (ice, Rose et al., 2003) or -1.5 (liquid cloud droplets, Penning de Vries and Wagner, 2011). Nevertheless, in (Penning de Vries and Wagner, 2011) UVAI values $\ll -1.5$ were regularly found in cloudy scenes, which could not be explained

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



using RTM calculations. In fact, the same phenomenon is observed in Fig. 1 (panels d–f), where strongly negative UVAI are found over clouded regions on the northern edge of the figure and directly south of the volcanic plume, and in Fig. 2 (panels e and f), where clouds cause negative UVAI to the north of the plume. Given the present data set and knowledge, it is not possible to say if the ice cloud merely enhances the UVAI signal from the sulphate plume or if the strong negative UVAI signals are caused by the ice cloud itself. Yet, a negative UVAI signal corresponding to the SO₂ plume can be discerned in SCIAMACHY and OMI UVAI data on 14 June, albeit much weaker, as shown in Fig. 2 (panels e and f), whereas the ice cloud has disappeared. Apart from the limb data, the most clear evidence that sulphate aerosols were detected by the UVAI comes from OMI, where the strong viewing angle dependence of UVAI indicates the presence of a high-altitude aerosol plume with very high single-scattering albedo (> 0.97). The strong viewing angle dependence complicates the interpretation of UVAI, but could be turned into an advantage with regard to the determination of aerosol properties. For elevated aerosol plumes large enough to cover the entire OMI swath, or plumes captured twice by OMI from opposing viewing angles (as in the case of the Nabro plume on 14 June) the viewing angle dependence can be exploited to determine the aerosol absorption and plume altitude. This approach can be applied to volcanic plumes, as shown here, but potentially also to extensive smoke or elevated pollution plumes.

5.2 Extinction profiles

One of the most challenging aspects of limb retrievals is determining where (the major part of) the detected signal comes from. For an optically thin atmosphere the tangent point is a good approximation, but the situation is more complex when an optically thick layer (e.g., a volcanic plume) affects the pathways of photons through the atmosphere. As shown in Figs. 6 and 7, the extinction and altitude of an aerosol layer can be severely underestimated (though not overestimated) if the spatial dimensions and the placement of the layer are not properly accounted for. Tests with model data point at a linear relationship between relative error in retrieved extinction and the “area” that the aerosol

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



layer covers within the total line-of-sight (indicated in green in Fig. 7). The retrieved altitude can be underestimated by several kilometers depending on the position of the aerosol layer with respect to the TP, as demonstrated in Figs. 6 and 7. Similar effects are expected for the colour index (ratio) and should be kept in mind when aerosol plumes or clouds (e.g. polar stratospheric clouds) of finite dimensions are the object of studies involving limb data.

During its overpass over the MPLNET (MicroPulse Lidar NETwork, Welton et al., 2001) station in Sede Boker, Israel, the AOT of Nabro's first volcanic plume was determined to be 0.17 by lidar measurements at 532 nm. When we integrate the profiles shown in Fig. 9, much smaller values of AOT are found, e.g. $4\text{--}5 \times 10^{-3}$ for the profiles from orbit 48 598 (state 16) on 16 June. The difference can be attributed to a number of effects: First, judging by the shape and column density of the SO₂ plume, the measured part of the plume (on 16 June) is not very dense in comparison with the middle of the plume, where the MPLNET AOT was measured. This is confirmed by CALIOP measurements from the orbit shown in Fig. 8, where an AOT of 0.03 is found for an aerosol layer at 16–17 km altitude (at 532 nm; assuming small particles, this corresponds to about 0.02 at 750 nm). Second, the finite dimensions of the plume cause the extinction from the limb retrieval to be on the order of a factor of 2 too low. Third, the retrieved extinction depends on the aerosol optical parameters assumed in the RTM calculations. Taking into account retrieval errors of CALIOP data (the uncertainty in extinction at 523 nm is given as 0.02), the agreement between SCIAMACHY limb measurements and satellite-borne lidar measurements is well within the uncertainty range.

5.3 Concluding remarks

Using nadir SO₂ and UVAI measurements from SCIAMACHY, GOME-2 and OMI, and limb measurements from SCIAMACHY we were able to detect aerosols in the stratosphere within the first volcanic plume from Nabro in the five days (13–17 June 2011) following the eruption. The bulk of the aerosol layer was found at 17–20 km by the limb retrieval; detailed analysis of nadir UVAI from OMI provided evidence for the non-

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



absorbing nature of the particles (single-scattering albedo > 0.97). We found that the viewing angle dependence of UVAI from high-altitude plumes can be large enough to change the sign of UVAI, which complicates the interpretation of UVAI, but also provides an opportunity to determine aerosol properties within extensive aerosol plumes.

In addition, we have shown that if the dimensions and the position of an aerosol plume (with respect to the tangent point) are not properly taken into account into the limb aerosol retrieval algorithm, both the retrieved extinction and layer height may be affected appreciably. In conclusion, the combination of nadir trace gas and aerosol data with profile information from limb measurements is well suited to study the characteristics of a high-altitude volcanic plume. Although SCIAMACHY is now lost, this type of analysis (at least for volcanic plumes) could be continued using an instrument like the Ozone Mapping and Profiling Suite instrument (OMPS) on NPP, which similarly measures SO₂ and UVAI in nadir and has limb-viewing capabilities.

Supplementary material related to this article is available online at
**[http://www.atmos-chem-phys-discuss.net/14/7739/2014/
acpd-14-7739-2014-supplement.pdf](http://www.atmos-chem-phys-discuss.net/14/7739/2014/acpd-14-7739-2014-supplement.pdf)**

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Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries
et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[⏪](#)[⏩](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

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Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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**Sulphate aerosols
from Nabro volcano**M. J. M. Penning de Vries
et al.[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

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Discussion Paper | Discussion Paper | Discussion Paper | Discussion Paper | Discussion Paper

ACPD

14, 7739–7775, 2014

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



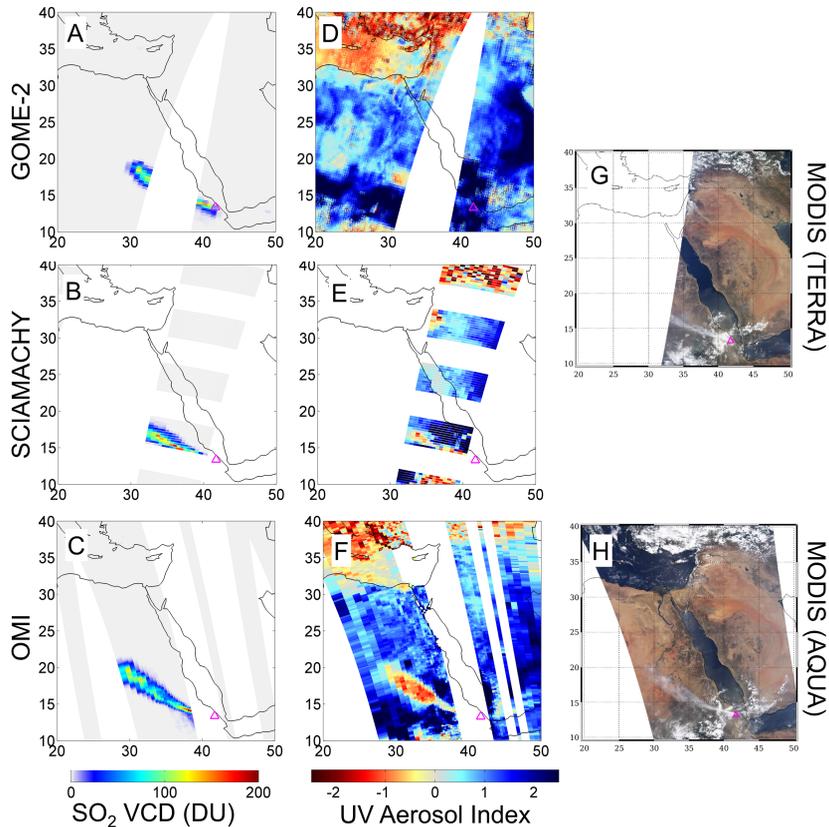


Fig. 1. SO₂ vertical column density in Dobson Units (**A–C**) and UV Aerosol Index (**D–F**) from GOME-2 (**A** and **D**), SCIAMACHY (**B** and **E**) and OMI (**C** and **F**) and true-colour images from MODIS on TERRA and AQUA (**G** and **H**) of the region around the Nabro volcano on 13 June 2011. Pink triangles indicate the location of the volcano. The overpass times (UTC) of the different instruments are: 06:25 and 08:05 for GOME-2, 07:40 for MODIS-TERRA, 07:50 for SCIAMACHY, 10:50 for MODIS-AQUA, and 11:00 for OMI.

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



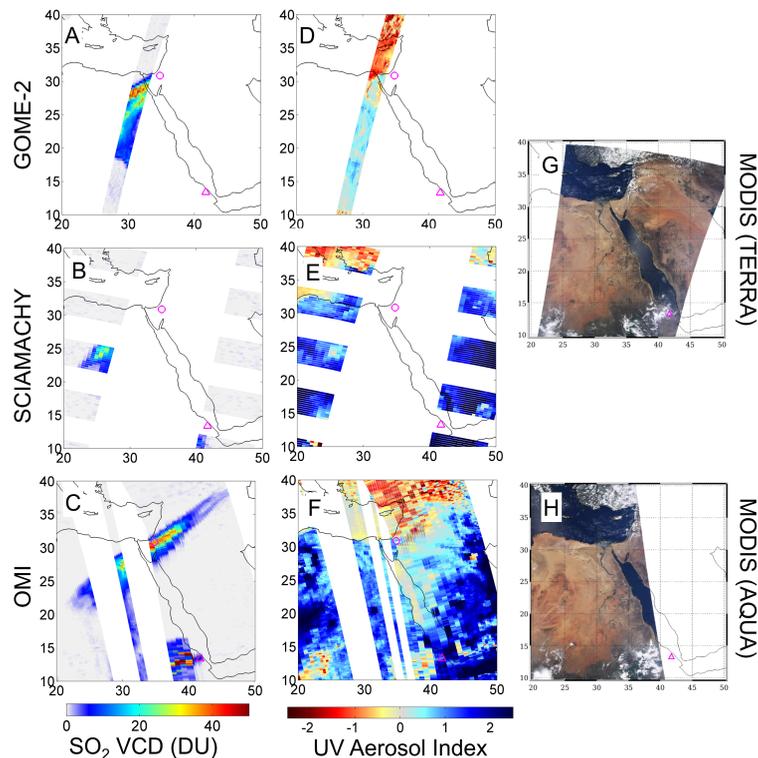


Fig. 2. SO₂ vertical column density in Dobson Units (**A–C**) and UV Aerosol Index (**D–F**) from GOME-2 (**A** and **D**), SCIAMACHY (**B** and **E**) and OMI (**C** and **F**) and true-colour images from MODIS on TERRA and AQUA (**G** and **H**) of the region around the Nabro volcano on 14 June 2011. GOME-2 was operating in narrow swath mode. Pink triangles and circles indicate the location of the volcano and of the Sede Boker MPLNET/AERONET station, respectively. The overpass times (UTC) of the different instruments are: 07:40 for GOME-2, 08:20 for MODIS-TERRA, 08:50 for SCIAMACHY, 10:00 and 11:40 for OMI, and 11:30 for MODIS-AQUA.

Sulphate aerosols from Nabro volcano

M. J. M. Penning de Vries et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

◀ ▶

◀ ▶

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



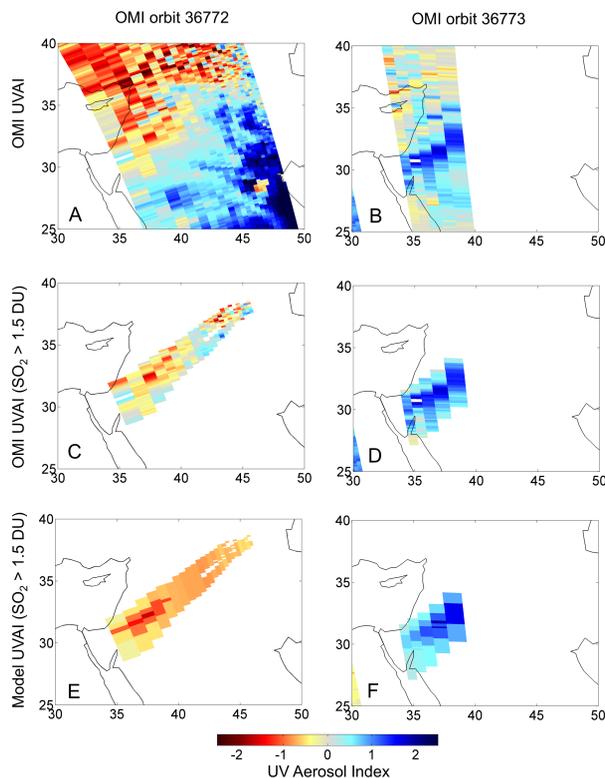
Sulphate aerosols
from Nabro volcanoM. J. M. Penning de Vries
et al.

Fig. 3. UV Aerosol Index from OMI (A–D) and from RTM calculations (E–F) for the volcanic plume detected on 14 June 2011. The right panels (A, C, and E) show data from orbit 36 772; on the left (panels B, D, and F) data from the following orbit, 36 773, are shown. The data in panels C–F are filtered by SO_2 vertical column density (≥ 1.5 DU) to more clearly show the volcanic plume. The UVAI in panels E and F were modelled using McArtim with aerosol parameters representative of sulphate particles at 18–19 km altitude (see text for details).

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



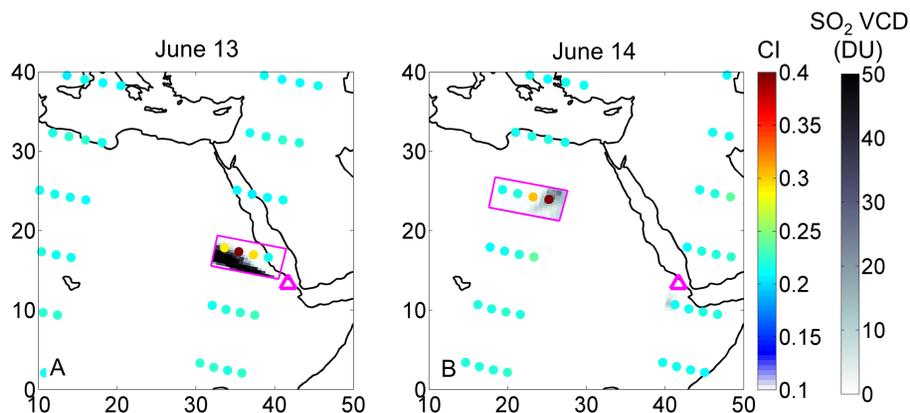
Sulphate aerosols
from Nabro volcanoM. J. M. Penning de Vries
et al.

Fig. 4. Maps of the SCIAMACHY Colour Index at a tangent height of 18.6 km for the region around the Nabro volcano on 13 June (**A**) and on 14 June 2011 (**B**). SO₂ columns from SCIAMACHY are shown in the background with a grey colour scale. The volcano is marked by a pink triangle; the pink boxes denote the two SCIAMACHY states to which the limb extinction retrieval was applied (Fig. 5).

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

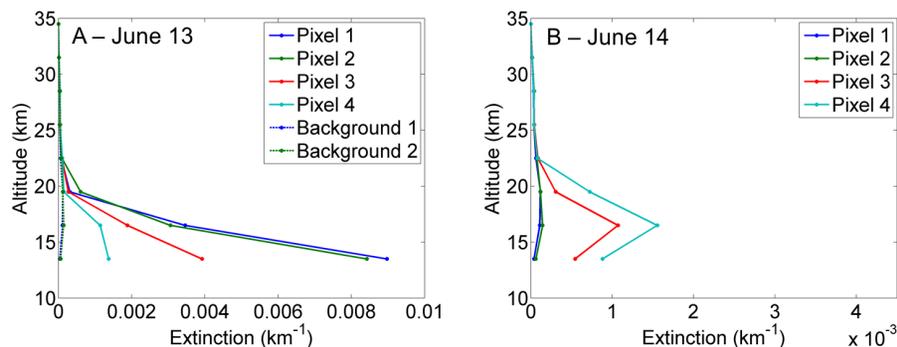
Sulphate aerosols
from Nabro volcanoM. J. M. Penning de Vries
et al.

Fig. 5. Limb extinction profiles of all limb pixels in the SCIAMACHY state containing parts of the SO₂ plume on 13 June (**A**) and on 14 June 2011 (**B**). The two profiles labelled “Background” (dashed lines in **A**) are the profiles from pixels 1 and 2 in on 14 June (**B**). Note that the panels have different x-axis scales. Pixel enumeration is from West to East.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

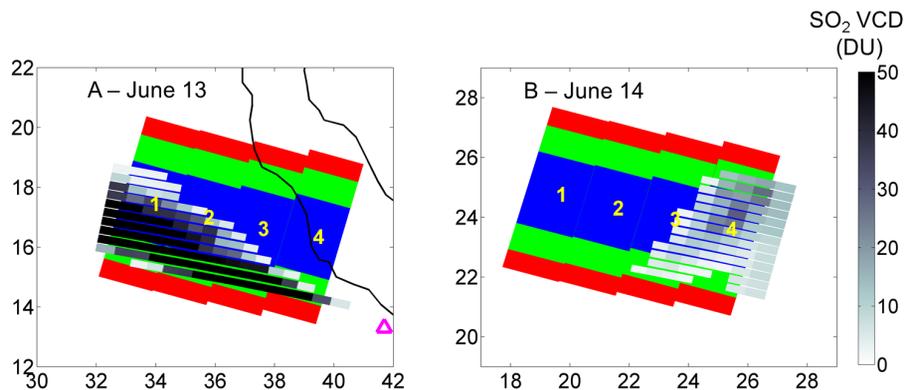
Sulphate aerosols
from Nabro volcanoM. J. M. Penning de Vries
et al.

Fig. 6. Areas contributing to the limb signal at TH = 19.5 km (blue), TH = 16.5 km (green), and TH = 13.5 km (red) for a homogeneous, infinite aerosol layer located at 18–21 km altitude. SCIAMACHY SO₂ columns (grey colour scale; only values > 2 DU are shown) are plotted to indicate the extent of the volcanic plume on 13 June **(A)** and 14 June **(B)**. The limb pixel numbers are shown in yellow at their respective TPs.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



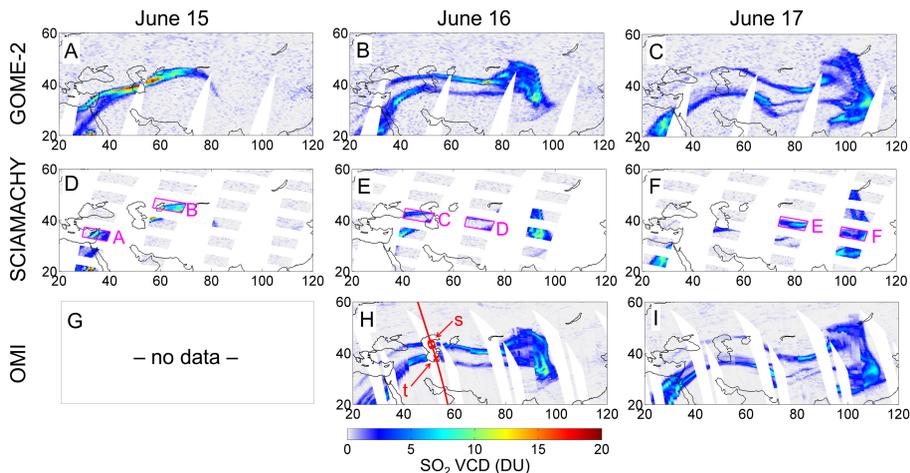
Sulphate aerosols
from Nabro volcanoM. J. M. Penning de Vries
et al.

Fig. 8. SO_2 time series from GOME-2 (A–C in the upper row), SCIAMACHY (D–F; centre), and OMI (G–I; lower row). Measurements from 15 June (left), 16 June (centre) and 17 June 2011 (right) are shown. There are no measurements available from OMI on 15 June. The extinction profiles from the states marked by labelled pink boxes are shown in the corresponding panels of Fig. 9. The red line indicates the ground track of one (near-simultaneous) CALIPSO orbit; the red symbols and arrows mark the locations of tropospheric (“t”, cross) and stratospheric (“s”, circles) aerosol layers detected by CALIOP.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



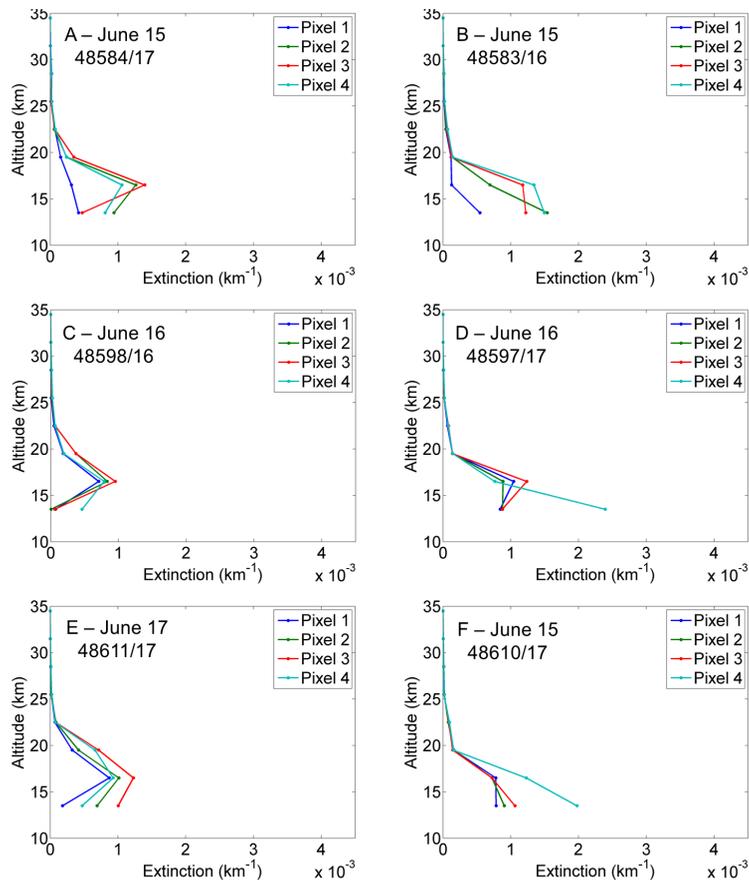
Sulphate aerosols
from Nabro volcanoM. J. M. Penning de Vries
et al.

Fig. 9. Limb extinction profiles of all limb pixels within the pink boxes in Fig. 8. First row (15 June): **(A)** orbit 48 584, state 17; **(B)** orbit 48 583, state 16. Second row (16 June): **(C)** orbit 48 598, state 16; **(D)** orbit 48 597, state 17. Third row (17 June): **(E)** orbit 48 611, state 17; **(F)** orbit 48 610, state 17. Pixel enumeration is from West to East.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

