

# Daily ozone cycle in the stratosphere: global, regional and seasonal behaviour modelled with the Whole Atmosphere Community Climate Model

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## Abstract.

The Whole Atmosphere Community Climate Model (WACCM) is utilised to study the daily ozone cycle and underlying photochemical and dynamical processes. The analysis is focused on the daily ozone cycle in the middle stratosphere at 5 hPa where satellite-based trend estimates of stratospheric ozone are most biased by diurnal sampling effects and drifting satellite orbits. The simulated ozone cycle shows a minimum after sunrise and a maximum in the late afternoon. Further, a seasonal variation of the daily ozone cycle in the stratosphere was found. Depending on season and latitude, the peak-to-valley difference of the daily ozone cycle varies mostly between 3–5 % (0.4 ppmv) with respect to the midnight ozone volume mixing ratio. The maximal variation of 15 % (0.8 ppmv) is found at the polar circle in summer. The global pattern of the strength of the daily ozone cycle is mainly governed by the solar zenith angle and the sunshine duration. In addition, we find synoptic scale variations in the strength of the daily ozone cycle. These variations are often anti-correlated to regional temperature anomalies and are due to the temperature dependence of the rate coefficients  $k_2$  and  $k_3$  of the Chapman cycle reactions. Further, the NO<sub>x</sub> catalytic cycle counteracts to the accumulation of ozone during daytime and leads to an anti-correlation between anomalies in NO<sub>x</sub> and the strength of the daily ozone cycle. Similarly, ozone recombines with atomic oxygen which leads to an anti-correlation between anomalies in ozone abundance and the strength of the daily ozone cycle. At higher latitudes, an increase of the westerly (easterly) wind cause a decrease (increase) in the sunshine duration of an air parcel leading to a weaker (stronger) daily ozone cycle.

## 20 1 Introduction

Despite being a minor constituent, ozone has a major impact on the Earth system. The stratospheric ozone layer absorbs solar radiation at short wavelengths and thus protects the biosphere from harmful ultraviolet radiation. Furthermore, the thermal state of the stratosphere mainly depends on absorption of downwelling ultraviolet and upwelling infrared radiation by ozone molecules. The associated  
25 production and destruction processes were described for the first time by the so-called Chapman cycle (Chapman, 1930) based on pure oxygen photochemistry. Later, the investigation of catalytic ozone depletion cycles revealed additional destruction processes of ozone. In particular, odd nitrogen ( $\text{NO}_y$ ) and odd chlorine ( $\text{Cl}_y$ ) cause catalytic ozone depletion in the stratosphere (Crutzen, 1970; Johnston, 1971; Stolarski and Cicerone, 1974).

30 In the 1970s, Molina and Rowland (1974) indicated that based on catalytic ozone depletion, anthropogenic emission of chlorofluorocarbons (CFCs) potentially damage the stratospheric ozone layer globally. Ozone observations in the mid-80s showed the surprising existence of the Antarctic ozone hole (Chubachi, 1985; Farman et al., 1985). To avert further depletion of stratospheric ozone, a ban of man-made CFC emissions was negotiated in the Montreal Protocol (Sarma and Bankobeza,  
35 2000).

In fact, present observations indicate a tendency for a slow recovery of the stratospheric ozone layer (Jonsson et al, 2009; Garny et al., 2013; Chehade et al., 2013; Kyrölä et al., 2014; Gebhardt et al., 2014). Projections of chemistry-climate models denote a return of stratospheric ozone concentrations to their 1980s level in the period between 2040 and 2050 (Eyring et al., 2007). Recent studies  
40 show that unregulated anthropogenic  $\text{N}_2\text{O}$  emissions are the dominant ozone-depleting emission in the 21st century (Ravishankara et al., 2009). The increasing  $\text{N}_2\text{O}$  fluxes from the troposphere and  $\text{CH}_4$  play an increasing role in determining stratospheric ozone and potentially delay the recovery of the ozone layer (Revell et al., 2012; Shindell et al., 1998).

The diurnal variation of stratospheric ozone was simulated with a photochemical box model at  
45 northern mid-latitudes by Herman (1979). This model depicted a strong decrease of ozone during daytime at 55 km in the mesosphere ( $-25\%$ ) and a weaker variation ( $\pm 5\%$ ) at 40 km in the stratosphere. Pallister and Tuck (1983) modelled the diurnal ozone cycle at  $34^\circ\text{N}$  with a photochemical box model and confirmed the results of Herman (1979). Furthermore, Fabian et al. (1982) indicated latitudinal and seasonal effects by utilisation of a two-dimensional, zonally averaged model. At that  
50 time the model utilised by Fabian et al. (1982) did not couple dynamics and diurnal photochemistry. Observations as rocket-based measurements of day- and nighttime ozone (Lean, 1982) and balloon-borne measurements (Aimedieu et al., 1981) seemed to support the simulation results of Fabian et al. (1982); Herman (1979) and Pallister and Tuck (1983). However, quality and sampling of the few measurements were not sufficient to give observational evidence for the small, daily cycle  
55 of stratospheric ozone.

Nowadays, there exist more observations of the daily cycle of stratospheric ozone which well

agree with the numerical simulation results of the photochemical models. Connor et al. (1994) presented diurnal variation in stratospheric and mesospheric ozone from 9 months of observations of a ground-based microwave radiometer at Oroville, California. Haefele et al. (2008) found a daily  
60 ozone cycle in the observations of the Ground-based Millimeter-wave Ozone Spectrometer (GROMOS) which operates at Bern, Switzerland in the frame of NDACC. Later, Studer et al. (2013) derived a monthly climatology of diurnal variation in stratospheric and mesospheric ozone from 17 years of GROMOS observation. Satellite-based observations from SMILES (Superconducting Submillimeter-Wave Limb-Emission Sounder) (Kikuchi et al., 2013), SABER (Sounding of the At-  
65 mosphere using Broadband Emission Radiometry) (Russell et al., 1999) and MLS (Microwave Limb Sounder) (Barth et al., 1983) showed the existence of a daily ozone cycle in the stratosphere at tropical and mid-latitudes (Huang et al., 2008; Sakazaki et al., 2013). There is also some evidence that this data show good agreement with model data at tropical latitudes (Sakazaki et al., 2013).

Trend studies on the global stratospheric ozone distribution are mainly based on ozone measurements from satellites (e.g. Stolarski and Frith, 2006). After years in operation satellite orbits potentially drift away from their initial orbits, so that the measurements are taken at local times differing by hours from the initial measurements (DeLand et al., 2012). In this context, diurnal ozone variation and its global pattern should be taken into account for a correction of the satellite ozone data (e.g. Bhartia et al., 2013). The recovery of the stratospheric ozone layer is expected to be of approx-  
75 imately 1 % per decade (Garny et al., 2013; Chehade et al., 2013) so that even small diurnal ozone variations can seriously bias the ozone trend estimate. To date, most observations based on satellites with different equator-crossing time are analysed under no consideration of diurnal ozone variations (Bhartia et al., 2013).

Aside the trend analysis, current research considers ozone as a source of excitation of atmospheric  
80 tides which in turn effect the daily ozone cycle (Sakazaki et al., 2013). Further, fundamental research on photochemistry by modelling of the diurnal variations of ozone, ozone depleting substances and other trace gases (Muncaster et al., 2012; Khosravi et al., 2013) is investigated as well as derivation of reaction rates (Kuribayashi et al., 2014).

The work presented here provides a first overview on global, seasonal and regional behaviour of  
85 the daily ozone cycle in the stratosphere. We look at the dependence of the daily ozone cycle on geographic longitude which was not discussed in previous observational or modelling studies. In order to understand and estimate the effects, we perform a 3-D high-resolution numerical simulation with the WACCM model. This model includes the physical and chemical processes which are relevant for the daily ozone cycle (e.g. detailed photochemistry, transport, vertical mixing, parametrized gravity wave fluxes). In addition, the model provides a realistic composition and circulation of the middle  
90 atmosphere.

In Sect. 2 the WACCM model and its setup for simulating the daily ozone cycle is described. Model output and methods of data analysis are specified. In Sect. 3 the chemistry and photochemistry

of the stratosphere are discussed. Section 4 shows the daily ozone cycle and its underlying chemical  
95 and photochemical processes at 5 hPa. Section 5 describes seasonal effects in the daily ozone cycle.  
Section 6 describes regional effects in the daily ozone cycle. The conclusions give a brief summary  
of the results and their implications for a correction of diurnal effects in ozone series.

## 2 The Whole Atmosphere Community Climate Model

The Whole Atmosphere Community Climate Model is a community access model which was devel-  
100 oped by the National Center of Atmospheric Research (NCAR) (Garcia et al., 2007; Marsh et al.,  
2007; Tilmes et al., 2007). The WACCM model consists of individual land, ice, ocean and atmo-  
sphere models which interact within the Community Earth System Model (CESM) software frame-  
work.

The atmosphere model fully couples chemistry, radiation and dynamics including the processes of  
105 chemical heating, gravity wave drag, ionization, and molecular diffusion. WACCM calculates ozone  
which feeds back to the thermal state and dynamics of the modelled atmosphere. The model domain  
ranges from the Earth’s surface to  $5.1 \times 10^{-6}$  hPa ( $\sim 150$  km) hence comprises the troposphere,  
stratosphere, mesosphere and the lower thermosphere. The implemented 3-D chemistry model is  
called Model of Ozone and Related Chemical Tracers (MOZART) (Horowitz et al., 2003). It cal-  
110 culates the abundance of 59 neutral species including the  $O_x$ ,  $NO_y$ ,  $HO_x$ ,  $Cl_y$  and  $Br_y$  families.  
Besides chemistry, the photolysis rates of the species are calculated (e.g.  $O_3$ ,  $N_2O$ ,  $H_2O$ , CFCs).

The model was intercompared with other chemistry-climate models during the CCMVal activity  
(SPARC Report, 2010) and performed very well. WACCM calculates accurate photolysis rates under  
consideration of radiative transfer, aerosol and cloud distributions. For this study we perform free-  
115 running simulations of WACCM without nudging of meteorological reanalysis data.

### 2.1 Numerical simulation

Numerical simulations for this work were performed utilising Version 4 of the WACCM model  
which was released in February 2012. The F 2000 WACCM scenario was utilised to simulate one  
year starting from 1 January. The atmosphere and the land model are free-running while ocean, land  
120 ice and sea ice models read climatologies.

The analysis deals with atmospheric effects of short time scales compared to common climate re-  
lated processes. On this account we perform the simulations with a coupling of dynamics and physics  
with a time step of 15 min which is half of the standard time resolution of WACCM. Although it is  
desirable to have even shorter time steps for photochemistry during sunrise, the simulations are  
125 limited by computing capacity.

All simulations were carried out with a horizontal resolution of  $4^\circ$  latitude by  $5^\circ$  longitude. The  
model has 66 vertical layers on hybrid-pressure ( $\sigma - p$ ) coordinates (Phillips, 1957) with a vertical

resolution ranging from 1.1–2.0 km in the middle atmosphere. The coordinates are terrain-following below the 100 hPa level and isobar above.

## 130 2.2 Model output and data analysis

The output data of the model are instantaneous fields of atmospheric parameters with an output interval of one hour. For the study of the daily ozone cycle, we often transfer from universal time (UT) to local solar time meridian (LSTM):

$$\text{LSTM} = \text{UT} + \phi \frac{24 \text{ h}}{360^\circ}, \quad (1)$$

135 where  $\phi$  refers to longitude in degree. The term “LT” is used as a synonym for the inconvenient “LSTM”.

In the middle stratosphere, ozone reaches a minimum and a maximum during a day at a given location. The difference between these extrema is the peak-to-valley difference  $D_{\text{O}_3}$

$$D_{\text{O}_3} = \text{O}_{3, \text{max}} - \text{O}_{3, \text{min}}, \quad (2)$$

140 where  $\text{O}_{3, \text{max}}$  and  $\text{O}_{3, \text{min}}$  refer to the extrema within the period of one day. In the following we often use the word “ozone” as a synonym for ozone volume mixing ratio. This convention applies also to other species within this study. Similarly, peak-to-valley differences in temperature are defined as

$$D_{\text{T}} = T_{\text{max}} - T_{\text{min}}, \quad (3)$$

with  $T_{\text{max}}$  and  $T_{\text{min}}$  referring to the maximum and the minimum temperature of a day.

145 Within the discussion of regional effects of the ozone variations, longitudinal anomalies in a species X are defined by

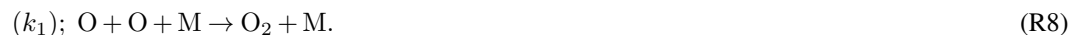
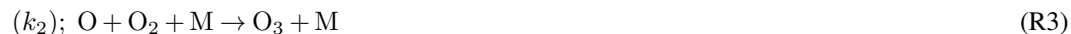
$$X'(\lambda, \phi, z, t) = X(\lambda, \phi, z, t) - \bar{X}(\lambda, z, t) \quad (4)$$

where  $\bar{X}$  is the zonal mean of X. The arguments  $\lambda$ ,  $\phi$ ,  $z$  and  $t$  refer to latitude, longitude, altitude and time respectively.

## 150 3 Chemical and photochemical reactions

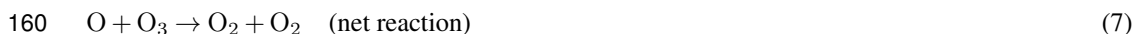
In a first approach, pure oxygen photochemistry explains the occurrence of stratospheric ozone. The so-called Chapman cycle (Chapman, 1930) is given by the following reactions derived from pure

oxygen photochemistry:



All reactions are given with their reaction rate coefficients  $k_i$  and photolysis frequencies  $J_i$  according to the notation of Brasseur and Solomon (2005). The photolysis of  $\text{O}_3$  is energetically allowed at wavelengths  $\lambda \leq 1180$  nm and mostly takes place in the spectral regions of the Hartley, Huggins and Chappuis bands. Photolysis of  $\text{O}_2$  occurs for ultraviolet radiation of wavelengths  $\lambda \leq 242.4$  nm  
 155 (Sander et al., 2011).

Stratospheric ozone is reduced by catalytic cycles shown in Eqs. (5–7) where  $X$  represents a generic catalyst.



In this scheme the catalyst destroys an ozone molecule while forming molecular oxygen and becoming oxidised. In a further reaction, the oxidised catalyst reacts with atomic oxygen resulting in a oxygen molecule and a reactivated catalyst. In sum these reactions transform an ozone molecule  
 165 and atomic oxygen into two oxygen molecules. A continuous repetition of the catalytic cycle can destroy a large number of ozone molecules by a single catalyst (Johnston, 1971). Catalytic cycles are mostly interrupted by deactivation of catalysts into reservoir species (e.g.  $\text{HCl}$ ,  $\text{N}_2\text{O}_5$ ,  $\text{ClONO}_2$ ) or by transport to the troposphere. For the stratosphere and the mesosphere the catalysts are essentially from the families of  $\text{NO}_y$  (Crutzen, 1970; Johnston, 1971) and  $\text{HO}_x$  (Bates and Nicolet, 1950). Cat-  
 170 alytic ozone depletion is also fostered by halogens of families such as  $\text{Cl}_y$  (Stolarski and Cicerone, 1974) and  $\text{Br}_y$  (McElroy et al., 1986; Tung et al., 1986; Wofsy et al., 1975).

Chemical and photochemical reactions are handled by the MOZART model which is embedded in WACCM. The contributions to the ozone budget given by Eq. (8) correspond to our model simulation. The reaction involving  $\text{NO}$ ,  $\text{NO}_2$ ,  $\text{H}$ ,  $\text{OH}$ ,  $\text{HO}_2$ ,  $\text{Cl}$  and  $\text{Br}$  are parts of catalytic ozone

175 depletion cycles which are limited by the rates of intermediate steps (Johnston and Podolske, 1978).

$$\begin{aligned} \frac{d(O_3)}{dt} = & k_2(M)(O)(O_2) - (J_3 + J_3^*)(O_3) - k_3(O)(O_3) - k_5(O(^1D))(O_3) \\ & - b_4(NO)(O_3) - b_9(NO_2)(O_3) - a_2(H)(O_3) - a_6(OH)(O_3) \\ & - a_{6b}(HO_2)(O_3) - d_2(Cl)(O_3) - e_2(Br)(O_3) \end{aligned} \quad (8)$$

The reaction and photolysis rates of Eq. (8) are again denoted as in Brasseur and Solomon (2005) and missing definitions of reaction rates are described in Appendix A. For more details about Eq. (8) the reader is referred to Emmons et al. (2010). The concepts as catalytic cycles, null cycles and consumptive sequences leading to differential equations of atmospheric constituents as Eq. (8) were reviewed by Johnston and Podolske (1978).

#### 4 The daily ozone cycle

185 Earlier studies on the daily ozone cycle of the stratosphere (e.g. Pallister and Tuck, 1983; Herman, 1979; Fabian et al., 1982) indicate a strong connection to chemistry and photochemistry. In this section, we analyse WACCM simulations of chemical and photochemical processes under their coupling to dynamics and the thermal state of the atmosphere. The analysis focuses on 5 hPa where strong diurnal variations of ozone occur in the stratosphere.

Figure 1 shows the simulated net ozone conversion rate (black line) for different latitudes at 21 March. Additionally, the coloured lines depict contributions of specific cycles. The sum of those cycles (coloured lines) yield the net conversion rate of ozone (black line). The label ‘‘Chapman’’ (blue line) refers to all ozone involving reactions of the pure oxygen photochemistry (Reactions R3–R7). Within the Chapman cycle, ozone production (Reaction R3) dominates during daytime over ozone reduction (Reactions R4–R7) in the stratosphere at 5 hPa. This result is not trivial since the partitioning between ozone and atomic oxygen reverses at mesospheric altitudes leading to minimal ozone during daytime.

Figure 1 demonstrates that the NO cycle (red line) is the dominant catalytic ozone depletion cycle at 5 hPa. At lower and middle latitudes, the NO cycle is approximately 2–4 times stronger than the Cl cycle. According to the model simulation, conversion rates of NO<sub>2</sub>, H, OH, HO<sub>2</sub> and Br catalytic cycles are more than 10 times less compared to the NO cycle at 5 hPa. The WACCM simulation confirms earlier results of Pallister and Tuck (1983) who identified the NO<sub>y</sub> cycle as the dominant catalytic cycle at altitudes from 26–43 km at noon.

At all latitudes shown in Fig. 1, the net conversion rate results in periods of ozone decrease and accumulation. For instance, at the equator ozone decreases from sunrise (approximately 06:00 LT) until about 09:00 LT. Subsequently, ozone is accumulated until about 16:00 LT. During the evening ozone decreases until about 21:00 LT.

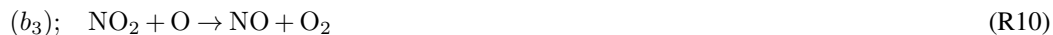
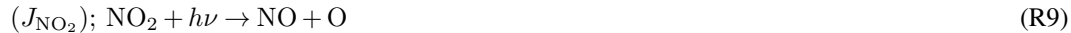
In Fig. 1 the Chapman cycle and the NO cycle show a dependence on latitude whereas the Cl cycle is almost constant over latitude. At the equator the conversion rates of the Chapman cycle and

the NO cycle are approximately two times stronger than at 74° N or 74° S at 21 March. The increase  
 210 of the Chapman cycle and the NO cycle from high to low latitudes is caused by the increase of solar  
 irradiance with decreasing solar zenith angle (Craig and Ohring, 1985). Further, in Fig. 1 the net  
 conversion rate of ozone shows a dependence on latitude. Thus the solar zenith angle is a regulator of  
 diurnal ozone decrease and accumulation at 5 hPa in the stratosphere. The almost constant Cl cycle  
 over latitude may be attributed to the global distribution of Cl species. For instance, the abundance  
 215 of ClO which is a source of active Cl is approximately two times higher than at 60° latitude in the  
 tropics in March (Solomon and Garcia, 1984). The increase of ClO at mid and high latitudes may  
 compensate the effect of the increased solar zenith angle at these latitudes.

Figure 2 shows the simulated daily ozone cycle at 5 hPa for March, June, September and Decem-  
 ber. The cycles comprise a morning minimum and a late afternoon maximum. The peak-to-valley  
 220 difference  $D_{O_3}$  (Eq. 2) is approximately 1.5 % (0.11 ppmv) of midnight ozone in December and 5 %  
 (0.3 ppmv) in June derived from zonal mean data at 50° N.

A similar increase of ozone during the afternoon was found in a climatology of the daily ozone  
 cycle observed by the ground-based microwave radiometer GROMOS in Bern, Switzerland (Studer  
 et al., 2013). Satellite-based observations from SMILES (Kikuchi et al., 2013), SABER and MLS  
 225 showed a morning minimum and a late afternoon maximum in the daily ozone cycle in the tropics  
 and subtropics (Huang et al., 2008; Sakazaki et al., 2013). In Fig. 2 the strength of  $D_{O_3}$  differs for  
 March, June, September and December which indicates a seasonal variation in the daily ozone cycle.  
 The existence of a seasonal variation of the daily ozone cycle was also suggested by Sakazaki et al.  
 (2013). We discuss the seasonal effects in more detail in Sect. 5.

The characteristic morning minimum and the late afternoon maximum mostly result from chemi-  
 cal and photochemical reactions. For instance, after sunrise NO is activated by photodissociation of  
 NO<sub>2</sub> at wavelengths  $\lambda \leq 402$  nm as shown in Reaction (R9) and to a minor extent by Reaction (R10).



230 The activated NO rapidly decreases ozone after sunrise by means of the NO cycle. During the  
 morning, ozone is rebuild and further increased due to photodissociation of oxygen molecules and  
 ozone production given by Reactions (R1) and (R3).

Figure 3 shows the NO conversion rates decomposed into chemical and photochemical reactions  
 on 21 March. Reactions with absolute conversion rates less than 5 ppbv h<sup>-1</sup> are omitted in Fig. 3.  
 According to the model simulation, NO is mainly activated by photodissociation of NO<sub>2</sub> given by  
 Reaction (R9) (orange line) and marginally by Reaction (R10) (red line). Most of the activated NO  
 directly decreases ozone (cyan line) by means of the NO cycle. To a minor extent NO couples to  
 the Cl<sub>y</sub> family and activates Cl by Reaction (R11) (blue line).





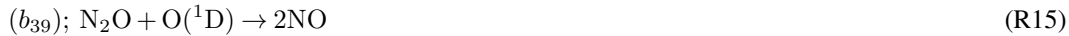
Similar results were derived from photochemical box models (e.g. Pallister and Tuck, 1983). The 3-D chemistry-climate model WACCM confirms these results on global scale and under inclusion  
 235 of dynamical and thermal effects in a fully coupled atmosphere model with realistic distributions of trace gases, transport processes, mixing, gravity waves, planetary waves, tides, and wave-mean flow interactions.

The NO production over daytime strongly depends on NO<sub>2</sub> at 5 hPa. NO<sub>2</sub> in turn reacts with further atmospheric trace gases as nitrogen compounds. For instance, N<sub>2</sub>O<sub>5</sub> is assembled during nighttime by removing NO<sub>2</sub> from the stratosphere by termolecular Reaction (R12).



During daytime N<sub>2</sub>O<sub>5</sub> is thermally decomposed or photodissociated by Reactions (R13) and (R14). In this context, N<sub>2</sub>O<sub>5</sub> is a temporary reservoir for reactive NO<sub>x</sub> (NO and NO<sub>2</sub>). The coupling of  
 240 NO<sub>2</sub> to the reservoir gives a contribution to the diurnal NO<sub>x</sub> cycle (Walker and Dudhia, 2011) and in turn influences the daily ozone cycle. For instance, ozone decrease by the NO cycle increases from the morning to the afternoon. NO<sub>x</sub> which is in form of N<sub>2</sub>O<sub>5</sub> in the early morning is slowly released from the reservoir and increases the conversion rates of the NO cycle over daytime.

Sources of NO<sub>x</sub> in the stratosphere are for instance natural and anthropogenic surface emissions of N<sub>2</sub>O. These emissions of N<sub>2</sub>O ascend into the stratosphere where NO<sub>x</sub> is increased by means of Reaction (R15).



For this reason N<sub>2</sub>O is an important regulator of stratospheric ozone. In the lower and middle  
 245 stratosphere the chemical and dynamical lifetimes of NO<sub>x</sub> are comparable so that transport plays a role. N<sub>2</sub>O passes the tropopause mostly in the tropics and is globally distributed in the stratosphere by the Brewer–Dobson circulation.

## 5 Seasonal variations of the daily ozone cycle

Figure 4 shows  $D_{\text{O}_3}$  over latitude for March, June, September and December. In March and September  
 250  $D_{\text{O}_3}$  is decreasing from the equator (0.35 ppmv) towards the poles (< 0.07 ppmv). Surprisingly, June and December show significantly enhanced  $D_{\text{O}_3}$  at high latitudes in the summer hemisphere. For instance,  $D_{\text{O}_3}$  peaks in the simulation data at 66° S (0.79 ppmv, 14.5 %) in December and at 66° N (0.74 ppmv, 11.7 %) in June.

The simulation data of March and September (blue and cyan lines) reinforce the assumed regulation  
 255 of ozone decrease and accumulation ( $D_{\text{O}_3}$ ) by the solar zenith angle at 5 hPa. Nevertheless, the solar zenith angle cannot explain the high values of  $D_{\text{O}_3}$  around the polar circle in summer.

Figure 2 illustrates the seasonal variation in  $D_{O_3}$  in the Northern Hemisphere at  $50^\circ$  N. In June the small solar zenith angle and the long sunshine duration (period of insolation over daytime) lead to stronger effects in  $D_{O_3}$  than a short day with high solar zenith angle in December. The sunshine duration strongly depends on season at higher latitudes and is constant at the equator. The seasonal altering of the sunshine duration impacts the photochemical system in the stratosphere. A long sunshine duration leads to a long period of ozone accumulation and hence to enhanced  $D_{O_3}$ .

Based on the sunshine duration, the solar zenith angle and auxiliary Fig. 5, we discuss the effects of  $D_{O_3}$  in December including the strong effect in the summer hemisphere at high latitudes. Figure 5 depicts the geometry of the Earth's axis and the solar radiation at December solstice. The solar zenith angle is the angle between the sun ray and the zenith ray as shown in Fig. 5. The threshold latitude ( $\lambda_{thr}$ ) assigned in Fig. 5 refers to the border of the polar day or polar night. Figure 5 includes four major regions of different characteristics with respect to sunshine duration and solar zenith angle:

- Polar night area:  
Absent solar radiation.
- Mid-latitudes and tropics:  
Sunshine duration increases from the northern winter to the southern summer hemisphere.  
Solar zenith angle at noon is zero at  $23.4^\circ$  S (subsolar latitude).
- Polar day area:  
Sunshine duration of 24 h with sun at low elevation and small changes in the solar zenith angle.
- South Pole:  
Sunshine duration of 24 h with constant solar zenith angle.

The four characteristic regions can be related to  $D_{O_3}$  over latitude in December as shown in Fig. 4 (black line). In December,  $D_{O_3}$  is marginal at the North Pole and the northern polar region due to absent solar radiation. Equatorwards of the Arctic circle the sunshine duration and the solar zenith angle show different behaviour. The solar zenith angle is decreasing until  $23.4^\circ$  S ( $\Theta = 0^\circ$ ) and increasing towards the South Pole. The sunshine duration is increasing from 0 h at the Arctic circle to 24 h at the Antarctic circle (polar day). At high latitudes the gradient in sunshine duration is higher than in the tropics. According to the model simulation this region shows an almost steady increase of  $D_{O_3}$  from the Arctic circle towards the Antarctic circle where  $D_{O_3}$  peaks. The model resolution is too low to specify whether the peak in  $D_{O_3}$  is poleward or equatorward of the Antarctic circle. We assume that the maximum of  $D_{O_3}$  is slightly shifted equatorwards of the Antarctic circle where ozone decreases during a short night period. Poleward of the Antarctic circle the sunshine duration is 24 h with small variations in the solar zenith angle. This situation leads to decreasing  $D_{O_3}$  and marginal  $D_{O_3}$  at the South Pole.

The dependence of the daily ozone cycle on the solar zenith angle and the sunshine duration mostly explains the complex picture of the seasonal variation of the daily ozone cycle over the globe

as described above by means of the WACCM simulation. Based on the solar zenith angle and the sunshine duration most seasonal effects of the ozone cycle in the stratosphere can be explained qualitatively. To the authors knowledge, the global seasonal behaviour of the daily ozone cycle was not comprehensively described in the literature. Previous simulation studies were focused on northern mid-latitudes (Pallister and Tuck, 1983). Observational studies (Kikuchi et al., 2013) presented the daily ozone cycle of the tropics and mid-latitudes or local data (Studer et al., 2013).

Figure 6 gives a complete view on the seasonal change of  $D_{O_3}$  as function of day of year (DOY) and geographic latitude as simulated by WACCM. In addition, the sunshine duration (dashed contours) and the solar zenith angle (solid contours) are given as contours in Fig. 6. The strong effects of  $D_{O_3}$  at the border of the polar day is present over a period of approximately 80–100 days and is confirmed by observations of a ground-based microwave radiometer at Ny-Ålesund, Svalbard (Palm et al., 2013).

With regard to satellite-derived ozone trend estimates, diurnal sampling effects are most critical at high latitudes during summer. However, since the expected ozone trends are of the order of 1 % per decade or less (Garny et al., 2013; Chehade et al., 2013), a thorough correction of the diurnal sampling effects in time series of stratospheric ozone are necessary at any location. Such a correction could be guided by a chemistry-climate model.

## 6 Regional effects of the daily ozone cycle

The daily ozone cycle has regional variations in  $D_{O_3}$  at 5 hPa. The origin of the regional variation is manifold. Chemical, dynamical and thermal effects result in synoptic scale variations of the daily ozone cycle. In order to reveal and quantify the mechanisms we analysed the 3-D WACCM monthly mean data within focus on regional modulations.

Figure 7a–d shows the strength of the daily ozone cycle in the stratosphere depending on latitude and longitude as simulated by WACCM for March, June, September and December. In addition, contour lines depict the strength of diurnal temperature variations  $D_T$ . Figure 7a–d comprises the seasonal characteristics of the daily ozone cycle previously discussed in Sect. 5. For instance, Fig. 7a shows increasing  $D_{O_3}$  from the poles towards the equator. The strong daily ozone cycle at high latitudes in the summer hemisphere appears in Fig. 7b and d which show  $D_{O_3}$  in June and December. Moreover, Fig. 7a–d contains regional effects of synoptic scale which are superposed on the global pattern of the daily ozone cycle. In the following subsections main sources of regional modulations are presented.

### 6.1 Anti-correlation of $D_{O_3}$ and $D_T$

Ozone has a major impact on the vertical temperature gradient of the atmosphere called lapse rate. The heating due to absorption of ultraviolet radiation explains the positive lapse rate of the

stratosphere. At the same time, ozone varies with ambient temperature since ozone photochemistry strongly depends on temperature. Thus we expect a close connection of diurnal temperature variations and the daily ozone cycle in the stratosphere.

330 In pure oxygen photochemistry, the temperature-dependent rate coefficients  $k_2$  and  $k_3$  of Reactions (R3) and (R4) are the main cause for the anti-correlation of ozone and temperature variations (e.g. Froidevaux et al., 1989; Craig and Ohring, 1985). In the stratosphere at 5 hPa ozone is approximately in photochemical equilibrium so that the determination of ozone reduces to Eq. (9) where M refers to the molecule concentration of air (Brasseur and Solomon, 2005). Based on the photochemical equilibrium, the impact of stratospheric temperature variations on ozone can be assessed  
335 by

$$O_3 = O_2 \sqrt{\frac{k_2 J_2}{k_3 J_3}} M. \quad (9)$$

Figure 8 shows the dependence of ozone on temperature variations according to a pure oxygen atmosphere described by Eq. (9). Temperature dependencies of oxygen, photolysis rates and air density  
340 are neglected. The reference temperature ( $T_0$ ) of 250 K is a realistic temperature of the stratosphere. The analytic expressions of the temperature-dependent rate coefficients  $k_2 \propto (300/T)^{2.4}$  and  $k_3 \propto \exp(-2060/T)$  correspond to the model simulation (Horowitz et al., 2003; Emmons et al., 2010). Our estimation shows an anti-correlation of ozone and temperature of approximately  $-2.15 \text{ \% K}^{-1}$  in the stratosphere (Fig. 8). Considering an average tropical ozone volume mixing ratio of 9.23 ppmv (derived from model data) we get an anti-correlation of  $-0.20 \text{ ppmv K}^{-1}$ . The  
345 temperature-induced ozone change is mainly based on Reaction (R4) (77 %) and less on Reaction (R3) (23 %). These results indicate that an anti-correlation of temperature and ozone appears in the stratosphere and that Reaction (R4) is sensitive to temperature variations.

The apparent diurnal temperature variations in the stratosphere are mainly due to migrating and  
350 non-migrating atmospheric tides. The WACCM data show temperature variations in the stratosphere of up to a few K (not explicitly shown) which are in a rough agreement to observations (e.g. Huang et al., 2010; Sakazaki et al., 2012). In addition, WACCM and the observations often show a temperature maximum in the late afternoon at 5 hPa. Thus diurnal temperature variations can affect the strength of the daily ozone cycle.

355 In order to assess these effects, it is assumed diurnal variations of ozone and temperature are in phase. In this case, the strength of the daily ozone cycle is damped by the temperature tide. Or more precise: we expect that a  $D_T$  value of 1 K leads to a 2.15 % decrease in  $D_{O_3}$  (as derived from Eq. 9).

On the other hand, we can directly derive from the temperature and ozone distributions of WACCM whether regional anti-correlations between  $D_T$  and  $D_{O_3}$  exist. In Fig. 7c we find regional anti-correlations of  $D_T$  (magenta contours) and  $D_{O_3}$  in the tropics in September. On average  
360 the strength of a tropical daily ozone cycle is 0.35 ppmv at 5 hPa and is attended by an average temperature variation of 1.34 K. A regionally damped daily ozone cycle occurs at the east coast of

Africa (Fig. 7c, Region A). In contrast, a regionally stronger daily ozone cycle resides in the Atlantic ocean east of the Caribbean (Fig. 7c, Region B). The values of  $D_T$  and  $D_{O_3}$  at Regions A and B and their deviations from the tropical average are summarised in Table 1. From the data of Table 1 we  
 365 derive an anti-correlation of  $D_T$  and  $D_{O_3}$  of approximately  $-0.11 \text{ ppmv K}^{-1}$ .

The simplified Eq. (9) and the WACCM simulation yield an anti-correlation of  $D_T$  and  $D_{O_3}$  which is of the same order of magnitude and certainly not negligible. In addition the WACCM simulation clearly shows strong, regional variations of  $D_T$ .

370 Our analysis demonstrates the impact of temperature variations on the daily ozone cycle. The interconnections of dynamics, chemistry and photochemistry in chemistry-climate models may help to correct these effects in satellite data utilised for ozone trend estimation. However, the high spatiotemporal variability of atmospheric tides is a challenge for a reliable correction of diurnal effects in ozone trend estimates.

## 375 6.2 Regional NO and O<sub>3</sub> variability and its impact on D<sub>O<sub>3</sub></sub>

In Sect. 4 it was argued that the Chapman cycle and the NO cycle cause most of the net ozone conversion rate at 5 hPa. Thus we expect that regional variations of O<sub>3</sub> and NO may have an impact on the strength of the daily ozone cycle. For understanding of the regional variations of O<sub>3</sub> and NO, it is desirable to study the families of O<sub>x</sub> and NO<sub>y</sub> instead of chemical reactions (e.g. Johnston and  
 380 Podolske, 1978). Model simulations with globally altered initial O<sub>x</sub> and NO<sub>y</sub> concentrations give information about influences of changes in trace gases on the resulting daily ozone cycle and  $D_{O_3}$ .

We performed model simulations with altered O<sub>x</sub> and NO<sub>y</sub> concentrations in the initial atmosphere of the 1 January at 00:00 UT (model date) and a control simulation without changes in the initial atmospheric conditions. We altered concentrations of NO<sub>y</sub> in order to preserve the concentration ratio of NO<sub>x</sub>/NO<sub>y</sub>. The ratio is often used to indicate the ability of NO<sub>x</sub> to destroy ozone  
 385 in the stratosphere (Brasseur and Solomon, 2005). Equations (10–12) define the sums of nitrogen compounds NO<sub>x</sub> and NO<sub>y</sub> as well as O<sub>x</sub>.

$$O_x = O + O_3 \quad (10)$$

$$NO_x = NO + NO_2 \quad (11)$$

390  $NO_y = N + NO + NO_2 + NO_3 + 2N_2O_5 + HNO_3 + HO_2NO_2 + ClONO_2 + BrONO_2 \quad (12)$

The ozone conversion rates for +10 % initial NO<sub>y</sub> concentration are shown in Fig. 9. The depletion rate of the NO cycle (dashed red line) is higher than in the control simulation data (solid red line) due to higher initial NO concentration. The slightly increased ozone conversion rate of the Chapman cycle (dashed blue line) is related to the initially enhanced NO<sub>2</sub> concentration. After sunrise, more  
 395 NO<sub>2</sub> is photodissociated by Reaction (R9) and provides more atomic oxygen for ozone production by means of Reaction (R3). Thus, the ratio of ozone production and depletion within the Chapman cycle slightly changes in favour of ozone production.

Figure 10 compares the daily ozone cycles at 1 January derived from simulations of  $\pm 10\%$  initial  
400  $\text{NO}_y$  concentration and the control simulation. All simulations comprise the morning minimum and  
the late afternoon maximum at comparable LT as discussed in a previous section. Lower (higher)  
 $\text{NO}_y$  concentration results in more (less) assembled ozone and hence increased (decreased)  $D_{\text{O}_3}$   
respectively.

In a similar manner we analyse influences of altered  $\text{O}_x$  concentrations in the initial atmosphere  
405 on the daily ozone cycle. Figure 11 shows the ozone conversion rates for  $+5\%$   $\text{O}_x$  concentration  
at 1 January. The conversion rate of the Chapman cycle is lower (dashed blue line) compared to  
the data of the control simulation (blue line). Hence, the ratio of ozone production and depletion  
within the Chapman cycle changes in favour of ozone depletion. Further, the depletion rate of the  
NO cycle (dashed red line) is higher compared to the control simulation data (solid red line). This  
410 behaviour can be derived from Eq. (8) where ozone depletion by NO is linear in ozone. Finally,  
Fig. 12 shows the daily ozone cycles derived from simulations with  $\pm 5\%$  initial  $\text{O}_x$  concentration  
and the control simulation. Lower (higher)  $\text{O}_x$  concentration results in more (less) assembled ozone  
and hence increased (decreased)  $D_{\text{O}_3}$ .

In summary, all model simulations with altered initial concentrations indicate that  $D_{\text{O}_3}$  is anti-  
415 correlated to changes in  $\text{O}_x$  and  $\text{NO}_y$  at 5 hPa. Considering that all loss terms in Eq. (8) are linear in  
ozone and  $\text{O}_x$  is mostly  $\text{O}_3$  this is a understandable behaviour. Based on these promising simulation  
results we assumed that an anti-correlation of  $D_{\text{O}_3}$  to  $\text{O}_x$  and  $\text{NO}_y$  also induces regional effects in  
the daily ozone cycle (comparable to Fig. 7a–d but not shown here).

We remark that the initial conditions of atmospheric fields were altered under no consideration of  
420 the correct partitioning of the day and night side of the Earth. The initial trace gases were altered  
simultaneously at 1 January, 00:00 UT at any point of the model grid. Since  $\text{O}_x$  is mostly  $\text{O}_3$  at 5 hPa  
and we preserved the  $\text{NO}_x/\text{NO}_y$  concentration ratio a potential bias in our simulation is considered  
to play a minor role.

Regional variations of atmospheric trace gases are described by anomalies as defined in Eq. (4).  
425 The regional anomalies in NO and  $\text{O}_3$  are of synoptic scale in the WACCM simulation. From  
tropical to middle latitudes these anomalies vary up to approximately  $\pm 10\%$ . In Fig. 13 we relate  
anomalies in  $\text{NO}_{\text{noon}}$  (NO at 12:00 LT) and  $\text{O}_{3,\text{mid}}$  ( $\text{O}_3$  at 24:00 LT) to anomalies in  $D_{\text{O}_3}$  for March,  
June, September and December. The used data in Fig. 13 is taken for the latitude range  $\lambda \pm 30^\circ$  where  
 $\lambda_{\text{sub}}$  is the subsolar latitude. Figure 13 shows that  $D_{\text{O}_3}$  is often anti-correlated to anomalies in  
430  $\text{NO}_{\text{noon}}$  and  $\text{O}_{3,\text{mid}}$ .  $D_{\text{O}_3}$  decreases from positive (red) to negative values (blue) when the anomalies  
 $\text{NO}'$  and  $\text{O}'_3$  increase from negative to positive values.

We find further effects at high latitude related to atmospheric composition. A strong effect based  
on the anti-correlation of anomalies in  $\text{O}_3$  and  $D_{\text{O}_3}$  occurs in the Pacific region in December (Fig. 7d,  
Region C) where an ozone anomaly of approximately  $-10\%$  leads to strong increases of  $D_{\text{O}_3}$  over  
435 a wide region. Nevertheless, the anti-correlation of NO and  $D_{\text{O}_3}$  remains weak due to the complex

partitioning of the  $\text{NO}_y$  family. Based on the reservoir coupling some of the reactive NO is in form of  $\text{N}_2\text{O}_5$  during the diurnal  $\text{NO}_x$  cycle (Walker and Dudhia, 2011). Owing to this fact, Fig. 14 gives an overview on distributions of NO,  $\text{NO}_2$  and  $\text{N}_2\text{O}_5$  showing the day and nighttime partitioning at 5 hPa. The  $\text{NO}_y$  partitioning as simulated by WACCM qualitatively agrees with satellite  
440 observations of MIPAS (Micheelson Interferometer for Passive Atmospheric Sounding) (Funke and López-Puertas, 2005; Fischer et al., 2008). However, it is difficult to give a comprehensive view on anomalies in reactive  $\text{NO}_x$  since it is never in a state of equilibrium with the reservoir  $\text{N}_2\text{O}_5$  (Brasseur and Solomon, 2005).

Aside anomalies, the abundances of NO,  $\text{NO}_2$  and  $\text{N}_2\text{O}_5$  shown in Fig. 14 are conspicuously low  
445 at the equator. Poleward drifting branches of the Brewer–Dobson circulation (Brewer, 1948) may explain the equatorial anomaly in Fig. 14. This might be an example how dynamics come into play and calling for 3-D chemistry-climate models to assess the daily ozone cycle.

### 6.3 Zonal wind-induced effects on $D_{\text{O}_3}$

In the stratosphere, westerly and easterly winds of up to  $100 \text{ m s}^{-1}$  occur. Due to the apparent move  
450 of the Sun from the east to the west, the sunshine duration of an air parcel decreases for westerly winds and increases for easterly winds. The photochemical system will respond to these changes with an altered daily cycle of ozone. The effect is more pronounced at high latitude, since the Earth's rotation velocity decreases at higher latitudes ( $v = \Omega a \cos(\lambda)$  with  $a$ : Earth radius,  $\Omega$ : Earth angular velocity,  $\lambda$ : latitude). Sonnemann (2001) formulated the zonal wind-dependence of the sunshine  
455 duration of an air parcel by means of the well-known Doppler effect. Numerical simulations showed that the quasi-periodic radiative forcing can induce photo-chemical oscillations of mesospheric trace gases (Doppler–Sonnemann effect).

For a rough estimate we consider an easterly wind with  $30 \text{ m s}^{-1}$  over a period of 7 h during daytime inducing a change in sunshine duration. The enhancement of sunshine duration is approxi-  
460 mately 7.5 % at  $60^\circ$  latitude and 3.8 % at the equator for an easterly wind. Respectively, a westerly wind will have the opposing effect, i.e. a reduction of sunshine duration. This significant change in sunshine duration manifests in a regional modulation of the daily ozone cycle. Regions with strong westerly winds tend to show weaker  $D_{\text{O}_3}$  whereas regions with easterly winds tend to show enhanced  $D_{\text{O}_3}$ .

465 For instance, in Fig. 7d strong westerly winds at high latitude can be found from the east coast of North America to central Siberia (Region D). These zonal winds correlate to regionally damped  $D_{\text{O}_3}$  (0.08 ppmv, 1.3 %).

### 6.4 Other potential sources of regional $D_{\text{O}_3}$ variability

The previous discussion did not consider regional variations of the actinic radiation in the strato-  
470 sphere as induced by changes in cloud coverage, aerosol distribution, surface albedo, and greenhouse

gases (Meier et al., 1997). We expect many effects on the daily cycle of stratospheric ozone and other parameters. Further it is well-known that turbulence and gravity wave flux have strong daily cycles which may effect the regional, seasonal and global behaviour of the daily ozone cycle. It is beyond the scope of the present study to investigate all effects. However, the WACCM simulations seem to  
475 be appropriate for advanced studies of the details of the daily ozone cycle.

Finally, it is desirable to link these results to satellite-based ozone trend analysis. In particular, corrections due to atmospheric constituents such as O<sub>3</sub> and NO and stratospheric temperature variations are challenging. In addition the spatiotemporal variability of thermal tides in the upper stratosphere may induce a variability of the daily ozone cycle.

## 480 7 Conclusions

We gave a comprehensive overview on the global, seasonal and regional behaviour of the daily ozone cycle in the stratosphere at 5 hPa modelled with the fully coupled chemistry-climate model WACCM. The daily cycle of stratospheric ozone is strongest at about 5 hPa where the peak-to-valley difference is approximately 3–5 % (0.4 ppmv) for mid-latitudes and tropics. Though during polar  
485 summer at high latitudes (66.6° lat) the daily ozone cycle can be up to 15 % (0.8 ppmv). Satellite-derived trend estimates of stratospheric ozone are certainly biased by the daily ozone cycle since the ozone trends are of the order of a few percent per decade and the satellite orbits are shifted in local time. Additionally, each satellite orbit slowly drifts away from its initial orbit during the years in operation. For a future correction of the diurnal biases in stratospheric ozone trend estimates it is  
490 important to derive the characteristics of the daily ozone cycle, to validate the characteristics and to understand the fundamental processes which are causing the daily ozone cycle. The present study shows that the global view of WACCM on the daily ozone cycle is invaluable for understanding and planing of a correction method.

Generally, the results of WACCM are in good agreement with previously reported results of photochemical box models at northern mid-latitudes (Herman, 1979; Fabian et al., 1982; Pallister and  
495 Tuck, 1983). WACCM and the box models show a daily cycle of stratospheric ozone with a 1 % decrease after sunrise and a slow ascent to approximately 3–5 % in the late afternoon (similar to Fig. 2). Thus it seems that WACCM should be quite reliable to inform about the daily ozone cycle at any region. In addition we find the daily ozone cycle of WACCM in a good agreement with observations by a microwave radiometer at Bern, Switzerland (Studer et al., 2013) which operates in  
500 the frame of NDACC and from satellite-based observation of SMILES (Sakazaki et al., 2013) and TIMED/SABER (Huang et al., 2008). The surprising strong daily ozone cycle at high latitudes in polar summer is confirmed by observations of a ground-based microwave radiometer at Ny-Ålesund, Svalbard (Palm et al., 2013). The simulation and observational result is of high importance for future  
505 trend studies of polar stratospheric ozone which is most disturbed, variable and difficult for model



projections of the expected recovery phase of the stratospheric ozone layer.

Comparing characteristics of the daily ozone cycle from our own model data and other relevant literature we note one remarkable difference: WACCM and the photochemical box models show a decrease of stratospheric ozone after sunrise while ground-based radiometers measure more or  
510 less constant values of stratospheric ozone after sunrise at mid-latitudes. The disagreement could be due to an observation error, a retrieval error, or a simulation error. This example shows how the daily ozone cycle could be utilised as a test signal for controlling and improving measurement and retrieval techniques as well as for quality assessment of atmospheric composition, photochemistry and dynamics in chemistry-climate models.

515 Our simulations show a manifold of variations in the daily ozone cycle at 5 hPa. Though the simple statement that the daily cycle of stratospheric ozone mainly depends on ozone photochemistry is true, ozone photochemistry itself depends on solar zenith angle, sunshine duration, temperature, zonal wind, ozone concentration and ozone depleting substances as  $\text{NO}_x$  and  $\text{Cl}_x$ . All together are responsible for the global, seasonal and regional characteristics of the daily ozone cycle presented  
520 in the entirety of our figures. On the other hand the WACCM model might underestimate dynamical effects near the polar winter vortices where effects based on photochemistry are weak and transport might play a major role. However, by means of the model it was partly possible to disentangle the different effects or to learn about their interconnections. For example, we find that temperature tides with amplitudes of approximately 1 K can lead to a reduction of 2 % in the strength of the daily ozone  
525 cycle. The high spatiotemporal variability of temperature tides could be a major hurdle for a future correction of diurnal biases in trend estimates of stratospheric ozone. Extensive simulations by chemistry-climate models could allow a realistic estimation of the diurnal biases in satellite-derived trend estimates of stratospheric ozone.

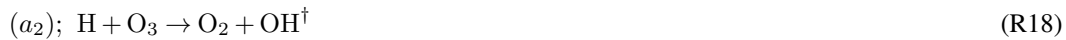
Our results may also help to understand the daily ozone cycle at different altitudes. For instance,  
530 in the upper stratosphere we assume a growing influence on ozone photochemistry from the  $\text{HO}_x$  and  $\text{Cl}_x$  cycles. From observations it is known, that the behaviour of the daily ozone cycle is different at other altitudes (Haefele et al., 2008; Studer et al., 2013). A comprehensive understanding can be achieved with further analysis of chemistry-climate simulation data.

In summary we conclude that a correction of the diurnal biases in stratospheric ozone series is  
535 more difficult than expected (e.g. regional anomalies, temperature tides) and more needed than expected (e.g. strong daily cycle of polar summer ozone). Further work on this research topic by observers and modellers may lead step-by-step to a reliable correction of diurnal biases of ozone series. As spin-off, we expect progress in modelling, instrument and retrieval technique.

## Appendix A

### 540 Definitions of reaction rates

Definitions of missing reaction coefficients of Eq. (8). The dagger symbol ( $\dagger$ ) denotes vibrationally excited molecules.



*Acknowledgements.* The research leading to these results has received funding from the European Community's Seventh Framework Programme (FP7/2007–2013) under grant agreement no. 284421 (see Article II.30. of the Grant Agreement). Also we would like to acknowledge the International Space Science Institute, Bern, Switzerland (ISSI Team #246, Characterizing Diurnal Variations of Ozone for Improving Ozone Trend Estimates, <http://www.issibern.ch/teams/ozonetrend/>) and the Ubuntu Foundation for providing open source software.

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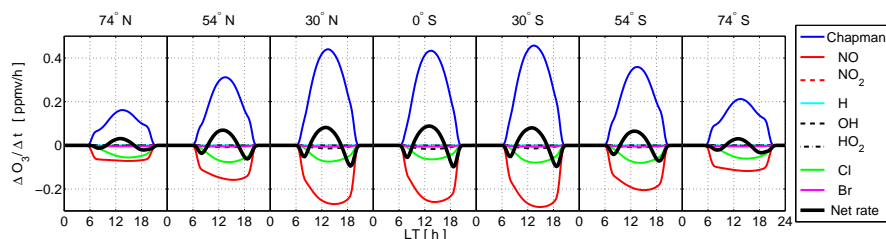
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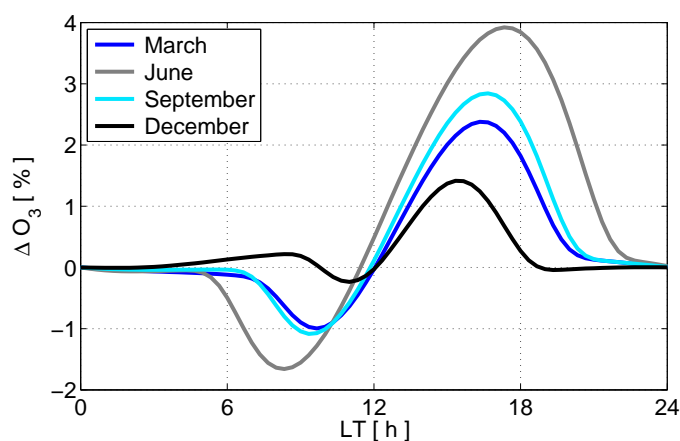
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**Table 1.** Relation of  $D_T$  and  $D_{O_3}$  in the tropics. The anomalies with respect to the tropical mean are given by  $D'_T$  and  $D'_{O_3}$ . Regions A and B are indicated in Fig. 7c.

Region	$D_{O_3}/D_T$	$D'_{O_3}/D'_T$
Tropical mean:	0.34 ppmv/1.34 K	
African coast (A):	0.20 ppmv/2.57 K	-0.14 ppmv/ 1.23 K
Atlantic ocean (B):	0.40 ppmv/0.76 K	0.06 ppmv/-0.58 K

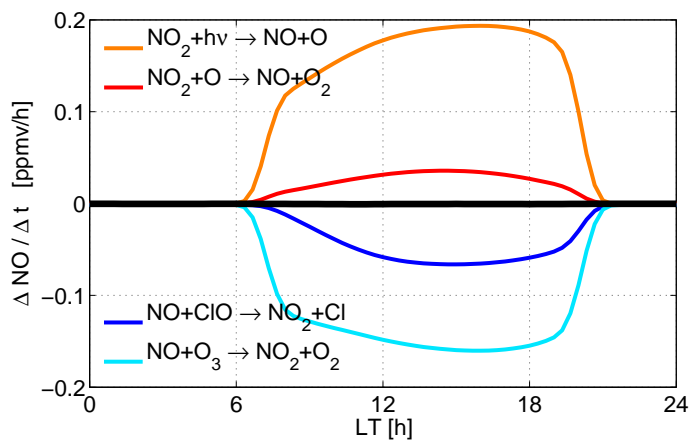


**Fig. 1.** Zonal mean ozone conversion rates at different latitudes modelled by WACCM (21 March, 5 hPa). Sunrise and sunset take place at about 06:00 LT and 18:00 LT. The conversion rates of the Chapman cycle (blue lines) include production and depletion of ozone according to Reactions (R3–R7). The red and green lines indicate conversion rates based on the NO and Cl catalytic cycles. Further catalytic cycles were taken into account by WACCM but showed negligible ozone conversion rates (catalytic cycles of  $NO_2$ , H, OH,  $HO_2$  and Br). The net conversion rate (black line) is the sum of the conversion rates of all cycles.

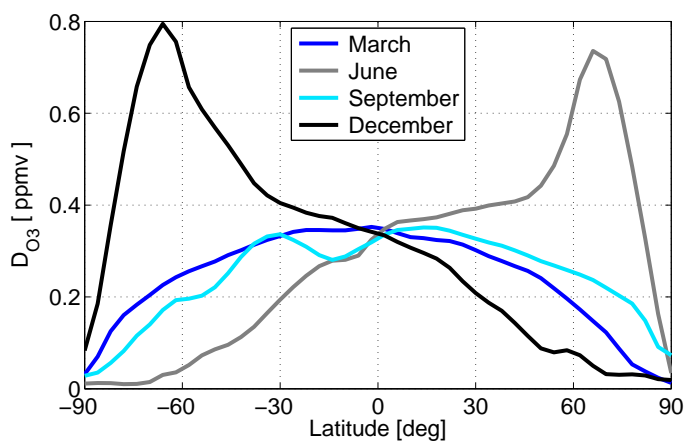


**Fig. 2.** Temporal evolution and seasonal diversity of the daily ozone cycle at  $50^\circ$  N. The data are zonal means modelled by WACCM at 5 hPa. The lines show changes in ozone relative to midnight in monthly means of March, June, September and December relative to the midnight value of ozone.

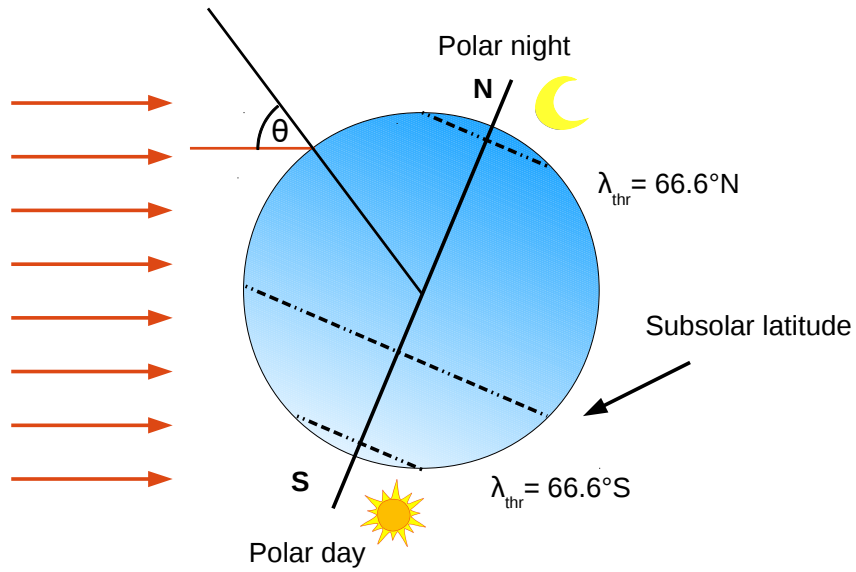




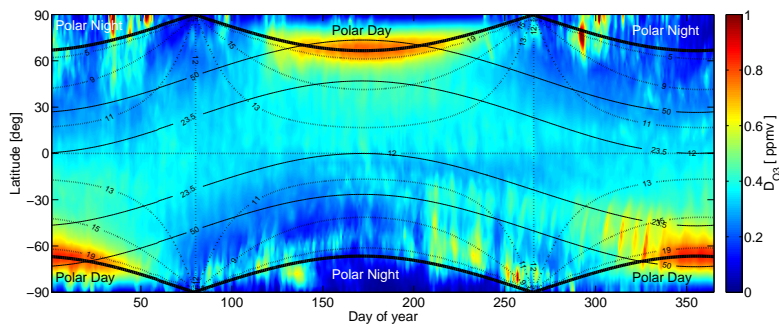
**Fig. 3.** Zonal mean conversion rates of NO at 5 hPa, 21 March, 50° N. The corresponding reactions are given next to the conversion lines. The net conversion rate of NO (black line) show negligible conversion rate since common NO changes at the solar terminator are from 0 to approximately 10 ppbv. Reactions with absolute conversion rates less than 5 ppbv h<sup>-1</sup> are omitted.



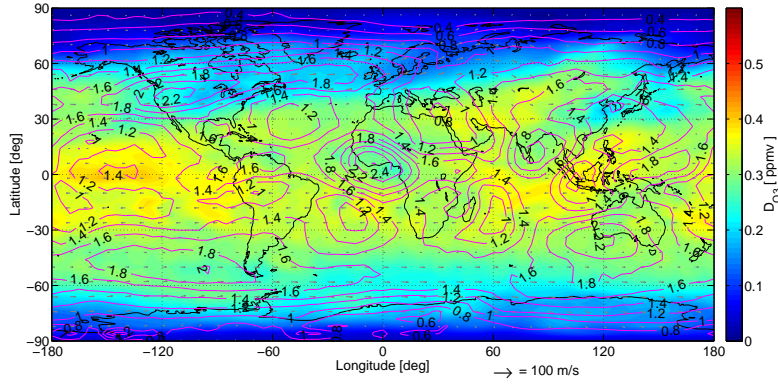
**Fig. 4.** Zonal mean of peak-to-valley difference  $D_{O_3}$  (Eq. 2) as function of latitude at 5 hPa modelled by WACCM. The lines show monthly mean  $D_{O_3}$  for March, June, September and December.



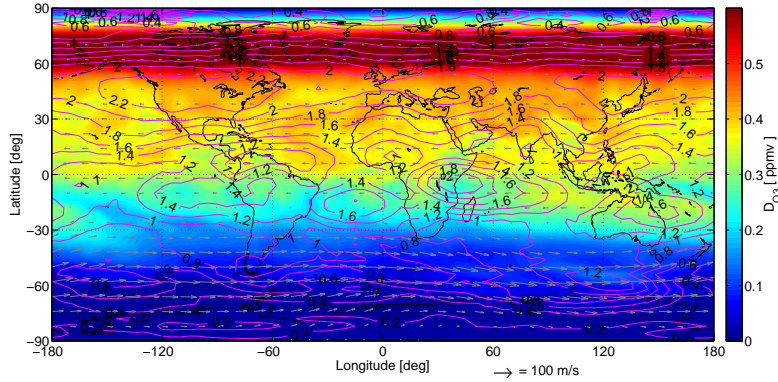
**Fig. 5.** Geometry of incident solar radiation and Earth rotation axis at the December solstice. The threshold latitudes  $\lambda_{thr}$  separates the polar day and polar night regions from regions of alternating day and night periods.  $\Theta$  is to the solar zenith angle.



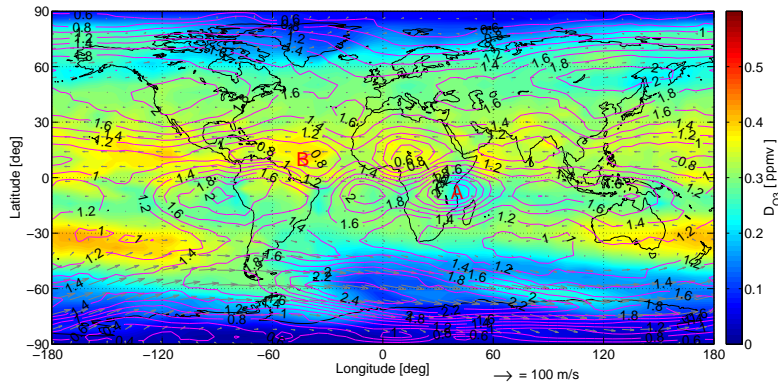
**Fig. 6.** Seasonal changes in the strength of the daily ozone cycle at 5 hPa. The figure shows zonal mean  $D_{O_3}$  as function of day of year and latitude derived from WACCM. The overlaid solid contours refer to the solar zenith angle at noon. Thick solid lines depict the borders of the polar day and polar night. Dashed contours show the sunshine duration given in hours.



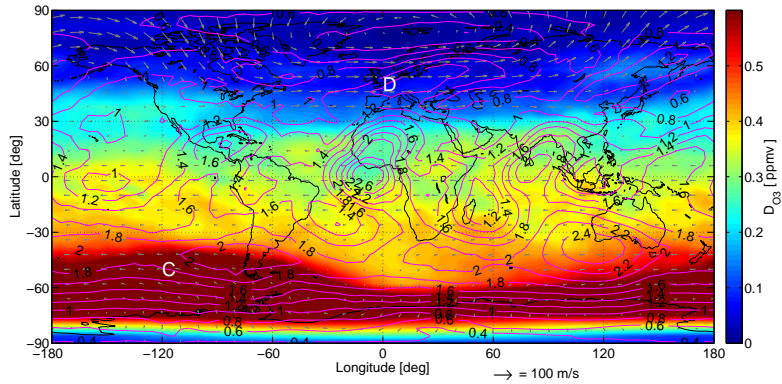
**Fig. 7a.** Monthly mean modelled  $D_{O_3}$  at 5 hPa in March as function of longitude and latitude. A daytime average of horizontal wind (grey arrows) from 09:00–15:00 LT is overlaid and a reference for wind speed is given near the x-axis. The magenta contour lines refer to the peak-to-valley difference  $D_T$  in temperature given in K.



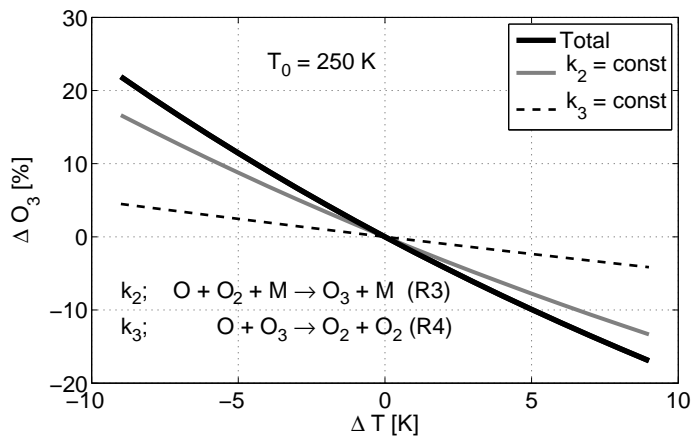
**Fig. 7b.** Similar to Fig. 7a with simulation data of June.



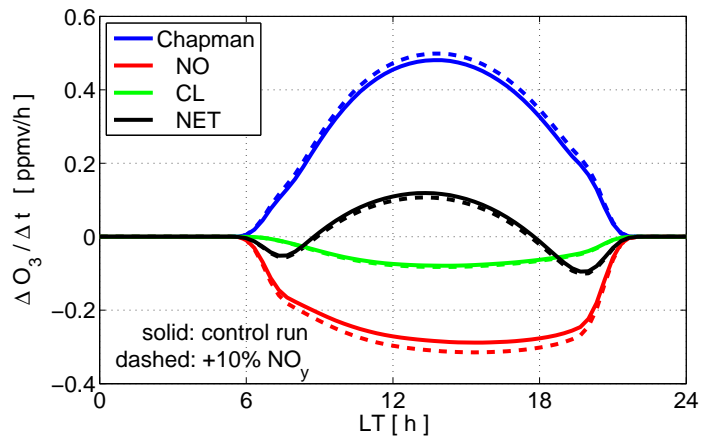
**Fig. 7c.** Similar to Fig. 7a with simulation data of September. Temperature-correlated effects of region A and B are discussed in Sect. 6.1.



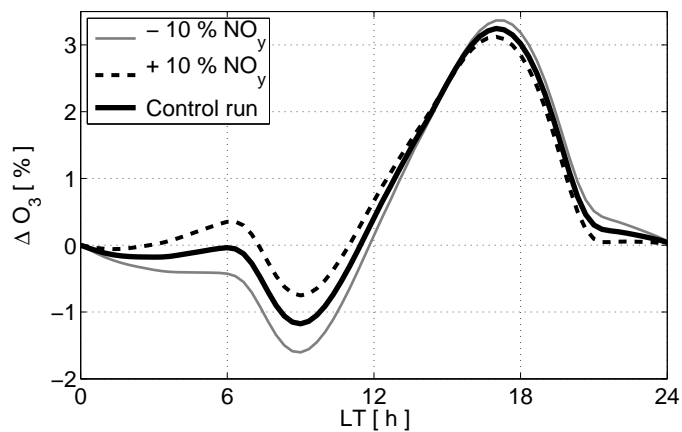
**Fig. 7d.** Similar to Fig. 7a with simulation data of December. The effects at Region C and D are discussed in Sects. 6.2 and 6.3.



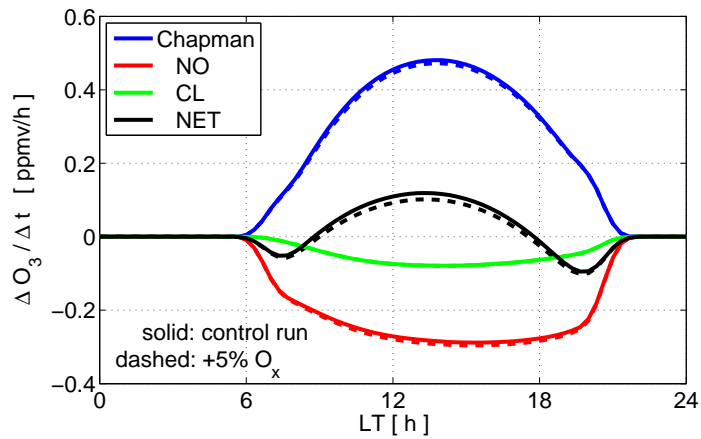
**Fig. 8.** Ozone variations induced by temperature variations in a pure oxygen atmosphere. “Total” denotes the result obtained for variable, temperature-dependent reaction rate coefficients  $k_2(T)$  and  $k_3(T)$ .



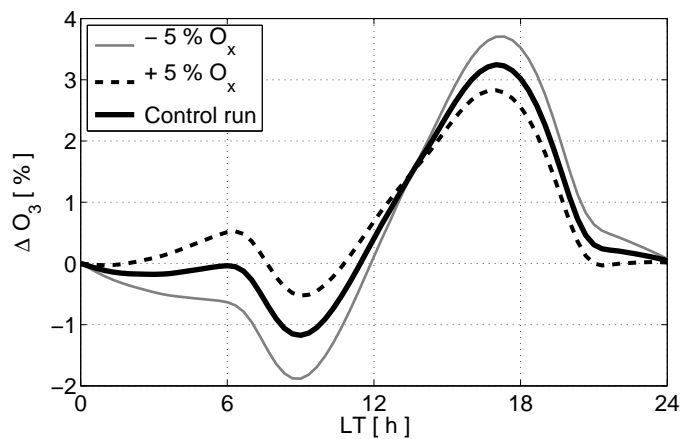
**Fig. 9.** Zonal mean conversion rates of ozone derived from a simulation with 10% more initial  $NO_y$  concentration and a control simulation at 1 January, 5 hPa, 22° N.



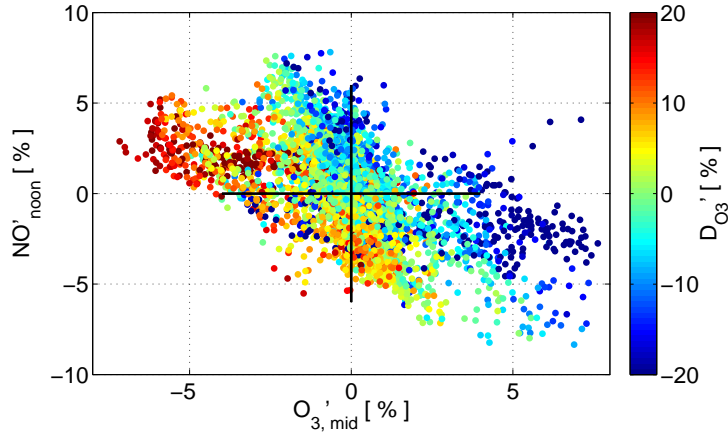
**Fig. 10.** Daily ozone cycles derived from simulations with  $\pm 10\%$  initial  $NO_y$  concentration and a control simulation shown at 1 January, 5 hPa, 22° N.



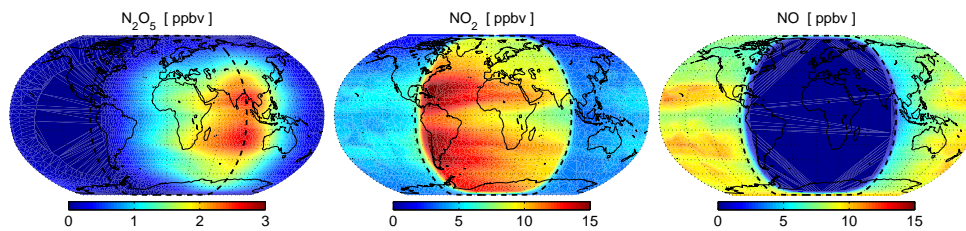
**Fig. 11.** Zonal mean conversion rates of ozone derived from a simulation with 5 % more initial  $O_x$  concentration and a control simulation at 1 January, 5 hPa, 22° N.



**Fig. 12.** Daily ozone cycles derived from simulations with  $\pm 5\%$  initial  $O_x$  concentration and a control simulation shown at 1 January, 5 hPa, 22° N.



**Fig. 13.** Dependence of  $D_{O_3}$  anomalies to anomalies in  $NO_{\text{noon}}$  (NO at 12:00 LT) and  $O_{3,\text{mid}}$  ( $O_3$  at 24:00 LT) for monthly means of March, June, September and December. Data points are taken from  $\lambda_{\text{sub}} \pm 30^\circ$  where  $\lambda_{\text{sub}}$  is the subsolar latitude. Two anti-correlations appear: increased (decreased)  $NO_{\text{noon}}$  or  $O_{3,\text{mid}}$  mostly lead to decreasing (increasing)  $D_{O_3}$ .



**Fig. 14.** Snapshot of  $N_2O_5$ ,  $NO_2$  and  $NO$  as simulated by WACCM at 5 hPa, 21 March, 0:00 UT. The solar zenith angle of  $90^\circ$  (dashed black lines) marks approximately the solar terminator. The nightside is in the centre of the maps. The reservoir  $N_2O_5$  is assembled during nighttime by Reaction (R12) until it peaks at the solar terminator. During daytime  $N_2O_5$  is dissociated into  $NO_2$  and  $NO_3$  by Reactions (R14) and (R13).  $NO_2$  in turn is photodissociated into  $NO$  and  $O$  or reacts with  $O$  by Reactions (R9) and (R10). Both reactions provide reactive  $NO$  which plays a major role for the daily ozone cycle at 5 hPa.