Investigating the Links Between Ozone and Organic
 Aerosol Chemistry in a Biomass Burning Plume from a
 Prescribed Fire in California Chaparral

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# 20 Abstract

Within minutes after emission, complex photochemistry in biomass burning smoke plumes 21 22 can cause large changes in the concentrations of ozone  $(O_3)$  and organic aerosol (OA). Being 23 able to understand and simulate this rapid chemical evolution under a wide variety of 24 conditions is a critical part of forecasting the impact of these fires on air quality, atmospheric 25 composition, and climate. Here we use version 2.1 of the Aerosol Simulation Program (ASP) to simulate the evolution of O<sub>3</sub> and secondary organic aerosol (SOA) within a young biomass 26 27 burning smoke plume from the Williams prescribed burn in chaparral, which was sampled 28 over California in November 2009. We demonstrate the use of a method for simultaneously

accounting for the impact of the unidentified intermediate volatility, semi-volatile, and 1 2 extremely low volatility organic compounds (here collectively called "SVOCs") on the formation of OA (using the Volatility Basis Set) and  $O_3$  (using the concept of mechanistic 3 reactivity). We show that this method can successfully simulate the observations of O<sub>3</sub>, OA, 4 NO<sub>x</sub>, C<sub>2</sub>H<sub>4</sub>, and OH to within measurement uncertainty using reasonable assumptions about 5 6 the average chemistry of the unidentified SVOCs. These assumptions were: (1) a reaction rate constant with OH of  $\sim 10^{-11}$  cm<sup>3</sup>/ s; (2) a significant fraction (up to  $\sim 50\%$ ) of the RO<sub>2</sub> + NO 7 reaction resulted in fragmentation, rather than functionalization, of the parent SVOC;  $(3) \sim 1.1$ 8 9 molecules of  $O_3$  were formed for every molecule of SVOC that reacted; (4) ~60% of the OH that reacted with the unidentified NMOC was regenerated as HO<sub>2</sub>, and (5) that ~50% of the 10 11 NO that reacted with the SVOC peroxy radicals was lost, presumably to organic nitrate 12 formation. Additional evidence for the fragmentation pathway is provided by the observed 13 rate of formation of acetic acid, which is consistent with our assumed fragmentation rate. 14 However, the model overestimates PAN formation downwind by about 50%, suggesting the 15 need for further refinements to the chemistry. This method could provide a way for 16 classifying different smoke plume observations in terms of the average chemistry of their SVOCs, and could be used to study how the chemistry of these compounds (and the  $O_3$  and 17 18 OA they form) varies between plumes.

### 19 **1** Introduction

Biomass burning is a major source of atmospheric trace gases and particles that impact air 20 21 quality and climate (e.g., Crutzen and Andreae, 1990; van der Werf, 2010; Akagi et al., 2011). 22 Within minutes after emission, rapid and complex photochemistry within the young biomass burning smoke plumes can lead to significant increases in the concentrations of secondary 23 24 pollutants such as ozone (O<sub>3</sub>, e.g. Mauzerall et al., 1998; Goode et al., 2000; Hobbs et al., 2003; Pfister et al., 2006; Lapina et al., 2006; Val Martin et al., 2006; Yokelson et al., 2009; 25 26 Jaffe and Widger, 2012; Akagi et al., 2012; 2013), peroxyacetyl nitrate (PAN, e.g. Jacob et al., 1992; Alvarado et al., 2010, 2011; Fischer et al., 2014), and organic aerosol (OA, e.g. 27 Hobbs et al., 2003; Grieshop et al., 2009a,b; Yokelson et al., 2009; Hennigan et al., 2011; 28 29 Heringa et al., 2011; Vakkari et al., 2014), while other smoke plumes can show little to no formation of O<sub>3</sub> (e.g. Alvarado et al., 2010, Zhang et al., 2014) or OA (e.g. Akagi et al., 30 2012). Understanding the atmospheric chemistry of these young smoke plumes, especially 31 32 which conditions can lead to the secondary formation of O<sub>3</sub>, PAN, and OA, is thus critical to

understanding the impact of these plumes on atmospheric composition and the resulting 1 2 impacts on air quality, human health, and climate. However, global- and regional-scale 3 Eulerian models of atmospheric chemistry artificially dilute biomass burning emissions into 4 large-scale grid boxes, which can result in large errors in the predicted concentrations of O<sub>3</sub> 5 and aerosol species downwind (e.g., Alvarado et al., 2009; Zhang et al., 2014). In contrast, plume-scale Lagrangian models allow us to examine the chemical and physical 6 7 transformations within these concentrated plumes in detail and can be used to develop 8 parameterizations for this aging process for coarser models (e.g., the parameterizations of 9 Vinken et al., 2011 and Holmes et al., 2014 for ship plumes).

10 Our understanding of the formation of ozone within biomass burning plumes is still poor, due 11 both to the limited observational data available on  $O_3$  formation in smoke plumes and the highly variable results seen in the available observations. Several aircraft and surface studies 12 13 of the chemistry of young biomass burning smoke plumes have found significant formation of 14 O<sub>3</sub> within smoke plumes. For example, Baylon et al. (2014) reported  $\Delta O_3/\Delta CO$  from 0.4% to 15 11%, corresponding to  $O_3$  enhancements of 3.8 to 32 ppbv in 19 wildfire plumes samples at 16 Mt. Bachelor Observatory. They note that plumes that have low values of  $\Delta O_3/\Delta CO$  can still correspond to significant  $O_3$  enhancements in concentrated plumes, with one event with a 17 18  $\Delta O_3/\Delta CO$  value of 0.81% corresponding to an  $O_3$  enhancement of 17 ppby. Akagi et al. 19 (2013) found significant O<sub>3</sub> formation ( $\Delta O_3/\Delta CO$  from 10-90%) within two hours for all of the South Carolina prescribed fires studied, and Parrington et al. (2013) found values of 20 21  $\Delta O_3/\Delta CO$  increased from 2.0±0.8% in boreal biomass burning plumes less than 2 days old 22 over Eastern Canada to 55±29% in plumes that were more than 5 days old. Similarly, Andreae 23 et al. (1994) found that aged plumes (over 10 days old) from the biomass burning regions of 24 South America and Africa had  $\Delta O_3/\Delta CO$  values between 20-70%. However, other studies, 25 mainly in boreal regions, have found little formation or even depletion of O<sub>3</sub> in some young 26 biomass burning plumes (e.g., Alvarado et al., 2010). This low  $O_3$  formation is likely due to a 27 combination of low emissions of  $NO_x$  from the boreal fires (Akagi et al., 2011), sequestration 28 of NO<sub>x</sub> in PAN and other organic nitrates (e.g., Jacob et al., 1992; Alvarado et al., 2010, 29 2011), and reduced rates of photochemical reactions due to aerosol absorption and scattering 30 (e.g. Jiang et al., 2012). Similarly, some studies have shown that fires can contribute to high 31 surface  $O_3$  events that exceed the US air quality standard for  $O_3$  (e.g., Jaffe et al., 2013), while 32 other studies suggest that this enhanced surface  $O_3$  is only present when the biomass burning 33 emissions mix with anthropogenic pollution (Singh et al., 2012; Zhang et al., 2014). However,

even given the observed variability among fires, it is likely that biomass burning has an impact on the concentrations of tropospheric  $O_3$ . For example, the recent review of Jaffe and Widger (2012) estimated that biomass burning could contribute 170 Tg of  $O_3$  per year, accounting for 3.5% of all global tropospheric  $O_3$  production. However, Sudo and Akimoto (2007) estimated that over a third of tropospheric  $O_3$  came from free troposphere chemical production due to biomass burning outflow from South America and South Africa.

7 The NO<sub>x</sub> emitted by biomass burning is rapidly converted into a wide variety of inorganic 8 nitrate (i.e. HNO<sub>3(g)</sub> and total aerosol inorganic nitrate, or NO<sub>3(p)</sub>) and organic nitrate species 9 (i.e. alkyl nitrates (RONO<sub>2</sub>) and peroxy nitrates (RO<sub>2</sub>NO<sub>2</sub>), including PAN; Jacob et al., 1992; Yokelson et al., 2009; Alvarado et al., 2010, 2011; Akagi et al., 2012). The rate at which this 10 11 conversion occurs and the relative production of inorganic nitrate, alkyl nitrates, and peroxy 12 nitrates are a key control of the impact of the biomass burning on  $O_3$  production and 13 atmospheric composition. Once NO<sub>x</sub> is converted to inorganic or organic nitrate, it is 14 generally unavailable for further O<sub>3</sub> formation near the fire source. Furthermore, while 15 conversion of NO<sub>x</sub> into inorganic nitrate (HNO<sub>3(g)</sub> + NO<sub>3(p)</sub>) is generally irreversible (except 16 for the slow reaction of HNO<sub>3(g)</sub> with OH), peroxy nitrate species like PAN can act as 17 thermally unstable reservoirs of  $NO_x$ , allowing transport of  $NO_x$  in the upper atmosphere far 18 from the original source and then producing  $NO_x$  via thermal decomposition as the airmass 19 descends to the surface (e.g., Fischer et al., 2010). This regenerated  $NO_x$  can thus impact  $O_3$ 20 formation far from the original source.

21 In addition, photochemistry within the smoke plume can rapidly oxidize non-methane organic 22 compounds (NMOCs), both those that were emitted in the gas phase and those emitted in the 23 particle phase, lowering their vapor pressure and thus leading to the formation of secondary 24 organic aerosol (SOA). As with  $O_3$  and PAN formation, the formation of SOA in smoke plumes is highly variable, with the ratio of OA to CO<sub>2</sub> increasing by a factor of 2-3 downwind 25 of some biomass burning fires (e.g. Hobbs et al., 2003; Grieshop et al., 2009a,b; Yokelson et 26 27 al., 2009; Hennigan et al., 2011; Heringa et al., 2011; Vakkari et al., 2014), while in others it can stay constant or even decrease (e.g. Capes et al., 2008; Akagi et al., 2012). For cases 28 29 where little net SOA formation was observed, it is likely that the NMOCs were still being oxidized. However, in these cases the fragmentation of the organic species after oxidation 30 (leading to higher volatility products) is likely more common than functionalization (i.e. the 31 32 addition of oxygen to the organic species, leading to lower volatility products).

Plume-scale Lagrangian parcel models can be used to investigate the evolution of  $O_3$ , PAN, 1 2 and OA in smoke plumes in detail, as their relatively simple parameterizations of plume 3 dispersion and transport allow detailed simulation of the chemical and microphysical 4 processes taking place within the young smoke plumes (e.g., Mauzerall et al., 1998; Mason et 5 al., 2001, 2006; Jost et al., 2003; Trentmann et al., 2005; Alvarado and Prinn, 2009; Arnold et al., 2014; Heilman et al., 2014). Previous plume-scale modeling studies have greatly advanced 6 7 our understanding of these transformations. Mauzerall et al. (1998) found that O<sub>3</sub> production 8 within biomass burning plumes was limited by the concentration of  $NO_x$  and that the 9 formation and subsequent degradation of peroxy acetyl nitrate (PAN) helped to maintain  $NO_x$ 10 concentrations. Mason et al. (2001) and Trentmann et al. (2003) showed that oxygenated 11 volatile organic compounds (OVOCs) were critical to the formation of  $O_3$  within the smoke 12 plumes. More recent work has suggested heterogeneous chemistry and currently unidentified 13 organic species as potential explanations for the rapid formation of O<sub>3</sub> and organic aerosol 14 seen within some smoke plumes (Trentmann et al., 2005; Mason et al., 2006; Alvarado and 15 Prinn, 2009).

16 The Aerosol Simulation Program (ASP) was developed to simulate the formation of ozone 17 and secondary organic aerosol (SOA) within young biomass burning plumes (Alvarado, 18 2008). ASP v1.0 was used to simulate several African and North American plumes (Alvarado 19 and Prinn, 2009) and to simultaneously simulate the chemistry, dynamics, and radiative 20 transfer within a smoke plume using a high-resolution three-dimensional plume model (Alvarado et al., 2009). Alvarado and Prinn (2009) showed while their initial ASP v1.0 21 22 simulations underestimated the formation of both OH and O<sub>3</sub> in the Timbavati savannah fire 23 smoke plume (Hobbs et al., 2003), if the OH concentration in ASP v1.0 was fixed at the estimated value of  $1.7 \times 10^7$  molecules/cm<sup>3</sup> then the model was able to reproduce the 24 observed concentrations of O<sub>3</sub>. This suggested that the model was missing an important 25 26 source of OH, and they proposed a heterogeneous reaction of NO<sub>2</sub> on aerosol particles 27 producing HONO, followed by the photolysis of HONO into NO and OH, as a candidate for 28 the missing source of OH within the smoke plume. Alvarado and Prinn (2009) also found that 29 including only SOA formation from known SOA precursors (mainly aromatic species like 30 toluene) underestimated the concentrations of organic aerosol observed downwind by  $\sim 60\%$ , 31 suggesting that the model was missing a large source of SOA. They proposed that the large 32 amount of gas-phase organic compounds that were unidentified by the then current 33 measurement techniques (Christian et al., 2003; Warneke et al., 2011) could include the precursors for the missing SOA. Assuming these compounds had SOA yields similar to
 monoterpenes gave the observed SOA formation.

In this paper, we describe recent updates to the gas-phase chemistry and secondary organic 3 4 aerosol (SOA) formation modules in ASP. We use this updated version (ASP v2.1) to simulate the chemical evolution of a young biomass burning smoke plume sampled over 5 6 California in November near San Luis Obispo (the Williams fire, Akagi et al., 2012). The 7 analysis of the  $O_3$ , PAN, and OA evolution in biomass burning plumes is complicated by the fact that a large fraction (30-50% by carbon mass, Christian et al., 2003; Warneke et al., 2011) 8 9 of the NMOCs present in smoke plumes are unidentified, and thus their oxidation chemistry is not well known. 10

Furthermore, while there was clear secondary formation of O<sub>3</sub> and PAN within the Williams fire plume, the dilution-corrected amount of OA in the plume decreased slightly (Akagi et al., 2012). Most current methods for modeling the OA evolution in smoke plumes lead to significant secondary growth of the OA (e.g., Grieshop et al., 2009a), but we need instead to modify ASP v2.1 to simulate both this slight loss of OA and the chemical formation of O<sub>3</sub>, PAN, and other gas-phase species.

Here we present a method for simultaneously accounting for the impact of the unidentified organic compounds (here collectively called "SVOCs") on the formation of OA and O<sub>3</sub>, drawing on the Volatility Basis Set approach (e.g., Robinson et al., 2007) for modeling OA and the concept of the mechanistic reactivity of a mixture of organic compounds (e.g., Carter, 1994). We show that this method can successfully simulate the Williams fire plume observations using reasonable assumptions about the chemistry of the unidentified SVOCs.

23 Section 2 describes the updates to the gas-phase chemistry and secondary organic aerosol 24 formation modules of ASP for version 2.1. Section 3 discusses our validation of the gas-phase 25 chemistry in ASP v2.1 against data from a smog chamber (Carter et al., 2005). Section 4 26 describes the Williams fire and summarizes the available observations of the smoke plume 27 from Akagi et al. (2012). Section 5 discusses the results of the ASP simulation of the 28 Williams fire, including sensitivity tests to investigate the chemistry of the unidentified 29 SVOCs and their impacts on  $O_3$ , PAN, other trace gases, and OA, while Section 6 gives the 30 conclusions of our study and directions for future work.

## **2** Updates to the Aerosol Simulation Program (ASP)

An overview of ASP v1.0 is given by Alvarado and Prinn (2009), and the routines are described in detail in Alvarado et al. (2008). Here we briefly discuss the modules of ASP that have not changed since Alvarado and Prinn (2009) in Section 2.1 before describing the updates to the gas-phase chemistry (Section 2.2) and SOA formation (Section 2.3) routines for ASP v2.1.

# 7 2.1 ASP Modules

8 Aerosols are represented in ASP by a single moving-center sectional size distribution, where 9 the aerosol concentrations are distributed over increments in radius space (Jacobson 1997, 10 2002, 2005). ASP includes modules to calculate aerosol thermodynamics, gas-to-aerosol mass 11 transfer (condensation/evaporation), and coagulation of aerosols. The thermodynamics 12 module in ASP uses the Mass Flux Iteration (MFI) method of Jacobson (2005) to calculate the 13 equilibrium concentration of gas and aerosol species. Equilibrium constants for the inorganic 14 electrolyte reactions match those of Fountoukis and Nenes (2007). Binary activity coefficients 15 of inorganic electrolytes are calculated using the Kusik-Meissner method (Kusik and Meissner, 1978; Resch, 1995), as are the mean activity coefficients. The water content of 16 17 inorganic aerosols is calculated with an iterative routine that calculates water activities for 18 aqueous solutions of a single electrolyte using a formula based on the Gibbs-Duhem equation 19 (Steele, 2004). Steele (2004) and Alvarado (2008) found this approach compares well with 20 other inorganic aerosol thermodynamics models such as ISORROPIA (Nenes et al., 1998; 21 Fountoukis and Nenes, 2007).

Mass transfer between the gas and aerosol phases is calculated in ASP using a hybrid scheme where the flux-limited kinetic equations governing the condensation/evaporation of  $H_2SO_4$ and organic species are integrated, whereas NH<sub>3</sub>, HNO<sub>3</sub>, and HCl are assumed to be in equilibrium (Alvarado, 2008). Aerosol coagulation is calculated using the semi-implicit scheme of Jacobson (2005) with a Brownian coagulation kernel.

# 27 2.2 Gas-phase Chemistry Updates

The gas-phase chemistry within the ASP model for Version 2.1 has been completely revised from ASP v1.0, which used the CalTech Atmospheric Chemistry Mechanism (CACM, Griffin et al., 2005). The revised ASP v2.1 gas phase chemical mechanism includes 1608 reactions between 621 species. Examples of the gas-phase species input file and the reaction
 mechanism input file for ASP v2.1, along with other key chemical input files, are included in
 the Supplemental Material.

All inorganic gas-phase chemistry within ASP v2.1 was updated to follow the IUPAC recommendations (Atkinson et al., 2004; updated data downloaded from <u>http://iupac.pole-</u> <u>ether.fr/</u>, accessed June 2012). We also tested the JPL recommendations (Evaluation #17, Sander et al., 2011) for these rate constants, but found that the differences between the recommendations generally made little difference to the model simulations, and as the IUPAC values were closer to those in ASP v1.0, these values were used.

All gas-phase chemistry for organic compounds containing 4 carbons or less has been
"unlumped," i.e. the chemistry for each individual organic compound is explicitly resolved.
This was done by following the reactions of the Leeds Master Chemical Mechanism (MCM)
v3.2 (http://mcm.leeds.ac.uk/MCM/, accessed June 2012; Jenkin et al., 1997, 2003; Saunders
et al., 2003; Bloss et al., 2005) for these species.
The chemical mechanism of isoprene within ASP v2.1 has been updated to follow the Paulot

16 et al. (2009a,b) isoprene scheme, as implemented in GEOS-Chem and including corrections 17 based on more recent studies (e.g., Crounse et al., 2011, 2012). The (lumped) chemistry for all other organic compounds in ASP has been updated to follow the Regional Atmospheric 18 19 Chemistry Mechanism (RACM) v2 (Goliff et al., 2013). We chose RACM2 over the SAPRC-20 07 (Carter, 2010) and CB05 (Yarwood et al., 2005) lumped chemical mechanisms as the 21 treatment of peroxy radicals in the RACM2 mechanism was more similar to the treatment in the Leeds MCM and the Paulot isoprene scheme, resulting in a more consistent chemical 22 23 mechanism for ASP v2.1.

24 Photolysis rates are calculated offline using the Tropospheric Ultraviolet and Visible (TUV) 25 radiation model version 5.0 (Madronich and Flocke, 1998) for 15 minute increments, which are then linearly interpolated in ASP. Alvarado and Prinn (2009) assumed a "clear sky" 26 27 radiation field that ignored the effect of aerosol absorption and scattering on the calculated 28 photolysis rates. Here we instead estimate the time-dependent aerosol,  $O_3$ ,  $SO_2$ , and  $NO_2$ 29 concentrations within the smoke plumes and calculate their effect on the photolysis rates at 30 different heights within the plume. In the TUV simulations, we assume no clouds and that the 31 initial smoke plume AOD at 330 nm decreases due to dilution assuming a background concentration of  $\sim 0$ , and the aerosol is assumed to have a constant (both with time and 32

8

wavelength) single scattering albedo (SSA) of 0.9 based on the review of AERONET biomass 1 2 burning smoke optical property retrievals by Reid et al. (2005a). We also dilute the initial 3 plume concentrations of the trace gases NO<sub>2</sub> and SO<sub>2</sub> assuming a background concentration of 4 ~0, as these species can also absorb ultraviolet and visible (UV-VIS) light and thus can 5 impact photolysis rates. For the photolysis rate calculations only,  $O_3$  is assumed to be 0 initially and increased after 15 minutes to a constant value based on the observed formation of 6 7  $O_3$  within the smoke plume. Section 5.1 has more details on the specific approach and 8 quantitative values used for the Williams Fire.

# 9 2.3 SOA Formation Updates

10 We have updated the SOA formation module to follow the semi-empirical Volatility Basis Set 11 (VBS) model of Robinson et al. (2007). Our implementation of this scheme followed the 12 approach used by Ahmadov et al. (2012) to link the VBS scheme with the RACM chemical 13 mechanism within WRF-Chem. We use 9 surrogates or "bins" for semi-volatile, intermediate 14 volatility, low volatility, and extremely low volatility organic compounds (hereafter 15 collectively referred to as "SVOCs" for simplicity) as in Dzepina et al., 2009, rather than only 16 4 as in Ahmadov et al. (2012). The saturation mass concentration at 300 K (C\*, see Robinson 17 et al., 2007) of each SVOC bin differs by a factor of 10, and covers the range from 0.01 to  $1.0 \times 10^6 \,\mu\text{g/m}^3$ . Note that "SVOC" as defined in this paper includes both semivolatile organic 18 compounds ( $C^*$  between  $10^{-2}$  and  $10^3 \mu \text{g m}^{-3}$ ) and intermediate volatility organic compounds 19 20 (species with C\* between  $10^4$  and  $10^6 \mu \text{g m}^{-3}$ ) as defined in Dzepina et al. (2009), but we refer to both of these species classes collectively as "SVOCs" rather than as "S/IVOCs" as in 21 22 Dzepina et al. (2009) for simplicity. Following the Model to Predict the Multiphase 23 Partitioning of Organics (MPMPO) of Griffin et al. (2003, 2005) and Pun et al. (2002), we 24 assumed that an aqueous phase and a mixed hydrophobic organic phase are always present in 25 the aerosol. Partitioning of organics between the gas and hydrophobic phase is governed by 26 Raoult's law (assuming that all hydrophobic-phase OM is quasi-liquid and can dissolve 27 organics as in Pankow, 1994a,b), while partitioning of organics into the aqueous phase is 28 governed by Henry's law. Following Pun et al. (2002), we assumed that (1) there is no 29 interaction between the aqueous phase inorganic ions and the aqueous phase organics, and 30 thus no organic salt formation, and (2) the activity coefficients of the organic ions (formed by 31 the dissociation of organic acids) are equivalent to those of the corresponding molecular 32 solute. We further assumed that the pH of the aqueous phase is dominated by the strong

inorganic acids and bases, and that the pH effects of the dissociating organic acids are
 negligible.

Like most organic compounds, SVOCs will react with OH. Most mechanisms for this
chemistry (e.g., Robinson et al., 2007; Dzepina et al., 2009; Grieshop et al., 2009a,b;
Ahmadov et al., 2012) parameterize this chemistry by assuming that the SVOCs react with
OH to form a lower volatility SVOC, as in the reaction:

7 
$$\operatorname{SVOC}_{i} + \operatorname{OH} \xrightarrow{k_{OH}} \mu \operatorname{SVOC}_{i-n}$$
 (R1)

8 where  $\mu$  is the relative mass gain due to oxidation (e.g. via O addition),  $k_{OH}$  is the reaction rate 9 with OH, and *n* is the "volatility shift", or by how many factors of 10 to lower the *C*\* of the 10 product with each OH reaction. This simplified chemistry can be extended to account for the 11 fact that the SVOCs could fragment during oxidation, leading to higher volatility products:

12 
$$\operatorname{SVOC}_{i} + \operatorname{OH} \xrightarrow{k_{OH}} \mu(1-\alpha) \operatorname{SVOC}_{i-n} + \mu\alpha \operatorname{SVOC}_{i+1} + \alpha \operatorname{VOC}_{j}$$
 (R2)

13 where  $\alpha$  is the fraction of SVOC<sub>i</sub> that fragments into SVOC<sub>i+1</sub> and VOC<sub>i</sub>. Shrivastava et al. 14 (2013) used a similar approach to show that adding SVOC fragmentation to WRF-Chem 15 simulations of the Mexico City Plateau improved the model's ability to simulate the observed 16 concentrations of SOA. However, the highly simplified chemistry of Reactions R1 or R2 is 17 not appropriate for situations where reactions with the SVOC compounds are a potentially 18 significant sink of OH, such as in a concentrated smoke plume. Thus in ASP v2.1, the 19 average, lumped chemistry of the SVOCs is instead parameterized in a more realistic manner 20 for a generic organic species, following the idea of "mechanistic reactivity" (e.g., Carter, 1994; Bowman and Seinfeld, 1994a,b; Seinfeld and Pandis, 1998). After reaction with OH 21 SVOCs produce peroxy radicals (RO<sub>2</sub>), which can react with NO to form NO<sub>2</sub> and HO<sub>2</sub>, 22 23 thereby regenerating OH and forming O<sub>3</sub>. Reactions R3 and R4 show this more general 24 chemical mechanism for the SVOCs:

25 
$$\text{SVOC}_i + \text{OH} \longrightarrow \text{RO}_2$$
 (R3)

26 
$$\begin{array}{c} \operatorname{RO}_{2,i} + \chi \operatorname{NO} \xrightarrow{\kappa_{RO_{2,i}}} \mu (1 - \alpha) \operatorname{SVOC}_{i-n} + \mu \alpha \operatorname{SVOC}_{i+1} \\ + \alpha \operatorname{VOC}_{i} + \beta \operatorname{NO}_{2} + \delta \operatorname{HO}_{2} \end{array}$$
(R4)

where  $k_{RO_{2}}$  is assumed to be  $4.0 \times 10^{-12}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> based on the reaction rate for the 1 2 peroxy radicals from long-chain alkanes and alkenes with NO in RACM2 (Goliff et al., 2013). 3 We can see that  $\gamma$ - $\beta$  is the number of NO<sub>x</sub> lost (implicitly via the addition of a nitrate group to 4 the product SVOCs), 1- $\delta$  is the number of HO<sub>x</sub> lost, and  $\beta + \delta$  is the number of O<sub>3</sub> made per 5 reaction (by subsequent reactions of  $NO_2$  and  $HO_2$  to generate  $O_3$ ). For example, the values 6 for long-chain alkanes (HC8) in the RACM<sub>2</sub> mechanism (Goliff et al., 2013) would be  $\chi = 1, \delta$ 7 = 0.63, and  $\beta$  = 0.74, such that 0.26 NO<sub>x</sub> and 0.37 HO<sub>x</sub> are lost and 1.37 O<sub>3</sub> are formed per reaction. Note that the mechanism of Reactions R3 and R4 is still highly simplified: we 8 9 assume that reaction of SVOC with OH always produces a RO<sub>2</sub> radical, and that the RO<sub>2</sub> produced does not react with HO<sub>2</sub> or another RO<sub>2</sub>. Also note that Reactions R3 and R4 10 11 represent the average chemistry of the unknown species collectively, and may not apply to 12 any individual species in that mixture. Our purpose is less to detail all the possible reactions 13 of the unidentified SVOCs and more to explore how their average chemistry might affect  $O_3$ 14 and OA evolution in smoke plumes. The specific combinations of parameters for Reactions 15 R3 and R4 that were evaluated in this study are shown in Table 2.

16 We also adjusted the calculation of aerosol water content to use the "kappa" ( $\kappa$ ) 17 parameterization of organic hygroscopicity (Petters and Kreidenweis, 2007) for the lumped 18 SVOCs. In this parameterization, the hygroscopicity parameter  $\kappa$  for the organic species is 19 defined as:

$$20 \qquad \frac{1}{a_w} = 1 + \kappa \frac{V_{s,i}}{V_{w,i}} \tag{1}$$

where  $a_w$  is the activity of water in the solution (equal to the relative humidity at equilibrium),  $V_{s,i}$  is the volume of the dry organic solute *i* and  $V_{w,i}$  is the volume of water in the solution. The water content calculated for each organic species, along with that calculated for the inorganic solution ( $V_{w,inorg}$  see Section 2.1 above) are then combined using the Zdanovskii, Stokes, and Robinson (ZSR) approximation (Zdanovskii, 1948; Stokes and Robinson, 1966):

$$26 V_w = \frac{a_w}{1 - a_w} \sum_i \kappa_i V_{s,i} + V_{w,inorg} (2)$$

## **3** ASP Photochemistry Evaluated with Smog Chamber Data

To evaluate the performance of the updated photochemical mechanism in ASP v2.1 in predicting the formation of ozone, several test simulations were performed to compare the results of the mechanism to laboratory smog chamber data. This comparison provides us with a baseline for interpreting the results of our simulation of  $O_3$  formation in the Williams Fire in Section 5. The data used for the comparison came from the EPA chamber of Carter et al. (2005), which consists of two collapsible 90 m<sup>3</sup> FEP Teflon reactors (chambers A and B).

8 Table S1 in the Supplemental Material shows the temperature and initial reactant 9 concentrations used in our model to simulate each chamber study. All model simulations were 10 performed at a pressure of 1000 mbar, a relative humidity of 1%, and a CH<sub>4</sub> concentration of 11 1800 ppby. The temperature and concentration data were provided by William P.L. Carter 12 (http://www.cert.ucr.edu/~carter/SAPRC/SAPRCfiles.htm, accessed March 2014). The EPA 13 chamber runs used an 8 compound surrogate for ambient VOC concentrations, consisting of 14 formaldehyde, ethylene, propene, trans-2-butene, n-butane, n-octane, toluene, and m-xylene 15 (Carter et al., 1995, 2005). The initial concentrations of HONO were extrapolated from CO-NO<sub>x</sub> and n-butane-NO<sub>x</sub> runs to account for the potential chamber radical source (Carter et al., 16 17 2005).

18 Table S2 in the Supplemental Material presents the rates of off-gassing (i.e., re-emission of 19 HCHO and other species from the walls of the reaction chamber), wall reaction rates, and selected photolysis rates for the chamber experiments considered here. The off-gassing rate 20 21 for HONO was determined as the rate that enabled the SAPRC-99 chemical mechanism 22 (Carter, 2000) to best predict the  $O_3$  formation observed in CO-air, HCHO-air and CO-HCHO-air experiments performed within the chamber (Carter et al., 2005). The rate in 23 24 Chamber A was found to be slightly higher than that in Chamber B, so different values are 25 used for the chambers. The off-gassing rate of HCHO was chosen to match the low but 26 measurable amount of formaldehyde found even in pure air and  $CO-NO_x$  experiments in the 27 chamber. Heterogeneous wall loss reaction rates for O<sub>3</sub>, NO<sub>2</sub>, and N<sub>2</sub>O<sub>5</sub> were also estimated 28 from reactor observations (Carter et al., 2005). The photolysis rate of  $NO_2$  in the chambers 29 was measured directly, and scaling factors for the other photolysis rates were calculated by 30 Carter et al. (2005) from the relative spectral intensity of the arc lamp.

Following Carter et al. (2005), we evaluated the ability of our mechanism to simulate the total amount of NO oxidized and O<sub>3</sub> formed in the experiments, measured as  $\Delta([O_3] - [NO])_t =$ 

 $([O_3]_t - [NO]_t) - ([O_3]_{initial} - [NO]_{initial})$ . The hourly results of the comparisons for  $\Delta([O_3] - [O_3]_{initial})$ . 1 2 [NO]) are presented in Figure 1. We can see that the ASP v2.1 mechanism tends to underestimate  $\Delta([O_3] - [NO])$ , with a mean absolute bias of -24.6 ppbv and a mean 3 4 normalized bias of -22.4%. Comparisons of the ASP calculations for O<sub>3</sub>, NO, and NO<sub>x</sub> (not 5 shown) show that this model underestimate of  $\Delta([O_3] - [NO])$  is primarily due to the model underestimating O<sub>3</sub> formation, rather than underestimating the loss of NO or NO<sub>x</sub>. Similarly, 6 7 the ASP v2.1 calculations for the concentrations of the organic gas species matches well with 8 the chamber measurements (not shown) except for formaldehyde (HCHO), where the 9 secondary formation of HCHO appears to be underestimated. Figure 2 shows the bias in 10  $\Delta([O_3] - [NO])$  versus the initial ratio of the mixing ratio of reactive organic gases (ROGs, 11 e.g., the concentration of the surrogate gases in ppm C) to the mixing ratio of NO<sub>x</sub> (in ppm 12 N). We can see that the bias is between 0% and -10% for ROG/NO<sub>x</sub> ratios greater than 30, but 13 increases to -40% to -50% for "high NO<sub>x</sub>" cases (ROG/NO<sub>x</sub> ratios  $\ll$  20). For comparison, 14 the initial ROG/NO<sub>x</sub> ratio in biomass burning smoke can range from  $\sim$ 10-100 ppm C/ppm N 15 (Akagi et al., 2011, assuming that the total NMOC mass is 1.6 times the mass of carbon in 16 these compounds). Both the general underestimation of  $\Delta([O_3] - [NO])$  and the increase of the negative bias at low  $ROG/NO_x$  concentrations is consistent with the behaviors of the 17 18 SAPRC-99 (Carter et al., 2005), SAPRC-07 (Carter, 2010), and CB05 (Yarwood et al., 2005) 19 mechanisms evaluated against the EPA chamber data. Carter (2010) noted that this under-20 prediction of O<sub>3</sub> at low ROG/NO<sub>x</sub> ratios was apparently linked to the presence of aromatics in 21 the surrogate mixture, with comparisons of SAPRC-07 with EPA chamber runs with a non-22 surrogate mixture showing a positive bias of about +25% for cases with low ROG/NO<sub>x</sub> ratios.

23 4 Williams Fire Data

The Williams Fire (34°41'45" N, 120°12'23" W) was sampled by the US Forest Service 24 (USFS) Twin Otter aircraft from 10:50-15:20 LT on November 17, 2009 (Akagi et al., 2012). 25 The fire burned approximately 81 hectares of scrub oak woodland understory and coastal sage 26 scrub. Skies were clear all day and RH was low (11-26%) with variable winds (2-5 m s<sup>-1</sup>). The 27 Williams Fire smoke plume showed significant secondary production of O<sub>3</sub> and PAN, but the 28 29 enhancement ratio of OA to CO<sub>2</sub> decreased slightly downwind (Akagi et al., 2012). In this study, we use the processed data from Akagi et al. (2012) that provided concentrations of 30 several trace gases and OA measured during several quasi-Lagrangian transects of the 31 32 Williams fire. Full details on the measurements made and the processing of the data for the

1 plume transects are given in Akagi et al. (2012); those used in this study are briefly described

2 here.

### 3 4.1 Airborne Fourier Transform InfraRed spectrometer (AFTIR)

4 The University of Montana AFTIR system and the instruments described below were 5 deployed on a US Forest Service (USFS) Twin Otter aircraft. The AFTIR (Yokelson et al., 6 1999, 2003) was used to measure 21 gas-phase species, including carbon dioxide (CO<sub>2</sub>), 7 carbon monoxide (CO), nitric oxide (NO), nitrogen dioxide (NO<sub>2</sub>), nitrous acid (HONO), 8 peroxy acetyl nitrate (PAN), ozone (O<sub>3</sub>), glycolaldehyde (HCOCH<sub>2</sub>OH), ethylene (C<sub>2</sub>H<sub>4</sub>), 9 formaldehyde (HCHO), acetic acid (CH<sub>3</sub>COOH), and formic acid (HCOOH). IR spectra were 10 collected at 1 Hz. "Grab samples" of air were selected by closing the values for 1-2 min to 11 allow signal averaging, and the resulting IR spectra were analyzed to quantify all detectable 12 compounds.

# 13 4.2 Aerosol Mass Spectrometer (AMS)

14 An Aerodyne compact time-of-flight (CToF) aerosol mass spectrometer (herein referred to as 15 AMS) measured aerosol chemical composition in a repeating cycle for 4 out of every 12 s 16 during flight, including within the smoke plume. The AMS has been described in great detail 17 elsewhere (Drewnick et al., 2005; Canagaratna et al., 2007). An isokinetic particle inlet 18 sampling fine particles with a diameter cut-off of a few microns (Yokelson et al., 2007; 19 Wilson et al., 2004) supplied the AMS. As the AMS does not measure super-micron particles, 20 the inlet transmission should not have affected the results. A collection efficiency of 0.5 21 (Huffman et al., 2005; Drewnick et al., 2003; Allan et al., 2004) was applied to the AMS data, 22 which were processed to retrieve the mass concentration for the major non-refractory particle 23 species: OA, non-sea salt chloride, nitrate, sulfate, and ammonium, with <36% uncertainty.

#### 24 4.3 Other Measurements

The ambient three-dimensional wind velocity, temperature, relative humidity, and barometric pressure were measured at a frequency of 1-Hz with a wing-mounted AIMMS-20 probe (Aventech Research, Inc., Beswick et al., 2008). A non-dispersive infrared instrument NDIR (LiCor model 7000) measured CO<sub>2</sub> (at 0.5 Hz) from the third channel on the isokinetic particle inlet that also supplied the AMS.

#### 1 5 ASP Simulation of Williams Fire

# 2 5.1 ASP Setup

As in Alvarado and Prinn (2009), we simulated the Williams fire smoke plume using ASP within a simple Lagrangian parcel model following Mason et al. (2001). We assume a Lagrangian parcel of fixed vertical extent (H, here assumed to be 1 km) and down-trajectory length (L), but variable cross-trajectory width y(t). The temperature and pressure of the parcel are assumed to be constant. The full continuity equations for the Lagrangian parcel model are then

9 
$$\frac{dC_q}{dt} = -\frac{4K_y}{\left(y_o^2 + 8K_yt\right)} \left(C_q - C_q^a\right) - \frac{v_d}{H}C_q + \left(\frac{dC_q}{dt}\right)_{cond} + \left(\frac{dC_q}{dt}\right)_{chem}$$
(3)

$$10 \qquad \frac{dn_i}{dt} = -\frac{4K_y}{\left(y_o^2 + 8K_yt\right)} \left(n_i - n_i^a\right) - \frac{v_d}{H}n_i + \left(\frac{dn_i}{dt}\right)_{cond} + \left(\frac{dn_i}{dt}\right)_{coag}$$
(4)

$$11 \qquad \frac{dc_{q,i}}{dt} = -\frac{4K_y}{\left(y_o^2 + 8K_yt\right)} \left(c_{q,i} - c_{q,i}^a\right) - \frac{v_d}{H} c_{q,i} + \left(\frac{dc_{q,i}}{dt}\right)_{cond} + \left(\frac{dc_{q,i}}{dt}\right)_{coag} + \left(\frac{dc_{q,i}}{dt}\right)_{chem}$$
(5)

where  $C_q$  is the concentration of gas-phase species (molecules cm<sup>-3</sup> air),  $n_i$  is the number concentration of particles in size bin *i* (particles cm<sup>-3</sup> air),  $c_{q,i}$  is the concentration of aerosol species *q* in size bin *i* (mol cm<sup>-3</sup> air),  $y_o$  is the initial plume width (m), and  $K_y$  represents the horizontal diffusivity of the atmosphere (m<sup>2</sup> s<sup>-1</sup>). The superscript *a* indicates the concentration of the given species in the atmosphere outside of the parcel (i.e., the background concentration).

18 The first term on the right-hand side of Equations 3-5 represents the effect of plume 19 dispersion on the concentrations. Note that  $y_o$  and  $K_y$  can be reduced to a single parameter, the 20 initial dilution time scale  $\tau_{mix,o}$ :

21 
$$-\frac{4K_{y}}{\left(y_{o}^{2}+8K_{y}t\right)} = -\frac{1}{\frac{y_{o}^{2}}{4K_{y}}+2t} = -\frac{1}{\tau_{mix,o}+2t}$$
(6)

The second term on the right hand side of Equations 3-5 is the effect of deposition on the 1 concentrations, where  $v_d$  is the deposition velocity (m s<sup>-1</sup>). We set the dry deposition velocity 2 equal to 0 for gas-phase species, as the plume did not touch the ground during the modeled 3 period, and use the size-dependent terminal velocity of the aerosol particles as the deposition 4 5 velocity for aerosol species assuming a 1 km thick plume. As submicron aerosol dominated the aerosol mass in the smoke plume, this deposition of aerosol species has a negligible effect 6 7 on the results, and given the low relative humidity during the Williams Fire, we also did not 8 include wet deposition of particles or gases. The remaining terms represent the change in gas-9 and particle-phase concentrations due to net mass transfer between the gas and aerosol phases 10 (cond), coagulation of particles (coag), and chemical production and loss (chem).

11 The observed changes in CO mixing ratio were used to determine the best-fit model initial dilution time scale ( $\tau_{mix,0} = 106.9$  s) as well as upper and lower limits of the time scale 12 13  $(\tau_{mix,o}(0) = 15.0 \text{ s and } 212.2 \text{ s, respectively})$ , as shown in Figure 3. Note that this dilution 14 fitting procedure neglects the impact of the chemical production and loss of CO on the 15 observed concentrations, but modeling results with an inert tracer suggest that this error is < 16 4%, much smaller than the dilution uncertainty represented by our upper and lower limit estimates. The temperature of the plume was set at a constant value of 288.4 K, pressure of 17 18 880 hPa, and relative humidity of 15.7% based on the observations of Akagi et al. (2012). The 19 parcel was assumed to be emitted at 11:00 Pacific Standard Time (PST) and the model was 20 integrated for 5 hours. The integration of the different terms of the continuity Equations 3-5 were operator split for computational efficiency. The chemistry and mixing time steps were 1 21 22 s for the first 10 minutes of model integration due to the rapid dilution and chemical changes 23 during this period, and were 60 s thereafter. The aerosol thermodynamics, condensation, and 24 coagulation time steps were 60 s throughout.

25 The initial and background concentrations for the gas-phase inorganic and NMOC species are 26 in Table S3 of the Supplemental Material, and Table S4 gives the initial and background 27 concentrations used for the aerosol species. Initial and background concentrations of trace 28 gases and aerosols in the smoke were taken from observations of the Williams Fire (Akagi et 29 al., 2012), where available. Emission ratios for other species were calculated using the 30 literature reviews of Akagi et al. (2011) and Andreae and Merlet (2001). Other background concentrations were taken from runs of the GEOS-Chem model (Bey et al., 2001), run for our 31 32 period as in Fischer et al. (2014).

The volatility distribution for the POA was taken from the wood smoke study of Grieshop et 1 2 al. (2009a,b). Table 1 shows the POA total mass fractions used for wood smoke in Grieshop 3 et al. (2009a) and the values used in this study for the Williams fire. At the measured 4 temperature (288.4 K) and initial concentration of organic aerosol in the Williams fire smoke plume (849 µg m<sup>-3</sup>), the Grieshop et al. (2009a) POA volatility distribution implies that 81% 5 of the total mass of SVOC species  $SVOC_1$  to  $SVOC_7$  is in the aerosol phase, leaving 200 µg 6 m<sup>-3</sup> of SVOC species in the gas phase. Note that the May et al. (2013) POA volatility 7 8 distribution (not shown in Table 1) is more volatile than Grieshop et al. (2009a), with 65% of 9 the total mass of SVOC species  $SVOC_1$  to  $SVOC_7$  in the aerosol phase, leaving about 460 µg m<sup>-3</sup> of SVOC species in the gas phase. 10

11 However, Grieshop et al. (2009a) and May et al. (2013) were only able to measure species with a saturation mass concentration ( $C^*$ ) of  $10^4 \ \mu g \ m^{-3}$  or less. Furthermore, Akagi et al. 12 13 (2011) provide emission factors for unidentified non-methane organic compounds (NMOCs) 14 from savannah/grassland and chaparral fires, with unidentified NMOCs estimated to be equal 15 in mass to the identified species. The savannah/grassland estimate is about twice as large as 16 the chaparral estimate, as fewer species have been identified in chaparral fires. Here we use the savannah/grassland estimate to calculate an emission ratio of 0.195 g unidentified 17 NMOC/g CO, but assign this value an uncertainty of ~50%, consistent with the lower 18 chaparral estimate. This implies that there is about  $2000\pm1000 \ \mu g \ m^{-3}$  of unidentified NMOCs 19 20 in the gas-phase of the smoke. So to be consistent with the EFs of Akagi et al. (2011) and the volatility distributions of Grieshop et al. (2009a) and May et al. (2013), there still needs to be 21 another 1500-1800  $\mu$ g m<sup>-3</sup> of unidentified NMOCs initially in the plume with C\*>10<sup>4</sup>  $\mu$ g m<sup>-3</sup> 22 23 which the techniques used by Grieshop et al. (2009) and May et al. (2013) would not have 24 been able to measure. These remaining unidentified NMOCs were included as  $SVOC_8$  ( $C^* =$ 10<sup>5</sup> µg m<sup>-3</sup>), as shown in Table 1 and Tables S3 and S4 of the Supplemental Material. Below 25 we also discuss sensitivity tests that were performed to see how the results change if the 26 remaining unidentified NMOCs are considered as SVOC<sub>9</sub> ( $C^* = 10^6 \ \mu g \ m^{-3}$ ) instead, as well 27 as for an increase or decrease of the estimated unidentified SVOC concentrations by 50%. 28

For all organic species, we assumed a constant  $\kappa = 0.04$ , corresponding to an O/C ratio of 0.25 (Jimenez et al., 2009) that is typical of biomass burning organic aerosol (Donahue et al., 2011). Since the relative humidity in the Williams fire plume was very low, this assumption had little impact on our results.

1 The initial smoke aerosol size distribution was assumed to be a log-normal with a geometric 2 mean diameter  $D_g$  of 0.10 µm and a standard deviation  $\sigma$  of 1.9 (unitless) based on Reid and 3 Hobbs (1998) for flaming combustion of Brazilian cerrado, which structurally is a similar mix 4 of shrubs and grasses as in the Williams fire. The initial total number concentration of aerosol particles  $(2.34 \times 10^6 \text{ particles cm}^3)$  was calculated such that the initial total organic aerosol 5 6 mass matched the  $\Delta OA/\Delta CO_2$  emission ratio from Akagi et al. (2012). The evolution of the 7 aerosol size distribution with time was simulated by ASP v2.1 using a center-moving 8 sectional size distribution with 10 bins, 8 bins for particles with volume-equivalent spherical 9 diameters between 0.05 and 2.0  $\mu$ m, one for particles with diameters smaller than 0.05  $\mu$ m, 10 and one for particles with diameters greater than  $2 \,\mu m$ .

11 Photolysis rates were calculated offline using TUV v5.0 (Madronich and Flocke, 1998) as noted in Section 2.2 above. The smoke aerosols were assumed to dilute with time according to 12 13 the three dilution rates derived above (see Figure 3). In the TUV simulations, we assumed no clouds and an initial AOD of 8.0 at 330 nm (consistent with the ASP v2.1 calculated initial 14 15 extinction coefficient and the assumed plume thickness of 1 km), which decreases due to 16 dilution assuming a background concentration of  $\sim 0$ . As noted above, we assumed a constant single scattering albedo of 0.9 based on Reid et al. (2005a). We also assumed initial plume 17 18 and background concentrations of the trace gases NO<sub>2</sub> (initial 295 ppby, background 0 ppby) and  $SO_2$  (initial 50.9 ppby, background 0 ppby). For the photolysis rate calculations only,  $O_3$ 19 was assumed to be 0 initially and increased after 15 minutes to a constant value of 100 ppbv 20 21 to account for the observed formation of  $O_3$  within the smoke plume. The overhead ozone 22 column was assumed to be 278 Dobson Units (DU), based on the average of values from the 23 Ozone Monitoring Instrument (OMI) for 11/16/2009 (276 DU) and 11/18/2009 (280 DU) 24 (accessed through http://jwocky.gsfc.nasa.gov/teacher/ozone overhead.html on June 2012, 25 now at http://ozoneag.gsfc.nasa.gov/tools/ozonemap/ozone overhead). The surface albedo was assumed to be 0.035 based on the GEOS-Chem data file for the 0.5°×0.667° North 26 27 American grid for November 1985, which is in turn based on data from the Total Ozone Mapping Spectrometer (TOMS). Photolysis rates were calculated for three altitudes: just 28 29 above the plume (i.e., at 2.1 km altitude), near the middle of the plume (1.6 km), and near the 30 bottom of the plume (1.1 km). This, combined with the three dilution rates, gave nine estimates of photolysis rates versus time. The nine values for the NO<sub>2</sub> photolysis rate  $(J_{NO_2})$ 31 were compared with the clear sky (no aerosol) case in Figure 4. In the middle of the plume 32

(1.6 km),  $J_{NO_2}$  was reduced from an initial clear-sky value of  $9 \times 10^{-3}$  s<sup>-1</sup> to an initial value of 1  $2 \times 10^{-3}$  s<sup>-1</sup>. However, by 15 minutes after emission  $J_{NO_2}$  in the middle of the plume increased 2 to  $6-8.5 \times 10^{-3}$  s<sup>-1</sup> depending on the dilution rate, showing that the plume reduced photolysis 3 rates by 5-33% after the initial, rapid dilution of the plume.  $J_{NO_{\gamma}}$  was slightly enhanced above 4 the plume (initially  $1.1 \times 10^{-2}$  s<sup>-1</sup>) over the clear sky value, and the photolysis rates were lowest 5 in the bottom of the plume. As expected, the impact of the plume was larger for lower dilution 6 7 rates, but the difference between the different dilution rates was largest for the bottom of the 8 plume. Note that, while our assumption of a constant SSA is questionable as aerosol 9 absorption is likely to change with both smoke age and with wavelength, our use of three 10 dilution rates and three altitudes in the plume results in a wide range of photolysis rates used 11 in this study, which can also account for uncertainties in the aerosol optical properties and 12 other parameters used to calculate the photolysis rates.

#### 13 **5.2** ASP Results with No Unidentified SVOC Chemistry

We first ran ASP assuming the unidentified SVOCs emitted by the fire are unreactive.
Deficiencies in these simulations provide information on what the average chemistry of the unidentified SVOCs needs to be in order to explain the observations.

17 Figure 5 shows the ASP v2.1 results and Akagi et al. (2012) observations for the enhancement 18 ratios (EnR, mol/mol) of O3 and PAN in the Williams fire smoke plume versus time after 19 emissions. The EnR is defined as the ratio of the enhancement of a species X within the 20 smoke plume ( $\Delta X = C_x - C_x^a$ , Akagi et al., 2011) to the enhancement of a less reactive species, such as CO<sub>2</sub> or CO. The choice of whether to use CO or CO<sub>2</sub> in the denominator of the EnR 21 22 was made on a species-by-species basis to match the choices made in Table 2 of Akagi et al. 23 (2012), which were in turn chosen to minimize the impact of measurement and sampling 24 errors on the EnRs. We can see that the range of dilution rates and photolysis rates simulated 25 for this case capture the general rate of the secondary formation of  $O_3$  and PAN, but ASP v2.1 appears to be overestimating the rate of formation of these compounds. This is in contrast to 26 27 Alvarado and Prinn (2009), who found that ASP v1.0 underestimated the much faster O<sub>3</sub> 28 formation in the Timbavati savannah fire smoke plume by about 50%.

The ASP v2.1 value for the best estimate dilution and photolysis case (i.e., the best-fit dilution combined with the middle of the plume photolysis rates, plotted as a solid black line in Figure

1 5) at 4.5 hours for  $\Delta O_3/\Delta CO$  is 0.116 mol/mol which is within the uncertainty associated with 2 the average value measured for the Williams fire  $(0.095\pm0.022)$ . This overestimate is similar 3 to the positive bias (~25%) of the SAPRC-07 mechanism versus the EPA smog chamber 4 results for low ROG/NO<sub>x</sub> ( $\leq 20$  ppb C/ppb N) ratios when aromatics are not part of the surrogate. As aromatics are a minor constituent in biomass burning smoke, and the  $ROG/NO_x$ 5 ratio for savannah/scrubland fires like the Williams fire (without including unidentified 6 7 species) is  $\sim 10$  ppb C/ppb N, we would expect the mechanism in ASP v2.1 to show a similar positive bias. ASP v2.1 predicts an "average" value of  $\Delta PAN/\Delta CO_2$  of 8.4×10<sup>-4</sup> at 4.5 hours 8 downwind, 65% larger than the observed value of  $(5.10\pm1.21)\times10^{-4}$ . 9

10 Figure 6 shows the ASP v2.1 results and observations for  $\Delta NO_x/\Delta CO_2$  and  $\Delta C_2H_4/\Delta CO$ 11 versus time after emission. Figure 6a shows that the  $\Delta NO_x/\Delta CO_2$  values are correctly simulated by ASP v2.1, with the best estimate dilution and photolysis case EnR of  $3.4 \times 10^{-4}$ 12 matching the observed value of  $4.6\pm2.3\times10^{-4}$  4 to 4.5 hours downwind. However, the 13 observations show a faster rate of decay in the first two hours after emission than is seen in 14 15 the model results. Figure 6b shows that the decay of  $C_2H_4$  is also well matched by the model 16 results, suggesting that the modeled OH is similar to the actual OH concentrations. This can 17 also be seen by comparing the modeled OH concentration for the best estimate dilution and 18 photolysis case  $(5.3 \times 10^6 \text{ molecules cm}^{-3})$  to that derived by Akagi et al. (2012) using the observed decay of  $C_2H_4$  (5.27±0.97×10<sup>6</sup> molecules cm<sup>-3</sup>). This is again in contrast with 19 20 Alvarado and Prinn (2009), who found that ASP underestimated the observed OH radical concentrations for the Timbavati smoke plume  $(1.7 \times 10^7 \text{ molecules cm}^{-3}, \text{Hobbs et al., 2003})$ . 21

We can explore this contrast further by looking at the rate of loss of HONO in the smoke 22 23 plume, shown in Figure 7. Note that unlike the previous figures, Figure 7 only shows the first hour after emission as the observations showed no detectable HONO further downwind. As 24 25 noted in Section 1, to explain the underestimate of O<sub>3</sub> and OH in the Timbavati fire, Alvarado and Prinn (2009) posited that a heterogeneous reaction of NO<sub>2</sub> to make HONO and HNO<sub>3</sub> was 26 27 taking place in that plume. However, the O<sub>3</sub> and OH results for the Williams fire show no 28 evidence of this chemistry, and the HONO decay seen in Figure 7 also shows little evidence 29 for a secondary source of HONO except for a few points within the first 12 minutes after 30 emission that have more HONO than is predicted by the model. While explaining the 31 discrepancy between the Williams and Timbavati results is beyond the scope of this paper, we 32 note that the Timbavati fire took place closer to the equator (24°S versus 35°N), earlier in the

1 year (7 September versus 17 November) than the Williams fire, and that the relative humidity 2 was higher as well (45.0% versus 15.7%). All of these differences would tend to increase 3 photolysis rates and the formation of OH. In addition, the higher actinic flux and RH in 4 Timbavati may have increased the speed of reactions for forming HONO from NO<sub>2</sub> that are 5 not included in standard chemical mechanisms, either via aqueous chemistry (Jacob et al., 2000), sunlight-activated humic acid surfaces (Stemmler et al., 2006, 2007), or photo-excited 6 7  $NO_2$  reacting with H<sub>2</sub>O (Ensberg et al., 2010). Though we find no evidence for secondary 8 HONO production in the Williams Fire data, this does not preclude that some HONO was 9 made, but remained below the AFTIR detection limit of 10 ppbv as the plume diluted.

10 Figure 8 shows the ASP results for two aldehydes, HCHO and glycoaldehyde (HCOCH<sub>2</sub>OH),

and two organic acids, formic acid (HCOOH) and acetic acid (CH<sub>3</sub>COOH), in the Williams fire plume in terms of EnRs to CO. We can see that ASP generally underestimates the formation of these species. Part of this underestimate may be due to errors in the chemical mechanism for known precursor compounds, as was seen for HCHO in the smog chamber results, but neglecting the chemistry of the SVOCs and their ability to form these smaller organic compounds is also likely responsible for this underestimate.

Figure 9 shows the modeled OA enhancement ratios ( $\Delta OA/\Delta CO_2$ , g/g) at 4.5 hours 17 downwind using the parameters listed in Table 2 in addition to the observed average OA 18 enhancement ratio  $(2.83 \pm 1.08 \times 10^{-3})$  and the modeled OA enhancement ratio for the case 19 where the chemistry of the unidentified SVOCs is not included  $(2.27 \times 10^{-3})$ . When SVOC 20 21 chemistry was not included, some of the original OA evaporated into the gas phase as the 22 plume diluted, and as there was no chemistry to make these SVOC species less volatile, they stayed in the gas phase leading to a net decrease in  $\Delta OA/\Delta CO_2$  with time. Changing the gas-23 phase concentrations of the unidentified SVOC by  $\pm 50\%$  has a small impact (~3%) on these 24 25 results, but the match between the model and observation could be improved by using a less 26 volatile POA distribution than that given by Grieshop et al. (2009a). However, the modeled 27 decrease without SVOC chemistry is larger (but still within the error bars) than the decrease 28 that was reported by Akagi et al. (2012). In addition, the assumption that the SVOCs do not 29 react is unrealistic – as large multi-functional organic compounds, they should have a 30 relatively fast reaction rate with OH (see Section 5.3 below). Thus in Sections 5.3 through 5.5 31 below we test different, more realistic implementations for the chemistry of these SVOCs.

# 1 5.3 OH Reaction Rate and Fragmentation Probability of the Unidentified 2 SVOCs

Here we evaluate the ability of the parameters from the original VBS paper of Robinson et al. 3 (2007), a study of SOA formation in wood smoke by Grieshop et al. (2009a,b), and the 4 5 implementation of the VBS scheme into WRF-Chem by Ahmadov et al. (2012) to simulate 6 the observed evolution of OA in the Williams fire plume. Table 2 shows the values for the 7 parameters in Reactions R3 and R4 that define these various SVOC mechanisms. Figure 9 8 shows that the SVOC mechanisms of Robinson et al. (2007) and Grieshop et al. (2009a,b) 9 overestimated the OA downwind by a factor of 1.8 and 3.7, respectively. This is primarily due 10 to their relatively large values for  $k_{OH}$ . For the Grieshop et al. (2009a,b) case, the 11 overestimation is also partially due to the large increase in mass ( $\mu$ ) and decrease in volatility (n) for each OH reaction. The OA formed using these mechanisms can be reduced by a further 12 25% if we assume the unidentified SVOCs are mainly the more volatile SVOC<sub>9</sub> ( $C^* = 10^6 \mu g$ 13 m<sup>-3</sup>) instead of SVOC<sub>8</sub> ( $C^* = 10^5 \ \mu g \ m^{-3}$ ), but are fairly insensitive to errors in the POA 14 volatility distribution. The scheme of Ahmadov et al. (2012), with  $k_{OH} = 10^{-11} \text{ cm}^3 \text{ molecule}^{-1}$ 15  $s^{-1}$ , was consistent with the uncertainty in the observations, but slightly higher than the 16 observed value  $(3.48 \times 10^{-3} \text{ versus the observed value of } 2.83 \pm 1.08 \times 10^{-3})$ . One approach to 17 further reduce the modeled OA would be to reduce  $k_{OH}$  even further. However, it seems 18 19 unlikely that the average OH reaction rate of the unidentified SVOC species would be less than  $10^{-11}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup>, as this is close to the reaction rate for large alkanes ( $k_{OH}$ (298K)) 20 =  $1.1 \times 10^{-11}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup>, Goliff et al, 2013) and the presence of other functional groups 21 22 (double bonds, aldehydes) would be expected to result in even higher reaction rates. For example,  $\alpha$ -pinene has a  $k_{OH}(298\text{K})$  of 5.0×10<sup>-11</sup> cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> (Goliff et al., 2013), and 23 other monoterpenes can have even faster reaction rates with OH. Thus, we think a more likely 24 25 explanation for the remaining overestimate is that a substantial fraction of the SVOC and OH reactions resulted in the fragmentation of the primary SVOC into more volatile compounds, 26 27 as in the 2D-VBS schemes of Jimenez et al. (2009) and Donahue et al. (2011).

Figure 9 shows that a  $k_{OH}$  of  $10^{-11}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> and a fragmentation probability of 50% (the "Half Fragmentation" case, see Table 2) provided a reasonably good match with the observed  $\Delta OA/\Delta CO_2$  4.5 hours downwind in the smoke plume ( $2.63 \times 10^{-3}$  versus the observed value of  $2.83\pm1.02\times10^{-3}$ ). Here we assumed that the SVOC fragmented into a small VOC and another, more volatile, SVOC, as in Reactions R3 and R4. While this is a relatively large

22

1 fragmentation probability, we note that it seems reasonable given the likely complex and 2 multifunctional nature of the unidentified SVOCs in a biomass burning smoke plume.

This fragmentation of the SVOCs after reaction with OH could also help to explain the 3 4 underestimate of aldehydes and organic acids seen in Section 5.2 when SVOC chemistry was neglected. For example, Figure 10 shows the ASP modeled EnR of acetic acid when we 5 6 assumed that the VOC fragment produced in Reaction R4 is acetic acid. This provided a 7 remarkably good match with the observed acetic acid formation, providing additional 8 evidence to support the fragmentation hypothesis. While we are not claiming to have proven 9 this is the source of the missing acetic acid, we note that the fragmentation hypothesis is thus consistent with the initial underestimate of the secondary formation of aldehydes and organic 10 11 acids in ASP v2.1. In addition, there is some evidence from biomass burning plume 12 observations that the formation of acetic acid and OA are inversely correlated with each other. In the Yucatan plume studied by Yokelson et al. (2009), a large amount of SOA was formed, 13 14 but acetic acid did not increase downwind, while in the Williams Fire, acetic acid increased, but OA did not. Thus, the limited amount of relevant airborne data in BB plumes is so far 15 consistent with the idea that the branching between functionalization and fragmentation in BB 16 17 plumes is variable and future work should identify what environmental and combustion 18 factors control the outcome.

19 An additional potential explanation for the SOA overestimate observed when the schemes of 20 the Robinson et al. (2007), Grieshop et al. (2009a,b), and Ahmadov et al. (2012) were used is that the OA was becoming more viscous and "glassy" with time (i.e., the particles had a lower 21 22 bulk diffusivity), thereby reducing the amount of quasi-liquid OA for SVOC compounds to 23 dissolve into (e.g., Kidd et al., 2014; Zaveri et al., 2014). There has been some recent 24 evidence for this process occurring in smoke plumes from biomass burning in the western US (A. Sedlacek, personal communication, March 2014). ASP v2.1 is not able to examine this 25 26 possibility in detail, but we do note that while the formation of "glassy" OA would reduce SOA formation, it likely would not increase the formation of aldehydes or organic acids as in 27 28 the fragmentation hypothesis.

# 29 5.4 HO<sub>x</sub> and NO<sub>x</sub> Chemistry of the Unidentified SVOCs

Section 5.3 showed that an SVOC mechanism following Reaction R2 with a  $k_{OH}$  of  $10^{-11}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup> and a fragmentation probability  $\alpha$  of up to 0.5 (the "Half Fragmentation" scheme in Table 2) could explain the observed evolution of OA in the Williams fire. However, neglecting the regeneration of  $HO_x$  and reaction of the peroxy radical with NO, as in Reaction R2, can lead to substantial underestimates of OH in the concentrated smoke plumes. This is because including Reaction R2 in ASP leads to a loss of OH with no corresponding regeneration of  $HO_2$ .

6 For example, Figure 11 shows that using the "Half Fragmentation" scheme reduced the ASP v2.1 estimates of the enhancement ratios of  $O_3$  and PAN downwind by 24% and 23%, 7 8 respectively (for the best estimate dilution and photolysis case, the black line in Figure 5), 9 while Figure 12 shows that it increased the ASP v2.1 estimates of  $C_2H_4$  and NO<sub>x</sub> downwind 10 by 33% and 151%, respectively. The "Half Fragmentation" O<sub>3</sub> and PAN estimates are more 11 consistent with the observations - the overestimate of PAN seen with the unidentified SVOC chemistry neglected in Section 5.2 has disappeared – but the large overestimate of  $NO_x$  (i.e., 12 13 underestimate of NO<sub>x</sub> loss in the plume presumably due to missing organic nitrogen 14 formation) is a serious problem. While this underestimate of NO<sub>x</sub> loss reduces the amount of 15 O<sub>3</sub> and PAN formed within five hours after emission, it would lead to large overestimates of 16 the impact of biomass burning plumes on  $O_3$  and PAN formation further downwind.

In addition, the chemistry of Reaction R2 is unrealistic, in that it implies a total loss of OH and no effect of the SVOC oxidation on NO<sub>x</sub>. One approach for addressing the first concern is to artificially regenerate the OH by simply adding it as an additional product to Reaction R2. While this makes sense as a "first do no harm" modeling approach to keep the gas-phase results the same regardless of the SVOC scheme, it is equally unrealistic, as it assumes that the SVOCs are oxidized without having any impact on NO<sub>x</sub> or HO<sub>x</sub>.

We prefer to approach this problem by recognizing that SVOCs are going to have impacts on the  $HO_x$  and  $NO_x$  radical budgets just like any other organic species, and that this chemistry can be approximated via Reactions R3 and R4. Including this more realistic, yet still simplified, chemistry allows ASP to simultaneously simulate the observed changes in OA and O<sub>3</sub> while still making reasonable, chemically plausible assumptions about the chemistry of the unidentified SVOCs emitted by the fire.

Our approach thus used the observations of OA,  $O_3$ , PAN,  $NO_x$ , and  $C_2H_4$  in the Williams fire as constraints on  $\beta$  and  $\delta$ , the amount of  $NO_2$  and  $HO_2$  produced in Reaction R4, respectively, while assuming that  $\chi$ =1 throughout. For the Williams Fire, we know from the above results that we want the optimized SVOC chemistry to (a) increase  $O_3$ , PAN, and OA formation as

little as possible, (b) increase the loss of NO<sub>x</sub>, either through organic nitrate formation or 1 2 increased OH concentrations, and (c) increase the OH concentration, thereby increasing  $C_2H_4$ loss. We found that using the parameters for large alkanes from RACM2 ( $\delta = 0.63$  and  $\beta =$ 3 4 (0.74) generally produced too much  $O_3$  and PAN and too little OH, but did a reasonable job for 5 NO<sub>x</sub> loss. However, attempts to increase OH by increasing  $\delta$  led to too much O<sub>3</sub> formation except for unrealistically low values of  $\beta$  (~0.1). Thus we set  $\delta = 0.6$  and reduced  $\beta$  to 0.5, 6 7 implying that 1.1 O<sub>3</sub> is formed per molecule of SVOC reacted. These parameters (arrived at 8 by trial and error) appear to give the best balance of reducing modeled  $NO_x$  and  $C_2H_4$  mixing 9 ratios while minimizing the increase in O<sub>3</sub>, PAN, and OA. The following section discusses 10 the ASP v2.1 model results for the Williams fire smoke plume using these parameters in 11 detail. Note that while slightly different, more precise parameters might provide a slightly 12 better match with observation, our goal here is not to derive exact best-fit parameters, but 13 rather to roughly classify the average chemistry of the SVOCs in the Williams fire smoke 14 plume, both for modeling this fire and for future comparisons with other smoke plumes.

#### 15 5.5 Results with Optimized SVOC Chemistry

16 Figure 9 shows the  $\Delta OA/\Delta CO_2$  4.5 hours downwind in the smoke plume using the optimized 17 SVOC chemistry discussed in Section 5.4. For the best estimate dilution and photolysis model case (i.e., the solid black line in Figure 5),  $\Delta OA/\Delta CO_2$  is  $2.75 \times 10^{-3}$  (g/g), very close to the 18 observed value of  $2.83 \pm 1.02 \times 10^{-3}$ . As in Section 5.2, changing the gas-phase concentrations 19 20 of the unidentified SVOC by  $\pm 50\%$  has a small impact (~3%) on these results, as does assuming that the unidentified SVOCs are mainly the more volatile SVOC<sub>9</sub> ( $C^* = 10^6 \text{ µg m}^{-3}$ ) 21 instead of SVOC<sub>8</sub> ( $C^* = 10^5 \ \mu g \ m^{-3}$ ). However, this result is still sensitive to the POA 22 volatility distribution – for example, moving all the mass in SVOC<sub>3</sub> ( $C^* = 1 \ \mu g \ m^{-3}$ ) to 23 SVOC<sub>7</sub> ( $C^* = 10^4 \,\mu \text{g m}^{-3}$ ) decreases the modeled  $\Delta OA/\Delta CO_2$  downwind by 12% for this case. 24 Figure 13 shows the enhancement ratios of  $O_3$  and PAN for the optimized SVOC chemistry. 25 26 and Figure 14 shows the results for  $NO_x$  and  $C_2H_4$ . The O<sub>3</sub> results are very similar to those 27 from Section 5.2 (where SVOC chemistry was not included in the model), while the PAN 28 results are slightly lower (and closer to the observed values) than in that case. For the best 29 estimate dilution and photolysis model case  $\Delta O_3/\Delta CO$  is 0.119 at 4.5 hours downwind, about 30 25% larger than the observed value of  $0.095\pm0.022$ , while the  $\Delta PAN/\Delta CO_2$  is now  $7.56\times10^{-4}$ 

31 at 4.5 hours downwind, about 48% larger than the observed value of  $(5.10\pm1.21)\times10^{-4}$ .

1 However, the O<sub>3</sub> and PAN values are reasonably close given the uncertainties in the 2 concentrations and in the estimated smoke ages for the observations.

3 The  $NO_x$  results were much improved from the "half frag" case in Section 5.4, with the best 4 estimate dilution and photolysis case  $\Delta NO_x/\Delta CO_2$  of 1.6×10<sup>-4</sup> being below the mean observed value of  $4.6\pm 2.3\times 10^{-4}$ , but consistent with the error bars of the individual samples as shown in 5 Figure 14. We could attempt to get a closer match by increasing  $\beta$ , but at the cost of increases 6 7 in the modeled  $O_3$ , PAN, and OA formation. The decay of  $C_2H_4$  is also better modeled than in 8 the Half Fragmentation case, suggesting the model OH is also improved. The modeled OH concentration for the best estimate dilution and photolysis case is now  $5.3 \times 10^6$  molecules cm<sup>-</sup> 9 <sup>3</sup>, matching the observed value of  $5.27 \pm 0.97 \times 10^6$  molecules cm<sup>-3</sup>. 10

# 11 6 Conclusions

We have used version 2.1 of the ASP model, which includes extensive updates to the gas-12 13 phase chemistry and SOA formation modules, to simulate the near-source chemistry within 14 the smoke plume from the Williams fire, as sampled by Akagi et al. (2012). We find that the assumptions made about the chemistry of the unidentified SVOCs emitted by the fire have a 15 16 large impact on the simulated secondary formation of  $O_3$ , PAN, and OA within the plume. We 17 showed that reasonable assumptions about the chemistry of the unidentified SVOCs can 18 successfully simulate the observations within the uncertainties of the measurements, the 19 estimated smoke ages of the samples, the plume dilution rate, and the vertical location of the samples in the plume. For the Williams fire, these assumptions were: (1) a reaction rate 20 constant with OH of  $\sim 10^{-11}$  cm<sup>3</sup>/ s; (2) a significant fraction (up to  $\sim 50\%$ ) of the RO<sub>2</sub> + NO 21 reaction resulted in fragmentation, rather than functionalization; (3)  $\sim 1.1$  molecules of O<sub>3</sub> 22 23 were formed for every molecule of SVOC that reacts; and (4) 60% of the OH that reacted 24 with the SVOC was regenerated as  $HO_2$  by the  $RO_2 + NO$  reaction, which implied (5) that 50% of the NO that reacted with the SVOC peroxy radicals was lost, likely due to organic 25 26 nitrate formation. However, this chemistry still overestimates PAN formation downwind by 27 about 50%, suggesting the need for further refinements to the chemistry and estimated 28 emission rates of PAN precursors like acetaldehyde. Furthermore, these specific, quantitative 29 results only apply to the Williams fire analyzed in this paper. Further analysis of other smoke 30 plume observations is needed to determine how these parameters vary between individual 31 smoke plumes.

The method used in this study can provide a way of classifying different smoke plume 1 2 observations in terms of the average chemistry of their unidentified SVOCs. Similar studies of 3 other young biomass burning plumes would allow us to see how the chemistry of the 4 unidentified SVOCs varies with combustion efficiency, fuel type, and other combustion and 5 environmental parameters, providing an additional constraint on the reactivities of the unidentified SVOCs. These constraints could then provide more insight into the formation of 6 7 O<sub>3</sub>, PAN, and OA in young biomass burning smoke plumes and serve as the basis of 8 parameterizing this process for regional or global scale models. Future field experiments, 9 focused on quasi-Lagrangian sampling of biomass burning smoke plumes, should ideally also 10 provide data beyond that available for the Williams Fire that will increase our understanding 11 of the chemistry of these plumes. These field experiments should include (a) observations of 12 changes in particle size distribution to test model simulations of condensational growth, 13 coagulation, and new particle formation; (b) observations of a larger suite of NO<sub>v</sub> species, 14 such as HNO<sub>3(g)</sub>, peroxy nitrates, and alkyl nitrates, for use in studying and constraining the 15 transformations of reactive nitrogen; (c) direct measurements of photolysis rates within the 16 smoke plumes; (d) measurements of organic aerosol volatility, viscosity, and mixing state with black carbon and inorganic aerosols; and (e) more detailed measurements of the 17 18 currently unidentified organic species present in the smoke plumes, including acetaldehyde, 19 an important PAN precursor.

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- 13

Species	<i>C</i> * (µg m <sup>-3</sup> @ 300 K)	$\Delta H_{vap}$ (kJ mol <sup>-1</sup> )	MW (g mol <sup>-1</sup> )	POA volatility distributions <sup>a</sup>			
				Grieshop et al. (2009a)	This study <sup>b</sup>		
SVOC <sub>1</sub>	10 <sup>-2</sup>	77	524	0	0		
SVOC <sub>2</sub>	10 <sup>-1</sup>	73	479	0	0		
SVOC <sub>3</sub>	$10^{0}$	69	434	0.1	0.038		
SVOC <sub>4</sub>	$10^{1}$	65	389	0.14	0.0532		
SVOC <sub>5</sub>	10 <sup>2</sup>	61	344	0.33	0.1254		
SVOC <sub>6</sub>	10 <sup>3</sup>	57	299	0.33	0.1254		
SVOC <sub>7</sub>	$10^{4}$	54	254	0.1	0.038		
SVOC <sub>8</sub>	10 <sup>5</sup>	50	208	0	0.62		
SVOC <sub>9</sub>	$10^{6}$	46	163	0	0		

Table 1. Definition of SVOC species following Grieshop et al. (2009a). 

<sup>a</sup>Relative mass emissions in each volatility bin. <sup>b</sup>Where the relative amounts of  $SVOC_i$  (*i*=1 to 7) are kept as in Grieshop et al. (2009a), but additional organic mass is added to  $SVOC_8$  to account for the unidentified NMOC mass 

reported by Akagi et al. (2011).

- 1 Table 2. SVOC chemistry parameters in the mechanisms studied here. See Reactions R3 and
- 2 R4 for definitions of the parameters.

Mechanism	$k_{OH} \times 10^{11}$ (cm <sup>3</sup> molecule <sup>-1</sup> s <sup>-1</sup> )	μ	n	α	χ	β	δ
Grieshop et al. (2009)	2.0	1.4	2	0	0	0	0
Robinson et al. (2007)	4.0	1.075	1	0	0	0	0
Ahmadov et al. (2012)	1.0	1.075	1	0	0	0	0
Half Fragmentation	1.0	1.075	1	.5	0	0	0
Optimized SVOC Chemistry	1.0	1.075	1	.5	1	.5	.6

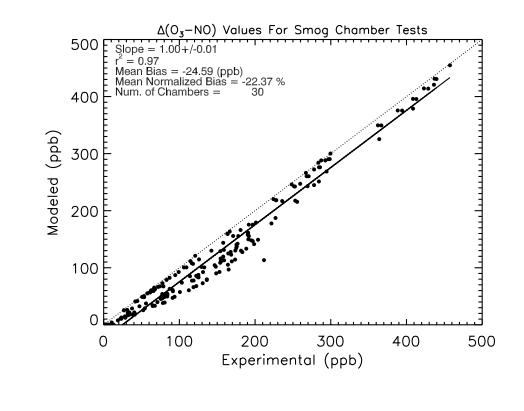


Figure 1. ASP calculated hourly values of  $\Delta([O_3] - [NO]) = ([O_3]_{final} - [NO]_{final}) - ([O_3]_{initial} - [NO]_{initial})$  versus the values measured in the EPA chamber of Carter et al. (2005) for 30 "full surrogate" experiments. Note that all time points for the 30 chamber experiments are plotted, not just the final values.

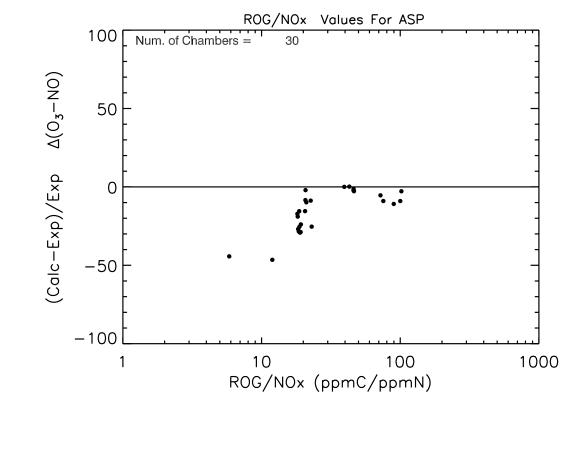


Figure 2. Percentage bias in final  $\Delta([O_3] - [NO])$  versus the initial ratio of reactive organic gases (ROG) to NO<sub>x</sub> (ppm C/ppm N) for the chamber experiments.

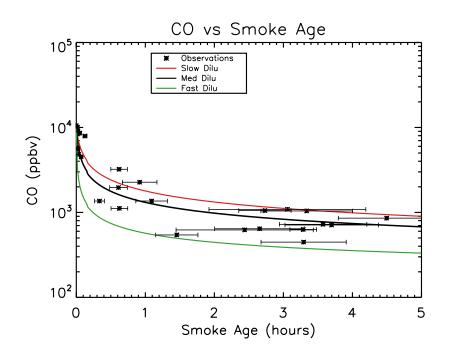


Figure 3. CO mixing ratio (ppbv) versus smoke age. Red, black, and green are for the slow,
best-fit (medium), and fast plume dilution rates. Asterisks are the measured mixing ratios,
with the horizontal error bars showing the uncertainty in the estimated age, which is much
larger than the uncertainty in the CO mixing ratio.

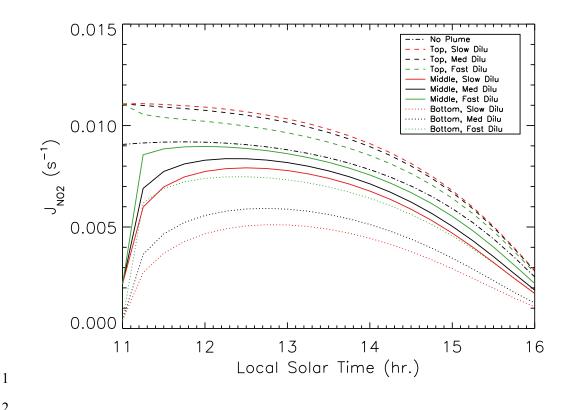


Figure 4. NO<sub>2</sub> photolysis rates (s<sup>-1</sup>) versus local time. Red, black, and green are for the slow, best-fit (medium), and fast plume dilution rates. Dashed lines are for photolysis rates above the plume, solid lines are for the middle of the plume, and dotted lines are for the bottom of the plume, as described in the text. The black dot-dashed line is the clear-sky (no plume) photolysis rate.

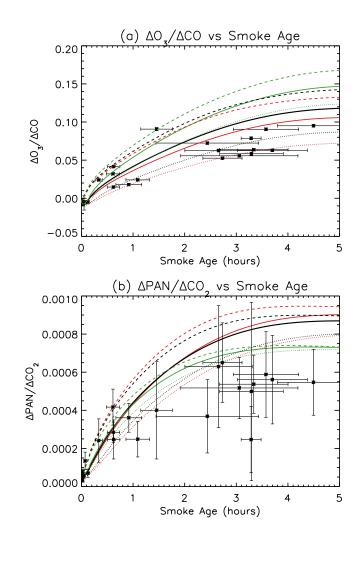
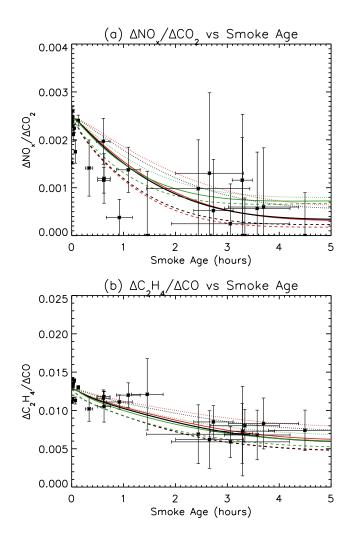
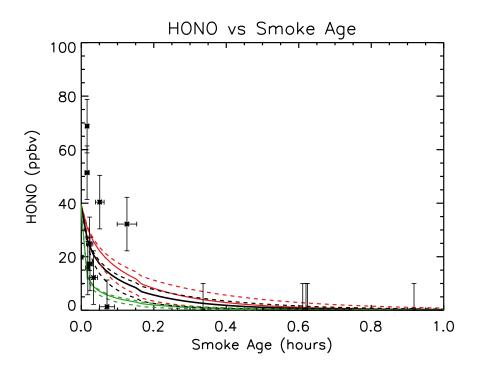




Figure 5. Enhancement ratios (mol/mol) of (a) O<sub>3</sub> to CO and (b) PAN to CO<sub>2</sub> versus estimated 3 4 smoke age when the chemistry of the unidentified SVOCs is not included in the model. 5 Asterisks are the measured mixing ratios, with the horizontal error bars showing the 6 uncertainty in the estimated age and the vertical error bars showing the uncertainty in the 7 measurement. Red, black, and green are ASP results for the slow, best-fit (medium), and fast 8 plume dilution rates. Dashed lines are for above-plume photolysis rates, while solid lines are 9 for the middle of the plume, and dotted lines are for the bottom of the plume (see the legend 10 in Figure 4).



3 Figure 6. (a) NOx enhancement ratio (EnR, mol/mol) to CO2 versus estimated smoke age 4 when the chemistry of the unidentified SVOCs is not included in the model. (b) EnR of C<sub>2</sub>H<sub>4</sub> 5 to CO versus estimated smoke age. Asterisks are the measured mixing ratios, with the 6 horizontal error bars showing the uncertainty in the estimated age and the vertical error bars 7 showing the uncertainty in the measurement. Red, black, and green are ASP results for the 8 slow, best-fit (medium), and fast plume dilution rates. Dashed lines are for above-plume 9 photolysis rates, while solid lines are for the middle of the plume, and dotted lines are for the 10 bottom of the plume (see the legend in Figure 4).



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3 Figure 7. HONO mixing ratio (ppbv) versus estimated smoke age for the first hour after 4 emission (note difference in x-axis scale from Figures 4-6) when the chemistry of the 5 unidentified SVOCs and a downwind HONO source is not included in the model. Asterisks 6 are the measured mixing ratios, with the horizontal error bars showing the uncertainty in the 7 estimated age and the vertical error bars showing the uncertainty in the measurement. Red, 8 black, and green are ASP results for the slow, best-fit (medium), and fast plume dilution rates. 9 Dashed lines are for above-plume photolysis rates, while solid lines are for the middle of the 10 plume, and dotted lines are for the bottom of the plume (see the legend in Figure 4).

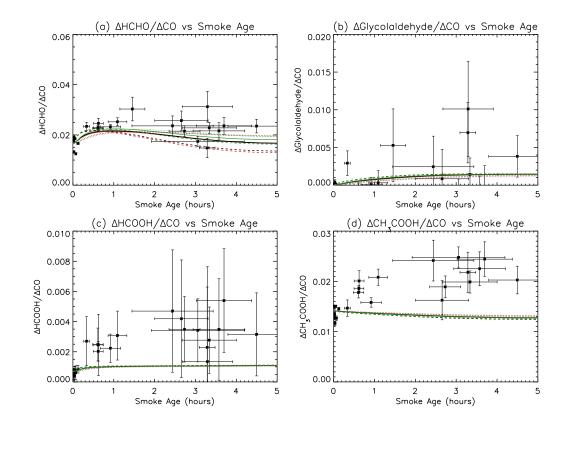


Figure 8. Enhancement ratio (EnR, mol/mol) of (a) HCHO, (b) glycoaldehyde (HCOCH<sub>2</sub>OH), (c) formic acid (HCOOH), and (d) acetic acid (CH<sub>3</sub>COOH) to CO versus estimated smoke age when the chemistry of the unidentified SVOCs is not included in the model. Asterisks are the measured mixing ratios, with the horizontal error bars showing the uncertainty in the estimated age and the vertical error bars showing the uncertainty in the measurement. Red, black, and green are ASP results for the slow, best-fit (medium), and fast plume dilution rates. Dashed lines are for above-plume photolysis rates, while solid lines are for the in-plume rates, as described in the text.

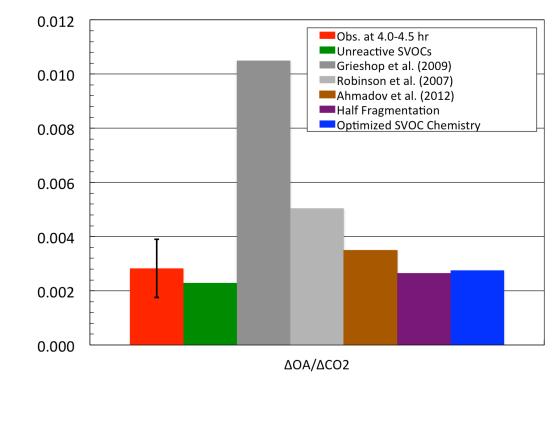


Figure 9. Enhancement ratio (EnR, g/g) of organic aerosol (OA) to CO<sub>2</sub> after 4 to 4.5 hr of smoke aging. The error bars on the observed values are based on the 36% uncertainty in the AMS observations of OA. All model results assume the best-estimate dilution rate and the photolysis rates corresponding to the middle of the plume (solid black line in Figure 4).

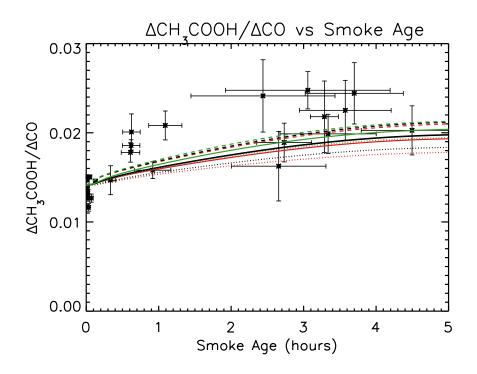
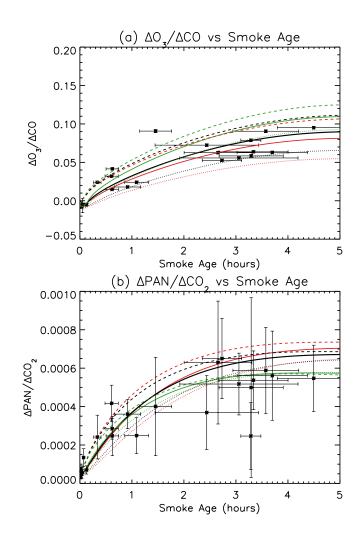


Figure 10. As in Figure 8d, but for the "Half Fragmentation" SVOC mechanism (see Table 1)
where the VOC fragment produced by fragmentation of the parent SVOC is assumed to
become acetic acid.



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3 Figure 11. As in Figure 5, but for the "Half Fragmentation" SVOC mechanism rather than no

4 fragmentation (see Table 2).

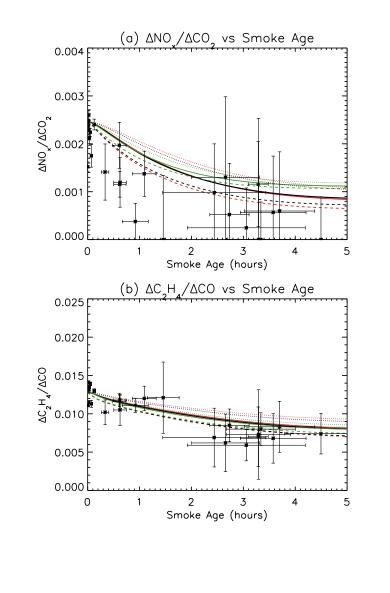
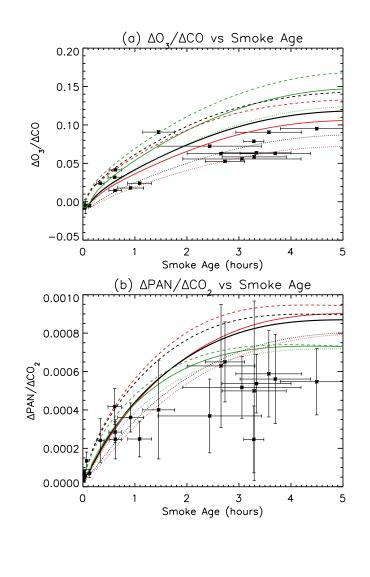
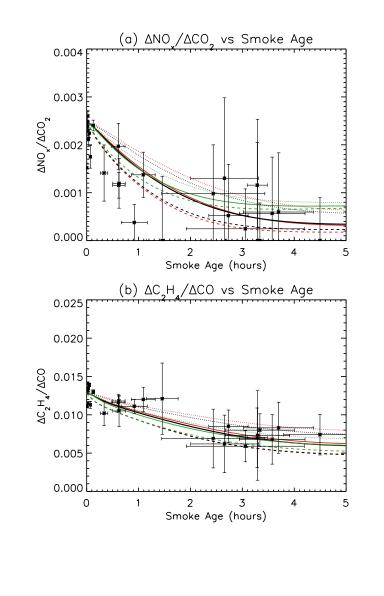


Figure 12. As in Figure 6, but for the "Half Fragmentation" SVOC mechanism (see Table 2).



3 Figure 13. As in Figure 5, but for the optimized SVOC chemistry (see Table 2).



3 Figure 14. As in Figure 6, but for the optimized SVOC Chemistry (see Table 2).