

# Simulation of the isotopic composition of stratospheric water vapour – Part 1: Description and evaluation of the EMAC model

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**Abstract.** This modelling study aims on an improved understanding of the processes, that determine the water vapour budget in the stratosphere by means of the investigation of water isotope ratios. At first, an additional (and separate from the actual) hydrological cycle has been introduced into the chemistry-climate model EMAC, including the water isotopologues HDO and H<sub>2</sub><sup>18</sup>O and their physical fractionation processes. Additionally an explicit computation of the contribution of methane oxidation to H<sub>2</sub>O and HDO has been incorporated. The model expansions allow detailed analyses of water vapour and its isotope ratio with respect to deuterium throughout the stratosphere and in the transition region to the troposphere. In order to assure the correct representation of the water isotopologues in the model's hydrological cycle, the expanded system has been evaluated in several steps. The physical fractionation effects have been evaluated by comparison of the simulated isotopic composition of precipitation with measurements from a ground-based network (GNIP) and with the results from the isotopologue-enabled general circulation model ECHAM5-wiso. The model's representation of the chemical HDO precursor CH<sub>3</sub>D in the stratosphere has been confirmed by a comparison with chemical transport models (1-D, CHEM2D) and measurements from radiosonde flights. Finally, the simulated stratospheric HDO and the isotopic composition of water vapour have been evaluated, with respect to retrievals from three different satellite instruments (MIPAS, ACE-FTS, SMR). Discrepancies in stratospheric water vapour isotope ratios between two of the three satellite retrievals can now partly be explained.

## 1 Introduction

30% of the temperature change since 1980 can be attributed to the radiative forcing of increased stratospheric water vapour (Solomon et al., 2010) and 10% of the global total ozone decline from 1960 to 1999 can be explained by the water vapour increase (Stenke and Grewe, 2005). However, the mechanisms driving long-term changes in stratospheric water vapour are not well understood (Füglister et al., 2009). Stratospheric water vapour is determined by in situ methane oxidation and the intrusion of water vapour through the tropical tropopause layer (TTL). The seasonal cycle of lower stratospheric water vapour in the tropics is characterised by the tape recorder (Mote et al., 1996), which exhibits a hydrated lower stratosphere in boreal summer and fall and a dry lower stratosphere in boreal winter and spring. Thus, most of the water vapour enters the stratosphere during boreal summer when the tropopause temperatures are higher and Monsoon systems (e.g. Gettelman and Kinnison, 2004) as well as enhanced deep convection over the tropics (e.g. Khaykin et al., 2009) cause augmented transport of water vapour into the TTL. The contribution of the individual mechanisms to the overall budget of stratospheric water vapour, however, is poorly quantified.

The application of water isotopologues in tracer studies has the potential to answer open questions concerning the stratospheric water vapour budget. The different vapour pressures and binding energies of the respective water isotopologues lead to fractionation effects during phase transitions and chemical reactions. The individual processes, which con-

60 trol the water vapour budget in the stratosphere, thus leave an isotopic fingerprint in the respective water vapour compound (Johnson et al., 2001). Direct injection of ice into the strato- 115 sphere through deep convection, dehydration of air during ascent by in situ formation of cirrus clouds and the in-mixing of older stratospheric air lead to an offset of the isotope ratios in the upper troposphere and lower stratosphere (UTLS) from 65 simple Rayleigh distillation (see Steinwagner et al., 2010). 120 The mechanisms, which are responsible for the intrusion of water vapour into the stratosphere, can hence be separated through the isotopic signature.

70 Recent satellite retrievals from three different instruments (SMR, MIPAS, ACE-FTS, see Urban et al., 2007; Steinwag- 125 ner et al., 2007; Nassar et al., 2007) allow to obtain a global picture of the water isotopologue HDO and the water vapour isotope ratio  $\delta D(H_2O)$  ( $\delta D(H_2O) = \left( \frac{[HDO]}{R_{VSMOW} [H_2O]} - 1 \right)$ . 75 1000;  $R_{VSMOW} = 155.76 \cdot 10^{-6}$ , Hagemann et al. (1970); 130 VSMOW: Vienna Standard Mean Ocean Water, IAEA (2009)) in the stratosphere and in the transition region to the troposphere. In some regions the different satellite retrievals, however, show considerable deviations when comparing pro- 80 files and annual cycles of HDO mixing ratios and isotope ratios, respectively (Lossow et al., 2011). Furthermore, Stein- wagner et al. (2010) found a tape recorder signal in  $\delta D(H_2O)$  in the tropical stratosphere in the MIPAS observations, com- 135 parable to the known tape recorder in water vapour mixing ratios. In contrast, Randel et al. (2012) did not find an analogue upward propagation of the seasonal cycle of wa- 85 ter vapour isotope ratios in the stratosphere when analysing ACE-FTS data. Speculations about this discrepancy mainly focus on the different sampling techniques and retrieval algo- 140 rithms. Consequently, in order to improve the understanding of the basic structure of  $\delta D(H_2O)$  in the stratosphere, comprehensive modelling of the physical and chemical isotope processes in a well resolved global chemistry climate model (CCM) with explicit stratospheric dynamics is necessary. 145

95 Hitherto, a GCM (General Circulation model) study by Schmidt et al. (2005) including HDO and focussing on the UTLS region only featured a relatively coarse vertical resolution in the stratosphere and only a prescribed ten- 150 dency for chemical influences on HDO. Here, the water isotopologue HDO is implemented into the EMAC (ECHAM MESSy Atmospheric Chemistry; MESSy: Modular Earth Submodel System) model (Jöckel et al., 2005, 2010). EMAC provides the opportunity to accurately analyse troposphere- 100 stratosphere exchange processes on a global scale. When used with 90 layers in the vertical, its base model (ECHAM5; 105 Roeckner et al., 2003) possesses a well resolved tropopause region (the vertical resolution near the tropopause is about 500 m; see Jöckel et al., 2006) and explicit stratospheric dynamics. Moreover, EMAC optionally represents chemical 160 processes, which are needed for the computation of methane oxidation, which is crucial for the representation of  $H_2O$  and HDO in the stratosphere.

Confirmation of the correct representation of the fractiona- tion processes during phase transitions in EMAC is achieved through the evaluation of the isotope ratios in precipitation. The consideration of the influence of methane oxidation on atmospheric HDO requires the computation of the methane isotopologue  $CH_3D$ . In a next step, this tracer is evaluated with respect to chemical transport models (1-D, CHEM2D, 115 Ridal et al., 2001; Ridal and Siskind, 2002) and measurements from stratospheric balloon flights (Röckmann et al., 2011). At last, the HDO mixing ratios and  $\delta D(H_2O)$  in the stratosphere are compared with the observations from satel- 120 lites. This approach yields a more complete picture of the isotopic composition of stratospheric water vapour and provides insights into the discrepancies between the respective satellite retrievals. Comprehensive analyses of stratospheric  $\delta D(H_2O)$  in EMAC will be presented in the companion part 2 of this article (Eichinger et al., 2014). These model results will be investigated with respect to the sensitivity and the ori- 125 gin of the  $\delta D(H_2O)$  tape recorder and the role of convective ice lofting on the pattern is analysed.

## 2 Model description and simulation setup

The MESSy submodel H2OISO ( $H_2O$  ISOtopologues) has been incorporated into the EMAC model system. This sub- model comprises tracers (see Jöckel et al., 2008) for the three stable water isotopologues  $H_2^{16}O$  ('normal' water, here- 135 after denoted as  $H_2O$ ),  $H_2^{18}O$  and HDO for all three phases (vapour, liquid and ice), respectively. Moreover, it contains an additional hydrological cycle, identical to the model's ac- tual hydrological cycle, which includes all processes that modify the tracers and the corresponding fractionation ef- 140 fects for the isotopologues during phase transitions. The modular approach of MESSy allows the optional usage of the H2OISO submodel for all users in future EMAC versions. Besides this structural difference, the implementation of the water isotopologues follows previous studies with ECHAM3 (Hoffmann et al., 1998), ECHAM4 (Werner et al., 2001) and 145 ECHAM5 (Werner et al., 2011). Supplementary to these previous studies, the chemical fractionation effects during the formation of water vapour through methane oxidation are considered. Since observations of water isotopologues in the stratosphere are mostly available for  $H_2O$  and HDO, the more elaborate accounting for the chemical fractionation of  $H_2^{18}O$  was not conducted in this study. Hence, here  $H_2^{18}O$  ba- 150 sically only serves the evaluation of the model in the tropo- sphere. The physical isotope effects of HDO and  $H_2^{18}O$  only differ by the corresponding fractionation factors.

### 2.1 Water isotopologues in the hydrological cycle of EMAC

The hydrological cycle in the H2OISO submodel reproduces the actual hydrological cycle of the EMAC model. The trac-

ers for  $\text{H}_2\text{O}$  in the gaseous ( $\text{H}_2\text{O}_{vap}$ ), the liquid ( $\text{H}_2\text{O}_{liq}$ ) and the ice ( $\text{H}_2\text{O}_{ice}$ ) phase were tested to be equal to the standard ECHAM variables for the water vapour (Q), the liquid water (XL) and the ice water (XI) content, respectively. Small numerical errors (the maxima are in the order of  $10^{-20}$  kg/kg/s for water vapour and  $10^{-22}$  kg/kg/s for liquid water and ice), which arise due to the coding design, are corrected after each time step, in order to prevent the two hydrological cycles from diverging. Moreover, we assured that these numerical errors are small enough to not deteriorate the calculation of the water isotopologues.

For the water isotopologues, fractionation effects occur during phase changes. Equilibrium and kinetic fractionation during the evaporation of water from oceans is described by the bulk formula of Hoffmann et al. (1998). The surface flux for the water isotopologues depends on the isotope ratio of the ocean, the wind speed above the surface, the sea surface temperature, the specific humidity and the isotopic composition of the vapour above the surface (Hoffmann et al., 1998). As in the study by Werner et al. (2011), the isotope ratio of the ocean is prescribed with a global gridded data set based on the  $^{18}\text{O}$  isotopic composition in sea water by LeGrande and Schmidt (2006). Since there is no equivalent data set for the deuterium isotopic composition, the HDO content is prescribed as eight times the  $\text{H}_2^{18}\text{O}$  isotope ratios. This is in accordance with global observations (Craig and Gordon, 1965). Due to the limitations of the applied land surface scheme in the EMAC model, we neglect any isotope fractionation from land surfaces. The land surface scheme in EMAC is identical to the scheme in the ECHAM5 model. A more detailed description of this simplification is given in Werner et al. (2011). As for water vapour, liquid water and ice in the actual hydrological cycle, the advection of the new water isotopologue tracers follows the flux form semi-Lagrangian (FFSL) scheme by Lin and Rood (1996).

The cloud and convection parameterisations (CLOUD and CONVECT) in EMAC include a number of phase transitions and therefore several different fractionation effects. Again, the implementation follows the study of Werner et al. (2011). During the formation of clouds, condensation of water vapour to liquid water and deposition of vapour to cloud ice take place. For condensation within clouds, a closed system is assumed. The condensate is taken on to be in contact, and hence in isotopic equilibrium, with the surrounding vapour during the entire process. This also applies for the evaporation of cloud water, where, in contrast to evaporation from the ocean, a closed system is assumed. An open system is used for the deposition of water vapour to ice. Due to the low diffusivities of the isotopologues in the ice phase, no exchange happens between ice and vapour. The effective fractionation factor  $\alpha_{eff}$  is used here, including a function for the supersaturation  $S$ . Werner et al. (2011) adjusted this function to  $S = 1.01 - 0.0045 \cdot T_{cond}$  (with  $T_{cond}$  as the condensation temperature during ice crystal formation, given in  $^{\circ}\text{C}$ ), in order to attain realistic isotope ratios

in Antarctic snow. Since the focus of the present study lies on the tropopause region, where also low temperatures have a major effect on kinetic fractionation through deposition, as a first estimate, the values from Werner et al. (2011) have been taken. During the melting of ice and the freezing of water, no fractionation is assumed. Other in-cloud processes like sedimentation of ice, autoconversion, accretion and aggregation include no fractionation effects either. The isotopic (non-)equilibrium factor  $\alpha_{eff}$ , which describes a fractionation process between a falling raindrop and the surrounding water vapour, may vary with time. Its value depends on the humidity, the temperature, the diffusivity of the water molecules and the droplet size. Since these processes are not resolved in GCMs, the fractionation during reevaporation of raindrops falling through undersaturated air can only be approximated. Following Hoffmann et al. (1998), an isotopical equilibration of 45% is assumed for large drops from convective rain and 95% for small drops falling from stratiform clouds. Due to their low exchange rates, snow and ice do not reequilibrate at all, which leads to more depleted isotope ratios in solid precipitation.

## 2.2 Stratospheric isotope chemistry for water and methane

The EMAC model contains the submodel CH4, which provides a simplified methane oxidation mechanism to take into account the chemical production of water vapour. It includes a tracer for methane (the  $\text{CH}_4$  tracer), which experiences a source from the surface (here as lower boundary conditions from the submodel TNUDGE, see next section; alternatively as methane fluxes, provided by the submodel OFFEMIS (see Jöckel et al., 2010)) and a sink in form of methane oxidation. Solutions are calculated for the four oxidation reactions, which are determined by the mixing ratios of the three oxidation partners (Cl, OH,  $\text{O}(^1\text{D})$ ) and the photolysis rate. The photolysis rate  $j_{\text{CH}_4}$  ( $= r_{hv}$ ) is here calculated in the MESSy submodel JVAL (for details, see Landgraf and Crutzen, 1998; Sander et al., 2014) and passed on to the submodel CH4 (alternatively it can be prescribed). The rates for the oxidation of methane with the reaction partners Cl, OH and  $\text{O}(^1\text{D})$  are calculated within the submodel CH4. First, the first order reaction coefficients  $k_{\text{OH}}$  for OH,  $k_{\text{Cl}}$  for Cl and  $k_{\text{O}1\text{D}}$  for  $\text{O}(^1\text{D})$  are determined. While  $k_{\text{O}1\text{D}} = 1.75 \cdot 10^{-10} \text{ cm}^3/\text{s}$  is constant,  $k_{\text{OH}}$  and  $k_{\text{Cl}}$  are temperature ( $T$  in K) dependent and are computed by

$$k_{\text{OH}} = 1.85 \cdot 10^{-20} \cdot \exp\left(2.82 \cdot \log(T) - \frac{987}{T}\right) \frac{\text{cm}^3}{\text{s}} \quad (1)$$

and

$$k_{\text{Cl}} = 6.6 \cdot 10^{-12} \cdot \exp\left(-\frac{1240}{T}\right) \frac{\text{cm}^3}{\text{s}}. \quad (2)$$

Subsequently the pseudo first order rate coefficients for the reactions with methane are

$$r_{O1D} = k_{O1D} \cdot c_{air} \cdot O(^1D) \quad (3)$$

$$r_{Cl} = k_{Cl} \cdot c_{air} \cdot Cl \quad (4)$$

$$r_{OH} = k_{OH} \cdot c_{air} \cdot OH, \quad (5)$$

with  $O(^1D)$ ,  $Cl$  and  $OH$  representing the prescribed mixing ratios (in  $mol/mol$ ) of the respective species and  $c_{air}$  the concentration of dry air molecules (in  $cm^{-3}$ ), which is calculated by

$$c_{air} = \frac{N_A \cdot 10^{-6} \cdot p}{R_{gas} \cdot T \cdot \left[ 1 + \left( \frac{M_{air}}{M_{H_2O}} - 1 \right) \cdot Q \right]}. \quad (6)$$

Here  $N_A$  denotes the Avogadro Constant ( $6.022045 \cdot 10^{23} \text{ mol}^{-1}$ ),  $p$  the pressure (in Pa),  $R_{gas}$  the universal gas constant ( $8.314409 \text{ J/K/mol}$ ),  $T$  the temperature (in K),  $M_{air}$  the molar mass of dry air ( $28.97 \text{ g/mol}$ ),  $M_{H_2O}$  the molar mass of water ( $18.02 \text{ g/mol}$ ) and  $Q$  the specific humidity (in  $kg/kg$ ).

The tendency for the methane tracer (in  $mol/mol/s$ ) is then given by

$$\frac{\partial(CH_4)}{\partial t} = -1 \cdot CH_4 \cdot (r_{O1D} + r_{Cl} + r_{OH} + r_{h\nu}), \quad (7)$$

where  $CH_4$  is the methane mixing ratio (in  $mol/mol$ ) of the previous time step and the  $-1$  accounts for the fact that this is a pure sink reaction for the methane tracer. To calculate the tendency for the specific humidity due to methane oxidation,

$$\left. \frac{\partial Q}{\partial t} \right|_C = \frac{-2 \cdot \frac{\partial(CH_4)}{\partial t}}{\frac{M_{air}}{M_{H_2O}} \left( \frac{1}{1-Q} \right)^2} \quad (8)$$

is applied. The subscript  $C$  denotes, that this is the chemical tendency of  $Q$ . This division is to convert the tendency from  $mol/mol/s$  to  $kg/kg/s$ . The negative sign here accounts for the fact that methane oxidation is a source for water vapour and the factor 2 for the reaction of the four hydrogen atoms of one methane molecule into two water molecules.

In order to take into account the chemical production of HDO, analogously a parameterisation for the oxidation of  $CH_3D$  has been devised. A tracer for  $CH_3D$ , the most abundant deuterium isotopologue of methane, was included for this purpose. The coefficients for the mass-dependent kinetic isotope effects (KIE) for the reactions of  $CH_3D$  with  $OH$ ,  $O(^1D)$  and  $Cl$  have been determined in laboratory measurements by Saueressig et al. (1996, 2001). They are partly temperature dependent and can be described with the function  $KIE(T) = A \cdot \exp(B/T)$ . The values for  $A$  and  $B$  and their temperature ranges are given in Tab. 1 (see also Röckmann

**Table 1.** Temperature dependent kinetic isotope fractionation coefficients for the reaction with  $CH_3D$ . The kinetic isotope effect is determined by  $KIE(T) = A \cdot \exp(B/T)$  for the given temperature range (see Röckmann et al., 2011).

Reactant	$T$ range	$A$	$B$
OH		1.097	$49 \pm 22$
$O(^1D)$	224-295	1.066	0
Cl	223-295	1.278	$51.31 \pm 19.1$

et al., 2011). The absorption cross-section of  $CH_3D$  is shifted 0.9 nm blueward relative to  $CH_4$  (Nair et al., 2005). For the photodissociation of  $CH_3D$ , this results in the fractionation factor  $KIE_{h\nu} = 0.995$  in the atmosphere of Mars (see also Nixon et al., 2012). This approach is here applied onto the Earth's atmosphere, since the photodissociation characteristics of methane do not differ from one planet of the solar system to another.

As for physical kinetic fraction processes, the Rayleigh equation

$$R = R_0 \left( \frac{N}{N_0} \right)^{KIE^{-1} - 1} \quad (9)$$

is applied for the fractionation processes in the chemical reactions. Inserting the isotope ratios ( $R_0$  and  $R$ ) and the total mixing ratios ( $N_0$  and  $N$ ) before and after the reaction, leads to

$$\frac{CH_3D - \frac{\partial(CH_3D)}{\partial t}}{CH_4 - \frac{\partial(CH_4)}{\partial t}} = \frac{CH_3D}{CH_4} \left( \frac{CH_4 - \frac{\partial(CH_4)}{\partial t}}{CH_4} \right)^{KIE^{-1} - 1} \quad (10)$$

Using Eq. 7 for  $\partial(CH_4)/\partial t$  and considering that the KIE is different for each of the reactions, the tendency of the  $CH_3D$  tracer is given by

$$\frac{\partial(CH_3D)}{\partial t} = CH_3D \cdot \left[ \left( 1 - (1 + r_{OH})^{KIE_{OH}^{-1}} \right) + \left( 1 - (1 + r_{Cl})^{KIE_{Cl}^{-1}} \right) + \left( 1 - (1 + r_{O1D})^{KIE_{O1D}^{-1}} \right) + \left( 1 - (1 + r_{h\nu})^{KIE_{h\nu}^{-1}} \right) \right]. \quad (11)$$

In order to calculate the tendency of the HDO tracer from the tendency of the  $CH_3D$  tracer (i.e. the chemical tendency of HDO), analogously to Eq. 8,

$$\left. \frac{\partial(HDO)}{\partial t} \right|_C = \frac{-1 \cdot \frac{\partial(CH_3D)}{\partial t}}{\frac{M_{air}}{M_{HDO}} \left( \frac{1}{1-HDO} \right)^2}, \quad (12)$$

is applied. Here, we use the specific humidity and the molar mass of HDO ( $M_{HDO} = 19.02 \text{ g/mol}$ ) instead of  $Q$  and  $M_{H_2O}$  and the factor 1 instead of 2, because the oxidation of one  $\text{CH}_3\text{D}$  molecule can only produce one HDO molecule. This approach, however, neglects the intermediate reactions between  $\text{CH}_3\text{D}$  and HDO with molecular hydrogen and its isotopologue HD. While  $\text{H}_2$  is fairly constant throughout the stratosphere and therefore can be neglected for the  $\text{CH}_4$  conversion to  $\text{H}_2\text{O}$ ,  $\delta\text{D}(\text{H}_2)$  increases with altitude due to chemical isotope fractionation effects (see Röckmann et al., 2003; Rhee et al., 2006). Hence, some deuterium is stored in HD during this process and therefore one  $\text{CH}_3\text{D}$  molecule translates into a little less than exactly one HDO molecule (see also, Rahn et al., 2003). In order to account for this, we apply the stratospheric relation

$$HD = -6.32 \cdot 10^{-5} \cdot CH_4 + 0.297 \quad (13)$$

derived from measurements by McCarthy et al. (2004). The time derivative of HD in dependence of the time derivative of  $\text{CH}_4$  can now be used to account for the stored deuterium in HD and thus for the reduction of chemically formed HDO. Therefore, Eq. 12 was extended to

$$\left. \frac{\partial(HDO)}{\partial t} \right|_C = \frac{-\frac{\partial(CH_3D)}{\partial t} + 6.32 \cdot 10^{-5} \cdot \frac{\partial(CH_4)}{\partial t}}{\frac{M_{air}}{M_{HDO}} \left( \frac{1}{1-HDO} \right)^2} \quad (14)$$

This approach still neglects a number of effects that may be important for the chemical production of HDO: Firstly, the other, rather rare methane isotopologues  $\text{CH}_2\text{D}_2$ ,  $\text{CHD}_3$  and  $\text{CD}_4$  as well as the reaction partner OD (an isotopologue of the hydroxyl radical) are not considered. Secondly, the fractionation effects during the formation of HDO by HD are not taken into consideration. Most of these, however, are poorly quantified (Zahn et al., 2006) and therefore neglected for this initial study. These simplifications have to be kept in mind when evaluating the model results in the stratosphere.

Emissions of  $\text{CH}_3\text{D}$  could be defined with the aid of an extension (for deuterium) of the MESSy isotope scheme of Gromov et al. (2010). However, this task requires the integration of the full chemistry scheme and therefore goes beyond the scope of our study, which mainly focuses on stratospheric and upper tropospheric processes. This leads to another simplification: In accordance with Rhee et al. (2006) the isotope ratio of methane was fixed to  $-86\text{‰}$  in the troposphere (i.e. here, below 500 hPa). This is in between the stratospheric entry values of  $\delta\text{D}(\text{CH}_4)$  ( $\delta\text{D}(\text{CH}_4)$  is also based on VS-MOW) found by Rahn et al. (2003) ( $-90\text{‰}$ ), Röckmann et al. (2003) ( $-80\text{‰}$ ) and Röckmann et al. (2011) ( $-81\text{‰}$ ) and also within the range of recent aircraft measurements presented by Umezawa et al. (2012).

### 2.3 Simulation setup

For this study, an EMAC (v2.42) model simulation, in T42L90MA resolution was performed. This corresponds to an approximate horizontal grid box size of  $2.8^\circ \times 2.8^\circ$ , 90 layers in the vertical and explicitly resolved stratospheric dynamics. The uppermost model layer is centered around 0.01 hPa. The MESSy submodels, which were applied in this simulation (in addition to the described H2OISO and CH4 submodels and the ECHAM5 base model from EMAC version 2.42) are listed and described in the appendix. The time step of the simulation was twelve minutes and the output was set to produce instantaneous values with an interval of eleven hours. The EMAC model provides the possibility to use several different convection schemes. In all the simulations conducted for this study, the ‘‘Tiedtke-Nordeng’’ convection scheme (Tiedtke, 1989; Nordeng, 1994) was applied. Before starting the actual simulation, a 20 year free running simulation was carried out to obtain steady-state initial values for water, methane and their isotopologues. From these initial conditions a simulation with specified dynamics (i.e. in ‘‘nudged’’ mode) was started, which means a Newtonian relaxation of the divergence, the vorticity, the temperature and the logarithm of the surface pressure towards reference data. Here, the relaxation is performed up to 1 hPa towards the ERA-INTERIM reanalysis data (ECMWF; Dee et al., 2011). This guarantees, that not only the climatic state, but also the meteorological situation of the model simulation corresponds to the actual states and allows a direct comparison of the model results with satellite or in situ measurements. The simulation starts at the beginning of the year 1982 and terminates at the end of the year 2010. From the steady-state conditions, which are used for the initialisation, the model again needs several years to adjust to the conditions of the ‘‘nudged’’ mode. The first eight years are hence not considered for the analysis. Only the 21 years from 1990 to 2010 are evaluated. Transient greenhouse gas concentrations are prescribed throughout the atmosphere. Methane is prescribed at the lower boundary through the submodel TNUDGE, based on observations. The mixing ratios of OH, Cl,  $\text{O}(^1\text{D})$  are prescribed (monthly averages) from a previous ‘‘nudged’’, transient EMAC simulation with full chemistry. The same applies for ozone, which is needed to calculate the photolysis rate in the submodel JVAL. As mentioned above, the methane oxidation is calculated in the CH4 submodel.

### 3 Model evaluation

Before evaluating the model simulation with respect to the isotopic composition of water vapour in the stratosphere, we show that the basis for this is established in the EMAC model. This includes the correct representation of the physical water isotope effects in the troposphere and the chemical HDO precursor  $\text{CH}_3\text{D}$  in the stratosphere, as well as its sink

440 reactions. Subsequently stratospheric HDO mixing ratios and  
 445  $\delta D(H_2O)$  are compared to satellite observations.

### 3.1 Water isotope ratios in precipitation

The isotopic composition in precipitation allows to evaluate various isotope fractionation effects (Dansgaard, 1964). Hence, the representation of the isotope physics in the model can be evaluated by examining the isotope ratios in precipitation. This has been conducted in many studies with isotopologue-enabled GCMs (e.g. Risi et al., 2010; Werner et al., 2011).

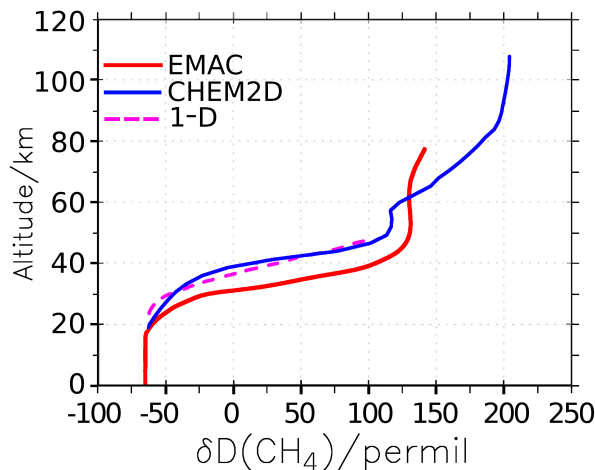
450 The GNIP (Global Network of Isotopes in Precipitation) measurement survey provides a comprehensive data base for this purpose. The International Atomic Energy Agency (IAEA) and the World Meteorological Organisation (WMO) have started GNIP in 1961. Since its start, more than 800 meteorological stations in 101 countries have collected samples of the water isotopologues HDO and  $H_2^{18}O$  in precipitation (IAEA, 2009). Some of these stations provided monthly precipitation measurements for several decades, most of them, however, operated for a much shorter period, only. This data set serves as a basis for the evaluation of the simulated isotope ratios in the troposphere. Good qualitative agreement was achieved when comparing the general global patterns in annual mean precipitation, as well as in the seasonal cycles between GNIP and EMAC. Details can be found in the supplement.

465 The isotopologue-enabled GCM ECHAM5-wiso (water isotopologues) has also been tested successfully against GNIP data, as well as against other water isotopologue measurements (see Werner et al., 2011). Several model resolutions have been applied for this study. The results of the lowest of these horizontal resolutions ( $T31 \sim 3.75^\circ \times 3.75^\circ$ ) was largely reproduced with the EMAC model (see supplement). Since the model physics and dynamics of the two models, including the hydrological cycle and the implementation of the water isotopologues (as described in Sect. 2), are almost identical, this is not surprising. However, it is a prerequisite for assuring the correct representation of the tropospheric isotope composition of water. A detailed comparison between EMAC and ECHAM5-wiso is given in the supplement. The good quantitative agreement of this comparison supports the conclusion, that the EMAC model with the H2OISO submodel represents the state of the art of GCMs with an explicit representation of the water isotopologues HDO and  $H_2^{18}O$  in the troposphere.

### 485 3.2 $CH_3D$ in the stratosphere

The simulated  $CH_3D$  is compared to results from the 1-D model by Ridal et al. (2001) and the CHEM2D model by Ridal and Siskind (2002). These models comprise an oxidation scheme, where  $CH_3D$  produces HDO through a number of chemical reactions. This oxidation scheme was developed for

the 1-D model. For the CHEM2D model, it was extended for higher altitudes and included into the Naval Research Laboratory two-dimensional chemical/dynamical model. In Ridal (2002) and Ridal and Siskind (2002) the two chemical transport models have shown good general agreement with measurements from the ATMOS (Atmospheric Trace MOlecule Spectroscopy) instrument (Irion et al., 1996). ATMOS provides global data for  $CH_3D$  and HDO, however, with large uncertainties. The equatorial values of  $\delta D(CH_4)$  of EMAC, CHEM2D and the 1-D model are presented in Fig. 1.



**Fig. 1.** Comparison of equatorial averages of  $\delta D(CH_4)$  with altitude between EMAC (red), CHEM2D (blue) by Ridal and Siskind (2002) and the 1-D model (dashed purple) by Ridal et al. (2001).

The tropospheric values of  $\delta D(CH_4)$  in the CHEM2D and the 1-D model are fixed to  $-68\text{‰}$  and  $-65\text{‰}$ , respectively. For a better comparison with these models, an additional EMAC simulation was carried out, with the only difference, that this value was set from  $-86\text{‰}$  to  $-68\text{‰}$  (see Sect. 2.2). Therefore, the model results do not differ below the tropopause. Moreover, the overall dependence of  $\delta D(CH_4)$  on altitude qualitatively agrees in all the three model simulations. Between 20 and 50 km altitude the methane isotope ratio increases from  $-68\text{‰}$  to around  $+120\text{‰}$  in CHEM2D and the 1-D model, and to around  $+130\text{‰}$  in EMAC. Especially the increase in the lower stratosphere is much stronger in EMAC, which leads to rather large discrepancies in the altitude range between 30 and 40 km. Between 50 and 60 km, both, CHEM2D and EMAC, show almost no change in  $\delta D(CH_4)$  with altitude, the 1-D model does not extend above 50 km. This is the transition region between the altitudes of the chemical and the photolytic methane oxidation. The photodissociation, which becomes important above 60 km and increases continuously above, is much stronger in CHEM2D. This is somewhat surprising, because there is no fractionation

included for the photolysis of  $\text{CH}_3\text{D}$  in the CHEM2D model. Even though the fractionation for photolysis in EMAC is very small, the photolysis of  $\text{CH}_3\text{D}$  is expected to be of similar strength as in CHEM2D. The discrepancy is hence likely caused by the differences in the calculation of the photolysis rates in EMAC and CHEM2D. Since the mid of the uppermost layer of the EMAC model in the applied resolution is at 80 km, a comparison further above is not possible.

Measurements of  $\text{CH}_3\text{D}$  in the stratosphere are sparse. Röckmann et al. (2011), however, collected 13 altitude profiles from stratospheric balloon borne measurements, which were provided by the Max-Planck Institute (MPI) for Solar System Research and by the Institute for Atmospheric and Environmental Sciences of the University of Frankfurt. The mixing ratios and the isotopic composition of  $\text{CH}_4$  were measured, using a high-precision continuous flow isotope ratio mass spectrometry system (Brass and Röckmann, 2010). Twelve of these balloon flights can be used for direct inter-comparison with the data from the EMAC simulation. One flight (Flight ID: HYD-87-03) was conducted in 1987. Since the model results are considered to be in steady-state only from 1990 on, this profile is not taken into account for the evaluation. In Fig. 2, twelve panels are presented showing the balloon borne data and the EMAC results of  $\delta\text{D}(\text{CH}_4)$  as function of altitude, between 5 and 35 km. The flight IDs, included in the panels provide information about the location, the month and the year of the balloon flights. To provide an estimate about the average and the annual variability of  $\delta\text{D}(\text{CH}_4)$  in the model simulation, additionally, the averages, maxima and minima of the 21 simulation years of the respective months at the location of the launch are included in the panels.

In general, a good qualitative agreement is achieved between the measured and the simulated data. Both simulation and measurements, show an increase of the methane isotope ratios from tropospheric values to values between 100‰ and 200‰ at 25 to 35 km in the Arctic region and to values between 0‰ and 100‰ in the mid-latitude and tropical regions. The balloon borne profiles are mostly lying within the extremes of the simulation and close to the simulated values from the same day and location. The measured  $\delta\text{D}(\text{CH}_4)$  values in the troposphere and lower stratosphere are in good agreement with the chosen value of  $-86‰$  from the study by Rhee et al. (2006). The steep increase of  $\delta\text{D}(\text{CH}_4)$  above 25 km in the simulation can only be seen in the measurements of the Arctic region (KIR). In the mid-latitude (GAP and ASA) profiles, the  $\delta\text{D}(\text{CH}_4)$  increase with height is smaller, which leads to partly large deviations at these altitudes. The differences in the KIR-00-01 and the KIR-03-03 profiles between simulation and balloon flights are exceptional. Apart from the two uppermost data points in the KIR-03-03 profile, the measured  $\delta\text{D}(\text{CH}_4)$  values are constantly higher than the simulated values. These two profiles are associated with special meteorological and thus chemical situations. The KIR-

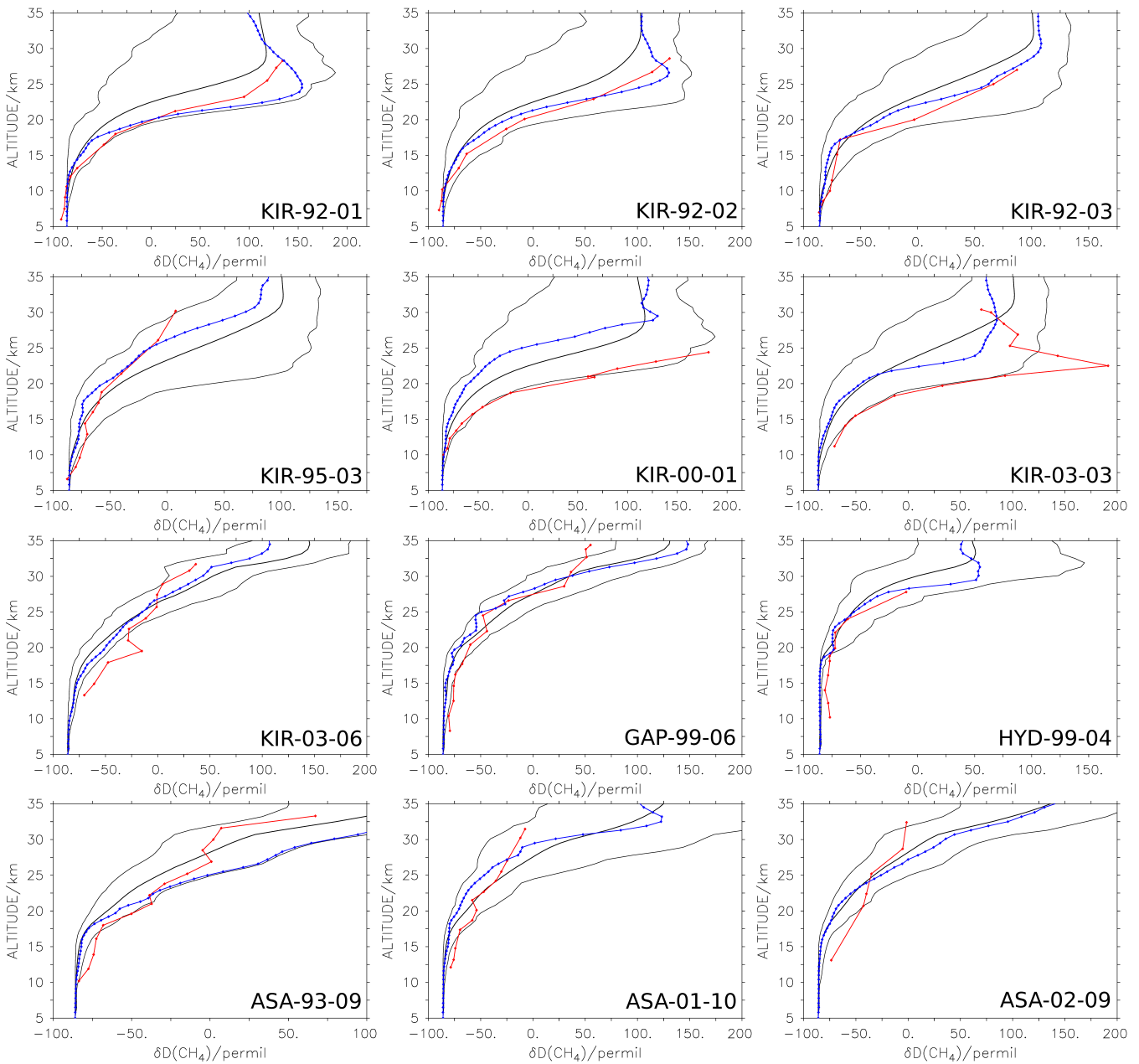
03-03 data comprises a mesospheric enclosure and during the sampling of the KIR-00-01 data, a strong Arctic vortex was present (Röckmann et al., 2011). These phenomena can also be observed in the simulation, due to the “nudging” these special meteorological situations and the associated chemical situations are broadly represented. However, the horizontal resolution of the model simulation is too coarse to resolve sharp horizontal gradients around the site of the balloon launch. Moreover, the balloon’s drift off the launching site while ascending can cause deviations of such magnitudes. Also, this may be related to omitting the chemical cycle of  $\text{H}_2$  and HD, since isotopically very heavy  $\text{H}_2$  might have been reconverted to  $\text{CH}_4$  in these profiles.

Another method for evaluating the methane isotope chemistry is assessing the relation of  $\delta\text{D}(\text{CH}_4)$  to the  $\text{CH}_4$  mixing ratio. The  $\delta\text{D}(\text{CH}_4)$  values of the same data as in Fig. 2, from 5 to 35 km altitude, are plotted versus the  $\text{CH}_4$  mixing ratios in Fig. 3. The figure is divided into the launches in the polar region (KIR) in the left panel and the launches in mid- (ASA, GAP) and tropical (HYD) latitudes in the right panel.

Again, overall consistency between the simulation and the measurements is visible. Apart from a single exception (the measurement from the HYD-99-04 profile, right panel) the simulated  $\delta\text{D}(\text{CH}_4)$  as function of the methane mixing ratios agree very well with the measurements. The slope of increasing isotope ratios with decreasing methane mixing ratios is almost identical. Since these compact tracer-tracer correlations are generally found for trace gases whose life times are longer than the transport times (Plumb and Ko, 2004), it implies that the chemical removal of the  $\text{CH}_3\text{D}$  tracer in relation to the removal of the  $\text{CH}_4$  tracer is well represented, despite the simplified chemistry parameterisation.

### 3.3 HDO in the stratosphere

During the first decade of the 21st century, three satellite missions collected data applicable for the retrieval of the water isotopologue HDO in the stratosphere. The MIPAS (Michelson Interferometer for Passive Atmospheric Sounding) instrument on Envisat (Environmental Satellite) allowed the retrieval of HDO by measuring the thermal emission in the mid-infrared. This high-resolution Fourier transform spectrometer measured at the atmospheric limb and provided data for HDO retrievals in full spectral resolution from July 2002 to March 2004, roughly in the altitude range between 10 and 50 km. It orbits the Earth sun-synchronously 14 times a day. The vertical resolution for the retrieval of HDO is around 5 km between the tropopause and 30 km altitude, above it degrades ( $\sim 8$  km at 40 km) (Steinwagner et al., 2007; Los-sow et al., 2011). The Odin satellite also orbits the Earth sun-synchronously and carries the SMR (Sub-Millimetre Radiometer) instrument, among other purposes to passively measure HDO on the global scale roughly on one day per week. It operates in the microwave range. Data has been retrieved from the start of the mission in 2001 until today,

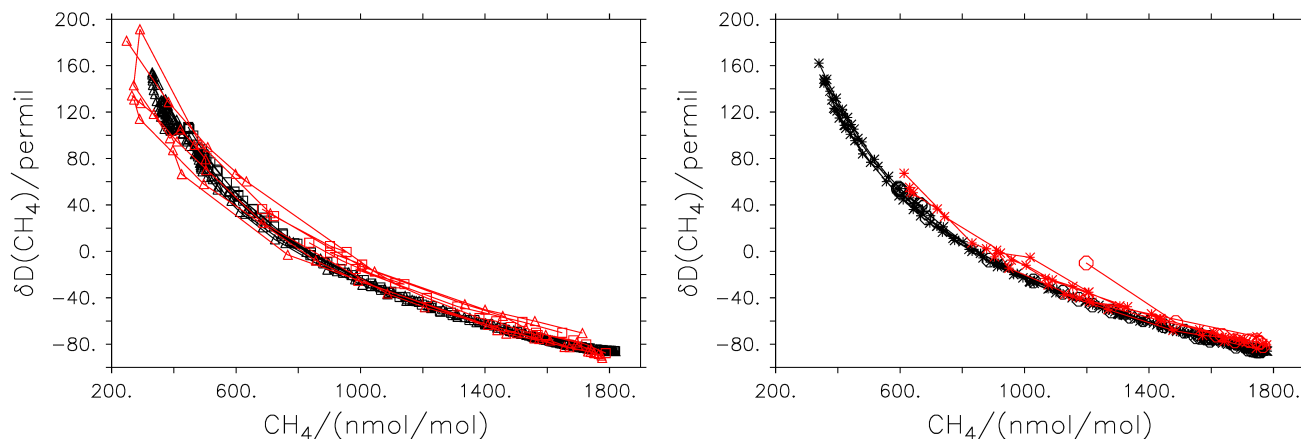


**Fig. 2.** Comparison of  $\delta D(\text{CH}_4)$  vertical profiles between EMAC and balloon borne data by Röckmann et al. (2011). The red lines show the observations and the blue lines show the EMAC data of the same day at the location of the balloon launch. The black lines represent the simulated averages, minima and maxima of the 21 monthly averages of the respective month at the location of the balloon launch. The flight IDs, included in the panels denote the location, the year and the month of the balloon flight. KIR: Kiruna, Sweden ( $67.9^\circ\text{N}$ ,  $21.10^\circ\text{E}$ ); GAP: Gap, France ( $44.44^\circ\text{N}$ ,  $6.14^\circ\text{E}$ ); HYD: Hyderabad, India ( $17.5^\circ\text{N}$ ,  $78.60^\circ\text{E}$ ); ASA: Aire sur l'Adour, France ( $43.70^\circ\text{N}$ ,  $0.30^\circ\text{W}$ ).

at altitudes between roughly 20 and 70 km with a vertical resolution of around 3 km (Urban et al., 2007). The ACE-FTS (Atmospheric Chemistry Experiment Fourier Transform Spectrometer) instrument circularly orbits the Earth on the SCISAT satellite and obtains Fourier transform absorption spectra from solar occultation measurements. It has a vertical

resolution between 2 and 6 km and a comparably limited spatial sampling. The ACE orbit is optimised for measurements in mid- and high latitudes, the tropics are only covered during the four months of February, April, August and October (Nassar et al., 2007; Randel et al., 2012).



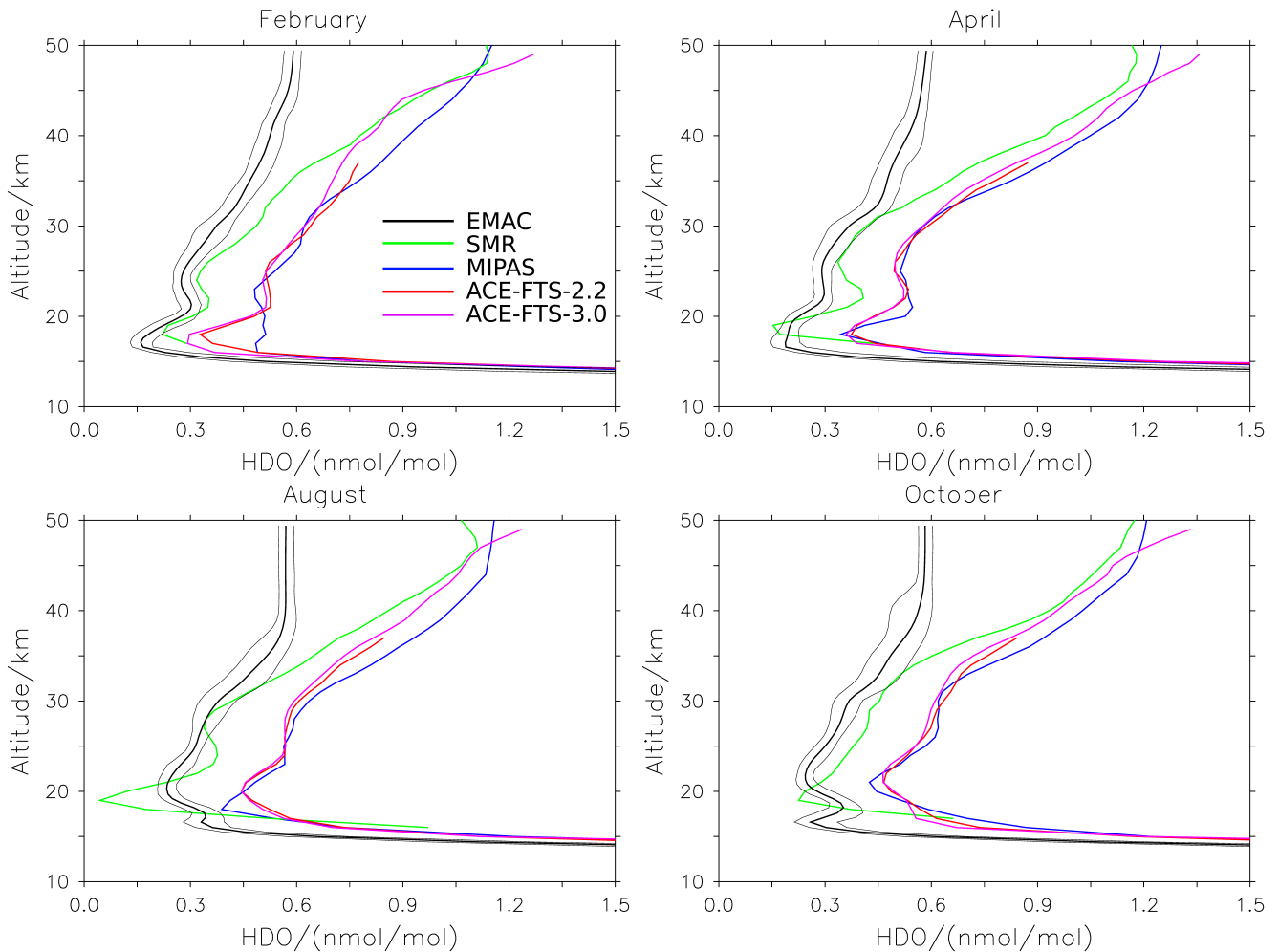


**Fig. 3.** Comparison of the relations of  $\delta D(\text{CH}_4)$  to  $\text{CH}_4$  between EMAC (black) and balloon borne data (red; data as in Fig. 2). The left panel shows the Arctic data (KIR,  $\triangle$ : vortex,  $\square$ : non-vortex) and the right panel shows the mid-latitude ( $*$ ) and tropical data ( $\circ$ ). The altitude range is 5 to 35 km.

Lossow et al. (2011) collected data of the three instruments <sup>675</sup> for intercomparison and concluded a good consistency between MIPAS and ACE-FTS at altitudes above 20 km. Below this altitude, issues like different cloud filtering and measurement techniques as well as different vertical resolutions cause large deviations. Generally, in the stratosphere the <sup>680</sup> MIPAS and the ACE-FTS data agree favourably, the SMR data shows considerably dryer conditions, especially below 30 km. This is mainly due to uncertainties in the different spectroscopies of the instruments (Lossow et al., 2011).

These data could now also be used to evaluate the HDO <sup>685</sup> simulated by EMAC. In Fig. 4, the tropical ( $15^\circ\text{S}$  to  $15^\circ\text{N}$ ) HDO mixing ratios of the three satellite instruments and the EMAC model are presented. Additionally to the data of MIPAS, SMR and version 2.2 of ACE-FTS shown by Lossow et al. (2011), here also version 3.0 of ACE-FTS is included, which reaches higher up in the stratosphere, compared to version 2.2. Since ACE-FTS only provides data for four months per year in this region, the panels show averages for February, April, August and October. Since the years of the ACE-FTS and the MIPAS satellite retrievals do not <sup>695</sup> overlap, a direct comparison is not possible. However, tests have shown, that the averages of the periods of the individual retrievals do not show substantial differences between each other and between the average of the entire simulation. In order to get a good estimate for the comparison of the model with all retrievals and for the annual variability, the averages, the minima and the maxima of the respective months have been taken from the entire 21 years of the EMAC simulation. Since, at first order, the HDO mixing ratios follow those of  $\text{H}_2\text{O}$ , the profiles of  $\text{H}_2\text{O}$  of EMAC and of the mentioned <sup>705</sup> satellite retrievals are shown in Fig. 5 in the same manner, in order to allow a more detailed discussion of the results.

The EMAC data is generally dryer in HDO and in  $\text{H}_2\text{O}$  compared to the MIPAS and the ACE-FTS profiles in each of the presented months at all altitudes. For water vapour in the UTLS, this offset has already been shown by Jöckel et al. (2006) and is consistent with the slightly too cold hygropause in the nudging data (see e.g. Liu et al., 2010). Only between 30 and 35 km the  $\text{H}_2\text{O}$  profiles of EMAC increase stronger than in the satellite data and reach the level of ACE-FTS  $\text{H}_2\text{O}$  mixing ratios. For HDO, this is not the case, its increase with height at these altitudes is rather similar between EMAC and the satellite retrievals and therefore the offset remains. In the altitude range between 16 and 30 km HDO in the EMAC simulation quantitatively corresponds well with the Odin retrieval. In this region also local maxima and minima, which can be seen in all four satellite profiles are reproduced qualitatively in EMAC. These reveal the seasonal cycle of HDO and  $\text{H}_2\text{O}$ . Due to the spectral band that is used to derive these parameters by SMR, in this special case,  $\text{H}_2\text{O}$  can not be retrieved below 20 km while the limit for HDO is between 17 and 18 km. Especially for April and August, the local HDO minimum between 25 and 30 km in the EMAC data is not as pronounced as in the satellite retrievals. Above the local minima the HDO mixing ratios of all satellite profiles increase strongly with altitude to values around 1.1 nmol/mol at 50 km, while the EMAC simulation shows HDO mixing ratios of only 0.6 nmol/mol at these altitudes. In  $\text{H}_2\text{O}$ , this behaviour is observable as well, however, only at altitudes above around 40 km and at a significantly smaller magnitude. Therefore, to some extent, this is most likely due to the assumptions made in the chemistry parameterisation for HDO, which does not include the influence of the fractionation effects during the reactions concerning molecular hydrogen. Also the simplified methane oxidation scheme itself, however, plays a role here.



**Fig. 4.** Comparison of HDO mixing ratio-altitude profiles between EMAC and various satellite observations. Black: Averages, minima and maxima of 21 monthly averages of the EMAC simulation for the respective months; green: Odin/SMR; blue: ENVISAT/MIPAS; red: SCISAT/ACE-FTS-2.2; purple: SCISAT/ACE-FTS-3.0.

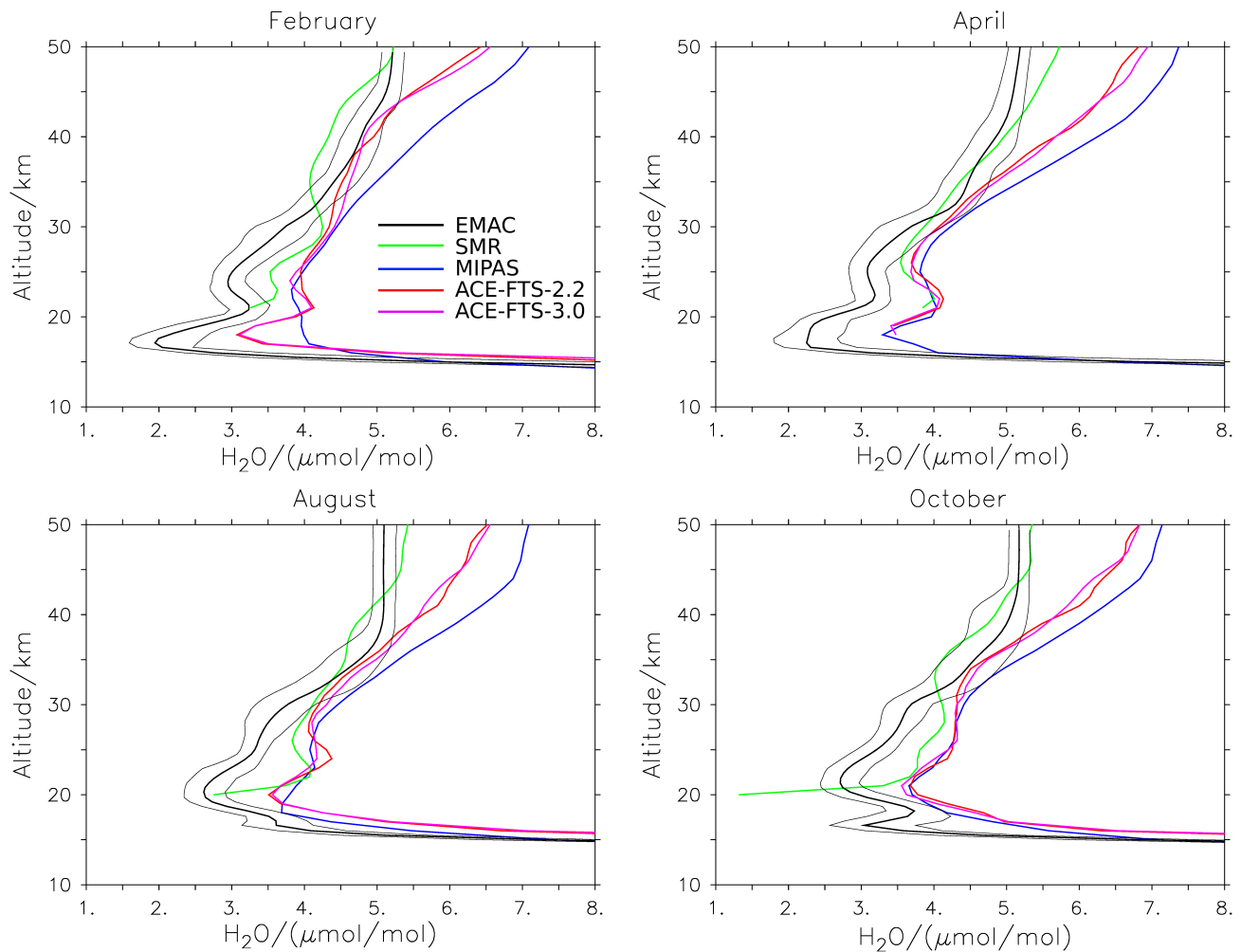
### 3.4 The stratospheric $\delta D$ tape recorder

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710 The tape recorder signal in  $H_2O$ , HDO and  $\delta D(H_2O)$  in the EMAC simulation is evaluated with respect to the MIPAS data. The satellite and the model data are compared in Fig. 6. The left panels show the EMAC results and the right panels 730 show the MIPAS retrieval.

715 Overall, there is only a rather weak agreement between EMAC and MIPAS in all three quantities. As already shown in the previous section, a persistent (all altitudes, all seasons) dry bias in  $H_2O$  and HDO is visible in the EMAC simulation. 735 Also  $\delta D(H_2O)$  is generally too low throughout the stratosphere. In the lower stratosphere, some of these differences can be explained with the coarse vertical resolution of the MIPAS retrieval, which smoothes the hygropause (see Steinwagner et al., 2007). In the upper stratosphere, the underes- 740 timation by the model can be associated to the assumptions

in the chemistry parameterisation. In the lower stratosphere, the strongest deviations in  $\delta D(H_2O)$  can be observed during NH summer when MIPAS observations show  $\delta D(H_2O)$  values around  $-500\text{‰}$  and the  $\delta D(H_2O)$  values in EMAC do not exceed  $-600\text{‰}$ . A tape recorder signal can be seen in all three quantities for both, model and observations, although with different amplitudes and a phase shift of two to three months. While the maxima of the tape recorder in the lower stratosphere in EMAC are found during summer, the satellite data show them at the beginning of autumn. This can partly (at maximum one month) be an artefact of the MIPAS retrieval and its coarse vertical sampling, but has to be kept in mind and compared with other data sets. The tape recorder signals in HDO and  $H_2O$  fade out at around 30 km in both, model and observations. The  $\delta D(H_2O)$  tape recorder signal in MIPAS reaches these altitudes as well, the



**Fig. 5.** Comparison of  $\text{H}_2\text{O}$  mixing ratio-altitude profiles between EMAC and various satellite observations. Black: Averages, minima and maxima of 21 monthly averages of the EMAC simulation for the respective months; green: Odin/SMR; blue: ENVISAT/MIPAS; red: SCISAT/ACE-FTS-2.2; purple: SCISAT/ACE-FTS-3.0.

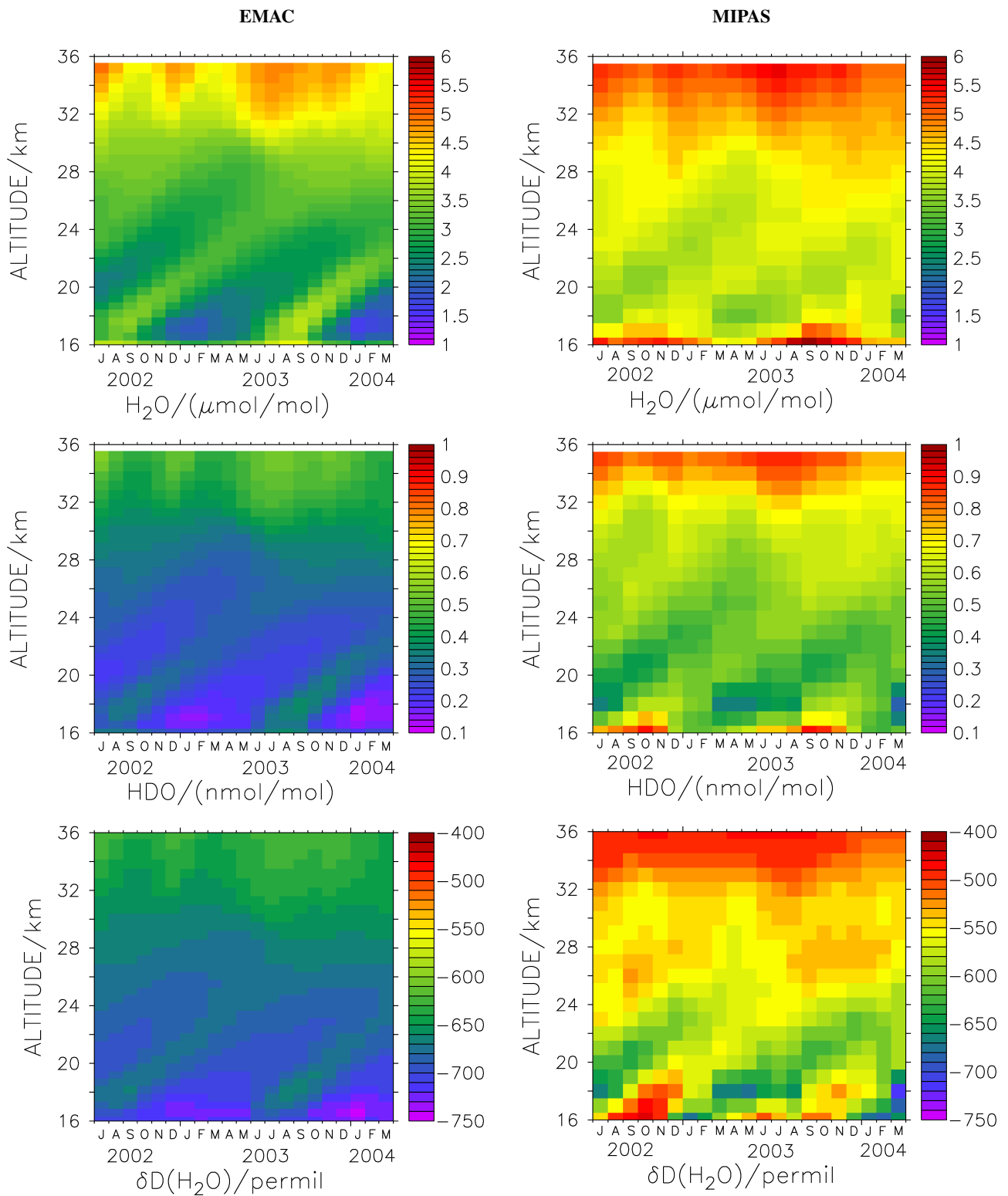
EMAC  $\delta\text{D}(\text{H}_2\text{O})$  tape recorder, in contrast, fades out somewhat lower already.

The amplitude of the tape recorder in EMAC is larger for  $\text{H}_2\text{O}$  and smaller for HDO and  $\delta\text{D}(\text{H}_2\text{O})$ , respectively, compared to the MIPAS data. Above 28 km, the  $\delta\text{D}(\text{H}_2\text{O})$  tape recorder becomes overshadowed by chemically produced high  $\delta\text{D}(\text{H}_2\text{O})$  values. This effect is further investigated in part 2 of the article (Eichinger et al., 2014). The MIPAS  $\delta\text{D}(\text{H}_2\text{O})$  tape recorder might be amplified artificially by the offsets in vertical resolution between  $\text{H}_2\text{O}$  and HDO (Orbe et al., 2013). A correction of this error is likely to reveal a  $\delta\text{D}(\text{H}_2\text{O})$  tape recorder signal with smaller amplitude in the MIPAS retrieval.

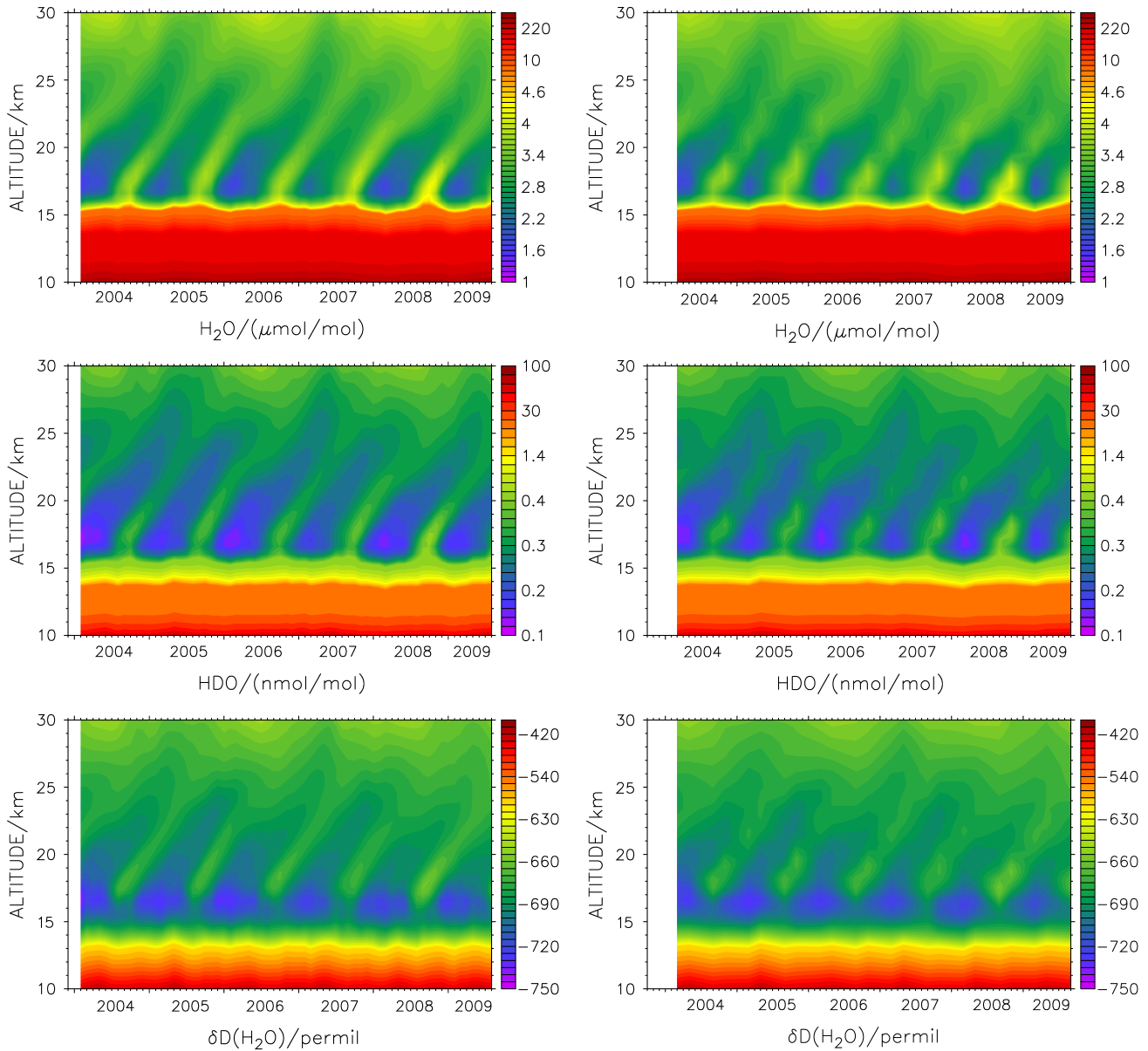
By analysing ACE-FTS data, Randel et al. (2012) found a tape recorder signal in  $\text{H}_2\text{O}$  and in HDO, but could not find

a corresponding pattern in  $\delta\text{D}(\text{H}_2\text{O})$ . The lower stratosphere in Fig. 7 of Randel et al. (2012) shows distinct seasonally varying maxima and minima of  $\delta\text{D}(\text{H}_2\text{O})$ , these, however, hardly propagate upwards in time. For comparison with the ACE-FTS data, the tropical  $\text{H}_2\text{O}$ , HDO and  $\delta\text{D}(\text{H}_2\text{O})$  in the stratosphere of the EMAC simulation are displayed in Fig. 7 for the same period and altitudes as in Randel et al. (2012).

The left panels show the monthly averages of the respective values. Qualitatively these agree fairly well with the ACE-FTS observations by Randel et al. (2012). The increase of HDO and therewith  $\delta\text{D}(\text{H}_2\text{O})$  with altitude in the stratosphere is too weak, however. At 30 km  $\delta\text{D}(\text{H}_2\text{O})$  exhibits values below  $-600\text{‰}$ , while in the ACE-FTS retrieval values of around  $-500\text{‰}$  can be seen. Similarly to the comparison with the MIPAS retrieval, also the lower stratosphere in EMAC is dryer in  $\text{H}_2\text{O}$  and in HDO. Still, a clear tape



**Fig. 6.** Altitude-time diagrams of  $\text{H}_2\text{O}$ ,  $\text{HDO}$  and  $\delta\text{D}(\text{H}_2\text{O})$  in the tropics ( $15^\circ\text{S}$ - $15^\circ\text{N}$ ). Left: EMAC simulation, right: MIPAS observations.



**Fig. 7.** Altitude-time diagrams of  $\text{H}_2\text{O}$ ,  $\text{HDO}$  and  $\delta\text{D}(\text{H}_2\text{O})$  in the tropics ( $15^\circ\text{S}$ - $15^\circ\text{N}$ ) from the EMAC simulation. Left: monthly averages; right: averages of only February, April, August and October, as in Randel et al. (2012). The plotting algorithm linearly interpolates between the available months.

recorder signal can be observed in all three panels, with minimum values in the lower stratosphere during boreal winter and maxima during boreal summer.

For the right panels, the EMAC data has been filtered, using only the four months (February, April, August and October), which are also available in the ACE-FTS retrieval, to estimate the influence of the sparse temporal sampling on the tape recorder signals. This filtering somewhat blurs the tape recorder in all three panels, compared to the full data set. Especially the tape recorder in  $\delta\text{D}(\text{H}_2\text{O})$  appears to lose some

of its upward motion at around 20 km and generally becomes less obvious. Therefore, it can be assumed, that the sparse temporal sampling of ACE-FTS data is an issue in the evaluation of the  $\delta\text{D}(\text{H}_2\text{O})$  tape recorder and may well contribute to the indistinctness of the signal in the study by Randel et al. (2012).

Due to its instrumental design, the SMR satellite instrument measures  $\text{H}_2\text{O}$  and  $\text{HDO}$  orbitally shifted. This means, that  $\text{H}_2\text{O}$  and  $\text{HDO}$  are never measured at the same time and implies inaccuracies for the calculation of  $\delta\text{D}(\text{H}_2\text{O})$ . A com-

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parison of the  $\delta D(H_2O)$  tape recorder signal between EMAC and SMR has therefore not been conducted.

### 3.5 Summary and discussion

The results of the EMAC simulations were first assessed with respect to the isotope ratios in precipitation, in order to ensure the correct representation of the physical processes in the troposphere. The EMAC results agree very well with GNIP (IAEA, 2009) data and are almost identical with results from the ECHAM5-wiso (Werner et al., 2011) model. Hence the physical processes and fractionation effects of the water isotopologues in the hydrological cycle of EMAC are represented satisfactorily (see supplement).

The chemical HDO precursor  $CH_3D$  in the stratosphere was evaluated next. Here, a good qualitative agreement with chemically elaborate transport models by Ridal et al. (2001) and Ridal and Siskind (2002), as well as with balloon flights by Röckmann et al. (2011) is achieved. Differences in the model dynamics as well as in the fractionation coefficients can contribute to disagreements between the model results. The coarse horizontal resolution of the model, the possible drift of the balloon and associated local meteorological features and especially the sparsity of the measurements are the main issues in the comparison of  $\delta D(CH_4)$  in EMAC with observations.

As desired by Röckmann et al. (2011), this modelling approach can assist further studies, especially concerning the investigation of the isotope effects of the chemical sink reactions of methane isotopologues. For instance, Kaiser et al. (2002) and Röckmann et al. (2003) discuss the application of apparent rather than laboratory based fractionation factors ( $KIE_{app} = \sqrt{KIE}$ ) for long-lived trace gases that are removed in the stratosphere. While in the EMAC simulation the laboratory based fractionation factors lead to very good agreement with measurements, according to Röckmann et al. (2001) turbulent diffusion and mixing of air masses lead to significantly smaller fractionation factors. Anyhow, also more extensive measurements are desired for further evaluation of the methane isotope ratios in the stratosphere.

The mixing ratios of HDO and  $H_2O$  were compared with the results of satellite retrievals. Inconsistencies between the individual satellite retrievals (see Lossow et al., 2011) render it difficult to define a distinct result. In the UTLS, the HDO profiles of EMAC agree well with the SMR satellite observations, but the  $H_2O$  retrieval of SMR only reaches down to around 20 km. Compared to the ACE-FTS and the MIPAS satellite profiles, retrieved from measurements in the mid-infrared, the UTLS appears to be too dry in  $H_2O$  and in HDO in the EMAC model. Since, for  $H_2O$  this dry bias had already been discussed (see Jöckel et al., 2006) and HDO is, at first order, determined by  $H_2O$ , this dry bias for HDO is

not surprising. Reasons for this are the slightly too cold hygropause in the nudging data (see e.g. Liu et al., 2010) and the coarse horizontal resolution of the model. However, due to the rather low vertical resolution of the satellite observations, the hygropause is blurred in the retrievals and therefore,  $H_2O$  as well as HDO are overestimated in MIPAS (see e.g., Steinwagner et al., 2007).

For  $H_2O$ , this offset weakens between 30 and 40 km and becomes stronger again further above. At around 35 km, the simulated  $H_2O$  increases stronger with altitude than the satellite observations show. Above 40 km, simulated  $H_2O$  increases at much smaller rates compared to the satellite observations. These inconsistencies are most likely caused by the simplified methane oxidation scheme, since they are obvious during all four months where observations of ACE-FTS are available.

HDO in contrast shows a too weak increase with altitude from around 30 km upwards. Largely, this is attributable to the lack of the intermediate reactions containing HD in the  $CH_3D$  oxidation chain in the model. Although we accounted for the deuterium storage in HD (using a relation by McCarthy et al., 2004), the isotope effects of the intermediate reactions on HDO are not taken into account in the model. As has been shown by Rahn et al. (2003) and Röckmann et al. (2003), stratospheric  $\delta D(H_2)$  increases with altitude, while the mixing ratios of molecular hydrogen are rather constant. The conversion of isotopically very heavy molecular hydrogen to water would therefore increase the HDO mixing ratios strongly, while leaving  $H_2O$  largely unaffected. Additionally, the influence of the oxidation of  $CH_3D$  itself also increases with altitude. This implies, that the importance of the intermediate reaction with HD on HDO, which is not considered here, increases with altitude too. Our simplifications can therefore explain the offset in HDO mixing ratios between the EMAC simulation and the satellite retrievals and the different behaviour of  $H_2O$  and HDO. This leads to too low  $\delta D(H_2O)$  values in the upper parts of the stratosphere.

In future studies, the chemical deficiencies of the H2OISO submodel can be reduced, by including an explicit representation of HD to take into account the intermediate reactions from  $CH_3D$  to HDO. However, according to Zahn et al. (2006) the reaction rates of HD and especially the fractionation effects of these reactions are poorly quantified and hence can be subject to large uncertainties. As an alternative, however more sophisticated modelling approach, the water isotopologues can be included to the isotope scheme by Gromov et al. (2010), by extending it to the deuterium cycle. This features a number of isotopic species and reactions, which also affect the hydrogen reactions.

The tape recorder signals of  $H_2O$ , HDO and  $\delta D(H_2O)$  of the EMAC simulation were compared to the MIPAS retrieval (see Steinwagner et al., 2007, 2010). In the lower stratosphere, EMAC and MIPAS are closest during winter, where the differences can mostly be explained with the

low vertical resolution of the MIPAS sampling. The summer months, however, show much stronger deviations between model and observations, with too low values in all quantities in the simulation, especially however, in HDO and  $\delta D(H_2O)$ . This suggests deficiencies in the model physics, for example the underrepresentation of overshooting convection in the convection scheme (Tiedtke-Nordeng; Tiedtke, 1989; Nordeng, 1994) applied here. According to studies by Dessler et al. (2007) and Bolot et al. (2013) a more complete representation of this effect can enhance the mixing ratios of HDO, while  $H_2O$  is not hardly affected. This would increase  $\delta D(H_2O)$  in the UTLS especially during NH summer, since convective events are more frequent during that time of the year. Thus, a better representation of the  $\delta D(H_2O)$  tape recorder in the model could be achieved. However, the inconsistent vertical resolutions between  $H_2O$  and HDO in the MIPAS retrieval can be the reason of a spurious amplification of the  $\delta D(H_2O)$  tape recorder signal in the observations (Orbe et al., 2013). The correction of this artefact in the MIPAS data is subject of current investigations and may also lead to a better agreement between the EMAC and MIPAS amplitudes of the  $\delta D(H_2O)$  tape recorders.

The influence of convective ice lofting on the isotopic composition of stratospheric water vapour in this simulation is analysed in part 2 of the article (Eichinger et al., 2014). In future studies, the quality of other convection schemes could be tested with respect to the results of HDO in the UTLS and a possibly better representation of overshooting convection. This, however, requires the implementation of the water isotopologues into the alternative convection schemes.

The phase shift of two to three months between the tape recorders of MIPAS and EMAC, which can be observed in all three quantities, is not easily explainable. Due to the coarse vertical resolution of the MIPAS retrieval there might be a delay in the retrieved tape recorder signals of up to one month. There is no evidence that the seasonality of convection in EMAC is in phase shifted. Simulated patterns of moisture in the UTLS compare well with observations and analyses of the seasonal cycle of zonal mean precipitation (convective and large-scale clouds) and integrated water vapour conducted by Hagemann et al. (2006) and Tost et al. (2006) also show accordance with observations.

The seasonal cycle of lower stratospheric  $\delta D(H_2O)$  in the ACE-FTS retrieval (see Randel et al., 2012) shows a different behaviour than that of the MIPAS retrieval and the EMAC simulation. The too low  $\delta D(H_2O)$  values in EMAC compared to ACE-FTS, especially during summer and in the upper stratosphere, are consistent, however, Randel et al. (2012) would not find the tape recorder signal in  $\delta D(H_2O)$ . The  $\delta D(H_2O)$  tape recorder in EMAC is weaker than the corresponding signals in  $H_2O$  and HDO. Still, the pattern is clearly recognisable. A possible reason for the lack of the upward propagation of the seasonal cycle of  $\delta D(H_2O)$  in the ACE-FTS observations is the sparse temporal sampling of

the instrument in the tropics. A filtering of the EMAC data to its sampling reduces the apparent temporal upward motion of the  $\delta D(H_2O)$  tape recorder in the UTLS between around 17 and 20 km. Since  $\delta D(H_2O)$  does not increase strongly enough with altitude in the EMAC simulation due to the incomplete chemistry, a weaker  $\delta D(H_2O)$  tape recorder signal is expected for the correction of this effect. It would blur the tape recorder signal in EMAC above 20 km and therewith suppress its upward propagation. Below around 25 km, however, the chemical influence is very weak and thus some tape recorder signal will remain. The effect of methane oxidation on the  $\delta D(H_2O)$  tape recorder will be analysed more deeply in the companion part 2 of the article.

For more quantitative comparisons between models and observations, more sophisticated methods will be applied in future evaluations. The vertical resolution of the EMAC data can be transformed to the resolution of the respective satellite retrieval using their averaging kernels. Also, the cloud filtering methods used for the satellite data can be applied to the model data. This elaborate evaluation can possibly assist to reduce the discrepancies between model results and observations and reveal the model and the measurement insufficiencies more precisely. In the present and in the companion (Part 2) study, however, a more comprehensive understanding of the  $\delta D(H_2O)$  tape recorder is desired, which can best be achieved with the complete data record.

#### 4 Conclusions

The EMAC model system has been extended with the submodel H2OISO, which contains a separate hydrological cycle, comprising the water isotopologues HDO and  $H_2^{18}O$  and their physical fractionation effects. The good agreement with measurement data and the ECHAM5-wiso model leads to the conclusion, that this expanded model system represents the state of the art of water isotopologue-enabled atmosphere GCMs. Moreover, a parameterisation of the oxidation of  $CH_3D$  was included to the H2OISO submodel. The sink reactions of  $CH_3D$ , which include reaction- and partly temperature-dependent kinetic fractionation effects, determine the chemical production of HDO. This rather simple parameterisation for the methane isotopologue  $CH_3D$  is apparently quite robust and can be applied for further studies concerning the isotopic composition of methane in the stratosphere. The comparison of stratospheric HDO profiles with satellite observations reveals a qualitatively good agreement. However, systematic discrepancies can be observed. These can be associated with uncertainties in the convection scheme, the simplified representation of methane oxidation and, especially in the upper stratosphere, the negligence of the fractionation effects involving molecular hydrogen. The comparison of the stratospheric tape recorder signal in  $H_2O$ , HDO and  $\delta D(H_2O)$  between EMAC and satel-

lite observations is difficult. Considering all the uncertainties in the measurements and in the model, the overall representations of the tape recorder signals in EMAC are reasonable. The  $\delta D(H_2O)$  tape recorder simulated by EMAC at least partly resolves the discrepancies between the divergent conclusions from the MIPAS and the ACE-FTS satellite retrievals. The EMAC  $\delta D(H_2O)$  tape recorder ranges between the pronounced tape recorder from MIPAS and the missing upward propagation of the seasonal signal in ACE-FTS observations. The revisions of the individual insufficiencies of the retrievals and the model, however, are expected to generate more compatible results. Despite the quantitative differences in stratospheric  $\delta D(H_2O)$  between EMAC and satellite observations, the conclusion can be drawn, that the new MESSy submodel H2OISO, used in the framework of the EMAC model, provides the possibility to attain additional insights into the mechanisms, which control the stratospheric water vapour budget. The physical and chemical properties of the isotopic composition of water allow new investigations, with respect to the processes and pathways which control the stratospheric water vapour budget. The H2OISO submodel will be available in future EMAC versions as an additional option for all users.

## Appendix A

### Applied MESSy submodels

**CLOUD:** Original **cloud** and cover routines from ECHAM5 as MESSy submodel (Roeckner et al., 2006, and references therein).

**CONVECT:** This submodel calculates the process of **Convection**. It consists of an interface to choose different convection schemes and the calculations themselves (Tost et al., 2006).

**CVTRANS:** The **Convective Tracer Transport** submodel calculates the transport of tracers due to convection. It uses a monotonic, positive definite and mass conserving algorithm following the bulk approach (Tost, 2006).

**GWAVE:** Hines non-orographic **gravity wave** routines from ECHAM5 as MESSy submodel (Roeckner et al., 2006).

**RAD4ALL:** ECHAM5 radiation scheme as MESSy submodel (Roeckner et al., 2006; Jöckel et al., 2006).

**TROPOP:** Submodel for **Tropopause** (WMO + PV) and other diagnostics (Jöckel et al., 2006).

**JVAL:** This submodel is for fast online calculation of **J-values** (photolysis rate coefficients) using cloud water content and cloudiness calculated by the base model and/or climatological ozone and climatological aerosol (see Landgraf and Crutzen, 1998; Sander et al., 2014).

**TNUDGE:** The submodel **'Tracer Nudge'** is used for Newtonian relaxation of user-defined tracers as pseudo-emissions (Kerkweg et al., 2006).

Supplementary material related to this article is available online at <http://www.gmd.net/journalurl/@pvol/@fpage/@pyear/@journalnameshortlower-/@pvol-/@fpage-/@pyear-supplement.pdf>.

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