The Impact of Dust Storms on the Arabian Peninsula and the Red Sea

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P. Jish Prakash, Georgiy Stenchikov, Stoitchko Kalenderski, Sergey Osipov, and Hamza Bangalath

Division of Physical Sciences and Engineering, King Abdullah University of Science and
Technology, Thuwal, Saudi Arabia

8 Correspondence to: Georgiy Stenchikov (georgiy.stenchikov@kaust.edu.sa)

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10 Abstract

11 Located in the dust belt, the Arabian Peninsula is a major source of atmospheric dust. 12 Frequent dust outbreaks and some 15 to 20 dust storms per year have profound effects on all 13 aspects of human activity and natural processes in this region. To quantify the effect of severe 14 dust events on radiation fluxes and regional climate characteristics, we simulated the storm 15 that occurred on 18-20 March 2012 using a regional weather research forecast model fully 16 coupled with the chemistry/aerosol module (WRF-Chem). This storm swept over a 17 remarkably large area affecting the entire Middle East, North-Eastern Africa, Afghanistan and 18 Pakistan. It was caused by a southward propagating cold front and associated winds activated 19 the dust production in river valleys of the lower Tigris and Euphrates in Iraq, the coastal areas 20 in Kuwait, Iran, and the United Arab Emirates, Rub al Khali, An Nafud and Ad Dahna 21 deserts, and along the Red Sea coast on the west side of the Arabian Peninsula. Our 22 simulation results compare well with available ground-based and satellite observations. The 23 total amount of dust generated by the storm reached 93.76 Mt. About 80% of this amount 24 deposited within the calculation domain. The Arabian Sea and Persian Gulf received 5.3 Mt, 25 and the Red Sea 1.2 Mt. Dust particles bring nutrients to marine ecosystems, which is 26 especially important for the oligothrophic Northern Red Sea. However, their contribution to 27 the nutrient balance in the Red Sea remains largely unknown. By scaling the effect of one storm to the number of dust storms observed annually over the Red Sea, we roughly estimate 28 29 the annual dust deposition to the Red Sea, associated with major dust storms, to be 6 Mt.

1 **1 Introduction**

Mineral dust is the dominant atmospheric aerosol (Buseck and Posfai, 1999). It plays an important role in the Earth's climate system, although the magnitude and even the sign of its radiative effect at the Top-of-the-Atmosphere (TOA) remain uncertain. Most airborne dust is generated in desert and semi-desert areas and transported across local-to-global scales.

Dust storms lift millions of tons of dust into the atmospheric boundary layer. Such large 6 7 quantities of dust cause severe air pollution, reduce visibility, cause airport shut-downs, and increase traffic and aircraft accidents (Morales, 1979; Hagen and Woodruff, 1973; Middleton 8 and Chaudhary, 1988; Dayan et al., 1991; Chung and Yoon, 1996). Other environmental 9 impacts of dust include reduced soil fertility and crop damage, reduced solar radiation on the 10 11 surface and, as a consequence, decreased efficiency of solar devices, damaged telecommunications and mechanical systems, increased occurrence of respiratory diseases and 12 13 other impacts on human health (Hagen and Woodruff, 1973; Mitchell, 1971; Fryrear, 1981; Squires, 2007; Jauregui, 1989; Liu and Ou, 1990; Chung and Yoon, 1996; Nihlen and Lund, 14 15 1995; Longstreth et al., 1995; Bennett et al., 2006; Bennion et al., 2007). Dust deposition to 16 oceans, however, provides nutrients to ocean surface waters and the seabed (Talbot et al., 1986; Swap et al., 1996 and Zhu et al., 1997). 17

18 Direct and indirect atmospheric radiative impact by dust has implications in global climate 19 change and presently is one of the largest unknowns in climate model predictions. The 20 African and Asian low-latitude deserts, the so-called dust belt, are the major sources of dust for the entire world (Pye, 1987; Prospero et al., 1987, Arimoto et al., 1989; Laurent et al., 21 2008; Uno et al., 2005). The dust belt includes the Sahara desert, arid and semiarid regions in 22 Arabia and Central Asia, and the Taklamakan and Gobi deserts in East Asia. During severe 23 24 dust storms, dust from East Asia can reach far beyond the continent, drifting over the Pacific Ocean to the west coast of North America (Husar et al., 2001; Tratt et al., 2001; Mc Kendry et 25 al., 2001). Similarly, Saharan dust crosses the Atlantic and affects both Americas (Prospero et 26 27 al., 1987; Reid and Maring, 2003; Haywood et al., 2003).

The Sahara is the world's largest dust source (Washington et al., 2003; Shao et al., 2011). The characteristics of Saharan dust were studied experimentally since the 1990s (Jayaraman et al., 1998; Satheesh and Ramanathan, 2000; Haywood and Boucher, 2000). The Puerto Rico Dust Experiment (PRIDE) (Reid and Maring, 2003) and the Saharan Dust Experiment (SHADE) (Haywood et al., 2003) focused on Saharan dust during long-range transport. Two comprehensive field campaigns were conducted in 2006 (Heintzenberg, 2009) and 2008 (Ansmann et al., 2011) under the framework of the Saharan Mineral Dust Experiment (SAMUM) to quantify the optical properties and the radiative impact of Saharan dust near source regions and in the aged mixed plume from Saharan dust and biomass burning aerosols transported west off the African continent. The Fennec 2011 campaign (Washington et al., 2012) made significant advances by providing new measurements close to dust sources over the remote Sahara.

8 It has long been recognized that radiative effects of mineral aerosols are an important driver of 9 the climate (Tegen and Fung 1994, 1995; Tegen et al. 1996; Lie et al., 1996; Andreas, 1996). 10 Early attempts at dust modeling using global models were made by Westphal et al. (1987, 11 1988), Gillette and Hanson (1989), Joussaume (1990), Tegen and Fung (1994, 1995). The 12 presence of dust affects atmospheric heating rates and thus atmospheric stability and 13 circulation, which together with changes in the surface energy balance affect the hydrologic 14 cycle (Miller and Tegen, 1998; Miller et al, 2004b).

15 Miller and Tegen (1998) conducted one of the first studies of the radiative effects of dust 16 using a Global General Circulation Model (GCM) coupled with a mixed ocean module to show that dust aerosols reduce the global average surface net radiation by nearly 3 W m^{-2} 17 during summer in the northern hemisphere. Tegen et al. (1996) showed that dust from 18 disturbed soils causes a decrease in the net surface radiation by about 1 W m^{-2} . Recently, 19 Balkanski et al. (2007) confirmed that the global average dust-related surface net radiative 20 flux change is in the range of -1.11 to 0.92 W m⁻². However, the radiative effect of dust in 21 desert regions during dust outbreaks could be two orders of magnitude stronger (Kalenderski 22 23 et al., 2013). A comprehensive analysis of simulated global dust distributions and their 24 radiative effects was recently conducted under the scope of the Aerosol Inter-comparison (AeroCom) project (Textor et al., 2006; Stier et al., 2007). 25

Regional-scale dust effects were addressed using fine-resolution limited area models (e.g., Nickovic et al., 2001; Uno et al., 2001; Lu and Shao, 2001; Kaskaoutis et al., 2008; Tegen et al., 2013). A regional modeling system was developed for simulations of Saharan dust emissions, transport and radiative effects (Heinold et al., 2008, 2009; Laurent et al., 2010; Tegen et al., 2010) during SAMUM-1. A regional model was also used to simulate the spatiotemporal evolution of the mixed plume of Saharan dust and biomass burning aerosols (Heinold et al, 2011a), the radiative effect, and the dynamic response of the atmosphere due to
 Saharan dust and biomass burning (Heinold et al., 2011b) during SAMUM-2.

3 The studies mentioned thus far mostly focused on western and central North Africa, the Sahel, 4 and Sahara regions. There are relatively fewer measurement and modeling studies focusing on eastern North Africa and the Arabian Peninsula. This region has been severely under-sampled; 5 6 there are few available in-situ observations, and very few international research campaigns 7 have been conducted in this area compared with the Sahara. In the Arabian Peninsula, dust 8 storms and blowing dust are frequent events during most of the year. The major dust sources 9 in the Arabian Peninsula include the Tigris and Euphrates Rivers, the alluvial plain in Iraq and 10 Kuwait, the low-lying flat lands in the east of the peninsula along the Persian Gulf and the Ad 11 Dahna and the Rub Al Khali deserts (Kutiel and Furman, 2003; Shao, 2008). Goudie and Middleton (2006) provided an extensive discussion of the causes of dust storms in the Middle 12 East. Eastern Syria, northern Jordan and western Iraq are the source for most of the fine dust 13 14 particles (less than 50 µm in diameter) found in Arabian dust storms (NOAA, 2002). This dust is transported far into southern Arabia where it may have contributed to loess deposits on 15 parts of the Arabian Shield and even in the Asir Plateau. Sandstorms containing larger 16 particles (150 to 300 µm in diameter) are also frequent throughout Arabia, but rarely reach 17 18 above 15 m altitude. Most sand grains move by saltation and some by surface creep; they 19 rarely move in suspension by storm winds, haboobs, or dust devils.

20 Dust storms in the southern Arabian Peninsula are more frequent in summer. In the northern 21 Arabian Peninsula they occur mainly in spring. The peak of dust storm activity occurs usually 22 during the day time, when intense solar heating of the ground generates turbulence and local 23 pressure gradients (Middleton, 1986a, 1986b). Across the Arabian Peninsula, remotely sensed 24 aerosol and dust activity peaks during May to August (Prospero et al., 2002; Washington et al., 2003; Barkan et al., 2004; Goudie and Middleton, 2006). The United Arab Emirates 25 Unified Aerosol Experiment (UAE²) focused on dust in the southern Arabian Gulf region in 26 August to September 2004 to evaluate the properties of dust particles that converge in the 27 UAE region from numerous sources and their impact on the radiation budget (Reid et al., 28 29 2008). Mohalfi et al. (1998) studied the effect of dust aerosols on the synoptic system and 30 showed that dust aerosol radiative heating strengthens Saudi Arabian heat low. Kalenderski et 31 al. (2013) simulated a winter dust event that occurred in January 2009 over the Arabian Peninsula and the Red Sea to study various dust phenomena. Notaro et al. (2013) investigated 32

the temporal and spatial characteristics of Saudi Arabian dust storms using trajectory analysis.
 Using MODIS data, they found that the highest aerosol optical depth (AOD) is achieved
 during dust storms that originate from the Rub Al Khali and Iraqi Deserts.

This study aims to quantify the impact of dust on the Arabian Peninsula and Red Sea focusing 4 5 on the severe dust storm observed during 18-19 March 2012. This storm likely drew at least 6 some of its dust from Iraq, Iran, Kuwait and the Arabian Peninsula's Empty Quarter or Rub' 7 al Khali desert, sprawling over parts of Saudi Arabia, Yemen, Oman and United Arab 8 Emirates. The simulation period was from 00:00 UTC (Universal Time Coordinate) on 5 9 March to 00:00 UTC on 25 March 2012 with the model spin-up during first six days. Using 10 model output and observations, we provide an improved estimate of the storm's impact on the terrestrial and oceanic environment. 11

The remainder of this paper is organized as follows. Section 2 discusses the model's configuration and methodology. Section 3 examines the environmental conditions that led to development of the dust storm, estimates the dust emissions, load and deposition, discusses the structure of the dust storm, compares model results with available Aerosol Robotic Network (AERONET) and satellite observations, presents the vertical structure of atmosphere and dust distribution, and calculates radiative effect of dust. We offer a summary of the results in Sect. 4.

19

20 2 Methodology

In this study, we combine advanced high-resolution modeling of meteorological and dust processes with analysis of ground-based and satellite observations of aerosols and meteorological fields.

24 2.1 Modeling

The Weather Research Forecast (WRF) system is a mesoscale forecast model with an incorporated data assimilation capability that advances both the understanding and prediction of weather (Skamarock et al., 2008). WRF has been utilized in a variety of research and operational projects, from the scale of convective storms to the scale of continental weather patterns (Michalakes et al., 2005). WRF-Chem extends WRF by incorporating a chemistry module that interactively simulates emissions of aerosols and gases, their transport, turbulent and convective mixing, and chemical and microphysical transformations of trace gases and
 aerosols (Grell et al., 2005).

3 Here, we configured WRF-Chem with the RADM2 (Regional Acid Deposition Model 2) 4 photochemical mechanism (Stockwell et al., 1990), the Fast-j photolysis scheme (Wild et al., 2000), the Modal Aerosol Dynamics Model for Europe (MADE) and the Secondary Organic 5 6 Aerosol Model (SORGAM) aerosol model (Ackermann et al., 1998; Schell et al., 2001). We 7 also employed the Goddard Global Ozone Chemistry Aerosol Radiation and Transport 8 (GOCART) dust emission scheme (Ginoux et al., 2001) to calculate the influx of dust into the 9 atmosphere. GOCART simulates dust emissions as a function of surface wind speed, surface 10 erodibility, and surface wetness (Chin et al., 2002). Emission flux, F_p, for a specific aerosol 11 size group, p, is expressed as

12
$$F_p = CSs_p U_{10m}^2 (U_{10m} - U_t)$$
 if $U_{10m} > U_t$ (1)

where C is a dimensional constant coefficient of proportionality; S is the space-dependent 13 14 dimensionless erodibility field taken from Ginoux et al. (2001) and has a spatial resolution of $0.25^{\circ} \times 0.25^{\circ}$; U_{10m} is the wind speed 10 m above the ground; U_t is the threshold velocity of 15 wind erosion, which depends on particle size and surface wetness; sp is a particular particle 16 sizemode mass fraction emited in the atmosphere. The original GOCART emission scheme 17 18 was coupled with the eight-bin aerosol model. But in this study WRF-chem is configured 19 with the MADE/SORGAM that uses the modal approach with three log-normally distributed 20 modes (Aitken, Accumulation, and Coarse) to represent sulfate, nitrate, ammonium, organic 21 matters, black carbon, and sea salt. Mineral dust is assumed to have only accumulation and 22 coarse modes. The GOCART emission scheme (1) has been modified to couple with the MADE/SORGAM. It calculates a dust mass flux from the surface (see Eq. (1)) assuming 23 24 that, by default, $s_p=0.07$ for accumulation mode, which for emitted particles has the modal diameter D=0.6 μ m, width σ =2 μ m; and s_p=0.93 - for coarse mode with D=6 μ m, σ =2.2 μ m. 25 26 According to MADE/SORGAM formulation the modal diameters change in atmosphere due 27 to microphysical processes but the widths of both distributions remain fixed. The total 28 emission flux is adjusted using constant C in Eq. (1) to fit AERONET observations, as discussed by Zhao et al. (2010) and Kalenderski et al. (2013). In our simulations, we used C =29 $0.8 \text{ mg s}^2 \text{m}^{-5}$ to achieve consistency between the simulated and AERONET aerosol optical 30 thickness (AOD) observed from 11 to 24 March 2012. 31

1 To cover the entire area affected by the storm and to account for processes that were 2 responsible for the storm's development, we ran the WRF-Chem model in the spatial domain from 4° N to 40° N in longitude and 25° E to 80° E in latitude with 10 km horizontal 3 4 resolution. In our calculations, we therefore used 495 grid points in the west-east direction and 396 grid points in the south-north direction, as well as 40 vertical levels with the top of the 5 model domain at the 10 hPa level. The following physical parameterizations were used to 6 7 configure the WRF-Chem simulations: the Lin microphysics scheme; the Rapid Radiative 8 Transfer Model (RRTMG) for both longwave (LW) and shortwave (SW) radiation; the 9 Mellor-Yamada-Janjic (MYJ) boundary layer scheme; the Noah land surface model; the Grell 10 cumulus parameterization. Lateral boundary and initial conditions for meteorological fields 11 were provided by the National Centers for Environmental Prediction's (NCEP) global analysis 12 (FNL). We used NCEP's daily global sea surface temperature (SST) analysis (RTG SST HR) 13 to update SST every six hours. The various model domain configurations and physics options 14 are summarized in Table 1.

15 **2.2 Observations**

16 To test and constrain the model simulations, we used reanalysis fields as well as17 meteorological and aerosol observations available in the Middle East and surrounding areas.

18 **2.2.1** Space-borne aerosol instrumental observations

Aerosol Optical Depth (AOD) data from MODIS (Moderate Resolution Imaging 19 20 Spectroradiometer) sensors on board the Terra and Aqua satellites are widely used in aerosol 21 studies (Salomonson et al., 1989). The MODIS instrument provides high radiometric 22 sensitivity (12 bit) in 36 spectral bands ranging in wavelength from 0.4 µm to 14.4 µm. Two 23 bands are imaged at a nominal resolution of 0.25 km at nadir, with five bands at 0.5 km, and 24 the remaining 29 bands at 1 km. A ±55° scanning pattern of the Earth Observing Satellite (EOS) orbit at 705 km altitude achieves a 2330 km swath and provides global coverage every 25 one to two days. The MODIS daily Level 2 AOD data are produced at the spatial resolution 26 27 of a 10 km \times 10 km (at nadir). There are two MODIS aerosol data products, one containing data collected from the Terra platform and the other containing data collected from the Aqua 28 platform. To maximize the observation coverage, we used AOD retrievals over land and sea 29 30 derived from the dark target product (Remer et al., 2005, 2008) and the "deep blue" product over bright land surfaces (Hsu et al., 2004, 2006). The dark target ocean and land AOD 31

products were available from both Terra and Aqua, but the deep blue retrievals were only available from Aqua. In this study, we have used daily deep blue (level 2) data in combination with the standard ocean algorithm for comparison with the simulated aerosol optical properties.

Spinning Enhanced Visible Infrared Radiometer (SEVIRI) (Aminou et al., 2002) is a line 5 6 scanning radiometer currently located on board the European geostationary meteorological 7 satellite, Meteosat-9; it was previously on Meteosat-8. This instrument provides data in four 8 visible and near infrared channels and 8 infrared channels with a resolution of 3 km at nadir. 9 In this study, we used the 0.6 µm channel. An advantageous feature of SEVIRI is its ability to 10 image the study area with high temporal resolution of 15 minutes. This allows SEVIRI to 11 track aerosol events, which offers a great advantage over polar orbiting instruments, like MODIS, that usually sees the particular seen once a day. The SEVIRI's spatial resolution is, 12 however, coarser than that of MODIS but the aerosol gridded products for the Arabian 13 14 Peninsula from both instruments are roughly on the 10 km \times 10 km grid.

15 2.2.2 Ground-based aerosol observations

The Aerosol Robotic Network (AERONET) uses CIMEL Robotic Sun Photometers and 16 provides observations of AOD on up to 8 wavelength channels between 0.340 and 1.640 µm 17 (Holben et al., 1998) and angular distribution of sky radiance at four wavelengths (0.440, 18 19 0.675, 0.870, and 1.020 µm). In addition to the direct measurement of AOD, the inversion 20 algorithm retrieves physical and optical properties of aerosols that comprise the aerosol 21 refractive index, column average size distribution, single scattering albedo, and asymmetry 22 parameter. The maximum AERONET uncertainty in AOD retrieval is estimated to be 0.02 23 with the highest error in the ultraviolet wavelength (Holben et al., 1998; Eck et al., 1999) and 24 the calibrated sky radiance measurements typically have an uncertainty less than 5% (Holben 25 Cimel Sun Photometers are calibrated annually by comparison with an et al., 1998). 26 AERONET master instrument.

The AERONET data are at high time resolution but have limited spatial coverage. During the study period, AOD observations were available at Kuwait University (29.32° N, 47.97° E), at Mezaira (23.14° N, 53.77° E), and at our own site established in February 2012 on the King Abdullah University of Science and Technology (KAUST) campus (22.30° N 39.10° E). We chose cloud screened (Level 1.5) AERONET AOD for our model validation since level 2 data (cloud screened and quality assured data) were not available during the simulation period. The 1 Angstrom power law was used for comparing the simulated optical properties output at 0.60

2 µm with the AERONET measurements as follows:

3
$$AOD(0.600) = AOD(0.675) \times \left(\frac{0.600}{0.675}\right)^{-\alpha}$$
 (2)

4 where α is the Angstrom exponent calculated from the AERONET measurement as:

5
$$\alpha = \frac{\ln\left[\frac{AOD(0.440)}{AOD(0.675)}\right]}{\ln\left(\frac{0.675}{0.440}\right)}$$
 (3)

6 2.2.3 Weather station meteorological observations

7 There are meteorological data from 108 surface weather stations and 12 upper air stations 8 available in the area of interest. For the purposes of this study, we focus on the upper air 9 radiosonde data. The radiosonde (RS) is a balloon-borne instrument platform with radio 10 transmitting capabilities. Data from the radiosonde is interpreted at the launching station and 11 entered into a worldwide communications network. The time sampling of RS data is usually 12 twice daily at 00:00 UTC (midnight) and 12:00 UTC (noon). RS profiles contain pressure, 13 temperature, relative humidity, wind speed, and wind direction. Depending on the sensor and balloon used at each location, the vertical resolution of RS data typically have 40 levels 14 15 between 1000 and 50 hPa. We used temperature profiles from three RS stations for a comparison with modeled heating rates at 12:00 UTC on 19 March 2012. Stations are 16 17 identified by their World Meteorological Organization (WMO) five digit codes. The three stations over the Arabian Peninsula selected for this study include Jeddah (41024), Abu Dhabi 18 19 (41217) and Kuwait (40582). The vertical temperature RS profiles were retrieved from the 20 upper air archive at the University of Wyoming 21 (http://weather.uwyo.edu/upperair/sounding.html).

22 2.2.4 Reanalysis output

To compare meteorological fields and spatial patterns, we use output from the European Centre for Medium-Range Weather Forecast (ECMWF) ERA-Interim (ERA-I) reanalysis (Dee et al., 2011). It is the latest global atmospheric reanalysis produced by the European Centre for Medium-Range Weather Forecast as a transition between ERA-40 and a future

1 reanalysis project. It provides information on a large variety of surface parameters (3-hourly), 2 describing weather as well as ocean-wave and land-surface conditions and 6-hourly upper-air parameters (37 pressure levels up to 1 hPa), on a $0.7^{\circ} \times 0.7^{\circ}$ grid. ERA-I uses an improved 3 atmospheric model and a more sophisticated data assimilation method (4D-Var) for 4 atmospheric analysis compared with ERA-40. Information about the current status of ERA-I 5 production online 6 and availability of data be found can at 7 (http://apps.ecmwf.int/datasets/data/interim full daily/).

8

9 3 Results

10 The dust storm on 18-22 March 2012 covered an extremely large area and disrupted human 11 activities in Iraq, Iran, Kuwait, Syria, Jordan, Israel, Lebanon, UAE, Qatar, Bahrain, Saudi 12 Arabia, Oman, Yemen, Sudan, Egypt, Afghanistan, and Pakistan. Many airports in these 13 countries were shut down because of dust impact on visibility and machinery.

14 **3.1** Meteorological conditions and storm spatial-temporal development

15 Figure 1 shows geopotential height from WRF-Chem simulations and ERA-Interim reanalysis 16 at 500 hPa at 00:00 UTC on 17-19 March 2012. Sea-level pressure for the same days at 00:00 UTC is shown in Fig. 2. The model results compare well with the reanalysis fields in both 17 18 Figures. On 17 March there was a trough in mid troposphere at 500 hPa stretching over 19 eastern Mediterranean Sea, Syria, Iraq, western Iran and northern Saudi Arabia. This upper air 20 disturbance caused a significant pressure decrease in the lower atmosphere and a low-pressure 21 system consequently developed in eastern Iraq and northwestern Iran with a central sea-level 22 pressure of 1008 hPa. A high pressure is seen over northeastern Afghanistan and Tajikistan. The spatial distribution of 10m wind bars at 09:00 UTC from 17 to 19 March is shown in Fig. 23 3. Strong surface winds of 15 m s⁻¹ are observed over northwestern Iraq on 17 March at 09:00 24 UTC (Fig. 3a). 25

On 18 March the high level trough deepened and moved southeastward into the northern Arabian Peninsula covering Iraq and western Iran. A surface high pressure builds up over the eastern Mediterranean Sea. The presence of a low pressure in southern Saudi Arabia, the Gulf of Oman and central Iran in combination with a high in the eastern Mediterranean Sea enhanced a strong pressure gradient along the Arabian Peninsula. As a result, a strong surface wind of 15 m s⁻¹ developed to the west of the low near the Persian Gulf in southeastern Saudi Arabia and Qatar and to the northeast of UAE (Fig. 3b). This is well above the generally
assumed threshold wind speed (5 to 6 m s⁻¹) for dust suspension (Gillette, 1978). These strong
winds forced the dust emission and transported dust to the south and southwest of the Arabian
Peninsula.

On 19 March the high level trough moved towards eastern Iran and western Afghanistan. The 5 6 low-pressure system in the southern Arabian Peninsula was overtaken by the high-pressure 7 system. A low-pressure system that developed in central Pakistan in combination with the 8 high pressure in the northern Arabian Peninsula increased surface winds in southern Iran up to 20 m s^{-1} (Fig. 3c). One of the most identifiable synoptic features in the Arabian Peninsula 9 10 during the simulation is the low-pressure system that moved southeastward through the Peninsula on 19 March 2012 suggesting that the dust storm was closely related to the south 11 and southwestward propagation of a cold front. Figure 4 shows the simulated near-surface 12 wind field and dust concentration at 08:00 UTC and 15:00 UTC on 18 March 2012. 13 14 Comparisons with synoptic observations confirm that the flow field is well simulated. The 15 cold front is a line-shaped narrow region stretching from the western coast of Saudi Arabia to Eastern UAE, as can be seen from the dense temperature contours (Fig. 4a). Strong 16 northeasterly wind (maximum 15 m s^{-1}) prevailed behind the cold front, resulting in strong 17 dust emissions and high dust concentrations. The dust was carried southwestward by the 18 19 northeasterly wind to converge on the frontal area, resulting in high dust concentrations near 20 the front. The dust was then transported westward over the red sea along the front and towards 21 the southern Arabian Peninsula. In southern Saudi Arabia, the dust front rapidly advanced to 22 the south and southwest. Figure 4 shows that within a time period of seven hours, the dust 23 front advanced more than 150 km. At 15:00 UTC 18 March (Fig. 4b), dust was wide spread, 24 with the same maximum dust concentration as that at 08:00 UTC.

Dust and sand lift both ahead of and behind cold fronts; winds tend to be stronger behind the
front than ahead of it (Sissakian et al., 2013). Behind an advancing cold front, southern Iraq
and northern Saudi Arabia experienced blowing dust or shamal-like conditions (Hamidi et al.,
2013).

3.2 The vertically integrated mass balance and optical depth of the dust

30 Dust emissions are calculated in the model interactively from Eq. (1) using simulated surface 31 winds that drive dust emissions. To ensure that the model could adequately generate dust, we

1 compare the simulated 10m wind field (U_{10m}) with the output from the ECMWF Interim 2 global reanalysis, ERA-I. The U_{10m} is extracted from ERA-I and interpolated to WRF projections for comparison. The simulated and reanalysis fields are well correlated (not 3 shown) with a time-averaged correlation coefficient of 0.79. Figure 5a shows the erodibility 4 5 field S used in Eq. (1) depicting the major dust emission areas. Figures 5b-d show simulated dust emissions averaged over 24 hour periods for 17, 18 and 19 March 2012, respectively. On 6 7 17 March (Fig. 5b) the most intense simulated emissions were located in the lower Tigris and 8 Euphrates river valleys and in newly dry lakebeds, which are filled with loose fine-grain soils 9 and have high susceptibility. These are known major sources of dust in Iraq. The maximum dust emission rate exceeded 300 μ g m⁻² s⁻¹. On 18 March (Fig. 5c) dust was generated in the 10 11 wide Arabian Desert area (The Rub' al Khali, An Nafud and Ad Dahna) with almost same maximum intensity as on the previous day. A dust source belt about 50 km wide along the 12 13 west coast (22° N to 25° N) of the Arabian Peninsula was also active on 18 March. On 19 14 March (Fig. 5d), the dust source regions were along the eastern coast of the Persian Gulf, 15 coastal areas of the Arabian Sea near Iran and Pakistan, and on the southern flanks of the 16 mountain chain along the Arabian Sea coast. The active dust source in this region was a small intermountain valley centered at 27.5° N, 59° E. At the center of this valley, there is a large, 17 1087 km², dry salt lake, Hamum-e-JazMurian. The majority of dust at the eastern coast of the 18 19 Persian Gulf was generated from this lake. Dust source regions along the western coast of the 20 Arabian Peninsula were also active. The area of dust generation was relatively compact but maximum dust emissions exceeded 500 μ g m⁻² s⁻¹ due to strong winds (Fig. 3). 21

22 The spatial distribution of daily average simulated dust deposition rates for 17, 18 and 19 23 March 2012 are shown in Fig. 6. There was no precipitation during the simulation period and 24 therefore the dominant dust removal mechanism was dry deposition. The deposition velocity 25 is higher for large particles therefore coarse dust particles tend to deposit closer to a source than the fine ones. Figure 6 shows that the maximum dust deposition exceeds 90 μ g m⁻² s⁻¹, 26 27 about one third of maximum dust emissions rates. On 17 March (Fig. 6a) a lot of dust was 28 deposited in northeastern Persian Gulf relatively far from the source regions. On 18 March the strongest dust deposition happened in wide surroundings of the dust sources, as well as in the 29 30 Persian Gulf and the Red Sea (Fig. 6b). On 19 March (Fig. 6c), dust was deposited in the Red 31 Sea, the Gulf of Oman and the Arabian Sea off the Iran and Pakistan coasts.

1 Total dust emission/deposition are computed by integrating emission/deposition over the 2 domain area. Figure 7a shows time series of daily total emission and deposition (in Mt day⁻¹) for the simulation period of 11 to 24 March, 2012. The daily emissions (F_s) and depositions 3 (F_d) are of the same order of magnitude. The maximum daily emission, $F_s=13.10$ Mt day⁻¹, 4 and maximum daily deposition, $F_d = 10.06$ Mt day⁻¹, occurred on 18 March. Dust was 5 dispersed and deposited over much wider areas than those where dust was generated. The 6 maximum total hourly dust emission and deposition rates of 0.85 Mt hr^{-1} and 0.52 Mt hr^{-1} , 7 8 respectively, occurred in the day time at about 07:00 UTC on 18 March (not shown). This is 9 expected because in the daytime the atmosphere is unstable and the momentum transfer from the free troposphere to the surface, which enhanced the surface winds, is more efficient. Dust 10 11 emissions are substantially weaker in the night when the boundary layer collapses.

The comparison with other similar studies shows that the dust storm of 18-22 March 2012 is 12 in the range of the most powerful dust events considered in the literature (Zhao et al., 2010; 13 14 Shao et al., 2003). A portion of the dust deposited in the ocean provides nutrients to marine 15 ecosystems and increases the ocean's net productivity. Dust emitted from southern Iraq, Iran 16 and the eastern Arabian Peninsula was mainly deposited in the Persian Gulf and the Arabian 17 Sea. Dust emitted from the western coast of the Arabian Peninsula was mostly deposited in 18 the Red Sea. Time series of daily total dust deposition over the ocean areas (including the Red 19 Sea, Persian Gulf, and parts of the Arabian Sea that are included in the domain) are plotted in 20 Fig. 7b. Over the ocean areas, the dust load increased in the course of the dust episode from 21 17 to 20 March, reaching its maximum on 20 March. This delay in time is expected, as dust was transported from land to ocean. The maximum dust deposition of 1.5 Mt day^{-1} over ocean 22 23 areas also occurred on 20 March. Over the Red Sea, the maximum dust deposition of 0.23 Mt day⁻¹ occurred on 19 March. The dust balance over the entire simulation period is 24 25 summarized in Table 2. Over the 11-day simulation period, about 93.76 Mt of dust were 26 emitted and about 73.04 Mt of dust were deposited over the entire study domain. The total 27 dust deposited on the land was about 66.48 Mt and over the ocean around 6.56 Mt, out of which 1.20 Mt were deposited on the Red Sea. 28

Although the magnitude of dust deposition over the oceans is not accurately known, there is evidence that dust deposition could have a significant impact on chemical and biological processes in the oceans (Martin, 1990; Watson et al. 2000; Fan et al. 2006; Sunda and Huntsman, 1997). In the present study, we have calculated the amount of dust deposited over the ocean and particularly in the Red Sea. Aeolian deposition is especially important for the
 Red Sea because this sea has very little fresh water discharge from the coastal areas that is
 mostly associated with flash flood events.

4 To calculate the total frequency and spatial distribution of severe dust events over the Arabian 5 Peninsula and the Red Sea region we choose to analyzed the MODIS satellite images for the 6 2002–2013 period, which were provided under the scope of natural hazard reports in this area 7 by NASA's rapid response system (http://earthobservatory.nasa.gov/NaturalHazards). For this 8 12-year period, the number of cases when a dust plume covered more than 20% of the Middle 9 East area reaches 237 or about 20 annually. The number of dust storms affected the Red Sea is 10 71 or about 5–6 annually. We counted dust events based on their extent and optical depth 11 estimate. The obtained number of regional storms is consistent with the results of the previous studies (e.g., Rezazadeh et al, 2013 and Prospero et al., 2002). Thus 5-6 severe storms per 12 year would deposit to the Red Sea about 6 Mt of mineral dust. This figure is a rough estimate. 13 14 An extended analysis of the intensity and variability of the dust storm activity over the 15 Arabian Peninsula, as well as contribution of dust deposition during fair weather conditions, 16 are required to make this estimate more certain.

Figure 8a shows the simulated spatial pattern of the dust load at 10:00 UTC on 19 March. 17 18 Dust plumes were seen spreading over the Gulf of Oman, the Red Sea and the southern coast of the Arabian Peninsula. The highest vertically integrated dust load exceeded 4.5 g m^{-2} over 19 20 the Gulf of Oman and southern Saudi Arabia bordering Yemen. The image taken by 21 Aqua/MODIS at 09:50-10:05 UTC (Fig. 8b) shows a significant dust load over the entire 22 southern Arabian Peninsula and its coastline along the Arabian Sea as well as over the Persian 23 Gulf. The simulated spatial distribution of the dust loads in Fig. 8a are quite consistent with 24 the Modis dust plume image in Fig. 8b at most locations, suggesting that the model correctly 25 predicted the dust event. Both modeled and satellite imagery clearly show dust aerosols covering not only the land but also very large areas of the oceans. 26

Figures 9a–c show a spatial distribution of AOD from WRF-Chem, MODIS and SEVERI retrievals, respectively. Since the number of passes of the MODIS instrument on board the Aqua satellite over the domain of interest during the time of interest is limited, two 0.55 μ m retrievals at 09:50 and 10:05 UTC on 19 March were used for comparisons with the simulated AOD (0.60 μ m) at 10:00 UTC. A combination of both standard-ocean and deep blue products was used to get the maximum spatial coverage over the simulated domain. The model captures the spatial distributions of AOD well. Both modeled and observed optical fields
 indicate a high dust plume over the Persian Gulf, Gulf of Oman, Red Sea and southern
 Arabian Peninsula.

4 Figures 9d–f compare hourly simulated and observed AODs at three AERONET sites 5 (KAUST campus, Kuwait University and Mezaira) during the simulation period. The 6 AERONET 15-minute observations were averaged hourly to compare with modeled values. 7 The time series match well with each other most of the time. Even though the magnitude of 8 the simulated AOD at the KAUST campus was lower compared with that of AERONET 9 during 19 March, the simulations were able to reproduce the peak and the temporal 10 dependence quite well.

3.3 Vertical mixing of dust within the boundary layer and free troposphere

The vertical distribution of dust aerosols is important for estimating their radiative effect and their impact on circulation and the climate (Forster et al., 2007; Claquin et al., 1998; Huang et al., 2009; Zhu et al., 2007). The evolution and diurnal cycle of the atmospheric boundary layer directly affects the life cycle and vertical redistribution of dust particles (Cuesta et al., 2009).

16 **3.3.1 Vertical structure of a dust layer**

17 Most of dust aerosols resides in the atmospheric boundary layer. The portion of aerosols that 18 moves above the boundary layer to the free troposphere is subject to long-range transport and 19 therefore is of particular interest. N'Tchayi Mbourou et al. (1997), Engelstaedter et al. (2006), 20 and Chaboureau et al. (2007) reported that the diurnal cycle of dust in the Sahel and Sahara is 21 the result of daytime boundary-layer dry convection affecting the uplift and transport of dust in the almost 6 km deep boundary layer (Gamo, 1996; Carolina and Martin, 2012). The 22 23 Planetary Boundary Layer (PBL) can be influenced by different mechanisms associated with the coastal breezes, terrain and surface properties (Warner and Sheu, 2000). The Asir 24 25 mountain range in Western Arabia runs along the Red Sea coast and is the highest in the south 26 where its altitude reaches 3 km. During the day, the surfaces of the mountains heat the air 27 higher up in the atmosphere, quicker than the ocean surface can. This leads to enhancing 28 subsidence over the Red Sea and suppresses the growth of daytime PBL (e.g., Kuwagata and 29 Kimura, 1995; Whiteman, 1982; Kimura and Kuwagata, 1995; Bader and McKee, 1983). 30 Upward motion over heated elevated terrain during the day causes the vertical temperature

profile to become less stable and thus possibly tends to increase the PBL height over the Asir
 mountains.

3 Figure 10 shows the diurnal cycle of simulated vertical profiles of dust concentration, 4 potential temperature, wind vector, and PBL height in the altitude/longitude cross-section at 21.7° N on 19 March 2012. The aerosol concentration reaches 2000 μ g m⁻³. The thick dust 5 plume stretches from 28.7° E to 70.2° E. At 06:00 UTC, heating has just started, and the PBL 6 7 height nearly follows the topography. Over mountain and desert areas, the PBL is shallow. 8 The inland convection increases and enhanced subsidence over the Red Sea suppresses the 9 growth of PBL. The sea breeze starts developing at about 7:00 UTC. At 12:00 UTC, this 10 process continues but much more slowly as the sea breeze front moves inland and the PBL height grows farther to the east. A well-mixed, deep boundary layer in our simulations 11 develops during the day up to the height of 3.5 km in the afternoon as seen in Fig. 10c. At 12 night, the PBL collapses, resulting in an elevated dust-laden residual boundary layer above the 13 14 surface inversion (Fig. 10d).

15 The temporal PBL variability is characterized by the more rapid PBL growth over the high 16 elevations in the western Arabian Peninsula early in the heating cycle. The high PBL then spreads towards the eastern Peninsula. During the night, the nocturnal jets contribute to the 17 18 dust mobilization close to the surface, while aerosol particles increasingly sediment from the 19 residual layer after the collapse of the PBL. The peak of dust vertical mixing is observed at 20 about 12:00 UTC (Fig. 10c) over the Asir mountains, in southwestern Saudi Arabia, where a 21 significant amount of dust penetrates in the free troposphere downwind of the Asir mountain 22 ridge.

23 To quantify the amount of dust mixed above the boundary layer, we calculate the domain 24 average loading below and above the diurnal maximum PBL height (Fig. 11) that approximates the top of a residual layer. Aerosols remaining in the residual layer will be again 25 26 entrained in the PBL during the next diurnal cycle therefore we assume that only aerosols 27 above the residual layer are disconnected from the surface. Because the PBL dynamics is 28 different over land and over the sea, we did these calculations separately for the land and 29 ocean areas within the model domain. Figure 11 shows that during the simulation period, 30 about 15% of dust over land is in the free troposphere. Over the sea, it is often more than 31 50%. The dust storm increased the dust loading over land by about 30% and almost tripled it 32 over the sea.

To assess the simulated vertical structure of the atmosphere we used RS soundings from the 1 2 Jeddah's King Abdulaziz Airport (21.7° N, 39.18° E). The model and RS profiles of the water vapor mixing ratio and potential temperature (Fig. 12a) show that the PBL height on 19 3 March 2012 at 00:00 and 12:00 UTC is captured well by the model. The simulated PBL 4 5 height was about 3 km in the afternoon as seen in Fig. 10c, which is in a good agreement with the position of the inversion layer from Jeddah's soundings (Fig. 12a). The potential 6 7 temperature at 12:00 UTC (Fig. 12a) is almost constant between the ground and 3 km altitude, 8 as a result of strong mixing in the PBL. The RS vapor mixing ratio profile also shows the 9 mixing height to be around 3 km. The vertical profile of the magnitude and direction of the 10 horizontal wind over Jeddah from RS profiles on 19 March (00:00 and 12:00 UTC) is shown 11 in Fig. 12b. The wind direction shows a strong diurnal change in the boundary layer at about 1 12 km altitude. The wind over Jeddah was found to be east/south-east in the lowest 1 km and 13 mostly west/north-west above 3 km. The wind speed exceeds 15 m s⁻¹ near the surface and is more than 50 m s⁻¹ at 12 km. The surface wind at 00:00 UTC is enhanced by the land breeze. 14 This is a very important factor for the transport and deposition of dust over the Red Sea 15 16 during the night-time (Fig. 10a,d).

17 **3.4** The direct radiative effect of dust

Aerosols directly influence the Earth's radiative budget by absorbing and emitting long-wave 18 19 (LW) radiation (Cautenet et al., 1991; Markowicz et al., 2003; Ackerman and Chung, 1992; Haywood et al., 2005) and scattering and absorbing short-wave (SW) radiation (direct effect) 20 21 (Tegen et al., 1996; Haywood and Boucher, 2000; Harrison et al., 2001; Haywood et al., 22 2001; Sokolik et al., 2001), and indirectly by altering cloud droplet size distribution (indirect 23 effect) (Twomey, 1977; Albrecht, 1989). Because there is little cloudiness over deserts, the 24 direct radiative effect of dust is of primary interest in this study. The surface SW cooling by 25 dust results in large reductions in latent and sensible heat fluxes from the surface to the atmosphere (Miller et al., 2004b; Shell and Somerville, 2007). This reduces the turbulent 26 27 energy within the PBL and the downward transport of momentum to the surface, which suppresses the surface wind speed and dust generation (Miller et al., 2004a). In the 28 29 troposphere, the absorption of both incoming and reflected solar radiation, as well as emission of thermal radiation by dust particles affect the air temperature and circulation 30 31 (Alpert et al., 1998; Miller and Tegen, 1998; Sathesh et al., 2007). Depending on their 32 physical and optical properties as well as their chemical composition, aerosols exert a cooling

or warming influence on the atmosphere and underlying surface (e.g., Chylek and Wong,
1995; Sokolik and Toon, 1996; Miller and Tegen, 1998; Ahn et al., 2007; Shell and
Somerville, 2007; Balkanski et al., 2007). Large uncertainties in the assessment of the
radiative impact of mineral dust on regional and global scales have been pointed out in a
number of early and recent studies (Sokolik and Toon 1996; Tegen et al., 1996; Carlson and
Benjamin, 1980; Sokolik and Golitsyn, 1993; Ackerman, 1997; Liao and Seinfeld, 1998a;
Claquin et al., 1998; Stier at al., 2007)

8 To assess the direct radiative effect (DRE) of dust aerosols we calculate the change of total 9 clear-sky radiative flux ΔF , at the top of the atmosphere (TOA) and at the bottom of the 10 atmosphere (BOA), in a simulation, when aerosols are present (F^A) and in a simulation where 11 aerosols are absent (F^C). Then the DRE can be defined at BOA and TOA as follows

12
$$\Delta F_{BOA} = F_{BOA}^{\downarrow A} - F_{BOA}^{\uparrow A} - F_{BOA}^{\downarrow C} + F_{BOA}^{\uparrow C}$$
(4)

13
$$\Delta F_{TOA} = F_{TOA}^{\uparrow C} - F_{TOA}^{\uparrow A}$$
(5)

14 where the arrows indicate the direction of the fluxes: \downarrow denotes a downward flux and \uparrow an 15 upward flux. This convention is used for both LW and SW radiation. Definitions (Eqs.4 and 16 5) imply that negative values of ΔF at the BOA and TOA are associated with an aerosol 17 cooling effect, while positive ΔF is associated with warming. The DRE, as defined here, is 18 different from a conventional instantaneous direct radiative forcing (DRF) because it includes 19 model responses (see, e.g., Stenchikov et al., 1998). However, the clear-sky SW DRE, 20 discussed below, should be a fairly good approximation of the clear-sky SW DRF.

21 **3.4.1** The shortwave radiative effect

The spatial distributions of daily average clear sky SW, LW and net (SW+LW) dust DRE at 22 BOA and TOA on 19 March are shown in Fig. 13. The SW DRE is mostly negative at the 23 surface, reaching -134 W m^{-2} (Fig. 13a). The domain-averaged SW DRE at BOA equals to 24 -16 W m⁻² and tends to cool the surface. The daily averaged SW DRE at TOA over desert 25 areas is about 10 W m^{-2} and tends to warm the surface-atmosphere system over desert (Fig. 26 13b). However the domain-average SW DRE at TOA equals to -4.25 W m⁻². The warming 27 28 effect of dust aerosols over desert surfaces is largely explained by two factors: absorption of 29 solar radiation by dust and high reflectance of the underlying surface. Dust absorbs and 30 scatters the SW radiation. Both these effects reduce the downward flux at BOA but tend to

compensate for each other at TOA. Therefore, the BOA cooling effect is robust, but the TOA 1 2 warming is very sensitive to the dust's single scattering albedo and surface reflectivity. Over the ocean, the SW DRE at TOA is negative due to the smaller ocean albedo. Over the ocean, 3 4 the dust layer increases the outgoing shortwave radiation since it is a much better reflector 5 than is the underlying dark water. The negative values of SW DRE at TOA over the ocean are consistent with the results from Perez et al. (2006) on Mediterranean dust. Daily averaged SW 6 BOA aerosol effect up to -21.1 Wm⁻² was reported by McFarlane et al. (2009) during the 7 8 atmospheric radiation measurement mobile facility deployment in Niamey and Niger. During 9 the winter monsoon season, Satheesh et al. (2006) reported daily averaged surface SW radiative effect from -23.6 to 12 W m^{-2} over the Arabian Sea for 8 years from 1995 to 2002 10 with a maximum instantaneous value of -51 W m^{-2} . During the summer monsoon season, the 11 daily averaged surface SW radiative effect was in the range of -24.2 to -13.1 W m⁻² with a 12 "maximum" instantaneous value of -45 W m^{-2} . At TOA, the SW radiative effect of dust was 13 in the range of -16.5 to -9.1 W m⁻² during the summer monsoon and -7.3 to -6.0 W m⁻² 14 during the winter monsoon. Haywood et al. (2003) revealed an instantaneous direct SW 15 16 radiative effect at BOA from a large Saharan dust plume advected off the coast of western Africa to be approximately -130 W m^{-2} . The maximum instantaneous SW DRE at surface in 17 this study was -449 W m^{-2} on 19 March 2012. 18

19 **3.4.2** The longwave radiative effect

20 Figures 13c and d show that the daily averaged LW DRE is mostly positive at TOA and BOA and is larger at the surface. Dust absorbs terrestrial radiation coming from the surface and 21 22 from the atmosphere and generally reemits it at lower temperatures than the underlying 23 surface. The presence of dust reduces the outgoing LW radiation and increases the downward TOA LW flux. The maximum daily average LW DRE at the surface is 43 W m^{-2} with a 24 domain average value of 6 W m^{-2} . At TOA, the maximum daily average value of LW DRE is 25 8.4 W m⁻² with a domain average value of 0.25 W m⁻². The LW warming can offset the SW 26 cooling effect of dust at TOA. Zhang and Christopher (2003) estimated the dust aerosol LW 27 DRE over the Saharan desert from instruments onboard a satellite and showed that there is a 28 strong warming effect of 7 W m⁻² over the cloud-free Sahara desert regions for September 29 2000 and when averaged over six regions from the study area, the LW DRE ranges from -1 to 30 15 W m⁻². Liao and Seinfield (1998a) reported LW DRF ranging from 0.9 to 1.4 W m⁻² for 31 32 the dust layer using a one-dimensional column radiation model. From low dust loading under

dry tropics atmospheric conditions, Sokolik et al. (1998) calculated the LW DRF at BOA in the range of 7 to 14 W m⁻² and the LW DRF at TOA in the range of 2 to 7 W m⁻². We found that under dust storm conditions, the LW DRE at the surface vary between 50 and 80 W m⁻² and the LW DRE at TOA varies between 15 and 25 W m⁻². It is smaller but comparable with the SW DRE.

6 3.4.3 Net DRE

7 Figures 13e and f show the net (SW+LW) DRE at BOA and TOA, respectively. At the surface, the net DRE is negative in most of the region with an extreme value of -94 W m⁻². 8 Strong negative net DRE at BOA is seen over the region where the model and satellites show 9 10 the highest dust atmospheric loading (Fig. 8). In Figure 13 the domain average value of net DRE is -10 W m^{-2} and -4 W m^{-2} at BOA and TOA, respectively. The SW DRE exceeds the 11 LW DRE at both BOA and TOA. Figures 13e and f show that net DRE causes cooling at the 12 13 surface, while at the TOA it causes warming over the land but cooling over the oceans. This 14 pattern in the net DRE agrees with that found by Yue et al. (2010), Ackerman and Chung 15 (1992), Woodward (2001), Liao et al. (2004). The sign of the net DRE at TOA could be either positive (warming) or negative (cooling), depending on several key variables, such as the 16 17 surface albedo, particle size, vertical distribution of dust, its optical depth, and the imaginary 18 part of the dust refractive index (Tegen and Lacis, 1996; Liao and Seinfeld, 1998a,b). In this 19 study, the imaginary part of the dust refractive index is set to 0.003 for the entire SW 20 spectrum. Table 3 summarizes the domain average SW, LW, and Net DRE at TOA and BOA 21 for 19 March 2012.

22 **3.4.4** Atmospheric heating rates

The rate of change of temperature in an atmospheric layer (K day^{-1}) due to radiative 23 heating/cooling is called the radiative heating/cooling rate. The simulated instantaneous dust 24 25 induced perturbations (similar to definitions Eqs.(4) and (5)) of SW heating rates over Jeddah, 26 Abu Dhabi, and Kuwait at 12:00 UTC on 19 March are shown in Fig. 14. The available 27 temperature profiles at 12:00 UTC from RS are also shown. The heating rates vary with the altitude following the vertical distribution of the aerosols. The maximum heating rates over 28 Kuwait (~4.4 K day⁻¹) and Abu Dhabi (~4 K day⁻¹) are at about 1 km altitude. The heating 29 rates reach their maximum value at about 3 km over Jeddah (~ 4.9 K day⁻¹). Over Mezaira (not 30 shown), the dust-generated heating rates in the lower troposphere reach 9 K day⁻¹. The 31

maximum heating over Kuwait and Abu Dhabi is at lower levels compared with that over
Jeddah. This is because Kuwait and Abu Dhabi are closer to the dust source regions and hence
there is more dust in the lower levels of the atmosphere.

4 The RS temperature profiles help to clarify the effect of aerosols on the vertical structure of 5 the atmosphere under storm conditions. All RS temperature profiles show strong inversions 6 above the layers with maximum heating rates. These inversions are formed by absorption of 7 SW radiation at the top of a dense aerosol layer and further mixing above this layer. The 8 temperature inversion is quite strong, e.g., over Jeddah at about 2.5 km altitude, it reaches 4.4 9 K and significantly affects the atmospheric and dust dynamics by preserving a dense aerosol 10 layer near the surface. Reported heating rates from the Arabian dust storm are consistent with 11 the figures in the literature for other locations. Lemaître et al. (2010) reported average dust heating rates to be between 1.5 and 4 K day⁻¹, depending on the altitude and location of a dust 12 event observed over Benin and Niger, which is comparable with the maximum heating rate 13 over Kuwait (~4.5 K day⁻¹) in our study. The instantaneous heating rates they observed were 14 as high as 8 K day⁻¹ in some limited areas, which is comparable with the strong heating of 9 15 K day $^{-1}$ observed at Mezaira near Abu Dhabi. 16

The radiative heating/cooling by dust must be taken into account to predict the overall impact of aerosols on atmospheric circulation. Karyampudi and Carlson (1988) showed that radiative heating by Saharan dust helps to maintain a warm and deep Saharan air layer over the ocean, enhancing the strength of the mid-level easterly jets and reducing the convection within the equatorial zone. Chen et al. (1994) showed that the radiative heating rates of dust can affect the evolution of a dust storm, leading to stronger surface frontogenesis. They suggested that the heating rates can significantly affect mesoscale weather systems in arid and desert regions.

24 **3.5** Meteorological responses to the dust DRE

Figure 15 presents our simulated spatial distributions of instantaneous changes in surface air temperature and latent and sensible heat fluxes caused by the dust over land at 12:00 UTC on 19 March 2012. The surface air temperature over the majority of the study domain cools in response to the net dust radiative effect. The temperature drops by up to -6.7 K. A number of previous studies demonstrated similar magnitudes of the short-term impacts of dust on surface temperature (Han et al., 2013; Wang et al., 2010).

1 The change in the surface latent heat flux is quite similar to the prediction by Miller et al. 2 (2004b). The instantaneous latent heat flux decreases over the study region with a domain average value of -1 W m^{-2} in the presence of dust aerosols over land. Figure 15c shows the 3 4 decrease in the instantaneous sensible heat flux over land due to dust aerosols with a domain average value of -23.6 W m^{-2} . The daytime cooling at the surface leads to reductions in the 5 upward transport of heat from the surface to the atmosphere. The decreases in surface latent 6 7 heat and sensible heat fluxes caused by dust aerosols have been reported previously by Yue et 8 al. (2010), Miller et al. (2004b), Shell and Somerville (2007).

9

10 4 Summary and Conclusions

11 A severe dust storm on 18-22 March 2012 was simulated using WRF-Chem to account for 12 interactive dust generation and radiative effects. The model is able to simulate the major 13 spatial features of meteorological fields in the study domain reasonably well. The synoptic 14 systems that generated these dust events were associated with a fast-moving cold front accompanied by a high level trough. Strong wind (velocities exceeding 15 m s^{-1}) occurred 15 16 behind the cold front, which entrained large quantities of dust particles into the atmosphere. 17 The meteorological conditions that led to the strong winds capable of producing the severe 18 dust event have been identified. There were several rich dust source areas in the region, 19 activated by the passage of the front, where the major plumes of dust storm originated. The 20 simulation suggests that the main dust sources during the dust storm include river valleys of the lower Tigris and Euphrates in Iraq, areas from Kuwait, Iran, UAE and within the basin of 21 22 the Arabian desert (which includes the Rub' al Khali, An Nafud and Ad Dahna). Dust sources 23 were also identified along the western coast of the Arabian Peninsula. Our simulation shows 24 that around 93.79 Mt of dust were emitted into the atmosphere during the simulation period. 25 About 18% (73.04 Mt) of the dust were deposited back to the study domain and around 22% (i.e., 20.72 Mt) were transported outside the model domain. About 6.56 Mt of dust was 26 27 deposited to the oceans within the model domain. The Red Sea received about 1.2 Mt of dust 28 during this event. Dust particles bring nutrients to marine ecosystems; by scaling the effect of 29 one storm to the number of dust storms observed annually over the Red Sea, we roughly 30 estimate the annual dust deposition, associated with dust storms, to be 6 Mt.

The variation in terrain elevation and surface properties produces large variability in the PBL depth over the Arabian Peninsula. The deepest PBLs developed over the higher elevations of the western Arabian Peninsula and over desert areas. The model predicted a well-mixed, deep boundary layer up to a height about 3.5 km in the afternoon. Over land, about 15% of dust were entrained in the free troposphere, while over the seas, it might be more than 50%. This is extremely important for understanding of dynamics and long-range transport of dust plumes.

6 The modeled AOD compares well with AERONET at the three sites and with satellite data. 7 The simulation shows that mineral dust heats the lower atmosphere (1-3 km) by SW absorption with a maximum rate up to around 9 K dav⁻¹ over Abu Dhabi at 12:00 UTC on 19 8 9 March. The lower atmosphere heating is also reflected by the temperature inversions observed 10 in the RS temperature profiles. The meteorological responses to the dust radiative effects 11 indicate a cooling at the surface due to the dust aerosols over the high dust concentration regions. The daily domain average values of net dust DRE at the surface and TOA under clear 12 sky conditions at the time of the dust storm on 19 March are -10 W m^{-2} and -4 W m^{-2} , 13 respectively. The shortwave direct radiative effect is negative at the surface over the domain 14 with a maximum daily average value of -134 W m⁻² and a domain-averaged value of -1615 W m⁻². The daily average shortwave radiative effect at TOA is positive (~10 W m⁻²) over 16 high albedo deserts and negative over oceans due to the smaller surface albedo. The maximum 17 daily averaged LW DRE at the surface is 43 W m^{-2} with a domain average value of 6 W m^{-2} . 18 At TOA, the maximum daily average value of LW DRE is 8.4 W m⁻² with a domain-19 averaged value of 0.25 W m⁻². At the surface, the SW DRE always leads to cooling and LW 20 21 DRE to heating. The dust radiative effects cause decreases in the surface air temperature with 22 maximum up to -6.7 K and a domain-averaged value of -0.26 K at 12:00 UTC on 19 March. 23 The change in the surface air temperature in the model is a combined effect of SW and LW DRE. The indirect effect of mineral dust is not implemented in the model for the present 24 25 study. The decrease in domain average values of latent and sensible heat flux are -1 and -23.6 W m^{-2} , respectively. 26

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- 27

Table 1. Model setup parameters.

Model Name	WRF-Chem (Grell et al.,2005)
Horizontal Resolution	10 Km
Dimensions (X, Y)	495, 396
Vertical Levels	40
Boundary conditions	NCEP Final Analysis FNL
Time Step	60 s
Dust-radiative feedback	Yes
Simulation period	11–24 March 2012
Physical schemes	Microphysics: Lin et al. Scheme (Lin et al.,1983) PBL: Mellor-Yamada-Janjic (MYJ) TKE scheme (Janjic, 2001) Surface: unified Noah land-surface model (Chen and Dudhia, 2001) Cumulus convection: New Grell Scheme (Grell and Devvenyi, 2002)
Emission scheme	GOCART Scheme (Ginoux et al., 2001)
Aerosol Model	MADE/SORGAM (Ackerman et al., 1998; Schell et al., 2001)
Photolysis scheme	Fast-J (Wild et al., 2000)
Gas-phase mechanism	RADM2 (Stockwell et al.,1990)
Dust SW Refractive Index	0.003

Table 2. Total emission and deposition over the simulated domain from 11 to 24 March 2012.

Dust Deposition/Emission	Dust Mass
	(Mt)
Total Dust Emission	93.76
Total Dust Deposition	73.04
Total Ocean Deposition	6.56
Red Sea Deposition	1.20
Arabian Sea Deposition	3.01
Land Deposition	66.48

Table3. Summary of domain average values of daily mean clear sky direct radiative effect
 (DRE) for shortwave (SW), longwave (LW) and net (SW+LW) radiation at top of the
 atmosphere (TOA) and bottom of atmosphere (BOA) on 19 March 2012.

Level	SW DRE (W m^{-2})	LW DRE (W m^{-2})	Net DRE (W m^{-2})
TOA	-4.25	0.25	-4
BOA	-16	6	-10

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Figure 1. Geopotential height (m) at 500 hPa at 00:00 UTC from WRF-Chem simulations on
(a) 17 March, (b) 18 March and (c) 19 March 2012 and from ERA-Interim reanalysis on (d)

- 4 17 March, (e) 18 March and (f) 19 March 2012.
- 5





Figure 2. Sea level pressure (hPa) with H and L markers indicating low and high pressure
systems at 00:00 UTC from WRF-Chem simulations on (a) 17 March, (b) 18 March and (c)
19 March 2012 and from ERA-I reanalysis on (d) 17 March, (e) 18 March and (f) 19 March
2012.







Figure 3. Simulated spatial distribution of 10m wind bars and wind magnitude (m s⁻¹) at
09:00 UTC on (a) 17 March, (b) 18 March and (c) 19 March 2012.





Figure 4. Simulated 10m wind bars (m s⁻¹), dust concentration at 900 hPa (μ g m⁻³), and 2m air temperature (K) at (a) 08:00 UTC and (b) 15:00 UTC on 18 March 2012.



Figure 5. (a) Dimensionless dust erodibility field (S) from the GOCART emission scheme used in simulations; (b) simulated daily mean dust emission rate ($\mu g m^{-2} s^{-1}$) on 17 March 2012; (c) Same as (b), but for 18 March 2002; (d) Same as (b) but for 19 March 2002.



2 Figure 6. Simulated daily mean dust deposition rate ($\mu g m^{-2} s^{-1}$) on (a) 17 March, (b) 18

3 March, and (c) 19 March 2012.



Figure 7. Simulated daily mean domain integrated (a) dust emission (black) and deposition
(red) (Mt day⁻¹) and (b) dust deposition (Mt day⁻¹) over all ocean areas within the domain
(black), over the Arabian Sea (red), and over the Red Sea (blue) as a function of time.



3 Figure 8. (a) Spatial distribution of simulated dust load (g m^{-2}) at 10:00 UTC on 19 March

4 2012, (**b**) combined image taken by Aqua/MODIS at 09:50 UTC and 10:05 UTC on 19 March

- 5 2012 showing dust storm passing over the Arabian Peninsula and Red Sea.
- 6





Figure 9. Spatial distribution of AOD on 19 March 2012 from (a) WRF-Chem simulations at
10:00 UTC, (b) MODIS (combined standard ocean and deep blue products) at 09:50–10:05
UTC, (c) SIVIRI at 10:00 UTC, and 0.60 µm AOD from the WRF-Chem simulations (black
solid line) at AERONET sites and AERONET observations (red dots) at (d) KAUST Campus,
(e) Kuwait University, (f) Mezaira.



Figure 10. The height-Longitude cross-section along 21.7° N showing dust concentration (μg m⁻³, shading) and PBL height (km) shown by the black solid line from WRF-Chem on 19

- 4 March 2012 at (a) 00:00 UTC, (b) 06:00 UTC, (c) 12:00 UTC, (d) 18:00 UTC.
- 5



Figure 11. Domain-averaged dust loading (g m⁻²) above (solid line) and below (dash line)
daily maximum PBL heights for land (red) and ocean (blue) areas as a function of time for the
entire period of simulations.



Figure 12. (a) Simulated (dashed lines) and RS profiles of potential temperature and water vapor mixing ratio (g kg⁻¹) at the Jeddah's King Abdulaziz International Airport at 00:00 and 12:00 UTC on 19 March 2012; (b) Vertical profiles of wind direction (degrees) (blue and green) and wind speed (m s⁻¹) (red and black) from radiosonde soundings at the Jeddah's King Abdulaziz International Airport at 00:00 and 12:00 UTC on 19 March 2012.



Figure 13. Simulated daily mean clear sky direct radiative effect (DRE) of dust aerosols (W m⁻²) on 19 March 2012; (a) SW DRE at BOA, (b) SW DRE at TOA, (c) LW DRE at BOA;
(d) LW DRE at TOA; (e) net DRE at BOA (f) net DRE at TOA. Positive values correspond to the heating of the atmospheric column and underlying surface.



Figure 14. Simulated (dashed lines) solar heating rates profiles in the atmosphere over Jeddah
(black), Abu Dhabi (red) and Kuwait (green) sites and radiosonding temperature profiles at
the same locations (solid lines) at 12:00 UTC on 19 March 2012.



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Figure 15. Simulated instantaneous differences in (a) air temperature (K), (b) latent heat flux (W m⁻²), and (c) sensible heat flux (W m⁻²) in the runs with and without dust aerosols at

4 12:00 UTC on 19 March 2012.