



**Determination and climatology of the planetary boundary layer height**

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# Determination and climatology of the planetary boundary layer height by in-situ and remote sensing methods as well as the COSMO model above the Swiss plateau

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## Abstract

The planetary boundary layer (PBL) height is a key parameter in air quality control and pollutant dispersion. The PBL height can however not be directly measured and its estimation relies on the analysis of the vertical profiles of the temperature, the turbulences or the atmospheric composition. An operational PBL height detection including several remote sensing instruments (windprofiler, Raman lidar, microwave radiometer) and several algorithms (Parcel and bulk Richardson number methods, surface-based temperature inversion, aerosol or humidity gradient analysis) were developed and the first year of application allowed validating these various detection methods against radio sounding measurements. The microwave radiometer provides convective boundary layer heights in good agreement with the radio sounding (median bias < 25 m,  $R^2 > 0.70$ ) and allows to fully analyzing the PBL height diurnal cycle due to its smaller time granularity. The Raman lidar also leads to good results whereas the windprofiler yields some more dispersed results. Comparisons with the numerical weather prediction model COSMO-2 were also established and point out a general overestimation by the model. Finally the seasonal cycles of the daytime and nighttime PBL heights are discussed for each instrument and each detection algorithm for two stations on the Swiss plateau.

## 1 Introduction

The height of the planetary boundary layer (PBL), also called atmospheric boundary layer is a key parameter for air quality analysis, pollutants dispersion and quantification of pollutant emissions and sources. It controls the interactions of the atmosphere with the oceans and the land and determines the air volume available for the dispersion of all atmospheric constituents (including anthropogenic pollution and water vapor) emitted at the earth surface, and hence contributes to the assessment of the pollutant concentration near the surface. The PBL height is therefore a key parameter of all air pollution

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models predictions (Baars et al., 2008; Seidel et al., 2012; Ketterer et al., 2013), the results depending on both the model and the measurement types.

Some of these studies were done on long enough time series (between 1 and 25 years) to analyze the PBL height climatology at some stations in Europe and US (Baars et al., 2008; Schmid and Niyogy, 2012; Beyrich and Leps, 2012; Granados-Muñoz et al., 2012; Sawyer and Li, 2013) or over continents (Seidel et al., 2010, 2012). For continental stations, a clear CBL seasonal cycle is usually found with summer maxima reaching 1000 to 2000 m above ground level (a.g.l.) and winter minima between 500 and 1200 m a.g.l. The seasonal cycle of the nocturnal SBL was to our knowledge only addressed on the basis of temperature ( $T$ ) profiles from RS measurements (Seidel et al., 2010; Beyrich and Leps, 2012; Seidel et al., 2012). Both authors found a minimum in summer and a maximum in winter explained by greater wind speeds and consequently stronger mechanical turbulences during winter. Few of the PBL height detections run operationally, meaning that they runs as fully automatic processes delivering continuous PBL estimation in real time. Some authors specify that a visual inspection is necessary to increase the results reliability.

In this study an operational system for PBL height detection has been developed based on the analysis of vertical atmospheric profiles of  $T$ , wind turbulence and atmospheric constituents, which are measured by different remote sensing techniques like radar windprofiler (WP), microwave radiometer (MWR) and lidar as well as by radiosounding (RS). One year (2012) of measurements was used to compare all these PBL height determination methods against a reference method being the PM applied either on RS or on MWR measurements. The PBL heights computed by the COSMO-2 model (numerical weather prediction model of the Consortium for Small Scale Modelling; see [www.cosmo-model.org](http://www.cosmo-model.org)) were also compared to the instrumental PBL height determination. A two-year climatology of the CBL, the cloudy-CBL and the different layers constituting the nocturnal SBL was moreover computed for Payerne (PAY) and Schaffhausen (SHA) situated both on the Swiss plateau.

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resolution remains constant (30 m). The effective time resolution of profiles is 30 min. No measurements are possible during precipitation and in presence of low clouds, i.e. the lidar powers down if the clouds are below 900 m and powers up as soon as the cloud base rises above 2000 m.

5 A ceilometer (CBME80 from Eliasson) measuring at  $\lambda = 905$  nm with a time resolution of a few seconds is interfaced to the lidar system to provide independent cloud information. This model was not configured to record backscatter profiles but only to provide the height of the cloud bases detected by a strong gradient in the backscattered signal.

10 In addition to the remote sensing instruments, the station of Payerne performs RS providing pressure ( $p$ ),  $T$ , humidity and wind speed and direction profiles up to 35 km. Meteolabor SRS 400 C34 radiosondes are launched twice a day at 00:00 and 12:00 LT. The horizontal displacement of the sonde can reach up to 200 km. However, only the first vertical 3500 m corresponding to approximately 12 min of rise are used to determine the PBL height allowing neglecting the RS horizontal displacement. RS has a constant height resolution of 5–6 m corresponding to a one second time resolution.

15 The SwissMetNet meteorological surface network provides surface  $T$ , humidity,  $\rho$ , wind direction and speed as well as sunshine duration and precipitations every 10 min. The wind components are measured at 10 m and all the other parameters at 2 m. In addition, the PAY station is equipped with a sonic anemometer on a 10 m mast measuring several parameters related to turbulence, including sensible heat flux that characterizes the thermal energy exchanged and is used to estimate the intensity of the convective forces.

20 The COSMO-2 model (<http://www.cosmo-model.org>) was used in assimilation mode. It has a horizontal grid spacing of 2.2 km and a total of 60 vertical levels, of which 15 lie within the first 500 m. The time step is 20 s and data are written out every 1 h. The bulk Richardson number method is used to estimate the boundary layer height in the model (see Sect. 2.2.1).

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The cloud cover is detected by Automatic Partial Cloud Amount Detection Algorithm (APCADA) that estimates in real-time the sky cloud cover from surface based measurements of long-wave downward radiation,  $T$  and humidity (Dürr and Philipona, 2004). APCADA does not take into account the cirrus clouds.

Measurements of both the MWR and the lidar are necessary to calculate the virtual potential temperature ( $\theta_v$ ), and they are combined with WP data to calculate the bulk Richardson number (see Sect. 2.2.1). These three instruments have however different vertical levels and time constants. For these cases, a vertical scale (35 levels of 100 m between 0 and 3500 m) is set and the mean of the parameters in each level are used. Despite the rather long integration times in the case of the windprofiler and the lidar all observational data have been assumed to be instantaneous. These different time granularities are sometimes visible by a time shift of the CBL growth measured by MWR/PM and WP/SNR or lidar/ASR.

## 2.2 Methods to determine PBL height

### 2.2.1 Methods based on $T$ profiles

The Parcel method (Holzworth, 1964; Fisher et al., 1998) defines the PBL height as the elevation to which an air parcel with ambient surface  $T$  can rise adiabatically from the ground by convection. As depicted in Fig. 2, the PBL height is set to the elevation  $z$  where the  $T$  profile crosses the dry adiabatic, or where the potential temperature  $\theta$  is equal to the surface  $\theta$ . The PM needs only the  $T$  profile and a precise surface  $T$  measurement. To apply the PM, the condition  $\theta(z_1) < \theta(z)$ , with  $z_1 > z$ , corresponding to unstable  $\theta$  vertical profile, has to be fulfilled. No excess  $T$  has been added to the surface  $T$ . The PM was applied to RS and MWR  $\theta$  profiles to detect daytime PBL in case of strong or weak convective conditions (CBL and cloudy-CBL).

The bulk Richardson number ( $Ri_b$ ) is a dimensionless quantity combining the potential energy and the vertical wind shear. It corresponds to the ratio of convective and wind shear produced turbulences and is widely used in turbulence characterization. In order

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to be consistent with the  $Ri_b$  used in the COSMO-2 model (Szintai, 2010), the following formulation was applied:

$$Ri_b = \frac{gz(\theta(z) - \theta(z_0))}{\bar{\theta}(U^2(z) + V^2(z))} \quad (1)$$

where  $z$  is the height ( $z > z_0$ ),  $U$  and  $V$  the two horizontal wind velocity components,  $g$  the Earth gravitational constant and  $\bar{\theta}$  the mean  $\theta$  between  $z_0$  and  $z$ . The PBL height corresponds to the first elevation  $z$  with  $Ri_b$  greater than a critical threshold taken as 0.22 or 0.33 in case of unstable (day) or stable (night) conditions, respectively (Fisher et al., 1998; Jericevic and Grisogono, 2006; Szintai, 2010). It has to be noted that, in most cases, the exact threshold value has only very little impact on the PBL height due to the great  $Ri_b$  slope in this interval (see for ex. Figure 2b). PBL height detected by bR is by definition higher than PBL height detected by PM, because both methods are similar if the threshold value is set to 0 involving  $\theta(z) = \theta(z_0)$ . The WP wind velocities were used to calculate the  $Ri_b$  from the MWR  $T$  profile. The bR method was applied on daily RS and MWR  $\theta$  and COSMO-2  $\theta_v$  profiles for CBL, cloudy-CBL and SBL detection.

For both PM and bR methods the surface  $T$  has a large impact on the determined PBL height and hence it is crucial to take a representative measurement that is not biased by micrometeorological effects. The surface  $T$  was therefore taken from the meteorological surface network at 2 m. If needed, a linear interpolation between two measured  $\theta$  is applied to determine the PBL height. Uncertainties in PBL height for both methods were calculated by varying the surface  $T$  by  $\pm 0.5^\circ$  and were found to be in the order of  $\pm 50$ – $150$  m around the PBL maximal height reached in the early afternoon. Far larger uncertainties were found for the PBL height decrease in the late afternoon. For RS,  $\theta$  and  $\theta_v$  were calculated using the  $\rho$  and RH provided by RS measurements, whereas for MWR RH was provided by the lidar and  $\rho$  was calculated from the MWR

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$T$  profile only by using the ideal gas equation and the hydrostatic equilibrium:

$$\rho(z, T) = \rho_o \cdot \exp\left(-\int_{z_0}^z \frac{M_a g}{RT(z')} dz'\right) \quad (2)$$

where  $M_a$  is the mass of the air,  $R$  the specific gas constant and  $\rho$  is measured at 2 m from the meteorological surface network.

The nocturnal SBL can only be detected by the  $T$  profiles measured by RS and MWR, since wind turbulence, aerosol and relative humidity (RH) profiles retrieve the RL height during the night. The SBI is defined as the height of the surface-based  $T$  inversion, where  $T$  first decreases with elevation ( $dT/dz = 0$ ) as depicted in Fig. 3a (Bradley et al., 1993; Stull, 1988). A surface-based  $T$  inversion is a clear indicator of a stable boundary layer that can be defined as a SBL height (Seidel et al., 2010). The SBL top can also be defined as the transition between the stable surface layer and the neutral residual layer (Stull, 1988). This height is detected by a vanishing  $\theta$  gradient ( $d\theta/dz = 0$ ), which will be called SBLpT (Fig. 3b and c). SBLpT is per definition higher than SBI since the  $\theta$  gradient is still positive at the height of the surface-based  $T$  inversion and does not correspond to the top of the stable layer.

## 2.2.2 Method based on wind turbulence profiles

The radar echo measured by the WP is generated by inhomogeneities in the refractive index, which are characterized by the structure constant  $C_n^2$ . It can be shown that the range corrected SNR is proportional to  $C_n^2$ , which has a maximum at the top of the capping inversion, which marks the PBL top (White et al., 1991; Angevine et al., 1994, and references therein). Therefore a peak in the SNR profile can be associated to the PBL height under convective conditions. However, turbulence as well as humidity and  $T$  gradients associated with clouds and other dynamical processes can generate high SNR values, which do not correspond to the PBL height, leading to an attribution problem. To get rid of part of the false PBL height attribution, a time continuity algorithm was

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applied: each SNR peak with local maximum greater than 75 % of the absolute maximum was weighted by a Gaussian function with mean equals to the PBL height of the former time step and a standard deviation  $\sigma$  depending on the hour of the day. The PBL height is then attributed to the maximum of the weighted SNR peak. The uncertainty of this method is considered equal to the full width at half maximum (FWHM) of the selected SNR peak after subtraction of the noise floor and is in the order of 100–500 m. CBL starting height at sunrise was set to ground height with a large  $\sigma$  attributed to the first hours after sunrise. Similar algorithms taking into account the SNR slope and curvature were tested but have shown a lower consistency with respect to the other PBL height detection methods and a higher rate of false detections. This WP/SNR method was used to detect the CBL during the day and the RL during the night, but cannot be used in case of precipitation.

### 2.2.3 Method based on concentration profiles

The aerosol scattering ratio (ASR) is the ratio between the total and the molecular backscatter coefficients. Since the PBL top is characterized by a sharp decrease in concentration of all pollutants, the absolute minima in the vertical gradient of the lidar/ASR and of the RS/RH profiles can be associated to the CBL height during day and to the RL during night. A continuity algorithm similar to the WP/SNR method (see Sect. 2.2.2) was applied, with the modified condition that the local minimum has to be lower than 10 % of the absolute minimum. According to the WP/SNR method the uncertainty is considered equal to the FWHM of the selected in the lidar/ASR gradient profile and is in the order of 100–250 m.



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peaks both at the same altitude of 500 m until 03:00 and decreases to about 200 m thereafter, (3) the SBL detected by the COSMO-2 bR (orange diamonds) stays constant at 250 m until sunrise, (4) the RL is detected by both the WP/SNR (light blue circles) and the lidar/ASR (green circles) at 1500 m, the WP catching another turbulent layer at 700–800 m between 03:00 and 09:00 corresponding to a jet of north-east wind ( $15 \text{ m s}^{-1}$ , not shown). These two layers measured by WP/SNR and lidar/ASR before sunrise are finally merged into the developing CBL at 07:00 and 09:00, respectively.

- The CBL development from sunrise to mid afternoon: (1) one hour after sunrise, the CBL height increase is very well caught by all the methods based on  $T$  profiles, MWR/bR and COSMO-2/bR showing a quicker CBL increase and a higher CBL height between sunrise and 09:00 than MWR/PM. This difference between the PM and bR methods is due to the horizontal wind component that is taken into account in the bR method. In this case, the air moisture seems to have minor influence since the bR method leads to similar results if applied on the  $\theta_v$  (COSMO-2) or on  $\theta$  profiles (MWR). (2) The CBL remains then constant from 12:00 to 15:00, when the temporal gradient of the ground  $T$  vanishes before becoming negative (see the lower panel of Fig. 4). This CBL height maximum is consistently measured by all methods. (3) The CBL decrease after 15:00 is also well depicted by the methods based on  $T$  profiles (MWR/bR and MWR/PM), whereas the RL is thereafter measured by the WP/SNR and the lidar/ASR. The PM method, which is devised for CBL detection, becomes non applicable as soon as the vertical sensible heat flux becomes negative (see the red curve in the lower panel of Fig. 4), generating a positive or vanishing gradient of  $\theta$ .
- The nocturnal SBL development: after 18:00, the bR method continues to follow the CBL decrease whereas the development of the nocturnal SBL can be detected by the MWR/SBI and MWR/SBLpT methods.



bR definition, its PBL heights are higher than the ones computed by PM, the median bias remaining however very low at about 20 m.

- The MWR results are somewhat more scattered, but with very small median bias (< 25 m) and interquartile ranges (100 m). The MWR/PM has the smallest interquartile ranges and whiskers size due to the same applied detection method that, contrary to bR, do not use the WP wind velocity.
- The WP/SNR method has the lowest correlation coefficients (0.47), the largest median bias (–63 m) and the largest interquartile range (–560 to 460) of all the experimental methods. It also contains several large positive outliers that may be explained by the detection of elevated cloud layers falsely attributed to PBL height.
- The comparison with lidar/ASR can be only done on a reduced dataset (61 cases) due to its lower data availability. Taking into account the very different detection methods based on  $T$  and aerosol profiles, the comparison with RS/PM is very good with a slope of 1.00, correlation coefficients of 0.81 and a median bias of –5 m.

Since the CBL may not always be at its maxima at 12:00, an inter-comparison on the same set of 119 convective days was performed with MWR/PM as reference for the 12:00–15:00 time interval corresponding to CBL height maxima for all seasons (Fig. 7). Similarly to the 12:00 case, the difference between PM and bR is rather small with interquartile ranges of 5 and 71 m and whiskers far below 200 m. The lidar/ASR also shows a very good agreement with a median bias of 20 m and an interquartile range of about  $\pm 150$  m. Finally the great number of false detections of the WP due to either cloud, high humidity layers or turbulences are visible in the WP/SNR larger median bias (71 m) and interquartile range of about 200 m.

Each of the considered method has its own uncertainties in the PBL height detection as explained in Sect. 2.2. The uncertainty minimum is usually obtained for fully developed CBL that is the easiest case to detect. These uncertainties provide a similar

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picture as the inter-comparison, with a greatest precision for methods based on  $T$  profile and the lowest one for WP/SNR.

Finally, considering all these statistical differences found between the various instruments and methods as well as their related uncertainties, one has however to remain conscious that the measured parameter (PBL height) is in reality not a fixed point but rather a transition layer between 2 states of the atmosphere, which thickness reaches probably several tens of meters, and that the remote sensing instruments measure an air volume and not a precise point. The results of the inter-comparison and the instrumental uncertainties are of course greater than the thickness of this transition layer, but they stay however in the same order of magnitude.

### 3.4 Comparison between PBL height measured and computed by COSMO-2

Table 3 and Figs. 6 and 7 show that the PBL height given by COSMO-2 model has a positive bias compared to experimentally determined PBL heights. The median biases are of 275 m and 299 m when compared to the RS/PM (12:00) and to the MWR/PM (12:00 to 15:00), respectively. The interquartile ranges reach 200 to 350 m and the maximal whiskers are higher than 1000 m. A detailed analysis of the individual plots (see Fig. 8 for example) reveals that COSMO-2 often overestimates the PBL height during the whole day and tends to show a too rapid PBL growth in the morning. This behavior is not limited to clear-sky convective days and is observed throughout the year. This significant positive bias compared to all experimental methods and the asymmetry of the distribution, which is obvious on the histograms of Figs 6 and 7, may be explained by several reasons:

- Contrary to all the experimental methods, COSMO-2 determines the PBL height from  $\theta_v$  profile, leading to a physically meaningful systematic positive bias. This bias of 3–8% (see Sect. 3.1) cannot however explain the large discrepancy with the experimental methods.

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- The use of the bR method also induces a positive bias compared to the PM method, but the difference does not exceed some tens of meters as demonstrated by the RS and MWR results.
- The bR method is very sensitive to the surface  $T$  and an overestimation of this parameter can induce a systematic positive bias of PBL height. Errors and uncertainties in both  $T$  and RH profiles of COSMO-2 could also explain the large observed bias.
- The occurrence of clouds, which may be missing in the model, can reduce for a while the surface heating and the convection of air masses leading to a lower measured PBL height. This phenomenon is clearly visible in some cases (not shown).

Further studies are necessary to assess the impact of these various parameters and determine the main causes of the PBL height overestimation by COSMO-2.

### 3.5 PBL height two-year climatology at PAY and SHA

#### 3.5.1 CBL climatology

The two years climatology of CBL heights calculated from all instruments and COSMO-2 is presented in Fig. 9 for PAY (256 days) and SHA (289 days). It has to be noted that the same subset of days was taken for the MWR, the WP and COSMO-2, whereas the lower availability of lidar/ASR data leads to a smaller dataset that still allows the comparison with the CBL heights estimated from the other instruments. The CBL heights have a clear annual cycle with a minima at 300–700 m in winter and a maxima at 1200–1500 m during the Mai–August period. It has to be noted that the CBL extremes occur at the solstices and not at the  $T$  extremes (January–February and July–August), allowing to conclude that the solar radiation would be a better climatic variable to predict CBL cycle than the  $T$ .



higher than for the CBL case. The WP/SNR cloudy-CBL heights are most of the time more than 500 m higher than MWR/PM ones and measures probably the cloud top in a number of cases (see for example Fig. 5). Despite the low amount of available data, the lidar/ASR results are very similar to MWR/PM PBL heights during summer and somewhat higher in winter, similarly to the CBL case.

### 3.5.3 SBL climatology

The SBL climatology was divided into clear-sky (Fig. 11) and cloudy nights (Fig. 12) in order to differentiate cases with large and small radiative cooling. Clear-sky (186 at PAY and 163 at SHA) and cloudy nights (126 at PAY and 151 at SHA) were selected as days without precipitation between 00:00 and 5:00 and with 0–2 and 7–8 octa of the sky covered by clouds estimated by APCADA, respectively. While some features of the SBL annual behavior can be deduced, the low number of cases for some months, particularly for cloudy conditions, does not allow us to draw strict conclusions on the effective seasonal cycle of the different layers forming the SBL. The following points can however be observed:

- During clear-sky nights, the complete SBL structure can be clearly observed at PAY with SBL heights being between 100 and 500 m during the whole year, SBLpT being lower than 500 m in winter and rising up to 800 m during the other seasons. The RL measured by the lidar/ASR has a seasonal cycle completely similar to the CBL one (see Fig. 9), so that the pollutant emitted during the preceding days remain at the altitude of the CBL maxima during the night.
- During clear-sky nights, the WP/SNR method, which is more frequently subjected to false attribution than the other methods, leads to much more scattered results and large quartiles. The WP/SNR results are however comparable to the RL heights measured by the lidar/ASR.

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nighttime RL heights detected by lidar/ASR between 500 and 1500 m in Neuchâtel (Switzerland) similarly to our results. Finally, Beyrich and Leps (2012) and Seidel et al. (2010) studied the 10 year climatology of PBL height detected by RS measurements (twice a day). The SBL seasonal cycles over Europe were found to depend on the method applied to the RS profiles: the PM method leads to almost constant SBL during the whole year, whereas SBI has a seasonal minima in summer and a maxima in winter. Unfortunately our two-year dataset restricted by the cloud coverage is not large enough to compare our SBL seasonal cycles with these results. Finally, similarly to our results, the gradient method applied to the RH or specific humidity profiles is maximal during summer and minimal during winter. As expected, they also found that SBI yields the smallest heights, followed by the PM method, while the humidity and the ASR profiles similarly lead to much greater heights, corresponding to RL top.

#### 4 Conclusion: strengths and limitations of an operational mode

A system for automatic real time detection of the PBL height based on several methods applied to various remote sensing observations was implemented and operated for two years (2012–2013) for two upper air stations on the Swiss plateau. The numerical weather prediction model COSMO-2 PBL height was also compared to instrumental results. All the remote sensing and model results were validated on a subset of 119 convective days, the RS/PM at 12:00 or the MWR/PM between 12:00 and 15:00 being taken as references. A two years climatology for daytime and nighttime PBL heights were calculated for convective days and clear-sky nights, as well as for non-convective days and for cloudy nights without precipitation. The system for automatic detection of the PBL height is now implemented in an operational environment and the data are visualized and provided to end users in real time.

The difficulty of the PBL height detection comes first from the complexity of the troposphere itself, which can be composed of several layers with different thermal structure, wind regimes and concentrations of atmospheric constituents. Secondly, each







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instruments and seems to follow the solar radiation cycle rather than the  $T$  cycle. In case of partial or total cloudy conditions, a similar annual cycle but with lower PBL heights is measured, the WP results being however largely influenced by wind turbulence at the cloud top. The nocturnal PBL structure can be clearly observed under clear-sky conditions, with the SBI height remaining rather constant during the year at 200–300 m, the top of the stable layer lying at 800 m for most of the non-winter months and finally the RL nocturnal seasonal cycle following the CBL diurnal maximal. In case of total cloud coverage, the SBI height is lower than in case of clear-sky, and the SBL layers seems to be compressed and not well-structured under the cloud base. Further meteorological phenomena such as fog, neutral boundary layer height, main pollutant advections or nocturnal jets will be further addressed either as case studies or statistically after a longer measurement period.

## References

- Angevine, W. M., White, A. B., and Avery, S. K.: Boundary-layer depth and entrainment one characterization with a boundary layer profiler. *Bound.-Lay. Meteorol.*, 68, 375–385, 1994.
- Baars, H., Ansmann, A., Engelmann, R., and Althausen, D.: Continuous monitoring of the boundary-layer top with lidar, *Atmos. Chem. Phys.*, 8, 7281–7296, doi:10.5194/acp-8-7281-2008, 2008.
- Beyrich, F. and Leps, J.-P.: An operational mixing height data set from routine radiosoundings at Lindenberg: methodology, *Meteorol. Z.*, 21, 337–348, 2012.
- Bianco, L. and Wilczac, J. M.: Convective boundary layer depth: improved measurement by doppler radar wind profiler using fuzzy logic methods, *J. Atmos. Ocean. Technol.*, 19, 1745–1758, 2002.
- Bradley, R. S., Keimig, F. T., and Diaz, H. F.: Recent changes in the North American Arctic boundary layer in winter, *J. Geophys. Res.*, 98, 8851–8858, doi:10.1029/93JD00311, 1993.
- Calpini, B., Ruffieux, D., Bettems, J.-M., Hug, C., Huguenin, P., Isaak, H.-P., Kaufmann, P., Maier, O., and Steiner, P.: Ground-based remote sensing profiling and numerical weather prediction model to manage nuclear power plants meteorological surveillance in Switzerland, *Atmos. Meas. Tech.*, 4, 1617–1625, doi:10.5194/amt-4-1617-2011, 2011.

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Cimini, D., De Angelis, F., Dupont, J.-C., Pal, S., and Haeffelin, M.: Mixing layer height retrievals by multichannel microwave radiometer observations, *Atmos. Meas. Tech. Discuss.*, 6, 4971–4998, doi:10.5194/amtd-6-4971-2013, 2013.

Cost Action 710 – Final report “Harmonisation of the pre-Processing of meteorological data for atmospheric dispersion models”, edited by: Fisher, B. E. A., Erbrink, J. J., Finardi, S., Jeannet, P., Joffre, S., Morselli, M. G., Pechinger, U., Siebert, P., and Thomson, D. J., European Communities, ISBN 92–828-3302-X, Belgium, 1998.

Degreane Horizon: Degrewind PCL 1300 Processing Computer User Manual, Cuers, France, 2006.

Dinoev, T. S., Simeonov, V. B., Calpini, B., Parlange, M. B., Monitoring of Eyjafjallajökull Ash Layer Evolution over Payerne Switzerland with a Raman Lidar, in: Proceedings of the TECO 2010, Helsinki, Finland, 30 August to 1 September, Keynote 2, 2010.

Dinoev, T., Simeonov, V., Arshinov, Y., Bobrovnikov, S., Ristori, P., Calpini, B., Parlange, M., and van den Bergh, H.: Raman Lidar for Meteorological Observations, RALMO – Part 1: Instrument description, *Atmos. Meas. Tech.*, 6, 1329–1346, doi:10.5194/amt-6-1329-2013, 2013.

Dürr, B. and Philipona, R.: Automatic cloud amount detection by surface longwave downward radiation measurements, *J. Geophys. Res.*, 109, D05201, doi:10.1029/2003JD004182, 2004.

Emeis, S.: Surface-based remote sensing of the atmospheric layer, in: Atmospheric and Oceanographic Sciences Library, vol. 40, edited by: Mysak, L. A. and Hamilton, K., Springer, Dordrecht, Heidelberg, London, New York, 171 pp., 2009.

Fisher, B. E. A., Erbrink, J. J., Finardi, S., Jeannet, P., Joffre, S., Morselli, M. G., Pechinger, U., Siebert, P., and Thomson, D. J.: Cost Action 710-Final report: Harmonisation of the Pre-processing of meteorological data for atmospheric dispersion models, Office for Official Publications of the European Communities, Belgium, 1998.

Granados-Muñoz, M. J., Navas-Guzmán, F., Bravo-Aranda, J. A., Guerrero-Rascado, J. L., Lyamani, H., Fernández-Gálvez, J., and Alados-Arboledas, L.: Automatic determination of the planetary boundary layer height using lidar: one-year analysis over southeastern Spain, *J. Geophys. Res.*, 117, D18208, doi:10.1029/2012JD017524, 2012.

Haeffelin, M., Angelini, F., Morille, Y., Martucci, G., Frey, S., Gobbi, G. P., Lolli, S., O’Dowd, C. D., Sauvage, L., Xueref-Rémy, I., Wastine, B., and Feist, D. G.: Evaluation of mixing-height

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retrievals from automatic profiling lidars and ceilometers in view of future integrated networks in Europe, *Bound.-Lay. Meteorol.*, 143, 49–75, doi:10.1007/s10546-011-9643-z, 2012.

Holzworth, G. C.: Estimates of mean maximum mixing depths in the contiguous united states, *Mon. Weather Rev.*, 92, 235–242, 1964.

5 Jericevic, A. and Grisogono, B.: The critical bulk Richardson number in urban areas: verification and application in a numerical weather prediction model, *Tellus A*, 58, 19–27, 2006.

Ketterer, C., Zieger, P., Bukowiecki, N., Collaud Coen, M., Maier, O., Ruffieux, D., and Weingartner, E.: Investigation of the planetary boundary layer in the Swiss Alps using remote sensing and in-situ measurements, *Bound.-Lay. Meteorol.*, 151, 317–334, doi:10.1007/s10546-013-9897-8, 2014.

10 Liu, S. and Liang, X.-Z.: Observed diurnal cycle climatology of planetary boundary layer height, *J. Climate*, 23, 5790–5809, doi:10.1175/2010JCLI3553.1, 2010.

Mahrt, L., Sun, J., Blumen, W., Delany, T., and Oncley, S.: Nocturnal boundary-layer regimes, *Bound.-Lay. Meteorol.*, 88, 255–278, 1998.

15 Martucci, G., Matthey, R., Mitev, V., and Richner, H.: Comparison between backscatter lidar and radiosonde measurements of the diurnal and nocturnal stratification in the lower troposphere, *J. Atmos. Ocean. Technol.*, 24, 1231–1244, 2007.

Milroy, C., Martucci, G., Lolli, S., Loaec, S., Sauvage, L., Xueref-Remy, I., Lavrič, J. V., Ciais, P., Feist, D. G., Biavati, G., and O'Dowd, C. D.: An Assessment of Pseudo-Operational Ground-Based Light Detection and Ranging Sensors to Determine the Boundary-Layer Structure in the Coastal Atmosphere, *Advances in Meteorology*, 18 pp., doi:10.1155/2012/929080, 2012.

Radiometer Physics GmbH: Technical Instrument Manual, Meckenheim, Germany, 2011.

Salmond, J. A. and McKendry, J. A.: A review of turbulence in the very stable nocturnal boundary layer and its implications for air quality, *Prog. Phys. Geog.*, 29, 171–188, 2005.

25 Sawyer, V. and Li, Z.: Detection, variations and intercomparison of the planetary boundary layer depth from radiosonde, lidar and infrared spectrometer, *Atmos. Environ.*, 79, 518–528, 2013.

Schmid, P. and Niyogi, D.: A method for estimating planetary boundary layer heights and its application over the ARM Southern Great Plains Site, *J. Atmos. Ocean. Technol.*, 29, 316–322, doi:10.1175/JTECH-D-11-00118.1, 2012.

30 Seidel, D. J., Ao, C. O., and Li, K.: Estimating climatological planetary boundary layer heights from radiosonde observations: comparison of methods and uncertainty analysis, *J. Geophys. Res.*, 115, D16113, doi:10.1029/2009JD013680, 2010.

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Seidel, D. J., Zhang, Y., Beljaars, A. C. M., Golaz, J.-C., Jacobson, A. R., and Medeiros, B.: Climatology of the planetary boundary layer over the continental United States and Europe, *J. Geophys. Res.*, 117, D17106, doi:10.1029/2012JD018143, 2012.

Stull, R. B.: *An Introduction to Boundary Layer Meteorology*, vol. 13, Kluwer Academic Publishers, the Netherlands, Dordrecht/Boston/London, 1988.

Summa, D., Di Girolamo, P., Stelitano, D., and Cacciani, M.: Characterization of the planetary boundary layer height and structure by Raman lidar: comparison of different approaches, *Atmos. Meas. Tech. Discuss.*, 6, 5195–5216, doi:10.5194/amtd-6-5195-2013, 2013.

Szintai, B.: Improving the turbulence coupling between high resolution numerical weather prediction models and Lagrangian particle dispersion models, Ph.D. thesis, Swiss Federal Institute of Technology, Zürich, Switzerland, 2010.

Wang, Z., Cao, X., Zhang, L., Notholt, J., Zhou, B., Liu, R., and Zhang, B.: Lidar measurement of planetary boundary layer height and comparison with microwave profiling radiometer observation, *Atmos. Meas. Tech.*, 5, 1965–1972, doi:10.5194/amt-5-1965-2012, 2012.

White, A. B., Fairrall, C. W., and Wolfe, D. E.: Use of 915-mHz wind profiler data to describe the diurnal variability of the mixed layer, AMS, in : *Proc 7th Joint Conf. Appli. Poll. Meteor.*, AMS, 14–18 January, New Orleans, LA, J161–J166, 1991.

Zahng, N., Chen, Y., and Zhao, W.: Lidar and microwave radiometer observations of planetary boundary layer structure under light wind weather, *J. Appl. Rem. Sensing*, 6, 063513-1–063513-8, doi:10.1117/1.JRS.6.063513, 2012.

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**Table 1.** List of abbreviations.

Atmospheric layers	
CBL	Convective Boundary Layer
cloudy-CBL	CBL for overcast conditions but without precipitations
NBL	Neutral Boundary Layer
PBL	Planetary Boundary Layer
RL	Residual Layer
SBL	Stable Boundary Layer
Instruments	
lidar	Raman lidar
MWR	Microwave radiometer
RS	Radio sounding
WP	Windprofiler
Methods	
APCADA	Automatic Partial Cloud Amount Detection Algorithm
ASR	Aerosol Scattering Ratio
bR	Bulk Richardson number method
COSMO-2	COnsortium for Small-scale MOdeling
PM	Parcel Method
SBI	Surface-based Temperature Inversion
SBLpT	Stable Boundary Layer detected by potential Temperature
SNR	Signal to Noise Ratio
Measuring sites	
PAY	Payerne
SHA	Schaffhausen

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**Table 2.** Linear regression of PBL height detected by  $\theta_v$  as a function of the PBL height detection by  $\theta$ : slope, intercept, coefficient of determination between the data and the fit ( $R^2$ ), coefficient of determination between the data and the 1 : 1 line ( $R_{th}^2$ ), root mean square error of the  $x - y$  difference (RMS), median of the difference between  $x$  and  $y$  coordinates (median bias) and the number of considered data ( $N$ ). The results are given for PM and bR methods applied on RS and MWR data.

Instrument/ method	Slope	Intercept	$R^2$	$R_{th}^2$	RMS [m]	Median bias [m]	$N$
RS/PM	1.08	-12	0.95	0.92	110	18	35
RS/bR	1.05	5	0.97	0.94	99	26	35
MWR/PM	1.03	104	0.95	0.84	161	117	437
MWR/bR	1.05	67	0.95	0.88	154	93	420

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**Table 3.** Linear regression of PBL height computed with various methods and instruments as a function of RS/PM. See Table 2 for parameters description.

Instrument/ method	Slope	Intercept	$R^2$	$R_{th}^2$	RMS [m]	Median bias [m]	$N$
RS/bR	1.02	46	0.95	0.93	122	21.5	118
RS/RH	1.01	3.64	0.86	0.90	154	0	118
MWR/PM	0.89	73	0.75	0.74	228	-25.5	100
MWR/bR	0.84	173	0.72	0.69	239	2.33	85
WP/SNR	0.73	210	0.49	0.41	351	-64	105
Lidar/ASR	1.00	-3	0.81	0.81	211	-50	61
COSMO-2	1.20	141	0.72	0.43	472	275	114

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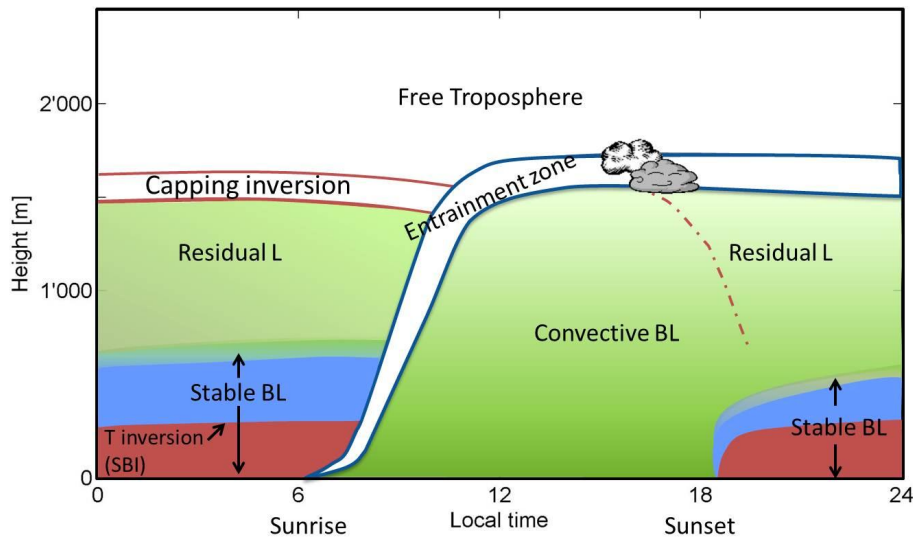
**Table 4.** Advantages and limits of detection methods and instruments to estimate the PBL height.

Method	Profiles	PBL height detected	Advantages	Limits
PM	$\theta$ or $\theta_v$	CBL, cloudy-CBL	<ul style="list-style-type: none"> <li>– also efficient under weak convective condition</li> <li>– early growth after sunrise until decrease when temporal gradient of surface <math>T</math> and vertical heat flux become negative</li> </ul>	<ul style="list-style-type: none"> <li>– requires negative gradient in <math>\theta</math> at the ground</li> <li>– not available during night</li> </ul>
bR	$\theta$ or $\theta_v$ + wind	CBL, cloudy-CBL, SBL	<ul style="list-style-type: none"> <li>– nighttime and daytime detection</li> <li>– transition between SBL and CBL at sunrise</li> <li>– CBL decrease also after the vertical heat flux and temporal <math>T</math> gradient become negative</li> <li>– SBL formation after sunset</li> <li>– describe the layer where the pollutants emitted during night are trapped</li> </ul>	<ul style="list-style-type: none"> <li>– requires wind profiles from WP or RS</li> <li>– often false SBL detection in case of constant <math>\theta</math> profile</li> </ul>
SBI	$T$	SBL	<ul style="list-style-type: none"> <li>– Formation and top of the stable nocturnal layer</li> <li>– measures the dynamics of aerosol dispersion</li> <li>– a real measure of the pollutants ML</li> </ul>	<ul style="list-style-type: none"> <li>– Not well defined limit of the SBL layered structure</li> <li>– no measure of the SBL</li> </ul>
SBLpT	$\theta$ or $\theta_v$	SBL	<ul style="list-style-type: none"> <li>– sometimes retrieves PBL height early growth after sunrise</li> <li>– only method based on the vertical structure of turbulence.</li> </ul>	<ul style="list-style-type: none"> <li>– large number of outliers due to false attributions</li> <li>– can also retrieve the cloud top</li> </ul>
Aerosol/humidity gradient	ASR, RH	CBL, RL		
SNR maxima	wind	CBL, (RL)		
Instrument	Profiles	PBL height detected	Advantages	Limits
Microwave radiometer	$T$ , RH, wind	CBL, cloudy-CBL, SBL	<ul style="list-style-type: none"> <li>– captures diurnal cycle</li> <li>– good data availability</li> <li>– good temporal resolution</li> <li>– daily cycle</li> <li>– can also retrieve the cloud top-based on the vertical structure of turbulence.</li> <li>– daily cycle</li> <li>– direct measurement of atmospheric composition</li> <li>– most accurate and precise data</li> <li>– best vertical resolution</li> </ul>	<ul style="list-style-type: none"> <li>– low vertical resolution</li> <li>– no PBL detection in case of precipitation</li> <li>– low data availability at low altitude</li> <li>– no data in case of fog, low clouds</li> <li>– needs maintenance</li> <li>– only twice a day at 00:00 and 12:00</li> </ul>
Windprofiler	Wind, SNR ratio	CBL, RL		
Lidar	ASR, RH	CBL, RL		
Radio Sounding	$T$ , $p$ , RH, wind	CBL, cloudy-CBL, SBL		



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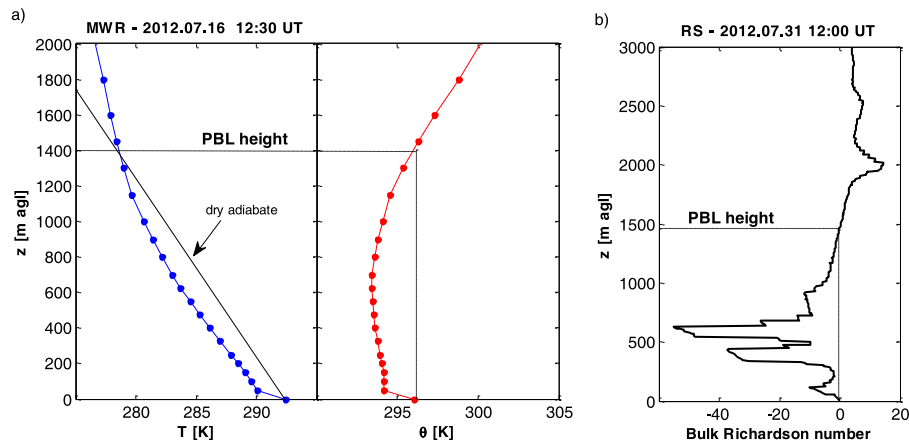


**Figure 1.** Diurnal cycle of the PBL height over land for a clear convective day (adapted from Stull, 1988).

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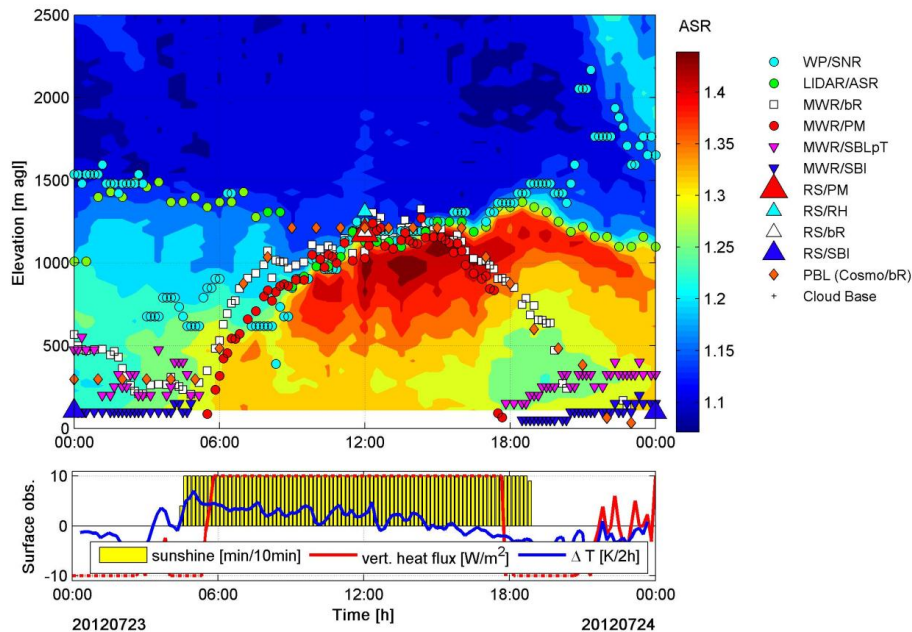


**Figure 2.** PBL detection methods based on  $T$  profiles: **(a)** Parcel method applied on MWR  $T$  and  $\theta$  profiles and **(b)** bulk Richardson number method applied on RS  $\theta$  profile. Both profiles were measured at about 12:00 on the 16 July 2012 in convective conditions.



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**Figure 4.** Upper panel: automatic detection of PBL height from all remote sensing instruments, RS and COSMO-2 model for a convective day in summer 2012 (23 July 2012) at PAY; the background signal corresponds to the lidar/ASR. Lower panel: sunshine duration, vertical heat flux and temporal gradient of surface  $T$ . Vertical heat flux greater than 10 or lower than  $-10 \text{ W m}^{-2}$  are limited to  $\pm 10$  with a dashed line.

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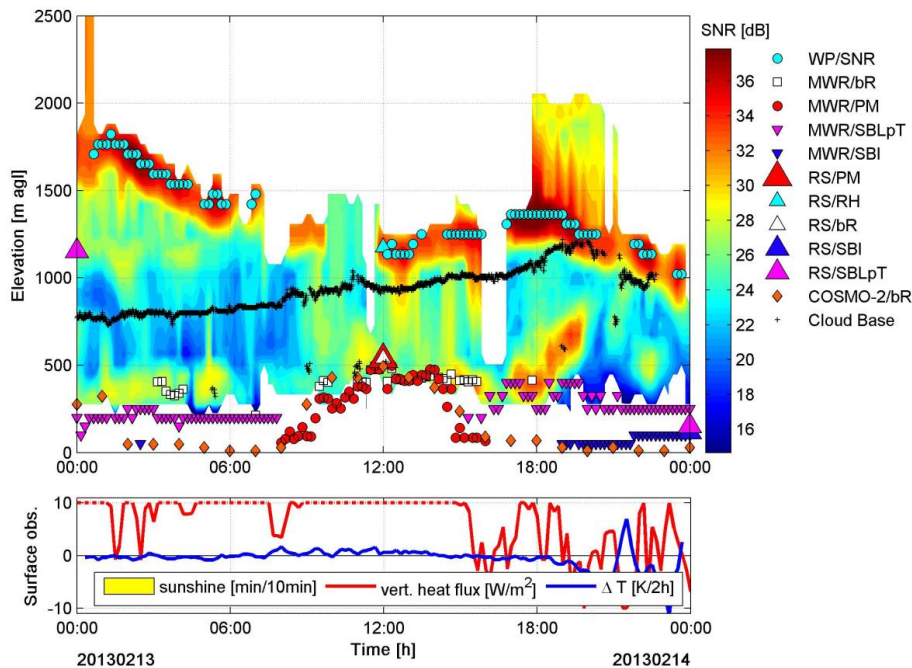
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**Figure 5.** Example of cloudy-CBL detection under cloudy conditions in winter (14 February 2013) plotted on WR/SNR signal as background. For symbol description see Fig. 4.

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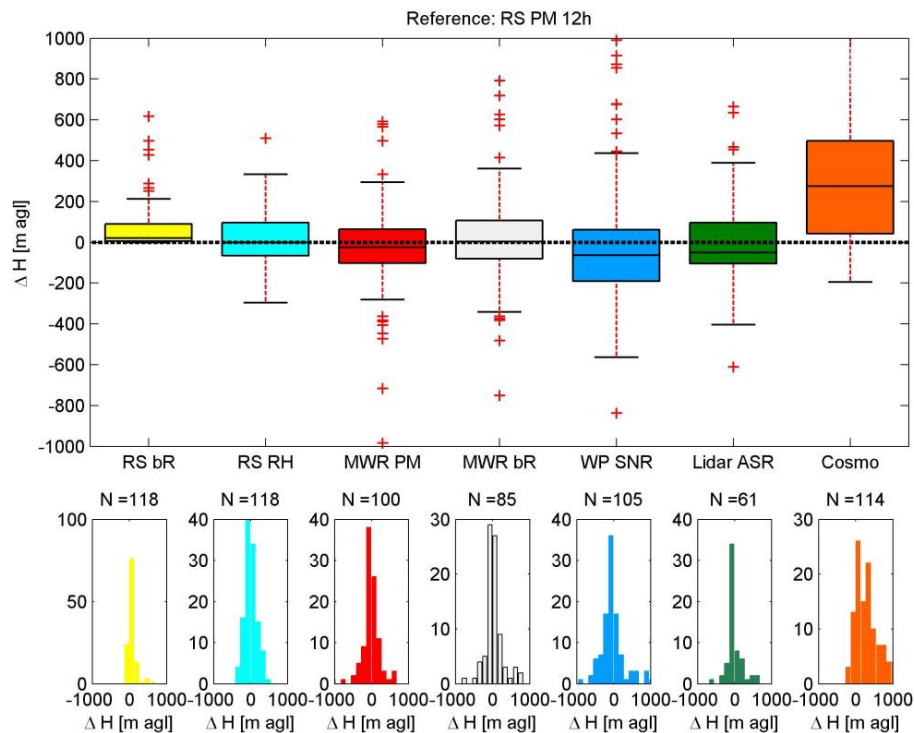
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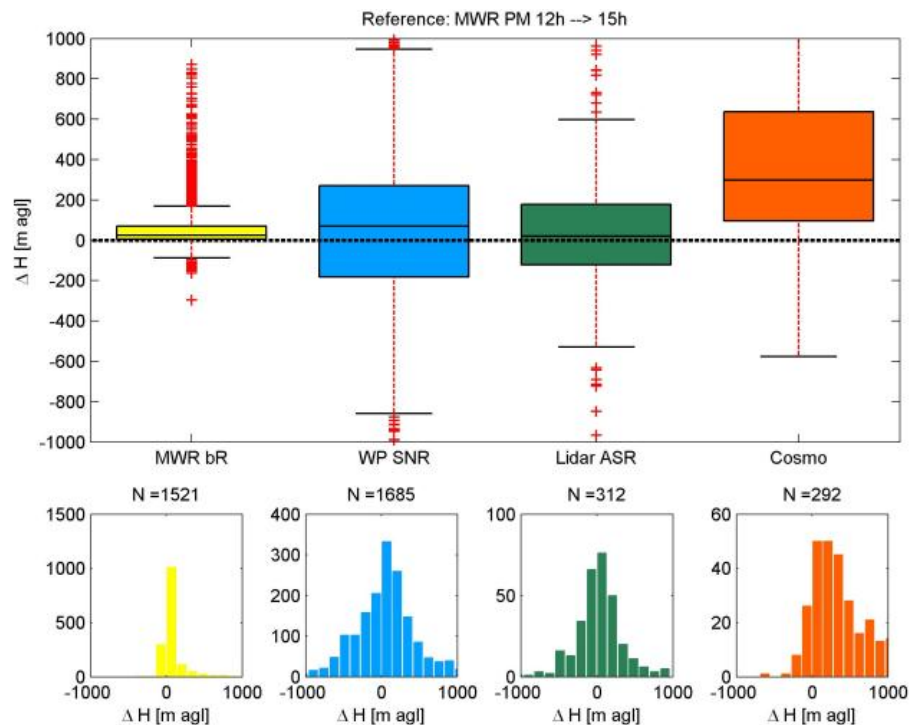
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**Figure 6.** Boxplots of PBL height differences  $\Delta H$  between RS/PM and other methods/instrumentation computed at 12:00. The central box line is the median, the edges of the box are the 25th and 75th percentiles ( $q_1$  and  $q_3$ ), the whiskers enclose all data points not considered outliers, and the red crosses are the outliers. Data are considered as outliers if they are larger than  $q_3 + 1.5 \cdot (q_3 - q_1)$  or smaller than  $q_1 - 1.5 \cdot (q_3 - q_1)$ , which means that whiskers cover 99% of data assuming a normal distribution. The  $\Delta H$  statistical distribution and the number of data points  $N$  for each boxplot are given in the sub plotted histograms, data points greater than 1000 m being displayed in the last column of the corresponding histogram.

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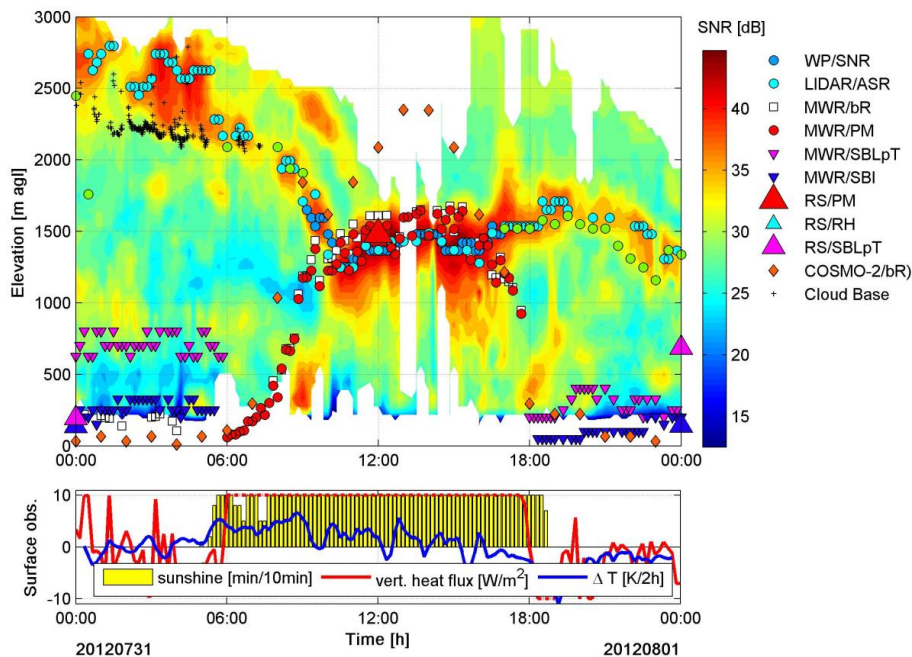


**Figure 7.** Same as Fig. 6 but between 12:00 and 15:00 UT and with MWR/PM taken as the reference.

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**Figure 8.** Example of CBL overestimation by COSMO-2, the background signal corresponds to the lidar/ASR. For a description of the symbols, see Fig. 4.

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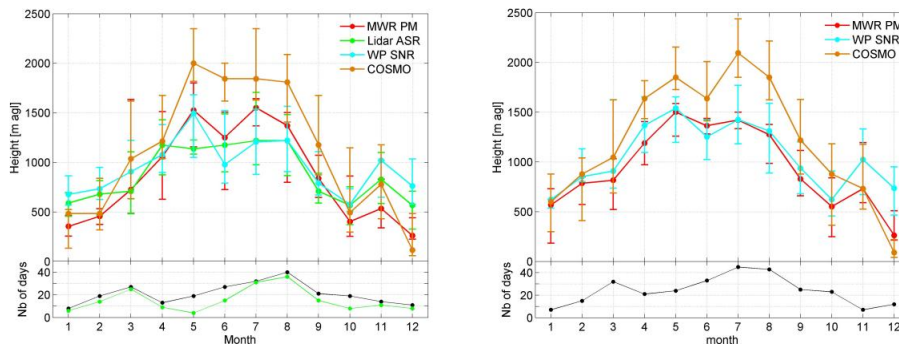
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**Figure 9.** Upper panel: CBL height two-years climatology at PAY (left) and SHA (right). The dots are the monthly median of the daily medians of the CBL height taken between 12:00 and 15:00; the error bars are the 25th and 75th percentiles. Lower panel: the number of convective days are given in black for MWR/PM, WP and COSMO-2 and in green for lidar/ASR.

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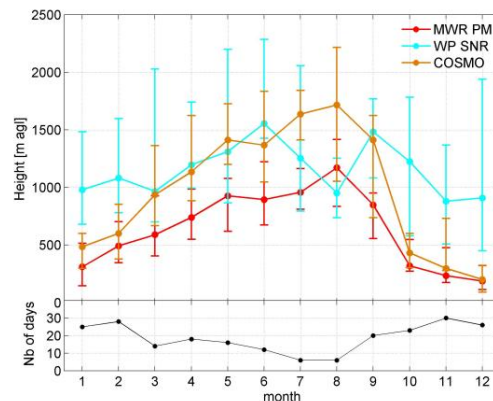
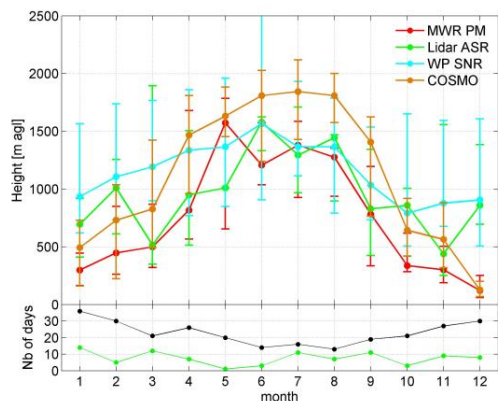
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**Figure 10.** Cloudy-CBL top height climatology at PAY (left) and SHA (right). Symbols as in Fig. 9.

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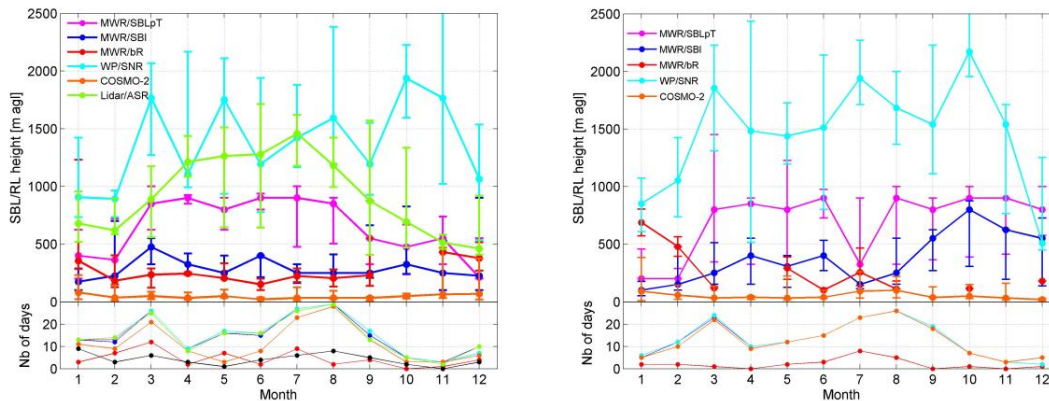
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**Figure 11.** SBL and RL heights for clear-sky conditions at PAY (left) and SHA (right). Symbols as in Fig. 9.

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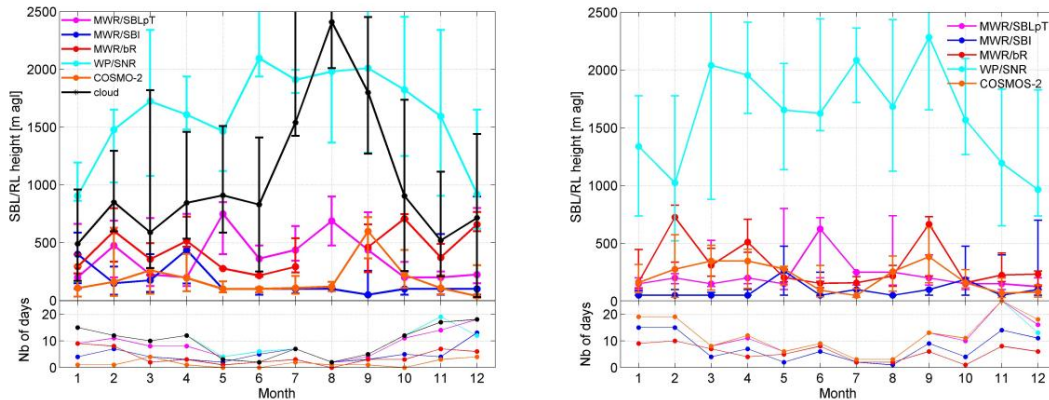
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**Figure 12.** SBL and RL heights for cloudy conditions at PAY (left) and SHA (right). Symbols as in Fig. 9.

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