

# A 2-year record of atmospheric mercury species at a background Southern Hemisphere station on Amsterdam Island

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Changes made in the revised manuscript have been highlighted in red.

## Abstract

Although essential to fully understand the cycling of mercury at the global scale, mercury species records in the Southern Hemisphere are scarce. Under the framework of the “Global Mercury Observation System” (GMOS) project, a monitoring station has been set up on Amsterdam Island (37°48'S, 77°34'E) in the remote southern Indian Ocean. For the first time in the Southern Hemisphere, a 2-year record of gaseous elemental mercury (GEM), reactive gaseous mercury (RGM) and particle-bound mercury (PBM) is presented. GEM concentrations were remarkably steady ( $1.03 \pm 0.08 \text{ ng/m}^3$ ) while RGM and PBM concentrations were very low and exhibited a strong variability (mean:  $0.34 \text{ pg/m}^3$  [range: < detection limit-4.07  $\text{pg/m}^3$ ] and mean:  $0.67 \text{ pg/m}^3$  [range: < detection limit-12.67  $\text{pg/m}^3$ ], respectively). Despite the remoteness of the island, wind sector analysis, air mass back trajectories and the observation of radonic storms highlighted a long-range contribution from the southern African continent to the GEM and PBM budgets from July to September during the biomass burning season. Low concentrations of GEM were associated with southerly polar and marine air masses from the remote southern Indian Ocean. This unique dataset provides new baseline GEM concentrations in the Southern Hemisphere mid-latitudes while mercury speciation along with upcoming wet deposition data will help improving our understanding of mercury cycle in the marine boundary layer.

## 33 1 Introduction

34

35 Due to its toxicity, persistence, bioaccumulative nature and long-range transport, mercury  
36 (Hg) is a global threat to ecosystems and human health. Since the 70's, multiple regulations  
37 **have been** implemented to tackle the **exposure** of populations to this contaminant. In 2013, the  
38 United Nations Environment Programme (UNEP) opened for signature a new legally-binding  
39 treaty on mercury, giving birth to the Minamata Convention on mercury (UNEP, 2013).  
40 **However, research gaps limiting mercury reduction policies at regional or global scale remain.**  
41 **For example, the policy effectiveness at reducing deposition of mercury requires a better**  
42 **knowledge of the chemistry of atmospheric mercury species (Selin, 2014).**

43 According to recent estimates (Amos et al., 2013) while 10% of annual global emissions of  
44 atmospheric mercury currently come from natural geological sources – e.g., volcanic  
45 emissions or mercury-containing rocks –, 30% are produced by a variety of anthropogenic  
46 activities – e.g., coal combustion, cement production, waste incineration or artisanal and  
47 small-scale gold mining –, and re-emissions of previously released mercury account for the  
48 remaining 60%. Gaseous elemental mercury (GEM,  $\text{Hg}^0$ ) is the dominant form of atmospheric  
49 mercury (Lindberg and Stratton, 1998). It can be oxidized by ozone or free radicals into  
50 highly reactive and water-soluble divalent species ( $\text{Hg}^{2+}$ ) and/or particle-bound mercury  
51 (PBM) (Lin and Pehkonen, 1999) that can be deposited through wet and dry processes  
52 (Lindqvist and Rodhe, 1985).

53 In remote areas far from any local sources, atmospheric deposition has been recognized as the  
54 main source of mercury to the ocean (Lindberg et al., 2007). Mercury can then be reemitted  
55 back to the atmosphere via gas exchange (Schroeder and Munthe, 1998) and modeling studies  
56 suggest that reemission from oceans is a major contributor to atmospheric concentrations of  
57 GEM, particularly in the Southern Hemisphere where oceans were shown to contribute more  
58 than half of the surface atmospheric concentration (Strode et al., 2007). **To better understand**  
59 **the cycling of mercury at the global scale a coordinated global monitoring network is needed**  
60 **(Pirrone et al., 2013), along with long-term records of atmospheric mercury species in the**  
61 **Southern Hemisphere and at background sites (Sprovieri et al., 2010).** To date observations in  
62 the Southern Hemisphere mainly rely on a few oceanographic campaigns (e.g., Lamborg et  
63 al., 1999; Temme et al., 2003a; Witt et al., 2010) and on-going ground-based monitoring  
64 surveys at the Cape Point station in South Africa (Slemr et al., 2008) and at Troll, Dumont  
65 d'Urville and Concordia stations in Antarctica (Pfaffhuber et al., 2012; Dommergue et al.,  
66 2013a; Dommergue et al., 2013b).

67 In this context, a monitoring station has been set up on Amsterdam Island, a remote island in  
68 the southern Indian Ocean, under the framework of the European Union-financed project  
69 “Global Mercury Observation System” (GMOS, <http://www.gmos.eu/>). The 2-year record of  
70 elemental, divalent and particle-bound mercury concentrations presented here is, to the best of  
71 the authors’ knowledge, the first reported in the Southern Hemisphere mid-latitudes. Along  
72 with mercury species, ancillary parameters were analyzed to categorize air masses reaching  
73 the station based on their source region. The main objective of this study is to investigate to  
74 what extent observations at Amsterdam Island could define Southern Hemisphere mid-  
75 latitudes background conditions and provide new constraints in multi-scale mercury species  
76 cycling models.

77

## 78 **2 Materials and Methods**

### 79 **2.1 Sampling site description**

80 Amsterdam Island is a small island (55 km<sup>2</sup>) located in the southern Indian Ocean (37°48’S,  
81 77°34’E), 3400 km and 5000 km downwind from the nearest lands, Madagascar and South  
82 Africa, respectively (see Figure 1). Instrumentation dedicated to the study of atmospheric  
83 mercury is located at the Pointe Bénédicte station, at the northwest end of the island, 55 m  
84 above sea level and 2 km west of the scientific base (30 residents at most). Other monitoring  
85 activities are performed at the station for various atmospheric compounds such as ozone (Gros  
86 et al., 1998), carbon monoxide (Gros et al., 1999), total aerosol number concentration (Sciare  
87 et al., 2001) or carbonaceous aerosol (Sciare et al., 2009).

### 88 **2.2 Ancillary parameters**

89 Meteorological data – air temperature, relative humidity, barometric pressure, wind speed, and  
90 wind direction – were provided by the local meteorological station. Radon 222 and 220 (Rn)  
91 are monitored with a precision of 10% on a 2-hour basis. The method is described in detail by  
92 Polian et al. (1986) and Kritz et al. (1990). It is assumed that <sup>222</sup>Rn and <sup>220</sup>Rn are in  
93 radioactive equilibrium with their short-lived daughters so that <sup>222</sup>Rn and <sup>220</sup>Rn concentrations  
94 can be calculated by measuring the concentration of their short-lived decay products. Upon  
95 formation these short-lived daughters are quickly and irreversibly scavenged by aerosols and  
96 sampled by filtration. The detection then relies on the measurement over time of the decrease  
97 of alpha radioactivity of these aerosols.

98 O<sub>3</sub> measurements have been performed at the Pointe Bénédicte station since 1994, halfway up  
99 on a 20-m high tower (Gros et al., 1998). Unfortunately, 2012 and 2013 data are not available  
100 due to technical problems.

101 A new cavity ring-down spectroscopy (CRDS) analyzer was installed at the Pointe Bénédicte  
102 station in 2012 for in-situ measurements of CH<sub>4</sub>. This instrument (G2301, Picarro) is  
103 calibrated once a month with four reference gases (NO11-2004 scale for CH<sub>4</sub>).

104 7-days air mass back trajectories were calculated at 60 meters above sea level, approximate  
105 sampling height for mercury analysis, using the HYSPLIT (HYbrid Single-Particle  
106 Lagrangian Integrated Trajectory) model accessed via NOAA Air Resources Laboratory  
107 READY (Real-time Environmental Applications and Display sYstem) website (Draxler and  
108 Rolph, 2013;Rolph, 2013). **Calculated back trajectories always have some uncertainty, arising  
109 for example from the possible errors in input meteorological fields and the numerical methods  
110 (Yu et al., 2009), and increasing with time along the path (Stohl, 1998). As suggested by Jaffe  
111 et al. (2005) it should be noted that back trajectories only give a general indication of the  
112 source region.**

113 Fire counts and chlorophyll-a data west of Amsterdam Island were obtained via the FIRMS  
114 MODIS Fire Archive Download (Davies et al., 2009) and the Giovanni online data system  
115 developed and maintained by the NASA GES DISC, respectively.

### 116 **2.3 Mercury analyzers**

117 Since January 2012 we have monitored **3 atmospheric mercury species**: GEM, PBM (< 2.5  
118 μm) and reactive gaseous mercury (RGM), the latter consisting of various oxidized gaseous  
119 Hg<sup>2+</sup> compounds and hereafter defined as all forms of mercury sampled using a KCl-coated  
120 denuder (Landis et al., 2002).

121 Atmospheric mercury species measurements were performed using a Tekran mercury  
122 speciation unit (Tekran 1130 and 1135) coupled to a Tekran 2537B analyzer (Tekran Inc.,  
123 Toronto, Canada). Concentrations are expressed in ng/m<sup>3</sup> (GEM) or pg/m<sup>3</sup> (PBM and RGM),  
124 at standard temperature and pressure (273.15 K, 1013.25 hPa). GEM was determined at sub  
125 ng/m<sup>3</sup> levels using a gas-phase mercury analyzer, based on the amalgamation of mercury onto  
126 a gold cartridge followed by a thermal desorption and a detection by an integrated cold vapor  
127 atomic fluorescence spectrometer (CVAFS) at 253.7 nm (Fitzgerald and Gill, 1979;Bloom and  
128 Fitzgerald, 1988). The presence of two gold cartridges allowed alternating sampling and  
129 desorption modes and thus a continuous analysis of GEM in the sample air stream. In order to  
130 protect the two gold cartridges against deleterious compounds such as acid gases and halogen

131 compounds, and against particulate matter, the sample air stream – **after exiting the speciation**  
132 **unit** – was pre-filtered through a sodalime trap and a 0.2 µm PTFE filter.

133 The speciation unit was located on the roof top of the station, the sampling inlet being at 6 m  
134 above the ground, and connected to the 2537B analyzer through a 10 m-long PTFE heated  
135 line (50°C). The sampling resolution was 5 min for GEM and **4 hours** for RGM and PBM,  
136 with sampling flow rates of 1 L/min and 10 L/min, respectively. Measurements were achieved  
137 through a multi-step procedure as described elsewhere (Lindberg et al., 2002) using an  
138 impactor inlet (2.5 µm cut-off aerodynamic diameter at 10 L/min), a KCl-coated quartz  
139 annular denuder in the 1130 unit, and a quartz regenerable particulate filter (RPF) in the 1135  
140 unit.

141

#### 142 *Quality assurance and quality control procedures*

143 Fortnightly to monthly routine maintenance operations on the denuder, RPF, sodalime trap  
144 and filters, along with thorough cleaning steps and weekly site visits enabled the collection of  
145 contamination-free air samples. The accuracy of flow measurements was checked twice a year  
146 with a calibrated flow meter (Definer 220). An automatic calibration step of the 2537B  
147 analyzer was carried out every 69 hours with an internal mercury permeation source. The  
148 accuracy of this permeation source was annually checked against manual injections of  
149 saturated mercury vapor taken from a temperature controlled vessel, using a Tekran 2505  
150 mercury vapor calibration unit and a Hamilton digital syringe, and following a strict  
151 procedure adapted from Dumarey et al. (1985). Both routine and exceptional maintenance  
152 were compiled and archived via a software program developed at the LGGE. This software  
153 program also enabled a rapid data processing in order to produce clean time-series of GEM,  
154 PBM and RGM. Screening criteria for data validation/invalidation **were inspired by standard**  
155 **operative protocols used by the Canadian Atmospheric Mercury Measurement Network**  
156 **(CAMNet) and the Atmospheric Mercury Network (AMNet) (Steffen et al., 2012). To ensure**  
157 **uniformity across the network, GMOS is currently developing a quality control (QC) software**  
158 **and an intercomparison with the AMNet QC software will be undertaken.**

159 **The best estimate of** the detection limit (DL) for GEM measurements was 0.10 ng/m<sup>3</sup> (Tekran,  
160 2011) and ten injections of 15 µL of saturated mercury vapor were performed to check the  
161 repeatability of the system response. The system gave a relative expanded uncertainty of 1%  
162 (95% confidence level). Validated 5-minute GEM data subsets were compiled into hourly-  
163 average data when the hourly recovery rate exceeded 50% (number of valid data records  
164 collected vs. that possible over the reporting period).

165 RGM and PBM sampling resolution was shifted from 3 to 4 hours after a few days due to  
166 very low concentrations, and the detection limit therefore decreased from 0.42 pg/m<sup>3</sup> to 0.28  
167 pg/m<sup>3</sup> based on the sampling volumes and GEM detection limit (Wang et al., 2014). The  
168 mean of the distributions was estimated using the Kaplan-Meier cumulative proportion-based  
169 method. It provides more reliable results for data sets containing below-detection limit values  
170 than the substitution method, i.e. replacement of below-detection limit values by a constant  
171 equal to 0, 0.5 DL or DL (Helsel, 2005). 75% and 50% of RGM and PBM measurements,  
172 respectively, were below the limits of detection resulting in differences for mean values up to  
173 60% and 15%, respectively, comparing Kaplan-Meier and normally averaged datasets. Values  
174 3.3 times above the stated detection limits, i.e. reliably quantified, will be discussed thereafter  
175 and referred to as RGM and PBM events (3% and 18% of RGM and PBM measurements,  
176 respectively).

177 There is growing evidence that RGM and PMB measurements might suffer from significant  
178 biases and interferences (Lyman et al., 2010;Gustin et al., 2013;Jaffe et al., 2014). Several  
179 studies highlighted the inefficient collection of gaseous oxidized mercury compounds with a  
180 KCl-coated denuder in the Tekran technique (Gustin et al., 2013;Huang et al., 2013), leading  
181 to an underestimation of reactive mercury concentrations by a factor 1.3 to 3.7 (Huang et al.,  
182 2013). Other studies suggested sampling artifacts for PBM measurements due to temperature  
183 or sampling duration (Malcolm and Keeler, 2007;Rutter et al., 2008). Moreover, the upper  
184 size cut-off diameter at 2.5 μm raises concerns about mercury associated with large (> 2.5  
185 μm) particle fractions (Kos et al., 2013), especially in the marine environment where mercury  
186 is likely mainly contained in coarse sea salt aerosols (Talbot et al., 2011;Feddersen et al.,  
187 2012). There is no robust calibration technique of the Tekran speciation unit and no certified  
188 reference material available. The precision of RGM measurements – shown to be of 15%  
189 under given conditions (Landis et al., 2002) – should be assessed in various sampling  
190 environments (e.g., varying ozone/relative humidity conditions). Given the limitations of the  
191 RGM and PBM measurements, data reported in this study should thus only be directly  
192 compared with the existing Tekran-based literature, as suggested by Wang et al. (2014). An  
193 extensive dataset has been gathered worldwide using the Tekran speciation technique, which  
194 is the best available automated method. Future interference and calibration tests are  
195 fundamental to validate measurements and quantify uncertainties (Kos et al., 2013), and might  
196 enable us to correct RGM and PBM data. Until then, orders of magnitude and variability in  
197 time and space of Tekran-based RGM and PBM concentrations can be used as first estimates  
198 by policy makers or to evaluate atmospheric models.

## 200 **3 Results and Discussion**

### 201 **3.1 Meteorological data**

202 Climate is mild oceanic, with frequent presence of clouds. Seasonal boundaries were defined  
203 as follows: winter from July to September and summer from December to February, in line  
204 with other studies performed at Amsterdam Island (NGuyen et al., 1990;Gros et al.,  
205 1998;Sciare et al., 2009). During the two-year period under discussion here (January 2012  
206 **until** December 2013), the monthly median air temperature ranged from 11°C in austral winter  
207 to 17°C in austral summer, while the monthly median relative humidity remained high and  
208 ranged from 65 to 85% most of the year. Precipitation was very frequent with total  
209 precipitation of 1262 mm in 2012 and 1128 mm in 2013, in good agreement with the 1124  
210 mm 40-year average reported by Miller et al. (1993). Wind speed remained comparatively  
211 high throughout the year (from 5 to 15 m/s); strong northwesterly winds peaked during winter  
212 months (**July to September**) when the roaring forties were at a maximum (see Figure 2).

### 213 **3.2 Gaseous elemental mercury concentrations**

#### 214 **3.2.1 Seasonality and contribution from biomass burning**

215 GEM concentrations were very steady with an average hourly mean concentration of  
216  $1.03 \pm 0.08 \text{ ng/m}^3$  (mean  $\pm$  standard deviation, range: 0.72 – 1.55  $\text{ng/m}^3$ ; n = 10 285, see Figure  
217 3). GEM data are lower than concentrations reported in the Northern Hemisphere but well  
218 within the expected range for a remote marine site in the Southern Hemisphere (**Sprovieri et**  
219 **al., 2010**). Indeed, Witt et al. (2010) measured a mean TGM (Total Gaseous Mercury defined  
220 as the sum of gaseous mercury species) concentration of  $1.24 \pm 0.06 \text{ ng/m}^3$  in the Indian Ocean  
221 at latitudes ranging from 9 to 21°S, while TGM concentrations ranged from 1.20 to 1.40  
222  $\text{ng/m}^3$  at the Cape Point station, South Africa (34°21'S, 18°29'E) between 1995 and 2004  
223 (Slemr et al., 2008) and GEM concentrations amounted to  $0.93 \pm 0.19 \text{ ng/m}^3$  at Troll station in  
224 Antarctica (Pfaffhuber et al., 2012). GEM concentrations measured at Amsterdam Island are  
225 furthermore highly consistent with data reported by Wang et al. (2014) in the marine  
226 boundary layer over the equator at Galápagos Islands (0°57'S, 90°58'W;  $1.08 \pm 0.17 \text{ ng/m}^3$ )  
227 despite occasional influence of northern hemispheric air at this station. A comprehensive  
228 comparison of mercury concentrations measured in the Southern Hemisphere is given in



229 Slemr et al. (2014).

230 Whereas TGM and GEM concentrations at the Cape Point station and at Galápagos Islands,  
231 respectively, showed a seasonal variation (Slemr et al., 2008; Wang et al., 2014), with  
232 minimum in winter and maximum in summer, GEM data at Amsterdam Island followed an  
233 opposite trend, with slightly but significantly higher concentrations in winter (July to  
234 September) than in summer (December to February) ( $1.06 \pm 0.09 \text{ ng/m}^3$  vs.  $1.04 \pm 0.07 \text{ ng/m}^3$ , p-  
235 value  $< 2.2 \cdot 10^{-16}$ , Mann-Whitney test, see Figure 4a).

236 This seasonality of GEM concentrations is in agreement with more frequent air masses  
237 originating from southern Africa (northwesterly winds) from July to September. In order to  
238 further test this hypothesis GEM data were sorted according to wind direction imposing a  
239 strict criterion on wind speed ( $> 8 \text{ m/s}$ ) to remove any local influence (Monfray et al., 1987).  
240 With north at  $0^\circ$ , southerly (S) winds ranged from  $100$  to  $250^\circ$  and northwesterly (NW) winds  
241 from  $250$  to  $300^\circ$ . GEM concentrations monitored during periods of NW flow were shown to  
242 be significantly higher than during S flow ( $1.05 \pm 0.08 \text{ ng/m}^3$  (n=2 833) vs.  $1.01 \pm 0.08 \text{ ng/m}^3$   
243 (n=822), p-value  $< 2.2 \cdot 10^{-16}$ , Mann-Whitney test, see Figure 4b).

244 The GEM budget on the island could be enhanced from July to September by long-range  
245 transport during the burning season in southern Africa. 2012 and 2013 satellite observations  
246 of fires west of Amsterdam Island (latitude ranging from  $3$  to  $53^\circ \text{S}$  and longitude from  $10$  to  
247  $73^\circ \text{E}$ ) showed that the burning season extended from May to October, peaking between June  
248 and September (see Figure 5a), in line with observations reported by Cooke et al. (1996). This  
249 observation is also supported by the concomitant seasonal maxima on Amsterdam Island of  
250 CO (Gros et al., 1999), equivalent black carbon, non-sea-salt potassium and oxalate (Sciare et  
251 al., 2009), the latter two being commonly used as tracers for biomass burning.

252

### 253 3.2.2 Short-time variations

254 The atmosphere at Amsterdam Island could define background marine boundary layer  
255 conditions, with minimum perturbation from anthropogenic influences. However, as reported  
256 by Balkanski and Jacob (1990), the rapid export of air from southern Africa to the  
257 subantarctic Indian Ocean could constitute a major source of pollution to southern mid-  
258 latitudes. The influence of continental air mass advection on GEM concentrations was thus  
259 investigated.

260 The background variability of GEM concentrations was assessed following a procedure  
261 adapted from Gros et al. (1999) calculating the difference, dGEM, between hourly GEM



262 concentrations and the monthly mean. dGEM events higher than  $\pm 0.18 \text{ ng/m}^3$ , 3 times the  
263 mean monthly standard deviation of dGEM measurements, were further investigated.  
264  $^{222}\text{Rn}$ , decay product of  $^{238}\text{U}$  with a 3.8 day half-life, is particularly well suited as a tracer of  
265 continental air over the oceans (Balkanski and Jacob, 1990). On the other hand,  $^{220}\text{Rn}$  and its  
266 daughter  $^{212}\text{Pb}$ , due to 54 s and 10.6 h half-lives, respectively, can only be attributed to local  
267 outgassing from Amsterdam Island's soil and not to any marine nor remote continental source  
268 (Polian et al., 1986). Therefore,  $^{222}\text{Rn}$  activities below  $100 \text{ mBq/m}^3$  are considered as typical  
269 for marine air (Brunke et al., 2004), whereas air with  $^{222}\text{Rn}$  activity above  $100 \text{ mBq/m}^3$  along  
270 with a  $^{220}\text{Rn}$  ( $^{212}\text{Pb}$ ) activity **below  $3.7 \text{ mBq/m}^3$**  is considered to be significantly influenced by  
271 a remote continent (Gros et al., 1999; Williams et al., 2001). Rapid and sharp variations of  
272  $^{222}\text{Rn}$  activity, referred to as "radonic storms" (Lambert et al., 1970), could be observed at  
273 Amsterdam Island, usually peaking around  $400 \text{ mBq/m}^3$ , and exceptionally reaching  $1000$   
274  $\text{mBq/m}^3$ . **The occurrence of radonic storms was about 4% in 2012 and 7% in 2013.**  
275 The local production of radon on Amsterdam Island can explain the radonic storm occurring  
276 on December 13, 2012 (see Figure 6a). As reported by Polian et al. (1986), it is associated  
277 with low wind speeds (below 5 m/s) and meteorological conditions corresponding to a low  
278 atmospheric eddy diffusion. The back trajectory ending on December 13, 2012 on Amsterdam  
279 Island (see Figure 7) meanders over the ocean, and low south-eastern/eastern winds prevailed  
280 ahead of  $^{222}\text{Rn}$  and  $^{220}\text{Rn}$  peaks, in good agreement with an influence from the island and local  
281 radon exhausts. The associated high GEM event ( $\text{dGEM} > 0.18 \text{ ng/m}^3$ ; see Figure 6b) was  
282 significantly positively correlated with  $^{222}\text{Rn}$  and  $^{220}\text{Rn}$  activities ( $r = 0.83$ ,  $p\text{-value} = 0.005$   
283 and  $r = 0.80$ ,  $p\text{-value} = 0.010$  respectively, Spearman test) and can therefore be attributed to a  
284 local non-anthropogenic source as no noteworthy activity occurred at the sampling station nor  
285 on the scientific base.  
286 **About 50%** of sharp high GEM events ( $\text{dGEM} > 0.18 \text{ ng/m}^3$ ) were associated with  $^{222}\text{Rn}$   
287 peaks, strong winds and  $^{220}\text{Rn}$  activities **below  $3.7 \text{ mBq/m}^3$** , and therefore ascribed to  
288 continental air mass advection. For example, the high GEM event occurring on September 21,  
289 2013 (see Figure 6d) was associated with a  $^{222}\text{Rn}$  peak of about  $200 \text{ mBq/m}^3$ , a  $^{220}\text{Rn}$  activity  
290 **below  $3.7 \text{ mBq/m}^3$**  and 15 m/s winds (see Figure 6c). This high GEM event was significantly  
291 positively correlated with  $^{222}\text{Rn}$  activities **but not correlated with  $^{220}\text{Rn}$  activities** ( $r = 0.81$ ,  $p\text{-}$   
292  $\text{value} = 9.3 \cdot 10^{-5}$  and  $r = -0.27$ ,  $p\text{-value} = 0.295$  respectively, Spearman test). The back  
293 trajectory ending on September 21, 2013 on Amsterdam Island (see Figure 7) passes near the  
294 southern African continent and Madagascar, confirming the continental origin of the air at the  
295 time of enhanced  $^{222}\text{Rn}$  activity and GEM concentrations. Although the back trajectory

296 indicates no direct passage over the continent, a continental origin of the air is not unlikely  
297 since that range of  $^{222}\text{Rn}$  activities has already been measured for trajectories calculated to  
298 have passed more than 100 km away from mainland Australia (Whittlestone et al., 1998).

299 As for high GEM events, continental induced radonic storms mainly occurred between June  
300 and October reflecting the seasonality of wind direction at Amsterdam Island and of long-  
301 range transport from the southern African continent.

302 As illustrated on Figures 6d and 7, low GEM events ( $\text{dGEM} < 0.18 \text{ ng/m}^3$ ) were well  
303 correlated with trajectories passing over Antarctica. This continent being almost all ice-  
304 covered, local emission of  $^{222}\text{Rn}$  and  $^{220}\text{Rn}$  is low (Polian et al., 1986) explaining why low  
305 GEM events are not associated with any  $^{222}\text{Rn}$  peak.

306 These short-term variations suggest that cleaner air masses originate from the remote southern  
307 Indian Ocean while NW air masses are influenced by continental southern Africa, as  
308 previously assumed by several authors (Miller et al., 1993; Gros et al., 1999; Williams et al.,  
309 2001).

310

### 311 **3.2.3 Southern Hemisphere mid-latitudes baseline concentrations for** 312 **modeling studies**

313 **The monthly means, medians and standard deviations are given in Table 1.** Over the 2 years  
314 **the occurrence of high and low GEM events was less than 1%** and these events did not  
315 significantly affect the monthly and annual means. Medians have been calculated for each  
316 month and differ in average from the arithmetical means by 0.90% which is less than the  
317 relative expanded uncertainty of the system response. **The difference between mean and**  
318 **median did not show any seasonal variation.**

319 GEM concentrations at Amsterdam Island can be considered as baseline concentrations and  
320 be used as is in modeling studies. However, a slight but significant seasonal cycle was  
321 highlighted and despite the remoteness of the station an influence of biomass burning was  
322 observed **from July to September**. Biomass burning affects all the mid-latitude belt of the  
323 Southern Hemisphere at least and can be considered as a widespread pollution (Fishman et al.,  
324 1991; Gros et al., 1999). Biomass burning slightly contributes to the background GEM level in  
325 this region and should be carefully considered in modeling studies when dealing with  
326 seasonality of GEM concentrations in the Southern Hemisphere mid-latitudes.

327

### 3.3 Oxidized mercury species: seasonality and sources

A 2-year record of RGM and PBM concentrations is presented hereafter – the longest ever reported in the Southern Hemisphere. Concentrations at Amsterdam Island were very low – at the lower end of the range reported during oceanographic campaigns (Laurier et al., 2003; Temme et al., 2003b; Laurier and Mason, 2007) –, and exhibited a strong variability (see Figure 3). RGM and PBM mean concentrations amounted to  $0.34 \text{ pg/m}^3$  [range:  $<DL-4.07 \text{ pg/m}^3$ ] and  $0.67 \text{ pg/m}^3$  [range:  $<DL-12.67 \text{ pg/m}^3$ ], respectively. Such low RGM and PBM concentrations at Amsterdam Island could be explained by the very frequent drizzle efficiently scavenging oxidized mercury species. To further investigate the latter assumption, a precipitation collector was set up on the island at the beginning of 2013 in order to analyze mercury species in rainwater.

#### 3.3.1 Reactive gaseous mercury

RGM can be emitted from point sources or originate from oxidation of GEM. Due to its short lifetime RGM can only be transported tens to hundreds of kilometers in the boundary layer (Schroeder and Munthe, 1998). Monitoring of primary emitted RGM at Amsterdam Island is therefore unlikely.

RGM in the marine boundary layer has been reported to originate from a photochemically driven oxidation of GEM (Hedgecock and Pirrone, 2001; Hedgecock et al., 2003; Laurier et al., 2003) or through entrainment from the free troposphere (Holmes et al., 2009).

Oxidation pathways of GEM involving ozone ( $\text{O}_3$ ), the hydroxyl radical (OH), atomic bromine (Br) or nitrogen dioxide ( $\text{NO}_2$ ), or a combination of them, have been suggested by modeling and field studies (Holmes et al., 2010; Wang et al., 2014). Our understanding of RGM production mechanisms at Amsterdam Island is still limited and no anti-correlation between RGM and GEM concentrations, suggesting an in-situ GEM oxidation, was found.

A slight but significant seasonal trend in RGM concentrations was highlighted ( $1.34 \pm 0.45 \text{ pg/m}^3$  in winter (July to September) vs.  $1.58 \pm 0.35 \text{ pg/m}^3$  in summer (December to February),  $p$ -value = 0.01, Mann-Whitney test), and RGM events occurred about 55% of the time between December and March (see Figure 5b), in line with an enhanced photochemistry in summer. While a significant negative correlation between  $\text{CH}_4$  and air temperature was observed ( $r = -0.638$ ,  $p < 2.2 \cdot 10^{-16}$ , Spearman test), consistent with its photochemical destruction by the hydroxyl radical (Khalil and Rasmussen, 1983), no correlation was found between RGM concentrations and air temperature, or any other meteorological parameter.

361 The lack of correlation between RGM concentrations and other parameters may come from  
362 the small number of RGM measurements above quantification limit (n=87).  
363 More frequent RGM events between December and March could also be in line with an  
364 enhanced biological activity in summer. The production of halogen species, photochemically  
365 oxidizing GEM, could be driven by biological activity (Gschwend et al., 1985). Unlike the  
366 oceanic region surrounding Amsterdam Island, an area located in a southwest upwind sector  
367 covering the subtropical front (see Figure 8) is highly productive, with a marine productivity  
368 (characterized by chlorophyll-a concentration) peaking from December to January and  
369 sometimes in March-April (Sciare et al., 2009), in agreement with peaks of RGM events.  
370 Similarly, marine organic aerosol concentrations at Amsterdam Island have been shown to be  
371 directly related to the seasonal cycle of chlorophyll-a (Sciare et al., 2009) and dimethylsulfide  
372 (DMS) concentrations peaking in summer have been reported on the island, in line with an  
373 enhanced biological activity (NGuyen et al., 1990;Sciare et al., 1999).  
374 While enhanced photochemistry and biological activity in summer might explain more  
375 frequent RGM events at Amsterdam Island between December and March, further field  
376 studies are needed to fully understand divalent mercury formation pathways.

377

### 378 **3.3.2 Particle-bound mercury**

379 PBM is associated with airborne particles – e.g., dust, soot, sea-salt aerosols or ice crystal –,  
380 or originates from the adsorption of reactive mercury onto atmospheric particles (Lu and  
381 Schroeder, 2004). Field and modeling studies (Rutter and Schauer, 2007;Amos et al.,  
382 2012;Steffen et al., 2014) suggested that the partitioning of mercury onto particles might be  
383 driven by air temperature and aerosol particle loadings.

384 PBM concentrations at Amsterdam Island followed a seasonal trend with significantly higher  
385 concentrations in winter (July to September) than in summer (December to February)  
386 ( $2.18 \pm 1.56 \text{ ng/m}^3$  vs.  $1.79 \pm 1.15 \text{ pg/m}^3$ , p-value = 0.027, Mann-Whitney test). Higher PBM  
387 concentrations were recorded during the strongest NW winds episodes (see Figure 9),  
388 suggesting an enhanced long-range transport of PBM from continental southern Africa during  
389 strong NW winds episodes. The higher number of PBM events in 2013 (see Figure 5c) is in  
390 good agreement with about twice as many continental induced radonic storms observed at  
391 Amsterdam Island in 2013 than in 2012.

392 PBM events occurred about 55% of the time between August and October in 2012 and 2013  
393 (see Figure 5c) and were significantly positively correlated with fire counts west of

394 Amsterdam Island ( $r = 0.56$ ,  $p$ -value = 0.005, Spearman test). This result is consistent with  
395 other observations of enhanced PBM concentrations during wildfires (Finley et al., 2009).  
396 **However, biomass fire counts reached a maximum between June and September while PBM**  
397 **events peaked later, between August and October.** The seasonality of aerosol optical depth  
398 (AOD) in the Southern Hemisphere Africa biomass burning region was extensively monitored  
399 by the Cimel sun-sky radiometer at the AERONET site in Mongu, Zambia. 1995 to 2009  
400 measurements highlighted that the monthly means of Level 2 direct sun-measured 500 nm  
401 AOD at Mongu reached a peak from August to October (Eck et al., 2013), in line with the  
402 PBM events peak observed at Amsterdam Island. This time lag between seasonal peaks in fire  
403 counts and emissions has already been pointed out by Swap et al. (2003), but its origin  
404 remains unclear. It has been attributed either to unusual synoptic conditions favoring eastward  
405 transport of pollution over measurement sites (Stein et al., 2003; Swap et al., 2003) or to  
406 undetected/highly emissive denser wooded vegetation burns at the end of the fire season  
407 (Edwards et al., 2006).

408

#### 409 **4 Conclusion**

410 Wind sector analysis, air mass back trajectories and the observation of radonic storms led to  
411 the important conclusion that despite the remoteness of Amsterdam Island the rapid export of  
412 air from the southern African continent during the biomass burning season contributes to  
413 GEM and PBM budgets on the island. Low GEM concentrations are associated with southerly  
414 polar and marine air masses from the remote southern Indian Ocean. This dataset provides a  
415 new insight into baseline concentrations of mercury species in the Southern Hemisphere mid-  
416 latitudes and **new measurement constraints on the mercury cycle**, opening the way for new  
417 avenues in future modeling studies. Our understanding of mercury cycle in the marine  
418 boundary layer over Amsterdam Island is still limited. It represents a real challenge given  
419 harsh weather conditions – with a very frequent drizzle most certainly scavenging oxidized  
420 species –, and technical and logistical limitations. Further studies involving wet deposition,  
421 simultaneous measurements of other trace gases, **and interference and calibration tests of the**  
422 **Tekran speciation unit** are needed to improve our understanding of deposition processes and  
423 oxidation mechanisms.

424

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436

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<b>Month</b>	<b>Mean (ng/m<sup>3</sup>)</b>	<b>Median (ng/m<sup>3</sup>)</b>	<b>Standard deviation (ng/m<sup>3</sup>)</b>	<b>Range (ng/m<sup>3</sup>)</b>	<b>n</b>
<i>2012</i>					
January	1.04	1.04	0.02	1.01-1.09	60
February	1.06	1.05	0.03	0.98-1.14	457
March	1.02	1.02	0.02	0.94-1.15	459
April	1.03	1.02	0.04	0.95-1.15	474
May	1.04	1.03	0.04	0.94-1.19	468
June	1.02	1.03	0.04	0.90-1.13	451
July	0.99	1.01	0.08	0.74-1.17	519
August	1.05	1.03	0.09	0.83-1.38	501
September	1.04	1.06	0.07	0.80-1.27	442
October	1.01	1.01	0.07	0.84-1.17	416
November	0.94	0.94	0.07	0.77-1.12	184
December	1.01	0.99	0.08	0.93-1.38	255
<i>2013</i>					
January	1.03	1.02	0.07	0.90-1.23	474
February	0.98	0.98	0.03	0.85-1.09	417
March	0.98	0.99	0.06	0.82-1.13	457
April	0.98	0.99	0.05	0.81-1.10	385
May	0.89	0.89	0.05	0.72-1.03	437
June	1.08	1.08	0.05	0.96-1.23	414
July	1.12	1.12	0.04	1.00-1.25	551
August	1.12	1.11	0.07	0.97-1.55	534
September	1.05	1.05	0.07	0.81-1.38	508
October	1.00	1.01	0.06	0.79-1.17	515
November	0.99	0.98	0.08	0.85-1.24	516
December	1.10	1.08	0.06	0.98-1.31	390
n: number of measurements					

Table 1: Summary of monthly GEM data at Amsterdam Island.

Figure 1: Location of Amsterdam Island (AMS), Cape Point (CPT), Troll (TRL), Dumont d'Urville (DDU), and Concordia (DMC) stations – projection centered over Amsterdam Island.

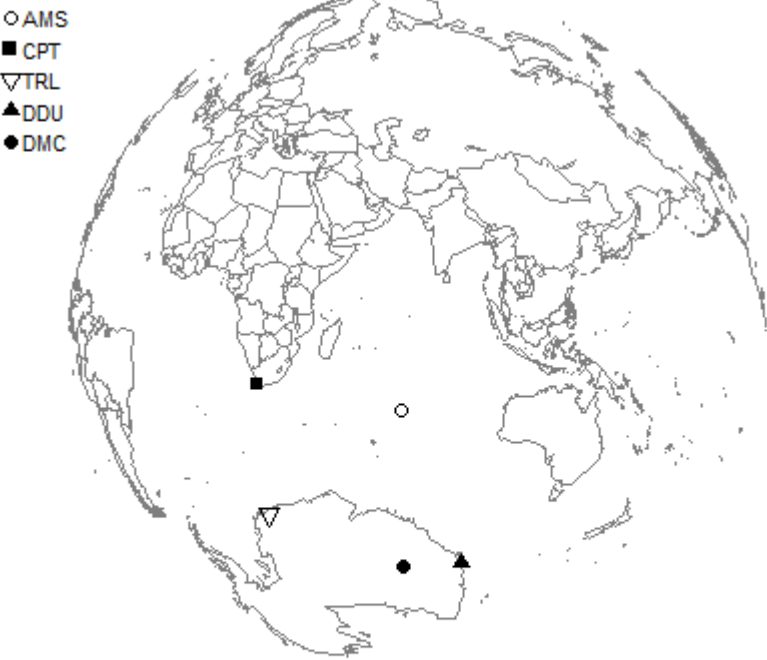


Figure 2: Wind direction and wind speed ( $\text{m}\cdot\text{s}^{-1}$ ) in winter (July to September) and summer (December to February) at the Pointe Bénédicte station. N: North, E: East, S: South, W: West.

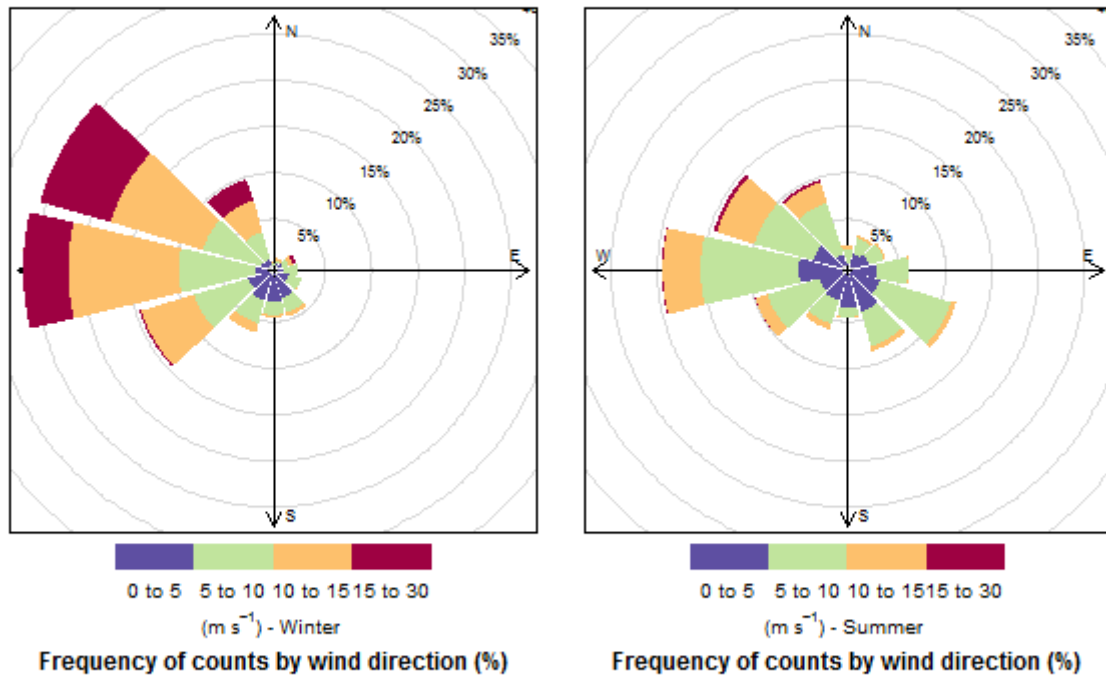


Figure 3: Particle-bound mercury (PBM, blue) and reactive gaseous mercury (RGM, green) concentrations ( $\text{pg}/\text{m}^3$ ), and hourly-average gaseous elemental mercury (GEM, red) concentrations ( $\text{ng}/\text{m}^3$ ) measured at Amsterdam Island from January 2012 to December 2013.

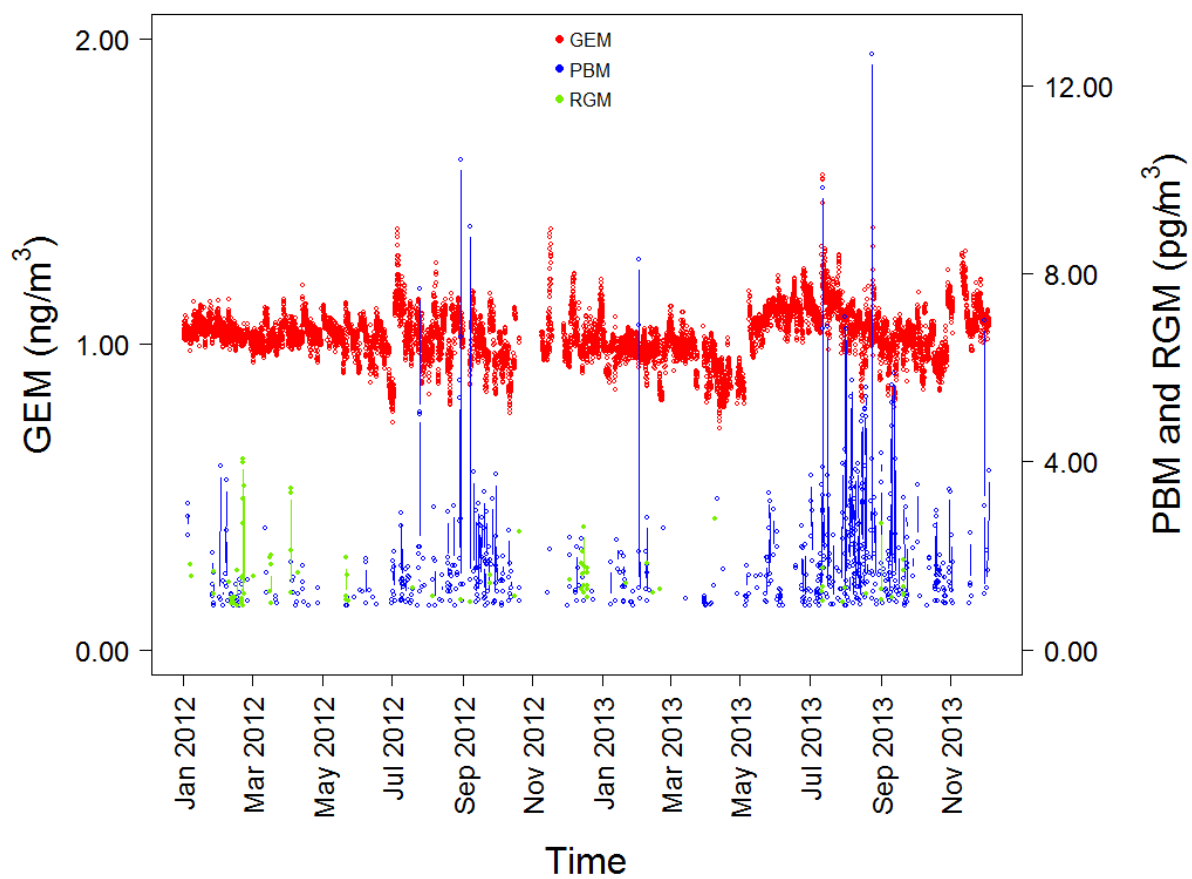


Figure 4: Hourly-average gaseous elemental mercury (GEM) concentrations ( $\text{ng}/\text{m}^3$ ) measured at Amsterdam Island: a) in winter (July-September) or summer (December-February), and b) under northwesterly (NW) or southerly (S) winds.  $\blacklozenge$ : mean, \*: statistically significant difference, n = number of hourly-average data, bottom and top of the box: first and third quartiles, band inside the box: median, ends of the whiskers: lowest (highest) datum still within 1.5 interquartile range of the lower (upper) quartile. Outliers are not represented

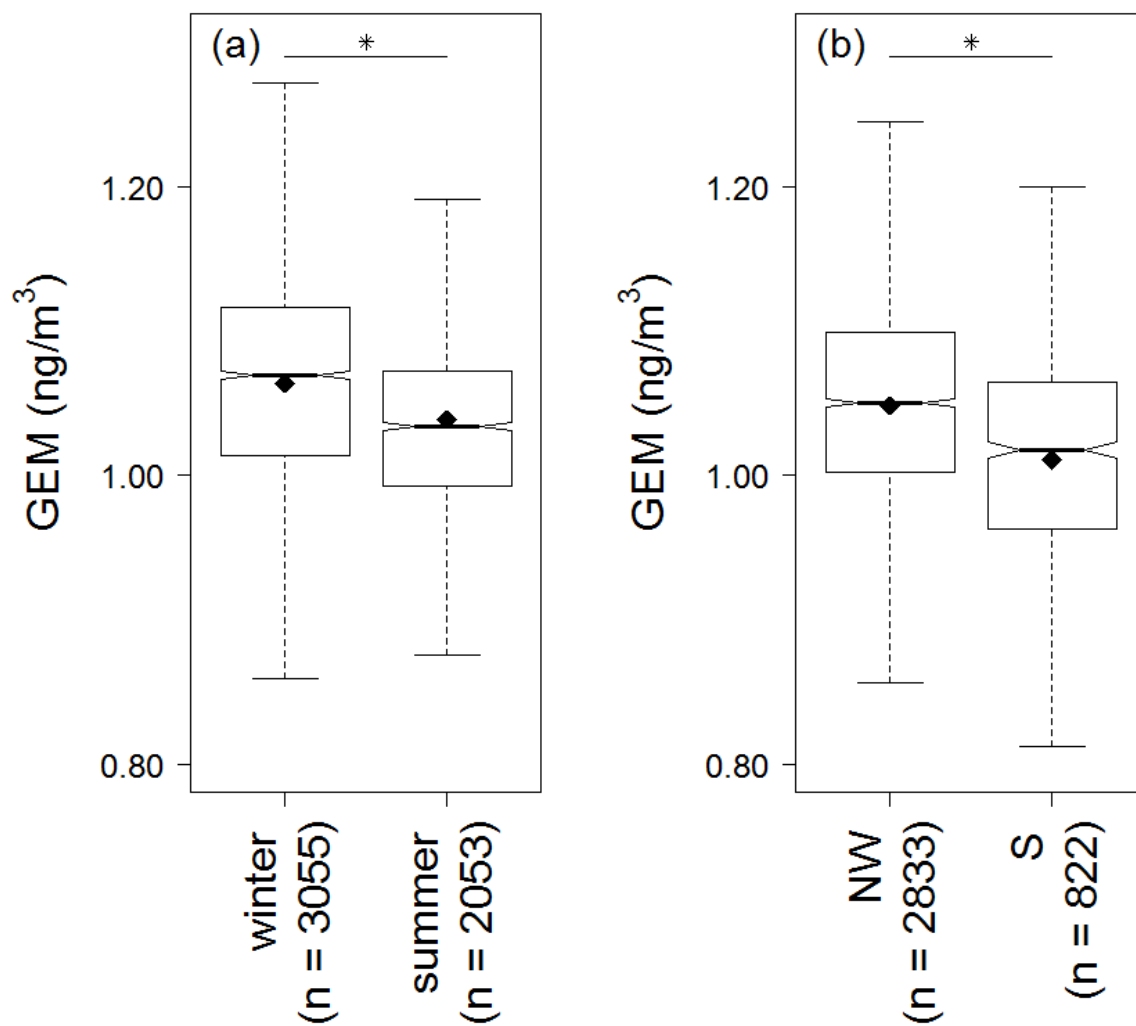
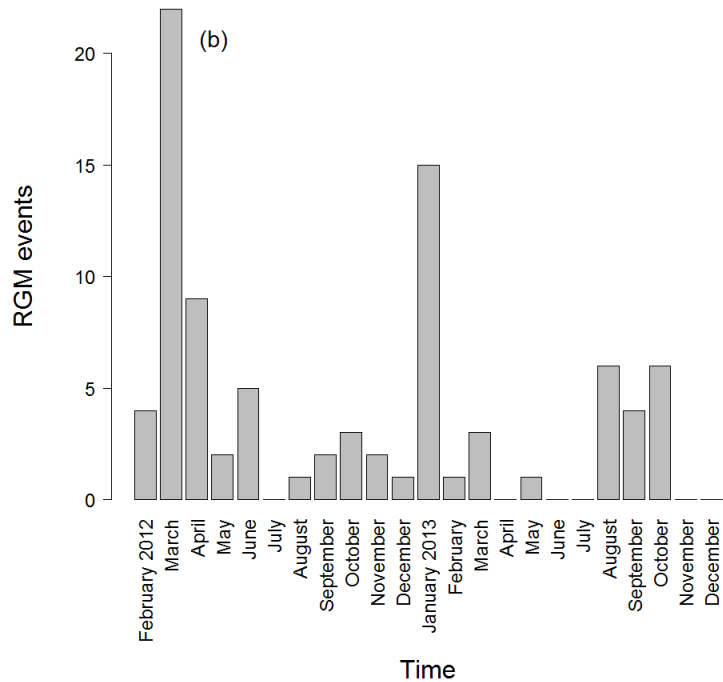
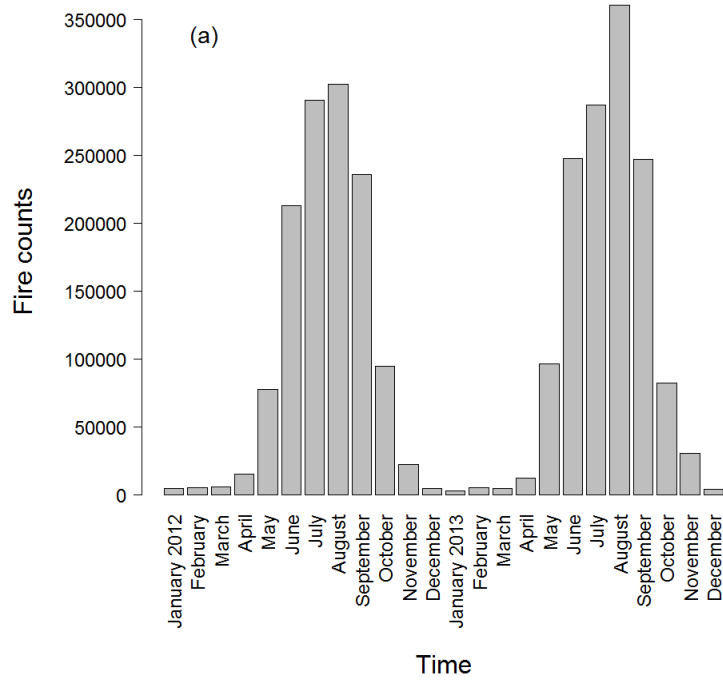




Figure 5: a) Fire counts west of Amsterdam Island (latitude: 3-53°S, longitude: 10-73°E) in 2012 and 2013. Data courtesy of FIRMS MODIS Fire Archive Download. b) Reactive gaseous mercury (RGM) events, i.e. number of measurements above quantification limit, at Amsterdam Island from February 2012 to December 2013. c) Particle-bound mercury (PBM) events, i.e. number of measurements above quantification limit, at Amsterdam Island from February 2012 to December 2013.



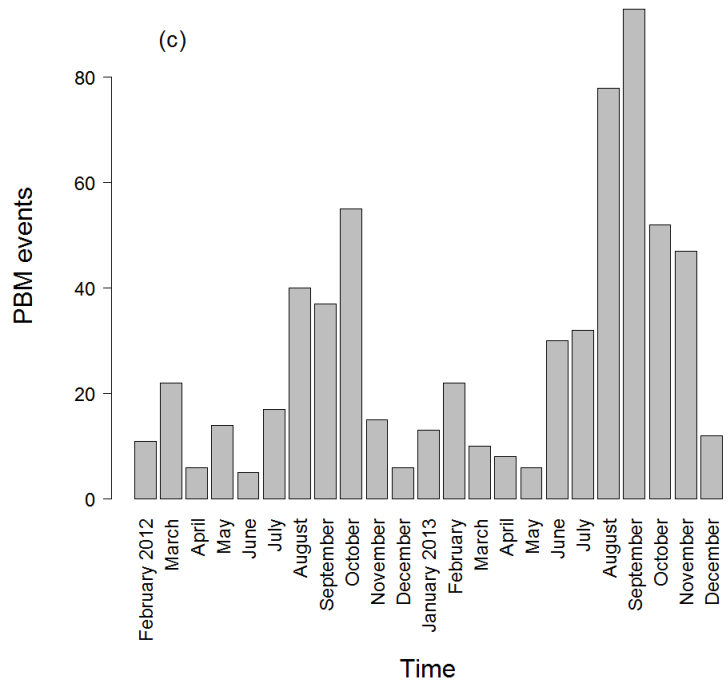


Figure 6: Examples of  $^{222}\text{Rn}$  ( $\text{mBq}/\text{m}^3$ ) peaks observed at Amsterdam Island along with  $^{220}\text{Rn}$  ( $\text{mBq}/\text{m}^3$ ) activity and wind speed ( $\text{m}/\text{s}$ ): a) in December 2012, c) in September 2013. Background variability ( $\text{dGEM}$ ,  $\text{ng}/\text{m}^3$ ) of GEM concentrations observed at Amsterdam Island in: b) December 2012, d) September 2013. Dotted lines represent 3 times the mean monthly standard deviation of  $\text{dGEM}$  measurements.

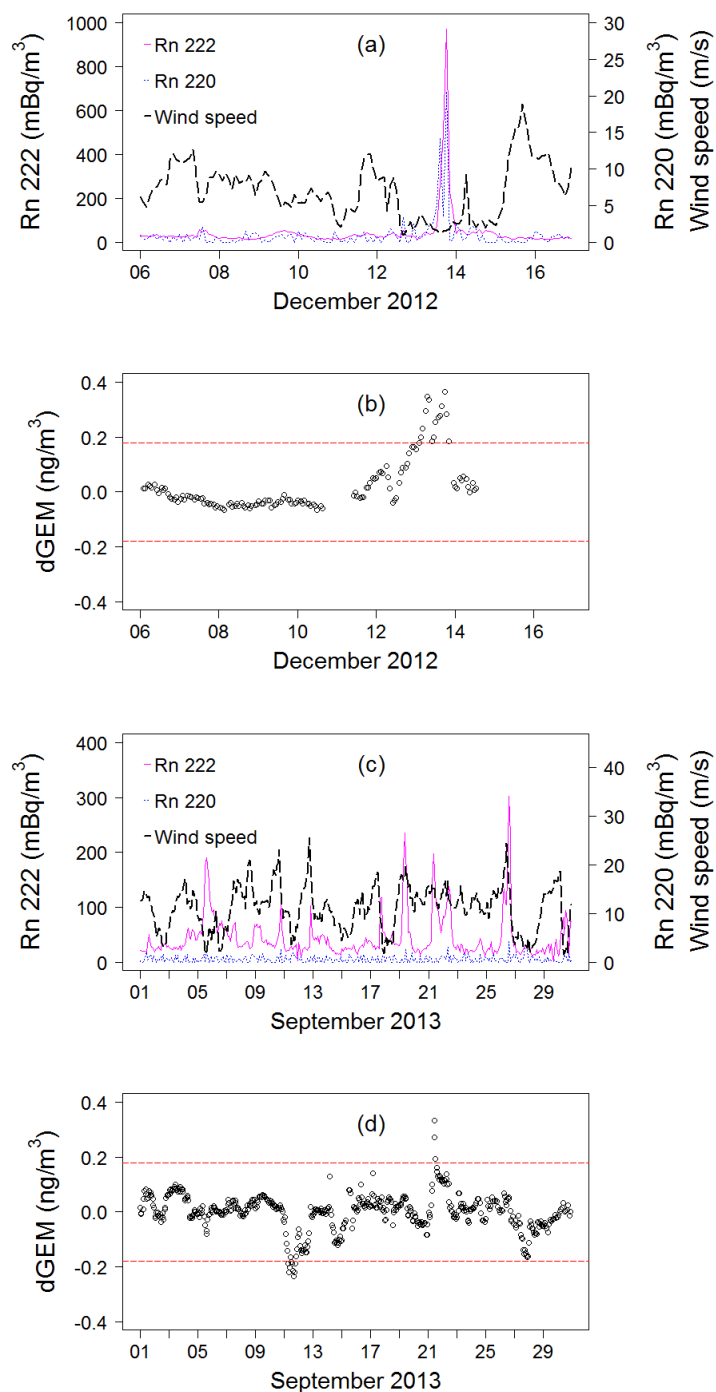


Figure 7: 7-days back trajectories ending on Amsterdam Island on December 13, 2012, September 11, 2013, September 21, 2013 and September 28, 2013. Data courtesy of NOAA.

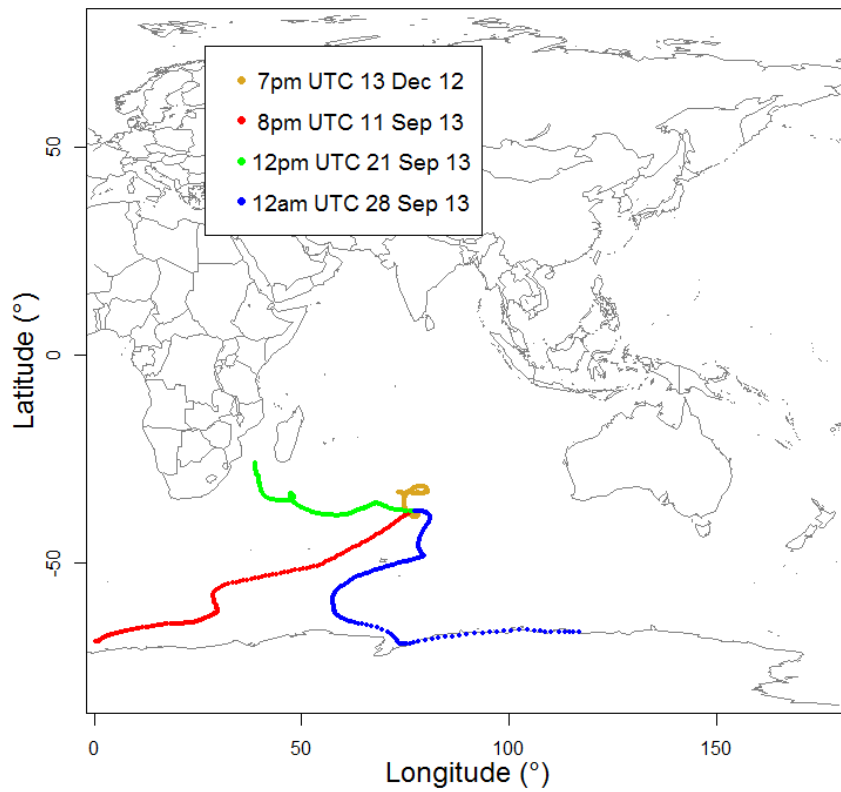


Figure 8: SeaWiFS chlorophyll-a map (January 2013) of the Indian sector of the Austral Ocean. The oceanic region located southwest of Amsterdam Island is highly productive in summer and potentially produces halogen species. Data courtesy of Giovanni online data system.

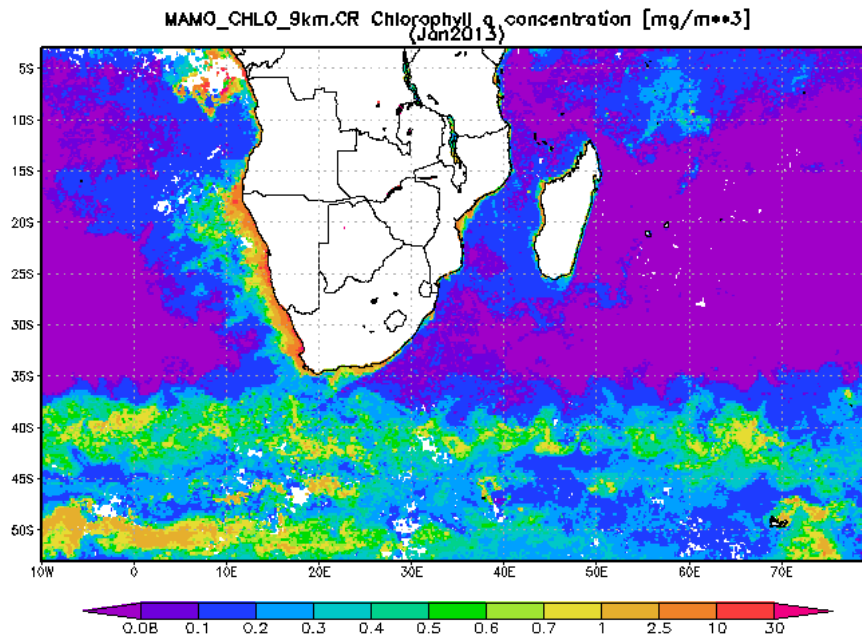


Figure 9: Particle-bound mercury (PBM) concentrations ( $\text{pg}/\text{m}^3$ ) according to wind speed ( $\text{m}\cdot\text{s}^{-1}$ ) and direction.

N: North, E: East, S: South, W: West.

