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## Satellite observations of cirrus clouds in the Northern Hemisphere

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# Satellite observations of cirrus clouds in the Northern Hemisphere lowermost stratosphere

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## Abstract

Here we present observations of the Cryogenic Infrared Spectrometers and Telescopes for the Atmosphere (CRISTA) of cirrus cloud and water vapour in August 1997 in the upper troposphere and lower stratosphere (UTLS) region. The observations indicate a considerable flux of moisture from the upper tropical troposphere into the extra-tropical lowermost stratosphere (LMS), resulting in the occurrence of high altitude optically thin cirrus clouds in the LMS.

The locations of the LMS cloud events observed by CRISTA are consistent with the tropopause height determined from coinciding radiosonde data. For a hemispheric analysis in tropopause relative coordinates an improved tropopause determination has been applied to the ECMWF temperature profiles. We found that a significant fraction of the cloud occurrences in the tropopause region are located in the LMS, even if a conservative overestimate of the cloud top height (CTH) determination by CRISTA of 500 m is assumed. The results show rather high occurrence frequencies ( $\sim 5\%$ ) up to high northern latitudes ( $70^\circ\text{N}$ ) and altitudes well above the tropopause ( $> 500\text{ m}$  at  $\sim 350\text{ K}$  and above) in large areas at mid and high latitudes.

Comparisons with model runs of the Chemical Lagrangian Model of the Stratosphere (CLaMS) over the CRISTA period show a reasonable consistency for the retrieved cloud pattern. For this purpose a limb ray tracing approach was applied through the 3-D model fields to obtain integrated measurement information through the atmosphere along the limb path of the instrument. The simplified cirrus scheme implemented in CLaMS seems to cause a systematic underestimation in the CTH occurrence frequencies in the LMS with respect to the observations. The observations together with the model results demonstrate the importance of isentropic, quasi-horizontal transport of water vapour from the sub-tropics and the potential for the occurrence of cirrus clouds in the lowermost stratosphere and tropopause region.

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## 1 Introduction

A large proportion of the uncertainties of climate change projections by general circulation models (GCMs) arises from poorly understood and represented interactions and feedbacks between dynamic, microphysical, and radiative processes affecting cirrus clouds (IPCC, 2014). Modelled climates are sensitive even to small changes in cirrus coverage or ice microphysics (Kärcher and Spichtinger, 2010). Fusina et al. (2007) point out that the net radiative impact strongly depends on ice water content and ice crystal number concentration. Small changes in the effective radius of the size distributions can substantially modify the surface temperatures.

Recent GCM studies (Sanderson et al., 2008; Mitchell et al., 2008) indicate that the climate impact of cirrus clouds depends in particular on the fall speed of ice particles, which in turn depends on ice nucleation rates (i.e. the concentrations of small ice crystals). The overall net warming effect of cirrus clouds can be substantially reduced by changing the concentrations of small ice crystals (i.e. the degree of bimodality) of the particle size distribution (PSD). The size distribution strongly affects the representative PSD ice fall speed. Mitchell and Finnegan (2009) investigated this sensitivity in more detail and concluded that cirrus clouds are a logical candidate for climate modification efforts.

The large uncertainties in climate prediction caused by processes involving cirrus clouds highlight the importance of more quantitative information on cirrus clouds by observations, especially for optically thin and small particle cirrus clouds like contrails close to the tropopause, which may have an overall cooling effect in contradiction to lower cirrus (Zhang et al., 2005). However, uncertainties for the climate feedback of cirrus clouds are still very large and a substantial reduction is needed (IPCC, 2014).

In particular, the altitude region of the Upper Troposphere and Lower Stratosphere (UTLS) plays an important role. Changes and variability of UTLS composition are major drivers of surface climate change. Even small changes of spatially highly variable

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concentrations of water vapour (H<sub>2</sub>O) have significant effects on the atmospheric radiation balance (e.g. Solomon et al., 2010; Riese et al., 2012).

Detailed understanding and modelling of the transport pathways of water vapour, and consequently the realistic representation of total water (gaseous and condensed form) in the UTLS region are therefore crucial for the correct representation of clouds and water vapour in climate models. Comprehensive analyses are published on transport processes from the troposphere into the stratosphere and tracer proportions of the extra-tropical UTLS region (e.g. Hegglin et al., 2009; Hoor et al., 2010; Ploeger et al., 2013). Rossby wave breaking (Ploeger et al., 2013) and mid-latitude overshoot convection (Dessler, 2009) result in transport and mixing of air masses into the extra-tropical UTLS on short time scales (weeks). On seasonal time scales, downwelling by the deep Brewer–Dobson circulation branch moistens the extra-tropical UTLS at altitudes above 450 K (Ploeger et al., 2013). Aged air masses transported into the extra-tropical lower stratosphere from above are moistened by methane oxidation and represent an important source for water vapour in the middle stratosphere (e.g., Jones and Pyle, 1984; Rohs et al., 2006).

The imprint of various water vapour transport ways into the lowermost stratosphere on cirrus formation has been investigated in a limited number of studies (Dessler, 2009; Montaux et al., 2010; Pan and Munchak, 2011; Wang and Dessler, 2012). The formation of cirrus clouds above the mid-latitude tropopause is discussed so far quite controversially. Dessler (2009) found relatively high occurrence rates of cirrus above the mid and high latitude tropopause in space borne lidar data of the Cloud and Aerosol Lidar (CALIOP) instrument on the Cloud Aerosol Lidar Infrared Pathfinder Satellite Observations (CALIPSO) (Winker et al., 2010). The analysis of Dessler (2009) shows cloud top height occurrences above the tropopause of up to 30–40 % for mid-high and tropical latitudes and still 0.1 % at 3 km or 40–50 K potential temperature above the tropopause. Pan and Munchak (2011) (in the following abbreviated as PM2011) showed that accurate tropopause definition and tropopause relative coordinates are important for this type of analysis and reach significantly different conclusions based on the same set of

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measurements. They find substantially fewer clouds above the tropopause and in their analysis the CALIOP data do not provide sufficient evidence of significant presence of cirrus clouds above the mid-latitude tropopause. The remaining but evidential events in the tropics show occurrences up to 24 % in the western Pacific and are usually located between the cold point and the lapse rate tropopause (up to 2.5 km above). PM2011 speculated that most of these clouds are triggered by gravity wave induced temperature disturbances, which typically are observed above deep convection areas (e.g. Hoffmann and Alexander, 2010).

In contrast, cloud observations by mid-latitude lidar stations show frequent events at and above the tropopause (e.g. Keckhut et al., 2005; Rolf, 2013). Many of them coincide with the observations of a secondary tropopause (Noël and Haeffelin, 2007). Isentropic transport and mixing of subtropical air masses with tropospheric high water values into the mid-latitude and polar LMS may cause such events and Montaux et al. (2010), in a case study, were able to reproduce the observation of such a cloud with an isentropic transport model by implementing a simple microphysical cloud model. However in this study, the cloud was observed just at the tropopause and not significantly above. Eixmann et al. (2010) investigated the dynamical link between poleward Rossby wave breaking (RWB) events and the occurrence of upper tropospheric cirrus clouds for lidar measurements above Kühlungsborn (54.1° N, 11.8° E). For three similar cirrus events they found a strong link between low values of potential vorticity (a proxy for RWB activity), enhanced up-draft velocities, and cloud ice water content. They concluded that based on the climatology of poleward RWB events following the method of Gabriel and Peters (2008) a parameterisation of the formation or occurrence of high and thin cirrus clouds seems to be possible.

Although there are a couple of ground based lidar observations suggesting the presence of cirrus clouds in the lowermost stratosphere (LMS), a region strongly influenced by isentropic (quasi-horizontal) transport of air masses from the tropics (Gettelman et al., 2011), there are open questions: which microphysical process and specific meteorological conditions foster the formation of ice particles in this specific region, how

frequently do these cirrus clouds occur on global scales, and are clouds tops or even complete clouds significantly above the tropopause.

The currently most sensitive sensor in space for cirrus cloud observations is the CALIOP lidar. Nonetheless, Davis et al. (2010) pointed out that the space lidar on CALIPSO might miss 2/3 of thin cirrus clouds with vertically optical depth  $\tau < 0.01$  in its current data products. Clouds with such a very low optical thicknesses have been observed by airborne lidars and in-situ instruments in the validation campaigns for CALIPSO. Frequently these cloud layers showed IWC values smaller than  $10^{-5} \text{ gm}^{-3}$  (Davis et al., 2010).

Here we argue that IR limb sounding from space provide an alternative measurement technique of high sensitivity for the detection of optically thin clouds (Mergenthaler et al., 1999; Spang et al., 2002; Massie et al., 2007; Griessbach et al., 2013), subvisible cirrus (SVC) defined by the extinction range  $2 \times 10^{-4} - 2 \times 10^{-2} \text{ km}^{-1}$  (Sassen et al., 1989), or the even thinner ultra-thin tropical cirrus (UTTC) (Peter et al., 2003; Luo et al., 2003). The detection sensitivity for clouds of IR limb sounders is in the range of spaceborne lidar measurements (Höpfner et al., 2009; Spang et al., 2012). A 100 km or even 1 km horizontally extended cirrus cloud is detectable by an IR limb sounder with an ice water content (IWC) of  $3 \times 10^{-6}$  and  $3 \times 10^{-4} \text{ gm}^{-3}$  respectively (Spang et al., 2012), presupposed that the cloud fills completely the field of view of the instrument. These values represent even better detection sensitivity than the current CALIOP cloud products.

In this paper we like to present new analyses of measurements from the Cryogenic Infrared Spectrometers and Telescopes for the Atmosphere (CRISTA) instrument during its 2nd Space Shuttle mission in August 1997 (CRISTA-2) (Grossmann et al., 2002). Due to its unique combination of moderate spectral resolution, high horizontal long track and cross track sampling, and good vertical resolution and sampling, the CRISTA measurements are a unique dataset for IR limb sounders in spite of information available from more modern satellite missions today. A reanalysis of the dataset can add complementary information especially for optically thin clouds compared to

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nadir passive and active instruments as well as limb sounders in the uv-vis and microwave wavelength region. The characterisation of frequent observations of Northern Hemisphere mid- and high-latitude cirrus clouds in respect to the tropopause (above or below) are in the focus of the present study.

The paper is organised as follows. First we introduce the CRISTA instrument and the applied data analysis methods followed by a section presenting the cloud top occurrence frequencies (COF) in respect to the tropopause. The comparison of CRISTA water vapour and cloud measurements presented in Sect. 4 is suggesting a strong influence of horizontal transport processes of high water vapour values from the subtropics to the latitude of cloud formation. A comparison with a global transport model can help to understand the origin and evolution with time of the cloud observations around the tropopause. This is investigated in Sect. 5 with a Lagrangian transport model containing a simple cirrus parameterisation.

## 2 Observations and analysis methods

### 2.1 CRISTA satellite instrument

The Cryogenic Infrared Spectrometers and Telescopes for the Atmosphere (CRISTA) instrument measured roughly one week in the UTLS during two space shuttle missions in November 1994 and August 1997 (Offermann et al., 1999; Grossmann et al., 2002). The measurements demonstrate the potential of the IR limb viewing technique to provide information on several trace gas constituents (Riese et al., 1999a, 2002) and clouds (Spang et al., 2002) with high spatial resolution. The spectral information in the wavelength ( $\lambda$ ) region 4–15  $\mu\text{m}$  is scanned with a resolution of  $\lambda/\Delta\lambda = \sim 500$ , which is equivalent to 1.6  $\text{cm}^{-1}$  at 830  $\text{cm}^{-1}$ . The vertical field of view (resolution) is in the order of 1.5 km and a typical vertical sampling of 2 km was used during CRISTA-2. A horizontal along-track sampling of 200 to 400 km was applied, depending on the measurement mode. An across-track sampling of  $\sim 600$  km was achieved by using three telescopes

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for three viewing directions simultaneously. A typical measurement net in the Northern Hemisphere is illustrated in Fig. 1. Due to the overlapping orbits and the  $57^\circ$  orbit inclination an even higher horizontal cross-track sampling becomes obvious at high northern latitudes ( $\sim 200$  km for north of  $60^\circ$  latitude). The spectrometers and optics were cryogenically cooled by helium to allow for measurements in the middle and far infra-red ( $4\text{--}70$   $\mu\text{m}$ ).

The instrument was hosted by the free-flyer system ASTRO-SPAS (Wattenbach and Moritz, 1997). The accuracy of the attitude system of the platform which was also used for two astronomic missions is excellent. The final pointing accuracy in the limb direction for the three viewing directions is in the order of 300 m (Riese et al., 1999a; Grossmann et al., 2002). The effect of refraction through the atmosphere in the limb direction is considered in the tangent height determination and can reduce the actual tangent height by up to  $\sim 300$  m at 12 km altitude. This correction is crucial for the cloud top height determination in the next section.

Here, we focus on the CRISTA-2 mission which took place from 8 to 15 August in 1997. Details on the instrument, the mission, and the specific water vapour retrieval are given in Grossmann et al. (2002); Offermann et al. (2002); and Schaeler et al. (2005) respectively. For the water vapour retrieval the continuous spectral scans of the spectrometers allow the selection of a spectral water vapour feature most suitable at tropopause altitudes (at a wavelength of  $12.7$   $\mu\text{m}$ ). An onion peeling retrieval is applied to the CRISTA measurements (Riese et al., 1999; Schaeler and Riese, 2001) and has the advantage of no upward propagation of errors to altitudes levels above optically thick clouds, even if the spectrum is not removed from the retrieval. The contamination by the strong cloud emissions and scattering processes in the corresponding IR spectra are too complex to model accurately in the retrieval process, and these measurements are not taken into account in the water vapour distributions presented later. The absolute accuracy of the water vapour retrieval is estimated to be 22 % (Offermann et al., 2002), though it is better at specific altitudes (10 % at 215 hPa). The precision is estimated to be 8–15 % (Schaeler et al., 2005).



## 2.2 CRISTA cloud detection

In the following special emphasis is put on cloud top height (CTH) observations at NH mid-latitudes in respect to the tropopause, where isentropic horizontal transport of water vapour from the subtropics to high latitudes may trigger cirrus formation in LMS.

5 The cloud detection for IR limb sounders has been investigated in detail over the last decade (Spang et al., 2012, and references therein). For spectrally resolved measurements simple colour ratio based methods are shown to be robust and accurate for the detection of cloudy spectra (e.g. Spang et al., 2001; Sembhi et al., 2012). For the following analyses the cloud index (CI) is defined by the colour ratio of the mean radiances from 788 to 796  $\text{cm}^{-1}$  divided by the mean radiances from 832 to 834  $\text{cm}^{-1}$ , which was already applied to various airborne and spaceborne limb IR instruments (e.g. Spang et al., 2002, 2004, 2007). The corresponding CTH is the first tangent height where CI falls below the defined threshold value ( $\text{CI}_{\text{thres}}$ ). The left panel of Fig. 2 illustrates a CI profile with the transition from clear sky conditions ( $8 > \text{CI} > 4.5$ ) to cloudy conditions ( $\sim 4 > \text{CI} > 1.1$ ), where optically thick conditions are in line with a CI-value of  $\sim 1.2$ .

10 Various studies have shown that the detection sensitivity is linked to the detection threshold and depends to some extent on the seasonal variation in the trace gas concentrations in the applied spectral windows (e.g. Spang et al., 2012). The main limiting effect of the cloud index method are high water vapour continuum emissions (for mixing ratios  $> 500$  ppmv) in the mid troposphere and below. Under such conditions a definite discrimination between clouds and high water becomes difficult (Spang et al., 2004, 2007).

20 Cloud top heights detected with a cloud index threshold value of  $\text{CI}_{\text{thres}} = 3$  are presented in Fig. 1. Inhomogeneities in the measurement net are caused by special measurement modes, where the free-flyer was pointed to specific regions of interest (e.g. the warm pool region over Micronesia or validation sites). The dense horizontal coverage of the measurement net over 24 h (here presented for 9 to 10 August 1997, noon to noon) allows to track the horizontal transport patterns of trace gases for example

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streamers (e.g. Riese et al., 1999a, 2002) in conjunction with meteorological parameters like the potential vorticity (PV). Figure 1 shows PV contours for the 350 K isentrope. All meteorological data used in this study are from the ERA Interim reanalysis dataset provided by the European centre for medium-range weather forecasts (ECMWF) (Dee al., 2011). The CTH distribution and PV contours are suggesting a link between the dynamical features of horizontal transport processes like Rossby wave breaking events (e.g. Juckes and McIntyre, 1987) and the presence of high cirrus clouds. Contours of low PV (4 and 8 PVU, with  $1 \text{ PVU} = 10^{-6} \text{ Km}^2 \text{ kg}^{-1} \text{ s}^{-1}$ ) are highlighting elongated air masses from the subtropics to mid and high latitudes where coincidentally high altitude clouds appear in the CRISTA observations.

High CTHs ( $> 12 \text{ km}$ ) are frequently present at mid ( $40\text{--}60^\circ \text{ N}$ ) and even at high geographic latitudes ( $> 60^\circ \text{ N}$ ) in regions of low PV. Nearly all of these high CTH locations show PV values greater 2 PVU, a common threshold for the dynamical tropopause in the subtropics (Holton et al., 1995), and marked in Fig. 2 by black circles. Whether these clouds are really formed in the LMS and not just below the tropopause is matter of particular interest in this study. In addition, for exploring clouds in the vicinity of the tropopause it is important to quantify the uncertainties in cloud top and tropopause height as good as possible. The following sections will present more details on the analysis method and error estimates of both parameters.

### 2.3 Uncertainties in cloud top height determination

The retrieved CTH from the CRISTA radiance profiles depends not only on the CI threshold values but also critically on the vertical sampling of the instrument (typically 2 km during CRISTA-2, see Fig. 2) and the vertical size of the field of view (FOV) of the instrument. The FOV of CRISTA is very well described by a Gaussian function with a full width half maximum (FWHM) of 2.624 arcmin, which is equivalent to  $\sim 1.5 \text{ km}$  at 15 km tangent height (Offermann et al., 1999) and corresponds to a standard deviation of  $\sigma_{\text{FOV}} = 625 \text{ m}$ .

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Uncertainties in the CTH are dominated by effects of the vertical FOV. If an optically thick cloud is only filling the lower part of the FOV, the attributed CTH may overestimate the real CTH. This potential error source was investigated in detail by modelled cloud index profiles for various cloud conditions and background atmospheres (Fig. 3).

We calculated radiance profiles of the analysed wavelength regions to compare the cloud index for varying cloud altitude (6–20 km) and layer thickness (0.5 and 2 km), and extinction  $\varepsilon$ . Simulations were carried out with the line by line radiative transfer code RFM (Dudhia et al., 2002). Scattering processes were neglected. For a full set of simulations a realistic parameter space was chosen, i.e. CTHs between 6 and 16 km, different reference gas atmospheres, cloud extinctions from  $10^{+1}$  to  $10^{-4}$  km<sup>-1</sup> at a wavelength of 12  $\mu$ m, and box type cloud layers with vertical extension of 0.5, 1, and 2 km. The calculations were performed on a 100 m vertical grid. The pencil beam radiance profiles were afterwards convolved with the FOV. By comparing the input CTH and the simulated CTH we estimate the maximum error in CTH for multiple CI thresholds (grey vertical lines in Fig. 3). Four different threshold values (1.2, 2, 3, 4) have been investigated for the various cloud layer extinctions and cloud vertical thickness. A detection threshold of 4 and 4.5 was applied for the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) instrument (Fischer et al., 2008) in various studies for the detection of the usually optically thin PSCs (e.g. Spang et al., 2003; Höpfner et al., 2005). Sembhi et al. (2012) showed that  $CI_{\text{thres}}$  values up to 6 are acceptable for the MIPAS measurements depending on the latitude and altitude region of interest.

Figure 3 shows the CI profiles of the pencil beam simulations and the CI profile with the FOV convolution for a homogeneous cloud layer for an optically thin ( $\varepsilon = 3 \times 10^{-3}$  km<sup>-1</sup>) cloud between 10 to 12 km (CTH = 12 km) and an optically thick ( $\varepsilon = 10^{-1}$  km<sup>-1</sup>) cloud between 11–13 km (CTH = 13 km). The conservative threshold  $CI_{\text{thres}} = 1.2$  detects only optically thick clouds and consequently the CTH is underestimated due to the FOV effect. Optically thinner clouds are only detectable with a less stringent threshold ( $CI > 1.8$ , Spang et al., 2005a). The analysis for  $CI_{\text{thres}} = 2$

shows a maximum possible CTH error ( $\Delta_{\max 2}$ ) of 0.6 km for all simulations. Higher thresholds result in higher detection sensitivity but cause higher uncertainties in CTH ( $\Delta_{\max 3} = 0.9$  km for  $Cl_{\text{thres}} = 3$  and  $\Delta_{\max 4} = 1.4$  km for  $Cl_{\text{thres}} = 4$ ) for optical thick clouds with  $\varepsilon > 10^{-1} \text{ km}^{-1}$ . Optically thinner clouds usually cause smaller maximum CTH errors.

However, the optically thin and thick example in Fig. 3 show that higher detection sensitivity will cause an overestimation of the CTH, in the examples for 0.8 and 1.25 km for  $Cl_{\text{thres}} = 4$ . In addition it should be noted that the actual CTH error of a measurement depends not only on  $Cl_{\text{thres}}$ , but also on the relative distance of the measured tangent heights to the actual “real” cloud top (sampling). For CRISTA measurements the distances of the tangent point to the “real” CTH in the atmosphere are statistically almost equally distributed with a maximum distance of  $\pm 1$  km due to the 2 km vertical sampling. The  $\Delta_{\max}$  values above are errors of a worst case scenario and represent the upper extreme of the FOV effect. Half of all detections of optically thick clouds will have a  $\Delta_{\text{CTH}}$  smaller than  $\Delta_{\max}/2$ , due to the fact that the FOV error for optical thick clouds declines linearly with declining difference between the sampled observation height and the “real” CTH (see Fig. 3).

In conclusion, the mean CTH error due to the FOV (for optical thick clouds)  $\Delta_{\text{FOV}}$  is consistent with  $\Delta_{\max}/2$ . For optically thinner clouds  $\Delta_{\max}$  becomes systematically smaller (e.g. compare Fig. 3a for thin and Fig. 3b for thick conditions) and for the more stringent  $Cl_{\text{thres}} = 2$  these clouds are only detectable at an observations height below the actually “true” CTH (negative CTH errors in Fig. 3a). In the following analyses we used two thresholds,  $Cl_{\text{thres}} = 2$  and  $Cl_{\text{thres}} = 3$ , for optimisation and best trade-off between quantified FOV effects and best detection sensitivity for optically thin clouds. The chosen  $Cl$  values are equivalent to extinction coefficients of  $\varepsilon > \sim 5 \times 10^{-3} \text{ km}^{-1}$  and  $\varepsilon > \sim 2 \times 10^{-3} \text{ km}^{-1}$  respectively, and correspond to  $\Delta_{\text{FOV}2} = 250$  m and  $\Delta_{\text{FOV}3} = 450$  m.

It is certainly a strong overstatement to postulate optical thick conditions for all potential cirrus cloud detections around the tropopause. How frequently cirrus clouds in the limb appear as optical thick or thin, equivalent to large and small positive biases,

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is unknown and therefore difficult to quantify. The slanted ice water path (IWP), the IWC integrated along the line of sight (Spang et al., 2012, and also Sect. 5.4), determines the actual optically thickness of the measured spectrum in the limb direction. Radiative transport model calculations for a characteristic diversity of cirrus size distribution parameter (Griessbach et al., 2014) under the assumption of vertically and horizontally homogeneous cloud layers show that the largest FOV induced CTH errors  $\Delta_{\max}$  (for  $\varepsilon > 10^{-1} \text{ km}^{-1}$ ) can be generated from a cloud layer with IWC  $> \sim 3 \text{ ppmv}$  ( $> 0.5 \text{ mg m}^{-3}$ ) (Spang et al., 2007; Fig. 5). This is a common value in IWC in-situ measurements in the upper troposphere, typically close and or slightly greater than the typical median values (Krämer et al., 2009) in the corresponding temperature range of the CRISTA measurements. Consequently, it is very likely that the CRISTA statistics also include a certain amount of underestimated CTHs by optically very thin clouds (IWC  $\ll 3 \text{ ppmv}$ ).

Uncertainties in CTH determination from broken cloud segments along the line of sight in combination with the horizontal integration of the limb information and from the cross track extension of the FOV (30 arcmin,  $\sim 15 \text{ km}$ ) are not considered in the present analysis. However, both effects cause a reduction in detection sensitivity and results in an underestimation of the CTH in respect to the true CTH. Consequently falsified detections above the local tropopause can be excluded by these two effects.

### 2.4 Radiosonde data

In this study we used the radiosonde station composite data from the University of Wyoming, Department of Atmospheric Science. For August 1997, a time period embedding the complete CRISTA-2 mission, around 5000 radiosonde launches were available for coincident comparison of tropopause location with respect to CTHs from CRISTA and for the global validation of an improved tropopause determination with the ERA Interim dataset (Sect. 2.5) at mid and high northern latitudes. Figure 2 illustrates a coincident radiosonde temperature, relative humidity (RH) and ice saturation temperature profile in comparison to the CRISTA cloud index profile. The cold point tropopause is

clearly visible in the blue temperature profile. The horizontal lines are indicating the retrieved CTH from CI and lapse rate tropopause based on coincident ERA Interim temperatures (details see Sect. 2.5). Usually radiosonde RH measurements around the tropopause have a low accuracy and act more like a qualitative measure of humidity.

5 Nevertheless, RH and the ice saturation temperature profile indicate atmospheric conditions around the tropopause that allow for the existence of ice. This is in coincidence with the slightly higher cloud layer detected by CRISTA, where CTH and tropopause height (TPH) are suggesting a cloud above the local tropopause.

## 2.5 Improved determination of the lapse rate tropopause for ERA Interim

10 Accurate tropopause height determination is crucial for the location of cloud events with respect to the tropopause as PM2011 already showed for the CALIPSO cloud detection. For our analyses we used the ERA Interim reanalysis data on hybrid coordinates (Dee et al., 2011) with original model resolution (60 levels and 600–1000 m resolution around the tropopause) for the computation of the tropopause height. A three step  
15 approach is applied to the data. (1) For each CRISTA tangent height the surrounding four ERA temperature and geopotential height profiles are determined. Then the lapse rate tropopause was defined for each profile as the lowest pressure (altitude) level at which the lapse rate is  $2 \text{ K km}^{-1}$  or less. The lapse rate should not exceed this threshold for the next higher levels within 2 km (WMO, 1957). The vertical resolution of the  
20 retrieved TPH cannot be better than the vertical grid resolution of the temperature data and hence, can produce a significant positive bias for analyses with tropopause related altitude coordinates (PM2011). In step (2) we applied a vertical spline interpolation with 30 m vertical resolution to the temperature profile around the actual TPH of step 1 and repeated the TPH computation with the artificially higher vertical resolution. Finally, in  
25 step (3) the weighted mean with distance of the four surrounding grid points of the observation point represents now the so-called high resolution tropopause height ( $\text{TPH}_{\text{hr}}$ ). By this approach of tropopause determination a more realistic lapse rate tropopause

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was found, because the single TPH values are not attached to the altitude grid points of the analysis data anymore.

Comparisons with radiosonde data set for August 1997 show a good correspondence between the  $TPH_{hr}$  and the lapse rate tropopause of the radiosonde for the coincidences with CRISTA profiles (see also Fig. 2). A statistical analysis of the difference in TPH between 5304 sonde profiles in the latitude range  $30^{\circ}$  N and  $80^{\circ}$  N and the ERA Interim based  $TPH_{hr}$  for August 1997 showed a mean difference of 0.015 km and a standard deviation of 650 m. The latter value is exactly the same value as found in the analysis of PM2011 for a comparison between radiosonde and National Center of Environmental Prediction Global Forecast System (GFS) data for a June-July-August season in 2007. The good correspondence gives us confidence that the described  $TPH_{hr}$  is the best possible and reliable approach for a tropopause determination for each CRISTA profile. In addition, the standard deviation of the differences appears to be a realistic estimate for the uncertainty of  $TPH_{hr}$ .

### 3 CRISTA cloud top height occurrence with respect to the tropopause

#### 3.1 Tropopause derived from radiosonde data

A comparison of tropopause heights determined from radiosonde measurements with the detected CTHs in the tropopause region allows a first quantitative assessment of the indication for frequent observation of optically thin clouds by CRISTA in the LMS like illustrated in Fig. 1. A miss time of 2 h and miss distance of 200 km coincidence criteria were chosen to minimise uncertainties in the comparison. For 158 CTH detections close to the tropopause and north of  $40^{\circ}$  N (with  $Cl_{thres} = 3$ ,  $CTH-TPH > -500$  m, and TPH defined from co-located ERA Interim temperatures) we found 188 coincident radiosonde profiles. The tropopause of 62 radiosonde profiles ( $\sim 33\%$ ) is pinpointed more than 500 m below the coincident CTH of CRISTA. This indicates a remarkable fraction of clouds well above the tropopause and emphasizes the question if these high

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altitude clouds in the LMS are a common feature or a rare incidence. However, the only limited number of good coincidences between CRISTA and radiosonde measurements with respect to adequate miss time and distance criteria allows no profound conclusions on this question. A statistical analysis of all cloud observed in the tropopause region together with the collocated tropopause heights based on ERA Interim may solve this problem.

### 3.2 Statistical analysis with tropopause derived from ERA Interim

For the statistical analysis of cloud occurrences around the tropopause we choose a vertical coordinate independent of the temporal and spatial location of the tropopause. Therefore thermal tropopause relative coordinates are applied in the following similar to PM2011. These coordinates are used extensively in chemical tracer analyses (e.g. Tuck et al., 1997; Pan et al., 2007; Kunz et al., 2013) and temperature profile analyses (e.g. Birner et al., 2006). The actual distance of a detected CTH to the tropopause is defined by  $\Delta_{\text{CTH}} = \text{CTH}_i - \text{TPH}_i$  where the index  $i$  is attributed to an individual measurement profile. For displaying results it is often more helpful to adjust the reference altitude to the mean tropopause height ( $\text{TPH}_{\text{mean}}$ ). The tropopause related altitude  $Z_r$  is then defined by:

$$Z_r = \bar{Z}_{\text{trop}} + (Z - Z_{\text{trop}}), \quad (1)$$

with  $Z$  the observations altitude or the CTH,  $Z_{\text{trop}}$  the individual tropopause, and  $\bar{Z}_{\text{trop}}$  for example a daily or monthly zonal mean tropopause. Here we use for  $\bar{Z}_{\text{trop}}$  the zonal mean tropopause during the CRISTA-2 measurement period.

Figure 4 presents the results of CTH height occurrence frequencies (COF) in respect to the single profile tropopause location defined by the approach above and a vertical grid size of 0.5 km. Latitude bands covering the whole Northern Hemisphere observations by CRISTA have been defined for the tropical (0–20° N), subtropical (20–40° N), mid (40–60° N) and high latitudes (60–75° N). The vertical distributions are very

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different for the various latitude bands. A joint feature is the location of the maximum in COF in a layer 0.75 to 1.25 km below the tropopause for almost all latitude bands. In parts this is caused by the limb geometry. The probability to detect a cloud along the line of sight located above the actual tangent height is enhanced with penetrating deeper into the troposphere. This effect causes artificially overestimated COFs below the real maximum in the COF distribution and is especially a large bias in the tropics, where horizontally small extended and patchy distributed cloud systems (< 50 km), e.g. by deep convection events, are generating unrealistic high COFs in limb observations (Kent et al., 1997; Spang et al., 2012). Note that CTHs assigned to such observations are actually low biased.

Only the subtropics show the COF maximum at slightly lower altitudes. But more remarkable, generally very low COF values are found around the tropopause compared to the three other latitude bands. This local minimum at 20–40° latitude in cloud probability has been already observed in various limb sounder observations around the tropopause (e.g. Wang et al., 1996; Spang et al., 2002, 2012).

Figure 4 presents the COF values for both CI threshold values. The frequencies for  $CI_{\text{thres}} = 3$  are systematically higher than for the less sensitive threshold  $CI_{\text{thres}} = 2$ , and more clouds are detected well above the predefined tropopause and also above the maximum in the COF distribution. For the more sensitive detection method  $CI_{\text{thres}} = 3$  a COF of 3 and 4 % (217 events in total) is found in the altitude grid box 500–1000 m for mid- and high-latitudes respectively. Even in the 1000–1500 m grid box above the tropopause COFs of ~ 1 % are observed in both latitude bands (59 events). These COF values above the tropopause indicate significantly larger occurrence rates than found in ground based lidar observations (e.g. Noel and Haefelin, 2007; Rolf, 2013). The typically observed frequencies of 4–10 % of cross tropopause cirrus were referred to the total number of lidar profiles with cirrus clouds and not to the total number of observations (cloudy and non-cloudy) such for the satellite data. An equivalent approach would result in significantly smaller COF values for the lidar measurements.

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In addition, the tropics show a pronounced local maximum well above the tropopause. The 100 m running mean statistic of 500 m grid boxes (red symbols) indicates that the feature is not a sampling artefact caused by an interplay of the measurement altitudes with the vertical gridding of the COF analysis. This interesting feature will be studied in more detail in a future study.

### 3.3 Significance tests of cloud top occurrence distribution

We have investigated in detail how the measurement uncertainties of the tangent point altitude, tropopause height, and cloud top height (Sect. 2) might influence or even falsify the COF statistics. Where TPH and tangent altitude errors are well described by Gaussian distribution (with corresponding standard deviations), the CTH error ( $\Delta_{\text{CTH}}$ ) caused by the vertical FOV effect for optically thick clouds acts like a positive bias in the cloud top height determination.

Monte Carlo (MC) simulations with the CRISTA measurement ensemble of cloudy and non-cloudy profiles have been performed taking into account (a) statistical and (b) systematic error sources. For both types of simulations all CTHs above the tropopause were excluded from the dataset and the remaining CTH observations have been modified by a randomly distributed statistical uncertainty or a systematic positive offset value with a Gaussian amplitude. The results show that a statistical (noise) errors like the TPH uncertainty with  $\sigma_{\text{TPH}} = 650$  m or even an overestimated value of 1000 m cannot reproduce the measured vertical COF distribution and cannot create the relative large COFs observed above the tropopause.

For testing the systematic errors we applied to all CTH observations below the tropopause a FOV-like, Gaussian-shaped, and only positive offset distribution. This approach is equivalent to the assumption that all clouds below the tropopause are optically thick (upper limit), are creating a positive offset, and are detected with the randomly distributed observation heights of CRISTA. The results for  $|\sigma_{\text{FOV}}| = 750$  m, a larger value than the real CRISTA  $\sigma_{\text{FOV}} = 625$  m, show comparable COF distributions for mid and high latitudes and can create similar COF values above tropopause

to the original statistic. Larger  $|\sigma_{\text{FOV}}|$  produces overestimated COF values well above the tropopause ( $> 1.5$  km), significant underestimates below the tropopause, and can be excluded. Surprisingly, the MC simulations were not able to reproduce the tropical COF distribution with the positive bias approach. These two negative tests, large  $|\sigma_{\text{FOV}}|$  and irreproducible tropical COF distribution, are additional constraints for a systematic and large FOV effect above the tropopause. In addition it is very unlikely that all clouds around the tropopause are optically thick (only  $< 50\%$ , see Sect. 2.3). Consequently it is very unlikely that this type of clouds is responsible for an artificial enhancement in COF above the tropopause like observed by CRISTA in Fig. 4.

An additional and substantial argument against a large systematic FOV effect is the different behaviour between tropical and mid/high latitudes in the measurements compared to the simulations. A positive bias for optical thick clouds should modify the pronounced maximum peak 1 km below tropical tropopause in the following way and is confirmed in the MC simulations: the observations would show a broader distribution and a significant extent of enhanced COF values in direction to higher altitudes similar to the mid and high latitudes. But this behaviour is not observed in the tropical measurements (Fig. 4a). A shift to higher altitudes can be reproduced in the error simulations with  $|\sigma_{\text{FOV}}| = 750$  m, but contrariwise it is creating strong enhanced COFs just below tropopause and reduced values below. The overall effect in the simulations, a positive shift of the whole distribution, is not observed in the original data.

In conclusion, taken all uncertainties into account the CRISTA COF distribution indicates a significant amount of cirrus cloud observations in the lowermost stratosphere. To our knowledge this is the first time such findings are reported for space-borne limb measurements.

### 3.4 Comparison with CALIPSO

For a brief comparison with the CALIOP lidar on CALIPSO results from PM2011 are superimposed in Fig. 4 for tropical and mid-latitude observations. For mid-latitudes CRISTA and CALIOP show very similar results at the tropopause and although time

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period and observations geometry are very different (7 day mean vs. multi annual seasonal mean and limb vs. nadir). In this region even the absolute COF values are in close agreement. However, in the grid boxes 500–1000 m and 1000–1500 m above the tropopause CRISTA detected  $\sim 2$  times more clouds than the CALIPSO climatology.

Below the tropopause the limb sounder statistic shows a substantial larger maximum in COF. This is primarily due to the long limb path integration and the artificially enhanced number of cloud detections well below the tropopause by higher altitude cloud fragments along the line of sight, and only secondary due to the better detection sensitivity for horizontally extended clouds in the limb compared to the short nadir viewing direction.

The tropical distribution looks more different, especially below the tropopause, where again the overestimation due the limb geometry plays a role. At the tropopause ( $\pm 500$  m), where this effect is negligible, both instruments show similar COFs. Well above the tropopause (500 to 1500 m) CRISTA COFs are again significantly larger than CALIPSO and indicate the presence of optically very thin clouds, which are currently not detected in the CALIPSO data products (see also Sect. 1). It should be noted that the CALIOP instrument may have observed these clouds, but the current detection threshold of the operational data products is not capable to detect ultra thin cirrus clouds. This was already shown by Davis et al. (2011) in a validation study with airborne lidar and in situ particle measurements. Modified detection schemes with larger horizontal averaging of the high resolution Level 1 profile data of CALIOP (e.g. 30 or 50 km instead of currently 5 km) will substantially improve the detection sensitivity for ultra thin cirrus clouds (M. Vaughan, personal communication, 2014). The current detection limit for cirrus clouds averaged horizontally to 5 km for the 532 nm extinction channel is in the range of 0.005 and  $0.02 \text{ km}^{-1}$ , which represents an equivalent IWC of 0.1 to  $4 \text{ mg m}^{-3}$  (Avery et al., 2012). However, the IR limb sounder detection limit for IWC is depending on the horizontal extent of the cloud considerably better. For a 100 km or 1 km horizontal extended cloud along the line of sight an IWC detection threshold of 0.003 and  $0.3 \text{ mg m}^{-3}$  is achievable (Spang et al., 2012).

## 4 Horizontal distribution of water vapour and clouds in the UTLs

### 4.1 Cloud top height distributions

Four days during CRISTA-2 have a measurement net dense enough for hemispheric analyses of horizontal structures in cloud and trace gas distributions. Figure 5 presents the daily cloud top height distribution detected with the algorithms described in Sect. 2.2 for 10 to 13 August at midnight  $\pm 12$  h. Potential temperature ( $\Theta$ ) has been used as vertical coordinate and CTHs in km are converted to  $\Theta$  by coincident ERA Interim temperature and geopotential height information. Only clouds detected in the altitude range 330 to 370 K are presented, an atmospheric layer usually located in the lower half of the lowermost stratosphere (LMS) at mid and high latitudes, and clearly in the upper troposphere and tropopause region for the tropics and subtropics.

The observations show frequent high cloud top  $\Theta$  ( $\Theta_{\text{CTH}}$ ) events in regions where elongated PV contours as well as horizontal winds (green contours) are suggesting strong horizontal transport and mixing processes extending from mid latitudes ( $\sim 40^\circ$  N) to high northern latitudes. Main regions are over the eastern pacific with extension in direction to Alaska, the north-eastern US directed to Greenland, from North-Atlantic and Central-Europe towards northern Scandinavia and the Baltic region, and from China towards Siberia. High altitude clouds are observed up to the northern edge of the CRISTA measurements at  $74^\circ$  N (e.g. over northern Scandinavia and the Baltic Sea on 10 August). During the four days of observations the frequency for high  $\Theta_{\text{CTH}}$  ( $> 350$  K) events north of the subtropical jet region seems to decline.

The highest detected  $\Theta_{\text{CTH}}$  (350–360 K) are in most cases above the local tropopause indicating an origin of the detected cloudy air masses in the LMS. CTHs below 350 K in regions dominated by high CTHs (e.g. in the Scandinavia-Baltic-Sea streamer) may be caused by a “real” lower altitude tropospheric cloud, but can also indicate underestimated CTHs caused by the vertical sampling of the CRISTA measurements. The constant 2 km vertical step size during CRISTA-2 in combination with a slightly drifting top altitude from one profile to the next with time and latitude result in

significant differences in the absolute tangent height between even subsequent orbits (up to 1 km) with close geographical co-location. Consequently a large cloud structure with nearly constant CTH may be detected at different altitudes in two subsequent orbits at higher latitudes or in the tropics on up and down leg of a single orbit.

## 5 4.2 Water vapour measurements during CRISTA-2

The horizontal distribution of the CRISTA water vapour at the 350 K isentrope is illustrated in Fig. 6 for the same days like in Fig. 5. For better visualisation the data of individual CRISTA H<sub>2</sub>O profiles have been first interpolated to a constant theta level (here 350 K). Vertical interpolation around the tropopause and below take always the risk that it create some numerical diffusion and artificially enhanced water vapour mixing ratios at the grid level due to the strong exponential gradient in the mixing ratio profile below the tropopause. This effect was minimized by logarithmic interpolation. In a second step a horizontal interpolation on a regular grid (1° × 1°) was performed by means of distance-weighted averaging. Data gaps in the tropics are mostly due to clouds.

The CRISTA water vapour measurements were validated with the Microwave Limb Sounder (MLS) and airborne in situ instruments (Offermann et al., 2002). The comparisons showed good agreement in the coincidence statistics and a retrieval accuracy of 10 % was estimated for the data. The precision of the data is 8 % for values > 10 ppmv, and 8–15 % for smaller values (Schaeler et al., 2005). Consequently, the horizontal structures in Fig. 6 are reliable.

Transport of water vapour from the tropical troposphere into the LMS on the 350 K isentropic surface seems evidential in Fig. 6. Shown are water vapour values as observed by CRISTA-2 on 10 to 13 August at midnight ±12 h. Rossby wave breaking events result in an erosion of the tropopause that can be identified by the cut-off of PV and water vapour contour lines over the Atlantic to Scandinavia and the Baltic sea, over the east northern pacific in direction to Alaska or from the eastern US in direction to Greenland. The intense wave processes are accompanied by fast isentropic transport

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of water vapour deep into the LMS. These horizontal structures of high water vapour are nearly coinciding with the “unusual” high CTH observation shown in Fig. 5.

The water vapour values do not include the locations of optically thicker cloud observations, because a corresponding filter ( $CI < 2$ ) was applied before the retrieval. However, it is evident that cloudy areas are embedded in regions of relatively high water vapour values. More quantitative analyses exclusively from the satellite observations are difficult, because for most cloudy CRISTA observations no adequate water retrieval is available.

## 5 Comparison with the CLaMS model

### 5.1 The CLaMS model

The Chemical Lagrangian Model of the Stratosphere (McKenna et al., 2002a, b; Konopka et al., 2007) is a chemistry transport model based on three-dimensional forward trajectories, describing the motion of air parcels. Additional to advection by winds, irreversible small-scale mixing between air parcels induced by deformation of the large scale flow is considered in the model (McKenna et al., 2002a; Konopka et al., 2004). The mixing intensity is controlled by the local Lyapunov coefficient of the flow, thus leading to stronger mixing in regions of large flow deformations. The sensitivity of simulated UTLS water vapour on the intensity of this quantity is discussed by Riese et al. (2012). The model uses a hybrid of pressure and potential temperature as vertical coordinate system first proposed by Mahowald et al. (2002).

### 5.2 Model setup

The CLaMS simulation was started in mid May 1997, three months in advance of the CRISTA observations, to give the model enough time for spin-up. The model was driven by six hourly ERA Interim re-analyses with a mixing time step of 24 h. The calculation of water vapour in CLaMS includes a simplified dehydration scheme, similar to that

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applied in von Hobe et al. (2011). Gas phase water in CLaMS is initialized at the beginning of the simulation utilizing the specific humidity taken from ERA Interim data. Boundaries are updated every CLaMS step from ERA Interim data as well. The lower boundary for this run is at 250 K with respect to the hybrid vertical coordinate, corresponding to approximately 500 hPa.

The formation of ice is parameterised either by using a fixed value of 100 % for saturation over ice (a value commonly used) or by a temperature dependent parameterisation for heterogeneous freezing (Krämer et al., 2009). The latter method was finally used in the model simulations presented below. This parameterisation results in saturation values between 120 and 140 % for the temperature range from 180 to 230 K. Water vapour with values above these saturation levels is removed from gas phase and added to the ice water content. Water vapour and ice water content are transported and mixed like any other tracer or chemical species. Evaporation at 100 % saturation and sedimentation of ice by assuming a uniform particle density and size distribution (Krämer et al., 2009) as well as parameterised processes like re- and de-hydration are considered, with the only exception of re-hydration by formerly sedimented particles. For sedimentation the terminal settling velocity is calculated. The corresponding sedimentation length is compared with a characteristic height defined by the vertical resolution of the model around the tropopause ( $\sim 650$  m for this simulation), and the related fraction of ice is removed. This mechanism was successfully used for long term studies with CLaMS (Ploeger et al., 2011, 2013). The horizontal resolution of the simulations is in the range of 70 km.

### 5.3 IWC and water vapour distribution in the model

In a first step we investigated CLaMS model results for water vapour and ice water content on synoptic maps of isentropic surfaces like in Fig. 7. Afterwards model output is analysed by applying the instrument specific limb geometry to the data, which is described in the next subsection.



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In the examples of Fig. 7 isentropic surfaces of  $\Theta = 350$  K are selected, which represent a nearly constant geometrical height with changing latitude, where isentropical transport can cross from the tropical UT into the mid latitude LS. In Fig. 7 both IWC and water vapour show distinctive streamer structures on two successive days (9 and 10 August). The streamers are elongated and spread out to mid and high northern latitudes. Especially the water vapour distribution suggests that regions of high water vapour are peeled off from the subtropical jet region, as typically observed in PV fields during Rossby wave breaking events (e.g. Homeyer and Bowman, 2013). These sub-tropical air masses are transported to and mixed in at high latitudes, where under favourable conditions the formation of cirrus clouds might be possible. Fine structures of water vapour and IWC can be observed similar to and more fine structured than the superimposed PV contours. The IWC structures are less pronounced and indicate significant less cloud formation at mid and high latitudes in contrast to the CRISTA observations (Fig. 5).

### 5.4 How to compare global model data and limb measurements?

For a quantitative comparison between the model data and limb measurements it is crucial to take into account the observation geometry and to apply averaging kernels of the instrument to the model data. In an optimised but very expensive process current investigations simulate the original measurement quantity of the instrument (e.g. here IR radiances) by a specific instrument simulator based on 3-D input parameter of a climate chemistry model (e.g. Bodas-Sacedo et al., 2011). This approach will reduce the uncertainties usually introduced by the complex retrieval process of the instrument target parameter (e.g. IWC, specific humidity or other trace gases) and is used for the validation of climate models. The detailed consideration of the observation geometry is especially important for comparisons with cloud measurements in the limb (e.g. Spang et al., 2012). Therefore we applied several processing steps for better representation of two major instrument-specific effects of the CRISTA measurements.

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In the first step, the temporal offsets between the asynoptic measurement times of the satellite and the synoptic time steps in the model output (every 24 h) have been compensated. For this purpose backward trajectories of all CRISTA observations below 25 km to the next synoptic model output time, usually every 24 h at 12:00 UTC were computed. Starting from these synoptic locations the cirrus module of CLaMS was run forward in time to the asynoptic time of the individual CRISTA observation. By this approach the formation and evaporation of ice clouds in a time frame of maximum 24 h between model output (12:00 UTC) and observation is taken into account.

Secondly, we implemented an integration of the signal along the line of sight. A limb ray tracing from the original position of the CRISTA satellite to the tangent point and the follow-on to deep space has been applied to sample the model data. In case of CRISTA the tangent height layer for the 1.5 km FOV has an extension of  $\sim 280$  km. Deeper tropospheric observations result in a factor of 2–3 longer (e.g. in the tropics) effective path lengths through the atmosphere where a cloud can occur. In the tropics for a tangent height at 10 km the maximum potential cloud occurrence altitude extends up to a height of  $\sim 18$  km and a corresponding line of sight segment of  $\sim 640$  km should be considered in the limb ray tracing.

Spang et al. (2012) showed that limb IR measurements of cirrus clouds are most sensitive to the integrated surface area density along the limb path (area density path, ADP) and that ADP is a useful quantity for comparisons with global models, where the limb path can be traced through the model output to generate the ADP quantity. The ADP and CI show an excellent correlation and ADP can be retrieved from the measured CI values (Spang et al., 2012). For a homogeneous limb path segment ADP and limb ice water path (IWP) are related by the simple equation:

$$\text{ADP} = 3 \cdot \text{IWP} / (R_{\text{eff}} \cdot \rho_{\text{ice}}) \quad (2)$$

with  $\rho_{\text{ice}}$  the mass density of ice, and  $R_{\text{eff}}$  the effective radius of the particle size distribution. More generally the limb IWP can be computed by the following relations,

depending which parameters are known from the model or measurements:

$$\text{IWP} = \int_0^{\infty} \text{IWC} dx = \int_0^{\infty} V \cdot \rho_{\text{ice}} dx = \frac{1}{3} \int_0^{\infty} A \cdot R_{\text{eff}} \cdot \rho_{\text{ice}} dx, \quad (3)$$

where  $R_{\text{eff}}$  and IWC are defined by the model,  $V$  and  $A$  represent volume and surface area density respectively (Spang et al., 2012).

Finally we have used the simulated IWC to compute the IWP for a CRISTA-like cloud detection in the CLaMS model fields. A 30 km step width along the line of sight over a distance of  $\pm 1000$  km with respect to the tangent point has been chosen, which is in line with the horizontal resolution of the model. Then  $\text{IWC}_{\text{CLaMS}}$  has been interpolated on the line of sight grid locations. A CTH detection is defined when the first (top) line of sight beam of a CRISTA profile shows an  $\text{IWP}_{\text{CLaMS}} > 0$ . The latter fact is neglecting any detection sensitivity threshold of the instrument and represents therefore an upper limit of what a CRISTA-like instrument would detect in the cloudy atmosphere modelled by CLaMS.

An example of the CRISTA-like limb IWP is given in Fig. 8. All tangent heights between the 330 and 370 K isentrope with  $\text{IWP} > 0$  are presented for 1997 10 August 00:00 UTC  $\pm 12$  h. The results can be compared with Figs. 1 and 5a, even though these figures show the CTH of a measured profile. For a perfect agreement between model and measurement the CLaMS cloud detections should be exactly at the location where CRISTA observed a cloud. Obviously the model field of IWP shows a good agreement with the cloud occurrence observed by CRISTA. Similar regions show the occurrence of clouds at mid and high latitudes. Some regions are extended larger in the observations than in the model (e.g. the east end of the North Atlantic to Baltic Sea streamer, the extension over Alaska or over Kamchatka). However there are also a few regions where the model shows clouds but the observation is cloud free. But care should be taken in the comparison of model grid and instrument grid data. For example, when the interpolation onto the line of sight is performed numerical errors can generate unrealistically

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the tropopause in the observations and slightly higher in altitude and percentages in the model. In addition, a reduced activity in the subtropics for both datasets and very similar COFs around the mid- and high-latitude tropopause are found. However, the CRISTA measurements show substantially increased cloud occurrence frequencies at altitudes well above the tropopause for all latitude bands.

The differences in COF between observation and model are illustrated in Fig. 9c. Obviously the model overestimates the COF in the tropical upper troposphere and up to slightly above the tropopause, a region defined with the term tropical transition layer (Fueglistaler et al., 2009). Between the isentropic surfaces 350 and 360 K the region of overestimation is extended up to latitudes of 40° N. At these altitudes horizontal transport is possible from tropical and sub-tropical air masses with high water vapour mixing ratios into the LMS to mid and higher latitudes. Since the subtropical jet acts as a transport barrier to meridional transport, this indicates a weakening of the subtropical jet, a condition typically coincident with RWB events in the jet region (Postel and Hitchmann, 1999). RWB events have been observed during the CRISTA-2 mission, like illustrated in Figs. 5 and 6 by the development of the PV contours during the mission. Such events are very typical in summer at this altitude and latitude location (e.g. Gabriel and Peters, 2008; Homeyer and Bowman, 2013).

Above 360 K in the subtropics and north of 40° N in the tropopause and LMS region CRISTA observations indicate higher COFs than the model and – as already presented in Fig. 4 – a large amount of high altitude cloud occurrence in the LMS, which is only weakly present in the model calculations. This is caused by the cirrus module in CLaMS which is including only a simplified approach for ice formation (mainly driven by temperature, the super saturation threshold in respect to ice, specific humidity, and IWC) and is not considering detailed microphysical background or constraint boundary conditions.

Sensitivity tests with the current setup of CLaMS with changing saturation thresholds for ice formation (100–150 %) and varied radius parameterisations for the sedimentation process did not improve the comparison in the LMS. Overall these variations in the

setup had little impact on the horizontal and vertical distribution of clouds in the LMS. The effect of sedimentation of ice particles in the model is so far not accurately considered. For large particles and therefore high sedimentation velocities IWC is not settling into a lower altitude CLaMS box and this settling part of available water is dehumidified from the model.

First tests with a more sophisticated model, a bulk model including microphysical formation mechanisms and more detailed cloud formation processes (Spichtinger and Gierens, 2009a, b) running on forward trajectories to single cloud locations of the CRISTA observations, were not capable to reproduce the formation of cirrus in the LMS. Only slight biases in temperature or the water vapour entrance value into the LMS may have a significant effect on the cloud formation process in the models. These parameters are mainly guided by the meteorological analysis data used for the model run. Hence they are hardly to improve and it is difficult to quantify if there are any biases in these datasets. In addition, temperature fluctuations by gravity waves causing high updraft velocities may play a key role in the nucleation process of ice particles (Spichtinger and Krämer, 2013). These temperature variations need to be considered accurately along the trajectory of an air parcel, but are currently not available for global modelling of cirrus clouds.

## 5.6 Limb ice water path comparison

This section presents the limb IWP estimated from CRISTA in comparison to CLaMS. As already outlined in Sect. 5.3 the parameters best suited to compare model results with a limb measurement of microphysical information of cirrus cloud are the limb integrated area density and ice water path.  $IWP_{CLaMS}$  is computed from the model data following Eq. (3) by using the CLaMS air parcel information of IWC and the limb ray tracing described in Sect. 5.4. The zonal mean  $IWP_{CRISTA}$  and  $IWP_{CLaMS}$  distributions presented in Fig. 10 show similar structures like the CTH occurrence frequencies (Fig. 9), but include all tangent heights with cloudy signals in the measurements ( $IWP > 0$ ) and in the observations ( $CI < 4$  equivalent to  $IWP > \sim 10^{-3} \text{ gm}^{-2}$ ) and not

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only the CTH locations. For the zonal mean calculation for CLaMS we shortcut all IWP values greater than the upper retrieval limit to  $IWP = 20 \text{ gm}^{-2}$ , because clouds with larger IWP are not discriminable in the CRISTA IWP retrieval. By this approach zonal mean values of model and observation become comparable. A good correspondence is found in the zonal mean of  $IWP_{\text{CRISTA}}$  and  $IWP_{\text{CLaMS}}$  for altitudes at and below the tropopause, although the CLaMS values tend to be larger than CRISTA when penetrating to lower levels of the troposphere. Only above the tropopause the significant cloud occurrence rates of CRISTA and the vanishing probability to create a cloud in the model become noticeable with means of  $IWP_{\text{CRISTA}} > IWP_{\text{CLaMS}}$ . The overall good correspondence below the tropopause is indicating the dominance of optically thicker clouds below the tropopause, whereby the probability for both, model and observation, is much higher to match a similar size of the retrieved IWP values than for the much more variable optical thicknesses of cirrus clouds at and above the tropopause.

## 6 Summary and conclusions

A re-analysis of cloud and water vapour measurements during the CRISTA-2 mission in August 1997 in conjunction with model calculations with the Lagrangian chemical transport model CLaMS were presented. Special emphasis was taken to quantify the cloud top altitude with respect to the local tropopause to demonstrate the potential importance of cirrus cloud formation in the lowermost stratosphere above the local tropopause. Little is known about the occurrence frequency and spatial distribution of this particular cloud phenomenon, for example how and why these clouds may form. The occurrence of LMS clouds was previously reported by some lidar stations in the Northern Hemisphere and occurrence frequencies above the tropopause of a few percent of the total number of cirrus profiles were reported. CRISTA observations extend these local observations to global coverage in the Northern Hemisphere and show strong indications for frequent cirrus occurrences in the LMS at mid and high latitudes.

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A reliable hemispheric picture of cloud occurrence in the LMS above the tropopause based on CRISTA-2 observations hinges on an accurate determination of the local tropopause. Here we applied a sophisticated new algorithm to the ERA Interim reanalysis dataset. This so-called high resolution tropopause was compared with TPHs from radiosonde data and showed a good agreement ( $\pm 650$  m) for more than 5000 selected station profiles in the time frame of the CRISTA-2 mission. This demonstrates the good quality of the ERA Interim temperature profiles around the tropopause.

Uncertainties in the determination of the TPH location, instrument specific effects like the potential overestimation of the CTH for optically thick clouds due to the vertical extent of the field of view or uncertainties in the tangent height location were addressed in the present analysis. By quantifying all potential error sources as accurately as possible and modelling the effect of these uncertainties on cloud occurrence frequency statistics in tropopause related coordinates the CRISTA observations show significant numbers of cirrus detections clearly above the local tropopause (500–1500 m) and consequently in the lowermost stratosphere.

In general, the cloud top height occurrence frequencies (COF) at the mid-latitude tropopause are in good agreement with the analysis of Pan and Mynchak (2011) based on the spaceborne lidar CALIOP, although these COFs are based on a multi-annual seasonal mean. The CRISTA-2 results show larger COFs than CALIOP above the mid-latitude tropopause ( $> 500$  m) and also above the tropical tropopause. Overall, rather high occurrence frequencies ( $\sim 5$ – $10$  % of all profiles) up to high northern latitudes ( $70^\circ$  N) and altitudes well above the tropopause ( $> 350$  K) were found in astonishingly large areas of the LMS. These numbers indicate a significantly larger occurrence frequency than in the ground based observations.

Further, the Northern Hemisphere CRISTA water vapour observations indicate a considerable isentropic flux of moisture (at  $\sim 350$  K) from the upper tropical troposphere into the extra-tropical lowermost stratosphere (LMS). This process is triggered by Rossby wave breaking events in the subtropical jet region accompanied by long range transport of high water vapour abundance in streamer and filament structures





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communication, 2014). Temperature fluctuations may play a key role for a realistic modelling of the formation process of cirrus (Spichtinger and Krämer, 2013), but are difficult to constrain, especially on global scales. Usually temperatures are not low enough and temperature variations are not fast enough to initiate ice formation under the model background conditions. Amongst others, these first attempts of modelling the observations of LMS cirrus indicate that already the meteorological analyses used for the initialisation of water vapour, IWC, and the temperatures along the trajectories, may not include the processes and variability (e.g. gravity waves or small scale and high frequency fluctuations) necessary to generate cirrus clouds in the LMS region. In addition, the formation processes of this specific cirrus cloud type may differ from the current knowledge and implementation in the current microphysical models.

Improvements and new developments in the cirrus modules in models as well as multi-instrumental analysis approaches are necessary to achieve progress concerning the questions and unknowns about the formation of cirrus clouds in the LMS. More accurate frequency distribution, seasonal cloud coverage of the northern and Southern Hemisphere, and microphysical information of LMS cirrus are necessary to quantify the potential radiation and climate impact of LMS cirrus. The unprecedented frequent and statistically significant observation of LMS cirrus by CRISTA may initiate more specific measurement campaigns or model studies with respect to LMS cirrus to quantify the importance of this still intriguing cloud type.

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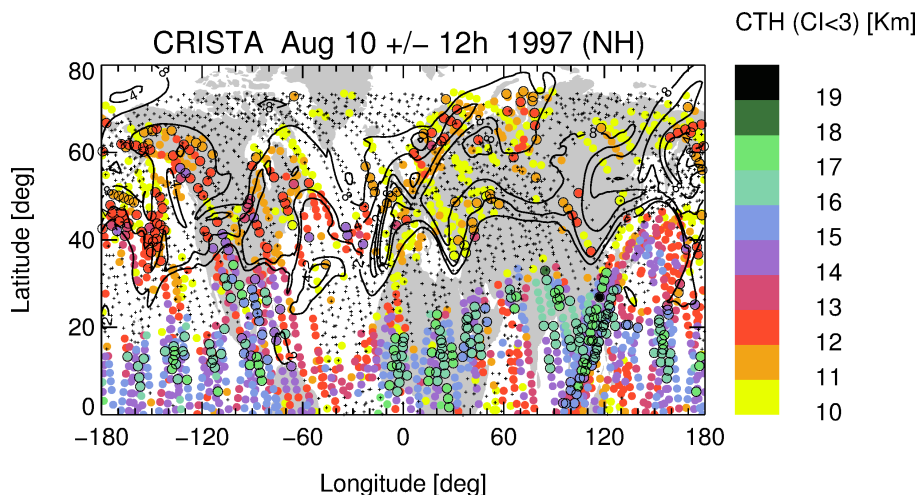
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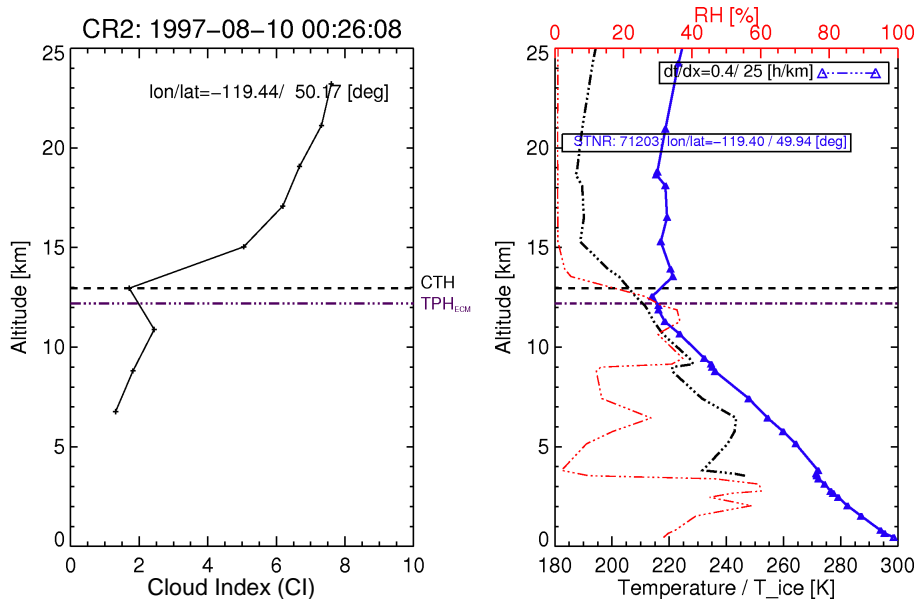
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**Fig. 1.** CRISTA cloud top height distribution for 10 August 1997 00:00 UTC  $\pm$  12h (coloured circles). Profiles with no cloud indications are marked by black crosses. Overlaid in black are contours of potential vorticity at 2, 4, and 8 PVU at the 350 K isentropic. Cloud tops with PV values greater than the proxy threshold 2 PVU for the dynamical tropopause are highlighted by black borders of the coloured circles for CTH.

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**Fig. 2.** Cloud index (CI) profile during CRISTA-2 (left) and coincident radiosonde measurements (right) at mid latitudes. The profiles on the right side show temperature (blue) and relative humidity (red). The ice saturation temperature curve is superimposed by the black-dash-dotted curve. Horizontal lines indicate cloud top height (CTH) from the CRISTA measurements (dashed) and the so-called “high resolution” tropopause height (TPH<sub>HR</sub>) from ERA Interim temperature data (dash-dotted), for details see Sect. 2.5.

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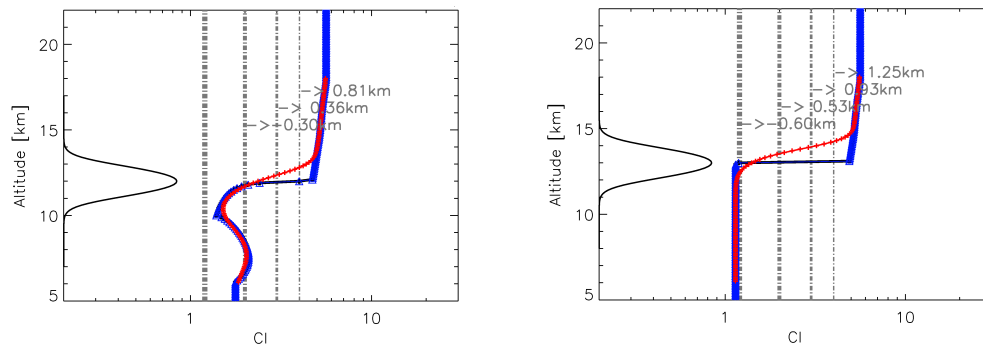
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**Fig. 3.** Examples of cloud index profile computed from modelled radiance profiles with high vertical resolution (0.1 km, in blue) for a 2 km thick cloud layer with cloud top at 12 and 13 km for optically thin (left) and thick (right) conditions respectively. The Gaussian shaped field of view was applied to the pencil beam simulations (blue) to simulate CRISTA profiles (red) (examples of the FOV function are centred at the cloud top at 12 and 13 km). The superimposed numbers indicate the maximum errors in CTH,  $dCTH = [\text{no detection}, -0.30, 0.36, 0.81]$  km (left) and  $dCTH = [-0.56, 0.53, 0.94, 1.28]$  km (right) for the corresponding CI threshold values [1.2, 2, 3, 4] applied in the cloud detection.

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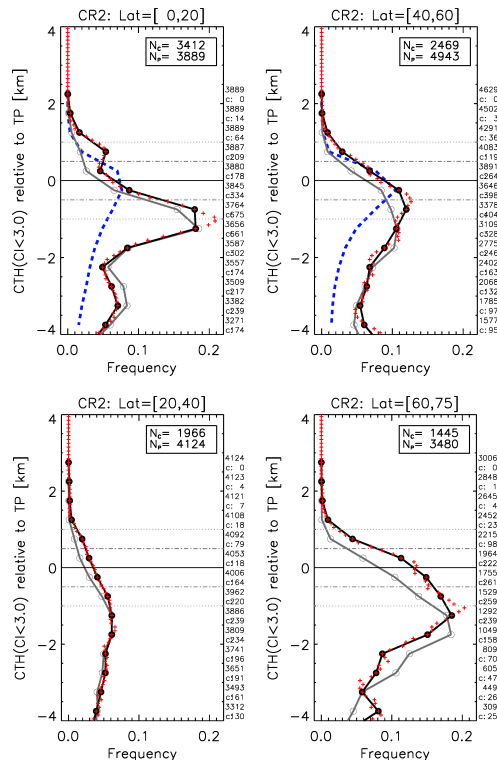
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**Fig. 4.** Cloud occurrence frequencies of CTHs relative to the tropopause for four latitude bands and the two detection thresholds  $Cl_{thres} = 2$  (grey circles) and  $Cl_{thres} = 3$  (black dots). The numbers at the right y-axis are highlighting the number of observations (top number) and the number of CTH counts (c:) in the corresponding altitude grid box (500 m) for  $Cl_{thres} = 3$ . Blue dashed line in the top two figures represents the JJA mean COF values for CALIPSO in the time frame June 2006 to May 2010 (taken from Fig. 10 in PM2011). Uncertainty limits of  $\pm 0.5$  and 1 km are presented by the horizontal lines. Red symbols represent a 100 m running mean statistic of 500 m vertical boxes.

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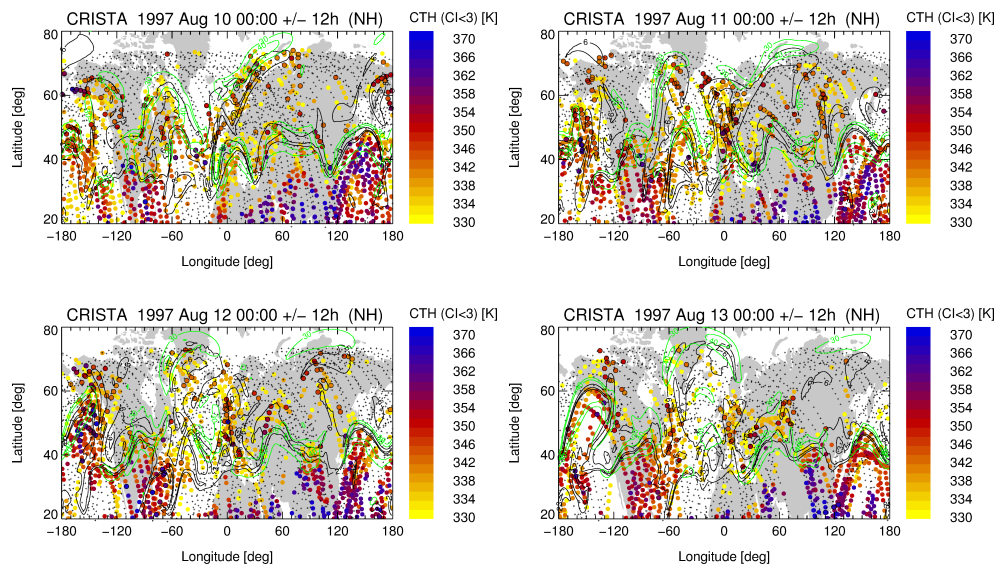
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**Fig. 5.** CRISTA-2 daily detection of cloud top heights on vertical  $\Theta$ -coordinates between 330 and 370 K (colour coded) from 10 to 13 August at 00:00 UTC  $\pm$  12 h. Black circles around the coloured symbols mark CTH observation above the local tropopause. Potential vorticity contours for 2, 3, and 6 PVU are overlaid in black at midnight conditions. In addition horizontal wind contours for 30 and 40  $\text{ms}^{-1}$  (in green) are highlighting the subtropical jet as well as regions with fast horizontal transport at higher latitudes. Crosses are marking non-cloudy profiles in the  $\Theta$ -range 330 to 370 K.

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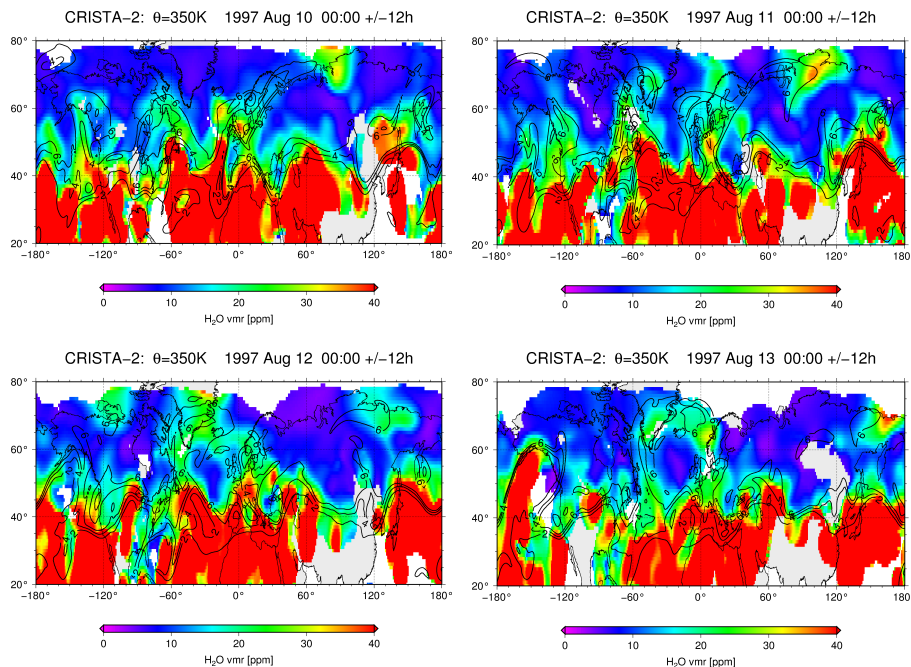
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**Fig. 6.** CRISTA 2 measurements of water vapour for 10 to 13 August 00:00 UTC  $\pm$  12 h. The asynoptic measurements are interpolated spatially to a regular grid (for details see text). In addition, PV contour lines of 2, 4, and 8 PVU are overlaid in black. White patches in the water distribution indicate data gaps due to clouds.

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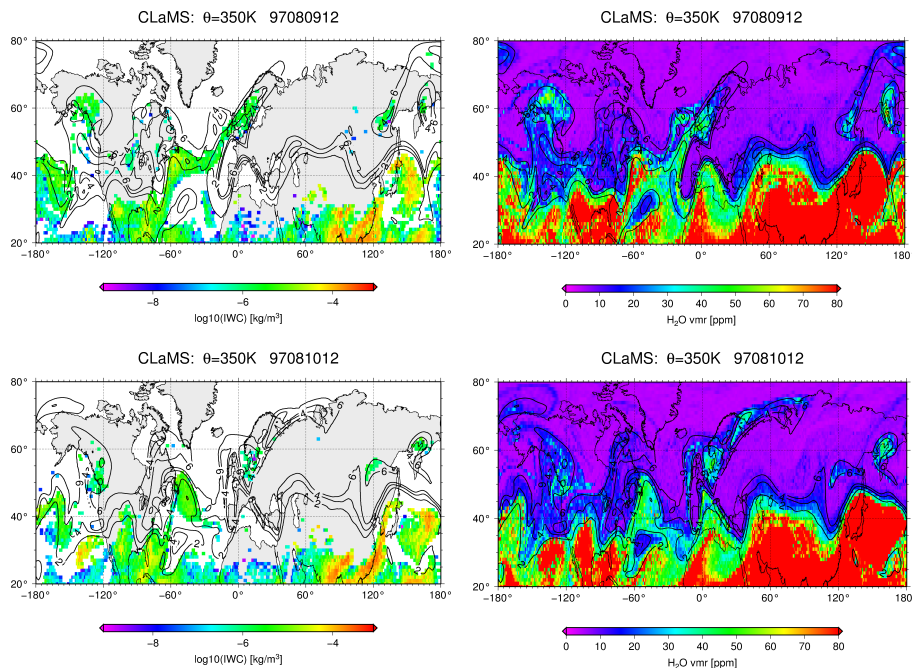
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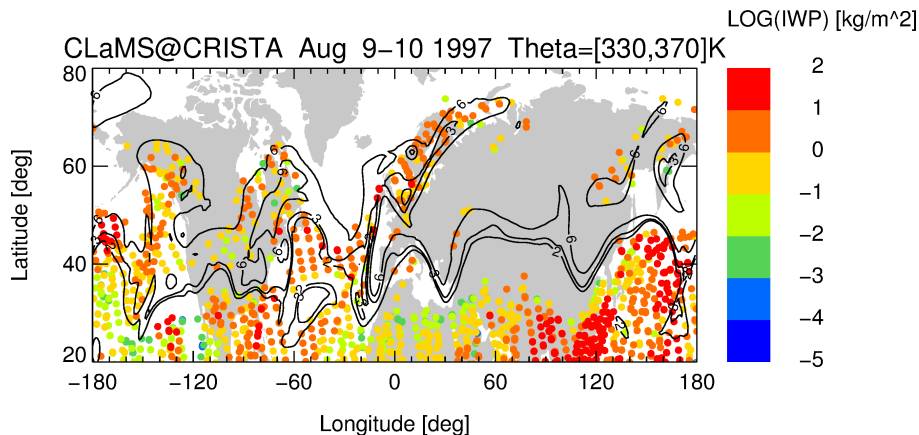
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**Fig. 7.** CLaMS model results of IWC (left) and water vapour (right) for meteorological input data based on ERA Interim, top row for the 9 August 12:00 UTC and bottom row for the 10 August 1997. CLaMS irregular data are interpolated to a  $1^\circ \times 1^\circ$  grid for the 350 K isentrope.

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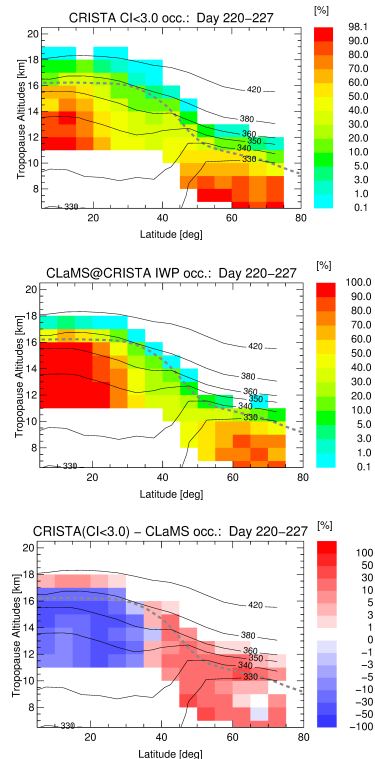


**Fig. 8.** Limb ice water path based on derived from CLaMS IWC integrated along CRISTA line of sight for 10 August 1997 00:00 UTC  $\pm$  12 h between the 330 and 370 K isentrope (for details see text). PV contours for 2, 3, and 6 PVU are superimposed.

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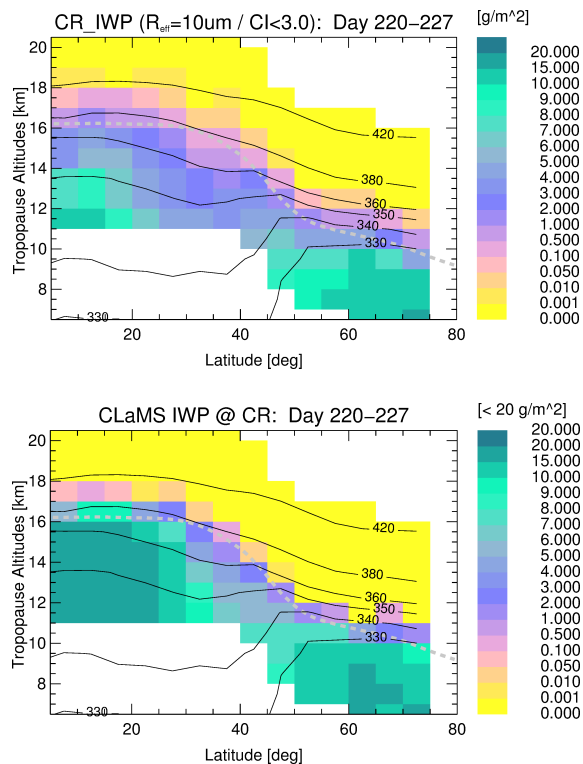


**Fig. 9.** Cloud top height occurrence frequencies in tropopause related vertical coordinates for the complete measurement period (7 days) of CRISTA-2 (top) based on  $CI_{\text{thres}} = 3$ , the corresponding CLaMS model is sampled with line of sights of CRISTA, and cloud top detection is based on  $IWP > 0$  (middle). Differences in cloud top height occurrence frequency between CRISTA and CLaMS sampled with CRISTA are presented in the bottom diagram. Contours for zonal mean isentropes (black) and for the zonal mean tropopause altitude (dashed grey) are superimposed.

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**Fig. 10.** Zonal mean CRISTA limb IWP retrieved from CI (top) and from the CLaMS model (bottom) with a shortcut of  $\text{IWP}_{\text{CLaMS}} > 20 \text{ g m}^{-2}$  at the upper detection sensitivity of CRISTA (right). For the computation of  $\text{IWP}_{\text{CR}}$  it is necessary to assume a constant effective radius for the ice particles ( $R_{\text{eff}} = 10 \mu\text{m}$  is applied, see also Eq. 3).