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## Differences in diabatic heating

J. S. Wright and  
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# Large differences in the diabatic heat budget of the tropical UTLS in reanalyses

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## Abstract

We present the time mean heat budgets of the tropical upper troposphere (UT) and lower stratosphere (LS) as simulated by five reanalysis models: MERRA, ERA-Interim, CFSR, JRA-25/JCDAS, and NCEP/NCAR. The simulated diabatic heat budget in the tropical UTLS differs significantly from model to model, with substantial implications for representations of transport and mixing. Large differences are apparent both in the net heat budget and in all comparable individual components, including latent heating, heating due to radiative transfer, and heating due to parameterised vertical mixing. We describe and discuss the most pronounced differences. Although they may be expected given difficulties in representing moist convection in models, the discrepancies in latent heating are still disturbing. We pay particular attention to discrepancies in radiative heating (which may be surprising given the strength of observational constraints on temperature and tropospheric water vapour) and discrepancies in heating due to turbulent mixing (which have received comparatively little attention).

## 1 Introduction

Meteorological analyses and reanalyses are best guesses of the true state of the atmosphere. As such, they are of eminent importance not only for initialisation of weather forecast model runs, but for process analyses and detection and attribution of changes in the climate system. The (re)analysis model state is optimised to observable quantities such as winds, temperature, trace gas mixing ratios (particularly water vapour), and clouds; accordingly, the “analysis step” may not strictly conserve the energy budget of the model state. For example, the moisture assimilation step may add or remove water substance (and hence latent heat), while the temperature assimilation step may add or remove sensible heat. Despite this caveat, (re)analyses are still superior to free-running general circulation models for a wide range of applications because the latter only allow comparison with observations in a statistical sense.

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Here, we present an overview of the diabatic heat budgets of the tropical upper troposphere and lower stratosphere (UTLS) in five widely used reanalyses: NASA's Modern-Era Retrospective Analysis for Research and Applications (MERRA; Rienecker et al., 2011), ECMWF's ERA-Interim (Dee et al., 2011), NCEP's Climate Forecast System Reanalysis (CFSR; Saha et al., 2010), JMA's JRA-25 Reanalysis and its continuation with the JMA Climate Data Assimilation System (Onogi et al., 2007), and the NCEP/NCAR Reanalysis 1 (Kalnay et al., 1996). Fueglistaler et al. (2009b) have previously provided a comparison of the ERA-Interim diabatic heat budget with that of the older ERA-40, and Ling and Zhang (2013) have examined the diabatic heat budgets of MERRA, CFSR, and ERA-Interim at lower levels in the troposphere.

We focus on the tropical UTLS between 300 hPa and 50 hPa. This region encompasses the important transition in regime from a balance between latent heating and radiative cooling in the troposphere to a balance between radiative heating (in the tropics) and radiative cooling (in the extratropics) in the stratosphere. Furthermore, transport into the stratospheric overworld (terminology follows Hoskins, 1991) occurs predominantly in the tropics, and conditions in the tropical tropopause layer (TTL, ~150–70 hPa) control the stratospheric abundance of a range of radiatively and chemically important trace gases (see review by Fueglistaler et al., 2009a). Uncertainties in the model diabatic heat budget in this layer therefore also affect confidence in model predictions of future stratospheric composition and its radiative impact on climate.

The diabatic heat budget of the TTL (and consequently transport pathways and mechanisms within the TTL) remains a subject of debate (see discussion in Fueglistaler et al., 2009a). In particular, the radiative effects of clouds and the impact of convection that overshoots its level of neutral buoyancy are not well quantified. Both of these processes pose challenges to global-scale numerical models due to the small scales involved. Additional processes, such as mixing related to shear-flow instability, have received comparatively less attention, but should also be discussed as they are highly parameterised and often serve as a “tuning parameter” that compensates for errors in other aspects of the model (see also discussion in Flannaghan and Fueglistaler, 2011).

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One may expect the radiative components of reanalyses to be comparatively more accurate because temperature and water vapour (in the troposphere) are constrained by observations. However, ozone concentrations are also radiatively important in this region and are generally not constrained by observations (in many cases, a prescribed climatology is used). Furthermore, cloud radiative effects may differ substantially between models.

Section 2 provides a description of the reanalysis models and data. Section 3 presents key aspects of the climatological mean diabatic heat budget in the tropical UTLS and describes the large differences in estimates of this budget according to different reanalyses. Section 4 discusses the sources and implications of some of the most pronounced differences among the reanalyses. Section 5 provides a summary of the current outlook.

## 2 Reanalysis data

The Modern Era Retrospective Analysis for Research and Applications (MERRA) has been produced by NASA's Global Modeling and Assimilation Office (Rienecker et al., 2011). MERRA was conceived as a reanalysis of the satellite era (1979–present), with the primary objective of improving the representation of the water cycle relative to previous reanalyses. Diabatic heating is reported on a  $1.25^\circ \times 1.25^\circ$  grid and includes components of heating due to long-wave radiation, short-wave radiation, moist physics, turbulent vertical mixing, gravity wave drag, and friction.

ERA-Interim (ERA-I) is, like MERRA, a reanalysis of the satellite era (1979–present). ERA-I has been produced by the European Centre for Medium-Range Weather Forecasting, and has been described in detail by Dee et al. (2011). We have interpolated the heating rates onto a  $1^\circ \times 1^\circ$  horizontal grid with 60 isobaric levels that correspond to the average pressure of the eta levels used in the forecast model. Individual components of diabatic heating due to long-wave and short-wave radiative heating are provided;

however, heating due to moist physics is not reported separately and must be inferred as a residual (see Fueglistaler et al., 2009b).

The Climate Forecast System Reanalysis (CFSR) was developed by the National Centers for Environmental Prediction (NCEP) as the first coupled atmosphere–ocean–land surface–sea ice reanalysis (Saha et al., 2010). CFSR technically only covers the period 1979–2009, but continuing output from the current (effectively identical) version of the Climate Forecast System model extends the CFSR data record through the present. Heating rates are provided on a  $1^\circ \times 1^\circ$  horizontal grid with 37 isobaric levels. In addition to total diabatic heating rates, CFSR provides heating rates due to long-wave radiation, short-wave radiation, deep convection, shallow convection, large-scale condensation, and vertical diffusion.

The Japanese 25 year Reanalysis (JRA) was initially commissioned to cover the period 1979–2004 (Onogi et al., 2007), but has since been extended to the present using the Japan Meteorological Agency (JMA) Climate Data Assimilation System. The diabatic heating rates used in this study are reported on a  $2.5^\circ \times 2.5^\circ$  horizontal grid with 23 isobaric levels, and include components due to long-wave radiation, short-wave radiation, convection, large-scale condensation, and vertical diffusion.

The NCEP/NCAR Reanalysis 1 (NCEP) covers the period from 1948 to the present (Kalnay et al., 1996). NCEP was among the first reanalyses to be produced and remains one of the most widely used for scientific research. The NCEP diabatic heating data are provided as monthly means on the original T62 model grid with 28 sigma levels. We have interpolated these data to fixed pressure levels that match the average pressure of the sigma levels used in the forecast model. The components of diabatic heating are identical to those reported by CFSR.

Table 1 provides an overview of the reanalysis forecast model frameworks, as well as key attributes of the radiation parameterisations. All of the diabatic heating rates used here are based exclusively on 6 h model forecasts except for the ERAI heating rates, which are based on 12 h model forecasts. The only observational constraint on

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reanalysis diabatic heating is the effect of the data assimilation on the forecast initial conditions.

### 3 Climatological diabatic heat budgets

The total diabatic heat budget and its major components are shown for each of the five reanalyses in Figs. 1–3. All quantities in these figures are averaged over the decade January 2001–January 2010. Figure 1 shows zonal mean distributions of total, radiative, and residual diabatic heating. Residual diabatic heating includes temperature tendencies due to all non-radiative physical processes, and is dominated by latent heating in the troposphere. Figure 2 shows the longitudinal distribution of total diabatic heating averaged over the inner tropics ( $5^{\circ}$  S to  $5^{\circ}$  N). The patterns are qualitatively similar (though more muted) when diabatic heating is averaged over  $15^{\circ}$  S to  $15^{\circ}$  N or  $30^{\circ}$  S to  $30^{\circ}$  N. Figure 3 shows the zonal mean distributions of the long-wave (LW) and short-wave (SW) components of radiative heating. Differences among these five data sets are substantial throughout the UTLS, and are not limited to any one component of the diabatic heat budget. These differences show that the representations of physical processes in this region are remarkably different between the reanalysis atmospheric models, and imply the existence of substantial differences in tracer transport and mixing.

For convenience, we divide the diabatic heat budget into two vertical domains within the UTLS: the upper troposphere (where latent heating is important) and the tropical tropopause layer and lower stratosphere (where latent heating largely disappears and the diabatic heat budget is dominated by radiative heating). The vertical boundary between these two domains is near 150 hPa for all reanalyses but NCEP, for which the boundary is approximately 100 hPa.

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### 3.1 The upper troposphere

The upper part of the overturning Hadley circulation is readily apparent between 300 and 150 hPa in the zonal mean diabatic heat budget (Fig. 1). Diabatic heating indicates an upward mass flux in potential temperature coordinates, while diabatic cooling indicates a downward mass flux. Surfaces of constant potential temperature roughly coincide with surfaces of constant pressure in the tropical UTLS, particularly in the inner tropics, so that regions of diabatic heating indicate regions of climatological mean ascent in pressure coordinates as well. In the context of the Hadley circulation, latent heating in the UT (predominantly associated with deep convection) is balanced by radiative cooling. Deep convection preferentially occurs in the inner tropics, leading (in a time-mean sense) to diabatic ascent over the inner tropics and diabatic descent over the subtropics.

Two peaks in tropical heating (located near 5° S and between 5 and 10° N) are identifiable in both the total and residual diabatic heating fields. These peaks correspond to seasonal shifts in the zonal mean location of the ascending branch of the Hadley cell, or intertropical convergence zone (ITCZ). The locations of these maxima are consistent among the reanalyses; however, the horizontal extent, vertical depth, and magnitude of zonal mean heating differ considerably, especially with respect to net diabatic heating. For example, MERRA indicates that the maximum in net diabatic heating in the inner tropics is relatively broad in horizontal extent but shallow in vertical extent, whereas ERAI indicates that it is weaker and narrower. The vertical extent of total heating associated with the ascending branch of the Hadley cell is also relatively shallow in ERAI, with a minimum in heating near 250 hPa rather than 200 hPa (as in MERRA, CFSR, and JRA) or 100 hPa (as in NCEP). It is remarkable that the climatological mean heating in this region is weakest in ERAI given that the residual heating (dominated by latent heat release) is stronger in ERAI than MERRA or CFSR (Fig. 1, bottom panels). This situation arises because ERAI radiative cooling is much stronger (Fig. 1, middle panels).

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The CFSR and JRA zonal mean diabatic heat budgets are qualitatively similar below 200 hPa: net heating in the inner tropics is narrower than MERRA but broader than ERAI, while radiative cooling is stronger than MERRA but weaker than ERAI. Quantitative differences between these two reanalyses are difficult to evaluate given differences in the spatial resolution of the data (1° for CFSR compared to 2.5° for JRA). The radiative heat budget of NCEP is similar to that of the other reanalyses in the upper troposphere, but the residual diabatic heating is much larger and extends higher (up to about 100 hPa) than in any other reanalysis.

Net diabatic cooling prevails throughout the zonal mean tropics in MERRA and CFSR between ~ 200 hPa and ~ 150 hPa. The extratropics do not provide a net diabatic heat source at these pressure levels either (see, e.g. Ling and Zhang, 2013, their Fig. 5). Although diabatic heating rates need not balance upon global averaging (as cross-isentropic mass fluxes must do), the presence of pressure surfaces with net negative heating rates everywhere indicates a deficit in the model's diabatic heat budget. This type of deficit may be corrected during the assimilation increment (the incremental adjustment of forecast temperatures during the data assimilation step; see, e.g. Fueglistaler et al., 2009b); however, simulations of atmospheric transport driven by these heating rates will encounter a vertical “transport barrier”. This layer of zonal mean diabatic cooling persists year-round in MERRA monthly means (not shown). Seasonally-varying regions of diabatic ascent connect the tropical UT to the tropical LS in each of the other four reanalyses, including CFSR.

Just as the zonal mean distribution of diabatic heating in the tropical UTLS is strongly tied to the Hadley circulation, the longitudinal distribution is strongly tied to the equatorial Walker circulation (Fig. 2). Regions of net diabatic heating are located in the well-known equatorial convective regions over South America (near 60° W), western and central Africa (near 20° E), and the maritime continent (centered near 120° E). Conversely, the absence of convection over the Eastern Pacific within 5° S–5° N leads to strong net subsidence. These geographical distinctions in the net diabatic heating field extend upward to approximately 150 hPa in ERAI, CFSR, and JRA and up

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to approximately 100 hPa in NCEP; however, the regions of UT heating are capped by strong net diabatic cooling at 200 hPa in MERRA. In fact, the meridional distribution of net diabatic heating at 200 hPa in MERRA is qualitatively opposite to that in ERAI, CFSR, JRA, or NCEP: tropical deep convective regions represent local minima in 200 hPa diabatic heating in MERRA but maxima in the other four reanalyses, while tropical subsidence zones represent local maxima in MERRA (although net heating is still negative) but minima in the other four reanalyses.

The time-mean transport barrier in MERRA appears to arise from a combination of two factors. First, the vertical extent of latent heating associated with moist convection appears to be considerably shallower in MERRA than in the other four reanalyses (see, e.g. the residual heating distribution shown in Fig. 1k relative to those shown in Fig. 1l–o). Second, there is a pronounced local maximum in UT radiative cooling at 200 hPa. This maximum is centered over the equator, directly above the peak latent heating associated with the ascending branch of the Hadley circulation, and is consistent with the maxima located over the ascending branches of the meridional Walker circulation. This distribution of radiative heating is unique to MERRA among the five reanalyses, and is related to LW radiative transfer (Fig. 3a). We will return to this topic in Sect. 4.

The distributions of SW radiative heating in the UT (Fig. 3f–j) are consistent with the hypothesis that these distributions arise from SW cloud radiative heating, which in this region is dominated by radiative heating associated with clouds of convective origin. This relationship can be deduced from the similarity between the zonal mean patterns of SW radiative heating and residual heating (Fig. 1k–o), the latter of which is dominated by deep convective latent heating. This suggests that differences in the distribution of SW radiative heating in this region are largely attributable to differences in model representations of deep convection and the associated clouds.

### 3.2 The tropical tropopause layer and lower stratosphere

The non-radiative diabatic heating terms are small above approximately 150 hPa in all reanalyses, with the notable exception of NCEP (where residual diabatic heating is

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large up to about 100 hPa). As expected, this shift in the dominant component of net diabatic heating roughly coincides with the level of zero net radiative heating (LZRH) in each reanalysis (see Fig. 3). The vertical location of the LZRH in pressure is largely consistent among the reanalyses (again with the exception of NCEP), but its zonal structure varies considerably. The LZRH is higher in the inner tropics than in the subtropics in MERRA and NCEP, whereas it is lower in the inner tropics than in the subtropics in ERAI and roughly isobaric in CFSR and JRA.

The meridional structure of the inner-tropical total diabatic heat budget in the TTL/LS is qualitatively similar among the reanalyses (Fig. 2); however, the zonal mean structure (Fig. 1) reveals differences that are important in the context of the forcing of tropical upwelling. In particular, ERAI shows a pronounced v-shaped structure of maximum heating rates, with maxima in the inner tropics below 100 hPa and in both subtropics near 70 hPa. This v-shaped structure is much less apparent in the other reanalyses, which simulate only weak local maxima in the subtropics near 70 hPa. These differences in the distribution of diabatic heating imply differences in the strength of upwelling in the inner tropics relative to upwelling in the subtropics, which in turn imply differences in the distribution of wave drag and the stratospheric overturning circulation (see, e.g. Plumb and Eluszkiewicz, 1999).

Figure 3 shows that many of the differences in the structure of the net diabatic heat budget in the TTL and LS arise from differences in longwave radiative heating, which is strongest (most positive) in ERAI and weakest in JRA and NCEP. Longwave radiative heating in this layer is strongly affected by temperature (with lower temperatures resulting in stronger radiative heating and vice versa; see, e.g. Fueglistaler et al., 2009b) and ozone. The differences in radiative heating are consistent with known temperature biases in ERAI (which is somewhat cold-biased; see Fueglistaler et al., 2013) and NCEP (which is somewhat warm-biased, particularly over the maritime continent; see Pawson and Fiorino, 1998). For ozone the situation is more complex: both ERAI and NCEP use older climatologies that do not resolve the ozone gradients in this layer well

(see Table 1). MERRA and CFSR use prognostic ozone simulated by the underlying forecast model, while JRA uses daily offline calculations.

Net diabatic heating rates in JRA are negative throughout much of the tropics between about 60 and 40 hPa (upper level not visible in Figs. 1–3), primarily due to a lack of LW heating in this layer relative to the other reanalyses (Fig. 3d). This layer of time-mean net diabatic cooling is again physically unrealistic. Moreover, it contradicts the generally accepted view of the stratospheric residual circulation, in which diabatic descent (cooling) over the extratropics is balanced by diabatic ascent (heating) at low latitudes. We will return to this topic in Sect. 4.

## 4 Discussion

The results presented in Sect. 3 indicate that the differences in the diabatic heating rates (and consequently the corresponding circulations) among the reanalyses are large. These differences arise in part due to differences in the models' representations of tropical deep convection, which is widely recognized as a challenge for global-scale models. The substantial differences in the radiative heating rates may be more of a surprise, and warrant a more detailed analysis.

Radiative heating rates in the 300 hPa to 100 hPa layer are particularly different between MERRA and ERAI (see Fig. 1, middle panel). These differences exist in both LW and SW radiation (see Fig. 3). Figure 4 shows profiles of clear-sky radiative heating and the radiative effects of clouds averaged over the inner tropics ( $5^{\circ}\text{S}$ – $5^{\circ}\text{N}$ ) for MERRA and ERAI. The vertical profile of clear-sky SW radiative heating is very similar in both models. Although there are some differences in the vertical distribution of clear-sky LW radiative heating, the large differences in tropical radiative heating rates evidently arise from differences in the radiative heating impacts of clouds. Cloud SW heating is greater in MERRA than in ERAI from about 300 to 200 hPa, but the opposite is true between 200 and 100 hPa. Similarly, cloud LW heating in MERRA is greater than that in ERAI below 250 hPa, while cloud LW heating in ERAI is greater than that in MERRA (which

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is in fact negative) between 250 and 100 hPa. Although these differences are most pronounced in the inner tropics (where cloud occurrence frequency is largest), the profiles are qualitatively similar when averaged over the entire study domain (30° S–30° N; not shown).

5 Fueglistaler and Fu (2006) and Mcfarlane et al. (2007) calculated the impacts of clouds on radiative heating in the UTLS based on Atmospheric Radiation Measurement (ARM) program observations at two sites in the tropical western Pacific (see also comparison with ERAI by Fueglistaler et al., 2009b). Their calculations showed that cloud impacts on net radiative heating at these two sites were positive up to about  
10 125 hPa. They also showed that peak net cloud radiative heating was shifted upward in several model estimates (including ERAI) relative to the ARM-based profiles; however, the inability of the ARM millimeter cloud radar to resolve optically-thin cirrus clouds introduces large uncertainties in comparisons above 200 hPa.

Yang et al. (2010) used a detailed radiative transfer model and observations of atmospheric composition and clouds (including thin cirrus) to calculate cloud radiative heating in the UTLS between 30° S and 30° N. Their calculations indicate that the impact of clouds on net radiative heating is positive below 100 hPa and negative above 100 hPa, with a positive SW cloud influence throughout the UTLS (peaking near 150 hPa) and a LW cloud influence that transitions from positive below ~ 165 hPa to negative above.  
20 The SW radiative effect of clouds in the UTLS in ERAI is similar to that derived by Yang et al. (2010); however, the LW cloud radiative heating is smaller below 200 hPa and larger above 200 hPa. The MERRA cloud radiative forcing qualitatively resembles that reported by Yang et al. (2010), but with peak SW cloud forcing shifted downward by ~ 100 hPa and the transition from positive to negative LW cloud forcing shifted downward by ~ 50 hPa. Moreover, the maximum in cloud LW cooling is quantitatively much  
25 stronger in MERRA. These differences lead to a mismatch in the sign of the net cloud radiative forcing between MERRA and the other two estimates in the layer between 200 and 100 hPa.

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In addition to latent heating and radiative heating, parameterised mixing contributes to the diabatic heat budgets of reanalyses. This term is generally about an order of magnitude smaller than the other terms and is often ignored; however, local variations in this term can be as large as variations in the radiative terms (see also Flannaghan and Fueglistaler, 2011). The physical realism of these terms is often questionable, as the mixing parameterisations are poorly constrained by observations and are often used to “tune” the model or compensate for errors in other aspects of the model formulation.

Figure 5 shows zonal mean time mean diabatic tendencies from parameterised mixing in CFSR, JRA, and NCEP. MERRA also provides a temperature tendency due to turbulent mixing; however, this term is approximately three orders of magnitude smaller than the radiative and latent heating components, and we have therefore chosen to omit it. ERAI does not provide separate moist physics and vertical mixing terms, though turbulent mixing due to shear-flow instability can be inferred from offline calculations (Flannaghan and Fueglistaler, 2011). Diabatic heating from parameterised mixing is much larger in NCEP than in the other reanalyses, with strong cooling between 100 and 50 hPa and strong warming between 200 and 100 hPa. This pattern implies substantial mixing of the upper troposphere with the lower stratosphere. The pattern in CFSR is similar to that in NCEP, but weaker in amplitude. Both models use a similar vertical diffusion scheme in the free atmosphere, but the diffusion coefficient in CFSR decreases exponentially with pressure while that in NCEP is constant. The relative strength of the turbulent heating in NCEP reinforces many of the unique features of the NCEP net diabatic heat budget, such as the greater vertical depth of the residual component and the cooling located immediately above the tropopause in the inner tropics. The pattern in JRA is similar in the inner tropics, but contains a weak warming near 100 hPa in the subtropics that is not seen in the other reanalyses.

It is certainly interesting from the perspective of troposphere–stratosphere exchange that all three reanalyses show local maxima in parameterised turbulent mixing across the tropical tropopause. Flannaghan and Fueglistaler (2011) have shown that different

formulations of mixing parameterisations may activate in response to different physical processes. In light of known challenges in reconciling temperature observations and general circulation model results in this layer (e.g. Fu et al., 2011; Seidel et al., 2012), it may be worthwhile to analyse the mixing terms in climate model simulations.

As mentioned in Sect. 3, net diabatic heating in JRA is negative through much of the tropics between approximately 60 and 40 hPa. This net diabatic cooling is inconsistent with climatological mean diabatic ascent over the inner tropics, and is attributable to a shortage of LW heating relative to the other reanalyses (Fig. 3, upper panels). JRA has a systematic cold temperature bias in the lower and middle stratosphere, which is related to known deficiencies in the radiation scheme (Onogi et al., 2007). A negative temperature bias is typically associated with a positive bias in LW heating (see discussion in Fueglistaler et al., 2009b); however, this may not be the case if the bias is caused by insufficient absorption of upwelling LW radiation. This situation might result if the LW flux reaching the stratosphere is too small; however, this explanation is inconsistent with a large positive bias in outgoing LW radiation (OLR) at the top of the atmosphere (TOA) in JRA relative to observations (Trenberth et al., 2011). The upward LW flux at the surface is also biased positive in JRA, but the bias in OLR is  $12 \text{ W m}^{-2}$  larger than the bias at the surface. The large difference between the biases at the surface and at the TOA suggests that the radiation scheme in JRA may simulate too little LW absorption in the atmosphere, and provides a plausible explanation for the layer of cooling in the stratosphere.

JMA has recently developed a new 55 yr reanalysis system, which is scheduled for release in mid-2013. This new reanalysis system includes substantial improvements to the radiation scheme that appear to largely eliminate the cold temperature bias in the stratosphere (Ebita et al., 2011). These improvements may also bring the stratospheric diabatic heat budget more in line with both expectations and other reanalyses. In the meantime, users of stratospheric diabatic heating rates from JRA-25/JCDAS should be aware of this difference between JRA and other reanalyses.

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## 5 Conclusions and outlook

We have compared the diabatic heat budgets of the tropical UTLS as simulated by five widely-used reanalyses. The differences in the distributions of net diabatic heating are large, as are the differences in the individual terms. These differences highlight the large uncertainties that remain in the processes that control the heat budget in the tropical UTLS, and have important implications for understanding transport and mixing in this layer. For example, the distribution of diabatic heating in the UT has implications for the transport of tracers from the surface to the stratosphere, while the magnitude of radiative heating in the stratosphere is important for transport and composition in the global stratosphere.

Tracer transport from the surface to the stratosphere typically occurs in two stages: rapid vertical transport from the surface to the UT or TTL in deep convection, followed by slow ascent into the stratosphere. Convection may also penetrate the stratosphere directly, but the impact of this very deep convection on the heat balance remains poorly quantified. As mentioned above, the existence of net diabatic cooling directly above deep convective regions in MERRA represents a barrier to transport into the TTL and LS. Bergman et al. (2013) compared estimates of vertical transport in the Asian monsoon region during boreal summer based on MERRA with estimates based on high-resolution operational analyses from ECMWF and NCEP. They reported that MERRA drastically underestimates vertical transport between 200 hPa and 100 hPa relative to the other datasets. This difference is particularly pronounced when MERRA diabatic heating is used to estimate vertical motion (rather than kinematic vertical velocity). By contrast, the strong net diabatic heating directly above convective regions in ERAI suggests that vertical transport from the surface to the LS may be particularly efficient in ERAI, with ascent through the TTL potentially occurring much more rapidly than in the other reanalyses. As discussed in Sect. 4, these differences between MERRA and ERAI are at least partially attributable to differences in simulations of clouds and their radiative impacts in the UTLS.

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Schoeberl et al. (2012) used winds and heating rates from MERRA, ERAI, and CFSR to drive ensembles of Lagrangian trajectories in the global stratosphere. They then used these trajectories to study interannual variability and trends in stratospheric water vapor over the period 1979–2010. Comparing the results to satellite observations, they reported that rates of vertical ascent in the tropical LS are approximately 30 % too fast when based on ERAI, close to observed when based on CFSR, and approximately 15 % too slow when based on MERRA. These differences in vertical ascent are apparent in the relative strengths of diabatic heating in the LS (Fig. 1), particularly in the LW radiative component (Fig. 3). We will explore the effects of differences in reanalysis diabatic heat budgets on transport and mixing in the tropical UTLS more fully in a follow-up publication.

The diabatic heat budget provides an additional perspective and physical constraint for process evaluation and studies of global change. We stress that more attention should be paid to this budget and its individual terms in the future, and therefore welcome the recent establishment of the Stratospheric Processes and their Role in Climate (SPARC) Reanalysis/analysis Intercomparison Project (S-RIP; <http://www.woa.ees.hokudai.ac.jp/~fuji/s-rip/>).

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S. Fueglistaler**Table 1.** Attributes of the reanalysis forecast models and radiation parameterisations.

	MERRA	ERA-Interim	CFRSR	JRA	NCEP
Period	1979–present	1979–present	1979–2011	1979–present	1948–present
Model grid	$(2/3)^\circ \times 0.5^\circ$	T <sub>L</sub> 255	T382	T106	T62
Model top	0.01 hPa	0.1 hPa	~ 0.266 hPa	0.4 hPa	~ 3 hPa
Model levels	72	60	64	40	28
UTLS $\Delta z^a$	~ 1.1 km	~ 1.1 km	~ 0.88 km	~ 1.3 km	~ 1.8 km
Heating from	6 h forecast	12 h forecast	6 h forecast	6 h forecast	6 h forecast
Long-wave <sup>b</sup>	CLIRAD-LW	RRTM	RRTMG	JMA GSM	GFDL-IR
Short-wave <sup>c</sup>	CLIRAD-SW	6-channel	RRTMG	CCM2	GFDL-SW
Cloud overlap	max/random	max/random	max/random	max/random	random
CO <sub>2</sub>	observed	fixed+trend	observed	fixed	fixed
Ozone <sup>d</sup>	prognostic	climatology	prognostic	offline CTM	climatology
Aerosols	climatology	climatology	climatology	fixed	–
Trace gases	climatology	fixed+trend	climatology	–	–

<sup>a</sup> Average vertical resolution between ~ 200 hPa and ~ 50–60 hPa (Fujiwara et al., 2012).

<sup>b</sup> CLIRAD-LW has been described by Chou et al. (2001), RRTM by Mlawer et al. (1997), RRTMG by Clough et al. (2005), JMA GSM by Sugi et al. (1990), and GFDL-IR by Fels and Schwarzkopf (1975) and Schwarzkopf and Fels (1991).

<sup>c</sup> CLIRAD-SW has been described by Chou and Suarez (1999), the ERA-Interim short-wave parameterisation by Fouquart and Bonnel (1980), RRTMG by Clough et al. (2005), CCM2 by Briegleb (1992), and GFDL-IR by Lacis and Hansen (1974).

<sup>d</sup> The MERRA and CFRSR radiation schemes use ozone concentrations calculated by the forecast model. ERA-Interim and NCEP use zonal and monthly mean climatologies (based on Fortuin and Langematz, 1994 and Fels et al., 1980, respectively). The ozone concentrations used in JRA are produced daily as described by Onogi et al. (2007).

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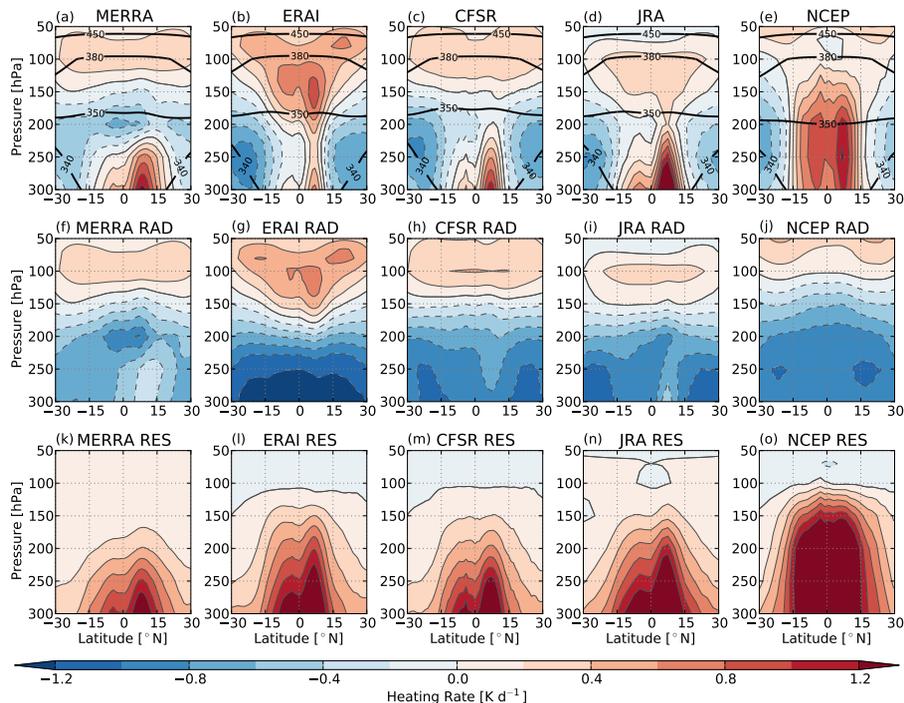
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**Fig. 1.** Top row: zonal mean total diabatic heating averaged over the period 2001–2010 according to the (a) MERRA, (b) ERAI, (c) CFSR, (d) JRA, and (e) NCEP reanalysis datasets. The 340, 350, 380, and 450 K isentropic surfaces from each reanalysis are shown as black contours. Middle row: zonal mean radiative components of diabatic heating averaged over the period 2001–2010. Bottom row: zonal mean residual components of diabatic heating averaged over the period 2001–2010. Residual diabatic heating includes all diabatic heating due to non-radiative physics (moist physics, turbulent mixing, etc.).

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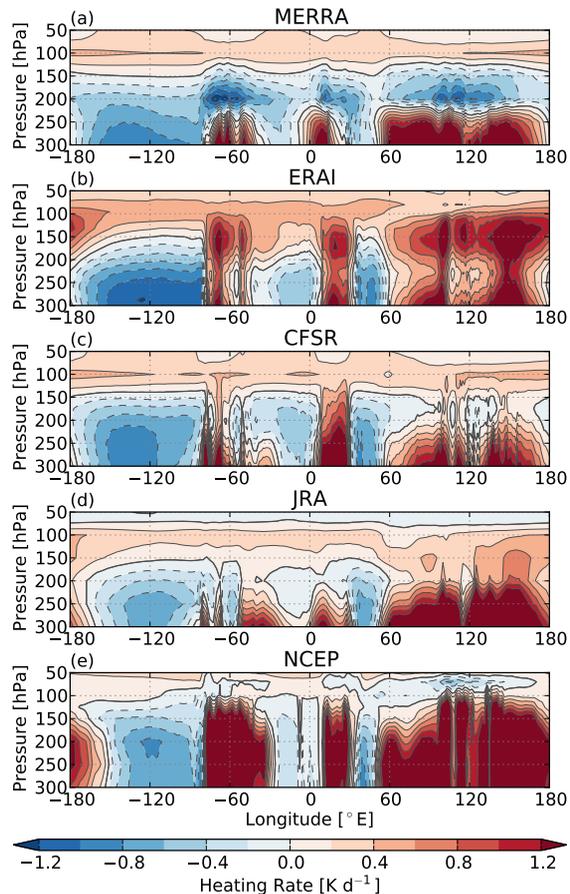
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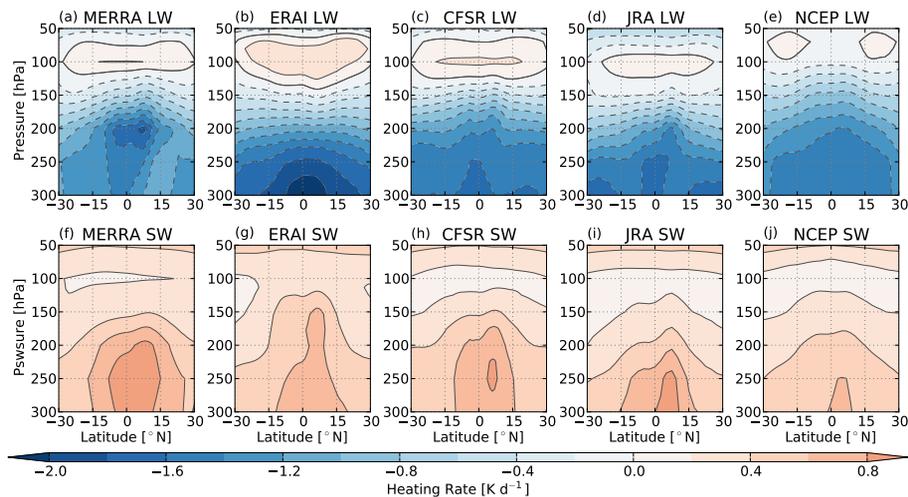
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**Fig. 2.** Longitude-pressure distributions of diabatic heating averaged over  $5^{\circ}$  S to  $5^{\circ}$  N for the period 2001–2010.

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**Fig. 3.** Zonal mean (a–e) long-wave and (f–j) short-wave components of radiative diabatic heating averaged over the period 2001–2010.

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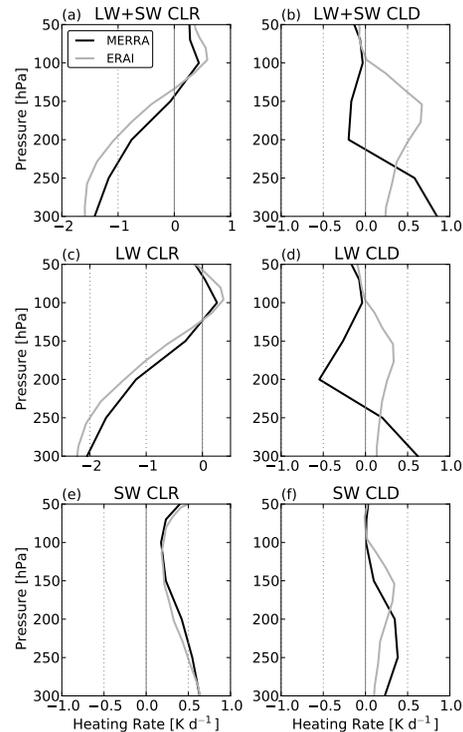
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**Fig. 4.** Profiles of clear-sky (left column) and cloud (right column) radiative heating rates for **(a–b)** long-wave plus short-wave, **(c–d)** long-wave, and **(e–f)** short-wave radiation according to MERRA (black line) and ERAI (grey line). Cloud radiative heating rates are calculated as all-sky minus clear-sky. Profile values have been averaged over 5° S to 5° N for the period 2001–2004.

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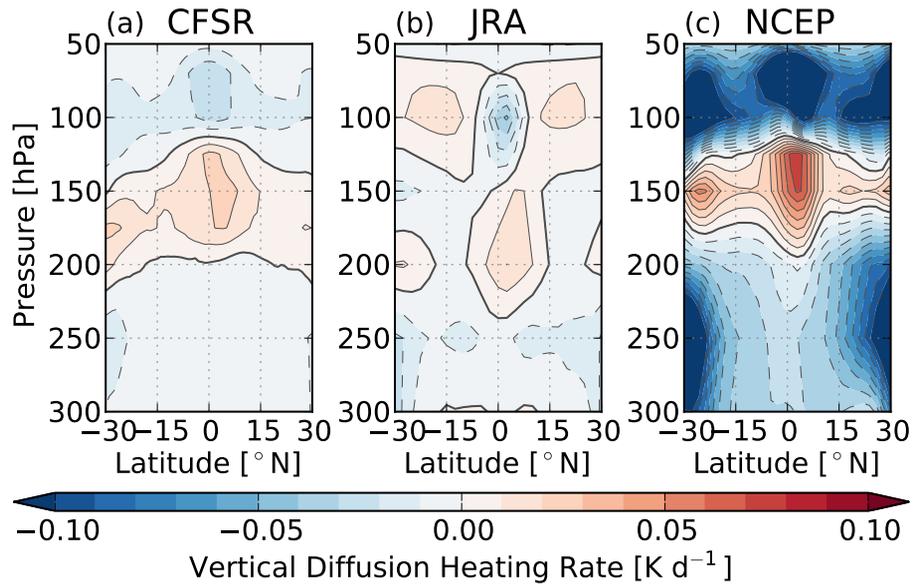
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**Fig. 5.** Zonal mean heating due to parameterised turbulent mixing in **(a)** CFSR, **(b)** JRA, and **(c)** NCEP.

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