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# Observations of I<sub>2</sub> at a remote marine site

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Discussion Paper

Discussion Paper

Discussion Paper

M. J. Lawler et al.

**ACPD** 

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

Title Page

**Abstract** 

Introduction

Conclusions

References

**Tables** 

**Figures** 

 $\triangleright$ 



Back



Full Screen / Esc

**Printer-friendly Version** 



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Inorganic iodine plays a significant role in the photochemistry of the marine boundary layer, but the sources and cycling of iodine are not well understood. We report the first I<sub>2</sub> observations in marine air that is not impacted by coastal macroalgal emissions or sea ice chemistry. The data clearly demonstrate that the very high I<sub>2</sub> levels previously reported for coastal air are not representative of open ocean conditions. In this study, gas phase I<sub>2</sub> was measured at the Cape Verde Atmospheric Observatory, a semi-remote site in the eastern tropical Atlantic, using atmospheric pressure chemical ionization tandem mass spectrometry. Atmospheric I<sub>2</sub> levels typically increased beginning at sunset, leveled off after midnight, and then rapidly decreased at sunrise. There was also a smaller midday maximum in I<sub>2</sub> that was at least partly due to a measurement artifact. Ambient I<sub>2</sub> mixing ratios ranged from < 0.02–0.6 pmol mol<sup>-1</sup> in May 2007 and < 0.03–1.67 pmol mol<sup>-1</sup> in May 2009. The sea-air flux implied by the nighttime buildup of I<sub>2</sub> is too small to explain the observed daytime IO levels at this site. Iodocarbon measurements made in this region previously are also insufficient to explain the observed 1–2 pmol mol<sup>-1</sup> of daytime IO. The observations imply the existence of an unknown daytime source of gas phase inorganic iodine. Carpenter et al. (2013) recently proposed that sea surface emissions of HOI are several times larger than the flux of I2. Such a flux could account for both the nighttime I<sub>2</sub> and the daytime IO observations.

#### Introduction

lodine in the marine atmosphere is ultimately derived from the iodide, iodate, and organically bound iodine contained in seawater. Iodine emitted from the sea surface undergoes rapid photochemical transformations involving both gas phase and heterogeneous reactions (Vogt et al., 1999; von Glasow, 2003; Saiz-Lopez et al., 2008). Iodine chemistry can potentially influence climate through the catalytic destruction of tropospheric ozone and through the formation of aerosols, but the mechanisms and mag-

Discussion Paper

Discussion Paper

Discussion Pape

### **ACPD**

13, 25911–25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page Introduction **Abstract** 

Conclusions

Back

References **Tables Figures** 

Close

Full Screen / Esc

**Printer-friendly Version** 



lodine can be emitted from the sea surface in several different forms. It has been known for some time that O<sub>3</sub> deposited to the sea surface can react with I<sup>-</sup> in the surface seawater to release I2. Rates of I2 release from this process have been estimated at around  $2 \times 10^6$  molec cm<sup>-2</sup> s<sup>-1</sup>, or 2.9 nmol m<sup>-2</sup> d<sup>-1</sup> (Garland and Curtis, 1981). More recent laboratory and modeling work has considered the release of HOI from this process as well, yielding estimates of daytime fluxes of 100 nmol m<sup>-2</sup> d<sup>-1</sup> and 10 nmol m<sup>-2</sup> d<sup>-1</sup> for HOI and I<sub>2</sub>, respectively (Carpenter et al., 2013). That represents an increase of roughly 20x the per-l-atom flux relative to what was previously thought . Because I is ubiquitous in the surface ocean, this process must occur globally and could account for a flux of elemental I on the order of 2Tgyr<sup>-1</sup>. Certain species of coastal macroalgae directly emit copious quantities of I2, particularly while under oxidative stress (Saiz-Lopez and Plane, 2004; Dixneuf et al., 2009; McFiggans et al., 2004; Palmer et al., 2005). In the open ocean, planktonic algae may also release I2, but the open ocean emission rate is not known (Amachi et al., 2000; Jones et al., 2010). Organoiodide compounds such as CH<sub>2</sub>I<sub>2</sub> and CHCII<sub>2</sub> are also emitted from the sea surface. Like I<sub>2</sub>, most of these compounds are rapidly photolyzed, releasing I atoms on timescales of minutes in the daytime marine boundary layer. The most abundant and longest-lived organoiodide compound, methyl iodide, has a photolysis lifetime of several days. The global I release attributed to organoiodides is on the order of 0.4-1 Tg I yr<sup>-1</sup>, with CH<sub>3</sub>I dominating (Jones et al., 2010, and references therein).

Once formed in air, iodine atoms can initiate catalytic ozone destruction cycles such as those below:

$$I + O_3 \rightarrow IO + O_2 \tag{R1}$$

Discussion Paper

**ACPD** 13, 25911–25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Discussion Paper

**Abstract** 

Introduction

Conclusions

References

**Tables** 

**Figures** 













Full Screen / Esc

**Printer-friendly Version** 

Interactive Discussion



$$HOI + hv \rightarrow I + OH$$
 (R3)

5 Net:

$$O_3 + HO_2 + hv \rightarrow 2O_2 + OH \tag{R4}$$

Catalytic destruction of tropospheric ozone by iodine is efficient in part because iodine radicals are relatively unreactive with organic trace gases.

Sea salt aerosols are thought to play a role in Ix cycling in marine air. I<sup>-</sup> present in aerosols can be oxidized to the volatile forms ICI, IBr, and I<sub>2</sub> via reaction with the hypohalous acids HOCI, HOBr, or HOI. The hypohalous acids are generated by gas phase daytime photochemistry (e.g. Reactions 1–2) and can enter aerosols to react with I<sup>-</sup>:

$$HOI + I^- + H^+ \rightarrow I_2 + H_2O$$
 (R5)

lodine radicals can also self-react to form iodine oxides such as  $I_2O_4$  and other larger iodine oxide clusters (McFiggans et al., 2004; Saiz-Lopez et al., 2006). If iodine levels are sufficiently high, the formation of iodine oxides leads to rapid particle nucleation. IOx-driven nucleation events have been observed downwind of  $I_2$ -emitting macroalgae beds. These events contribute to aerosol number and can potentially influence cloud properties in coastal regions (Saunders et al., 2010; Mahajan et al., 2010).

Relatively little is known about the emissions and cycling of iodine over the open oceans, because there are few observations in marine air that are not influenced by macroalgal emissions. Long path DOAS measurements of IO on Cape Verde found daytime IO levels of 1–2 ppt, which are sufficient to contribute significantly to ozone

**ACPD** 

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Discussion Paper

Discussion

Discussion Pape

Introduction

Conclusions

References

Tables

Figures

I₫



•



Back



Full Screen / Esc

Printer-friendly Version

Interactive Discussion



destruction (Read et al., 2008). The source of reactive iodine at CVAO remains an open question. Jones et al. (2010) measured air-sea dihalomethane fluxes in open ocean and upwelling Atlantic waters near Cape Verde. They observed iodocarbons in these waters at levels too low to account for the observed IO at CVAO and proposed that a large  $I_2$  flux might account for the discrepancy. Using a 0-D box, model, they incorporated a constant  $I_2$  flux large enough to achieve the IO levels observed at Cape Verde. This resulted in nighttime  $I_2$  levels of up to 7 pmol mol<sup>-1</sup>. Mahajan et al. (2010) used the same observations and a vertically resolved model with an  $I_2$  source which was allowed to vary. They concluded that an additional MBL source of  $I_2$  may be required to explain the observed IO levels, but that the source may have a strong diel cycle (Mahajan et al., 2010). Mahajan et al. (2012) observed slant column IO densities corresponding to roughly 1 pmol mol<sup>-1</sup> during shipboard measurements in the eastern Pacific.

In this manuscript we report nighttime levels of gas phase  $I_2$  at Cape Verde during measurement campaigns over three weeks in May–June 2007 and one week in May 2009. The only other reported observations of  $I_2$  in marine air have been conducted in regions influenced by emissions from coastal macroalgae or by  $I_2$  production on sea ice surfaces (Saiz-Lopez et al., 2012; Atkinson et al., 2012). These observations provide estimates of  $I_2$  levels and production rates that are more representative of open ocean conditions than previous studies.

## 2 Study site and observational methods

# 2.1 Setting and ancillary observations

The field site was the Cape Verde Atmospheric Observatory (CVAO), located at 16.864° N, 24.867° W on the island of São Vicente. The site is situated about 50 m from a northeast facing coastline, about 10 m above sea level. The island is volcanic in origin and has a steep continental shelf. Meteorological conditions were very similar in the

ACPD

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures













Full Screen / Esc

Printer-friendly Version



two study years, 2007 and 2009. Winds were consistently onshore during both years, and 15 min averaged windspeeds were  $7.67 \pm 1.47 \, \mathrm{m \, s^{-1}}$  (1 std dev) with a  $3.6-10.4 \, \mathrm{range}$  in 2009 and  $6.17 \pm 1.35 \, (2.2-8.7 \, \mathrm{range})$  in 2007. See Carpenter et al. (2011) for a thorough site overview.  $\mathrm{NO_x}$  ( $\mathrm{NO} + \mathrm{NO_2}$ ) was measured by chemiluminescence at a height of 3 m above the ground. The instrument had a detection limit of < 14 ppt for reported 15 min means.  $\mathrm{NO_x}$  levels were typically below 25 ppt in 2007, without a discernible diel pattern. In 2009 the daytime  $\mathrm{NO_x}$  levels were higher and had a distinct diel cycle, with a typical midday  $\mathrm{NO_2}$  maximum of 40–50 pmol  $\mathrm{mol}^{-1}$  and a late afternoon NO maximum of about 15 pmol  $\mathrm{mol}^{-1}$ . The 2009  $\mathrm{NO_x}$  measurements may have been influenced by the site diesel generator. The generator was located about 100 m away from the site in 2007, but it was moved directly adjacent to the downwind side of the site in 2009.  $\mathrm{O_3}$  was measured at 3 m by UV absorption.  $\mathrm{O_3}$  was  $33.4 \pm 5.3 \, \mathrm{nmol \, mol}^{-1}$  in 2007 and  $36.2 \pm 5.0 \, \mathrm{nmol \, mol}^{-1}$  in 2009 (mean  $\pm 1 \, \mathrm{std}$ . dev.). Lee et al. (2010) give an overview of  $\mathrm{NO_x}$  and  $\mathrm{O_3}$  observations at the site.

### 2.2 l<sub>2</sub> detection by chemical ionization mass spectrometry

 $I_2$  detection was carried out by chemical ionization triple quadrupole mass spectrometry using a modified Thermo TSQ Quantum instrument, with procedures similar to those in previously published studies of  $CI_2$ ,  $Br_2$ , and  $I_2$  (Lawler et al., 2009; Finley and Saltzman, 2008). The mass spectrometer was fitted with a  $^{63}$ Ni beta-emitting source, and ambient  $I_2$  molecules were ionized to  $I_2^-$ . The  $I_2^-$  ions were mass-selected in the first quadrupole (Q1), then dissociated by collision with Ar in Q2 (20 eV) to form  $I^-$ , which was in turn mass selected (Q3) and detected by a electron multiplier after impaction on a dynode. Mass transitions for  $ICI^- \rightarrow CI^-$  and  $IBr^- \rightarrow Br^-$  were also monitored.

The instrument background signal was assessed every hour by sampling ambient air which was scrubbed of  $I_2$ . The scrubber was a carbonate-coated plug of glass wool. Automated pneumatic gate valves controlled whether the instrument sampled scrubbed air or unmodified ambient air. This type of scrubber removes  $I_2$  with high

**ACPD** 

13, 25911-25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

l4 ►I

Back Close

Full Screen / Esc

Printer-friendly Version



An  $I_2$  gas standard was generated by flowing  $N_2$  over an  $I_2$  permeation tube in a temperature-controlled PFA housing. This  $I_2/N_2$  flow was subsampled and diluted with additional  $N_2$  to a final sampled mixing ratio of 2–16 pmol mol<sup>-1</sup> in the method of Gallagher et al. (1997).  $I_2$  in the dilution system was only exposed to PFA and PTFE surfaces. The standard gas flowed continuously through a length of PFA tubing up to a pneumatic PTFE valve near the top of the inlet. The valve was kept near the inlet to minimize equilibration time in the tubing. The outdoor tubing was shielded from radiation by opaque insulation to prevent photolysis of the standard  $I_2$ . During regular ambient sampling intervals, the  $I_2$  standard was routed to a container of scrubber material. The output of the  $I_2$  permeation tube was gravimetrically calibrated in the laboratory after the campaign. There were no standards for ICI or IBr.

Instrument sensitivity for  $I_2$  was assessed by performing multipoint calibrations and one-point standard additions. Sensitivity was assumed to vary linearly between calibrations and standard additions. In both 2007 and 2009, the one-point  $I_2$  standards were added on a 4 h schedule on every third day of measurements. Multipoint calibrations were run less frequently but consistently showed linear instrument response in the range of observed  $I_2$ . For these calibrations and single point standards,  $I_2$  in  $I_2$  was added near the front of the inlet.

### 2.3 Sample inlets

The air sampling inlet used in 2007 is the same as that described previously for  $\text{Cl}_2$  observations made at this site (Lawler et al., 2009). The setup is illustrated in Figure 1. Ambient air was drawn through a nominal 2 in (5.1 cm) acrylonitrile-butadiene-styrene (ABS) pipe from a height of 3 m. The flow in this pipe was dynamically controlled at 63 LPM (STP: 273 K, 1 atm) using a butterfly valve (MKS) and mass flow meter (TSI), and PID controller (MKS).

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# **ACPD**

13, 25911–25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

T Abstract

Discussion Paper

Discussion Paper

Discussion Paper

Introduction

Conclusions

References

Tables

Figures

I₫



- 4



Back

Close

Full Screen / Esc

Printer-friendly Version



There were minor differences in the standardization procedure between the two studies.  $I_2$  gas standards were added in scrubbed air in 2007, and in ambient air in 2009. No significant difference in sensitivity has been found for scrubbed vs. unscrubbed air. The sensitivity was assumed to vary linearly between one-point standards, except for obvious step changes in sensitivity.

#### 3 Results

(Fig. 1).

# 3.1 May-June 2007

Measured  $I_2$  ranged from <0.02–0.6 pmol mol<sup>-1</sup> with a regular diel pattern (Fig. 2).  $I_2$  climbed over the course of the night and reached its highest values either shortly before dawn or a couple of hours before.  $I_2$  levels always dropped dramatically at dawn, but usually remained at detachable levels. Surprisingly,  $I_2$  levels increased during the daytime, reaching a peak around midday and then declining again before nighttime.  $I_2$  levels did not covary with  $O_3$  levels. The daytime 'blank'  $I_2$  signals (scrubbed air) showed large  $I_2$  signals that closely tracked the daytime  $I_2$  increase, but blanks were consistently low and fairly constant over the night (Fig. 4).

# **ACPD**

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Discussion Paper

Discussion Paper

Discussion Paper

Introduction

Conclusions

References

Tables

Figures









Back



Full Screen / Esc

Printer-friendly Version



The ICI sample and blank signals showed similar diel cycles and were statistically indistinguishable from one another. The two ICI mass transitions observed (162  $\rightarrow$  35 and 164  $\rightarrow$  37) had a sample signal ratio close to 1, rather than the 3 ratio expected for the two chlorine isotopes. These observations indicate that ICI was not the dominant species observed at these transitions. The sample and blank signals for IBr also matched one another, and IBr remained below detection. The actual detection limits for ICI and IBr were not assessed, but detection limits for Cl<sub>2</sub>, BrCl, Br<sub>2</sub>, and I<sub>2</sub> for this instrument are in the range of 0.1–2 pmol mol<sup>-1</sup>, and ICI and IBr detection limits are also expected to be in this range.

# 3.2 May 2009

The 2009  $I_2$  levels were very similar to those in 2007, ranging from <0.03–1.67 pmol mol<sup>-1</sup> over one week of observations (Fig. 3). The diurnal pattern was also very similar, with a nighttime maximum and a smaller daytime peak. There was some day-to-day variability in the absolute levels, with the highest levels occurring late in the night on days 141 and 142. Late on day 141 there was a sudden reduction in wind-speed and shift to more northerly flow. Unfortunately the meteorological instruments were not operating on day 142, when the very highest  $I_2$  levels were measured. There was some difference in the instrument response to scrubbed air between the two deployments. In the 2009 deployment,  $I_2$  blank signals remained low throughout the day, and did not exhibit the daytime peak observed in the ambient measurements.

### 3.3 Macroalgal emissions

This is the first study to examine the behaviour of atmospheric  $I_2$  in an environment not impacted by iodine-emitting macroalgae. To confirm this, macroalgae were collected from near-site tidal pools and held before the instrument inlet in 2007. No enhancement in  $I_2$  levels was observed for any of the few species found. A container of coastal surface seawater held before the inlet also resulted in no enhancement of  $I_2$  signal.

ACPD

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

14

►I

- 4

•

Close

Back

Full Screen / Esc

Printer-friendly Version



#### 4 Discussion

### 4.1 Nighttime I<sub>2</sub> emission rates

At night, I<sub>2</sub> mixing ratios increase steadily after dusk, then level off around midnight (Fig. 5). This suggests I<sub>2</sub> emissions that are relatively constant during the buildup, and negligible later in the night. The Cape Verde observations indicate an average nighttime increase of about 0.17 pmol mol<sup>-1</sup>. Assuming the sea surface is the source and that there are no atmospheric losses, this is a rate of roughly 0.5 pmol mol<sup>-1</sup> d<sup>-1</sup> over the period of constant increase. This corresponds to a sea-to-air flux of 30 nmol m<sup>-2</sup> d<sup>-1</sup>, assuming that emissions are diluted into a 1000 m boundary layer and there are no losses. This is about three times the rate estimated in the Carpenter et al. (2013) study.

## 4.2 I<sub>2</sub> as a source for daytime IO

Midday maximum IO levels ranging from 1–2 pmol mol<sup>-1</sup> were observed by long path DOAS at the CVAO site during May 2007 (Read et al., 2008). Mahajan et al. (2010) and Jones et al. (2010) proposed sea surface emissions of I<sub>2</sub> as the principal source of iodine at Cape Verde because air/sea fluxes of organoiodide compounds in the eastern tropical Atlantic were considerably lower than required to account for the observed IO. Model simulations showed that a constant I<sub>2</sub> flux of 170–320 nmol I<sub>2</sub> m<sup>-2</sup> d<sup>-1</sup> was needed to explain the observed IO (Jones et al., 2010). However, this scenario also led to a nighttime maximum of approximately 7 pmol mol<sup>-1</sup> of I<sub>2</sub> and a resultant spike in IO at sunrise, neither of which are evident in the Cape Verde data. Our I<sub>2</sub> observations do not support such a large nighttime source. Mahajan et al. (2010) implemented a diurnally varying I<sub>2</sub> source tuned to best replicate the observed IO levels. This source

ACPD

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

Discussion Paper

Discussion Paper

References

Tables

Figures

14







Back



Full Screen / Esc

Printer-friendly Version



### 4.3 Observations of daytime l<sub>2</sub>: positive artifact

Midday I<sub>2</sub> levels of 0.2 pmol mol<sup>-1</sup> are much larger than would be predicted by the nocturnal emission rate. The noon photolytic lifetime of I<sub>2</sub> under CVAO conditions is roughly 5 s. An I<sub>2</sub> flux large enough to support these levels would result in higher IO levels than observed (see Modeling section below). These considerations, and the fact that the daytime blank I<sub>2</sub> signals closely tracked the ambient signals in the 2007 study (see Results section) lead us to conclude that the daytime measurements were affected by a positive experimental artifact.

We hypothesize that the daytime I2 is generated via heterogeneous reactions of a photochemically produced oxidant with I on the walls of the instrument inlet. This would require the oxidant to have the following characteristics; (1) to occur at levels of several pmol mol<sup>-1</sup> (or greater), comparable to the measured I<sub>2</sub>, (2) to be present at significant levels only during daytime, and (3) to be transmitted at least partially through the carbonate scrubber. The first consideration eliminates OH and the second eliminates O<sub>3</sub> as likely candidates. The hypohalous acids HOI, HOBr, and HOCI are photochemically generated daytime oxidants that are present in air at pmol mol<sup>-1</sup> levels, and are capable of oxidizing I<sup>-</sup>. HOI can produce I<sub>2</sub> directly upon reaction with I<sup>-</sup> in solution, e.g. (Allen and Keefer, 1955; Vogt et al., 1999). HOBr and HOCl might be expected to initially form IBr and ICI from reaction with I<sup>-</sup>, and further reaction would be required to form I<sub>2</sub>. However, the mixed halogens were not observed in this study. Previous studies have observed the production of Br<sub>2</sub> from HOBr and HOCl even on relatively clean inlets, apparently without formation of BrCl (Neuman et al., 2010; Lawler et al., 2011). Analogously, inlet I<sub>2</sub> production without concomitant ICl or IBr production may be possible.

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# **ACPD**

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Discussion Paper

Discussion Paper

Introduction

Conclusions

References

Tables

Figures









Back



Full Screen / Esc

Printer-friendly Version

Interactive Discussion



The difference in behaviour of the daytime blanks between the two deployments is likely due to the changes in inlet configuration. In 2007, scrubbed air was exposed to the same 3 m long, 2 in ID flow path as the ambient air so the walls of the pipe were coated with ambient aerosols. In this configuration, daytime blanks covaried with the ambient I<sub>2</sub> signal. In 2009, scrubbed air was exposed only to the last portion of the inlet, not to the large 6m long 6 ID pipe. In this configuration, the daytime blanks remained low. We speculate that this is because ambient aerosols were deposited near the turbulent sample inlet region of the large pipe, and the scrubbed air encountered only the relatively clean last stage of the inlet.

### 5 Modeling

Four model simulations (BASE, FLAT, PHOTO, and HOI, as defined below) were conducted to investigate the possible sources of reactive iodine at Cape Verde, given the constraints of observed daytime IO and nighttime I2 (Fig. 6). The one-dimensional chemical transport model THAMO was used for these simulations (Saiz-Lopez et al., 2008). The model includes gas phase reactive halogen, HO<sub>x</sub>, NO<sub>x</sub>, and hydrocarbon chemistry, including a module treating ultrafine particle formation by coagulation of iodine oxides. It also accounts for recycling of reactive halogens through marine aerosols. In the past, this model has been used to study IO, HO<sub>x</sub>, and HCHO observations at Cape Verde (Mahajan et al., 2010, 2011). The model reaction scheme and structure have been described in detail previously (Mahajan et al., 2010; Saiz-Lopez et al., 2008). For all the simulations, the model was constrained with observations of O<sub>3</sub> (Carpenter et al., 2011; Lee et al., 2010),  $HO_{\nu}$  (Whalley et al., 2010),  $NO_{\nu}$  (Lee et al., 2009),  $CH_{\Delta}$ and HCHO (Mahajan et al., 2011), and NMHCs (Read et al., 2009) from CVAO. The vertical mixing was calculated using the measured wind speed (Carpenter et al., 2010), using a description detailed by Saiz-Lopez et al. (2008). The model was allowed to run for ten days, usually reaching steady state in 3-4 days. Day 9 of each simulation was used for comparisons between individual scenarios.

ACPD

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

<u>Re</u>ferences

Tables

Figures

I₫

►I

- 4



Back



Full Screen / Esc

Printer-friendly Version



Discussion Pa

I<sub>2</sub> in the MBL

M. J. Lawler et al.

**ACPD** 

13, 25911-25937, 2013

Title Page

Abstract Introduction

Conclusions References

**Tables** 

Figures

14 1

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



In the BASE simulation, standard model chemistry was employed and average sea-air fluxes of halocarbons measured by Jones et al. (2010) near the Cape Verde site were included:  $CH_2I_2$  (13.0 nmol m<sup>2</sup> d<sup>-1</sup>),  $CH_2IBr$  (10.9 nmol m<sup>2</sup> d<sup>-1</sup>),  $CH_2ICl$  (16.2 nmol m<sup>2</sup> d<sup>-1</sup>d),  $CH_3I$  (48.5 nmol m<sup>2</sup> d<sup>-1</sup>),  $C_2H_5I$  (4.1 nmol m<sup>2</sup> d<sup>-1</sup>), and 1- $C_3H_7I$  (0.9 nmol m<sup>2</sup> d<sup>-1</sup>). This simulation significantly underpredicts both IO and  $I_2$  observations made at the site (Fig. 6). This result implies that there are additional important sources of reactive iodine in this environment, or that models currently underestimate the rate of aerosol recycling of reactive iodine.

In the FLAT simulation, a constant  $I_2$  flux of 14.3 nmol m<sup>2</sup> d<sup>-1</sup> from the ocean surface was added to the MBL, in addition to the halocarbon flux already present. This was intended to represent an  $I_2$  source from  $O_3$  deposition to the ocean surface, or a hypothetical biological background source. The flux was tuned to achieve nighttime  $I_2$  levels comparable to observations ( $\sim 0.2-0.3\,\mathrm{pmol\,mol^{-1}}$ ). The  $I_2$  profile matched the nighttime  $I_2$  observations well, but the flat daytime  $I_2$  source was not sufficient to achieve the mean IO levels observed during the day. Modeled IO was  $\sim 0.8\,\mathrm{pmol\,mol^{-1}}$  during the daytime, compared to observed levels of  $\sim 1.5\,\mathrm{pmol\,mol^{-1}}$ . Daytime  $I_2$  levels remained very low, peaking at  $0.02\,\mathrm{pmol\,mol^{-1}}$ .

In the PHOTO simulation, the prescribed flux of  $I_2$  was retained for the nighttime, and an additional  $I_2$  flux was included in the daytime. This was intended to simulate a light-dependent  $I_2$  source, which could be due to photochemical reactions in aerosols or at the sea surface, or due to daytime biological production in the sea surface. The daytime source was tuned to achieve average observed daytime IO levels. The total prescribed  $I_2$  flux reached a maximum of 125 nmol  $I_2$  mol  $I_2$  most a tenfold increase compared to the nighttime flux necessary to reproduce the observed  $I_2$ . This maximum flux was reached at about 09:20 h and was set back to the constant (nighttime) flux by shortly after 16:00 (Fig. 7). This simulation matches the daytime IO and nighttime  $I_2$  observations reasonably well and predicts a midmorning  $I_2$  peak (Fig. 6). The model still does not reproduce daytime  $I_2$  levels as high as observed, particularly in the afternoon. The model predicts midmorning  $I_2$  levels over half as large as observed. Reproduction

Discussion Paper

of the observed afternoon I<sub>2</sub> levels using an extra source of I<sub>2</sub> from the surface would result in a late evening peak in IO. Such a peak was not observed. To summarize, the daytime IO and nighttime I<sub>2</sub> may be explained by the addition of a time-varying surface I<sub>2</sub> source, but the observed daytime IO and I<sub>2</sub> cannot simultaneously be explained simply by an additional source. We think this lack of agreement is most likely explained by the positive measurement artifact in the daytime I<sub>2</sub>.

In the HOI simulation, sea-air fluxes of both HOI and  $I_2$  were prescribed without modification or optimization after Carpenter et al. (2013). Their sea-air fluxes of HOI are highest at night and the  $I_2$  fluxes are highest during the day due to the different air-sea gradients of the two species between night and day. In the daytime,  $I_2$  levels are extremely low due to rapid photolysis and no known gas phase source. HOI, on the other hand, photolyzes more slowly and can be formed in the gas phase. HOI is also thought to have a significant aerosol sink during the nighttime, as opposed to  $I_2$ . This simulation also does a good job representing both the daytime IO and the nighttime  $I_2$  mixing ratios, but yielding slightly higher daytime IO and slightly lower nighttime I2 than the PHOTO simulation. This simulation does not predict a morning  $I_2$  increase. The HOI case predicts the highest levels of ICI of all the runs, up to 6 pmol mol<sup>-1</sup>.

### 6 Discussion: Ix sources and halogen cycling

The modeling scenarios which best represent the observed inorganic iodine at CVAO are the PHOTO and HOI cases. The laboratory observations of HOI production from seawater ozonolysis provide a strong case that this process should be considered in MBL halogen chemistry. The good agreement between the lab-estimated fluxes and the observational data supports the idea that sea surface-derived HOI is the major source of gas phase inorganic iodine at CVAO. This is therefore also likely the case for all ocean regions not impacted by unusually strong localized emissions (such as near macroalgae beds). The photolytic I<sub>2</sub> source considered in the PHOTO case would require some novel chemistry and possibly complex kinetics to generate the peculiar

ACPD

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

I◀

\_\_\_\_

4



Close

Back

Full Screen / Esc

Printer-friendly Version



Both the PHOTO and HOI cases generate at least 3 pmol mol<sup>-1</sup> of ICI at night. This is a consequence of HOI uptake in particles followed by reaction with Cl<sup>-</sup>. ICI was not detected either in 2007 or in 2009, despite monitoring the relevant mass transitions for several days in each case. There is no reason to suspect that the instrument was not similarly sensitive to ICI as to I<sub>2</sub>, despite the lack of a specific calibration standard for ICI. Similarly, BrCl has not been observed at CVAO despite evidence for active bromine cycling (Lawler et al., 2009; Read et al., 2008). These observations strongly suggest that current models overpredict the conversion of HOBr and HOI to the interhalogens BrCl and ICI in aerosols. It remains unclear whether all hypohalous acids (including HOCI) undergo significant losses to organic species in particles, or whether the equilibrium reactions among the various dissolved halogen species need to be reexamined.

#### 7 Conclusions

We report the first MBL  $I_2$  observations in marine air that is not impacted by coastal macroalgal emissions or sea ice chemistry. The data clearly demonstrate that the very high  $I_2$  levels previously reported for coastal air are not representative of open ocean conditions. The very low observed nighttime levels provide an upper bound for  $I_2$  production by reaction of  $O_3$  on the surface ocean in this region. The "dark" processes producing  $I_2$  at night are too slow to explain the levels of IO observed in the day-time, given known iodine recycling mechanisms. The sea-air flux of HOI generated by ozonolysis of seawater and recently proposed by Carpenter et al. (2013) could explain the observations. If this is the case, then IO levels similar to those at CVAO should occur over most of the world oceans.  $O_3$ -stimulated release of iodine from the sea surface induces catalytic ozone destruction, limiting the lifetime of  $O_3$  over the oceans. The impacts of this chemistry on the evolution of continental outflow in the marine boundary layer should be investigated further. The apparent lack of significant levels of ICI

**ACPD** 

13, 25911-25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Discussion Paper

Discussion Paper

Introduction

Conclusions

References

Tables

Figures

-14



4



Back



Full Screen / Esc

Printer-friendly Version



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20

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13

Discussion

Discussion Paper

Discussion Paper

Discussion

**ACPD** 

13, 25911-25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

T Abstract

Introduction

Conclusions

References

Tables

Figures

-14

►I

- 4



Back

Close

Full Screen / Esc

**Printer-friendly Version** 

Interactive Discussion



Discussion Pape

[■

Back



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13, 25911–25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page **Abstract** Introduction Conclusions References

> **Tables Figures**

> > $\triangleright$

Close

Full Screen / Esc

**Printer-friendly Version** 

Discussion



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13, 25911–25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Introduction **Abstract** 

Conclusions References

> **Tables Figures**

[■  $\triangleright$ 

Back

Full Screen / Esc

Close

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**ACPD** 

l<sub>2</sub> in the MBL

13, 25911–25937, 2013

M. J. Lawler et al.

Title Page Introduction **Abstract** 

Conclusions References

**Tables Figures** 

[■  $\triangleright$ 

Back Close

**Printer-friendly Version** 

Full Screen / Esc

Interactive Discussion



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Discussion Paper

Discussion Paper

Discussion Paper

**Discussion Paper** 

13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

I<sup>4</sup>



M





Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Discussion Paper

Discussion Paper

Discussion Paper

Full Screen / Esc

Close

Back

**Printer-friendly Version** 

Interactive Discussion



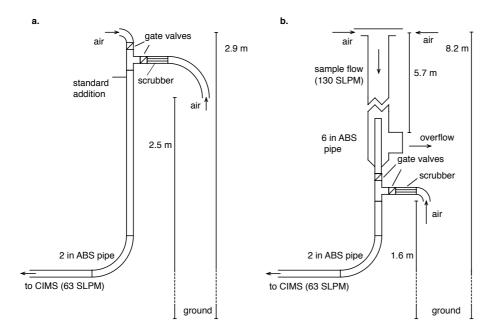


Fig. 1. (a) Sampling inlet used May-June 2007. (b) Sampling inlet used May 2009.

**ACPD** 13, 25911–25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page **Abstract** Introduction Conclusions References **Tables Figures** ┫◀ M



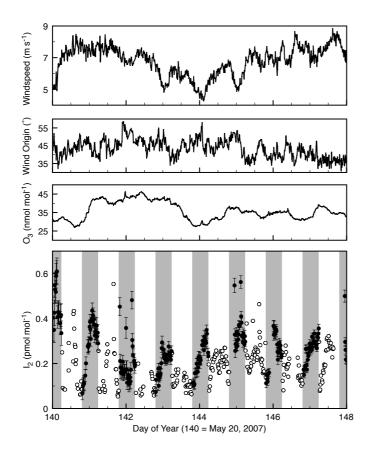


Fig. 2. Wind speed, wind direction, O<sub>3</sub> mixing ratios, and I<sub>2</sub> mixing ratios at CVAO during May 2007 time at CVAO. I<sub>2</sub> error bars are one standard error, and open circles daytime points which are considered upper limits. The dark vertical bars represent nighttime hours.

13, 25911–25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page **Abstract** Introduction Conclusions References **Tables Figures** [■  $\triangleright$ Close Back

**Discussion Paper** 

Full Screen / Esc

Back



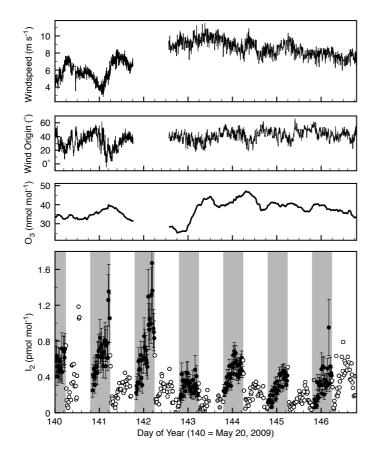


Fig. 3. Wind speed, wind direction, O<sub>3</sub> mixing ratios, and I<sub>2</sub> mixing ratios at CVAO during one week in May 2009. I<sub>2</sub> error bars are one standard error, and open circles are daytime points which are considered upper limits. Grey filled circles are points below the detection limit. The dark vertical bars represent nighttime hours. The gap in non-l<sub>2</sub> data about day 142 was due to power loss to some of the instruments.

**ACPD** 

13, 25911–25937, 2013

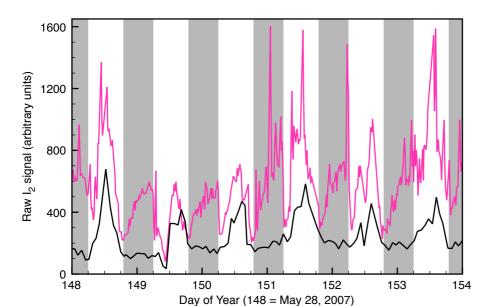
l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page **Abstract** Introduction Conclusions References **Tables Figures** I M

Full Screen / Esc

Close



**Fig. 4.** Example sample (magenta) and blank (black) I<sub>2</sub> signal during May 2007.

13, 25911-25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

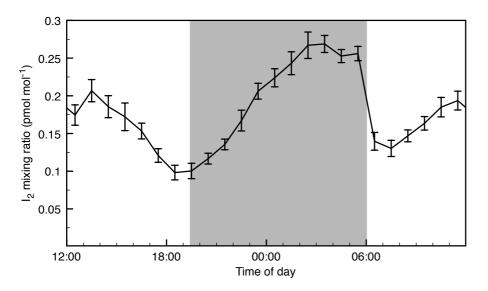
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■ ► Back Close

Full Screen / Esc

Printer-friendly Version





**Fig. 5.** Mean hourly binned I<sub>2</sub> mixing ratios measured over three weeks at CVAO in 2007. Error bars are one standard error. Shaded area represents nighttime hours. The mean observed solar cycle over the measurement period was used to assign sunrise and sunset.

13, 25911-25937, 2013

### I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures









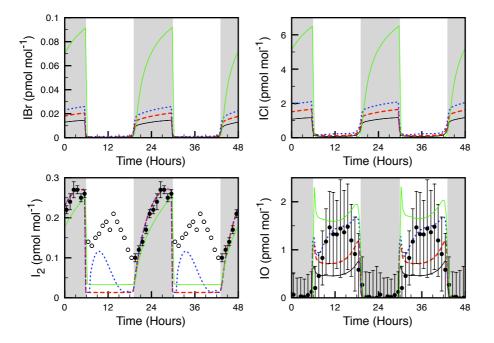






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**Fig. 6.** THAMO box model results for four simulations. BASE (halocarbons only): black solid line. FLAT (constant  $I_2$  flux): red dashed line. PHOTO (enhanced daytime  $I_2$  source): dotted blue line. HOI (added flux of HOI): green solid line. Average  $I_2$  and IO observations plotted as circles with one standard error bars. The upper limit daytime  $I_2$  observations are plotted as open circles. Dark shading represents nighttime hours.

13, 25911-25937, 2013

l<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

rigules





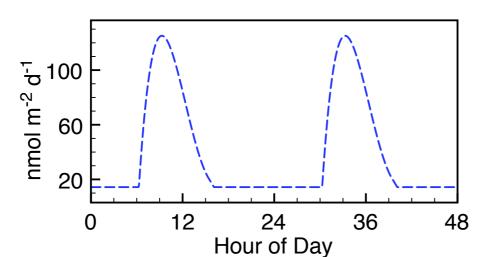






**Printer-friendly Version** 





**Fig. 7.** Flux of I<sub>2</sub> added in PHOTO model run.

13, 25911-25937, 2013

I<sub>2</sub> in the MBL

M. J. Lawler et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures















Printer-friendly Version

