

The regime of desert  
dust episodes in the  
Mediterranean

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This discussion paper is/has been under review for the journal Atmospheric Chemistry and Physics (ACP). Please refer to the corresponding final paper in ACP if available.

# The regime of desert dust episodes in the Mediterranean based on contemporary satellite observations and ground measurements

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Received: 9 March 2013 – Accepted: 24 May 2013 – Published: 17 June 2013

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Published by Copernicus Publications on behalf of the European Geosciences Union.

**ACPD**

13, 16247–16299, 2013

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## Abstract

The regime of desert dust (DD) episodes over the broader Mediterranean basin is studied for the period 2000–2007. The novelty of this work lies in its complete spatial coverage of the region. An objective and dynamic algorithm has been set up, which uses daily measurements of various aerosol optical properties taken by different satellite databases, enabling the identification of DD episodes and their classification into strong and extreme ones. The algorithm's performance was tested against surface based (in situ) Particulate Matter (PM) and (columnar) sun-photometric AERONET measurements from stations distributed across the Mediterranean. The comparisons have shown the reasonable ability of the algorithm to detect the DD episodes taking place within the study region. The largest disagreements with PM data were found in summer and western Mediterranean, when African dust transport has a great vertical extent that cannot be satisfactorily captured by surface measurements.

According to our results, DD episodes in the Mediterranean basin are quite frequent (up to 11.4 episodes/year) while there is a significant spatial and temporal variability in their frequency of occurrence and their intensity. Strong episodes occur more frequently in the western Mediterranean basin whilst extreme ones appear more frequently over central Mediterranean Sea areas. Apart from this longitudinal variation, there is a predominant latitudinal variability in both frequency and intensity, with decreasing values from south to north. A significant seasonal variation was also found for the frequency of DD episodes, with both strong and extreme episodes being more frequent during summer in the western Mediterranean basin, but during spring in its central and eastern parts. In most cases (> 85 %) the Mediterranean dust episodes last a bit longer than a day, although their duration can reach 6 days for strong episodes and 4 days for extreme episodes. A noticeable year by year variability was also found, especially for the frequency of the episodes. The spatial and temporal patterns of Mediterranean DD episodes can be explained based on surface pressure and precipitation spatio-temporal distribution patterns over the study region, as well as by the year by

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year variability of North Atlantic Oscillation (NAO). In this context, a decreasing frequency of appearance of DD episodes over the Mediterranean basin has been revealed over the period 2000–2007, especially over land surfaces, in line with decreasing NAO Index over the same period. Our findings demonstrate the reasonable ability to detect desert dust outbreaks in the Mediterranean basin from satellites.

## 1 Introduction

Desert dust aerosols are coarse particles with size ranging from 1  $\mu\text{m}$  to 5  $\mu\text{m}$  (Tanré et al., 2001), mainly produced by natural processes and only to a small extent (< 10 %) by agricultural activities (Tegen et al., 2004). They affect climate since they perturb the radiation budget of the Earth-atmosphere system interacting (through scattering and absorption) mainly with solar (shortwave, SW) but also with terrestrial (longwave, LW) radiation, producing, respectively, cooling and warming of the Earth-Atmosphere, either as planetary cooling (e.g. Christopher and Jones, 2007; Xia and Zong, 2009) or warming (e.g. Hatzianastassiou et al., 2004; Papadimas et al., 2012), depending also on surface albedo (Osborne et al., 2011; Yang et al., 2009).

Because of its high loadings and large radiative impacts, dust should be considered in climate and weather studies. More specifically, it has been shown in literature that dust can affect components of the hydrological cycle (Lau et al., 2006; Mallet et al., 2009), cloud properties (e.g. Huang et al., 2006a,b) and precipitation (e.g. Rosenfeld et al., 2001; Hui et al., 2008). Moreover, it has been documented that dust modifies sea surface temperature (e.g. Foltz and McPhaden, 2008) and ocean productivity (e.g. Neff et al., 2008) while having adverse health effects (e.g. Perez et al., 2008; Karanasiou et al., 2012 and references therein). All these impacts of dust along with those on radiation become maximum under conditions of extreme loadings in the atmosphere, namely in case of dust episodes (or events, Gkikas et al., 2010, 2011; Benas et al., 2011).

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The main sources of dust particles are located in the arid and semi-arid regions of the planet, especially in the Northern Hemisphere (Sahara, Arabian Peninsula, Gobbi, Taklamakan deserts) and less in the Southern Hemisphere. The deserts of Sahara and Arabian Peninsula have been identified, based on satellite measurements, as the major dust sources areas over the globe (Prospero et al., 2002; Washington et al., 2003) with the most intense production of dust particles being recorded in Chad, W. Sahara and southern Algeria (Middleton and Goudie, 2001; Barkan et al., 2004).

Aerosol optical depth (AOD) is a realistic measure of columnar aerosol loading and can appropriately describe its spatial and temporal distribution and variability. Many researchers (e.g. Remer et al., 2008; Zhang and Reid, 2010) used satellite observations, taking advantage of their geographically wide view, in order to describe globally the spatial distribution of AOD. According to their results, it is evident that significant dust aerosol loads, apart from being observed over the continental source regions, are also transported over oceanic areas, namely towards the Atlantic Ocean from the north African deserts (e.g. Huang et al., 2010) and to Pacific Ocean from Asian deserts (e.g. Eguchi et al., 2009).

The broader Mediterranean basin is one of the most affected worldwide regions by dust transport. This is because of its proximity to Sahara desert, southwards, and to the Middle East and Arabian Peninsula deserts to the east-southeast (Fig. 1). Dust aerosol over the Mediterranean, in case of intense low-pressure systems, can travel long enough to reach the northern parts of European continent (Papayannis et al., 2008; Barkan and Alpert, 2010; Bègue et al., 2012) and even up to northern Europe (e.g. Ansmann et al., 2003; Klein et al., 2010) as well as in the Arctic (Barkan and Alpert, 2010). Dust transport towards the Mediterranean is characterized by a seasonal cycle (e.g. Moulin et al., 1998; Querol et al., 1998 and 2009; Rodríguez et al., 2001; Escudero et al., 2005; Papadimas et al., 2008; Pey et al., 2013) being more evident in its eastern parts in winter and spring, in its central parts during spring and in the western parts in summer. This geographical (longitudinal) shift is driven by the prevailing synoptic conditions (Gkikas et al., 2012). Moreover, it has been documented

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that the phase of North Atlantic Oscillation (NAO) is a determinant factor for dust abundance in the Mediterranean (Moulin et al., 1997; Pey et al., 2013). Apart from its horizontal movement, dust is also lifted vertically to a significant extent. According to ground/satellite measurements, dust particles have been indentified up to 9–10 km in the Mediterranean atmosphere (e.g. Gobbi et al., 2000; Mona et al., 2006) although desert dust layers mainly extend between 1.5 and 6.5 km (Papayannis et al., 2005). It has been also found that the most intense Mediterranean dust transport is observed at around 3 km (Kalivitis et al., 2007).

Apart from studies focusing on the spatial and temporal distribution of dust loads in the Mediterranean basin, dust episodes or events have been also studied but only on an individual basis (e.g. Pérez et al., 2006) or over a few years period (e.g. Toledano et al., 2007) and for specific locations (e.g. Kubilay et al., 2003; Meloni et al., 2007). Nevertheless, a complete and comprehensive study dealing with the identification of dust episodes in the whole Mediterranean basin, and the determination of their spatio-temporal characteristics, is still missing to date. A study of this kind has been conducted by Gkikas et al. (2009), based on a developed algorithm, but dealt with the entire particulate matter, i.e. including other aerosol types and not specifically dust. In the present study, this algorithm is further developed and improved, aiming to characterize, identify and study the regime of dust episodes, solely.

The main scope of the present paper is to describe the characteristics of desert dust episodes over the broader Mediterranean basin. Its novelty is that the dust episodes are homogeneously determined over the entire Mediterranean basin, thanks to satellite (MODIS, EP-TOMS, OMI) data used, for a relatively long period (7 yr, from March 2000 to February 2007) thus providing a first realistic climatological-like database of the episodes. The identification of episodes is made with an objective and dynamic algorithm, set up here, which uses aerosol optical properties derived from different satellite platforms (Sect. 2). The algorithm is first validated (Sect. 3) through comparison against accurate ground-based: (i) particulate matter ( $PM_{10}$ ) concentration measurements, and (ii) aerosol optical properties from AEROSOL ROBOTIC NETWORK (AERONET)

stations which are located within the study area. Subsequently, the main characteristics of desert dust episodes, namely their frequency, intensity and duration are examined at the local scale (pixel-level), by means of geographical distributions (Sect. 4.1), as well as on a mean regional basis (Sect. 4.2). Finally, the pathways of air masses associated with the Mediterranean dust episodes are determined based on calculated back trajectories (Sect. 4.3) before drawing conclusions in Sect. 5.

## 2 Data and methodology

### 2.1 MODIS Terra

MODIS Terra Level-3 daily gridded atmospheric data product (MOD08\_D3) acquired from the MODIS web site <ftp://adsweb.nascom.nasa.gov/> is used in our analysis. Since February 2000 the MODIS instrument on board the Terra satellite, with daytime equator crossing time at 10:30 a.m. and 2330 km viewing swath, providing almost daily global coverage, has been continuously acquiring measurements at 36 spectral bands between 0.415 and 14.235  $\mu\text{m}$  with spatial resolution of 250 m, 500 m and 1000 m. MODIS-Terra data were chosen here rather than Aqua ones first because of consistency with our previous studies (Gkikas et al., 2009, 2012) and second in order to cover a longer period (starting from 2000). Nevertheless, it should be noted that Aqua data are already used in an extension of the present study and algorithm relying only to A-Train satellites and emphasizing on the vertical profile of dust episodes. The retrieval of MODIS aerosol data is performed by special algorithms (see e.g. Kaufman et al., 1997, 2001; Tanré et al., 1997; Levy et al., 2003; Remer et al., 2005), which are different over land and ocean, because of their different surface characteristics, and are continuously improved. The MODIS derived aerosol properties have been extensively validated against AERONET Sun-photometer measurements (e.g. Remer et al., 2008; Levy et al., 2010). The accuracy of algorithm for MODIS Collection 005 AOD is  $\pm 0.05 \pm 0.15 \times \text{AOD}$  over land (Levy et al., 2010), and  $\pm 0.03 \pm 0.05 \times \text{AOD}$  over ocean

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(Remer et al., 2002). More specifically, over the Mediterranean basin, the MODIS Collection 004 AOD values are biased high by 0.06 with respect to AERONET values (Papadimas et al., 2009).

Recently, the Collection 005 (C005) MODIS aerosol data were released, derived from a significantly improved retrieval algorithm, being in better agreement with surface-based AERONET measurements (Levy et al., 2007; Remer et al., 2008). This improvement was also seen over the Mediterranean basin (Papadimas et al., 2009, bias of C004 AOD values has disappeared and correlation coefficient with AERONET increased from 0.66 to 0.76). Here, the following C005 MODIS-Terra data are used: (i) AOD at  $\lambda = 550 \text{ nm}$  ( $\text{AOD}_{550 \text{ nm}}$ ), (ii) Ångström exponent over land ( $\alpha_{470-660 \text{ nm}}$ ), (iii) Ångström exponent over ocean ( $\alpha_{550-865 \text{ nm}}$ ), (iv) Fine mode Fraction (FF) of AOD over land and ocean, and (v) Effective radius over ocean ( $r_{\text{eff}}$ ). The relevant gridded aerosol data are stored in MODIS Level 3 (MOD08.D3) files, each corresponding to daily averages, and reported on a  $1^\circ \times 1^\circ$  latitude-longitude spatial resolution, based on statistics applied to the original 500 m resolution data. Their quality is satisfactory, since over the course of a day MODIS views the same  $1^\circ$ -square (grid) with a large variety of view angles, which improves the accuracy of calculated flux (Remer and Kaufman, 2006) and aerosol retrievals. The time-series of daily MODIS-Terra aerosol data cover the period March 2000–February 2007.

It should be noted that C005 MODIS data do not cover the highly reflecting desert areas of northern Africa (Sahara) due to restrictions in the associated retrieval algorithm. More recently, an improved version of the algorithm, enabling the retrieval of AOD above arid regions, led to the creation of the Deep Blue MODIS database (Collection 051) which covers Sahara. Nevertheless, the specific database has not been used in the present study mainly for two reasons: (i) because it is not yet enough validated (e.g. against AERONET) and (ii) in order to be consistent with our previous study, dealing with the identification of aerosol episodes (Gkikas et al., 2009), in which MODIS C005 data were used.

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## 2.2 Earth Probe-OMI Aerosol Index data

Absorption Aerosol Index (AI) data were taken from the Earth Probe (based on Total Ozone Mapping Spectrometer, TOMS) for the period 2000–2004 and OMI-Aura (based on Ozone Monitoring Instrument, OMI, measurements, since 2005) NASA's satellite databases (<ftp://toms.gsfc.nasa.gov/>), together covering the 7 yr period 2000–2007. They are both sequel to TOMS data, based on TOMS measurements onboard Nimbus-7 (1978–1993), Meteor3 (1991–1994), and Aeos (1996–1997) satellites. All together provide the longest available global aerosol record over land and oceans, starting with TOMS from 11 January 1978 and continued with the deployment of the EOS-Aura OMI in 15 July 2004 (Torres et al., 2007). This aerosol record is obtained at near UV spectral wavelengths, where the albedo of land surfaces (including arid and semi-arid areas) is very low, thus making possible retrievals of AOD, especially of absorbing aerosols.

AI is the primary TOMS aerosol product (Herman et al., 1997), provided on a daily basis, and it is a qualitative parameter associated with the presence of UV absorbing aerosols (e.g. desert dust). The near-UV aerosol retrieval method (full description provided by Torres et al., 1998, 2002) has been applied to observations by TOMS sensor onboard the Earth Probe (EP) platform (1996 to 2001). In version 2 of TOMS AI data, the TOMS algorithm, using observations at 331 and 360 nm, has been modified to make it consistent with the inversion procedure used by the Ozone Monitoring instrument (OMI) sensor (Version 8.5 Algorithm-TOMS), which was launched onboard the EOS-Aura satellite (1:38 p.m. equator crossing time, ascending mode) and its 2600 km viewing swath width provides almost daily global coverage. The basic algorithm for OMI uses 2 wavelengths (317.5 and 331.2 nm under most conditions, and 331.2 and 360 nm for high ozone and high solar zenith angle conditions). The OMI AI data are calculated from radiance residuals at 360 nm. AI is very nearly proportional to the aerosol absorption optical depth at 360 nm (Stammes and Noordhoek, 2002). However, the proportionality constant varies with the altitude (of the center of mass) of aerosol layer – the lower the altitude the smaller the constant. Most aerosols have stronger absorp-

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tion in the UV than in the visible, including mineral dust from deserts and carbonaceous aerosols containing organic and black carbon.

The mean daily AI data used in the present study are a combination of AI values from the Earth Probe and OMI-Aura satellites, mostly covering different time periods. Li et al. (2009) showed that mean annual cycles in the two data sets agree very well both globally and regionally, indicating a consistency between the AI products from Earth Probe and OMI. Since Earth Probe raw data are given at  $1^\circ \times 1.25^\circ$  spatial resolution, they have been re-gridded to  $1^\circ \times 1^\circ$  resolution in order to match that of the other satellite databases used here (OMI, MODIS). It should be noted that AI values from OMI-Aura have been adjusted down one half  $n$  value for consistency with the TOMS data record (<ftp://toms.gsfc.nasa.gov/pub/omi/data/aerosol/1README.txt>).

### 2.3 Surface PM data

PM<sub>10</sub> data from 21 regional background and sub-urban background sites were used in this study. These monitoring sites (see Fig. 1) are located from the west to east in the Mediterranean, as follows: 16 cover the entire Iberian Peninsula and the Balearic islands; 3 are in Italy, one being close to Rome, another in Sardinia and another one in Sicily; 1 is found in Crete; and 1 is in Cyprus. PM<sub>10</sub> concentrations were obtained in most cases from gravimetric determinations on filters, whereas in few cases they were determined by real-time instruments (Querol et al., 2009; Pey et al., 2013). All the data used in this study were obtained from public European databases: Airbase (<http://acm.eionet.europa.eu/databases/airbase/>), EMEP ([www.emep.int/](http://www.emep.int/)) and EUSAAR (<http://www.eusaar.net/>).

### 2.4 AERONET data

The aerosol columnar properties used in this work have been obtained from sun-photometric observations performed by the CIMEL sun-sky radiometer. The instruments are part of the Aerosol Robotic Network (AERONET, Holben et al., 1998) global

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network of stations (<http://aeronet.gsfc.nasa.gov>). More specifically, data are used from 9 AERONET stations found within the limits of the study region (Fig. 1).

The CIMEL data used in this study are Level 2.0 and provide information about the columnar AOD and aerosol size distribution. AOD data from the sun-photometers have been derived using the direct sun measuring mode, while the size distribution is calculated based on sky radiance measurements on specific angles along the almucantar and principle solar plane. The channel wavelength configuration is instrument version dependent, but for the above measurements filters at 440, 675, 870 and 1020 nm wavelengths were always present.

The AERONET technical specifications and the uncertainties of the CIMEL instrument are given in detail in Holben et al. (1998). More specifically, the total uncertainty of the AOD is influenced by various factors (instrumental, calibration-related, atmospheric and methodological). The AERONET instrument uncertainty for AOD is  $< \pm 0.01$  for wavelengths higher than 440 nm, and  $< \pm 0.02$  for UV wavelengths (Eck et al., 1999), or about 10% for a nominal aerosol optical depth of 0.1. The uncertainty of the sky radiance data and resulting size distributions are determined based on the calibration uncertainty that is assumed  $< \pm 5\%$  for all four wavelength channels (Holben et al., 1998).

## 2.5 Methodology

The procedure used to identify and characterize the desert dust (DD) episodes in the Mediterranean basin is depicted in the flowchart of Fig. 2. It is consisted of the following steps:

1. *First step*: mean and associated standard deviation (STDV) AOD<sub>550 nm</sub> values are first computed, for each  $1^\circ \times 1^\circ$  geographical cell, from time series of the pixel's AOD values over the entire study period.
2. *Second step*: threshold AOD levels are then defined for each geographical cell, separately for strong ( $\text{Mean} + 2\text{STDV} \leq \text{AOD}_{550 \text{ nm}} < \text{Mean} + 4\text{STDV}$ ) and extreme

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( $AOD_{550\text{nm}} \geq \text{Mean} + 4\text{STDV}$ ) DD episodes. This step, as well as the previous one, aim at the determination/identification of not only DD but in general aerosol episodes based on their loads (by means of AOD), and are described in detail by Gkikas et al. (2009) who applied them for identifying aerosol episodes over the Mediterranean basin.

3. *Third step*: it is the key step and consists in the discrimination and characterization of Mediterranean DD episodes. To this aim, the following other than AOD aerosol optical properties are used: (a) Ångström exponent ( $\alpha$ ), (b) Aerosol Index (AI), (c) Fine Fraction (FF) and (d) Effective Radius ( $r_{\text{eff}}$ ). The aerosol characterization algorithm is slightly modified for use over land and sea, since daily values of  $\alpha$ , FF and  $r_{\text{eff}}$  are derived by MODIS algorithms separately over land and sea.

A key issue is the selection of appropriate thresholds (cut-off levels) for each one of the algorithm aerosol optical properties. These thresholds have been defined in the present study based on self done sensitivity tests, taking into account literature results on aerosol characterization which are summarized, as to  $\alpha$ , FF, AI and  $r_{\text{eff}}$ , in the following few paragraphs. It should be noted that either one of those parameters or a combination of them has been used in literature.

### 2.5.1 Ångström exponent ( $\alpha$ )

Information on Ångström exponent is useful for separating aerosol particles between fine-mode and coarse-mode ones. The thresholds  $\alpha$  values are different among various studies and regions. Thus, for example, the value of 1 was used for classifying aerosols into fine ( $\alpha > 1$ ) and coarse ( $\alpha < 1$ ) ones (Eck et al., 1999; Holben et al., 2001). Takemura et al. (2002) found that  $\alpha$  values are less than 0.4 for Saharan dust (coarse) particles over subtropical Atlantic. Dubovik et al. (2002), using AERONET data, computed  $\alpha < 0.9$  in case of mineral dust particles.

More specifically, as far as the Mediterranean basin is concerned, Fotiadis et al. (2006), reported that desert aerosols over Crete island (eastern Mediterranean

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basin) are characterized by  $\alpha_{440-870\text{ nm}}$  values smaller than  $< 0.5$  whilst  $\alpha$  values smaller than  $0.6$  were reported by Bryant et al. (2006) for dust particles in Finokalia station (again in Crete island) during summer. Pace et al. (2006) reported that at Lampedusa Island (central Mediterranean) air masses arriving from Sahara desert resulted in  $\alpha$  values equal to  $0.42$ . Moreover, during Saharan dust outbreaks, Tafuro et al. (2006) computed a mean  $\alpha_{440-870\text{ nm}}$  value equal to  $0.2 \pm 0.1$  for five different AERONET sites in central Mediterranean, whereas Toledano et al. (2009) reported that dusty conditions are characterized by  $\alpha < 0.4$  during the SAMUM2006 campaign in Ouarzazate (Morocco).

### 2.5.2 Mean effective radius ( $r_{\text{eff}}$ ) and fine fraction (FF)

The mean effective radius and fine mode fraction are also used in the literature as another criterion for discriminating between fine and coarse aerosols. Thus, Tanré et al. (2001) used the critical value  $r_{\text{eff}} = 0.6$  to discriminate accumulation mode aerosols from the coarse mode ones and found that at three different locations affected by desert dust aerosols,  $r_{\text{eff}}$  values were equal to  $2.19 \pm 0.12$  (Banizoumbou),  $2.15 \pm 0.10$  (Sal Island) and  $3.01 \pm 0.24$  (Sede Boker). A similar discrimination of dust aerosols was also made by Jones et al. (2007) over ocean areas, based on a synergistic use of satellite observations and GOCART (Goddard Chemistry Transport Model) model, concluding that mean values of FF and  $r_{\text{eff}}$  for desert dust particles were equal to  $0.45 \pm 0.05$  and  $0.68$ , respectively.

### 2.5.3 Aerosol Index (AI)

A useful and widely used parameter in aerosol studies is the Aerosol Index (AI), which is a good and probably the only currently available indicator of aerosol absorptivity. Therefore, it is essential for distinguishing between the two most common types of coarse aerosols, namely dust and sea-salt, and determining dust source areas over the globe (Middleton et al., 2001; Prospero et al., 2002; Barkan et al., 2004; Washington

et al., 2003). Theoretically (Herman et al., 1997; Torres et al., 1998) negative values of AI indicate the presence of non-absorbing aerosols (e.g. sulfate or sea-salt particles) whereas positive AI values (mostly  $> 1$ ) indicate absorbing aerosols (dust or smoke).

#### 2.5.4 Combination of $\alpha$ , $r_{\text{eff}}$ , FF, and AI

5 Apart from using thresholds for each one aerosol parameter, independently, to distinguish between different aerosol types, a combination of more than one can lead to more satisfactory results. To this aim, several scientists have used combined data of aerosol optical properties such as AOD,  $\alpha$  and FF. Barnaba and Gobbi (2004) used one-year Level 2 MODIS-Terra data to study the intra-annual variability of specific  
10 aerosol types (maritime, continental, dust) over the Mediterranean basin. To this end, they applied a specific “aerosol mask” to identify desert dust aerosols ( $\text{FF} < 0.7$  and  $\text{AOD}_{550\text{nm}} > 0.3$ ). Kalivitis et al. (2007) identified the presence of desert dust particles over Crete, by combining  $\text{PM}_{10}$ , satellite (TOMS) and AERONET data and reported that they occur whenever  $\text{AOD}_{870\text{nm}} > 0.2$  and  $\alpha_{440-870\text{nm}} < 0.6$ . Toledano et al. (2007)  
15 also detected dust based on routine 6 yr (2000–2005) aerosol measurements at El Arenosillo station (Huelva, Spain), by setting thresholds for  $\text{AOD}_{440\text{nm}} (> 0.25)$  and  $\alpha_{440-870\text{nm}} (< 0.8)$ .

#### 2.5.5 Present algorithm

The available information from literature has been taken into account in our algorithm, and appropriate thresholds have been defined for each one of the four aerosol parameters ( $\alpha$ , AI, FF, and  $r_{\text{eff}}$ ) used to identify Mediterranean DD aerosol episodes. These thresholds are outlined in Fig. 2.

20 According to our algorithm, desert dust episodes are identified when  $\alpha \leq 0.7$ ,  $\text{AI} > 1$ ,  $\text{FF} \leq 0.4$  and  $r_{\text{eff}} > 0.6 \mu\text{m}$ . It should be noted that  $r_{\text{eff}}$  is used in the algorithm only above sea and not land areas because MODIS  $r_{\text{eff}}$  values are available only there. Given  
25 the availability of slightly different values in literature, questions may arise concerning

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the selection of the specific thresholds. To address the sensitivity of the algorithm to individual size parameters ( $\alpha$ , FF, and  $r_{\text{eff}}$ ) several sensitivity tests have been performed changing the associated thresholds. The results show that modifications up to 5 % are found except for FF for which, only for strong DD episodes over sea, changes can be larger.

### 3 Evaluation of the algorithm

The developed satellite based algorithm for the determination of DD episodes in the Mediterranean basin has been evaluated through comparison of its outputs, i.e. the identified strong and extreme DD episodes, with surface data for aerosol physical and optical properties, which are commonly considered as more reliable. *In a first step*, algorithm's outputs are compared against surface measurements of PM<sub>10</sub> for 21 stations located across the Mediterranean basin (Fig. 1, yellow coloured). For each station we have selected/used the satellite measurements for the pixels where stations are found, for cases (days) identified as strong or extreme DD episodes with the algorithm, yielding a number of 333 DD episodes in total. *In a second step*, the similar analysis was performed, but comparing MODIS-Terra aerosol properties against data from 9 AERONET stations across the Mediterranean basin (Fig. 1, red coloured) for 58 DD episodes.

#### 3.1 Particulate matter (PM) measurements

Figure 3i displays the computed correlation coefficients ( $R$ ) between surface PM<sub>10</sub> concentrations ( $\mu\text{g cm}^{-3}$ ) and satellite AODs (unitless) at 550 nm, under desert dust episodes conditions, for 21 Mediterranean stations. The size of circles, increasing with  $R$ , indicates how well these two measurements are correlated, while colors are indicative of data availability (red/blue for more/less than 10 episodes). The results show that in most locations  $R$  values are lower than 0.4 (poorest correlation, not statisti-

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cally significant, in Monagrega, NE Spain) while in 6 stations  $R$  values are larger than 0.5. The best correlation, statistically significant at 95 % confidence level, is found in Sicily ( $R = 0.90$ ), Crete ( $R = 0.75$ ) and Cyprus ( $R = 0.89$ ) island stations. In Fig. 3ii are depicted the scatterplot comparisons between ground PM and satellite AOD measurements for the selected 333 DD episodes, also obtained on a seasonal, apart from the annual, basis separately. The overall comparison is relatively satisfactory taking into account the different nature of compared data, i.e. surface PM measurements against columnar satellite AOD products. Thus, the computed overall  $R$  value for the 333 DD Mediterranean episodes is equal to 0.64 (statistically significant at 95 % confidence level), while it seems that PM values are biased low with respect to AOD ones. Nevertheless, more information is obtained on seasonal basis. More specifically, the correlation in winter and spring is good, with  $R$  values equal to 0.69 in winter (14 DD episodes) and 0.74 in spring (95 DD episodes), while the correlation is even better in autumn ( $R = 0.81$ , 35 DD episodes). It must be mentioned that these three correlation coefficients are statistically significant at 95 % confidence level. Therefore, it appears that if we except summer, correlation between surface PM and satellite AOD products is very good, since  $R > 0.7$ . However, the overall  $R$  value drops below 0.7 (0.64), because the major percentage of examined DD episodes (189 out of 333) occurs in summer, when the computed correlation coefficient is very poor ( $R = 0.09$ ) and not statistically significant. The smallest summer  $R$  values in Fig. 3ii are essentially in line with the low  $R$  values over the Iberian Peninsula stations in Fig. 3i, since dust transport in the western parts of the Mediterranean basin is mainly observed in summer (e.g. Toledano et al., 2007; Papadimas et al., 2008; Querol et al., 2009; Pey et al., 2013).

Despite the relatively good performance of our algorithm, in terms of comparison with surface PM products, differences are found, especially in summer. The amplified differences in summertime can be explained by the differences in vertical extension between the two types of products, given that dust transport in this season mainly occurs in the free troposphere. Such dust events can be detected by satellite observations but not by surface based ones, based on their nature. The vertical extension of dust loads



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in the Mediterranean basin has been the subject of several studies (e.g. Mona et al., 2006; Di Iorio et al., 2009; Papayannis et al., 2009; Sicard et al., 2011) at different locations in the Mediterranean basin, mainly based on lidar observations. All these lidar based studies have documented that transported desert dust particles are more wide vertically extended in summer months, which is meaningful due to stronger convection. Under such conditions, particulate matter can be removed away from the surface, thus restricting the ability of ground PM measurements to record DD episodes and leading to poor correlation between PM and satellite measurements with the former being biased low. In addition, Kalivitis et al. (2007) using trajectory analyses found that free tropospheric and boundary layer transport of dust (dust observed only at 3000 m and 1000 m, respectively) in the E. Mediterranean become maximum in summer, which also prevents surface PM measurements from recording DD episodes. Finally, CALIOP data for the period 2007–2011 (V. Amiridis, personal communication, 2013) shows that dust scale height, i.e. the height above ground where the 63 % of the columnar dust load is contained, has values up to 3 km in summer, being clearly higher (by 1–2 km) than in other seasons. On the other hand, however, it should be also noted that satellite AOD values can be overestimated due to the presence of clouds (cloud contamination) especially under total cloud cover larger than 80 % (Zhang et al., 2005; Remer et al., 2008). In order to investigate this we have re-computed the correlation coefficient between ground (PM) and satellite (MODIS) AOD values excluding points corresponding to cloud fraction larger than 80 % (results not shown here). As a result, the  $R$  value drastically increased from 0.64 to 0.82 (statistically significant). Finally, a third factor which can affect the quality of our results is that there are stations where the common pairs of ground-satellite measurements is small and the computed correlation coefficients for those stations (Fig. 3i) can be no representative.

### 3.2 AERONET measurements

By their nature AERONET data have more common characteristics with MODIS satellite ones than PM measurements, since they are both columnar AOD products based

on remote sensing from ground and space, respectively. A total number of 58 identified DD episodes have been examined in this case. Given that for the examined AERONET stations the available AOD values are mostly reported at 440 nm in order to match the MODIS wavelength (550 nm), we derived AERONET AODs at 550 nm from original values at 440 nm and Ångström exponent between the wavelengths 440 nm and 870 nm. The overall scatterplot comparison is shown in Fig. 4i revealing a relatively good agreement ( $R = 0.65$ , statistically significant) between our algorithm and AERONET in case of dust events. Again, MODIS AODs seem to be overestimated (bias = 0.22), especially for low AODs, while the situation is improved for stronger dust episodes.

The validity of the algorithm products was checked in another way by examining the aerosol volume size distributions for the 58 DD episodes. To this aim, and also in order to highlight the difference made under episode conditions, data of volume size distribution were taken from AERONET and subsequently have been averaged for all observations (blue colour) and for only the 58 DD episodes (red colour), both for the 9 Mediterranean stations. The results (Fig. 4ii) make evident that under DD episodes in the Mediterranean, the aerosol coarse mode is strongly increased, by a factor of  $\sim 10$ , with a volume distribution peak at  $2.24 \mu\text{m}$ , due to the predominance of coarse dust particles. Thereby, the overall bi-modal aerosol volume size distribution (blue curve) becomes mono modal (red curve) under Mediterranean desert dust episodes.

## 4 Desert dust episodes regime

### 4.1 Geographical patterns

#### 4.1.1 Frequency of occurrence

The averaged (2000–2007) geographical distributions of the frequency of occurrence (episodes/year) of strong and extreme DD episodes in the Mediterranean basin are presented in Fig. 5i and ii, respectively. It is evident a dominant south-to-north decreasing

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gradient, along with a smaller west-to-east gradient. Thus, strong DD episodes occur more frequently in the western parts of the Mediterranean (11.4 episodes/year over the northwestern coasts of Africa), dropping to less than 3.5 episodes/year in the eastern part of the basin. On the other hand, extreme DD episodes are mostly observed in the central Mediterranean (up to 3.9 episodes/year in Gulf of Sidra), whereas their maximum frequencies are systematically observed over sea surfaces opposite to strong episodes mostly occurring over land. The south-to-north gradient has been also observed from a ground-based monitoring perspective (Querol et al., 2009; Pey et al., 2013). On the contrary, Pey et al. (2013) reported a slight east-to-west decreasing gradient and even higher frequencies of dust events compared to ours, which can be both attributed to the use of surface PM measurements in that study and also to the different definitions of dust events with the present study.

The predominant south-to-north increasing gradient of DD episodes frequency is reasonable since it is known (e.g. Mona et al., 2006; Israelevich et al., 2012) that Mediterranean regions in the vicinity primarily of Sahara and secondarily of Middle East deserts are frequently affected by the transport of desert particles. Our frequencies when compared to those of literature for specific locations throughout the study region maybe found somewhat smaller but this can be explained by the different nature of data used and to the different applied methodologies. According to our results continental central European areas exhibit low frequencies ( $< 1$  episodes/year) indicating that desert dust can impact remote areas, which can be even more distant based on literature (Ansmann et al., 2003; Borbely-Kiss et al., 2004; Klein et al., 2010).

The seasonal maps of DD episodes frequency indicate that there is a longitudinal shift of the activity of both strong and extreme DD events (Fig. 6i and ii, respectively). Thus, during autumn, strong DD episodes occur more frequently over the central Mediterranean Sea (but with relatively low frequencies of 1.9 episodes/year) while in spring and winter they occur more often in the eastern Mediterranean with higher frequencies (up to 3.9 episodes/year). On the other hand, the highest frequencies of strong DD episodes are observed in summer (8.1 episodes/year) and take place in the



the different nature of these two parameters, and also by their different determinant factors and different degrees of their action. For example, the patterns of frequency of DD episodes can be generally regulated by the distribution patterns of pressure systems, while their intensity is mostly related to individual strong pressure systems.

### 4.1.3 Duration

The computed mean duration (in days) of strong and extreme DD episodes in the Mediterranean basin is presented in Figs 8i and 8ii, respectively. The strong DD episodes last slightly longer in the western (up to 1.4 days) than other parts of the basin, with maximum duration over Morocco. The extreme episodes are characterized by a reverse gradient, with longer episodes in the eastern Mediterranean basin (up to 1.5 days in E. Mediterranean) and maximum duration in the southern Levantine Sea. Long extreme DD episodes are also found in the eastern parts of the Atlantic Ocean.

Averaged duration results over 2000–2007 presented in Fig. 8 mask specific quite longer individual DD episodes (more for strong than extreme episodes). Thus, during our study period three strong dust episodes (in Morocco, July 2003) lasted up to 6 days, while a single extreme DD episode lasted four days and has taken place between 2 and 5 April 2000 in the sea region between Cyprus and Turkey (Kubilay et al., 2003). More than 85 % of all identified DD episodes in the broader Mediterranean basin last about one day, yielding a regional mean duration equal to 1.12 days for strong DD episodes over continental, and 1.08 days for extreme DD episodes over maritime areas. The computed duration of DD episodes in this study may seem somewhat short, but duration here is given in a day timescale, which implies that half a day corresponds to 12 h (and a decimal point to about 2.5 h, which is meaningful).

## 4.2 Temporal variability

Apart from the geographical, the temporal variation of Mediterranean DD aerosol episodes is possible to be studied here, for the entire region and specific sub-regions,

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thanks to the allowed satisfactory spatial coverage by the satellite data used in the present work. Thus, the intra-annual (seasonal) and inter-annual variations of strong and extreme DD episodes has been studied for the broader Mediterranean basin, but also separately for its western, central and eastern parts, and for land and sea surfaces.

#### 5 4.2.1 Intra-annual variability

The seasonal frequencies of occurrence (relative percent terms) of strong and extreme DD episodes are given in Table 1. Apart from the whole Mediterranean basin, results are separately computed over continental (land) and maritime (sea) areas of the basin as well as for its western, central and eastern sections.

#### 10 Continental areas

Over continental areas of the Mediterranean, strong DD episodes occur more frequently in summer (51.4%). This is actually observed in the western (67.3%) and central (51.5%) sections of the basin, but not in the eastern one, where maximum frequency occurs in spring (71%). Strong Mediterranean DD episodes take place secondarily in spring, with percentages of 25.5–40.4%, except for the eastern basin (summer secondary maximum). The seasonal regime of DD episodes is different for extreme than strong ones. Thus, extreme episodes occur more frequently in spring, both in central (60.4%) and eastern (87.7%) parts of the basin, as well as in the entire basin (61.1%), while in the western basin they are more frequently observed in summer (54.2%).

#### 20 Maritime areas

Over maritime areas of the Mediterranean basin the seasonality of DD episodes shows a slightly different behaviour than over continental regions. Thus, for the entire basin, the maximum frequencies of occurrence are recorded in spring, both for strong (43.4%) and extreme (57.2%) episodes, opposite to more frequent continental episodes in sum-

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mer. This specific seasonality, i.e. spring maxima both for strong and extreme episodes, is observed for the central (46.8 % strong DD, 63.8 % extreme DD) and eastern (61.9 % strong DD, 67.4 % extreme DD) parts of the Mediterranean Sea, while only in its western parts DD episodes occur more frequently in summer (49.5 % and 42.2 % for strong and extreme, respectively). The lowest frequencies are found in winter, except for extreme DD episodes in the eastern Mediterranean, due to the efficient wet removal in this season, associated with enhanced precipitation (Marriotti et al., 2002) which is the major atmospheric aerosol removal process (Pruppacher and Klett, 1997)

### Other studies

Our findings cannot be easily compared to other observational ones since they are the first with a complete spatial coverage for the studied period, namely subsequent to 2000, opposite to existing studies performed either for specific locations (Toledano et al., 2007; Meloni et al., 2007) or previous periods (Moulin et al., 1998). The seasonal spatial characteristics of dust transport in the broader Mediterranean basin has been studied by a few more researchers using satellite data but only over short periods (e.g. Barnaba and Gobbi, 2004) or indirectly, i.e. through total AOD and size parameters distributions (e.g. Papadimas et al., 2008). Moreover, this was also attempted based on model simulations (e.g. Basart et al., 2012). In summary, the findings of the aforementioned studies are in agreement with our results as to the aforementioned seasonal variability of dust in the Mediterranean.

### 4.2.2 Inter-annual variability

Previous studies (e.g. Papadimas et al., 2008; Hatzianastassiou et al., 2009; Yoon et al., 2011) have shown that there is a significant year by year variability of aerosol loadings in the Mediterranean basin. Moreover, Gkikas et al. (2009) have shown that such a significant inter-annual variability, in terms of frequency and intensity, also exists for maximum aerosol loadings, i.e. aerosol episodes.

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Here it is attempted to examine the inter-annual variability of dust episodes (outflows) in the Mediterranean basin, more specifically. The aim is to quantify this variability in terms of frequency of occurrence and intensity of DD episodes, which are both very important in many aspects. Computations are performed separately over Mediterranean land and sea areas and for strong and extreme DD episodes. The frequency is normalized to be expressed in number of events per unit area ( $1^\circ \times 1^\circ$  latitude-longitude or about  $100\text{ km} \times 100\text{ km}$ ) while the intensity is expressed by means of  $\text{AOD}_{550\text{ nm}}$ .

### Frequency of occurrence of DD episodes

The Mediterranean DD episodes depict a seasonal cycle with maximum frequencies, as already noted in Sect. 4.2.1, during the dry period of the year (spring–summer) and minimum appearances in the wet period (autumn–winter). The spring maxima are usually observed in April, whereas summer ones in July. The relative strength of maximum spring/summer frequencies is different when examined for different sub-regions, e.g. western-central-eastern basins, or for strong and extreme DD episodes, separately. Thus, sub-regional results (not shown here) indicate that summer maxima are primary in the western Mediterranean basin, whereas in the central and especially in eastern Mediterranean the spring maxima become primary, in agreement with the findings of Table 1. Also, the two maxima are about equivalent for strong DD episodes, while the spring maxima clearly become primary in case of extreme episodes. Differences in frequencies are also encountered between strong and extreme episodes, as well as between land and sea Mediterranean areas. More specifically, frequencies of strong episodes reach values up to about 1.2 episodes/pixel, whereas those of extreme episodes hardly exceed 1.0 episode/pixel being higher over sea than land areas. There is a significant year by year variability of frequencies of DD episodes, especially over land, with an absolute maximum of 1.2 episodes/pixel for strong episodes in July of 2003 (similar high frequencies were also observed in July 2002 and April 2000). A more distinct absolute maximum frequency of extreme DD episodes (1.2 episodes/pixel, again) is found over land in April 2000, but high frequencies are also



observed in April 2003. Therefore, according to our results, years 2000 (April) and 2003 (July) seem to have been marked by most frequent DD episodes over Mediterranean land areas. More specifically, according to the reproduced AOD maps with the Giovanni tool (<http://disc.sci.gsfc.nasa.gov/giovanni/overview/index.html>) for each day of April 2000 and July 2003, a significant portion of the eastern Mediterranean basin in April 2000, and of western Mediterranean in July 2003, are characterized by AOD values higher than 0.5.

The applied linear regression fit to time series of Fig. 9 (blue lines) indicate that strong and extreme DD episodes occurred in the Mediterranean basin over 2000–2007 with decreasing frequencies. This is valid over both land and sea Mediterranean areas although it is slightly more clear over land. The tendencies of frequencies of strong and extreme DD episodes have been also checked using anomalies instead of absolute frequency values, and the results, i.e. decreasing frequencies (not shown here), remain the same. This is also valid when using annual frequencies (again not shown here). Nevertheless, it should be noted that the decreasing tendencies over sea are uncertain as indicated by the computed error of associated slope values. According to Fig. 9, it appears that the decreasing tendency of occurrence of DD episodes in the Mediterranean is mainly caused by low spring and summer frequencies in years 2005–2007, especially over land, combined with high frequencies in years 2000 and 2003.

## Intensity of DD episodes

As to the inter-annual variation of the intensity of DD aerosol episodes (red lines in Fig. 9), our results do not indicate a clear seasonal cycle, both for the strong and extreme episodes. In general, the intensity of Mediterranean DD aerosol episodes is higher over maritime than continental areas. The maximum regional mean intensity of strong DD episodes is equal to 0.96 over land (in December 2000) and 1.23 over sea (in November 2004). The corresponding values for extreme DD episodes are equal to 1.5 (May 2001) and 4.86 (January 2004). The linear regression analysis, reveal decreasing tendencies (up to  $-7.5\%$ ), over continental parts, while over sea-areas the

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intensity of DD episodes is found to have decreased ( $-5.4\%$ ) for strong episodes and increased ( $13.5\%$ ) for extreme episodes. Similar results are found using anomalies of intensity instead of absolute values. It should be noted that the tendencies over Mediterranean Sea areas are non significant (the error of the slope is greater than the slope value). The changing intensity of Mediterranean DD aerosol episodes is possibly related with modifications of the position and strength of pressure systems, and the instability of the atmosphere, since there is strong connection between atmospheric circulation and associated desert dust concentrations. Of course, this deserves further and more thorough investigation which is beyond the scope or the present work.

### 4.3 Air masses and pathways

Identification of the origin and tracks of air masses associated with DD aerosol episodes in the Mediterranean basin is important, and contributes to the understanding of the properties of aerosol events and possibly to their forecasting. This has been attempted here by reproducing the back trajectories for air masses arriving at locations (geographical cells) undergoing the maximum frequencies of our determined extreme DD episodes over land and sea (shown in Fig. 5ii). Such back trajectories were finally derived for 3 cells at Morocco, Tunisia and Gulf of Sidra, corresponding to 57 DD episodes in total. However, the relevant results are representative of all events yielding a much higher number (about 15 000). The back trajectory analysis has been made for the specific cells only, because: (i) it is practically impossible to do the same for all aerosol episodes that were identified at pixel level, and (ii) in general, aerosol episodes are events occurring over spatially extended areas rather than specific locations, and therefore the obtained results here are more or less representative of a much higher number of pixel-level episodes. According to our procedure, for the selected cases, five-day back trajectories ending at the specific locations (cross points in Fig. 10) were calculated with the HYSPLIT (Hybrid Single Particle Lagrangian Integrated Trajectory Model) Dispersion Model (Draxler and Rolph, 2013). The computed trajectories end at three different levels, which are representative of the planetary boundary layer (PBL)

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containing most of aerosol particulate matter: (i) 500 m above ground level (a.g.l.) (trajectories in red colour), (ii) 1500 m a.g.l. (blue colour), and (iii) 3000 m a.g.l. (green colour).

The trajectories of extreme DD episodes ending at 500 m (in red, Fig. 10i) seem to be quite complicated, and originate and remain mostly over Europe and Africa, with few trajectories originating from the Atlantic Ocean. A more thorough look reveals that they circulate for at least 2 days (the last 48 h) over the deserts of North Africa. The situation is different for the trajectories ending at 1500 and 3000 m (in blue and green colours, Fig. 10ii and iii, respectively), with their entire pathways residing above North Africa. In particular, most 3000 m trajectories originate and travel over the arid regions of west and central Sahara. In cases of air masses ending at 1500 and 3000 m, some DD trajectories originate from remote locations in North Atlantic Ocean or even from the eastern coasts of North America. Even in these cases, however, they remain above N. Africa during the last 1–2 days of their path, thus being enriched by dust particles from the world's major desert of Sahara (e.g. Middleton and Goudie, 2001; Prospero et al., 2002; Washington et al., 2003; Barkan et al., 2004). This enrichment is realized through dust uplift and transport at high altitudes induced by cyclonic systems (e.g. Kalivitis et al., 2007; Papayannis et al., 2009).

An important factor determining the nature of aerosol masses and associated episodes is the altitude of back trajectories during the 5-day travel period. This information was derived by computing the mean (from the 57 studied extreme DD episodes) altitude of air masses ending at 500, 1500 and 3000 m, and the associated standard deviations (Fig. 10iv). According to our results, 1500 and 3000 m air masses inducing extreme DD episodes in the Mediterranean follow a gradually sinking motion from the beginning of their path till about 24–48 h before the end, reaching altitudes of 1000 m and 2000 m, respectively. At these altitudes, 1–2 days before the episode, air masses are enriched by uplifted desert dust particles under prevailing cyclonic conditions in N. Africa. At the same time, the uplift results in a subsequent increase of the altitude of

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air masses during the last 1–2 days of their path, to end at 1500 m and 3000 m (see Fig. 10iv).

## 5 Summary and conclusions

The present study aimed to describe, for the first time to our knowledge, the regime of desert dust episodes that take place over the entire Mediterranean basin. To this end an objective and dynamic algorithm has been set up which identified the DD episodes that occurred over the period 2000–2007. The algorithm uses as input data daily aerosol optical properties, derived from MODIS, Earth Probe and OMI satellite measurements, for the following parameters: Aerosol Index (AI), Fine Fraction (FF), Ångström exponent ( $\alpha$ ) and Effective radius ( $r_{\text{eff}}$ ). The identified DD episodes were classified into strong and extreme ones, based on their intensity by means of aerosol optical depth (AOD). The algorithm determined the main characteristics of DD episodes, namely their frequency, intensity and duration, at various spatial (from pixel-level to regional mean) and temporal (from daily to 7 yr means) scales.

First, the performance of the algorithm has been tested against quality surface measurements. This was done for 333 pixel level DD episodes collocated with selected stations found within the study region. An initial evaluation was performed using daily  $\text{PM}_{10}$  concentrations data from 21 stations across the Mediterranean basin. The results revealed a very good agreement between ground and satellite measurements for central and eastern Mediterranean stations (correlation coefficient,  $R$ , values up to 0.91) against lower  $R$  values over the western basin. More information was obtained performing the validation on seasonal basis. More specifically, there was a moderate-to-good agreement in winter ( $R = 0.69$ ), spring ( $R = 0.74$ ) and autumn ( $R = 0.81$ ), contrary to poor correlation in summer ( $R = 0.09$ ), which can be attributed to the higher vertical extension of dust loadings during this season. This vertically extended dust transport, far above the boundary layer, does not allow them to be captured by ground stations measuring PM, since their in-situ measurements are restricted to the surface air layer.

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In a second approach, the algorithm's performance has been also evaluated against aerosol optical properties from 9 AERONET stations across the Mediterranean basin. This has been done for 58 DD episodes, and the comparison revealed: (i) a relatively good correlation ( $R = 0.65$ ) between MODIS and AERONET AOD values and (ii) a significant increase (by a factor of  $\sim 10$ ) of coarse mode particles (dominated by dust, centred at  $2.24 \mu\text{m}$ ) in case of DD episodes relative to all cases.

According to our computed *geographical distributions*:

- There is a longitudinal gradient in the frequency of occurrence of Mediterranean DD episodes. This gradient is more evident for strong episodes, which occur more frequently (up to 11.4 episodes/year) in the western Mediterranean basin, and less evident for the (more sparse) extreme episodes which occur more frequently in the central basin (up to 3.9 episodes/year).
- There is also a clear and predominant latitudinal gradient, with frequencies of both strong and extreme DD episodes decreasing from south to northern, as the distance from the northern African desert areas increases while a west-east gradient is evident only for the strong ones, determined probably by the climatology of the pressure systems.
- The geographical variability of DD episodes' frequency in the Mediterranean, and specifically the longitudinal one, appears to be driven by the prevailing pressure systems in the area and precipitation patterns. This is confirmed by the spatial variability of frequency of dust episodes throughout the course of year. Thus, in the eastern Mediterranean basin, the DD episodes occur more frequently in spring (up to 3.9 episodes/year) and winter (up to 1 episodes/year) having in both seasons the maximum frequencies over the entire basin. Respectively, in central parts of the Mediterranean basin maximum DD episode frequencies are observed in spring (up to 3 episodes/year) and in the western parts of the basin in summer (up to 8.1 episodes/year).

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– The intensity of strong DD episodes (in terms of  $AOD_{550\text{nm}}$  values) in the Mediterranean can reach values up to 1.5 (Gulf of Sidra, Libyan Sea). The intensity of extreme DD episodes is however significantly higher, up to 4.1 (Libyan Sea). Although there is an apparent south-to-north gradient in the intensity of DD episodes, similarly to their frequency, there is no similarity between intensity and frequency in terms of their spatio-temporal patterns.

– As for their persistence, it is found that the strong DD episodes last up to 1.4 days (34 h) in the western Mediterranean basin, while the extreme ones can be more persistent, lasting up to 1.5 days (36 h), in the eastern parts.

The analysis of *intra-annual (seasonal) and inter-annual (year by year) variation* of Mediterranean DD episodes reveals that:

– For the entire Mediterranean basin, on average, there is a different seasonality over land and sea areas. Thus, strong Mediterranean DD episodes over land occur more frequently in summer (51.4 %) whilst over sea they are more frequent in spring (61.1 %). Respectively, extreme desert dust episodes are by far more frequent in spring over both continental (43.4 %) and maritime (57.2 %) areas of the Mediterranean basin.

– When examining the seasonality of DD episodes at smaller spatial scale, i.e. for different sub-regions, namely western, central and eastern parts of the basin, differences are found. Thus, our results indicate that in the western Mediterranean basin, DD episodes (both strong and extreme) occur more frequently in summer. On the contrary, in the central and eastern parts of the Mediterranean basin, DD episodes are more common in spring, except for strong episodes over land areas of central basin which occur more frequently in summer.

– In general, it appears a quite stable year by year seasonality of DD episodes, in line with that described above. On the contrary, there are distinct fluctuations in the

magnitude of DD episodes from one year to another, which are partly explained by inter-annual changes of NAO Index and precipitation amounts.

- Our results indicate that DD episodes in the Mediterranean have been decreased from 2000 to 2007, more over land than sea.

Finally, the obtained *5-day back-trajectories of air masses* leading to extreme DD episodes in the Mediterranean basin indicate that:

- The air masses ending at higher altitudes, especially at 3000 m a.g.l., mostly originate from the western and central parts of the Saharan desert, while those arriving in the lower boundary layer can also start their path from Europe. There are also some cases in which air masses originally come from distant areas of the northern Atlantic Ocean.
- The back trajectories show that during their travel the air masses gradually sink till 24–48 h before they complete their path, being enriched by mineral particles over the northern African deserts, lifted upwards by prevailing cyclonic conditions. Subsequently, air masses increase their height until they complete their path.

It is planned to supplement and improve the results of this study in the near future in various aspects. Thus: (i) a larger spatial coverage of the study region will be possible using MODIS Deep Blue aerosol data, which are also available over the highly reflecting deserts of Sahara and Middle East, after these data will be adequately validated, (ii) a larger temporal coverage is also desirable, which will strengthen our findings as to seasonal and inter-annual variability of Mediterranean DD episodes, (iii) the accumulation of enough data derived from satellite based lidar systems over time, will offer a robust evaluation of the outputs of our algorithm through comparisons against vertically resolved CALIOP aerosol information, which directly specifies the existence of various aerosol types with height, (iv) finally, the study will be repeated in other dust dominated key world regions, for example Gobi and Taklimakan deserts.

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*Acknowledgements.* This research has been co-financed by the European Union (European Social Fund - ESF) and Greek national funds through the Operational Program “Education and Lifelong Learning” of the National Strategic Reference Framework (NSRF)-Research Funding Program: **ARISTEIA**. Investing in knowledge society through the European Social Fund.

5 The Earth Probe (TOMS) and OMI aerosol climatologies are available from the NASA's Web sites ([http://acdisc.gsfc.nasa.gov/opendap/EarthProbe\\_TOMS\\_Level3/TOMSEPL3.008/contents.html](http://acdisc.gsfc.nasa.gov/opendap/EarthProbe_TOMS_Level3/TOMSEPL3.008/contents.html)) and (<http://ozoneaq.gsfc.nasa.gov/OMIAerosol.md>) respectively. The Collection 005 MODIS-Terra data were obtained from NASA's Level 1 and Atmosphere Archive and Distribution System (LAADS) website (<ftp://ladsweb.nascom.nasa.gov/>). The authors gratefully acknowledge the NOAA Air Resources Laboratory (ARL) for the provision of the HYSPLIT transport and dispersion model and/or READY website (<http://ready.arl.noaa.gov>) used in this publication.

10 We would like to thank the principal investigators maintaining the 9 AERONET sites used in the present work. We would like to acknowledge the EMEP program and the public European databases, Airbase and ACTRIS, which supplied PM<sub>10</sub> data used in this study.

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**Table 1.** Seasonal frequency of occurrence (in percent values) of strong and extreme DD aerosol episodes over land and sea surfaces of the entire Mediterranean basin, and over western, central and eastern Mediterranean, during the period 2000–2007. Maximum frequencies are indicated in bold and italic, the second largest frequencies in bold and the third ones in italic

LAND	Mediterranean		W. Mediterranean		C. Mediterranean		E. Mediterranean	
	strong	extreme	strong	extreme	strong	extreme	strong	extreme
Winter	1.6%	1.5%	1%	1.2%	0.4%	1.5%	3.3%	2%
Spring	<b>40.4%</b>	<b>61.1%</b>	<b>25.5%</b>	<b>37.4%</b>	<b>38.4%</b>	<b>60.4%</b>	<b>71%</b>	<b>87.7%</b>
Summer	<b>51.4%</b>	<b>30.9%</b>	<b>67.3%</b>	<b>54.2%</b>	<b>51.5%</b>	<b>31.4%</b>	<b>19%</b>	4.6%
Autumn	6.6%	6.5%	6.2%	7.2%	9.7%	6.7%	6.7%	<b>5.7%</b>
SEA	Mediterranean		W. Mediterranean		C. Mediterranean		E. Mediterranean	
	strong	extreme	strong	extreme	strong	extreme	strong	extreme
Winter	9.1%	13.3%	4.1%	6.3%	6.2%	8.4%	<b>20.6%</b>	<b>26.3%</b>
Spring	<b>43.4%</b>	<b>57.2%</b>	<b>27%</b>	<b>37.3%</b>	<b>46.8%</b>	<b>63.8%</b>	<b>61.9%</b>	<b>67.4%</b>
Summer	<b>29%</b>	<b>18.4%</b>	<b>49.5%</b>	<b>42.2%</b>	<b>24.7%</b>	<b>14.5%</b>	6%	0.9%
Autumn	18.5%	11.1%	19.4%	14.2%	22.3%	13.3%	11.5%	5.4%

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**Fig. 1.** The study region and stations whose data are used for comparison with the outputs of the present algorithm. Shown are stations providing: **(i)**  $\text{PM}_{10}$  measurements (yellow colour) and **(ii)** aerosol optical properties (AERONET, red colour).

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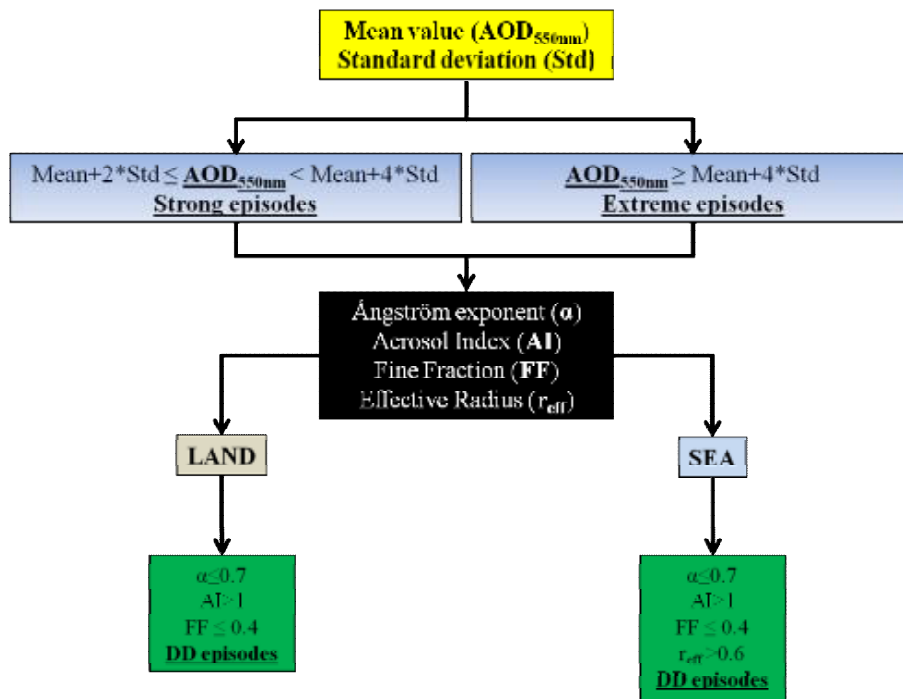
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**Fig. 2.** Flowchart of the developed methodology and algorithm for the identification and characterization of desert dust (DD) aerosol episodes in the Mediterranean basin.

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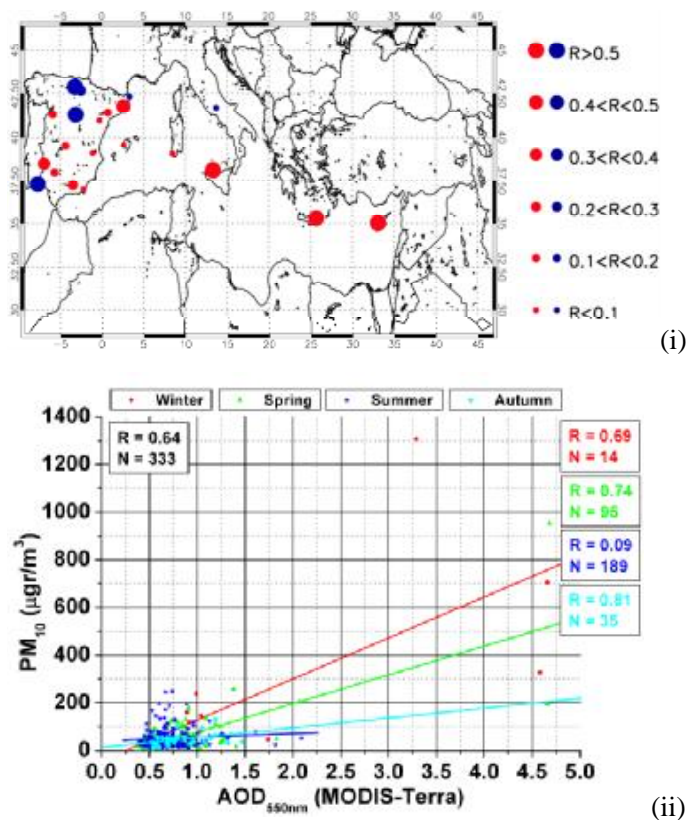
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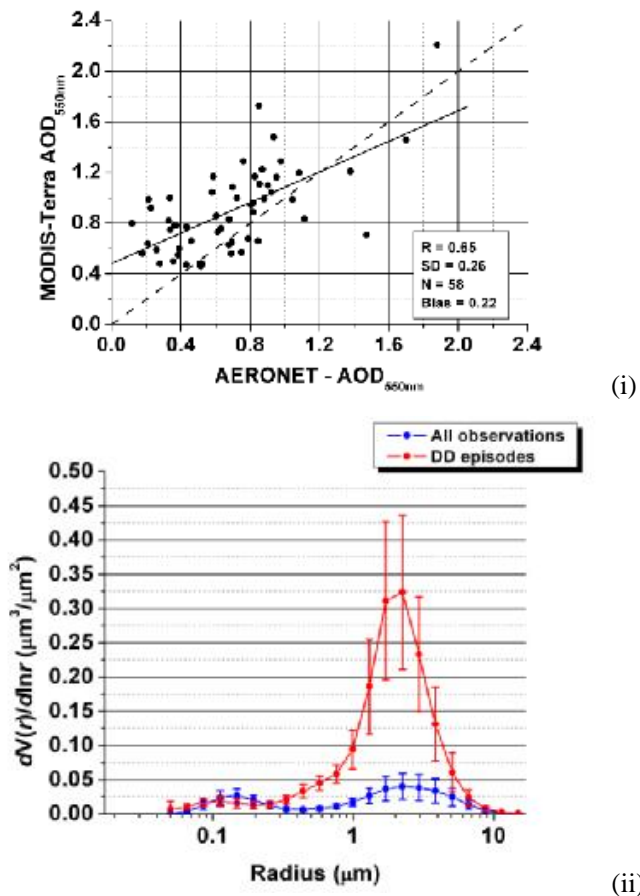
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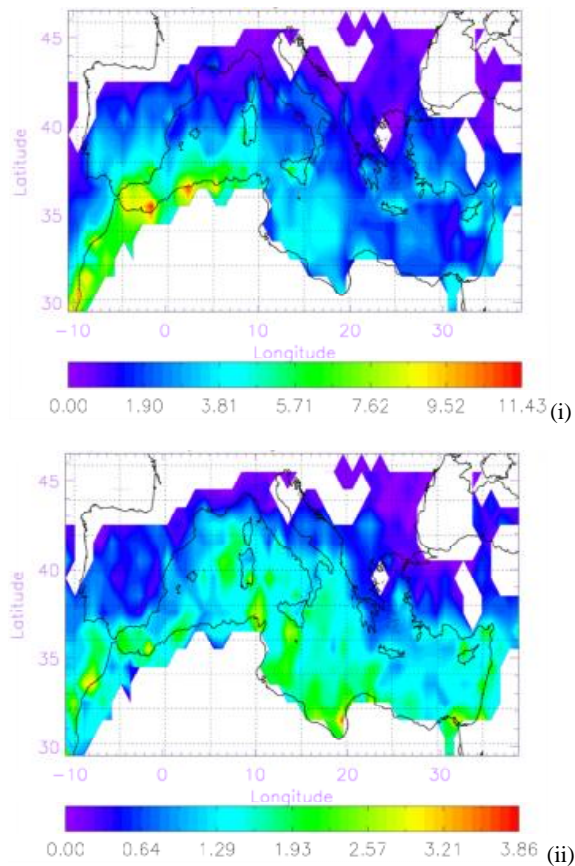
**Fig. 3.** (i) Computed correlation coefficients, between surface  $PM_{10}$  concentrations and AODs, during desert dust episodes, for various stations in the Mediterranean basin. Red and blue circles represent stations where the selected DD episodes are greater/equal and less than 10, respectively. (ii) Seasonal scatterplots, for all stations, between surface  $PM_{10}$  concentrations and AODs, during desert episodes.

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**Fig. 4.** (i) Scatterplot between AERONET and MODIS-Terra AODs at 550 nm under desert dust episode conditions and (ii) Average volume size distribution for all observations (blue curve) and desert dust episode conditions (red curve). The error bars represent the standard deviation values.

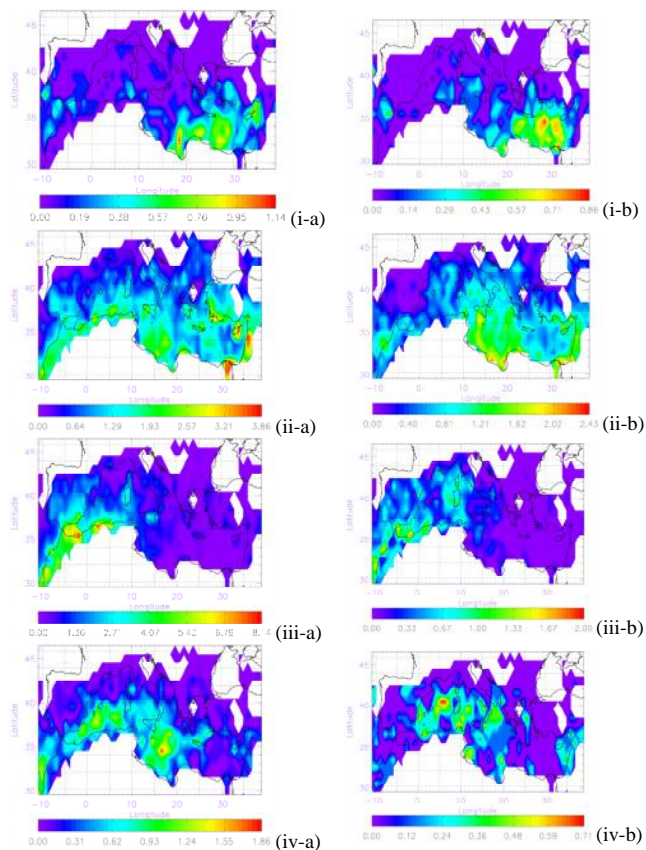


**Fig. 5.** Geographical distributions of the frequency of occurrence (episodes/year) of: **(i)** strong and **(ii)** extreme DD episodes over the broader Mediterranean basin for the period 2000–2007.



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**Fig. 6.** Seasonal geographical distributions of the frequency of occurrence (episodes/year) of: **(a)** strong and **(b)** extreme DD episodes taking place in: **(i)** winter, **(ii)** spring, **(iii)** summer and **(iv)** autumn, over the Mediterranean basin for the period 2000–2007.

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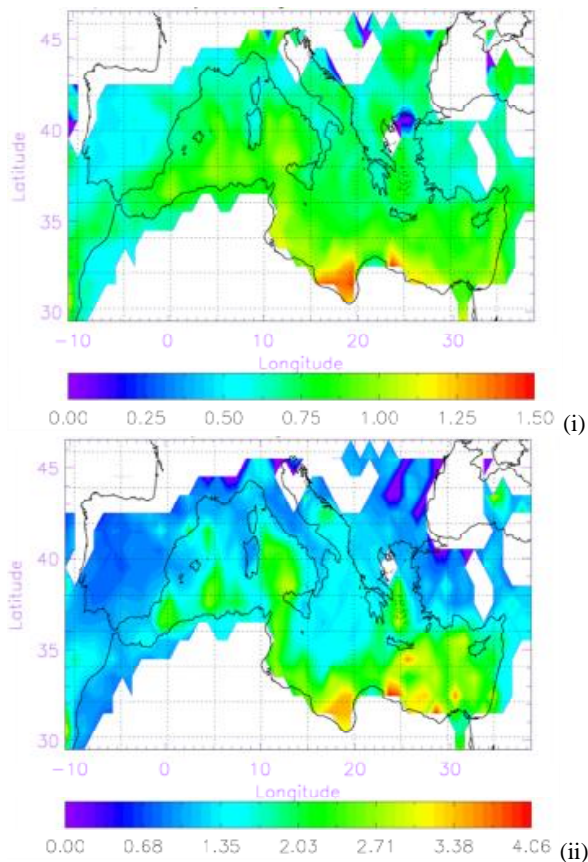
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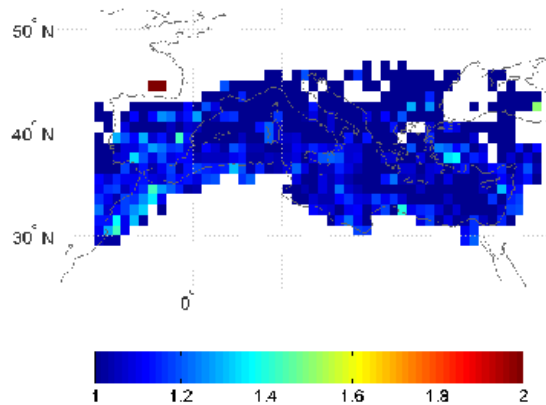
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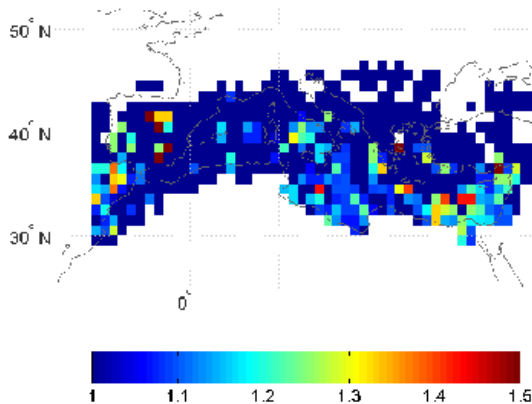
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**Fig. 7.** Geographical distributions of the intensity, in terms of AOD<sub>550nm</sub>, of: **(i)** strong and **(ii)** extreme DD episodes over the broader Mediterranean basin for the period 2000–2007.



(i)



(ii)

**Fig. 8.** Geographical distributions of the duration (in days) of: **(i)** strong and **(ii)** extreme DD episodes over the broader Mediterranean basin for the period 2000–2007.

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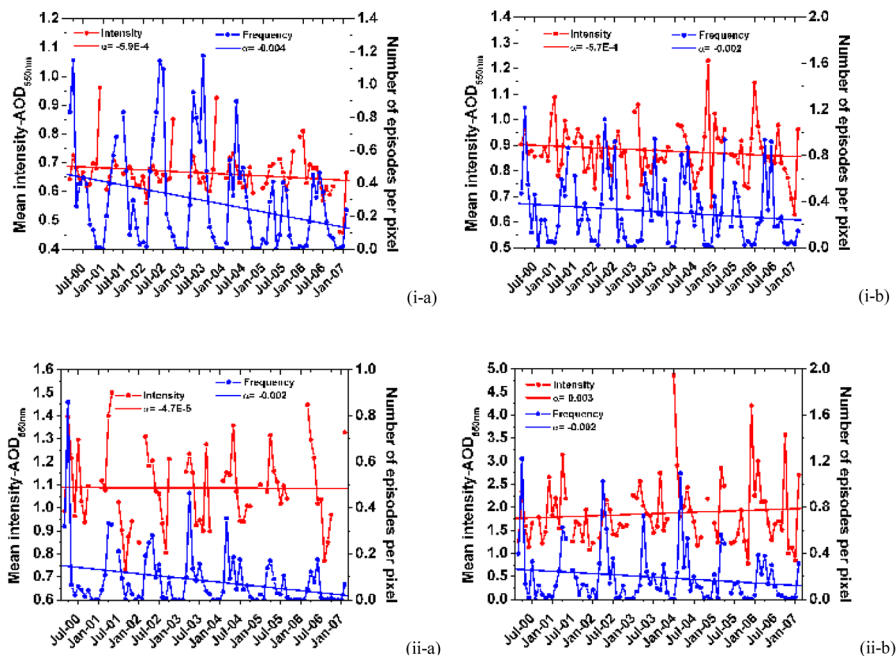
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**Fig. 9.** Inter-annual variation and tendencies of mean frequency (blue curves) and intensity (red curves) of: **(i)** strong and **(ii)** extreme DD aerosol episodes taking place over: **(a)** land and **(b)** sea surfaces of the broader Mediterranean basin for the period 2000–2007.

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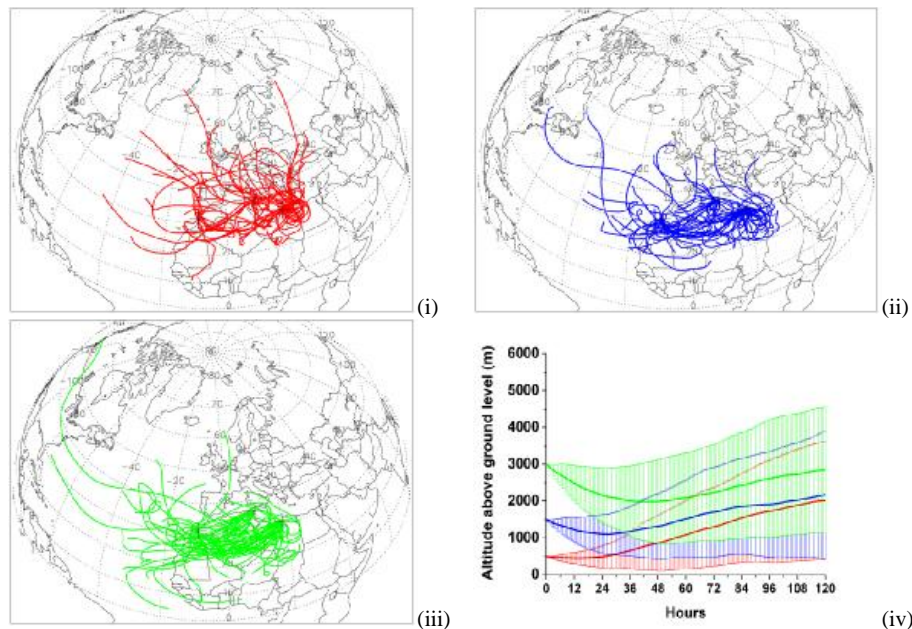
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**Fig. 10.** Five (5) day back trajectories ending at locations of the Mediterranean basin having undergone the maximum frequencies of extreme DD aerosol episodes over the period 2000–2007. Shown are: back trajectories ending at different heights a.g.l., namely at: **(i)** 500 m (red lines, 10i), **(ii)** 1500 m (blue lines, 10ii) and **(iii)** 3000 m (green lines, 10iii) above ground level, and **(iv)** the average altitude and associated standard deviations of all trajectories ending at 500 m, 1500 m and 3000 m, during their 5 day course.

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