Atmos. Chem. Phys. Discuss., 11, 2549–2609, 2011 www.atmos-chem-phys-discuss.net/11/2549/2011/ doi:10.5194/acpd-11-2549-2011 © Author(s) 2011. CC Attribution 3.0 License.



This discussion paper is/has been under review for the journal Atmospheric Chemistry and Physics (ACP). Please refer to the corresponding final paper in ACP if available.

## Airborne observations of mineral dust over Western Africa in the summer monsoon season: spatial and vertical variability of physico-chemical and optical properties

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Received: 31 December 2010 - Accepted: 10 January 2011 - Published: 24 January 2011

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Published by Copernicus Publications on behalf of the European Geosciences Union.





### Abstract

We performed airborne measurements of aerosol particle concentration, composition, size distribution and optical properties over Western Africa in the corridor  $2-17^{\circ}$  N and  $3-5^{\circ}$  E. Data were collected on board the French ATR-42 research aircraft in June–July

<sup>5</sup> 2006 as part of the African Monsoon Multidisciplinary Analysis (AMMA) intensive field phases in June–July 2006 using the AVIRAD airborne aerosol sampling system.

The aerosol vertical distribution was documented on an almost daily basis. In particular, the vertical distribution of mineral dust emitted locally by Mesoscale Convective Systems (MSC) was distinguished from that of mineral dust that was transported from

- the Saharan by the African Easterly Jet (AEJ). Mineral dust emitted in the Sahel by convection-driven erosion was mostly confined in the boundary layer. One episode of injection of Sahelian mineral dust in the AEJ was observed. The single scattering albedo of Sahelian dust is lower than that of Saharan dust, owing to different mineral-ogy and size distribution. Nonetheless this difference is within 0.02 (single scattering
- albedo unit) and does not significantly alter the radiative perturbation at the surface or at the top of the atmosphere. Our investigation provides with further experimental evidence of the limited dynamical evolution of the number size distribution during two days of transport. Whereas the chemical composition is clearly depending on the origin of mineral dust, no indications are of a possible dependence on the number size distribution.

### 1 Introduction

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The role of mineral dust on climate change is prominent but still largely uncertain, as pointed out by the latest Intergovernmental Panel on Climate Change's report (Forster et al., 2007). Dust directly impacts the radiative budget by scattering and absorption of solar and infrared radiations, and by altering the microphysics and radiative properties of clouds and ice layers (DeMott et al., 2003; Koehler et al., 2009). Mineral dust also





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represents a significant input of nutrients (e.g., iron and phosphorous) to remote marine and terrestrial ecosystems (Jickells et al., 2005).

Africa is the world's largest source of mineral dust aerosols (Delmas et al., 2005; Laurent et al., 2008). Mineral dust steaming out of Africa towards the Atlantic Ocean and

the Mediterranean Sea is emitted from various localized sources in the arid and semi-5 arid areas of the Sahara and Sahel, differing in intensity, seasonality and frequency (Goudie and Middleton, 2001; Prospero et al., 2002; Laurent et al., 2008). Because of their intensity and persistence throughout the year, up to very recently most of the research has focused in studying the emission and properties of dust emitted from arid areas in the Sahara (Pye, 1987; Soklik et al., 2001; Laurent et al., 2008).

More recently, various authors have invoked the emission of mineral dust by wind erosion over the semi-arid areas of the Sahel in order to explain the summertime maximum in the aerosol load observed by satellites over Western Africa (Tegen and Fung, 1994; Yoshioka et al., 2005; Forster et al., 2007). The long-time record of horizontal visibility from meteorological stations (N'Tchayi et al., 1994; N'Tchayi Mbourou et al., 1997; Goudie and Middleton, 2001) shows that the atmospheric content of mineral

dust in the Sahelian part of Western Africa is elevated and shows a very pronounced seasonal cycle despite high variability, both at the daily and inter-annual scales. This is largely explained by the alternance of two meteorological regimes. In boreal winter, the

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"Harmattan", a northeastern dry cold wind, is responsible for intense dust emissions 20 in desert areas and very efficient transport towards the gulf of Guinea. At contrary, during summer, due to the northern displacement of the Inter Tropical Convergence Zone (ITCZ), the Sahel experiences, in the lower layers, the wet monsoon flow from south-west. This shift induces a progressive organization of convection from isolated cells to large-scale propagating systems. 25

At the beginning of the rainy season, these systems, often not precipitating, are efficient in producing mineral dust by aeolian erosion over bare soils. As the monsoon season progresses, the precipitation is enhanced inducing the scavenging of dust transported from remote sources and the local wind intensity is reduced. These phenomena





prevent local aeolian erosion in the middle of the rainy season (Abdourhamane Touré et al., 2011).

The summertime column dust content over the Sahel results therefore from the superimposition of local emission and remote transport of dust emitted in the Sahara.

- <sup>5</sup> Because these two large source areas have different mineralogy (Pye, 1987; Claquin et al., 1999; Caquineau et al., 2002), mixing or layering of dust transported from the Sahara and locally emitted by convection over the Sahel should be detectable through differences in the com-position over the atmospheric column. The different residence times should also result in different size distribution, locally emitted dust being in prin-
- ciple richer in coarse particles than long-range transported. The impact of transported and locally emitted dust is therefore expected to be different. In particular, differences in physico-chemical properties of Saharan and Sahelian mineral dust are expected to be evident in the optical properties of scattering and absorption, and should result in differences in the direct radiative impact.
- In this paper, we present new data on the physico-chemical and optical properties that were acquired during twenty research flights over Western Africa between Niamey and Cotonou as part of the African Monsoon Multidisciplinary Analysis (AMMA) intensive field phases in June–July 2006. AMMA is an international project to improve our knowledge and understanding of the West African Monsoon (WAM) and its variability
- with an emphasis on daily-to-interannual time scales (Redelsperger et al., 2006). It provided with a first time opportunity of performing observation of Sahelian mineral dust in a systematic way, that is, coupling ground-based, airborne, spaceborne observations and modeling.

The experimental strategy of the AMMA program was based on embedded multiyear, seasonal and intensive observation periods (Janicot et al., 2008; Lebel et al., 2010). The summertime aircraft observations described in this paper were conducted from Niamey, in Niger, as part of the Special Observing Periods (SOPs) dedicated to the fine description of aerosols and convection. In particular, the present data were collected during the pre-monsoon and early-monsoon periods (called SOP1 and





SOP2\_a1, respectively), when summertime dust emissions from the Sahel are at their most. These observations are to be contrasted with those performed in wintertime during the so-called SOP0 period when local emissions do not occur and dust is purely transported from various sources in the Sahara (Haywood et al., 2008; Rajot et al., 2008; Osborne et al., 2008; Formenti et al., 2008).

Aircraft exploration of the atmospheric column up to 6 km was performed in the area between  $13^{\circ}$  and  $17^{\circ}$  N to illustrate the vertical layering of the dust content under different emission and transport conditions. In addition, exploration of the area south of  $13^{\circ}$  N is also presented to illustrate the time evolution of the meridian extent of the dust transport region with the progression of the monsoon season.

#### 2 Instrumentation

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Instruments have been operated onboard the Service des Avions Français Instruments pour la Recherche en Environnement (SAFIRE) ATR-42. With the necessary payload for AMMA, the aircraft had a maximum endurance of 4 h. Minimum flight altitude is 300 m over land. Ceiling is approximately 7 km. The aircraft was equipped with basic sensors for measuring the radiative, dynamics and thermodynamics properties of the atmosphere and also turbulence (although not used here) (Saïd et al., 2010).

Aerosol sampling was performed using the AVIRAD aerosol sampling system newly developed for AMMA. This system is described in detail in Formenti et al. (2011). It consists of an iso-axial and isokinetic inlet whose 50% passing efficiency has been

<sup>20</sup> consists of an iso-axial and isokinetic inlet whose 50% passing efficiency has been estimated at  $9\,\mu\text{m}$  in diameter. At the cruise speed of the ATR-42 ( $100\,\text{m}\,\text{s}^{-1}$ ) the delivered flow rate is  $350\,\text{l}\,\text{min}^{-1}$ .

This allows connecting the inlet to multiple instruments in parallel. Various straight sampling lines depart from the inlet to provide: (i) straight-levelled run 2-min averaged <sup>25</sup> measurements of the particle absorption coefficient at 370, 470, 520, 590, 660, 880 and 950 nm (aethalometer model AE31, Magee Sci.); (ii) 1-s measurements of the particle scattering and backscattering coefficients at 450, 550 and 700 nm (nephelometer



model 3596, TSI Inc.); (iii) 6-s resolution number size distribution between 0.3 and  $20 \,\mu\text{m}$  in diameter (optical particle counter model 1.108, GRIMM); and (iv) bulk aerosol composition and mass concentration by on-line filter sampling on polycarbonate membranes (Nuclepore, Whatman).

#### 5 2.1 Aerosol particle chemical composition

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Bulk samples of aerosol particles were obtained by filtration onto 42-mm diameter polycarbonate membranes (nominal pore size  $0.4 \,\mu$ m Nuclepore, Whatman) on polyethylene supports. Samples were collected only during horizontal flight legs lasting not less than 20–30 min in order to guarantee sufficient loading of the filter samples. Immediately after each flight, the loaded filters were stored in Petri dishes.

Samples were analyzed by wavelength-dispersive X-ray fluorescence (WD-XRF). WD-XRF analyses have been performed using a PW-2404 spectrometer by Panalytical. Excitation X-rays are produced by a Coolidge tube ( $I_{max}$ =125 mA,  $V_{max}$ =60 kV) with a Rh anode; primary X-ray spectrum can be controlled by inserting filters (AI, at different thickness) between the anode and the sample. Each element was analyzed three times, with specific conditions (voltage, tube filter, collimator, analyzing crystal and detector), lasting 8 to 10 s.

Data were collected for 9 elements (Na, Mg, Al, Si, P, K, Ca, Ti, Fe) using SuperQ software. The elemental mass thickness (μg cm<sup>-2</sup>), that is, the analyzed elemental
mass per unit surface, was obtained by comparing the filter yields with a sensitivity curve measured in the same geometry on a set of certified geo-standards (ANRT GS-N). These geo-standards were crunched in order to reduce and homogenize the grain size. Then they were deposited at different concentrations (<150 μg cm<sup>-2</sup>) on Nuclepore filters. The visual inspection of the deposit by electron microscopy suggested that
grains are generally smaller than 5 μm diameter. The atmospheric elemental concentrations were finally calculated by multiplying the analyzed elemental mass thickness by the ratio between the analyzed and the collection surfaces of each sample (16 and





uniform. Finally, for the AMMA samples, the limit of quantification of the technique ranged between  $3 \times 10^{-3}$  and  $13 \times 10^{-3} \,\mu g \, cm^{-2}$  depending on element.

In addition, the iron oxide content was determined with the adapted CBD-method developed by Lafon et al. (2004). The iron oxide content is defined as the fraction of iron which is not structural, that is accurring in the structural lattice of silicates (Kariakhoff

- <sup>5</sup> iron which is not structural, that is occurring in the crystal lattice of silicates (Karickhoff and Bailey, 1973). There are two major iron oxides in mineral dust (hematite, FeO3, and goethite, FeO·OH) and these are responsible for its light-absorption properties in the UV-visible (Sokolik and Toon, 1999; Alfaro et al., 2004; Lafon et al., 2006). This method is an adaptation for aerosol filters (with typical material mass less than 500 μg)
- <sup>10</sup> of the classical method of Mehra and Jackson (1960) for soil analysis. The method uses the CBD reagent to dissolve iron oxides selectively via reduction.

# 2.2 Aerosol particle size distribution, total number concentration, scattering and absorption coefficients (dry state)

The particle number size distribution between 0.3 and 20 µm was measured on 15 size classes in diameter by a GRIMM optical particle counter (GRIMM Inc., model 1.108) operated at 6-s time resolution. The instrument was operated by the external pump in order to overcome the limitations of the internal pump in overcoming large pressure differences. Values were acquired in terms of the equivalent optical diameter of polystyrene sphere latex (PSL) used for calibration (refractive index  $\tilde{n}$ =1.59–*i*0 at

- <sup>20</sup> 780 nm). The complex refractive index needed to correct the data to geometric diameter (Liu and Daum, 2000) was set to 1.53–0.002*i*, in accordance with the published values available in the literature (see the compilation by Redmond et al., 2010). The number fraction of particles smaller than 0.3 µm diameter was not probed in the absence of an available instrument. This is not a major issue for the observations on
- <sup>25</sup> mineral dust we conducted during the AMMA project (see Sect. 5.1). Otto et al. (2007) have demonstrated that up to 90% of light extinction of mineral is due to particles larger than 0.3 µm in diameter. When needed, the missing fraction was accounted for by performing log-normal fitting of the measured size distribution. This way of estimating the





magnitude of the accumulation mode is of course not very well constrained because the first particle size bin of the GRIMM optical counter is already on the decreasing slope of the mode. Sensitivity tests performed by calculation light extinction for a variable height and the geometric diameter of this fine mode confirmed that the calculated s extinction is not very sensitive (within 10–15%) to the exact shape of it.

The particle volume scattering and backscattering coefficients (dry state) were measured at 450, 550 and 700 nm wavelength using a three-wavelength integrating nephelometer (model 3563, TSI Inc., St. Paul, Minnesota). The instrument was operated at 301 min<sup>-1</sup> and data were acquired at 1-s resolution. It was factory-calibrated and checked prior and after the campaign. The measured scattering and backscattering coefficients have to be corrected for angular truncation and Lambertian non-idealities (Anderson et al., 1996). Truncation corrections are related to the restricted angular opening of the nephelometer sensing volume with respect to the ideal case (Anderson et al., 1996). Corrections were performed by Mie-calculations using the measured size distributions corrected for the effect of the refractive index and the missing particle fraction in the accumulation mode. The real part of the refractive index was left unvaried

with the wavelength (1.53), whereas the imaginary part was set to 0.003 at 450 and 550 nm and 0.002 at 700 nm (Redmond et al., 2010).

A spectral aethalometer (Magee Sci., model AE31) was used to measure the particle absorption coefficient at 370, 470, 520, 590, 660, 880 and 950 nm. The instrument measures the light attenuation through a particle-laden quartz filter spot compared to a portion of the same filter medium that is left unexposed to the air stream and that is used as a reference. Because of its operating principle, the instrument can be operated only on straight and levelled runs at constant altitude when the air flow through the filter

is constant. On those occasions, the reading of the internal mass flow meter remained constant within 2%. Data were acquired at 2-min resolution and the flow rate was set to 201 min<sup>-1</sup>. The use of the aethalometer on pressurised aircraft is commonly critized because the instrument is not fully leak-tight and also because of slowness due to the cycling of the seven LED-light sources. These are surely serious shortcomings.





A serious screen on the data acquired at low-concentration was performed in order to identify possible contamination from the cabin.

The spectral light attenuation measured by the aethalometer as a function of time were corrected to yield the aerosol absorption coefficient according to Weingartner et

- <sup>5</sup> al. (2003). The empirical correction factor by Weingartner et al. (2003) allows to take into account the effect of multiple scattering of the light beam at the filter fibres of the unloaded filter (C factor in the Weingartner et al., 2003, paper), and the "shadowing effect" of absorbing particles progressively masking the incoming light whilst accumulating on the filter (R(ATTN) in Weingartner et al., 2003). This second effect depends
- on the aerosol single scattering albedo, and it is lower for scattering than for absorbing particles. Conversely, the Weingartner et al. (2003) correction scheme does not allow taking into account the miscounting of particle scattering as absorption, which we accounted for by applying the correction factor proposed by Collaud Coen et al. (2010). In this work we set the C factor to 2.14 (±0.21). For the range of single scattering albedo
   values expected to mineral dust (Redmond et al., 2010), R(ATTN) varied between 1
- (first measurement after the spot change, ATTN=0) and 0.84 (last measurement after the spot change, ATTN=100), independently of the wavelength.

Particle number concentrations, and scattering and absorption coefficients were converted to standard temperature and pressure (STP) using T=273 K and P=1013.25 hPa. The scattering, backscattering and absorption coefficients are expressed in Mm<sup>-1</sup> (1 Mm<sup>-1</sup>=10<sup>-6</sup> m<sup>-1</sup>).

### 3 Area and period of study

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The large and regional scale features of the early monsoon season in 2006 over Western Africa are discussed by Janicot et al. (2008). Their analysis shows that the 2006 season was nearly-average but characterised by a later onset of the monsoon (about 10 days) than expected from climatology. From June to July, the Intertropical Convergence Discontinuity (ITD) at the surface between the monsoon wind and the Harmattan wind, moved from about 15° N to about 20° N (Janicot et al., 2008; Lothon et al., 2008).





This resulted in an increase influence of moist air transported in the south-westerly monsoon flow over the study area.

With respect to the dust content and origin, the analysis of satellite images, groundbased and airborne observations indicated that during SOP1 and SOP2\_a1 mineral dust dominated the aerosol content north of 10° N. This is shown by the time series of aerosol optical depth (AOD) at 870 nm and the Angstrom exponent (Fig. 1) measured by the sunphotometers of the Aerosol Robotic Network (AERONET) (Holben et al., 1998; 2001) located in the area.

Three stations located in the proximity of the area of observation had a significant
number of observations: Banizoumbou (Niger; 13° N, 2.5° E), Cinzana (Mali; 13° N, 5° W), and Djougou (Benin; 9° N, 1° E). Banizoumbou and Cinzana are located in semi-arid areas where mineral dust dominates, whereas Djougou represents vegetated areas mostly under the influence of the monsoon flow. These characteristics are reflected in the sunphotometer data showing higher AOD values and lower Angstrom exponent
values in the two dust-dominated stations.

Dust in the Banizoumbou area was either produced by local erosion (by convective systems or monsoon flow, Sow et al., 2009) or transported. Convective systems were either isolated cells of very local origin or large-scale organised systems, most of the time originating in the Aïr Mountains and on the Jos Plateau (Janicot et al., 2008).

- <sup>20</sup> During the flying periods, sixteen erosion events were detected in Banizoumbou on the basis of the record of the grain saltation (horizontal movement of the dust grain induced by increased wind speed). These corresponded to isolated cells or more organised convective systems, fifteen of which were followed by heavy rain (Desboeufs et al., 2010). These events mostly occurred in the late afternoon or at night and produced were bigh dust expectatorians measured at the graund level (Martinerane et al. 2010).
- very high dust concentrations measured at the ground level (Marticorena et al., 2010). That is why the increase of AOD due to local dust emission does not really show up in the level 2.0 AERONET daytime series. Furthermore, the latter tends to be biased to lower values since events characterised by extremely high concentrations and associated with clouds are discarded.





Transported dust originated from the Sahara, from sources north east of Niger, mostly the Aïr/Teneré, Tibesti/Ennedi corridor, and the Darfour areas (Kalu, 1979; Goudie and Middleton, 2001; Prospero et al., 2002). In summertime easterly transport occurs above 1.5 km in the Saharan Air Layer (SAL) (Carlson and Prospero, 1972) through the African Easterly Jet (AEJ) that blows over 1.5 km and is strongest at around 4.5 km at the latitude of Niamey (Kalapureddy et al., 2010). This contrasts with the surface transport pattern occurring during wintertime (Kalu, 1979; Chiapello et al., 1995; Osborne et al., 2008; Heese and Wiegner, 2008; Kim et al., 2009; Léon et al., 2009).

The operating region of the ATR-42 during the AMMA SOP1 (1–15 June) and SOP2\_a1 (1–15 July 2006) intensive phases are shown in Fig. 2.

The aircraft, based in Niamey (13° N, 2° E), the capital of Niger, performed twenty measurement flights. The general flight strategy included several fast ascents and descents to sound the vertical atmospheric column up or down to the maximum/minimum flight level (from 6 km to 300 m above ground level (a.g.l.)), in order to detect interesting meteorological and particle layering. In addition, 20–30 min long straight and levelled runs (SLRs) or, less frequently, closed triangular patterns were flown at fixed altitudes for aerosol particle collection. An example of flight plan is shown in Saïd et al. (2010). As described in Reeves et al. (2010) flights from SOPs 1 and 2a\_1 served various objectives of the AMMA science plans via three Intensive Observational Periods (IOPs): IOP 1.1 (surface-atmosphere: inter-tropical front and heat low surveys); IOP 1.2 (surface-atmosphere-aerosol: squall-line related to aerosol emissions surveys);

IOP 1.3 (surface-atmosphere: north-south land-ocean-atmosphere interactions surveys). A summary of flight information (date, location, aim) is provided in Table 1.

Most of the flights (fourteen out of nineteen) were conducted in coordination with the AMMA ground-based super-site of Banizoumbou (13.5° N, 2.5° E, 250 m a.s.l.), located at about 60 km east of Niamey (Rajot et al., 2008; Marticorena et al., 2010). In the attempt of accommodating the various IOPs scientific objectives, clear-sky conditions prior and post the occurrence of mesoscale convective events were sampled preferentially. Due to flight security restrictions, flights were conducted during daytime only.





## 4 Results

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## 4.1 Aerosol spatial variability

The north-south gradient of the aerosol particle load observed during the nineteen ATR-42 flights in SOP1 and SOP2\_a1 can be explored by looking at the variability of the particle number concentration as a function of latitude (Fig. 3).

Data are number concentrations for particles of optical diameter larger than  $0.3 \,\mu m$  (dN<sub>0.3</sub>) and for particle of optical diameter larger than  $1 \,\mu m$  (dN<sub>1.0</sub>) averaged in bins of 0.25 degrees in horizontal resolution and 100 hPa in vertical resolution. Data are then presented for three atmospheric levels: below 1 km, corresponding to the mean bound-

- ary layer depth during the campaign; between 1 and 3 km, corresponding to the mean extent of the African Easterly Jet (AEJ) and to the expected maximum dust concentration in the Saharan Air Layer (SAL); and above 3 km, for decreasing concentrations. This distinction has been based on the examination of the aerosol and thermodynamic vertical profiles which will be presented later in the paper.
- <sup>15</sup> The total number of dN<sub>0.3</sub> particles remained lower than 160 cm<sup>-3</sup> STP during June 2006 (SOP1, Fig. 3a). North of 12° N, comparable concentrations were observed in the boundary layer and above. South of 12° N, the north-south particle concentration at the surface showed no significant gradient but a slight increasing trend towards the north. An isolated maximum at the surface was observed around 6° N in correspondence to
- the position of the town of Cotonou. With over 1.2 million inhabitants and a large active industrial harbor with transiting oil and metals, Cotonou is the economic capital of Benin (Minga et al., 2010).

The dN<sub>0.3</sub> particle loading during July 2006 (SOP2\_a1, Fig. 3a) behaved differently. Values remained lower that 100 cm<sup>-3</sup> STP with the exception of an isolated maximum (~220 cm<sup>-3</sup> STP) observed at 2° N above 1 km. This corresponds to the vertical sounding which was performed on 04 July during flight V030, in an area where satellite images (OMI, MODIS, not shown) suggested transport of biomass burning plumes from the Southern Hemisphere. The CO and O<sub>3</sub> concentrations measured onboard the





aircraft reached 420 ppbv and 100 ppbv at 1500 m, instead of 150 ppbv and 40 ppbv in the boundary layer. Concentrations north of 12° N were lower than during SOP1 (below 50 cm<sup>-3</sup> STP), and that almost independently of altitude. South of 12° N, concentrations increased between 6° and 10° N. A possible explanation for this increase is the presence of numerous small forest fires which were detected during the transect flights.

On the opposite, the comparison of the dN<sub>1.0</sub> particle number concentration values measured during the SOP1 and the SOP2\_a1 periods shows that in the boundary layer the north-south gradient remained consistent during the two flying periods. Highest concentrations were observed towards the north of the domain, decreasing towards the south. This gradient reflects the northward displacement of the ITD from June to July. Concentrations higher that 5×10<sup>-2</sup> cm<sup>-3</sup> STP were found down to 10° N in June, but only north of 13° N during July. The threshold is indicated with the horizontal black line in Fig. 3b. This corresponds to the progressive northward advance of the monsoon layer, characterized by smaller particles. Observations between 13.5° N and 15.5° N, where the atmospheric column was more frequently sounded, show that concentra-

tions can be elevated also above the surface layer.

Because the elemental composition by bulk analytical techniques such as XRF is largely determined by the coarse fraction, the dN<sub>1.0</sub> particle distribution agrees well with the spatial distribution of the elemental concentrations which have been obtained by filter sampling and XRF analysis. As an example, the north-to-south gradient of elemental AI is shown in Fig. 4.

The lower number of data point is due to the coarser resolution of the filter sampling. In the absence of other significant sources, elemental AI can be used as a nonambiguous tracer of mineral dust (Mason, 1966; Taylor, 1986). The crustal origin of the other elements measured above detection limit (Mg, Si, P, K, Ca, Fe, and Ti) is put into evidence by calculating their linear regressions with respect to AI. A 2-tailed test of significance shows that the Pearson regression coefficients were comprised between 0.88 and 0.99 for Mg, Si, P, K, Ti and Fe, but only 0.57 for Ca. This is consistent with





the coexistence, in mineral dust, of aluminosilicate species (such as clays, feldspars) traced by AI and calcium carbonate species (such as calcite, dolomite gypsum) traced by Ca (Kandler et al., 2007; Rajot et al., 2008; Formenti et al., 2008). Good correlations were found between Ca and Mg (Pearson regression coefficient of 0.84) and between
 5 Ca and K (Pearson regression coefficient of 0.72).

In order to provide with the range of variability for dust conditions only, Table 2 summarizes the range of variability for samples collected north of 12° N. For sake of comparison with previous campaigns in Western Africa (Formenti et al., 2003, 2008; Rajot et al., 2008; McConnell et al., 2008), Table 2 also reports on the descriptive statistics of the total dust mass.

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The calculations have been restricted to the 51 samples for which Ti was above detection limit. The total dust mass has been estimated in three ways: (i) as the sum of major elemental oxides; (ii) as the sum of elemental concentrations; and (iii) from the AI elemental concentration considering that AI accounts for about 8.13–8.36% of

- the earth's crust mass (Mason, 1966; Taylor and McLennan, 1985). The three estimates of the mass are highly correlated (Pearson regression coefficient>0.96) due to the fact that the mass is dominated by the contribution of Al. Nonetheless, the differences between the three estimates are higher than the expected variability in the crustal composition. In the absence of a direct gravimetric measurement, it is not pos-
- sible to constrain in a robust way the conversion rule allowing estimating the total dust mass. Nonetheless, our and other group previous experience (Prospero et al., 1981; Prospero, 1999; Klaver et al., 2011) suggests that the conversion from the Al elemental concentration is a good representation of the dust mass. The uncertainties in the estimate of the dust mass has important implications when one wants to use measured concentrations for comparison with output of emission models, or to coloulate the mass.
- <sup>25</sup> concentrations for comparison with output of emission models, or to calculate the mass scattering or absorption efficiencies of the aerosol.

The elemental concentrations show the same large variability than during previous airborne campaigns in Western Africa (see Table 1 in Formenti et al., 2008). However, they are lower than those measured during wintertime in the Niamey area during the





SOP0 phase (Formenti et al., 2008) and only comparable to those observed for shortrange transport conditions such as those encountered offshore Senegal during the Saharan Dust Experiment (SHADE) and Dust Outflow and Deposition to the Ocean (DODO1) airborne campaigns (Formenti et al., 2003; McConnell et al., 2008). As

- a matter of fact, the long-term record of surface concentration measurements reported by Marticorena et al. (2010) illustrates that dust concentrations in June and July are characterised by mean values of the same order of magnitude as wintertime ones, but by a much higher variability due to the episodic character of both emission (by MSCdriven high winds) and deposition (by MSC-induced precipitation). These authors show
- that concentrations might vary by five orders of magnitude on a few-minute scale. It is clear that the limited frequency of aircraft sampling, which mostly was conducted at solar midday, at least height hours later than MSC-driven emissions, under-samples the full variability of the elemental concentrations. In addition, at times, the altitude of SLRs for particle collections had to be accommodated to other scientific requests, and might not have been chosen at the height where the highest concentrations were
- and might not have been chosen at the height where the highest concentrations were observed.

#### 4.2 Aerosol vertical profiles

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The chronological sequence of the vertical profiles of  $dN_{0.3}$  and of the spectral scattering coefficient  $\sigma_s$  at 450, 550, and 700 nm are presented in Fig. 5a and b for the June (SOP1) and July (SOP2\_a1) periods. The information relative to the time and geographical location of those profiles is summarized in Table 3.

The variability of the shape and magnitude of the profiles is evident. Particle concentrations varied between 25 and  $200 \text{ cm}^{-3}$ , whereas plumes where the scattering coefficient  $\sigma_s$  reached  $800 \text{ Mm}^{-1}$  were observed. The spectral dependence of the aerosol scattering provides with information on the size distribution and therefore indirectly on the nature of the aerosol. So, spectrally-neutral plumes are indicative of aerosols such as mineral dust with a significant coarse mode (Formenti et al., 2003; Osborne et al., 2008) whereas plumes for which the spectral dependence of the particle





scattering coefficient is well pronounced are due to aerosols dominated by the accumulation mode such as pollution or biomass burning (Formenti et al., 2002; Johnson et al., 2008).

Plumes of elevated concentrations were observed at different altitudes. Surface layers of thickness variable between 1 and 3.5 km were observed during flights V017, V018, V019, V026, V028 and V031. Elevated aerosol plumes were observed above 1.5 km up to approximately 5 km, either as a single layer (flights V025, V026, V030, V033, V036) or by a more stratified structure (flights V020, V021, V022, V024, V029, V032, V034). Layer of very low concentrations decreasing monotonically with height were also observed (profiles during flights V024, V027 and V035). Such diversity is consistent with the climatology of aerosol profiles in the area proposed by Cavalieri et al. (2010) based on ground-based and spaceborne lidar observations. Over land, the aerosol layers were always spectrally neutral, indicating than mineral dust dominated the aerosol load. Over sea (flights V024 and V030) plumes with a more pronounced spectral dependence were observed either close to the surface or above 1 km.

The aerosol vertical structure is strongly related to the thermodynamic structure of the atmosphere. This is depicted by the height of the top of the boundary layer ( $Z_i$ , solid line in Fig. 5a and b) and by the shear level ( $Z_s$ , dotted line in Fig. 5a and b) corresponding to the top height of the transition zone between the monsoon and the <sup>20</sup> Harmattan wind regimes. If  $Z_i$  is lower than  $Z_s$  the boundary layer is well isolated from the SAL. There is no possibility of entrainment, therefore aerosols emitted locally at the surface are confined to the boundary layer and cannot mix with dust transported in the SAL by the AEJ. Aerosol mixing by entrainment is possible when  $Z_i$  is equal to  $Z_s$ .

The examination of these thermodynamical discontinuities for the SOP1 and SOP2\_a1 periods indicates that,  $Z_i$  was generally slightly higher lower than  $Z_s$  at the time when the vertical soundings were performed (Fig. 5). The vertical soundings were usually made around midday and the boundary layer went on growing after the sounding. During SOP1, the boundary layer growth made  $Z_i$  reach  $Z_s$  (Saïd et al., 2010) which produced strong moisture exchanges at small scale between the monsoon layer





and the SAL (Canut et al., 2010). However, these exchanges cannot be detected on scalars (such as temperature and water vapour mixing ratio) or particles, when sampled at time resolutions larger than 1 s. Even if there is any exchange between the mixed layer and the SAL, only the average content of the mixing will be detected. The

<sup>5</sup> only mean to estimate the exchanges would have been to sample the time evolution of the concentration, which was not done. During SOP2\_a1,  $Z_i$  less often reached  $Z_s$ due to weaker surface heat fluxes, especially when the flights occurred just after rainy episodes (see V033, V034 and V036). It was also the case for flights performed over the sea (V024 and V030).

#### 10 4.3 Discussion of case studies

The vertical profiles of particle concentration and particle scattering coefficient and the chemical composition will be used to identify the aerosol origin and residence time in the atmosphere.

- Some elemental ratios are particularly informative in this sense. The Si/Al ratio represents the content of alumino-silicate types of minerals (clays and feldspars) and quartz (SiO<sub>2</sub>), whereas the Fe/Ca ratio that of iron-rich minerals (i.e., iron oxides) with respect to calcium carbonates (calcite, dolomites) and calcium sulphates (gypsum). Elemental Fe and Ca are present in alumino-silicate minerals too, but in much lower proportions, and this does not alter the use of the Fe/Ca ratio to trace the proportions of iron-rich and calcium-rich minerals. The regional variability of these ratios is demonstrated in various publications (Chiapello et al., 1997; Claquin et al., 1999; Kandler et al., 2007; Formenti et al., 2003, 2008; McConnell et al., 2008; Rajot et al., 2008). The Fe/Ca ratio allows distinguishing Saharan from Sahelian dust, the former having lower Fe/Ca ratio
- (<~2) than the latter (>~2). In particular, the Bodélé depression, considered to be the
   largest single source region in Africa (Prospero et al., 2002; Bristow et al., 2010), is depleted in Fe and therefore is characterised by a typical Fe/Ca ratio of about 1 (Formenti et al., 2008). Whenever possible, the iron oxide content will be provided. However, owing to relatively low concentrations, the latter could be determined only on a very limited number of samples.





#### 4.3.1 Background oceanic and continental conditions

Reference clean background conditions were encountered on a few occasions during the field campaign.

- Vertical soundings during flight V024 (Fig. 5a) represent pristine air oceanic conditions over the Gulf of Guinea at 2° and 4° N. In this figure, the particle number concentration is merged with the 400 nm  $\sigma_s$  curve. Aerosol number concentrations dN<sub>0.3</sub> were almost constant around 50 cm<sup>-3</sup> STP up to 1–1.5 km. So the marine boundary layer that was only 500 m deep exhibited the same particle composition as the whole monsoon layer that extended up to 1 km, suggesting that the boundary layer was deeper on the previous days and that the air mass had not changed. Above, the aerosol number concentrations decreased rapidly to zero. In the same way, the particle scattering coefficient averaged at around 40 Mm<sup>-1</sup> at 550 nm within the monsoon layer. The
- accentuated spectral dependence indicates the predominance of accumulation mode particles in the size distribution. The small peaks at 1.5, 1.8 and 2.8 km correspond
- to peaks in the O<sub>3</sub> vertical profiles and to a slight increase in the CO sounding. This suggests that these are aged plumes of combustion origin. As it will be discussed in the following section, an episode of transport of biomass burning was observed at 2° N during flight V030. In the monsoon layer this flight otherwise represented background oceanic condition.
- <sup>20</sup> Profiles collected during flights V017, V027, and V035 (5 June, 1 and 12 July) represent background continental conditions north of 13° N. Conditions sampled during flights V017 and V035 corresponded to the cleaning up of the atmosphere following the passage of rather localized precipitating convective systems which occurred over Banizoumbou on 4 June and 11 July (6 and 20 mm of cumulated rain, respectively). These profiles were characterized, at all altitudes, by dN<sub>0.3</sub> and  $\sigma_s$  values lower than 50 cm<sup>-3</sup> and 75 Mm<sup>-1</sup>, respectively. These values are comparable to the pristine oceanic conditions, except for the fact that the ratio  $\sigma_{s,550 \text{ nm}}/dN_{0.3}$  was lower than 1 for





oceanic conditions but higher than 1 for continental conditions. As it will be shown in

Sect. 4.3, this suggests that the number size distributions are different. As a matter of fact, background continental conditions are enriched in particles larger than 0.5  $\mu$ m in diameter whereas background oceanic conditions are enriched in fine particles smaller than 0.5  $\mu$ m. This suggests that the marine influence is minor at Banizoumbou.

Notable differences are also reflected in the mean Al concentrations. In the boundary layer, these were 0.26 (±0.09) μg m<sup>-3</sup> for flights over ocean (V024 and V030) but varied between 2.0 (±0.4) μg m<sup>-3</sup> and 4.5 (±0.2) μg m<sup>-3</sup> for flights V017, V027 and V035. The mean Si/Al and Fe/Ca values for these three flights were close to typical values for mineral dust and also very similar. For Si/Al values, they were 2.1 (±0.1), 2.2 (±0.1) and 2.2 (±0.1), respectively; whereas for the Fe/Ca ratio, 2.3 (±0.6), 2.2 (± 0.5) and 1.8 (±0.6). Only the Si/Al ratios were above detection limits on flights over the ocean and equal to 2.5 (±0.2).

#### 4.3.2 Advection of southern hemispheric biomass burning

An episode of advection of southern hemispheric biomass burning in the Northern <sup>15</sup> Hemisphere was detected at 2° N during flight V030 on 4 July (Fig. 5b).

Above 1.5 km, particle concentration  $dN_{0.3}$  was as high as  $350 \text{ cm}^{-3}$  STP, and the particle scattering coefficient at 550 nm averaged at  $100 \text{ Mm}^{-1}$ . Unfortunately the layer was only sounded during the vertical profile, and not sampled at constant altitude, disabling further investigation of its properties, notably of the absorption coefficient and of

- the aerosol composition. Assuming the average single scattering albedo value of 0.91 obtained at 550 nm by Haywood et al. (2003b) for Southern Hemisphere aged biomass burning, the scattering profile can be converted in extinction, which then can be integrated between 1.5 and 2.9 km to yield a lower limit estimate of the column aerosol optical depth. By doing so, the integrated extinction is 0.14, lower than the aerosol optical
- depth values reported for advected biomass burning during the dry season (Johnson et al., 2008). By comparing the profiles in this work and that of Johnson et al. (2008), it is nevertheless clear that the values of extinction are comparable. On the contrary, the profile ceiling is too low with respect to the full extent of the biomass burning plumes,





which has been found to extend up to 6 km (Fiedler et al., 2010). The bottom altitude of the biomass burning plume is in agreement with that shown for transported Northern Hemisphere biomass burning in the dry season (Kim et al., 2009; Léon et al., 2009). This is also in agreement with the vertical profile of the excess in  $O_3$  concentrations presented by Mari et al. (2008) and Thouret et al. (2009), both indicating that transported plumes are generally found around 2 km.

#### 4.3.3 Local erosion events

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Of the sixteen erosion events detected in Banizoumbou, only two were followed by a flight of the ATR-42.

- <sup>10</sup> Flight V018 on 6 June followed a non-precipitating, rather local, MCS which occurred at 03:00 UTC on the previous night. Erosion was only moderate, the number of saltating grains measured at Banizoumbou only reaching 2000 counts per 10 s. In correspondence, the record of particle backscattering provided by the Micro-Pulsed Lidar (MPL) of the Atmospheric Mobile Facility (AMF) of the Atmospheric Radiation
- <sup>15</sup> Measurement (ARM) at the Niamey airport (available online at http://www.arm.gov/ sites/amf/nim/instruments) showed a layer of uplifted aerosols below 1 km between 03:00 and 07:00 UTC, and then a more complex structure with two layers below and above 1 km converging into one only layer as time progressed and the solar warming increased the height of the boundary layer. By 16:00 UTC, only one layer was evident
- <sup>20</sup> between 1 and 4 km. In fact the boundary layer wind had turned from monsoon to Harmattan during the afternoon, following the southern retraction of the ITD, due to surface convection (Saïd et al., 2010).

Flight V028 on 2 July followed a larger organised MCS originated on the Aïr massif. This system reached Banizoumbou just after midnight on 1 July, provoking intense particle saltation (number of saltating particles up to 5000 counts per 10 s), but also intense rain (17 mm of accumulated precipitation).

These observations are consistent with the vertical profiles of dN<sub>0.3</sub> particle number concentrations and of the Fe/Ca and Si/Al ratios shown in Fig. 6, as well as of that of the scattering coefficient  $\sigma_s$  in Fig. 5a and b.





For flight V018 the enrichment in  $dN_{0.3}$  in the boundary layer below 0.9 km is evident when compared to the background continental conditions. These are shown by the red line illustrating the  $dN_{0.3}$  particle concentration during flight V017. The enrichment in particle concentration is less marked for flight V028, possibly due to enhanced depletion by wash-out and rain-out associated with the MCS. In this case, the scattering profile shown in Fig. 5b, and its comparison to that obtained on the previous day (flight V027, 1 July) is more informative.

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Nonetheless, in the two cases, it is clear that the aerosol layer is confined to the surface below 1 km and well separated from the free troposphere. The strong inversion on top of the mixed layer restricts mixing and vertical redistribution of locally-emitted dust (the transition layer between the mixed layer and the SAL is at least 600 m deep for flight V018 during the aircraft vertical ascent, even if there is some mixing afterwards (Saïd et al., 2010). The proximity of the ITD noticed before is also a source of local dust lifting, since the ITD is an area of convergence between the two opposite flows.

- <sup>15</sup> For both flights, differences in the surface and the elevated layers are also evident in the vertical gradient of the Fe/Ca and Si/Al ratios (Fig. 6). For these two flights, the Fe/Ca ratio averaged at 7.4 in the mixed layer. This should indicate enrichment in Fe and depletion in Ca which is consistent with the local Sahelian origin of mineral dust. Observation on samples collected during the day at the ground-based site Banizoum-
- <sup>20</sup> bou are also consistent in showing elevated values of the Fe/Ca above 10. In addition, the quantification of the iron oxide content was possible for two samples, one collected in the local erosion dust layer and one in background conditions. The ratio between Fe in the form of oxide and total Fe is 0.69 and 0.52, respectively. These values are consistent with those observed on the ground at Banizoumbou (J. L. Rajot, unpublished)
- data), and also with those reported by Formenti et al. (2008) concerning the wintertime ground-based and airborne AMMA campaigns. Depending on the way that the total dust mass is estimated, the Fe content in the local dust produced by erosion vary between 3 and 9%, the lower boundary obtained when the total mass is estimated from elemental AI and the upper boundary is obtained when the total mass is estimated as





the sum of the elemental concentrations. This is of course a very large spread which has implication for model calculation of the optical and radiative properties of mineral dust. Above, in the free troposphere, the Fe/Ca ratios were much lower, around 1.2 and 1.5 for flight V018 and V028, respectively, as expected for Saharan dust (Kandler

- and Schütz, 2007). No estimate of the Fe oxide content could be achieved in this layer. It might be surprising to notice that, despite the Fe/Ca ratios being very close, the Si/Al ratios in the mixed layer are different (3.1 and 1.9, respectively for flight V018 and for flight V028). These differences might be linked to the different size distribution, which will be discussed in Sect. 4.3. As a matter of fact, the local dust layer of flight V018 and for flight void be discussed in Sect. 4.3. As a matter of fact, the local dust layer of flight
- <sup>10</sup> V018 is richer in coarser particles than that of flight V028, possibly because of wet deposition which occurred prior this second flight (see Sect. 4.3). As a consequence, during flight V018 the fraction of large quartz particles, which are characteristic of the mineralogy of Sahelian soils (Caquineau et al., 2002), are enriched with respect to the dominantly clays, whereas during flight V028 they were depleted by wash-out. The
- Fe/Ca ratio remained invariant as iron oxides are mostly associated to micron clay particles, mostly kaolinite (Greenland et al., 1968). As a consequence Ca is also supposed to remain invariant. Because there is no calcite in the Sahelian soil, Ca is also associated to clay particles.

#### 4.3.4 Saharan dust transport

An episode of transport of Saharan dust occurred from 9 to 15 June, when the Niamey-Banizoumbou area experienced a persistent easterly flow, transporting mineral dust from the Bodélé depression in Chad and the Sudanese region (Flamant et al., 2009). These dust laden air masses were transported over Niger where, starting from 11 June, they were sampled by the ATR-42 (flights V021, V022, V025, and V026). The relevance of this episode of transport at the regional scale is highlighted also by the satellite images provided in Cuesta et al. (2008) showing that the dust from Bodélé reached also Tamanrasset in Algeria (~22.8° N). The episode is also evident in the times series for Banizoumbou and Cinzana (AOD values up to 2 after DOY 162) and to

a lesser extent in that measured at Djougou (AOD up to 1).





The presence of an elevated dust layer extending from 1 to 4 km, in agreement with the indications of back-trajectory calculations reported by Flamant et al. (2009), is evident in the corresponding vertical profiles in Fig. 5a. A closer look up of the particle concentration and the chemical composition is shown in Fig. 7, illustrating the situation prior the beginning of the episode (8 June, flight V020), on the first two days of the episode (11 and 12 June, V021 and V022), and at the end of the episode (14 and 15 June, V025 and V026).

The complexity and magnitude of the dust layer evolved with time. Highest particle concentrations (up to 250 cm<sup>-3</sup>) and scattering values (up to 600 Mm<sup>-1</sup>) were observed during 11 and 12 June, more than twice as much as the values observed prior the beginning of the period on 8 June (flight V020). The dust layers during 11 and 12 June also showed a complex structure with several sub-plumes less than 1-km deep. Lower values (less than 100 cm<sup>-3</sup> and 200 Mm<sup>-1</sup>) were observed then during 14 and 15 June (Fig. 5a), when the dust layer appearance became less stratified and more resembling

- <sup>15</sup> the classical structure of the SAL as described by Prospero and Carlson (1972) and Carlson and Prospero (1972). From the dynamic and thermodynamic point of view, the bottom of the elevated dust layers coincided with  $Z_s$ , the shear level between the monsoon flow and the AEJ, which most of the time also coincided with  $Z_i$ , the top of the boundary layer. As expected for the SAL, the dust plumes were associated with <sup>20</sup> strong easterly winds, low humidity and stable conditions. They were delinked from the
- monsoon conditions below.

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Figure 7 also shows the vertical profile of the  $dN_{0.3}$  particle concentration and the chemical composition for flight V026, which was conducted above Banizoumbou on 15 June 2006. In this case, despite the enrichment in particle concentration in the surface

<sup>25</sup> layer, the Fe/Ca and Si/Al ratios are close to values expected for Saharan dust and are not supportive of the contribution of local emission to the aerosol load in the boundary layer. The surface peak during flight V026, as that during flight V021, are due to local pollution possibly from the Niamey area.





These considerations are confirmed by the analysis of the elemental composition. As previously noted, the Si/Al mean value for flight V021 is outstanding (3.7±0.3), reflecting the Bodélé origin of the dust. The soil of the Bodélé depression is made of clays, diatom frustules and detrital quartz sand grains (Mounkaila, 2006; Moreno et al., 2006; Todd et al., 2008). The enrichment of Si in the elemental composition of dust transported from Bodélé to the ground-based site of Banizoumbou has been reported

- by Rajot et al. (2008) and Formenti et al. (2008) for the wintertime AMMA SOP0 field phase. On the following days, 12–15 June, the Si/Al dropped to values closer to typical Saharan dust (~2.5), reflecting the fact that the intensity of dust emission from Bodélé 10 was decreasing (Flamant et al., 2009). The corresponding values for the Fe/Ca ratio
- were close to one. As a matter of fact, the Bodélé source is rather depleted in iron oxides which would result in an increase of the Fe/Ca ratio (Formenti et al., 2008). Unfortunately, no direct measurements of the iron oxide content are available for this episode. One should note than the Fe/Ca ratio measured in the boundary layer prior
- the beginning of the episode was higher than 3, likely as the consequence of some minor local dust emission which had occurred the previous night.

Finally, because of the ubiquitous presence of those tracers in the major minerals constituting mineral dust, the use of the elemental composition might mask some of the differences in the dust composition in the boundary layer and the SAL.

#### 20 4.3.5 Sahelian dust transport

An episode of uplift and transport of Sahelian dust was sampled north of Banizoumbou in the period between 6 and 10 July. Flights V032, V033 and V034 were conducted on two consecutive days (6 and 7 July for V032 and V033, respectively) and then on 10 July for V034, in the area east of Banizoumbou between 13° and 15.5° N. During this period the ITD remained north of 17° N. These flights were preceded by a convective system which had passed on the flight area the previous day and had caused 7 mm of rain and by a large MSC which had appeared over the north of Nigeria in the late afternoon of 5 July and that caused intense and long-lasting (more than 1 h) saltation





in Banizoumbou. This first local erosion episode was followed daily by other similar but less intense occurrences until 11 July, when a large precipitating MSC (19 mm of cumulated rain) washed the aerosol out. On the next day (13 July, flight V036), dust transport within the SAL was reinitiated.

- <sup>5</sup> The vertical profiles of dN<sub>0.3</sub> particle concentration and the Fe/Ca and Si/Al ratios are shown in Fig. 8. This figure illustrates the vertical variability in shape and magnitude in the particle concentration profiles. The same variability appears when examining the vertical profiles of Fe/Ca and Si/Al ratios, in particular regarding flight V032. The dust layer for this flight extended between 1.5 and 4.8 km, with a three sub-layer structure
- <sup>10</sup> between 1.5 and 2 km, between 2 and 3 km, and then between 3 and 4.8 km. The Fe/Ca and Si/Al ratios for the first two sub-layers were around 5.5 and 1.8, respectively. Above 3 km, the Fe/Ca and Si/Al ratios were around 2.0 and 2.7. Differences are found also in the volume size distribution as it will be shown in Sect. 4.4. All together with the fact that the lowest two sub-layers were moister than the upper one, this suggests
- that the dust layer is not homogeneous in origin but is in reality a mixture between Saharan dust transported by the easterly winds within the SAL and Sahelian dust that has been emitted by erosion by the MCS and uplifted from the surface by convective updraft. Note in fact that the Fe/Ca and Si/Al ratios in the monsoon layer below 1 km are of the same order of magnitude than within the lowest Sahelian dust sub-layer
- which proves that the Saharan-Sahelian mixing occurred upstream and is not a local phenomenon. The fact that the number concentration is different below from above the monsoon top will be discussed in Sect. 4.4. On the following days, the shape of the vertical profile of the dN<sub>0.3</sub> particle concentration changed. However, the Fe/Ca and Si/Al ratios remained unchanged to values around 4 and around 2, respectively. Unfortunately no iron oxide measurement was possible in these layers.

### 4.3.6 Saharan atmospheric boundary layer (SABL)

Profile collected during flights V019 (07 June) corresponded to the northern excursion of the ATR-42 relative to the ITD position. Flight V019 was entirely dedicated to the





investigation of the vertical and horizontal structure of the ITD. With this aim, the aircraft was flown north of the ITD (located at  $15.25^{\circ}$  N at the Niamey longitude) up to  $16.5^{\circ}$  N in Mali, where a vertical profile was performed.

The resulting vertical profile (Fig. 5a) is strikingly different from that observed during the previous cases. The surface dust layer extended up to about 2.5 km, above the top of the mixed layer  $Z_i$  found at 1.9 km. The concentrations decreased by a factor of 3 above 2.5 km, but still the dust layer was present up to at least 4.5 km. These dust layers corresponded to moderately-to-high easterly winds and very low water vapour mixing ratio. The lower layer was unstable and slightly moister that the upper one, which was also characterised by more elevated wind speed. These observations agree with the description of dry convection conditions north of the ITD described by Cuesta et al. (2008).

From the chemical point of view, flight V019 was characterised by the highest concentration values of elemental AI. For logistic reasons, unfortunately filter samples were collected only at the very bottom (below 1 km a.s.l.) or at the very top (~3.8 km) of the profile and are not very informative of the vertical variability of the chemical composition. These are shown in Fig. 9.

Samples collected at the very bottom showed ratios Si/Al~2 and Fe/Ca~4. This is consistent with the rather local South Saharan origin of the dust, as confirmed by satellite images (MSG/SEVIRI, not shown). In this layer, the ratio between Fe in the form of oxide and total Fe was 0.64 at the lowest altitude and 0.58 at 3.8 km. Depending on the way that the total dust mass is estimated, the Fe content varies between 3.3 and 11%, the lowest value obtained when the dust mass is estimated from elemental Al and the upper value is obtained when the dust mass is estimated as the sum of the sate concentrations.

It is interesting to notice that a similar vertical structure was observed on the end of flight-profile which was performed on 4 July during flight V031. This flight was ended by a vertical sounding of opportunity at  $14^{\circ}$  N, well south of the ITD (located at around  $17^{\circ}$  N), as the entire flight was rather dedicated to the exploration of the north-south





gradient of the surface monsoon layer over Benin. The examination of the thermodynamic profiles and satellite indicates that the entire dust layer is included in the monsoon layer (and  $Z_s = Z_i$ ) which extended up to 2.5 km (westerly winds, elevated relative humidity).

<sup>5</sup> This vertical profile can be compared to the one observed the day before at the same location (3 July, flight V029 in Fig. 5b), when the ITD was located at around 15° N. In this case the top height of the monsoon layer  $Z_s$  was found at 1.6 km, and dust was only observed in the SAL above 2 km.

## 4.4 Particle size distribution

<sup>10</sup> The normalised volume distributions  $dV/d\log D$  corresponding to vertical profiles discussed in the previous sections are shown in Fig. 10.

All volume size distributions are characterised by a tri-modal structure: a first mode of particles of diameter smaller than 0.6  $\mu$ m, not very well defined because of the lower size cut of the GRIMM OPC; a second mode of particles of diameter comprised between 0.6 and 2  $\mu$ m; and a third, dominant, above 2  $\mu$ m. Whereas the position of the

second mode does not vary from a sample to the other, the modal diameter of the third mode does.

The modal geometric diameter of the coarse mode is found between 4.2 and 5.8 μm, in very good agreement with one of the modal diameter indicated by Rajot et al. (2008)
 for transported dust. No systematic trend towards larger modal diameters is found depending on time after transport and origin. Their diversity (background, erosion, transport) makes it difficult to conclude on its significance also due to the large uncertainty that affects the particle size distribution at these large diameters. Note that Rajot et al. (2008) have indicated in 8.7 μm the modal geometric diameter of the particle volume distribution of locally-produced dust. It is likely that those differences are due to

<sup>25</sup> ume distribution of locally-produced dust. It is likely that those differences are due to differential uplift above the surface level.

There is nonetheless some dependence on altitude. Figure 11 shows the volume size distribution for the three case studies of transport from Bodélé (flight V021), dust in the SABL (V019) and Sahelian dust uplift and transport (V032).





The progressive depletion with altitude of coarse particles larger than 2 µm is evident for flights V019 and V021. The particle size distribution for flight V032 confirms that the layer is a mixture of dust of different origin, and therefore residence time in the atmosphere, as it had been concluded based on the vertical variability of the chemical composition. As a matter of fact, the upper layer, which was also characterised by Si/Al and Fe/Ca ratios characteristics of Saharan dust, has a significantly different volume distribution with a less important coarse fraction that the layers below. This is consistent with the hypothesis of a more distant origin and therefore a higher transport time to the sampling point. The fact that the finest particles are more numerous in the highest SAL layer (2.5 km) also reinforces the transport hypothesis and explains the discrepancy previously noticed between the number concentration and the chemical composition (Fig. 8).

#### 4.5 Optical properties

The spectral single scattering albedo at seven wavelengths between 370 and 950 nm is shown in Fig. 12 for some of the case studies discussed above. Data correspond to straight and levelled runs which were performed within the dust layers of particular interest.

Values are in the range 0.82–0.93 at 370 nm, and increase with wavelengths up to 0.91–0.99 at 950 nm. The spectral dependence is due to both the significant presence

of a coarse particle mode and of that of light-absorbing minerals such as iron oxide in the dust composition. This range of values is consistent with that reported by Redmond et al. (2010) who performed a synthesis of measurements from recent campaigns. It is worth noting that values in the mid-visible are lower than those reported by Osborne et al., 2008 for the single scattering albedo corresponding to accumulation-mode particle
 only for dust collected in the same region. This is consistent with the underestimation of particle absorption (in the absolute, but also in relative terms with comparison to

scattering) because of the exclusion of the coarse mode.





Differences in the relative magnitude of values in Fig. 12 are significant even after taking into account the experimental errors on the scattering and absorption measurements and therefore results can be discussed with respect to the expected origin of mineral dust. The highest values, independently of the wavelengths, are observed for

- <sup>5</sup> dust within the SABL. The lowest values are observed for Saharan dust transported within the SAL during flight V036, as well as for flight V034. In between are values for flights V032 and V033, which are practically indistinguishable. This is rather surprising knowing that Sahelian mineral dust is supposed to be more absorbing due to the higher iron oxide content of Sahelian soils (Claquin et al., 1999).
- To understand the relative contribution of changes in particle size and composition, we have performed Mie-calculation of the measured scattering and absorption coefficients based on the measured number size distributions, and used these calculations to estimate the spectral single scattering albedo. Particles smaller than 0.3 μm diameter which are not measured by the GRIMM OPC are accounted by using the lognormal fitting as reported in Sect. 4.4.

The complex refractive index has been left as free parameter and fixed when the difference between the measurement and the calculation was within 5%. Because the sensitivity of the calculations to the real part of the refractive index is very minor, this was fixed to 1.53 according to the recommendation by Osborne et al. (2008), Müller et al. (2009) and Schladitz et al. (2009). An error of 5% in the real part results in an error of 3–7% in the spectral range from 300 to 800 nm.

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The imaginary part has been varied between 0.002 and 0.008 at 370 nm and between 0.001 and 0.004 at 950 nm (Table 4).

These values are consistent with most of the available in situ measurements for <sup>25</sup> African dust (Patterson et al., 1977; Osborne et al., 2008; McConnell et al., 2008, 2010; Müller et al., 2009; Kandler et al., 2009; Redmond et al., 2010; Klaver et al., 2011).

The highest values of the imaginary part are found for the same flights for which the single scattering albedo is lower (V034 and V036), confirming that the composition of





the aerosol encountered during those flights should be significantly different from that encountered in the other cases. It is difficult to imagine how this could be without invoking some mixing or processing of dust particles. The enhanced absorption could be explained by supposing that particles smaller than 0.5  $\mu m$  are an external mixture

- <sup>5</sup> where soot would represent up to 20% of the particle number (spectrally neutral refractive index of 1.77–0.44*i*), the rest being constituted by dust. These proportions are supported by the visual examination of the samples by electron microscopy. The existence of small absorbing particles mixed with dust is evident also in most of the in situ observations which had been conducted in the Western Africa atmosphere both in
- <sup>10</sup> the dry and the summertime seasons (Chou et al., 2008; Rajot et al., 2008; Müller et al., 2009; Schladitz et al., 2009; Kandler et al., 2009), and in particular with those by Matsuki et al. (2010) in the Niamey-Banizoumbou area later in the wet season. The optical effect of these absorbing particles is surely more evident for flights when the total particle concentration is lower than for the other cases. Because the contribution is rather uncertain, no corrections were performed.

Finally, the measured number size distribution and calculated refractive index obtained by optical closure are used to estimate the asymmetry factor at the seven wavelengths (Table 5).

Values range between 0.75 and 0.78 at 370 nm and between 0.68 and 0.71 at 950 nm. Values at 520 nm are in the 0.73–0.75, in agreement with previous investigations (Haywood et al., 2003a; Osborne et al., 2008; McConnell et al., 2008). There is no straightforward dependence on the origin of dust.

#### 5 Discussion and conclusions

In this paper we have presented data on concentrations, composition, size distribution

and optical properties of mineral dust of Saharan and Sahelian origin observed during the first two summertime intensive field phases of the AMMA program in June–July 2006.





Flights during these intensive research periods illustrated the variability of the dust distribution and vertical profile depending on meteorological conditions and residence time after emission. Dust was ubiquitous north of 12° N. South of 12° N, intense aerosol plumes were observed over the coastal city of Cotonou and at approximately 2° N in correspondence of the advection of a biomass burning plume from the Southern Hemisphere.

Surface transport is the common transport pathway in wintertime (Chiapello et al., 1995; Rajot et al., 2008; Heese and Wiegner, 2008; Osborne et al., 2008) but in summertime it was only observed during a few hours after a local erosion episode. Trans-

- port in the free troposphere was observed in single or multiple sub-layers, depending on time after emission. Generally, multi-plume layering was observed following turbulent uplifting. The lowermost extent of the dust plumes above the boundary layer varies between 1.5 and 2 km, whereas the uppermost extent was found to be remarkably constant at 5 km. The vertical extent, but also the structural and temporal variability of the
- <sup>15</sup> dust vertical profiles is in agreement with summertime observations by various authors (Prospero and Carlson, 1972; Carlson and Prospero, 1972; Karyampudi et al., 1999; Formenti et al., 2003; Reid et al., 2002; McConnell et al., 2008; Tesche et al., 2009; Léon et al., 2009; Cavalieri et al., 2010; Ben-Ami et al., 2009). In particular, our data show the day-to-day variability of the dust vertical profile, in good agreement, albeit
  <sup>20</sup> with lower statistical representation, with the long-time series of lidar data presented by Léon et al. (2009) and Cavalieri et al. (2010). Sub-daily scale variability related to the development of thermal turbulence or the perturbation by convective systems is also likely to be relevant but it is beyond the scope of this paper.

Dust layers of differing origin were identified on the basis of the meteorological record of occurrence of MSCs and by the examination of the simple inter-elemental ratios between major dust tracers, i.e., Al, Si, Ca and Fe.

Various types of situations were encountered: (1) enhanced dust concentrations in the boundary layer below 1-1.5 km corresponding to local erosion; (2) enhanced dust concentrations in the boundary layer this time extending up to 3 km and corresponding





to the dust laden Saharan atmosphere north of the ITD; (3) enhanced single-layer dust concentrations in the free troposphere between 1.5 and 5 km corresponding to the dust transport in the Saharan Air layer (SAL); and (4) enhanced single and multi-layer dust concentrations in the free troposphere between 1.5 and 5 km corresponding to the dust

- <sup>5</sup> transport from the Bodélé depression or to the turbulent uplift of Sahelian dust. Surface local erosion and Bodélé dust were characterised by elevated Si/AI ratios (Si/AI of the order of 3–4) due, in the first case, to the injection of large quartz particles, and in the second case, to the presence of diatomite debris. Conversely, the Si/AI ratio for Sahelian dust in the free troposphere is lower than 2, suggesting that large quartz
- <sup>10</sup> grain are not significantly uplifted after emission. Both at the surface and in the free troposphere, Sahelian dust is also characterised by mean Fe/Ca ratios higher than 4. This is due both to the enhancement of the iron content and to the depletion of the calcium content, linked to the Sahelian soil mineral mineralogy (Claquin et al., 1999; Kandler et al., 2007; Formenti et al., 2008). The quantification of the oxide fraction
- in iron was possible on a few samples only collected during the erosion event of flight V028 and on the SABL layer north of the ITD and revealed values comparable with very recent observations that our research group has performed for Sahelian and Saharan dust (Lafon et al., 2004; Formenti et al., 2008; Klaver et al., 2011). We have indicated a certain spread in the estimate of the percent fraction of iron oxides in our samples
- <sup>20</sup> owing to the lack of a direct measurement of the total dust mass. Nonetheless, the percent fraction of iron oxide in dust is of the order of 3% when estimating the total dust mass as proportional to elemental AI. Klaver et al. (2011) has recently demonstrated that this assumption allows approximating very well the gravimetric mass of Saharan and Sahelian aerosols.
- The vertical profile of the elemental composition illustrates the positive answer to one of the major questions that had motivated the research activities on mineral dust within AMMA, i.e., whether dust emitted by wind-erosion by MCS over Sahelian soils could have been uplifted to the SAL where the AEJ would favour long-range transport towards the Atlantic Ocean. The export of Sahelian dust could have implication for the





estimation of the direct radiative effect at the global scale as well as for ocean and soil fertilisation (Balkanski et al., 2007; Jickells et al., 2005). Our data indicate that only one episode (6–10 July) of Sahelian dust uplift and mixing within the SAL occurred during the AMMA SOP1 and SOP2\_a1 field phases, suggesting that the summertime

- <sup>5</sup> export of Sahelian dust might be minor with respect to the outflow of Saharan dust, which is also more persistent through the year. Not forgetting that this frequency might be under-sampled due to difficulties in flying soon after the occurrence of the MCSs, our results are in agreement with the conclusions from Desboeufs et al. (2010) based on the chemical characterisation of rain events during the monsoon period. The role
- <sup>10</sup> of the precipitating event surely also conditions the Sahelian dust transport in the SAL in scavenging dust uplift. The occasional frequency of Sahelian transport at the continental scale is also suggested by individual particle analysis of Reid et al. (2003) on samples of transported dust collected in the Caribbean.

The analysis of the volume size distribution of dust of different origin does not seem to depend on the source region but rather on the residence time in the atmosphere. This is in accordance with the surface observations by Rajot et al. (2008) in the wintertime. On the contrary, Reid et al. (2008) suggested that dust of different origin has initially different size distributions, which become indistinguishable after two days of transport in the atmosphere. The ensemble of these observations is not necessarily in conflict.

- We, as Rajot et al. (2008), mostly sampled dust freshly emitted or within 1–2 days of transport. The similarity of the size distribution of local emitted dust at the surface during 6 June (V018) to that of Sahelian dust at about 2 km on 6 July (V032) might not being significant as the area where emission had happened is not the same but nonetheless close (within the Sahel). In this sense, these observations are not able
- to show a possible dependence on source region. More measurements of the size distribution of the vertical flux at emission, like those performed in Banizoumbou by Sow et al. (2009) are needed to investigate this point, in comparison to the modelling results of Alfaro and Gomes (2001). The evolution of the size distribution during the episode of transport of Sahelian dust tends to confirm the conclusion of Reid et al. (2008)





concerning the relative invariance with time after 1-2 days of transport.

We also present the first measurements of the single scattering albedo of Sahelian dust. Our estimates are in good agreement with previous in situ observations for African dust when the particle coarse mode is taken into account (summarized by Red-

- <sup>5</sup> mond et al., 2010). Beside the cases of dust transport in the SAL, whose absorption properties might be biased towards low values due to mixing with pollution aerosol, it is demonstrated that at wavelengths smaller than 500 nm Sahelian dust has lower single scattering albedo values than Saharan dust found north of the ITD, although the iron oxide content is rather the same. Klaver et al. (2011) have demonstrated that the
- single scattering albedo cannot be parameterised linearly by the iron oxide content but the full mineralogical composition is necessary to estimate the spectral variability of the complex refractive index, and by consequence of the single scattering albedo of mineral dust.

The significance of the differences in the estimated optical properties (single scattering albedo and asymmetry factor) needs to be evaluated with respect to the implication for the dust radiative effect. In our data, the measured single scattering albedo at 500 nm is 0.94 and 0.96 for Sahelian and Saharan dust, respectively. A simple calculation indicates that over bare surfaces with surface albedo larger than 0.5, both Sahelian and Saharan dust will warm the atmospheric column (positive forcing at the top of the

- atmosphere). The warming effect will be moderate but twice as high for Sahelian than for Saharan dust. This case is relevant also in the case when mineral dust is transported over clouds, as it is frequent, especially as soon as the dust plumes leave the African continent. For lower surface albedos, both Sahelian and Saharan dust will cool the atmospheric column. In this case, the cooling effect will be higher for Saharan dust
- <sup>25</sup> but will tend to equality over dark surfaces such as the ocean. Despite the relative infrequency of the long-range transport of Sahelian dust, the implication of such differences on the redistribution of energy within the atmospheric column, on the convective stability and on the cloud cover of mineral dust needs to be further evaluated. Mesoscale modelling should be pursued to address these issues.





Acknowledgements. Based on a French initiative, AMMA was built by an international scientific group and is currently funded by a large number of agencies, especially from France, the UK, the United States, and Africa. It has been the beneficiary of a major financial contribution from the European Community's Sixth Framework Research Programme. Detailed information on

scientific coordination and funding is available on the AMMA International Web site at http: //www.amma-international.org.

We thank Didier Tanré (Laboratoire Optique Atmosphérique, Lille) and Bernadette Chatenet (LISA) for theirs effort in establishing and maintaining the Banizoumbou, Cinzana, and Djougou sites.

The authors wish to thank the personnel of the Division Technique of the CNRS/INSU (DT-INSU) for supporting the technical development of the AVIRAD sampling system. F. Baudin, former director of this unit, is particularly acknowledged. The instrument scientists, pilots and ground crew of SAFIRE are thanked for facilitating the instrument integration and conducting flying operations. A. Gaudichet is thanked for support to the experimental field phases.



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The publication of this article is financed by CNRS-INSU.

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**Table 1.** Summary of information (date, location, aim) of the ATR-42 flights performed the AMMA SOP1 and SOP2\_a1 field phases.

Flight ID	Date	Location	Objective	IOP
V017	5 Jun 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V018	6 Jun 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V019	7 Jun 2006	North of Niger towards Algeria	Transition monsoon/Harmattan	1.1
V020	8 Jun 2006	Local over Banizoumbou	Dust vertical structure	1.2
V021	11 Jun 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V022	12 Jun 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V023	13 Jun 2006	Niamey-Cotonou	Regional variability of the monsoon flux	1.3
V024	14 Jun 2006	Over sea, Guinea gulf	Regional variability of the monsoon flux	1.3
V025	14 Jun 2006	Cotonou-Niamey	Regional variability of the monsoon flux	1.3
V026	15 Jun 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V027	1 Jul 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V028	2 Jul 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1
V029	3 Jul 2006	Niamey-Cotonou	Regional variability of the monsoon flux	1.3
V030	4 Jul 2006	Over sea, Guinea gulf	Regional variability of the monsoon flux	1.3
V031	4 Jul 2006	Cotonou-Niamey	Regional variability of the monsoon flux	1.3
V032	6 Jul 2006	North of Niger towards Algeria	Transition monsoon/Harmattan	1.1
V033	7 Jul 2006	Local over Banizoumbou	Dust vertical structure	1.2/2.0
V034	10 Jul 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1/2.0
V035	12 Jul 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1/2.0
V036	13 Jul 2006	Local over Banizoumbou	Atmospheric vertical structure	1.1/2.0

**Table 2.** Range of variability of crustal tracers (Mg, Al, Si, P, K, Ca, Fe, and Ti) measured for samples collected north of  $12^{\circ}$  N. The statistics is restricted to samples for which concentrations are above the average minimum detection limit (MDL). For sake of comparison with previous campaigns in Western Africa, the descriptive statistics of the total dust mass estimated from the Al concentrations, as the sum of elemental concentrations and as the sum of elemental concentrations have been performed only on samples for which elemental Ti was above the MDL. Concentrations are expressed in  $\mu g m^{-3}$ .

	Average minimum detection limit (MDL)	Number of data points higher than MDL	Median	Min	Max
Mg	0.01	100	0.4	0.05	3.3
Al	0.02	100	3.8	0.7	38.4
Si	0.03	100	8.6	2.4	76.0
P	0.01	24	0.2	0.04	0.4
К	0.06	73	0.7	0.3	4.9
Ca	0.03	95	1.0	0.2	7.6
Fe	0.04	97	2.0	0.6	18.2
Ti	0.04	51	0.6	0.2	3.1
Mass calculated from the concentration of elemental Al <sup>a</sup>	-	51	73.7	8.0	468.2
Mass calculated as sum of elemental concentrations	-	51	22.4	5.7	147.0
Mass calculated as sum of elemental concentrations (expressed as oxides)	-	51	41.6	10.1	282.3

<sup>a</sup> Estimated assuming a mean conversion factor of 8.2%.

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Table 3.	Summary of information	(date, location	i, elevation) of	the vertical pro	files performed
by the AT	R-42 during the research	n flights of the A	AMMA SOP1 a	nd SOP2_a1 fie	ld phases. The
height of	the top of the boundary la	ayer $(Z_i)$ and the	ne height shear	r level ( $Z_{\rm s}$ ) is als	o indicated.

Date	Flight ID	Location	Start time (UTC)	End time (UTC)	Min altitude (km a.s.l.)	Max altitude (km a.s.l.)	<i>Z<sub>i</sub></i> (km a.s.l.)	Z <sub>s</sub> (km a.s.l.)
5 Jun	V017	Banizoumbou	12:37	13:01	0.3	4.3	0.86	0.82
6 Jun	V018	Banizoumbou	11:18	11:42	0.3	4.3	0.92	0.88
7 Jun	V019	16.5° N, 2.5° E	13:40	14:07	0.3	5.0	0.82	Not defined
8 Jun	V020	Banizoumbou	11:18	11:42	0.3	5.3	0.87	0.86
11 Jun	V021	Banizoumbou	11:48	12:09	0.3	5.3	0.86	0.86
12 Jun	V022	Banizoumbou	11:31	12:10	0.3	5.6	0.86	0.89
14 Jun	V024	2.1° N, 2.5° E	8:32	8:48	0.3	4.3	0.95	0.90
		4.2° N, 2.5° E	9:48	10:01	0.3	4.3	0.97	0.90
14 Jun	V025	Banizoumbou	15:12	15:30	0.7	5.3	0.89	0.84
15 Jun	V026	Banizoumbou	10:19	11:03	0.3	6.0	0.89	0.84
01 Jul	V027	Banizoumbou	11:38	12:04	0.3	4.7	0.89	0.89
02 Jul	V028	Banizoumbou	11:52	12:27	0.3	6.0	0.93	0.93
03 Jul	V029	Niamey to 14° N, 2.2° E	11:34	11:50	0.3	2.8	0.84	0.83
04 Jul	V030	2° N, 2°5E	9:39	9:57	0.3	3.0	0.96	0.87
04 Jul	V031	14° N, 2.2° E	16:50	17:20	0.3	5.3	0.76	0.76
06 Jul	V032	15.2° N, 2.3° E	12:27	12:35	0.3	5.2	0.88	0.80
07 Jul	V033	13.5° N, 2.5° E	13:51	14:21	0.3	5.2	0.85	0.75
10 Jul	V034	4.5° N, 3° E	10:45	11:16	0.3	5.0	0.88	0.82
12 Jul	V035	Banizoumbou	13:02	13:35	0.3	5.6	0.89	0.81
13 Jul	V036	Banizoumbou	12:15	12:41	0.3	5.0	0.85	0.80





**Table 4.** Spectrally-resolved imaginary part of the complex refractive index obtained as a closure parameter between the measured single scattering albedo and the Mie calculation using measured number size distributions for a number of case studies of dust of different origin and residence time.

	370	470	520	590	660	880	950
SABL (V019), 0.7 km a.s.l.	0.002	0.001	0.001	0.001	0.001	0.001	0.001
Saharan transport, Bodélé	0.004	0.003	0.002	0.002	0.002	0.002	0.002
(V021), 0.7 km a.s.l.							
Erosion (V018), 0.7 km a.s.l.	0.003	0.002	0.002	0.001	0.001	0.001	0.001
Sahelian uplift (V032), 1.7 km a.s.l.	0.003	0.002	0.002	0.001	0.001	0.001	0.001
Sahelian uplift (V033), 2.7 mm a.s.l.	0.004	0.002	0.002	0.001	0.001	0.001	0.001
Sahelian uplift (V034), 2.3 km a.s.l.	0.008	0.007	0.005	0.003	0.003	0.002	0.002
Saharan transport (V036), 2.5 km a.s.l.	0.008	0.006	0.005	0.005	0.005	0.004	0.004





**Table 5.** Spectrally-resolved asymmetry factor calculated using the estimated complex refractive indices (real part=1.53; imaginary part=see Table 5) by Mie calculation using measured number size distributions for a number of case studies of dust of different origin and residence time.

	370	470	520	590	660	880	950
SABL (V019), 0.7 km a.s.l.	0.77	0.74	0.75	0.73	0.73	0.71	0.71
Saharan transport, Bodélé	0.76	0.74	0.73	0.72	0.72	0.69	0.69
(V021), 0.7 km a.s.l.							
Erosion (V018), 0.7 km a.s.l.	0.78	0.75	0.75	0.73	0.73	0.70	0.70
Sahelian uplift (V032), 1.7 km a.s.l.	0.77	0.75	0.75	0.73	0.73	0.70	0.70
Sahelian uplift (V033), 2.7 mm a.s.l.	0.77	0.74	0.75	0.73	0.72	0.69	0.70
Sahelian uplift (V034), 2.3 km a.s.l.	0.77	0.75	0.74	0.72	0.71	0.68	0.68
Saharan transport (V036), 2.5 km a.s.l.	0.75	0.74	0.73	0.72	0.71	0.68	0.68





















**Fig. 3.** Average number concentrations for particles of optical diameter larger than  $0.3 \,\mu\text{m}$  (dN<sub>0.3</sub>) and for particle of optical diameter larger than  $1 \,\mu\text{m}$  (dN<sub>1.0</sub>) as a function of the latitude. Data have been averaged in bins of 100 hPa of atmospheric pressure and are expressed in cm<sup>-3</sup> STP (standard temperature and pressure, *p*=1013.25 hPa; *T*=20 °C). The upper panel **(a)** corresponds to the SOP1 sampling period and the lower panel **(b)** to the SOP2\_a1 one. The horizontal black dotted line indicates a threshold of 50 cm<sup>-3</sup> STP for the dN<sub>0.3</sub> particle concentration, used to check the ITD shift in latitude as explained in the text. The black solid line refers to the 0.05 cm<sup>-3</sup> STP for the dN<sub>1.0</sub> particle concentration.







**Fig. 4.** Elemental AI concentration ( $\mu$ g m<sup>-3</sup>) as a function of latitude as determined by filter sampling during the SOP1 and SOP2\_a1 field campaigns. Data have been colour-coded according to the altitude of sampling.







**Fig. 5.** Vertical profiles of  $dN_{0.3}$  particle number and of the spectral scattering coefficient  $\sigma_s$  at 450, 550, and 700 nm (blue, green and red crosses, respectively) observed during the SOP1 (a) and SOP2\_a1 (b) field phases. Data are expressed in cm<sup>-3</sup> and Mm<sup>-1</sup> STP (standard temperature and pressure, p=1013.25 hPa; T=20 °C). The heights of the top of the boundary layer ( $Z_i$ , solid line) and of the shear level ( $Z_s$ , dotted line) are also indicated.







**Fig. 6.** Vertical profile of the dN<sub>0.3</sub> particle number concentrations at STP conditions (left panel), Fe/Ca elemental ratio (central panel) and Si/Al elemental ratio (right panel) for the local dust erosion episodes observed over Banizoumbou on 6 June (flight V018, upper row) and 2 July (flight V028, lower row). As term of comparison, the red lines and points represent the the dN<sub>0.3</sub> particle number concentrations at STP conditions, and the Fe/Ca and SI/Al elemental ratios for flights V017 and V027 conducted the day before each of those flights.







**Fig. 7.** Vertical profile of the  $dN_{0.3}$  particle number concentrations at STP conditions (left panel), Fe/Ca elemental ratio (central panel) and Si/Al elemental ratio (right panel) for the Saharan dust episode observed in the proximity Banizoumbou in the period 9–15 June. **(a)** shows the vertical profiles for 8 June (flight V020, upper row) and 11 June (flight V021, lower row), whereas **(b)** shows the vertical profiles for 12 June (flight V022, lower row), and 14–15 June (flights V025, in red, and V026, in black).









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**Fig. 8.** Vertical profile of the  $dN_{0.3}$  particle number concentrations at STP conditions (left panel), Fe/Ca elemental ratio (central panel) and Si/Al elemental ratio (right panel) for the local dust erosion episodes observed north of Banizoumbou on 6 July (flight V032, upper row), 7 July (flight V033, mid-row) and 10 July (flight V034, lower row) during a Sahelian uplift and transport event.



**Fig. 9.** Vertical profile of the  $dN_{0.3}$  particle number concentrations at STP conditions (left panel), Fe/Ca elemental ratio (central panel) and Si/Al elemental ratio (right panel) for the local dust erosion episodes observed north of the ITD 7 June (flight V019). As term of comparison, the red line represent the  $dN_{0.3}$  particle number concentrations at STP conditions for flight V031 conducted on 4 July north of Banizoumbou but south of the ITD. The elemental ratios are not available as only a vertical sounding of opportunity was conducted in the Niamey/Banizoumbou area whereas the remaining flight was flown south of it.















**Fig. 11.** Normalised mean volume size distributions as a function of altitude for the episode of Saharan transport from Bodélé (9 June, flight V021, left panel), dust in the SABL (7 June, flight V019, mid-panel) and the episode of Sahelian dust uplift (6 July, flight V032, right panel).







