Atmos. Chem. Phys. Discuss., 10, 26045–26071, 2010 www.atmos-chem-phys-discuss.net/10/26045/2010/ doi:10.5194/acpd-10-26045-2010 © Author(s) 2010. CC Attribution 3.0 License.



This discussion paper is/has been under review for the journal Atmospheric Chemistry and Physics (ACP). Please refer to the corresponding final paper in ACP if available.

Retrievals of chlorine chemistry kinetic parameters from Antarctic CIO microwave radiometer measurements

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Received: 10 September 2010 – Accepted: 1 November 2010 – Published: 4 November 2010 Correspondence to: S. Kremser (s.kremser@niwa.co.nz)

Published by Copernicus Publications on behalf of the European Geosciences Union.

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Abstract

Key kinetics parameters governing the partitioning of chlorine species in the Antarctic polar stratosphere were retrieved from 28 days of chlorine monoxide (CIO) microwave radiometer measurements made during the late-winter/early spring of 2005 at Scott Base (77.85° S, 166.75° E). During day-time the ratio of the photolysis rate, J, and the 5 dimer formation rate, $k_{\rm f}$, determines the partitioning of CIO and its dimer, chlorine peroxide (CIOOCI). This ratio $J/k_{\rm f}$ remains uncertain and, as a complement to ongoing laboratory studies, the goal of this work is to provide a constraint on that uncertainty based on observations of CIO profiles in the Antarctic. First an optimal estimation technique was used to derive $J/k_{\rm f}$ ratios for a range of $K_{\rm eq}$ values. The optimal estimation forward model was a photochemical box model that takes J, $k_{\rm f}$, and $K_{\rm eq}$ as inputs, together with a priori of activated chlorine ($CIO_x = CIO + 2 \times CIOOCI$), ozone, temperature and pressure profiles. JPL06 kinetics are used as a priori in the optimal estimation and for all other chemistry in the forward model. Using the more recent JPL09 kinetics results in insignificant differences in the retrieved value of $J/k_{\rm f}$. A complementary approach was used to derive the optimal kinetic parameters; the full parameter space of

J, $k_{\rm f}$, $K_{\rm eq}$ and ClO_x was sampled to find the minimum in differences between measured and modelled ClO profiles. Furthermore, values of $K_{\rm eq}$ up to 2.0 times larger than recommended by JPL06 were explored to test the sensitivity of the $J/k_{\rm f}$ ratio to changes ²⁰ in $K_{\rm eq}$. The results show that the retrieved $J/k_{\rm f}$ ratios bracket the range of 1.97 to $1.23 \times (J/k_{\rm f})_{\rm JPL06}$ over the range of $K_{\rm eq}$ values considered. The retrieved $J/k_{\rm f}$ ratios lie in the lower half of the large uncertainty range of $J/k_{\rm f}$ recommended by JPL06 and towards the upper portion of the smaller uncertainty range recommended by JPL09.

1 Introduction

²⁵ The photochemistry of the chlorine monoxide dimer, chlorine peroxide (ClOOCl; hereafter "the dimer") in the polar stratosphere is central to quantitatively describing polar



ozone depletion and hence the Antarctic ozone hole. The formation and photolysis of CIOOCI, known as the CIO-dimer cycle (Molina and Molina, 1987), is typically responsible for 55% to 70% of the spring-time ozone depletion in the Antarctic stratosphere (Frieler et al., 2006; SPARC, 2009). When high chlorine monoxide (CIO) concentrations are present in the polar stratosphere the CIO-dimer cycle:

R1)
R2)
R3)
R4)

10 Net: $2O_3 + hv \rightarrow 3O_2$

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becomes an extremely efficient ozone loss process. The terms k_f and k_r refer to the reaction rate constants for the formation and the thermal dissociation of the dimer, respectively. M represents any molecule to remove excess energy and *J* is the photolysis rate of the dimer, which is directly related to the dimer absorption cross-section. Re-

actions (R2) and (R3) are the most important for ozone destruction since they release chlorine (CI) atoms that then react with ozone (Reaction R4). Reaction (R2) is the rate limiting step in the cycle.

If CIOOCI decomposes thermally (R1) rather than being photolyzed (R2), or if the photolysis reaction produces CIO, a null cycle results that leads to no change in ozone.

²⁰ During day-time, when ClOOCl loss occurs mainly by photolysis, the partitioning between ClO and ClOOCl, and the overall rate of the catalytic cycle, is controlled by the dimer formation rate (k_f) and photolysis rate (*J*). The equilibrium between ClO and ClOOCl is then given by the expression:

 $\frac{[\text{CIO}]^2}{[\text{CIOOCI}]} = J/k_{\text{f}}$



(1)

During night-time the temperature dependent thermal equilibrium constant (K_{eq}) governs the partitioning between CIO and CIOOCI:

$$K_{\rm eq} = \frac{k_{\rm f}}{k_{\rm r}} = \frac{[{\rm CIOOCI}]}{[{\rm CIO}]^2}$$

Uncertainties in these kinetics parameters (*J*, k_{f} and K_{eq}) contribute to differences ⁵ between measured and modelled polar stratospheric CIO (e.g., Stimpfle et al., 2004), affect our ability to accurately describe polar ozone destruction (e.g., Frieler et al., 2006; von Hobe et al., 2007), and our ability to confidently project the response of polar ozone to future changes in stratospheric chlorine loading (e.g., SPARC, 2009).

Models of polar ozone photochemistry have traditionally used values of the CIOOCI absorption cross-section and k_f recommended either by the JPL data panel (e.g., Sander et al., 2003, 2006, 2009) or the IUPAC Subcommittee on Gas Kinetic Data Evaluation panel (Atkinson et al., 2007). Values of the CIOOCI cross section recommended by the various panels typically fell between laboratory measurements by Huder and DeMore (1995) (low range in photolytically active region) and Burkholder et al.

- (1990) (high range). In 2007, Pope et al. published CIOOCI absorption cross-sections that were considerably lower than either the Sander et al. (2006) (hereafter: JPL06) or Atkinson et al. (2007) recommendations. This smaller photolysis rate would make it impossible to quantitatively explain observed ozone loss, suggesting that as yet unknown processes were active. More recent laboratory studies by von Hobe et al. (2009), Chen
- et al. (2009), Papanastasiou et al. (2009), and Wilmouth et al. (2009) concluded that the correction applied by Pope et al. to account for Cl_2 contamination in their CIOOCI sample may have been too large. This in turn would lead to cross-sections which are too small in the atmospherically important wavelength region (>300 nm). That said, these studies published in 2009 do not agree on the absorption cross-sections for the
- ²⁵ dimer, and so there remains uncertainty on the photolysis rate for the ClO-dimer. Comparisons between measured and modelled ClO (von Hobe et al., 2007; Schofield et al., 2008) concluded that no combination of $k_{\rm f}$ and $K_{\rm eq}$ were compatible with the absorption cross-sections measured by Pope et al. (2007).



(2)

Furthermore, field measurements of CIO and CIOOCI by von Hobe et al. (2007) and Stimpfle et al. (2004) and satellite measurements of CIO by Santee et al. (2010) indicate discrepancies in the equilibrium constant K_{eq} for CIOOCI compared to the JPL06 recommendation, and the more recent JPL09 recommendation. Therefore, there is a need to reduce the uncertainties in these reaction kinetics parameters, and thereby reduce uncertainties in modelled polar ozone loss.

In this study, two methods, both using CIO microwave radiometer measurements from Scott Base (78.85° S, 166.75° E) during the late-winter and early spring of 2005 (described in Sect. 2), are used to constrain the kinetic parameters governing CIO chemistry in the polar stratosphere, viz.:

- 1. An optimal estimation approach (Sect. 3), using a photochemical box model as a forward model, is used to derive optimal $J/k_{\rm f}$ ratios consistent with the measured CIO profiles.
- 2. An exploration of the J, $k_{\rm f}$, $K_{\rm eq}$ and CIO_x (CIO_x=CIO+2×CIOOCI, i.e. active chlorine) parameter space within physically plausible limits to find where the mini-
- rine) parameter space, within physically plausible limits, to find where the minimum of measurement-model differences occur.

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The first approach is significantly faster than the second and provides statistical uncertainties on the derived parameters. The second, in addition to showing which set of parameters minimizes the measurement-model difference, also shows the regions within the 4-D-space where differences of a similar magnitude are obtained. The second method also shows whether there are local minima in the 4-D-space that might provide a set of viable kinetic parameters consistent with the field observations. The results obtained are presented and compared with previously published results in Sect. 4. These results are discussed in Sect. 5, followed by the conclusions of this study.



2 Observations and model data

2.1 Observations

Stony Brook University and National Institute of Water and Atmospheric research (NIWA) have jointly operated a ground-based CIO microwave radiometer at Scott Base,

- Antarctica, since February 1996. The instrument and the CIO measurements are described in more detail in Solomon et al. (2000) and Connor et al. (2007). Briefly, the microwave radiometer measures the thermally-excited emission of CIO at 278.6 GHz within a frequency band of 506 MHz. An ozone line within the CIO microwave emission band interferes with the measurement. To eliminate this interference, and instrumental
- artefacts, the day minus night CIO spectrum is obtained by subtracting the night-time spectrum from the day-time spectrum. At night, Reaction (R1) dominates Reaction (R2), leading to most CIO_x being in the form of CIOOCI and, as a result, depending on the day of the year, night-time CIO is less than ~ 20% of day-time CIO. While the day/night subtraction is essential to retrieve the daily CIO profile, the CIO concentra-
- tions themselves are unaffected by this step in the retrieval. CIO profiles from 10 to 56 km are derived from the CIO day minus night spectra. Retrieved CIO concentrations on 20 altitude levels from 11 to 30 km are used in this analysis since it is over this altitude range that CIO concentrations maximize. 28 days of CIO profiles, made during the period from 21 August to 20 September 2005 (Fig. 1), are used in this study, as 2005
- offered the most comprehensive dataset. Due to bad weather conditions there were no CIO measurements available from 14–16 September 2005 and therefore these days are not considered in this study. At the beginning of the season (late August) the peak CIO mixing ratio occurs at around 22 km. In early September the peak CIO mixing ratio shifts downwards, occurring at 20–21 km. The maximum CIO concentration increases
- from 1.2 part per billion by volume (ppbv) to 2.3 ppbv within the 28 day period. Between 17 and 23 km the temperatures reach their seasonal minimum (<190 K) during this time period, as also shown in Fig. 1.



2.1.1 Model data – SLIMCAT 3-D CTM

Output from the SLIMCAT 3-D off-line chemical transport model was used to provide estimates of the abundance of active chlorine ($CIO_x = CIO + 2 \times CIOOCI$), active bromine (BrO_x=BrO+BrCI), and ozone over Antarctica. The model is described in detail in

- ⁵ Chipperfield (1999, 2006). Briefly, SLIMCAT contains a detailed stratospheric chemistry scheme, including heterogeneous reactions of liquid and solid polar stratospheric clouds (PSCs). The model uses a limited number of small families in the chemistry module. Cl, ClO and ClOOCl form a family and the partitioning between these species is found assuming instantaneous photochemical equilibrium. All other inorganic chlo-
- rine species (e.g. chlorine dioxide (OCIO), hypochlorous acid (HOCI) etc.) are integrated separately. Similarly bromine (Br) and bromine monoxide (BrO) are solved as a family while all other inorganic bromine species are integrated separately (e.g. bromine monochloride (BrCl)).

For this study output was taken from the SLIMCAT 3-D CTM. The run used here (for reference, run 509) has a resolution of $5.6^{\circ} \times 5.6^{\circ}$ with 32 levels from the surface to about 60 km and started on 1 January 1977. The run was forced using European Centre for Medium-Range Weather Forecast (ECMWF) analyses (ERA-Interim after 1989). Daily profile output from the model run was stored for the location of Scott Base. The run included a source of bromine from very short-lived species (VSLS), which contributed about 6 part per trillion by volume (pptv) to total stratospheric inorganic bromine (Br_v) in 2005 (Feng et al., 2007).

3 Retrieval algorithm/optimal estimation

An optimal estimation (OE) approach (Rodgers, 2000) is used to retrieve the kinetic parameters that optimize the agreement between the measured CIO profiles and pro-

files generated by the OE forward model. The quantities to be retrieved, $J/k_{\rm f}$ and ClO_x, from the ClO measurements (Y) are represented, together with $K_{\rm eq}$ and a dataset of



28 a priori CIO_{x} profiles, by the state vector (*X*). The resulting state vector has 563 elements, 3 kinetics parameters plus 28 days × 20 altitudes for a priori CIO_{x} . Rather than working with J/k_{f} and K_{eq} explicitly, they are specified as scalings (multiplicative factors) relative to the JPL06 recommendations. Hereafter, all quantities subscripted ⁵ with scale refer to scaling with respect to JPL06. JPL06 kinetics are the point of reference for this study because these kinetic parameters are most commonly used in recently published studies of polar ozone photochemistry. In some cases, results are also compared with the more recent JPL09 recommendation. The JPL09 K_{eq} value is ~70% of the JPL06 value, while J/k_{f} for JPL09 shows essentially no difference from JPL06.

The relationship between the CIO measurements (Y) and the state vector is described by a forward model (**F**) that calculates CIO as a function of the state vector X and other parameters, b, which include BrO_x , ozone, temperature, and pressure profiles:

15 $Y = \mathbf{F}(X, b) + \varepsilon$

where ε is the measurement error. The performed error analysis for the Antarctic CIO measurements used in this study, are described in depth in Solomon et al. (2000). The forward model (**F**) is a photochemical steady state box model, based on polar ozone photochemistry and described in detail by Salawitch et al. (1993). The Salawitch

- et al. box model is used to calculate CIO at 30 min resolution for the 28 day period considered, for various specifications of the kinetic parameters, which are set to JPL06 values if not otherwise noted. Daily day minus night CIO profiles, hereafter modelled CIO (\hat{Y}), are calculated from the 30 min time resolution output of the box model, to provide compatibility with the measured CIO profiles. For **b**, the daily BrO_x profiles
- ²⁵ were obtained from the SLIMCAT 3-D CTM (see Sect. 2.2). Because vertical ozone profiles from ozonesondes were not available for each day of the analysis period, ozone profiles were extracted from the SLIMCAT simulation. Linearly interpolated differences between measured and modelled ozone profiles were added to the SLIMCAT profiles

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so that they matched the ozonesonde measurements (taken from Hassler et al., 2008) on the days when measurements were available and provided best-estimates of ozone profiles on intervening days. The temperature and pressure profiles were provided from the National Centers for Environmental Prediction (NCEP) (Kalnay et al., 1996).

⁵ The inverse of Eq. (3) produces *X* as an output for a given *Y*. There is an infinite set of such state vectors *X* that could form solutions to Eq. (3), reproducing the prescribed measurements. The retrieval problem is said to be under constrained, as there are more elements in the state vector than there are measurements. Given an a priori state vector (*X*_a), uncertainties on the elements of *X*_a (the covariance matrix **S**_a) and the measurement error, the OE retrieves an optimal state vector (\hat{X}) that minimizes the difference between measured and modelled CIO. *X*_a constrains the OE algorithm. The iterative equation for the solution of the inverse of Eq. (3) is:

$$\hat{X}_{i+1} = X_{a} + (\mathbf{S}_{a}^{-1} + \mathbf{K}_{i}^{T} \mathbf{S}_{e}^{-1} \mathbf{K}_{i})^{-1} \mathbf{K}_{i}^{T} \mathbf{S}_{e}^{-1} \cdot [\mathbf{Y} - F(\hat{X}_{i}) + \mathbf{K}_{i}(\hat{X}_{i} - X_{a})]$$
(4)

where **K** is the weighting function matrix, with each element representing the partial derivative of the modelled CIO with respect to a state vector element; i.e. $\mathbf{K}_{ij} = d\mathbf{F}_i(\mathbf{X})/dX_j$. As such, the weighting functions describe the sensitivity of the modelled CIO to variations in the state vector and were determined by consecutively perturbing each state vector quantity and recalculating CIO using the forward model. \mathbf{S}_e in Eq. (4) is the covariance matrix of the measurement errors (Fig. 1b), constructed by placing the CIO measurement errors along the diagonal, and set to zero everywhere else. X_a was constructed by inserting values for $k_{f scale}$, $K_{eq scale}$ and J_{scale} , all set to 1.0 so that the a priori for the kinetics are the UPLOE recommendations. A set of 28 a priori CIO

the a priori for the kinetics are the JPL06 recommendations. A set of 28 a priori CIO_x profiles, corresponding to each day of measurement, is provided as a required input to the forward model.

The uncertainties on the elements of X_a are expressed along the diagonal of a covariance matrix S_a , which is elsewhere set to zero. Selection of the S_a settings requires some subjective judgement, noting that if excessively small values are chosen, the retrieval will be constrained too tightly. In this case the retrieval will make little use of the



information provided by the measurements and will differ little from the prescribed a priori X_a . On the other hand, if excessively large values for S_a are chosen, this results in an unrealistic retrieval where the measurement noise is interpreted as information. Here the S_a values for $k_{f \text{ sale}}$, J_{scale} and the a priori CIO_x profiles were numerically determined as described in Schofield et al. (2004) and are set to $\sigma_{k_{f \text{ scale}}} = 0.043$ and $\sigma_{J_{\text{scale}}} = 0.173$. The S_a values for the a priori CIO_x were set to 20% of the daily maximum CIO_x value, individually for each day, also as done by Schofield et al. (2004).

5

During day-time the partitioning of CIO and its dimer is governed by J/k_f (see Eq. 1). K_{eq} becomes important primarily for the night-time partitioning of CIO_x. Because the night-time spectrum was subtracted from the day-time spectrum to obtain a daily measured CIO profile, the CIO measurements only give information on J/k_f , not on these parameters individually, and the measurements are only weakly sensitive to K_{eq} . Therefore, with the OE approach we retrieve $J_{scale}/k_{f scale}$ and make no attempt to retrieve K_{eq} . Rather, to test the sensitivity of our results to the choice of $K_{eq scale}$, the OE is done 19 times, varying $K_{eq scale}$ from 0.2 to 2.0× $K_{eq JPL06}$ in steps of 0.1.

OE is done 19 times, varying K_{eq scale} from 0.2 to 2.0×K_{eq JPL06} in steps of 0.1. Complementary to the OE approach, a second method was used to derive the kinetic parameters of the CIO-dimer cycle. In this approach the full parameter space was explored by varying the values of the *J*, k_f and K_{eq scale} factors within physically plausible ranges. Furthermore, to estimate the dependence of the kinetic parameters on changes in the prescribed CIO_x profile, the CIO_x profile was scaled between 0.5 and 1.0 times the a priori CIO_x profile used in the OE approach, in steps of 0.1 (scalings above 1.0 were not used since this resulted in CIO_x occasionally exceeding the total stratospheric inorganic chlorine (CI_y) loading). The photochemical model described above in Sect. 3 was also used here, with the same input variables (i.e. BrO_x,

ozone) to calculate day minus night profiles of CIO (again, to be compatible with the CIO measurements). The kinetic values that give the smallest difference between the modelled and the measured CIO are compared to the results from the OE and previous publications.



4 Results

4.1 Retrieved kinetic parameters from optimal estimation

The measurements were well described by the forward model for each OE run, with root mean square (RMS) differences between measured and modelled CIO ranging from 0.063 to 0.066 ppbv, which is smaller than the maximum error of the CIO mea-5 surements. The results from the retrievals performed as outlined in Sect. 3, expressed as scale factors relative to the kinetic parameters recommended by JPL06, are shown together with 1 σ uncertainties in Fig. 2. In addition to prescribing J and $k_{\rm f}$ from JPL06 as a priori in the forward model, JPL06 kinetics were used for all other chemistry in the forward model. The derived RMS values for the 19 OE runs do not vary much (only by 10 ~5%), but the change in the retrieved $J/k_{\rm f}$ ratios over the prescribed $K_{\rm eq}$ range (0.2 to 2.0) is significant. The ratio retrieved by prescribing $K_{eq scale}$ =2.0 is about a factor of 1.6 smaller than the $J_{\text{scale}}/k_{\text{f scale}}$ ratio retrieved for the smallest $K_{\text{eq scale}}$ =0.2 used in this study. Tests using JPL09 kinetics showed only very slightly ($\sim 2\%$) larger values for the $J_{\text{scale}}/k_{\text{f scale}}$ ratio, which is an insignificant difference given other sources of 15 uncertainty. A $J_{\text{scale}}/k_{\text{f scale}}$ value of 1.0 refers to the JPL06 recommendations, where the photolysis rate of the dimer (J) is calculated using the absorption cross-sections recommended by JPL06. $J_{\text{scale}}/k_{\text{f scale}}$ values relative to JPL06 determined in earlier studies are indicated with vertical bars along the Y-axis in Fig. 2. The salient features of Fig. 2 and their causes are: 20

Negative slope: an increase in K_{eq} can be obtained either by increasing k_f , decreasing k_r , or both, or decreasing k_f and k_r but decreasing k_r relatively more. The k_r values associated with the prescribed $K_{eq \text{ scale}}$ shown in Fig. 2 decrease with increasing $K_{eq \text{ scale}}$, as does $J_{scale}/k_{f \text{ scale}}$. However, the decrease in k_r is small such that $k_{f \text{ scale}}$ needs to increase to cover the large range prescribed for $K_{eq \text{ scale}}$ (0.2–2.0). The combination of increasing $k_{f \text{ scale}}$ and decreasing k_r shifts the equilibrium towards



CIOOCI, reducing the calculated concentration of CIO. To restore the equilibrium such that the measured CIO is obtained, the photolysis of the dimer, *J*, must increase. The observed decrease in $J_{\text{scale}}/k_{\text{f scale}}$ can then only be achieved if $k_{\text{f scale}}$ increases more than J_{scale} . This can be explained by the greater effectiveness of the photolysis rate in forming CIO: i.e. a doubling of *J* is not equivalent to a doubling of k_{f} . Small relative changes in *J* have a significant impact on the CIO/CIOOCI partitioning, whereas larger relative changes in k_{f} are required to achieve an equivalent change in the CIO/CIOOCI partitioning.

- Asymptotic behaviour of $J_{\text{scale}}/k_{\text{fscale}}$ at high $K_{\text{eq scale}}$: the results show that the retrieved values of $J_{\text{scale}}/k_{\text{f scale}}$ do not vary much at high values of $K_{\text{eq scale}}$, 1.6 to 2.0, used in the analysis. An increase in $K_{\text{eq scale}}$ requires $k_{\text{f scale}}$ to increase (see above). However, the measured CIO imposes a limit on the amount of CIOOCI that results from the increase in $k_{\text{f scale}}$. If much of the available CIO is transformed into CIOOCI a further increase in k_{max} affects CIOOCI concentrations less than if
- ¹⁵ into CIOOCI, a further increase in $k_{\rm f\ scale}$ affects CIOOCI concentrations less than if larger abundances of CIO were available. Once the amount of CIOOCI maximizes, the photolysis rate also encounters a threshold within OE, such that $J_{\rm scale}/k_{\rm f\ scale}$ plateaus at high values of $K_{\rm eq\ scale}$. Additional increases in $K_{\rm eq\ scale}$ are then accomplished within the model by changing $k_{\rm r}$. This result suggests a lower limit for $J_{\rm scale}/k_{\rm f\ scale}$ of ~ 1.24±0.12, in combination with high values of $K_{\rm eq\ scale}$, to explain the Antarctic
- $\sim 1.24 \pm 0.12$, in combination with high values of $X_{eq scale}$, to explain the Antarctic measurements of CIO.

The $J_{\text{scale}}/k_{\text{f scale}}$ ratio of 1.39 interpolated to the K_{eq} obtained by Solomon et al. (2000) ($K_{\text{eq scale}}$ =0.86) agrees very well with their $J_{\text{scale}}/k_{\text{f scale}}$ ratio (see Fig. 2). Solomon et al. used measurements from the same CIO microwave radiometer at the same site. The measurements they used were taken from a different period (1996– 2000) when stratospheric temperatures were, on average, 4 K lower than in 2005. The formation rate of the dimer and the equilibrium constant are both known to be temperature dependent. A decrease in temperature leads to an increase in the dimer formation rate k_{f} , which in turn leads to a decrease in the J/k_{f} ratio. However, Solomon et al.



used the JPL97 functional form for the temperature dependence of $k_{\rm f}$, which is different from the JPL06 expression used here. The difference between JPL97 and JPL06 then partially offsets the effect of the temperature difference between the two studies.

The retrieved $J_{\text{scale}}/k_{\text{f scale}}$ ratios from all 19 OE runs lie within the lower half of the large uncertainty range given by JPL06, within the upper half of the smaller uncertainty range given by JPL09, and agree with most of the $J_{\text{scale}}/k_{\text{f scale}}$ ratios determined in previous studies (Fig. 2). The results suggest that to explain the Antarctic measurements of CIO made in 2005, J/k_{f} is unlikely to be smaller than the JPL06/09 recommendations. This result would exclude the lower uncertainty limit on J/k_{f} for both JPL06 and JPL09.

The best fit to the CIO measurements, i.e. the smallest RMS of the difference between modelled and measured CIO, is obtained at $K_{eq \ scale}=2.0$. In this case $J_{scale}/k_{f \ scale}$ is 1.23 ± 0.11 . In contrast to previous studies, our results suggest that to best represent the CIO measurements, $K_{eq \ scale}$ must lie towards the upper end of the prescribed range (0.2 to 2.0), in agreement with the recent laboratory study of Ferracci and Rowley (2010). However, the derived K_{eq} values in Ferracci and Rowley were obtained at higher temperatures than in this study. To explain the CIO measurements, the results presented in this study (Fig. 2) do not preclude higher K_{eq} values than currently recommended. That said, the decrease in RMS from $K_{eq \ scale}=0.2$ to 2.0 is small (5%), suggesting that K_{eq} is not well constrained by the measurements of CIO to make a robust evaluation of K_{eq} (see Sect. 3).

The RMS values obtained from the OE are elevated as a result of the inclusion of two days of measurements. Between 9 and 10 September 2005, CIO dropped from \sim 2.2 ppbv to \sim 0.8 ppbv before increasing back to \sim 2 ppbv two days later (see Fig. 1).

²⁵ The suppressed values of CIO over this two day period are caused by dynamical variability of the polar vortex. While SLIMCAT shows some reduction in CIO_x over this period, the reduction is not commensurate with the observed reduction in CIO apparent in the disagreement between measured and SLIMCAT modelled CIO profiles on these days. With the given resolution of the SLIMCAT run (see Sect. 2.2), the model



can not capture the observed strong gradients at the edge of the vortex. As such, the a priori CIO_x profiles provided by SLIMCAT are too high on those two days. The value of $J_{\text{scale}}/k_{\text{f scale}}$ is determined by the partitioning of CIO and CIOOCI on all days and is not determined separately for every day. The amount of CIO_v, on the other hand, can vary for every day independently within the given uncertainty range, which is determined by 5 the set up of the a priori uncertainty (S_a) . The selection of the magnitude of the elements of S_a determines the freedom of the OE, i.e. how well constrained is the OE. The selected uncertainties on the a priori CIO_x profile (see Sect. 3) are likely to be too small for these two days and therefore the OE is constrained too tightly. The retrieval does not have enough freedom to modulate CIO_x , for the given $J_{scale}/k_{f scale}$, to reproduce the 10 measurements. The combination of $J_{\text{scale}}/k_{\text{f scale}}$ with a higher amount of CIO_x leads to more CIO than observed and thereby contributes to the higher RMS values. While excluding the two days from the analysis reduces the RMS such that it is smaller than the measurement uncertainty at every altitude level, the retrieved $J_{\text{scale}}/k_{\text{f scale}}$ increases slightly, from $1.35 \pm 0.14 \times (J/k_f)_{JPI 06}$ (all data points) to $1.44 \pm 0.18 \times (J/k_f)_{JPI 06}$ (9 and 15 10 September excluded), if $K_{eq} = K_{eq} JPL06$. The difference between the two calculations is not statistically significant. This result shows the sensitivity of the applied OE method to the a priori CIO_x information. Since there is no valid scientific reason for excluding these two days in questions, and because they have only a small effect on the retrieved

value of $J_{\text{scale}}/k_{\text{f scale}}$, these two days were retained in the prior results of this section as well as in the OE analyses discussed further in Sect. 5.

4.2 Exploring the full parameter space

In contrast to the OE approach, the $k_{\rm f\ scale}$, $J_{\rm scale}$, and $K_{\rm eq\ scale}$ parameter space was explored to derive the optimal CIO-dimer cycle kinetic parameters required to explain the measured CIO. Noting, however, that the CIO_x values obtained from SLIMCAT may be uncertain, sensitivity to CIO_x was also explored by scaling CIO_x from 0.5 to 1.0 (as mentioned in Sect. 3, scalings above 1.0 were not considered since this resulted in CIO_x occasionally exceeding Cl_y). $J_{\rm scale}$ was varied between 0.1 and 2.5, while $k_{\rm f\ scale}$



was varied between 0.4 and 1.6. An upper bound of $K_{eq scale} = 1.6$ was used since JPL09 and most other studies (with the notable exceptions of Nickolaisen et al., 1994; Ferracci and Rowley, 2010) suggest $K_{eq scale} \leq 0.7$. Within this space, the kinetic values that give the smallest RMS of the difference between modelled and measured CIO 5 were determined.

The minimum RMS values (from the 6 CIO_x scalings) are shown for each combination of J_{scale} and $k_{\text{f scale}}$ and for $K_{\text{eq scale}}$ =0.4, 0.8, 1.2, and 1.6 in Fig. 3. In all cases the lowest RMS values are found for CIO_{xscale}=1.0 (unshaded regions in Fig. 3) and the absolute minimum RMS for the highest $K_{eq scale}$ (Fig. 3d), in agreement with the results obtained from OE. While lower RMS values might have been found for CIO_{xscale}> 1.0, this would be physically unrealistic for the reasons detailed above. The calculated minimum RMS in the J_{scale} - $k_{\text{f scale}}$ plane, for each $K_{\text{eq scale}}$, shows little dependence on $K_{eq scale}$, confirming again that our measurements do not provide a strong constraint on K_{eq} .

The RMS values define a "valley" of minima in each $J_{\text{scale}} - k_{\text{f scale}}$ plane. When 15 $K_{eq scale}$ is fixed and if $k_{f scale}$ increases, k_r must also increase (see Eq. 2). However, the resultant increase in k_r is not sufficient to produce enough CIO to match the measurements and hence J_{scale} too must increase. This explains the orientation of the valley of minimum RMS values seen in Fig. 3. The angle that this valley makes with the X-axis decreases slightly with increasing $K_{eq scale}$, indicating a decrease in $J_{scale}/k_{f scale}$ with 20 increasing $K_{eq scale}$, in agreement with the results from the 19 OE runs described in Sect. 4.1. The CIO measurements only give information on $J/k_{\rm f}$, not on these parameters individually (see Sect. 3). Therefore, the $J/k_{\rm f}$ ratio that best explains the CIO measurements can be derived from the orientation of the valley with an uncertainty constrained by the width of the valley. For higher $K_{eq scale}$ values (panels b–d in Fig. 3), 25 the orientation of the valley suggests a $J_{\text{scale}}/k_{\text{f scale}}$ ratio smaller than 1, which would disagree with the OE result. However, taking the absolute minimum RMS in each panel



is in good agreement with the retrieved ratios of 1.62 ± 0.24 , 1.40 ± 0.16 , 1.32 ± 0.13 and 1.27 ± 0.28 from OE, for $K_{\text{eq scale}}$ prescribed at 0.4, 0.8, 1.2, and 1.6, respectively. However, the resulting valley of RMS minima using $J_{\text{scale}}/k_{\text{f scale}}$ ratios of 1.66 and 1.25 would require $\text{ClO}_{\text{xscale}}$ values smaller than 1.

The RMS values obtained from exploring the whole parameter space are more than twice as large as those resulting from OE. Therefore, the results from the OE approach lead to a better fit to the CIO measurements than the results from exploring the whole parameter space. Contrary to the OE approach, the sampling of the whole parameter space does not readily permit changing the CIO_x value from day to day or from one altitude level to another. The parameter space exploration scales CIO_x values on all days and at all levels equally. This most likely explains the higher RMS values for the parameter space exploration.

5 Discussion and conclusion

The results presented above show that OE is a reliable method for investigating the kinetics of the CIO-dimer cycle. The results derived by exploring the whole parameter space agree with the retrieved kinetics from the OE runs. OE has the advantage that it is much faster than sampling the parameter space and provides quantitative estimates of the uncertainties on the derived parameters. Furthermore, the OE retrieves a quantitative estimate of the daily CIO_x abundances required to explain the observed CIO concentrations. Exploring the whole 4-D space is computationally expensive and changing CIO_x individually for every day would make it even more so. OE gives one set of parameters and, as shown above, these results provide a reliable fit to the measurements.

The results presented above indicate that the retrieved $J_{\text{scale}}/k_{\text{f scale}}$ ratios from all 19 OE runs agree with most values reported in previous studies and lie within the lower half of the large uncertainty range reported by JPL06. The recently updated JPL09 recommendation reports a much smaller uncertainty range on the CIOOCI absorption



cross-sections than JPL06. As a result, the uncertainty range on the CIOOCI photolysis rate (*J*) is reduced, which in turn places tighter constraints on the J/k_f ratio than given by JPL06. The retrieved $J_{\text{scale}}/k_{\text{f scale}}$ for prescribed K_{eq} values of 0.4 to 1.2 times JPL06, which corresponds to 0.6 to 1.7 times JPL09, lies within the upper part of the uncertainty range given for J/k_f by JPL09.

The 2005 ground-based microwave CIO measurements made over Scott Base, Antarctica, can be explained using the JPL06 recommendations of J, k_f , and K_{eq} with the given range of uncertainties (Fig. 2). Our study is therefore consistent with the JPL06 recommendation for the kinetic parameters that control the partitioning of CIO_x, but at the same time, the results suggest a larger value of J/k_f (i.e., either faster photolysis of CIOOCI, slower rate of formation of CIOOCI by the self reaction of CIO, or some combination of these two perturbations) than the JPL06 value of this ratio.

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The best fit to the CIO measurements is achieved by using a higher $K_{eq scale}$ value ($K_{eq} = 2.0 \times K_{eq JPL06}$) than currently recommended by JPL06/09, leading to a $J_{scale}/k_{f scale}$ value of $1.23 \pm 0.11 \times (J/k_{f})_{JPL06}$. While this finding agrees with the higher K_{eq} values found by Ferracci and Rowley (2010), we caution that our retrieved values of K_{eq} are not well constrained by the measurements of CIO.

In agreement with Stimpfle et al. (2004), using a J_{JPL06} value together with a smaller $k_{\rm f}$ value relative to JPL06 (Trolier et al., 1990) was found to reproduce the CIO mea-²⁰ surements. Furthermore, Stimpfle et al. concluded that the laboratory measurement of $K_{\rm eq}$ from Cox and Hayman (1988) agrees best with CIO and CIOOCI observations (190 K<7 <200 K). Prescribing the $K_{\rm eq}$ value derived by Cox and Hayman results in a $J_{\rm scale}/k_{\rm f}$ scale ratio of 1.54 ± 0.21 . This ratio would also include such a combination of Jand $k_{\rm f}$ as mentioned above, i.e. $J_{\rm JPL06}$ and $k_{\rm f}$ Trolier, to explain the CIO measurements. ²⁵ $K_{\rm eq}$ determined by Plenge et al. (2005) would lead to a $J/k_{\rm f}$ ratio of $1.75 \pm 0.51 \times (J/k_{\rm f})_{\rm JPL06}$. This result suggests that the photolysis rate of the dimer is higher than currently recommended and lies within the range of 1.25 to $2.25 \times J_{\rm JPL06}$ if $k_{\rm f}$ JPL06 is correct. If the photolysis rate determined by JPL06 is used, then $k_{\rm f}$ has to be smaller than currently recommended (between 45% and 80% of $k_{\rm f, JPL06}$). These results sug-



gest that using $K_{eq Plenge}$ would lead to higher modelled CIO abundances than using the current JPL recommendations.

If K_{eq} is fixed to JPL06 then our results indicate that $J/k_f = 1.35 \pm 0.14 \times (J/k_f)_{JPL06}$. Using J_{JPL06} requires a smaller k_f value than currently recommended and our results show that k_f must lie in the range of 0.6 to $0.95 \times k_{f JPL06}$ to explain the measurements. Therefore, the results presented by Bloss et al. (2001) and Trolier et al. (1990) for k_f together with the JPL06 recommendation for J and K_{eq} are consistent with the CIO measurements used in this study.

Using a combination of $K_{eq \ scale}=0.2$ (which corresponds to the lowest $K_{eq \ scale}$ value given the prescribed K_{eq} range in this study) and a *J* value obtained from the Burkholder et al. (1990) cross-sections, which agrees with the value of *J* found using cross sections of the more recent study by Papanastasiou et al. (2009), requires a $k_{f \ scale}$ value greater than 0.55 to be consistent with the CIO measurements. An increase in both *J* and k_{f} relative to the JPL06 recommendation would lead to an intrease in the calculated ozone loss by the CIO dimer cycle, because chlorine is cycled more quickly through the various steps of the cycle.

 K_{eq} governs the partitioning of CIO and its dimer primarily during the night. Therefore, CIO day-time measurements are only weakly sensitive to K_{eq} , which likely explains the small change in the RMS values over the relatively large prescribed range of $K_{eq \ scale}$ (0.2 to 2.0). Day-time and night-time measurements of CIO are required to determine a reliable value of K_{eq} . Although this study does not provide a tight constraint on K_{eq} , the retrieved J/k_{f} obtained here is robust against the uncertainty in K_{eq} . However, day-time and night-time measurements would also provide information on Jand k_{f} , and their uncertainties, individually.

²⁵ Night-time CIO measurements were used in the past to determine the equilibrium constant (Stimpfle et al., 2004; Santee et al., 2010). Given these studies and studies performed in the laboratory (e.g., Plenge et al., 2005), it is more likely that $K_{eq scale}$ lies between 0.27 and 0.5 than being far greater than recommended by JPL06. Furthermore, the study by von Hobe et al. (2007) found that their stratospheric CIO ob-



servations were described well using a similar scaling for K_{eq} . The JPL09 recommendation also provides a smaller K_{eq} value ($0.7 \times K_{eq JPL06}$) than previously recommended in JPL06. The range of $K_{eq scale}$ between 0.27 and 0.7, together with the results from our OE approach, suggest that J/k_f values lie between 1.75±0.29 and 1.44±0.17×(J/k_f)_{JPL06}. The analyses presented above show that using these combinations of the kinetic parameters would include J and k_f values which are consistent with the JPL06 recommendations within the given uncertainty ranges of J_{JPL06} and $k_{f JPL06}$. Furthermore, this result would preclude J values smaller than currently recommended (e.g., Pope et al., 2007) and would exclude the lower uncertainty range on the J_{JPL06} value.

 K_{eq} is known to be temperature dependent (e.g., Santee et al., 2010); an increase in temperature leads to a decrease in K_{eq} . As shown in Fig. 1c, stratospheric temperatures above Scott Base during the period of CIO measurements, and over the altitude range where CIO concentrations maximize, varied from 183.7 K to 217.6 K. In contrast, most of the earlier studies took place in conditions with temperatures above 190 K. 15 Schofield et al. (2008), for example, examined Arctic CIO measurements made at temperatures between 200 and 207 K. The study by Schofield et al. resulted in a very small K_{eq} value of $0.2 \times K_{eq JPL06}$. The differences in the underlying temperature fields could, in part, explain why K_{eq} values higher than currently recommended (JPL06/09) and previously published (Santee et al., 2010) can be used to reproduce the Antarc-20 tic CIO measurements used in this study. For temperature to be partially responsible for the differences in $K_{\rm eq\ scale}$ across these studies, the formulation of the temperature dependence (i.e. $\exp(B/T)$) would need to be incorrect. This highlights the need for a greater number of laboratory and field studies at the low temperatures typical of the

²⁵ Antarctic stratosphere.

Acknowledgements. We thank Alan Parrish for his significant input in both, the development of the CIO microwave radiometer and the initiation of the ground-based measurement program. We thank S. Wood, M. Kotkamp and Antarctica New Zealand for their help in maintaining the CIO microwave radiometer and their help in obtaining the CIO measurements. The



ground-based measurement program is funded by NASA, grant NNX09AF40G. R. Schofield was funded by European Union (EU) project WaVES (MIF1-CT-2006-039646). S. Kremser thanks the German Academic Exchange Service (DAAD) for their support throughout the Doktorandenstipendium, allowing this work to be conducted.

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Fig. 1. Measured CIO in ppbv (a), the corresponding errors (b) and the coincident temperatures in K (c; referred to in the discussion) plotted as a function of altitude and day of the year in 2005. There are no CIO measurements available for the period of 14–16 September 2005 (day 257 to 259) due to bad weather conditions (white bar).



Vertical bars along the Y-Axis: results from earlier studies for $J_{\text{scale}}/k_{\text{f scale}}$ at 190 K. Blue symbols along the X-axis: $K_{eq \ scale}$ relative to JPL06 determined from earlier studies (190< T <210 K). Light grey area: JPL06 uncertainty range on K_{eq} and J/k_{f} . Hatched area: JPL09 uncertainty range on K_{eq} and J/k_{f} (for more details see text).

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