

**Ozone production in  
a dissipating tropical  
convective cell**

G. A. Morris et al.

**Observations of ozone production in  
a dissipating tropical convective cell  
during TC4**

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Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Abstract

From 13 July–9 August 2007, 25 ozonesondes were launched from Las Tablas, Panama as part of the Tropical Composition, Cloud, and Climate Coupling (TC4) mission. On 5 August, a strong convective cell formed in the Gulf of Panama. World Wide Lightning Location Network (WWLLN) data indicated 563 flashes (09:00–17:00 UTC) in the Gulf. NO<sub>2</sub> data from the Ozone Monitoring Instrument (OMI) show enhancements, suggesting lightning production of NO<sub>x</sub>. At 15:05 UTC, an ozonesonde ascended into the southern edge of the now dissipating convective cell as it moved west across the Azuero Peninsula. The balloon oscillated from 2.5–5.1 km five times (15:12–17:00 UTC), providing a unique examination of ozone (O<sub>3</sub>) photochemistry on the edge of a convective cell. Ozone increased at a rate of ~1.6–4.6 ppbv/hr between the first and last ascent, resulting cell wide in an increase of  $\sim(2.1\text{--}2.5)\times 10^6$  moles of O<sub>3</sub>. This estimate agrees to within a factor of two of our estimates of photochemical lightning O<sub>3</sub> production from the WWLLN flashes, from the radar-inferred lightning flash data, and from the OMI NO<sub>2</sub> data ( $\sim 1.2$ ,  $\sim 1.0$ , and  $\sim 1.7\times 10^6$  moles respectively), though all estimates have large uncertainties. Examination of DC-8 in situ and lidar O<sub>3</sub> data gathered around the Gulf that day suggests 70–97% of the O<sub>3</sub> change occurred in 2.5–5.1 km layer. A photochemical box model initialized with nearby TC4 aircraft trace gas data suggests these O<sub>3</sub> production rates are possible with our present understanding of photochemistry.

## 1 Introduction

Numerous studies in the latter half of the 20th Century have examined the role of lightning in global and regional ozone (O<sub>3</sub>) (e.g., Kroening and Ney, 1962; Orville, 1967; Griffing, 1977; Levine, et al., 1981; Pickering et al., 1990; Lelieveld and Crutzen, 1994; Lawrence et al., 2003; Doherty et al., 2005; Cooper et al., 2006; see also Table 1) and in particular, its photochemical production resulting from lightning-produced reactive

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## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

⏪

⏩

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



nitrogen ( $\text{NO}_x$ ) (Noxon, 1976; Chameides et al., 1977, 1979; Peyrous and Lapeyre, 1982). With refinements in regional and global models and lightning flash data from satellites and ground-based networks, the global lightning  $\text{NO}_x$  ( $\text{LNO}_x$ ) budget has been reliably set at 2–8 Tg N/yr (Pickering et al., 2009; Schumann and Huntrieser, 2007; Lawrence et al., 1995). Photochemical reactions involving  $\text{LNO}_x$  lead to  $\text{O}_3$  formation. Grewe (2007) estimates the global lightning contribution to tropospheric  $\text{O}_3$  at >30%, with the fractional source greatest in the tropics.

Laboratory studies have suggested that direct production of  $\text{O}_3$  occurs mainly during the pre-discharge period of storms (Peyrous and Lapeyre, 1982), and Franzblau (1991) found (1) little  $\text{O}_3$  production except at very low energies, (2) large decreases in  $\text{O}_3$  immediately after discharges, and (3) nearly full recovery to pre-discharge levels after ~10 min (see his Fig. 1).

A number of field campaigns and modeling studies have attempted to identify the role of lightning in the formation and/or destruction of tropospheric ozone, and the role of convective activity in the redistribution of  $\text{NO}_x$ ,  $\text{O}_3$ , and other trace species within the troposphere. The results from these studies vary widely, with some suggesting strong enhancements of  $\text{O}_3$  (e.g., Schlanta and Moore, 1972; Dickerson et al., 1987) and others suggesting no such enhancements and potential losses of  $\text{O}_3$  (e.g., Ridley et al., 2006; Ott et al., 2007; Salzmann et al., 2008). Results from these previous studies are summarized in Table 1.

The TC4 campaign (Toon et al., 2010) was an excellent vehicle to further investigate wave activity, lightning, and  $\text{O}_3$  responses in a highly-convective environment, closer to the Intertropical Convergence Zone than the August–September Intercontinental Transport Experiment (INTEX)-B Ozonesonde Network Study (IONS-06) soundings (e.g., see Thompson et al., 2008). The concentration of new satellite products and aircraft instrumentation focused on convection permits us to quantify lightning, to relate flashes to improved representations in models, and to link the latter to validation of  $\text{NO}_2$  from the Ozone Measuring Instrument (OMI) aboard the NASA Aura satellite (e.g., Bucsela et al., 2010).

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

⏪

⏩

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



In addition to O<sub>3</sub> measurements on the three TC4 aircraft, the Southern Hemisphere Additional OZonedondes (SHADOZ) (Thompson et al., 2003) Costa Rican station (Heredia, 10° N, 84° W) and the TC4 site at Las Tablas, Panama (7.75° N, 80.25° W; see also (Thompson et al., 2010) provided O<sub>3</sub> profiles. Virtually all the Costa Rican and Panama soundings contained free-tropospheric (FT) and lower-stratospheric wave signatures, with an incidence in the tropical tropopause layer (TTL) of >40% (see Fig. 2a in Thompson et al., 2010). Case studies of O<sub>3</sub> within segments affected by gravity waves demonstrated a clear link to convective activity.

The 25 Panama sondes, in a region with a high convective frequency during TC4 (Toon et al., 2010), displayed prominent wave activity associated with convection near the beginning of the mission (13–22 July 2007) and after 2 August 2007, when the TC4 aircraft coordinated sampling south of Costa Rica, in the vicinity of the Panama Bight, and as far south as the Galapagos. On 5 August, a day when all three TC4 aircraft flew over the Panama Bight, the ozonesonde launched from Las Tablas was caught in a convective system that kept it oscillating between 2.5–5.1 km for nearly two hours before it resumed normal ascent. Though many previous studies provided profiles before and after convection, this ozonesonde data set is unique in providing insights into changing O<sub>3</sub> concentrations inside a dissipating tropical convective cell. Ozone increased throughout the oscillatory period, and we trace the cause to lightning-induced photochemistry. The following sections provide background with experimental details (Sect. 2), observations (Sect. 3), photochemical model results (Sect. 4), and a summary (Sect. 5).

## 2 Background

### 2.1 NPOL radar

The NASA polarimetric Doppler weather radar (NPOL) is an S-Band system operating at a frequency of 2.8 GHz (10 cm wavelength). It has horizontal and vertical beam

widths of  $1.4^\circ$ . The antenna is a flat, passive array instead of the typical parabolic dish. The design allows the system to be transported easily and allows NPOL to operate in conditions with strong winds (e.g., minimum wind loading). One of the main drawbacks of this design is that the signal deteriorates when the antenna is wet (Theisen et al., 2009). Therefore, the final quality-controlled dataset excludes periods when there is precipitation at the radar site.

For TC4, NPOL was deployed next to the Nittany Atmospheric Trailer and Integrated Validation Experiment (NATIVE) mobile laboratory near Las Tablas. NPOL operated almost continuously 16 July–12 August 2007, with the exception of the period 18:00 UTC 19 July–02:00 UTC 21 July. NPOL scanned with a temporal resolution of 10 min and spatial resolution of 200 m using a 12-tilt scanning strategy with elevation angles ranging from  $0.7^\circ$  to  $23.3^\circ$ . A volume scan with a maximum range of 150 km preceded each long-range (to 275 km) surveillance scan. NPOL-measured or derived quantities included standard radar parameters: radar reflectivity (DZ), Doppler velocity (VR), and spectral width (SW); and polarimetric parameters: differential reflectivity (ZDR), differential phase (PhiDP), specific differential phase (KDP), and cross correlation (corrHV) between horizontal and vertical polarizations. Nearly 3500 volume scans are available for studying the convective properties in and around Panama, a data set that includes a variety of interesting events, ranging from short-lived unorganized convection to long-lived mesoscale convective systems.

## 2.2 World wide lightning location network

In 2007 the World-Wide Lightning Location Network (WWLLN) (Rodger et al., 2006) consisted of approximately 25 sensors detecting lightning flashes at VLF frequencies of 3–30 kHz. WWLLN Director, Robert Holzworth, of the University of Washington provided near-real-time flash data (primarily cloud-to-ground or CG flashes) to NASA Goddard Space Flight Center (GSFC). Bucsela et al. (2010) estimated the detection efficiency of the WWLLN for total flashes (CG+intracloud or IC flashes) in the TC4 region (over open ocean near Costa Rica and Panama) through comparisons of flash rates

### Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



from the Costa Rica Lightning Detection Network (CRLDN, which uses the same sensors as the United States National Lightning Detection Network, NLDN, see Cummins et al., 1998) and the Lightning Imaging Sensor (LIS) on the Tropical Rainfall Measuring Mission (TRMM) satellite for six storms. The mean detection efficiency was  $0.22 \pm 0.08$ , in reasonable agreement with, though somewhat higher, than the estimate of Rodger et al. (2006). There is some indication, however, that the detection efficiency is greater over the ocean than over the land in this part of the world (Lay et al., 2009).

### 2.3 OMI

OMI has been collecting data since its launch in July 2004 (Levelt et al., 2006). The instrument is a nadir-viewing spectrometer with a CCD-array having wavelength and spatial dimensions, the latter comprising 60 pixels across the flight track. The pixel area at nadir is  $13 \times 24 \text{ km}^2$ , although this value increases by approximately an order of magnitude near the edges of the track. Overpass time is  $\sim 13:45 \text{ LT}$  in the tropics, improving the capability of OMI for observing afternoon convective events as compared with GOME and SCIAMACHY (which have morning overpass times).

Bucsela et al. (2006, 2008), Celarier et al. (2008), and Wenig et al. (2008) describe the retrieval algorithm for  $\text{NO}_2$  from OMI. It employs a spectral fitting procedure to obtain  $\text{NO}_2$  slant column densities (SCDs) from the OMI spectra. Vertical column densities (VCDs) result from dividing SCDs by air mass factors (AMFs), which are derived through radiative transfer calculations. The tropospheric component of the vertical column, including both pollution and lightning sources, is obtained by removing an unpolluted (here simply called “stratospheric”) component, using a wave-2 analysis in narrow latitude bands. Because of the method used to derive it, the stratospheric VCD is contaminated by small amounts of tropospheric  $\text{NO}_2$ , which we remove in the present study using output from the GMI model (Duncan et al., 2007). The corrected stratospheric VCD differs from the uncorrected value by approximately 5%.

For the 5 August 2007 analysis, OMI  $\text{NO}_2$  data provide estimates of  $\text{LNO}_x$  in the region near the Gulf of Panama. Bucsela et al. (2010) provide details of the procedure,

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



but we outline the method briefly. The first step is the calculation of the tropospheric SCD due to  $\text{LNO}_2$ , which is given by the total SCD minus the sum of the corrected stratospheric SCD and the tropospheric SCD due to sources other than lightning. This non-lightning tropospheric SCD is obtained from the OMI observations in the TC4 regions on days with a minimum of convective activity. In these calculations, all SCDs and VCDs are related through AMFs derived using radiative-transfer calculations and climatological  $\text{NO}_2$  profiles. We convert the  $\text{LNO}_2$  slant column to a vertical column of  $\text{LNO}_x$  using a modified AMF that accounts for the vertical distribution of  $\text{LNO}_2$  and the photolysis ratio of  $[\text{NO}_2]/[\text{NO}_x]$ . The former is obtained from TC4 aircraft data, and the latter from model calculations. Beirle et al. (2009) describe a similar approach.

The uncertainties in the OMI estimate of  $\text{LNO}_x$  are dominated by uncertainties in the global stratospheric and local background concentrations. These are treated as systematic errors, as opposed to random errors. The pixel-scale errors include variations in cloud parameters, photolysis ratios and the local a priori  $\text{NO}_2$  profile, and make a negligible contribution to the total error budget. Other more significant errors include the uncertainties associated with the size of the region of interest and the number of lightning strokes contributing to the observed  $\text{LNO}_x$ . As a result, the overall uncertainty in this estimate is large, though comparable to uncertainties associated with the other  $\text{LNO}_x$  estimates described below.

## 2.4 Ozonesondes

Launched from the NATIVE facility, the electrochemical concentration cell (ECC) type (Komhyr, 1986; Komhyr et al., 1995) En-Sci 2Z ozonesonde instruments with 0.5% buffered, KI cathode solution provided  $\text{O}_3$  profiles during TC4 at the Las Tablas site. The Jülich Ozone Sonde Intercomparison Experiment (JOSIE) found biases  $<5\%$ , a precision of 3–5%, and an accuracy of 5–10% up to 30 km for such sondes (Smit et al., 2007). With a typical rise rate of  $\sim 5$  m/s and a measurement time constant of  $\sim 25$  s, the effective vertical resolution of  $\text{O}_3$  features is  $\sim 125$  m (see also Smit et al., 2007). Most launches occurred from 12:00–15:00 LT (17:00–20:00 UTC) to coincide

### Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



with the ~13:45 local solar time overpass of NASA's Aura satellite (Schoeberl et al., 2006).

Vaisala RS80-15N radiosondes on each payload (described in Thompson et al., 2003 and 2007a) measured pressure, temperature, and relative humidity (RH). Payloads also contained global positioning systems (GPS) that provided latitude, longitude, altitude, wind speed, and wind direction data. Comparisons of pressure altitude with GPS altitude provide validation of the pressure measurements. When pressure offsets occur, they are usually <2 hPa, meaning that tropospheric O<sub>3</sub> mixing ratios are adjusted <~2% (<~0.2% at the surface). Post-flight processing corrects pressure errors so that at burst altitude, pressure and GPS altitudes agree to within 200 m. For 7 of the launches during TC4, RH data above 300–500 hPa appear unreliable. All the ozonesonde data can be found at: <http://physics.valpo.edu/ozone/tc4data.html>.

Although each ozonesonde is internally calibrated before each flight, Fig. 1 shows a comparison of surface O<sub>3</sub> readings with pre-launch ozonesonde readings from the 23 flights with good O<sub>3</sub> data during TC4. The Thermo Electron Corporation (TECO) Model 49C Ozone Analyzer, using the United States Environmental Protection Agency standard measurement technique (EQOA-0880-047) and located on the Nittany Atmospheric Trailer and Integrated Validation Experiment (NATIVE-data at [http://ozone.met.psu.edu/Panama\\_Data/](http://ozone.met.psu.edu/Panama_Data/)) trailer in Las Tablas, provided the pre-flight surface O<sub>3</sub> data. The mean bias (sonde-NATIVE) was  $-0.4 \pm 1.2$  ppbv, with a root mean square difference of  $1.05 \pm 0.76$  ppbv.

We compare ozonesonde columns with OMI total column ozone (Bhartia, 2007; McPeters et al., 2008) for nearby overpasses (<50 km from launch site). To determine the ozone column from the ozonesonde data, we integrate ozone profiles to the burst altitude, then augment the column with either a constant mixing ratio assumption for the upper stratosphere or the Solar Backscatter Ultra-Violet (SBUV) balloon-burst climatology of McPeters et al. (1997). For the former case, the difference is  $17.5 \pm 3.8$  DU, while for the latter, the difference is  $16.7 \pm 6.2$  DU, with the sondes higher than OMI by ~6%. These results are consistent with the finding that the Paramaribo SHADOZ

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



station (5.8° N, 55.2° W) reports columns ~10% higher than OMI (Thompson et al., 2007c).

### 3 Observations

#### 3.1 NPOL

NPOL observed a large convective system off the coast of Panama on 5 August 2007. The system developed during overnight hours in the Gulf of Panama and slowly propagated westward toward the Azuero Peninsula. NPOL data indicated that the precipitating area covered several hundred kilometers in the north-south direction and ~100 km in the east-west direction, with the convective core having a mean area of  $5300 \pm 2400 \text{ km}^2$  between 09:00–17:00 UTC. The peak convection occurred ~13:11 UTC. Figure 2 shows the reflectivity (DZ) field for 12:55–13:05 UTC. Reflectivity values ranged from about 0 dBZ in the lighter precipitation areas to a maximum of about 55 dBZ in the embedded convection.

Estimated vertical velocities are derived from NPOL Doppler velocity field through a technique called volume velocity processing (VVP) (Boccippio, 1995). For the 5 August cell, the derived wind in the lower atmosphere was easterly at speeds on the order of 10 m/s. With each time step, however, moderate directional shear appeared, with directions fluctuating from southeast to northeast (in agreement with the ozonesonde observations, see Sect. 3.4 below).

This convective system generated a significant number of lightning strikes. Figure 2 indicates the location of flashes between 12:55 and 13:05 UTC detected by the WWLLN (see Sect. 3.3 below) are well correlated with areas active convection indicated by the NPOL data. We use radar reflectivity data from NPOL to estimate lightning flash rates using the parameterization of Futyan and DelGenio (2007), which predicts the total flash rate  $IC+GC=F$  (flashes/min/300 km<sup>2</sup>) to be

$$F = 0.208(H_{17\text{dB}} - H_{0\text{°C}})^{1.8} \quad (1)$$

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

⏪

⏩

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



where  $H_{17\text{dB}}$  is the storm-averaged height (km) of the 17 dB radar-return signal and  $H_{0^\circ\text{C}}$  is the height of the  $0^\circ\text{C}$  (freezing) level. From the ozonesonde data, we find  $H_{0^\circ\text{C}}$  is  $\sim 4.5$  km. From the radar data, we determine hourly averages of  $H_{17\text{dB}}$  and the area of the storm. Table 2 shows the resulting hourly flash estimates for 5 August. During the period from the genesis of the cell to the final ascent of the balloon (09:00–17:00 UTC), this parameterization predicts a total of  $2300 \pm 300$  flashes.

We can use the NPOL-derived flash data to estimate the photochemical lightning production of  $\text{O}_3$  ( $\text{LpcO}_3$ ) from LNO. Estimates of LNO vary widely (Pickering et al., 2009; Huntrieser et al., 2008; Koike et al., 2007; Hudman et al., 2007; Drapcho et al., 1983) from a low of 43 moles/flash (Skamarock et al., 2003) to a high of 1100 moles/flash (Price et al., 1997; Winterath et al., 1999). Pierce (1970) and Prentice and Mackerras (1977) estimate the ratio of intra-cloud (IC) flashes to cloud-to-ground (CG) flashes for  $\sim 8^\circ\text{N}$  to be 6.5–8.0. Estimates of the NO production efficiency of IC-to-CG flashes varies from 0.1 (Price et al., 1997) to 1.4 (Fehr et al., 2004). Lin et al. (1988) estimate the production of 30 moles of  $\text{O}_3$  per mole of  $\text{NO}_x$ , (i.e., an ozone production efficiency, OPE, of 30) in box model studies of surface ozone pollution while more recent studies find OPE in the range 4–12. For example, Shon et al. (2008) found an OPE of 4.5 for air downwind of an industrial complex air in Mexico and 8.5 for free tropospheric marine air; Wood et al. (2009) found an OPE of  $\sim 7$  over the course of a day at a mountain observatory near Mexico City; and Zaveri et al. (2003) used aircraft and surface observations during the Southern Oxidant Study, finding OPEs increased with plume age from  $\sim 2$ –5.6. (Note: for our calculations of  $\text{LpcO}_3$ , we scale the OPE to account for the limited 2-h period between the first and final ascent and assume that enough sunlight was available on the edge of the cell to drive the photochemistry).

Using a Monte Carlo approach to combine the various parameters cited above, we estimate  $\text{LpcO}_3$  of  $(4.2 \pm 4.0) \times 10^6$  moles from this cell. The large uncertainty in this estimate owes to the remaining high uncertainty in all of the quantities that go into the calculation. Recalculating with values that we feel are most representative of the conditions for this cell (tropical, marine, etc.: OPE=2; moles  $\text{NO}/\text{flash}=227$  as sug-

gested by Bucselá et al. (2010); IC:CG NO production efficiency=1), we determine our best estimate of LpcO<sub>3</sub> to be 1.0×10<sup>6</sup> moles based on the NPOL-based lightning flash rates.

If we assume O<sub>3</sub> is produced directly from the lightning strikes (LdO<sub>3</sub>) at a rate of 300–3000 moles/flash, as in Minschwaner et al. (2008), we predict (3.7±1.8)×10<sup>6</sup> moles of O<sub>3</sub> for the flashes between 09:00–17:00 UTC. Since the ozonesonde observations suggest that the O<sub>3</sub> production was ongoing during at least the 108-min period after launch, only the LdO<sub>3</sub> after launch is relevant. Therefore, if we restrict our calculation to the flashes during the two-hour period after launch (see Table 2), we find possible (1.39±0.70)×10<sup>6</sup> moles of LdO<sub>3</sub>.

### 3.2 WWLLN

WWLLN reported frequent lightning in association with the convective cell over the Gulf of Panama on 5 August. Table 2 summarizes the number of flashes per hour detected in the box defined by the latitude range 7.25°–8.75° N and longitude range 78.75°–81.25° W, the region of the Gulf of Panama through which the convective cell passed. Figure 2 shows good correlation between the locations of the lightning flashes and the areas of active convective seen by the NPOL radar for 12:55–13:05 UTC. Notably, the CRLDN observed few if any lightning flashes over the Gulf of Panama on this day. Given the spatial distribution of lightning flashes observed by the CRLDN (not shown), it appears that the Gulf of Panama fell in a shadow of the network.

Using the WWLLN flash data between 09:00–17:00 UTC, we can estimate the associated total LpcO<sub>3</sub>, as we did for the NPOL data above with one further modification. Bucselá et al. (2010) estimating lightning detection efficiency of 0.22±0.08 for the WWLLN in the TC4 region. Accounting for this factor, we find 2560±930 flashes between 09:00–17:00 UTC, in reasonable agreement with the NPOL estimate of 2300±300 flashes. (Hour-by-hour estimates of the efficiency corrected WWLLN flashes can also be found in Table 2.)

Using a Monte Carlo approach to combine the factors in Sect. 3.1 above with the

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



WWLLN flash estimate, we find  $(3.3 \pm 4.5) \times 10^6$  moles of  $\text{LpcO}_3$ , with our best estimate of  $1.2 \times 10^6$  moles (using values associated with conditions more likely to be found in the present case, as we did for the NPOL calculation above), consistent with the NPOL estimate.

If we assume direct production of  $\text{O}_3$  of 300–3000 moles/flash, as in Minschwaner et al. (2008), and use the WWLLN efficiency corrected lightning strikes, we project  $(4.2 \pm 2.2) \times 10^6$  moles of  $\text{LdO}_3$  for the flashes between 09:00–17:00 UTC. Although the cloud parameterization used to estimate lightning frequency from the NPOL data (see Table 2) suggests lightning continued during the sonde oscillation period (15:00–17:00 UTC), the WWLLN suggests little-to-no lightning during this period. This mechanism, therefore, is not indicated by the WWLLN data for post-launch  $\text{O}_3$  production.

### 3.3 OMI $\text{NO}_2$

It is difficult to discern the  $\text{LNO}_2$  signal from an examination of Level 2 OMI tropospheric  $\text{NO}_2$  and cloud fraction data products that result from the standard retrieval conducted at NASA Goddard Space Flight Center (Bucsela et al., 2006), in part because the Level 2 data have not been cloud screened. With reprocessing that includes removing an estimate of background  $\text{NO}_2$  and applying an air mass factor more appropriate for convective outflow (Bucsela et al., 2010), the  $\text{LNO}_2$  becomes more evident. Figure 3 shows a map of the  $\text{LNO}_x$  field near the Gulf of Panama on 5 August 2007 after reprocessing (considering the  $\text{NO}_x$  to  $\text{NO}_2$  ratio at cloud-outflow levels). The boxed area ( $\sim 54\,000 \text{ km}^2$ ) contains  $1020 \pm 860 \text{ kmol LNO}_x$ , which would result in  $(1.2\text{--}8.9) \times 10^6$  moles of  $\text{O}_3$  (depending on the OPE selected). If we scale this estimate to the size of the core of the convective cell observed on the NPOL radar (average area of  $\sim 5300 \text{ km}^2$  from 09:00–17:00 UTC), the estimated  $\text{LpcO}_3$  becomes  $(1.2\text{--}8.6) \times 10^5$  moles, with a best estimate of  $(2.0 \pm 1.7) \times 10^5$  moles, whereas if we scale it to match the area of flashes detected by the WWLLN ( $\sim 46\,000 \text{ km}^2$ ), we find a range of  $(1.0\text{--}7.6) \times 10^6$  moles with a best estimate of  $(1.7 \pm 1.5) \times 10^6$  moles. Scaling by the

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[⏪](#)[⏩](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

larger area of the WWLLN estimate results in the best agreement with the NPOL and WWLLN LpcO<sub>3</sub> estimates detailed above. While the uncertainties are large, the values appear to be above background.

### 3.4 Ozone profiles

5 The ozonesonde launch on 5 August occurred at 15:05 UTC to coincide with the scheduled arrival of the NASA aircraft in the Panama area. At the time of launch, rain fell as part of the convective cell that had just moved ashore from the east, although ground observers reported no visible lightning. The surface temperature was ~24 °C with RH of 96% and a surface pressure of ~1003 hPa.

10 About 20 min into the flight, the balloon reached 5.1 km and began to descend. About 15 min later, it began to ascend again. Between launch and ~17:00 UTC, the balloon oscillated up and down through the air mass between ~2.5 and ~5.1 km five times, as shown in Fig. 4. (Note: each ascent is color-coded so subsequent figures can be analyzed more easily.) Although the detailed explanation for this behavior is beyond the scope of this paper (see Morris et al., 2010), it appears to be a combination of 15 downdrafts on the southern side of the westward-moving convective cell (based on NPOL radar data) and changing mass due to repeated condensation/evaporation of water and freezing/melting of ice on the surface of the balloon.

20 Figure 5 shows the O<sub>3</sub> concentrations measured on each ascent. Over the ~2h between the original ascent and the final ascent, O<sub>3</sub> in the layer between ~2.5 and ~5.1 km increased 4–12 ppbv, with a mean increase of 7.9±4.7 ppb. Integrating the change in O<sub>3</sub> between the first and last profiles from 2.55–5.11 km (the range of the oscillation), and assuming uniform production across the area of the storm (5300±2400 km<sup>2</sup> as indicated by the NPOL data from 09:00–17:00 UTC) in this layer, 25 we find a potential of ~3.0×10<sup>6</sup> moles of O<sub>3</sub> (with ~50% uncertainty) created as part of this cell, a number that agrees within roughly a factor of two with the best estimates from the lightning data (see Sects. 3.1 and 3.2) and the OMI data (see Sect. 3.3 above). We note that the estimates from the lightning data and from OMI represent the

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



total  $\text{LpcO}_3$  throughout the cloud, whereas our sonde estimate is based only on the enhancements in the 2.5–5.1 km layer.

Geostationary Operational Environment Satellite (GOES) data indicate cloud top temperatures consistent with heights of 14–15 km as the cell developed over the Gulf of Panama, descending to 11–12 km as the cell dissipated when moving ashore. Moderate-resolution Imaging Spectrometer (MODIS) data for this cell suggest the cloud tops reach 10–12 km. Both satellite data sets indicate the clouds reached altitudes far above the top of the layer in which the balloon oscillated and made repeated samples. To get a better estimate of  $\text{LO}_3$  throughout the cloud and estimate changes in  $\text{O}_3$  above 5.1 km, therefore, we use the  $\text{O}_3$  data from the DC-8 flight around the Gulf of Panama on 5 August. The mean DC-8 in situ  $\text{O}_3$  profiles before 17:17 UTC and after 17:17 UTC can be seen in Fig. 6. Figure 7 shows the locations of the “before”/“after” data points, with the former marked by green stars and the latter by green diamonds. If we integrate the difference between the final ozonesonde ascent profile and the mean aircraft profiles (before and after 17:17 UTC) from 2.5–12.0 km, we find  $4.2 \times 10^6$  and  $3.1 \times 10^6$  moles, respectively. Using the combination of the in situ DC-8 measurements and the ozonesonde data, we find that 70–97% of the ozone molecules were created in the layer in which the balloon oscillated (i.e., 2.5–5.1 km).

The DC-8 differential absorption lidar (DIAL) observations, shown as the small colored dots in Fig. 6, provide evidence for higher  $\text{O}_3$  downwind of convection as compared to observations upwind in clear-sky regions around the Gulf of Panama. The colors are coded to match the locations of these observations along the flight track shown in Fig. 7. Integrating the upwind-downwind differences from 1.6–8.5 km, we find an increase (post minus pre) of  $1.3 \times 10^7$  moles of  $\text{O}_3$ , with roughly two-thirds of this change occurring between 1.6–6.0 km.

Returning to the ozonesonde observations, the total change in  $\text{O}_3$  as observed by the balloon is given by

$$\frac{d[\text{O}_3]}{dt} = \frac{\partial[\text{O}_3]}{\partial t} + \mathbf{v} \cdot \nabla[\text{O}_3] \quad (2)$$

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

where the first term represents in situ photochemical production (loss) and the second term represents changes due to advection. If the balloon had remained in the same air parcel within the cloud or if no wind shear existed over the vertical range of oscillation, we could assume that the observations were Lagrangian, meaning the advection term would vanish. Figure 8 shows the wind speeds and directions on each ascent of the balloon as determined from on board GPS data. Because of the vertical wind shear within the cell (seen by NPOL and in the balloon data), the advection term (resulting from the difference between the trajectories of the balloon and the air parcels in which observations were taken) may be non-negligible, so we investigate further below.

Table 3 shows the calculated differences between the GPS-recorded balloon positions and the estimated subsequent positions of the air masses sampled on the first ascent at three levels (2.75, 3.75, and  $4.75 \pm 0.15$  km). To make these estimates, we calculated trajectories based upon GPS wind speed and direction vertically averaged in each of the three layers on successive ascents. We multiplied the resulting  $u$  (east-west wind) and  $v$  (north-south wind) values by the time difference between successive ascents to get longitude and latitude displacements. Since the balloon did not oscillate through all three layers each time, the table contains some “No data” entries. These calculations suggest separations of 15–30 km between the originally- and finally-sampled air masses, providing a constraint on the horizontal scale of potential  $O_3$  gradients.

Figure 9 shows the change in  $O_3$  with time as a function of altitude, calculated as the difference between the  $O_3$  at a given altitude as measured on each ascent with that measured on the first ascent. Changes of 3–10 ppbv/h are derived, with the highest rates at 2.5–3.0 km between the first and second ascents, and at 3.5–4.5 km between the first and third ascents. We note from Table 3 that for the former case, the balloon is located about 5 km upwind from the original air mass, somewhat farther from the center of the storm as it comes ashore. For the latter case, the balloon is located about 13 km downwind of the original air mass, closer to the center of the cell. Since both profiles suggest somewhat large  $O_3$  changes with time, with one being upwind and the

other being downwind of the original air masses, it seems likely that the advection term over these spatial scales is relatively small.

One last component of the advection term to investigate is the vertical term. While some of the change in  $O_3$  is due to descent, most of the change appears to be due to other factors. First, the original ascending profile in Fig. 5 (purple) joins with near perfect continuity the final ascending profile between 5 and 5.5 km, suggesting the  $O_3$  enhancements are not due to air descending from above 5.5 km. Second, Fig. 10 shows the change in potential temperature ( $\theta$ ) between the first and final ascent as a function of altitude between 2.55–5.11 km, with warming between 2.55 and ~3.40 km of 1.25–2.25 K, and between 3.4–5.11 km of 0.25–1.0 K. If we assume that the change in  $\theta$  in the lower layer is due solely to descent, the air found in the 2.55–3.40 km layer on the last pass was originally between 3.09–3.55 km. The mean  $O_3$  in this layer (defined by  $\theta$ ) on the first pass was  $32.1 \pm 1.3$  ppb while on the last pass, it was  $37.9 \pm 2.2$  ppb, a difference of  $5.8 \pm 2.6$  ppb. If we assume no descent at all, the mean  $O_3$  in this layer (defined by altitude) was  $29.5 \pm 1.9$  ppbv on the first pass, a difference of  $8.4 \pm 2.9$  ppbv. Thus, we attribute  $2.6 \pm 3.9$  ppbv (~31%) of the change to descent of air between the first and last ascents, with air descending at an average rate of 6.2 cm/s. Reducing the observed change in  $O_3$  of 4–12 ppbv over ~108 min (the time between the first and final ascent) by the 30% due to descent, the balloon data project 1.5–4.6 ppbv/h change in  $O_3$ .

Figure 11 shows  $O_3$  as a function of  $\theta$  rather than height for the first and last ascents. Integrating the change in  $O_3$  as a function of  $\theta$  between 311.75–320.00 K, and assuming the cell size as before, we find an increase of  $2.5 \times 10^6$  moles of  $O_3$ , agreeing within roughly a factor of two with the estimates from the lightning data detailed above.

To establish characteristics of the large scale flow, we calculated back trajectories from 2.75–4.75 km using both the Hysplit model (Draxler and Rolph, 2010) (using Global Data Assimilation System (GDAS) meteorological fields with a horizontal resolution of  $1^\circ$  latitude  $\times$   $1^\circ$  longitude) and a kinematic version of the NASA Goddard Space

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



Flight Center trajectory model (GTM, Schoeberl and Sparling, 1995) using National Centers for Environmental Prediction (NCEP) reanalysis meteorological fields (horizontal resolution of  $1^\circ$  latitude  $\times$   $1^\circ$  longitude and a 6-h time resolution). The results from both models suggest that air masses between 2.75 and 4.75 km moved across the Gulf of Panama from the ENE to Las Tablas over the 8 h prior to the balloon observation, consistent with the motion of the cell seen in the NPOL radar data. The trajectory data, therefore, suggest that the air mass was not a mixture of air from vastly different source regions.

#### 4 Photochemical modeling

In an effort to simulate the  $O_3$  changes observed by the 5 August 2007 ozonesonde, we ran a zero-dimensional photochemical box model with Regional Atmospheric Chemical Mechanism (RACM, Stockwell et al., 1997) to calculate the concentrations of radicals and other reactive intermediates. Rate coefficients of the reactions in RACM were updated by Jet Propulsion Laboratory (JPL) data evaluation (Sander et al., 2006) as applicable. The model was constrained by a set of 1-min resolution merged observations of temperature, pressure,  $H_2O$ ,  $O_3$ ,  $CO$ ,  $NO$ ,  $NO_2$ ,  $HNO_3$ , NMHC, ethanol, acetone, MACR, MVK, and benzene from NASA DC-8 aircraft at 16:45–17:00 UTC on 5 August 2007. During this time period, the aircraft was spiraling downward from 5.2 to 0.3 km near  $5.7^\circ$  N and  $78.0^\circ$  W, in a clear-sky region to the southeast of the active region of convection in which the balloon measurements occurred. Unmeasured photolysis frequencies were calculated for clear sky condition from the Tropospheric Ultraviolet and Visible (TUV) model (<http://www.acd.ucar.edu/TUV>) or based on the solar zenith angle equations in Jenkin et al. (1997) and then scaled by the photolysis frequency of  $NO_2$  calculated from the NCAR CCD Actinic Flux Spectroradiometers (CAFS) measurements made on the NASA WB-57 aircraft that flew in the vicinity of active convection around the same time as the DC-8. The model was run with the FACSIMILE software (MCPA Software Ltd.) for sufficient time to reach instantaneous steady state of the

### Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



intermediates.

The instantaneous net O<sub>3</sub> production,  $P_{\text{net}}(\text{O}_3)$ , can be calculated by

$$\begin{aligned} P_{\text{net}}(\text{O}_3) &= P(\text{O}_3) - L(\text{O}_3) \\ &= k_{\text{NO}+\text{HO}_2}[\text{NO}][\text{HO}_2] + \sum k_{\text{NO}+\text{RO}_2,i}[\text{NO}][\text{RO}_2] \\ &\quad - k_{\text{OH}+\text{NO}_2+\text{M}}[\text{M}][\text{NO}_2][\text{OH}] - k_{\text{O}^1\text{D}+\text{H}_2\text{O}}[\text{O}^1\text{D}][\text{H}_2\text{O}] \\ &\quad - k_{\text{HO}_2+\text{O}_3}[\text{HO}_2][\text{O}_3] - k_{\text{OH}+\text{O}_3}[\text{OH}][\text{O}_3], \end{aligned} \quad (3)$$

where the various  $k$ 's represent the rate coefficients of the corresponding reactions associated with the production and loss of O<sub>3</sub>. The resulting calculated net ozone production as a function of altitude is shown in Fig. 12a, with a mean value of 0.84 ppbv/h between 2–5 km. Because this aircraft spiral occurred in a region with much less convective activity than the cell in which the ozone sonde made its measurements, we performed four additional model runs with NO<sub>x</sub> concentrations 150%, 200%, 300% and 500% of those detected in this spiral (Fig. 12a). The runs resulted in 2–5 km averaged  $P_{\text{net}}(\text{O}_3)$  of 1.5, 2.2, 3.3, and 5.3 ppbv/h, respectively.

Table 2 summarizes the WWLLN flash detection for 5 August in a box defined by 5.0–6.5° N, 75.5–78.5° W, a box of roughly the same area as used for the analysis of the convective cell in the Gulf of Panama discussed above. This box is located upwind of the area of the DC-8 spiral on 5 August. Hysplit trajectory calculations (not shown) suggest that air sampled by the DC-8 in its spiral would have been located in this box earlier in the day. The WWLLN data indicate that the air mass sampled during the DC-8 spiral saw only ~10% as much lightning from 09:00–17:00 UTC and only ~34% as much lightning for the entire day as the area near the convective cell. Increases of NO<sub>x</sub> by a factor of 2–3 for use in modeling LpcO<sub>3</sub>, therefore, are reasonable, if not conservative.

Figure 12b examines the sensitivity of the calculated net ozone production rates on the photochemical rate constants. DeCaria et al. (2005) suggest a linear scaling of the clear sky photolysis rate constants using adjustment factors ranging from 0.1 at the base to 2.0 at the top of a very dense cloud, and from 0.4 at the base to 1.7 at the top

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



of a dense cloud. Because we do not know the precise structure of the cloud in which the balloon oscillated, we perform a sensitivity study, scaling the photolysis rates by the amounts suggested by DeCaria et al. (2005) and using the 200% NO<sub>x</sub> case described above. The average net ozone production for the 0.1, 0.4, 1.7, and 2.0 cases are 0.83, 1.5, 2.6, and 2.7 ppb/h, respectively.

The Cloud Physics Lidar (CPL, McGill et al., 2004) and Cloud Radar System (CRS, D. Hlavka et al., unpublished manuscript) data from the DC-8 flight on 5 August (shown in Fig. 10 of Thompson et al., 2010) suggest some structure topping out in the 4–6 km altitude range, near the top of the range of the balloon oscillation. The production rates in the cloud sampled on 5 August, therefore, should move from the 0.1 or 0.4 scaled curves in Fig. 12a near 2.5 km toward the 1.0 to 2.0 scaled curves near 5.0 km, resulting in estimated production rates of 1.0–1.7 ppb/h near 2.5 km to 1.3–1.6 ppb/h near 5.0 km. Thus, the photochemical model results are consistent with ~1.0–2.5 ppbv/h in the 2–5 km altitude range sampled by the oscillating balloon. (We note that by using the 500% NO<sub>x</sub> case for this sensitivity analysis, we find ~2.1–4.7 ppb/h at the cloud base and ~3.5–4.5 ppb/h near 5.0 km, results that are quite consistent with the rates calculated from the balloon data above).

Figure 13 shows NO data, using a chemiluminescence technique (Fontjin et al., 1970; Pearson and Stedman, 1980), and NO<sub>2</sub> data, using the thermal dissociation-laser induced fluorescence technique (Thornton et al., 2000), from all DC-8 flights during TC-4. Only measurements recorded at least 10 min after take-off, at most 10 min before landing, and at least 50 km away from the airport in Aljujuela, Costa Rica (10.0° N, 84.2° W) are included in the plot. Relatively high values of NO<sub>x</sub> can be found throughout the troposphere. Tables 4 and 5 summarize all occasions when the NO and NO<sub>2</sub> concentrations exceeded 1.0 ppbv. We note two of these occasions in particular: 1) on 21 July 2007, a peak of 1.0–3.1 ppbv of NO is observed over the Caribbean Sea between 3.7–6.2 km; 2) on 29 July 2007, a peak of 1.0–1.5 ppbv of NO between 6.3–7.4 km and a peak of 1.0–1.8 ppbv of NO<sub>2</sub> between 5.5–7.4 km is observed over the Pacific Ocean. Observation 1 occurred as the aircraft was descend-

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



ing after passing through an area of fresh convection, as indicated by the GOES IR data (anlger.larc.nasa.gov/tc4) for 21 July 2007. Observation 2 occurred just off the west coast of Costa Rica immediately following take-off. The aircraft encountered a region of fresh convection, as indicated by the GOES IR data for 29 July 2007. These two examples from the TC4 aircraft data support the possibility that  $\text{NO}_x$  production associated with fresh convection can be observed in the mid-troposphere.

## 5 Discussion and summary

This work has presented a unique ozonesonde profile over Las Tablas, Panama on 5 August 2007. The balloon initially ascended on the southern side of a dissipating convective cell as it came ashore from the east. Between 09:00 and 17:00 UTC, WWLLN data indicate 563 flashes ( $\sim 2600$  flashes accounting for the lightning detection efficiency of this network) in and around the Gulf of Panama, while estimates of lightning flash rates using NPOL radar height data result in  $\sim 2300$  flashes associated with this cell. The ozonesonde oscillated 5 times between  $\sim 2.5$  and  $\sim 5.1$  km for  $\sim 108$  min and measured 4–12 ppbv  $\text{O}_3$  increases. Examination of the meteorological data gathered on the sonde flight suggests that  $\sim 30\%$  of the increase may be due to descent of higher  $\text{O}_3$  from above. As a result, we calculate  $\text{LO}_3$  from the balloon measurements of 1.5–4.6 ppbv/h. Using data gathered on DC-8 flight in the Gulf of Panama region on 5 August 2007, we initialized a run of the RACM to evaluate  $\text{LpcO}_3$  from the background conditions. Although the aircraft samples used to initialize the RACM run came from a region with only 10–34% as many lightning flashes, the base run showed a  $\text{LpcO}_3$  of 1.5 ppbv/h between 2–5 km. Increasing  $\text{NO}_x$  by a factor of 2–3 and compensating for changes in photolysis rates due to cloud cover, as suggested by DeCaria et al. (2005), led to net  $\text{LpcO}_3$  of 1.0–2.5 ppbv/h (increasing  $\text{NO}_x$  by a factor of 5 yields net ozone production rates of 2.1–4.7 ppbv/h). The RACM model results, therefore, indicate that most of the  $\text{LpcO}_3$  observed by the ozonesonde can result from known photochemistry.

A number of previous studies have suggested that strong  $\text{O}_3$  increases are seen at

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



higher altitudes downwind of convective activity (see Table 1 for a summary of previous findings). For example, DeCaria et al. (2005) found  $O_3$  production of 10 ppb/day near 9 km, and Pickering et al. (1996) and Thompson et al. (1997) found 7–8 ppb/day in the 8–12 km.

5 Other studies suggest  $O_3$  losses due to titration by NO near 5 km in association with convective cells, including 6–8 ppbv losses found in Ridley et al. (2006), 9 ppbv losses during a convective event in Europe at 5.5 km reported by Ott et al. (2007), and maximum  $O_3$  losses at 5 km during a storm in tropical Pacific reported by Salzmann et al. (2008).

10 But numerous previous studies also have reported large increases in  $O_3$  in the lower troposphere. Dickerson et al. (1987) found a peak in  $O_3$  at 5 km in association with convective activity, while Pickering et al. (1990) suggest the highest potential  $O_3$  production of 7–17 ppbv/day below 5 km. Viewed in this context, our ozonesonde observations provide evidence for this layer of enhanced  $O_3$  in the lower troposphere.

15 The rapid increase of  $O_3$  seen in our sonde data may suggest that much of the  $O_3$  production occurs soon after daylight returns to the air mass affected by the cell rather than over the course of a day, a hypothesis consistent with the laboratory work of Franzblau (1991, see his Fig. 2 which shows rapid recovery of  $O_3$  in <10 min. when a UV lamp illuminates a controlled chamber after electrical discharges result in loss of  $O_3$  through reactions with  $LNO_x$ ). Although our balloon payload contained no instruments with which direct measurements of radiation could be made, the meteorological observations from the ozonesonde flight are consistent with increasing radiation at 2.5–5.1 km over the time of the oscillations. In particular, the sonde reported higher temperatures (by  $0.92 \pm 0.46$  °C) and lower humidity (by  $9.2 \pm 4.8$ %) on the final ascent as compared to the initial ascent between 2.6 and 5.1 km, both of which are consistent with increased sunlight. This explanation also is consistent with Ott et al. (2007), who reported 9 ppbv losses of  $O_3$  during a storm with large amounts of  $LNO_x$ , but 5 ppbv/day production of  $O_3$  downwind at 5.5 km.

25 Significant changes in the lower free tropospheric  $O_3$  associated with convective

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**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

---

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



activity are consistent with the lower portion of the C-shaped LNO<sub>x</sub> profile of Pickering et al. (1998), which shows 15.9% of the LNO<sub>x</sub> below 5 km, and the updated profiles of Ott et al. (2010), which shows 14.7% of LNO<sub>x</sub> below 5 km. While the ICARTT data shown in Hudman et al. (2007) do not indicate a significant enhancement of NO<sub>x</sub> in the 2–6 km altitude range, the DC-8 did not sample in this altitude range near convective activity; nevertheless the higher variability in NO<sub>x</sub> and corresponding lack of a similar increase in variability in the CO data at 3.5 km (see their Fig. 3) may be explained by lower tropospheric LNO<sub>x</sub>. As pointed out by Ridley et al. (2004), for cases in which CG flashes occur outside of the convective core or in regions of the convective cell influenced by downdrafts, enhancements of LNO<sub>x</sub> are likely in the lower troposphere. In fact, the DC-8 data from at least two TC4 flights near active convection found NO<sub>x</sub> enhancements in the 3.5–7.5 km altitude region (21 and 29 July 2007, see Tables 4 and 5). Given the location of our ozonesonde profile on the southern edge of the dissipating convective cell, either of these possibilities may explain our sonde observation.

Furthermore, the changes observed by the ozonesonde during the 108 min of its oscillation do not represent typical variability within this layer during a single day of the TC4 mission. To assess this variability, we examine the differences between ascending and descending ozonesonde measurements from 2.5–5.1 km, measurements that are typically separated by ~2 h in time and ~100 km. Using all flights with both ascending and descending data within this layer taken during TC4, we find that only ~30% of data in this altitude range showed absolute changes >4 ppbv, only ~5% of data in this altitude range showed absolute changes of >12 ppbv; only ~13% of all data showed differences >+4 ppbv and <1% showed changes >+12 ppbv. We also note that for the 5 August flight, 100% of the data from 2.6–5.1 km show O<sub>3</sub> increases. Of the 14 flights for which we have descending data at more than 50% of the levels from 2–6 km, only 3 flights (other than the 5 August flight) had at least 80% of the levels in this altitude range with increased O<sub>3</sub>, with mean increases of 5.8±3.4 ppbv, 3.7±2.6 ppbv, and 2.8±1.3 ppbv. These differences are less than those seen on the 5 August, which showed a mean increase of 7.8±2.2 ppbv. Thus, the 5 August flight appears unique

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[⏪](#)[⏩](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

among all of our TC4 ozonesonde flights in terms of ozone changes observed over  $\sim 2$  h between 2 and 6 km.

Wind profiles from the sonde and NPOL data indicate some vertical divergence within the layer of observed  $O_3$  changes. Subsequent trajectory calculations suggest the possible separation of the originally sampled air mass from that sampled on the final ascent by 15–30 km over the  $\sim 108$  min. Thus, horizontal  $O_3$  gradients with a scale  $< 30$  km and/or in-situ photochemistry may explain some of the sonde-observed changes in  $O_3$ . In the former case, the original air mass at 2.75 km ends up  $\sim 15$  km to the southwest of balloon trajectory (see Table 3), so the balloon ends up falling backward relative to the westward moving center of the cell and the original air mass. As  $O_3$  increases, the dynamical explanation would require higher  $O_3$  concentrations outside of the cell in less dense clouds in which photochemical  $O_3$  production may take place and in which loss of  $O_3$  through reactions with NO from lightning may not have occurred. At 3.75 and 4.75 km, however, the trajectories suggest the original air mass ends up 25–30 km southeast of the balloon trajectory, resulting in the balloon sampling air closer to the center of the cell. After the last oscillation, the balloon position is on the southeastern edge of the storm, whereas for the first oscillation it was just west of the center of the cell. At these higher altitudes, it would seem that the balloon is sampling air more reflective of the cell core, which suggests  $O_3$  levels within the cell have actually increased. Thus, we argue that the changes observed by the ozonesonde cannot be explained entirely by entrainment into or detrainment from the convective cell.

Assuming the sonde-observed changes in  $O_3$  are representative of the entire cell ( $\sim 5300 \text{ km}^2$ ), an increase of  $\sim 3.0 \times 10^6$  moles of  $O_3$  between 2.5–5.1 km may be associated with this storm. Based on comparisons of the ozonesonde profile with in situ profiles from the DC-8, the total increase in  $O_3$  from 2.5–12 km may have been  $(3.1\text{--}4.2) \times 10^6$  moles, a result that suggests 70–97% of the increase occurred in the 2.5–5 km layer.

Lidar data from the DC-8 also indicated large  $O_3$  differences between the profiles taken downwind and upwind of the cell: the lidar data report  $\sim 1.3 \times 10^7$  more moles of

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



$O_3$  after the passing of the cell between 1.6–8.5 km, with  $6.3 \times 10^6$  of those between 2.5 – 5.1 km and  $8.6 \times 10^6$  between 1.6–6.0 km. These differences could be due to spatial gradients in  $O_3$ , but the horizontal scale over which those gradients appear is  $\sim 550$  km, much greater than the  $\sim 30$  km horizontal distance (measured relative to the convective cell motion) over which the ozonesonde detected similarly large changes.

Furthermore, if we reduce our ozonesonde  $LpcO_3$  estimate by 30% to account for changes in  $O_3$  that can be attributed to descending air with higher  $O_3$ , we find  $\sim (2.1\text{--}2.5) \times 10^6$  moles of  $O_3$  created, an estimate which agrees within a factor of two of our best estimates from the WWLLN data ( $\sim 1.2 \times 10^6$  moles), from the NPOL data ( $\sim 1.0 \times 10^6$  moles), and from the OMI  $LNO_2$  data ( $\sim 1.7 \times 10^6$  moles, provided we scale by the larger area considered with the WWLLN data rather than the convective core area indicated by the NPOL data), although the estimates from the WWLLN and NPOL data carry large uncertainties due to the large uncertainties associated with the parameters included in the photochemical production calculations.

While the precise mechanism responsible for the  $O_3$  changes observed by the oscillating ozonesonde remains uncertain, several possibilities are suggested by the data. (1) If lightning and photochemistry are responsible for the changes in  $O_3$ , our ozonesonde observations are consistent with the hypothesis that shortly after the lightning strikes, NO reacts with  $O_3$  forming  $NO_2$  and leading to  $O_3$  loss within the clouds. This hypothesis is supported by the modeling studies of Ott et al. (2007) and Salzmann et al. (2008), the thunderstorm observations in Ridley et al. (2006), and the laboratory data of Franzblau (1991). By the time the convective cell reaches the Panama coast, the lightning has subsided (as suggested by the WWLLN data in Table 2) and the clouds have begun to dissipate. Ozone within the clouds, therefore, may be relatively depleted compared to its pre-storm values, and our balloon measurements may represent measurements at various stages of recovery as  $NO_x$  photochemistry begins to favor production of  $O_3$ . (2) New  $O_3$  may have been produced within the cloud, as suggested by Winterrath et al. (1999), who report a 62% increase in  $O_3$  within thunderstorm clouds; Shlanta and Moore (1972), who found  $O_3$  levels 2.6 times higher

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



at 6 km inside a cloud than pre-storm readings at the surface; and Clark and Griffing (1985), who reported 250% increases downwind of thunderstorms near Baltimore in 1980. If the dissipating cell was still producing lightning after launch, as suggested by the NPOL hourly flash estimates (Table 3), it is possible that direct production of O<sub>3</sub> from lightning occurred within the cloud, as suggested by Minschwaner et al. (2008). However, the absence of detected lightning strokes in the WWLLN record as well as the rapid dissipation of the cell as it came over land argue against the importance of direct production in explaining the changes observed by the sonde. (3) Mixing of air lower in ozone from near the surface lofted by the convection and air higher in O<sub>3</sub> from above sinking near the edge of the convection. (4) Some combination of these processes. Future measurements campaigns and modeling studies will be required to identify the relative importance of these processes. Nevertheless, our ozonesonde dataset provides the modeling community with new and important constraints that can be applied to future studies of O<sub>3</sub> production associated with tropical convective cells.

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**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures



Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## References

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### Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



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**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



**Table 1.** Summary of previous studies of O<sub>3</sub> production.

Observation	Location	Instrument/model	Citation
Convection enhances O <sub>3</sub> <6.5 km at 5° N (Fig. 1)	Global analysis	Global 3-D model	Lelieveld and Crutzen (1994)
Most LNO <sub>x</sub> is produced below 5 km by CG flashes	Global analysis	International Satellite Cloud Climatology Project	Price et al. (1997)
+12% of tropospheric O <sub>3</sub> due to convection	Global analysis	MATCH-MPIC model	Lawrence et al. (2003)
10–20% of tropospheric O <sub>3</sub> from lightning at 2–6 km; 20–30% at 6–12 km at ~10° N	Global analysis	MOZART model	Zhang et al. (2003)
In the tropics, convection reduces O <sub>3</sub> by 1–3 ppbv <1.5 km and increases O <sub>3</sub> by 1–5 ppbv from 1.5–5.5 km (Fig. 3), but globally, convection reduces tropospheric O <sub>3</sub> by 13% 80% of O <sub>3</sub> above background due to LNO <sub>x</sub> ; P(O <sub>3</sub> )=+3–4 ppbv/day; 16–24 ppbv O <sub>3</sub> in UT	Global analysis	STOCHEM-HadAM3	Doherty et al. (2005)
+10–15 DU of O <sub>3</sub> downwind of convection and biomass burning	Southeast USA and Texas	Ozonesondes (IONS-04)+Flexpart model	Cooper et al. (2006)
LNO <sub>x</sub> leads to +25–30 ppbv of O <sub>3</sub> at 10 km	Tropical Atlantic	SCIAMACHY	Martin et al. (2007)
10% of tropospheric O <sub>3</sub> from lightning	Southeast USA	Ozonesondes (IONS-06)+Flexpart model	Cooper et al. (2007)
O <sub>3</sub> aloft=2–3×pre-storm surface O <sub>3</sub>	North America	Ozone sonde (IONS-04)+CTM model	Pfister et al. (2008)
+300 ppb of O <sub>3</sub> near 1 km in thunderstorm	New Mexico	Ozone sonde	Schlanta and Moore (1972)
Peak O <sub>3</sub> detected at 5 km, secondary peak at 10 km	Maryland	Aircraft	Clark and Griffing (1985)
The highest O <sub>3</sub> production potential found below 5 km following deep convection; dissipating cell 7–17 ppbv/day at 1–6 km (Fig. 14)	Midwest USA	Aircraft (PRE-STORM project)	Dickerson et al. (1987)
7–8 ppbv/day of O <sub>3</sub> produced downwind of thunderstorms at 8–12 km	South Central USA	Aircraft (PRE-STORM)+1-D photochemical model	Pickering et al. (1990)
+38 ppbv O <sub>3</sub> in convective cloud	Brazil	DC-8, Ozonesondes+GCE model	Pickering et al. (1996) and Thompson et al. (1997)
+2 ppbv of O <sub>3</sub> during storm, +10 ppbv/24 h in UT (5–10 km)	France	DOAS+GOMETRAN model	Winterrath et al. (1999)
6–8 ppbv lower O <sub>3</sub> in convective cell	USA high plains	STERAO-A+CSCTM and GCE models	DeCaria et al. (2005)
9 ppbv lower O <sub>3</sub> during 3h of storm, but +5 ppbv/day downwind at 5.5 km	Tropical Pacific	Aircraft (PEM-Tropics B)	Ridley et al. (2006)
+7 DU of O <sub>3</sub> ; 80% O <sub>3</sub> from storms at 2.5–10 km	Europe	European Lightning Nitrogen Oxides Project+GCE model	Ott et al. (2007)
Direct production of 0.2–2.0×10 <sup>27</sup> molecules O <sub>3</sub> per flash; mechanism produces ~21% as much O <sub>3</sub> as photochemistry	Tropical Atlantic	Ozonesondes+LIS+RegCM3 model	Jenkins et al. (2008)
Maximum O <sub>3</sub> loss at 5 km during storm	New Mexico	Ozone sonde	Minschwaner et al. (2008)
	Western Pacific	CSRMC model for TOGA COARE/CEPEX	Salzmann et al. (2008)

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

⏪ ⏩

◀ ▶

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

**Table 2.** Flashes detected by the World Wide Lightning Location Network (WWLLN Flashes), estimated flashes using the WWLLN detected flashes and the network flash detection efficiency (WWLLN Flashes\*), and flashes calculated from radar heights estimated by the NPOL radar (uncertainties in parenthesis) near the Gulf of Panama on 5 August 2007. WWLLN Flashes 2 refer to flashes detected in the region upwind of the DC-8 spiral on 5 August 2007. See text for details of each.

Hours (UTC)	WWLLN flashes	WWLLN flashes*	NPOL flashes	WWLLN flashes 2	WWLLN flashes 2*
00:00–08:00	9	41 (15)	0	103	470 (170)
08:00–09:00	4	18.2 (6.6)	0	51	230 (84)
09:00–10:00	36	163 (60)	9 (10)	26	118 (43)
10:00–11:00	61	277 (101)	0	9	41 (15)
11:00–12:00	34	155 (56)	44 (72)	13	59 (21)
12:00–13:00	174	791 (288)	240 (61)	13	59 (21)
13:00–14:00	160	727 (264)	450 (240)	0	0
14:00–15:00	92	418 (152)	670 (110)	0	0
15:00–16:00	6	27 (10)	600 (100)	0	0
16:00–17:00	0	0 (0)	240 (78)	0	0
17:00–18:00	1	4.5 (1.7)	11 (13)	0	0
18:00–19:00	2	9.1 (3.3)	1.7 (2.1)	5	22.7 (8.3)
19:00–20:00	9	41 (15)	4.3 (6.3)	0	0
20:00–24:00	150	682 (248)	108 (63)	34	155 (56)
<b>Total 09:00–17:00</b>	<b>563</b>	<b>2560 (440)</b>	<b>2300 (300)</b>	<b>61</b>	<b>277 (55)</b>
<b>TOTAL</b>	<b>738</b>	<b>3400 (500)</b>	<b>2400 (310)</b>	<b>254</b>	<b>1160 (210)</b>

Title Page

Abstract Introduction

Conclusions References

Tables Figures

◀ ▶

◀ ▶

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

**Table 3.** Balloon trajectory and estimated air mass trajectories at three different altitudes within the 2.5–5.1 km layer in which the balloon oscillated. “No data” are reported when the balloon did not oscillate through the level of the air mass trajectory calculation. The “Direction” column indicates the compass heading from the balloon location to the estimated air mass location. Color-coding matches that for the balloon data shown in Figs. 12–15. See text for details.

Trajectory Altitude (km)	Time from launch (s)	Balloon lat (deg)	Air mass lat (deg)	Balloon lon (deg)	Air mass lon (deg)	Separation (km)	Direction (deg)
2.75 km	614	7.734	7.734	-80.269	-80.269	0.000	0
	2555	7.677	7.661	-80.337	-80.382	5.246	250
	no data	no data	no data	no data	no data	no data	no data
	4629	7.660	7.597	-80.418	-80.492	10.816	230
	6208	7.654	7.550	-80.463	-80.551	15.103	220
3.75 km	859	7.724	7.724	-80.277	-80.277	0.000	0
	2766	7.674	7.683	-80.346	-80.262	9.320	84
	3459	7.671	7.657	-80.374	-80.260	12.672	97
	5010	7.654	7.605	-80.431	-80.261	19.494	106
	6535	7.655	7.565	-80.469	-80.235	27.614	111
4.75 km	1134	7.722	7.722	-80.284	-80.284	0.000	0
	no data	no data	no data	no data	no data	no data	no data
	3956	7.667	7.635	-80.393	-80.307	10.116	111
	5335	7.658	7.582	-80.440	-80.300	17.591	118
	6806	7.659	7.544	-80.476	-80.269	26.130	119

[Title Page](#)
[Abstract](#)
[Introduction](#)
[Conclusions](#)
[References](#)
[Tables](#)
[Figures](#)
[Back](#)
[Close](#)
[Full Screen / Esc](#)
[Printer-friendly Version](#)
[Interactive Discussion](#)


**Table 4.** Measurements of NO<sub>x</sub> >1 ppbv, >50 km away from Alajuela, Costa Rica airport (10.0° N, 84.2° W) from DC-8 flights during TC4. Enhanced NO<sub>x</sub> <8 km are highlighted orange over Columbia, blue over open water, and yellow over the west coast of Costa Rica.

Date	Altitude (km)	Latitudes	Longitudes	NO (ppbv)
070721	10.7	4.3 <sup>0</sup> N	76.1 <sup>0</sup> W	1.0 – 1.1
070721	4.6 – 6.4	4.0 <sup>0</sup> – 4.4 <sup>0</sup> N	73.4 <sup>0</sup> – 73.9 <sup>0</sup> W	1.0 – 1.8
070721	0.5 – 0.6	4.2 <sup>0</sup> N	73.6 <sup>0</sup> W	1.3 – 2.5
070721	10.9	10.3 <sup>0</sup> – 11.5 <sup>0</sup> N	75.6 <sup>0</sup> – 77.9 <sup>0</sup> W	1.0 – 2.1
070721	3.7 – 6.2	12.0 <sup>0</sup> – 12.2 <sup>0</sup> N	80.2 <sup>0</sup> – 80.3 <sup>0</sup> W	1.0 – 3.1
070721	2.9 – 5.8	9.5 <sup>0</sup> – 9.9 <sup>0</sup> N	84.3 <sup>0</sup> – 84.6 <sup>0</sup> W	1.0 – 4.8
070722	11.0 – 12.2	12.2 <sup>0</sup> – 15.1 <sup>0</sup> N	78.4 <sup>0</sup> – 79.1 <sup>0</sup> W	1.0 – 1.6
070724	10.6	5.3 <sup>0</sup> N	85.1 <sup>0</sup> – 85.4 <sup>0</sup> W	1.0 – 1.3
070724	7.9 – 8.6	8.4 <sup>0</sup> – 8.6 <sup>0</sup> N	84.6 <sup>0</sup> W	1.4 – 3.6
070728	11.6 – 11.9	13.6 <sup>0</sup> – 14.6 <sup>0</sup> N	81.8 <sup>0</sup> – 82.8 <sup>0</sup> W	1.0 – 2.2
070728	10.4 – 11.9	8.7 <sup>0</sup> – 9.7 <sup>0</sup> N	86.4 <sup>0</sup> – 87.3 <sup>0</sup> W	1.0 – 2.0
070729	6.3 – 7.4	8.8 <sup>0</sup> – 9.2 <sup>0</sup> N	84.4 <sup>0</sup> W	1.0 – 1.5
070729	9.7 – 10.7	1.2 <sup>0</sup> – 7.5 <sup>0</sup> N	84.3 <sup>0</sup> – 84.9 <sup>0</sup> W	1.0 – 2.2
070729	11.0 – 12.1	6.5 <sup>0</sup> – 9.9 <sup>0</sup> N	78.8 <sup>0</sup> – 81.8 <sup>0</sup> W	1.0 – 2.1
070803	9.4 – 12.2	7.7 <sup>0</sup> – 9.3 <sup>0</sup> N	80.1 <sup>0</sup> – 83.8 <sup>0</sup> W	1.0 – 2.3
070805	11.0 – 11.5	6.0 <sup>0</sup> – 8.9 <sup>0</sup> N	78.0 <sup>0</sup> – 82.0 <sup>0</sup> W	1.0 – 1.3
070806	9.8 – 10.3	2.3 <sup>0</sup> – 6.1 <sup>0</sup> N	88.8 <sup>0</sup> – 92.2 <sup>0</sup> W	1.2 – 2.2
070808	9.4 – 12.2	5.8 <sup>0</sup> – 8.9 <sup>0</sup> N	81.2 <sup>0</sup> – 84.2 <sup>0</sup> W	1.0 – 2.2
070810	8.9 – 11.0	11.8 <sup>0</sup> – 16.3 <sup>0</sup> N	85.5 <sup>0</sup> – 87.5 <sup>0</sup> W	1.0 – 1.8

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

◀ ▶

◀ ▶

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion



## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

**Table 5.** Measurements of  $\text{NO}_2 > 1$  ppbv,  $> 50$  km away from Alajuela, Costa Rica airport ( $10.0^\circ \text{N}$ ,  $84.2^\circ \text{W}$ ) from DC-8 flights during TC4. Enhanced  $\text{NO} < 8$  km are highlighted orange over Columbia and blue over open water.

Date	Altitude (km)	Latitudes	Longitudes	$\text{NO}_2$ (ppbv)
070721	10.4 – 10.7	$2.5 - 6.3^\circ \text{N}$	$76.1 - 77.5^\circ \text{W}$	1.0 – 1.3
070721	0.5 – 0.6	$4.2^\circ \text{N}$	$73.6^\circ \text{W}$	1.0 – 1.8
070722	0.3	8.3	$84.8^\circ \text{W}$	7.2 – 14.2
070722	10.5 – 11.6	$6.1^\circ - 7.9^\circ \text{N}$	$80.7^\circ - 84.2^\circ \text{W}$	1.0 – 2.3
070722	12.2	$11.3^\circ - 12.6^\circ \text{N}$	$78.2^\circ - 84.2^\circ \text{W}$	1.0 – 1.3
070724	0.3	$5.4^\circ \text{N}$	$85.8^\circ \text{W}$	1.3 – 2.5
070729	5.5 – 7.4	$8.8^\circ - 9.4^\circ \text{N}$	$84.4^\circ \text{W}$	1.0 – 1.8

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

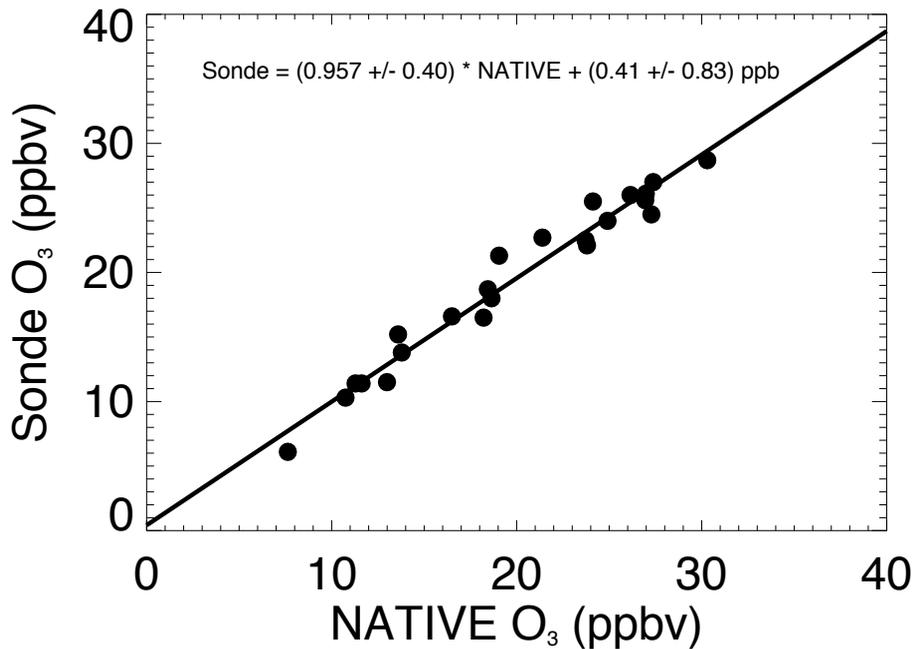
Full Screen / Esc

Printer-friendly Version

Interactive Discussion



# Las Tablas Surface O<sub>3</sub>



**Fig. 1.** A comparison between the ozonesonde readings and NATIVE surface O<sub>3</sub> measurements at the time of launch. Agreement is good to within ~5% during the TC-4 campaign.

## Ozone production in a dissipating tropical convective cell

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

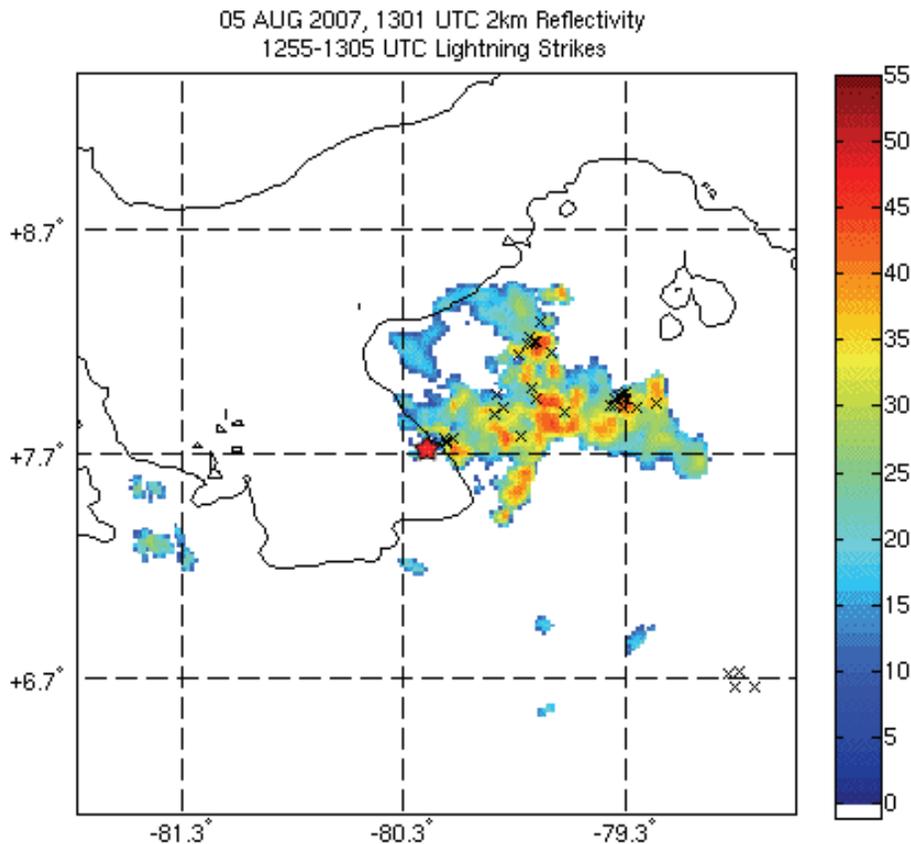
Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion





**Fig. 2.** Low-level (0.7 elevation) PPI images of convection observed east of NPOL at 13:11 UTC 5 August 2007. The color scale has units of radar reflectivity (dBZ). Each “x” marks the location of a detected WWLLN lightning strike. The red star marks the location of Las Tablas.

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

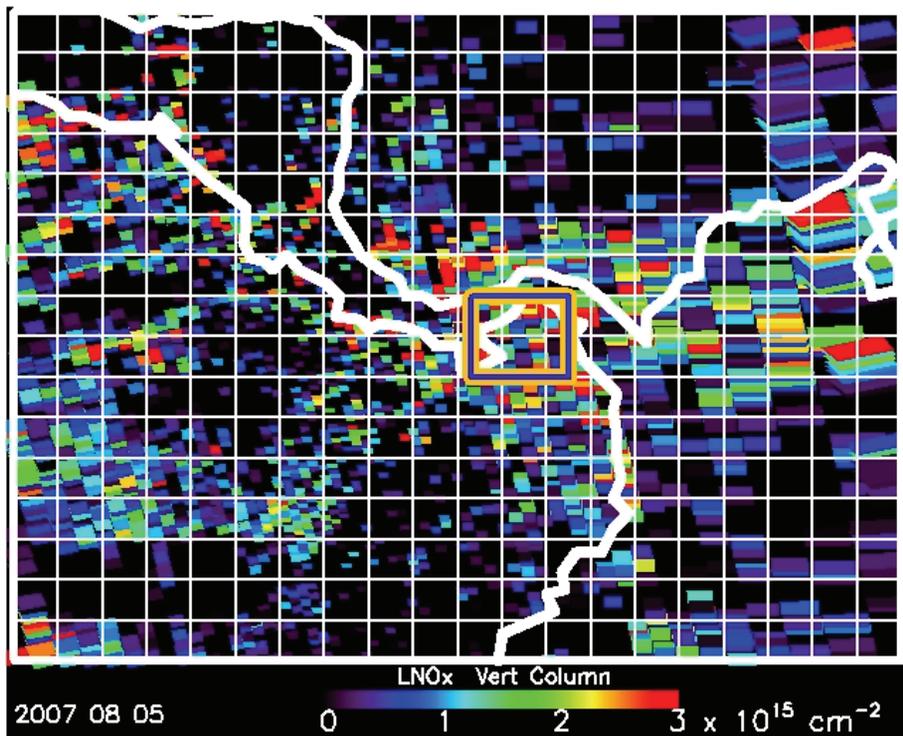
Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion





**Fig. 3.** Gridded ( $1^\circ \times 1^\circ$ ) OMI lightning  $\text{NO}_x$  in the Panama area on 5 August 2007. **(a)** Shows the larger region used for the first estimate, while **(b)** shows the smaller region in the Gulf of Panama used for the second estimate (see text).

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

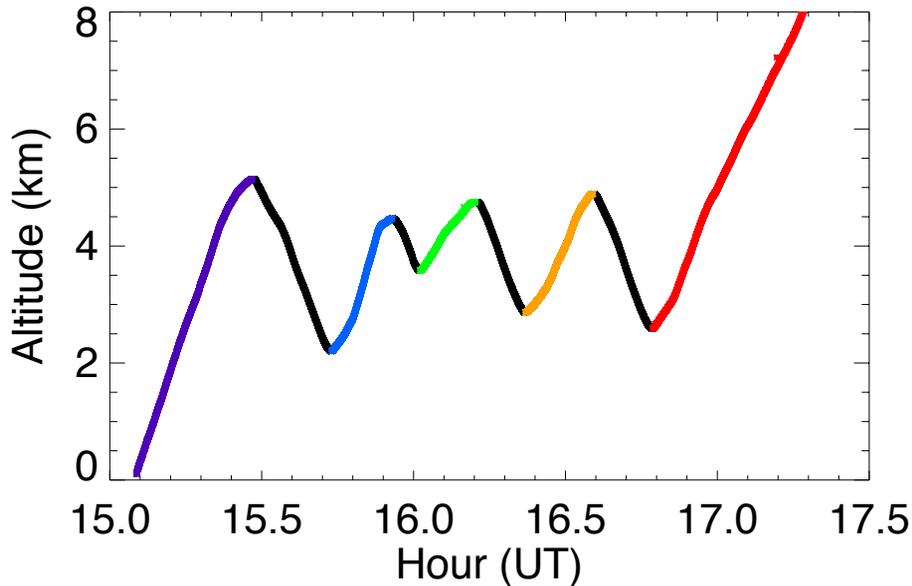
Full Screen / Esc

Printer-friendly Version

Interactive Discussion

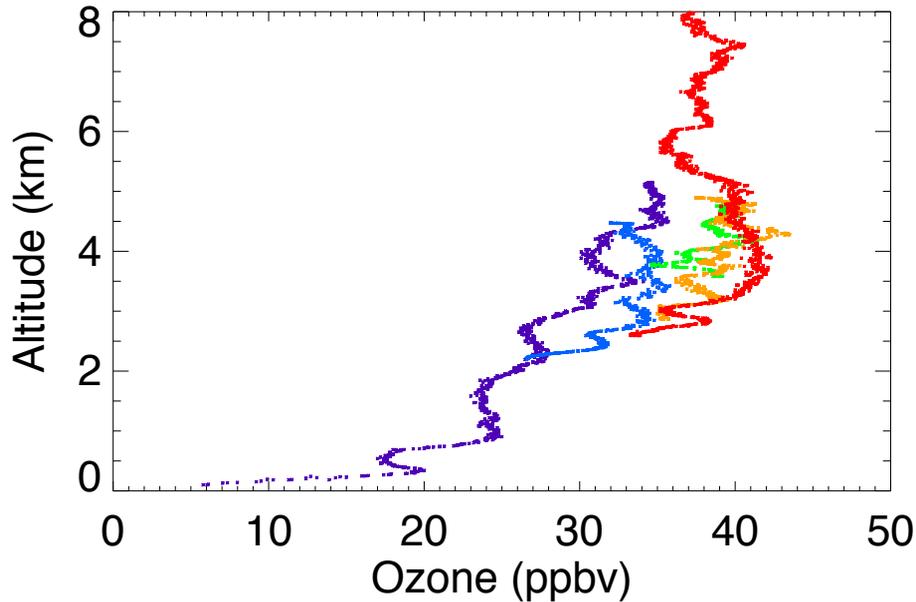


## Panama - 5 Aug 2007



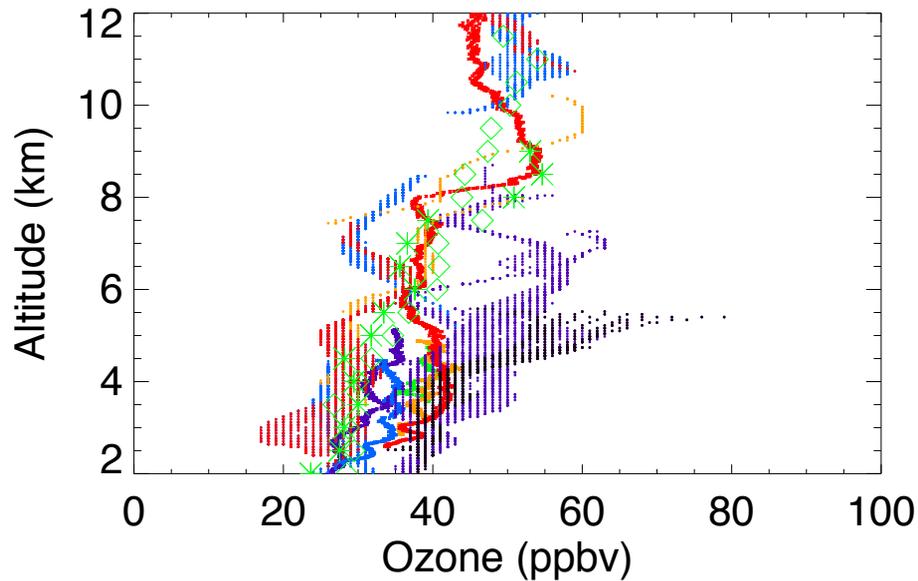
**Fig. 4.** The altitude vs. time of the ozonesonde flight on 5 August 2007 from Las Tablas, Panama shows the balloon oscillating 5 times between  $\sim 2.5$  and  $\sim 5.0$  km. The color coding of each ascent will be used in successive plots to identify changes with time of other measured parameters.

## Panama - 5 Aug 2007



**Fig. 5.** Ozone vs. altitude on the ascents as the balloon oscillated on the 5 August 2007 flight, with color coding to match the ascents identified in Fig. 4.

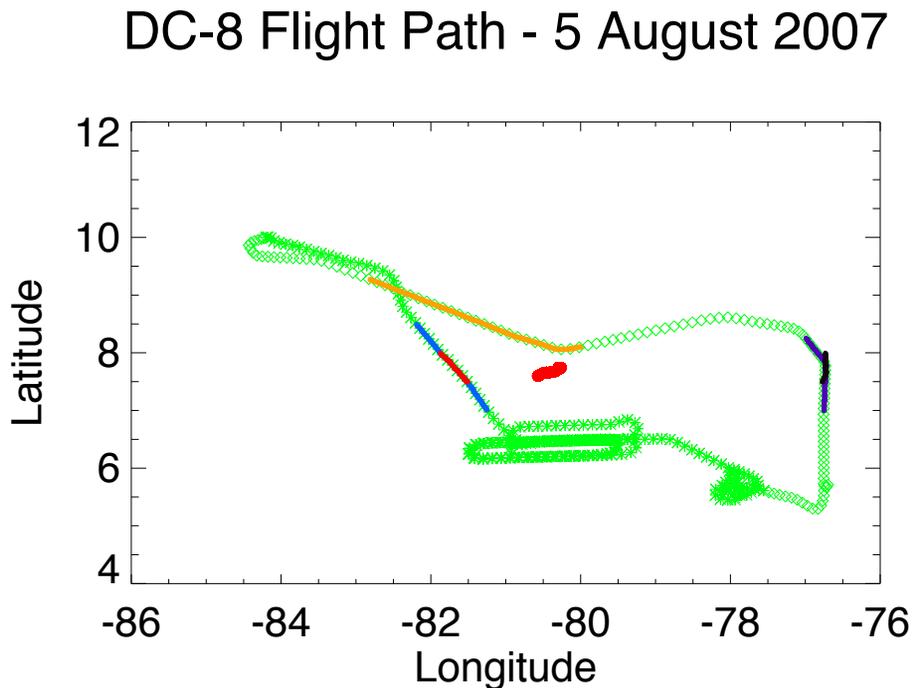
## Panama - 5 Aug 2007



**Fig. 6.** Ozone profiles on 5 August 2007 in the Gulf of Panama region from sonde (thick colors), DC-8 DIAL (colored dots), and mean DC-8 in situ (green diamonds before 17:17 UTC and green stars after 17:17 UTC). See text for details.

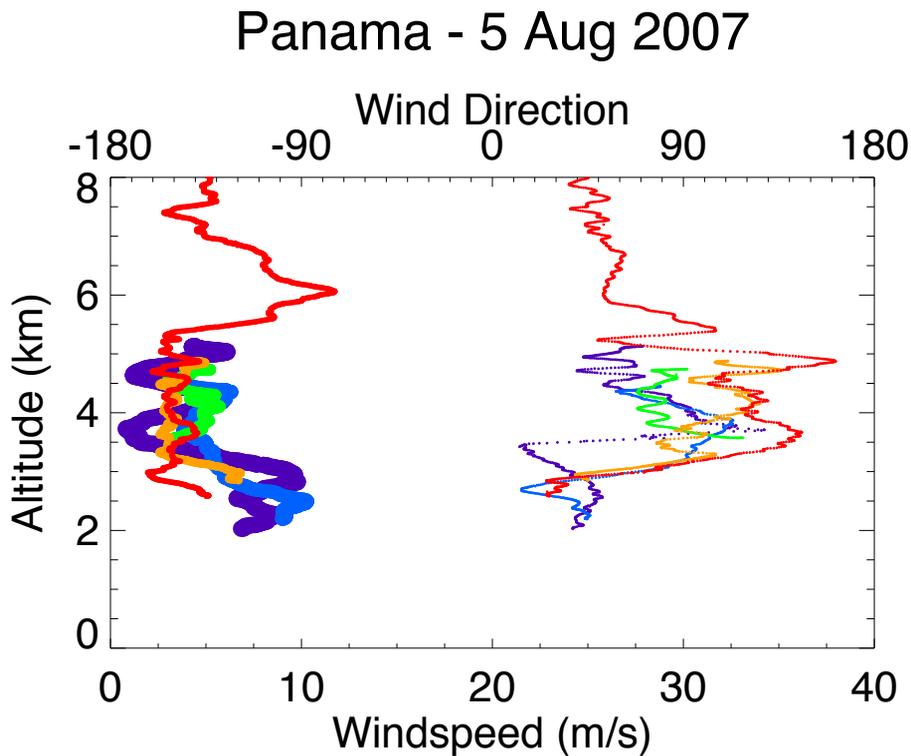
**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.



**Fig. 7.** DC-8 flight path with green stars marking the flight path prior to 17:17 UTC and the green diamonds marking the flight path after 17:17 UTC. The colored dots superimposed on the flight path correspond to DC-8 DIAL measurement locations, with the color-coding to match the profiles seen in Fig. 6.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)



**Fig. 8.** Wind speed (thick) and direction (thin) as determined from the ozonesonde GPS data. The balloon moved through a region of vertical shear, resulting in the separation of the balloon trajectory from the trajectories of the air masses the sonde sampled. See text and Table 3 for further details.

**Ozone production in a dissipating tropical convective cell**

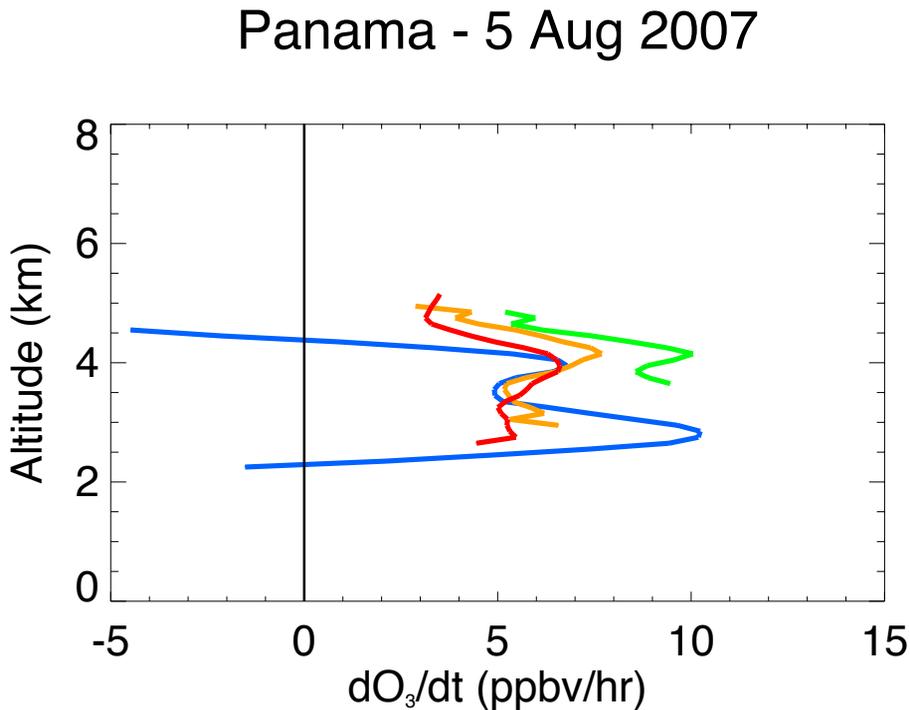
G. A. Morris et al.

<a href="#">Title Page</a>	
<a href="#">Abstract</a>	<a href="#">Introduction</a>
<a href="#">Conclusions</a>	<a href="#">References</a>
<a href="#">Tables</a>	<a href="#">Figures</a>
<a href="#">◀</a>	<a href="#">▶</a>
<a href="#">◀</a>	<a href="#">▶</a>
<a href="#">Back</a>	<a href="#">Close</a>
<a href="#">Full Screen / Esc</a>	
<a href="#">Printer-friendly Version</a>	
<a href="#">Interactive Discussion</a>	



**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.



**Fig. 9.** Calculated  $dO_3/dt$  vs. altitude for the ascents of the flight on the 5 August 2007 flight, with color coding to match the ascents identified in Fig. 4. See text for details.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

◀ ▶

◀ ▶

Back Close

Full Screen / Esc

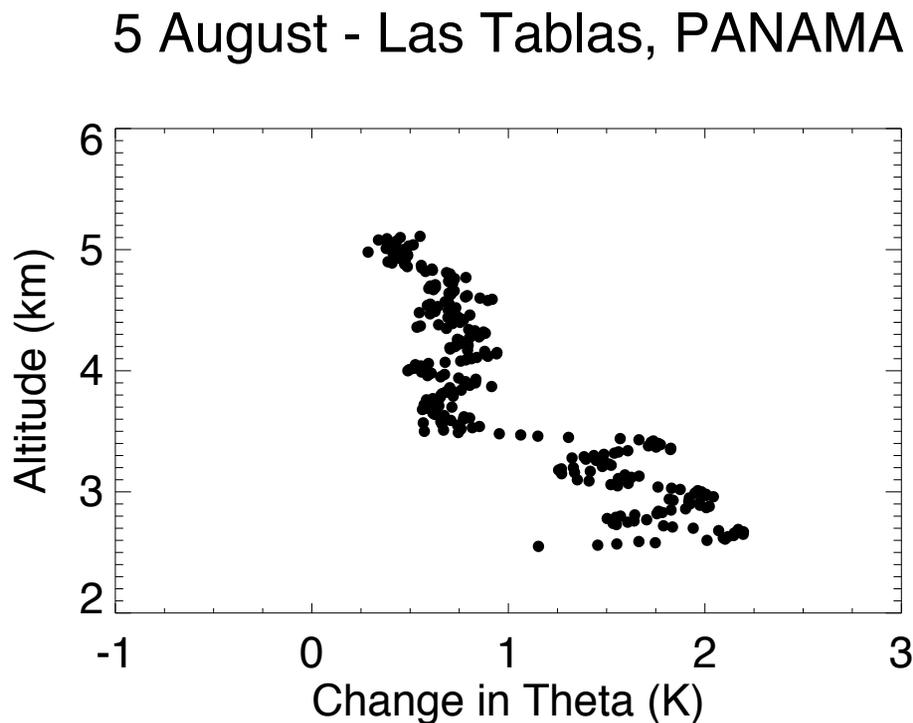
Printer-friendly Version

Interactive Discussion



**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

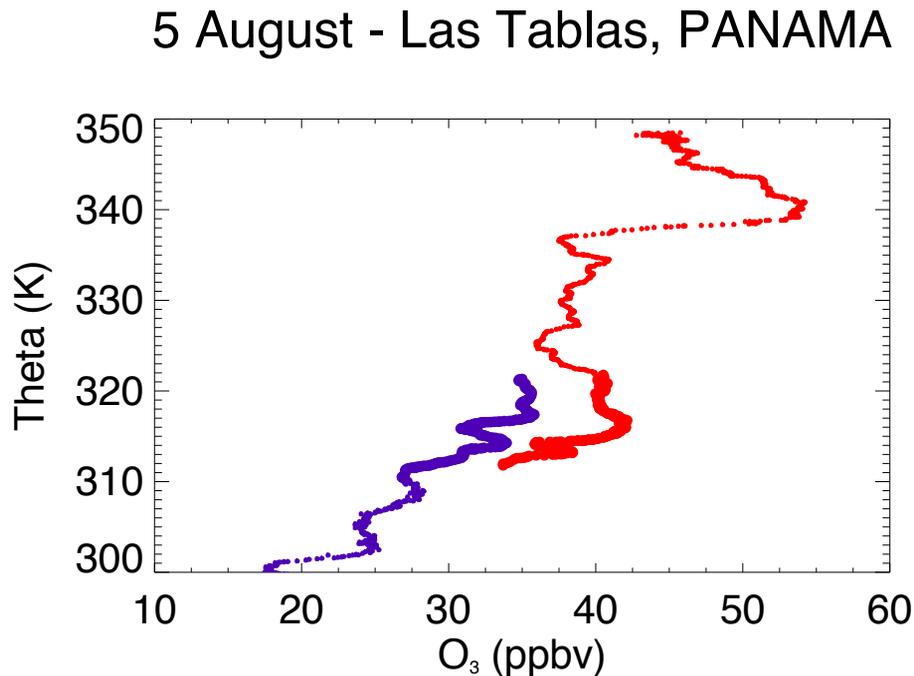


**Fig. 10.** The change in potential temperature from the first to the last ascent vs. altitude for the flight of 5 August 2007.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

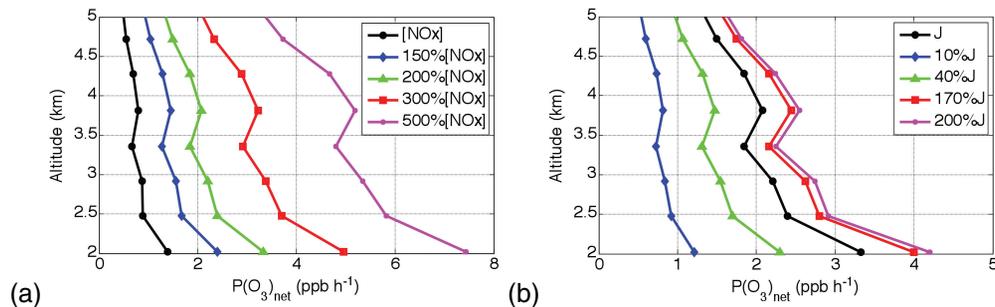


**Fig. 11.** Ozone vs. potential temperature for the first and last ascents of the flight of 5 August 2007 shows that the changes cannot be due to descent of the air mass alone.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.



**Fig. 12.** Calculated ozone net production  $P_{\text{net}}(\text{O}_3)$  within the altitude of 2–5 km based on the modeling results from the aircraft data at 16:45–17:00 UTC on 5 August 2007 with initial and controlled **(a)**  $\text{NO}_x$  concentrations and **(b)** photolysis rates ( $J$  values) with 200% $[\text{NO}_x]$  case.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

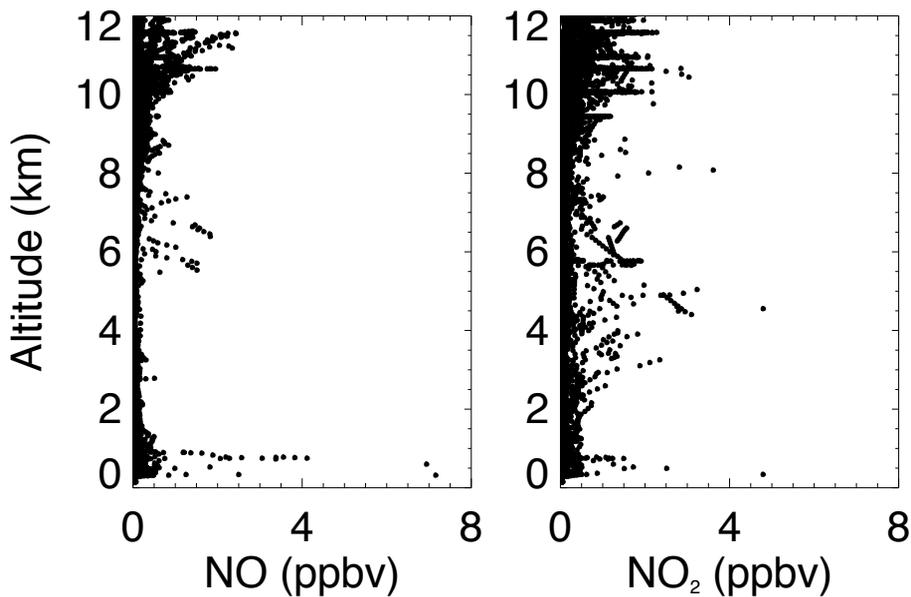
Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion





**Fig. 13.** NO and NO<sub>2</sub> measurements aboard the DC-8 aircraft during TC4. All measurements taken at least 10 min after take-off and 10 min before landing at a distance of >50 km from the airport in Alajuela, Costa Rica airport (10.0° N, 84.2° W) are shown.

**Ozone production in a dissipating tropical convective cell**

G. A. Morris et al.

Title Page

Abstract Introduction

Conclusions References

Tables Figures

◀ ▶

◀ ▶

Back Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

