



Supplement of

Dust storms from the Taklamakan Desert significantly darken snow surface on surrounding mountains

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1 S1 Method

2 S1.1 Attribution of the spatial variation in snow albedo reduction As noted above, the snow albedo reduction depends mainly on the dust content, snow 3 optical effective radius (Reff), snow depth (SD) and solar zenith angle (SZ). Snow 4 optical effective radius and snow depth can be categorized as the snow properties (SP). 5 Here we choose these three variables to discuss their fractional contributions to the 6 7 spatial variability in snow albedo reduction. $\Delta \alpha$ can be expressed as $\Delta \alpha = f(\text{dust}, \text{SP}, \text{SZ}).$ 8 (S1) 9 The spatial variability about snow albedo reduction due to dust can be described as $\Delta \alpha$ (dust)=f(dust, <u>SP</u>, <u>SZ</u>), 10 (S2) where \overline{SP} and \overline{SZ} indicate spatial-mean values of SP, and SZ. Similarly, we can 11 obtain the following equations: 12 $\Delta \alpha(SP) = f(\overline{dust}, SP, \overline{SZ}),$ 13 (S3)

14
$$\Delta \alpha(SZ) = f(\overline{dust}, \overline{SP}, SZ).$$
 (S4)

15 We fit the $\Delta \alpha$ based on multiple linear regression, we can express it as

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$$\Delta a_{\rm fit} = a \times \Delta a({\rm dust}) + b \times \Delta a({\rm SP}) + c \times \Delta a({\rm SZ}),$$
 (S5)

17 where $\Delta \alpha_{\text{fit}}$ is the fitted snow albedo reduction, and *a*, *b*, *c* represent the regression 18 coefficients. As a result, we can use $\Delta \alpha_{\text{fit}}$ to replace $\Delta \alpha$ to access the contribution to 19 spatial variation of individual variables, Eq. (14) can be written as follows:

20
$$\Delta \alpha_{\text{fit}} - \overline{\Delta \alpha_{\text{fit}}} = a \times (\Delta \alpha(\text{dust}) - \overline{\Delta \alpha(\text{dust})}) + b \times (\Delta \alpha(\text{SP}) - \overline{\Delta \alpha(\text{SP})}) + c \times (\Delta \alpha(\text{SZ}) - \overline{\Delta \alpha(\text{SZ})}), (\text{S6})$$

- 21 where we use $\Delta \alpha_{\text{fit}} \overline{\Delta \alpha_{\text{fit}}}$ represent snow albedo reduction anomaly
- 22 ($\Delta \alpha_{\text{fit}}^{\text{anomaly}}$). After, Eq. (15) can be written as

23
$$\Delta \alpha_{\rm fit}^{\rm anomaly} = a \times \Delta \alpha_{\rm anomaly} ({\rm dust}) + b \times \Delta \alpha_{\rm anomaly} ({\rm SP}) + c \times \Delta \alpha_{\rm anomaly} ({\rm SZ}).$$
(S7)

24 According to Pu et al. (2019) and Cui (2021), the fractional contribution of dust to the

spatial variability in snow albedo reduction (F_{dust}) can be written as 25

26
$$F_{\text{dust}} = \frac{1}{n} \sum_{i=1}^{n} \frac{\left(a \times \Delta \alpha_{\text{anomaly}}(\text{dust})\right)^2}{X_i},$$
 (S8)

27
$$X_{i} = (a \times \Delta \alpha_{\text{anomaly}} (\text{dust})_{i})^{2} + (b \times \Delta \alpha_{\text{anomaly}} (\text{SP})_{i})^{2} + (c \times \Delta \alpha_{\text{anomaly}} (\text{SZ})_{i})^{2}.$$
(S9)

where n represent the length of the data set. Similarly, we can get the F_{SP} and F_{SZ} . 28



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Figure S1. Variations in spectral snow albedo due to (a) dust content ($\mu g g^{-1}$), (b) snow depth (m), (c) snow grain size (µm), and (d) solar zenith angle (o) while the 31 32 other three parameters are kept constant.

33

(a) 15 May 2019



(c) 17 May 2019





(d) 18 May 2019



(e) 19 May 2019

(f) 20 May 2019





(g) 21 May 2019





- Figure S2. Satellite observations during the 15–22 May 2019 severe dust event across the Tien Shan (a-h). Satellite images (a-h) are from Terra/MODIS (https://worldview.earthdata.nasa.gov).
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Figure S3. Spatial distribution of the ratios between 15 May 2019 and 22 May 2019 40 of (a) dust, (b) albedo reduction, (c) radiative forcing and (d) Reff. The background

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400

200

C

Tien Shan

Kunlun

Qilian

44

20

0

Tien Shan

45 Figure S4. Statistics for regionally averaged (a) dust, (b) albedo reduction, (c) radiative forcing and (d) Reff for Tien Shan, Kunlun Mountains and Qilian 46 47 Mountains. The pink color shows 15 May, 2019, 23 August, 2019 and 01 November, 2012 before the dust storm in three regions. The blue color shows 22 May, 2019, 48 06 Sep, 2019 and 04 November, 2012 after the dust storm in three regions. The 49 boxes denote the 25th and 75th quantiles, and the horizontal lines represent the 50 50th quantiles (medians); the averages are shown as blue triangle; the whiskers 51 denote the 5th and 95th quantiles. 52

*

Kunlun

Qilian

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- 54



3000 3300 3600 3900 4200 4500 4800 5100 5400 5700 6000 6300 6600 6900 Elevation (m)

- 56 Figure S5. The overall lower bound and upper bound of the uncertainty value of
- 57 snow albedo reduction retrieval due to atmospheric correction in Tien Shan (a-d),
- 58 Kunlun Mountains (e-h) and Qilian Mountains (i-l).
- 59



Figure S6. Satellite observations during the 05–11 March 2020 severe dust event across the Kunlun Mountains. (a, c) Terra/MODIS satellite true-color images acquired on 05 March 2020, prior to the dust storm. (b, d) Terra/MODIS satellite images acquired on 11 March 2020, with significant snow darkening across the Kunlun Mountains after the dust storm. Satellite images (a-d) are from Terra/MODIS (https://worldview.earthdata.nasa.gov).





- Figure S8. Spatial distribution of the ratio between 23 Aug 2019 and 06 Sep 2019
 of (a) dust, (b) albedo reduction, (c) radiative forcing and (d) R_{eff}. The background
 images in (a-d) show the elevation of Kunlun Mountains.



Figure S9. Spatial distributions of the average values and differences of R_{eff} across
 the (a-c) Tien Shan, (d-f) Kunlun Mountains and (g-i) Qilian Mountains,
 respectively.



event across the Qilian Mountains (a-d). Satellite images (a-d) are from Terra/MODIS (https://worldview.earthdata.nasa.gov).





Figure S11. Spatial distribution of the ratio between 01 Nov 2012 and 04 Nov 2012 of (a) dust, (b) albedo reduction, (c) radiative forcing and (d) Reff. The background images in (a-d) show the elevation of Qilian Mountains.



- 96 Figure S12. Spatial distributions of the averaged MERRA-2 (a) BC, (b) dust and
- 97 (c) OC deposition rate from March to August 2019.
- 98







(d) 29 Sep 2015





99

100Figure S13. Satellite observations during the 29 September to 02 October 2015101severe dust event across the Tien Shan (a-f). Satellite images (a-f) are from

- 102 Terra/MODIS (https://worldview.earthdata.nasa.gov).
- 103



- 105 Figure S14. Satellite observations during the 01–06 March 2016 severe dust event
- 106across the Tien Shan (a-f). Satellite images (a-f) are from Terra/MODIS107(https://worldview.earthdata.nasa.gov).
- 108





110 Figure S15. Satellite observations during the 04–11 March 2023 severe dust event

- 111 across the Tien Shan (a-f). Satellite images (a-f) are from Terra/MODIS
- 112 (https://worldview.earthdata.nasa.gov).
- 113



- 114
- 115 Figure S16. Satellite observations during the 28 January to 03 February 2019
- severe dust event across the Kunlun Mountains (a-f). Satellite images (a-f) are
- 117 from Terra/MODIS (https://worldview.earthdata.nasa.gov).
- 118



120 Figure S17. Satellite observations during the 16–23 March 2019 severe dust event

- 121 across the Kunlun Mountains (a-f). Satellite images (a-f) are from Terra/MODIS
- 122 (https://worldview.earthdata.nasa.gov).



- 125 Figure S18. Satellite observations during the 23–27 July 2019 severe dust event
- 126 across the Kunlun Mountains (a-f). Satellite images (a-f) are from Terra/MODIS
- 127 (https://worldview.earthdata.nasa.gov).
- 128



- 130 Figure S19. Satellite observations during the 03–12 December 2014 severe dust
- event across the Qilian Mountains (a-f). Satellite images (a-f) are from
 Terra/MODIS (https://worldview.earthdata.nasa.gov).
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135 Figure S20. Satellite observations during the 25–30 December 2017 severe dust

event across the Qilian Mountains (a-f). Satellite images (a-f) are from
 Terra/MODIS (https://worldview.earthdata.nasa.gov).



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140 Figure S21. Satellite observations during the 31 January to 03 February 2019

- severe dust event across the Qilian Mountains (a-f). Satellite images (a-f) are from
- 142 Terra/MODIS (https://worldview.earthdata.nasa.gov).
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