



# Supplement of

# Influence of intense secondary aerosol formation and long-range transport on aerosol chemistry and properties in the Seoul Metropolitan Area during spring time: results from KORUS-AQ

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## 1 Tables

- 2 Table S1. Average (± 1 standard deviation), minimum and maximum concentrations of the
- 3 particulate matter  $(PM_1)$  species and the total  $PM_1$  mass over the whole campaign, and the average
- 4 contribution of each of the  $PM_1$  species to the total  $PM_1$  mass.

	Average conc.	5th	95th	Fraction	Detection limit
	$\pm$ one standard	percentile	percentile	of total	(3min) (µg m <sup>-3</sup> )
	deviation (µg m <sup>-3</sup> )	conc.	conc.	PM1 (%)	
		(µg m <sup>-3</sup> )	(µg m <sup>-3</sup> )		
Organics	$9.76 \pm 5.27$	3.12	19.3	44	0.04
Nitrate	$3.78\pm4.20$	0.33	13.4	17	0.01
Sulfate	$4.40\pm3.26$	1.10	11.0	20	0.01
Ammonium	$2.56\pm2.16$	0.53	7.46	12	0.03
Chloride	$0.04\pm0.05$	ND	0.14	0	0.01
Black carbon	$1.52\pm0.82$	0.44	2.94	7	0.1
Total PM <sub>1</sub>	$22.1 \pm 13.0$	6.79	50.1	100	0.05

1 Table S2. Correlation coefficient (Pearson's r) for the linear regressions between organic aerosol

2 (OA) factors (including the sum of primary factors (primary OA (POA) = hydrocarbon like OA

3 (HOA) + cooking OA (COA) + biomass burning OA (BBOA)), as well as the sum of the oxidized

factors (oxidized OA (OOA) = semi-volatile OOA (SV-OOA) + low volatile (LV-OOA)), and 4

5 various particle- and gas-phase species, and ions.

r	НОА	COA	POA (HOA+ COA)	SV-OOA	LV-OOA	OOA (SV-OOA+ LV-OOA)
Nitrate	0.28	0.13	0.23	0.21	0.02	0.11
Sulfate	0.05	-0.05	0.00	-0.06	-0.03	-0.04
Ammonium	0.19	0.04	0.12	0.07	0.00	0.03
Chloride	0.47	0.13	0.32	0.05	-0.10	-0.05
K (AMS)	0.50	0.68	0.52	0.03	0.55	0.70
Primary pollutants	0.00	0.00	0.70	0.72	0.00	
BC	0.54	0.38	0.52	0.42	0.22	0.34
CO	0.36	0.30	0.38	0.23	0.03	0.12
NO <sub>2</sub>	0.41	0.57	0.58	0.27	0.13	0.21
AMS tracer ions $(m/z \text{ value})$						
$CO_{2}^{+}(44)$	0.22	0.31	0.31	0.61	0.78	0.80
$C_2H_5N^+(43)$	0.10	0.18	0.17	0.64	0.58	0.68
$C_2H_4O_2^+(60)$	0.39	0.62	0.60	0.67	0.75	0.81
$C_{3}H_{5}O_{2}^{+}(73)$	0.35	0.61	0.58	0.66	0.81	0.85
$C_{3}H_{3}O^{+}(55)$	0.39	0.75	0.69	0.73	0.67	0.78
$C_{3}H_{5}O^{+}(57)$	0.37	0.61	0.59	0.79	0.75	0.87
$C_{3}H_{7}^{+}(43)$	0.87	0.80	0.97	0.43	0.17	0.31
$C_{3}H_{7}N^{+}(57)$	0	0.06	0.04	0.26	0.33	0.34
$C_4H_7^+(55)$	0.85	0.78	0.98	0.52	0.24	0.40
$C_4H_9^+(43)$	0.81	0.84	0.96	0.42	0.00	0.24
$C_5H_{11}^+(57)$	0.96	0.60	0.89	0.28	0.04	0.15
$C_5H_8O^+(84)$	0.46	0.89	0.82	0.60	0.41	0.55
$C_6H_{10}O^+(98)$	0.46	0.96	0.87	0.38	0.20	0.31
$C_7H_{12}O^+(112)$	0.40	0.70	0.66	0.44	0.30	0.41
C <sub>9</sub> H <sub>7</sub> <sup>+</sup> (115)	0.54	0.71	0.74	0.72	0.59	0.73
CHN <sup>+</sup> (27)	0.23	0.25	0.28	0.37	0.47	0.49
CN <sup>+</sup> (26)	0.15	0.29	0.26	0.11	0.08	0.11
$CH_2SO_2^+(77)$	0.20	0.09	0.16	0.06	0.04	0.05
$CH_3SO_2^+(78)$	0.21	0.09	0.17	0.05	0.03	0.04

6 7 BC, black carbon; AMS, aerosol mass spectrometer; PAH, polycyclic aromatic hydrocarbons

Value that are r > 0.7 are boldfaced

8 **Table S3.** Comparison of the average O/C, H/C, and OM/OC ratios of total OA and the four OA

9 factors identified from PMF analysis calculated using the Aiken-Ambient method (Aiken et al.,

10 2008) and the improved Canagaratna-Ambient method (Canagaratna et al., 2015).

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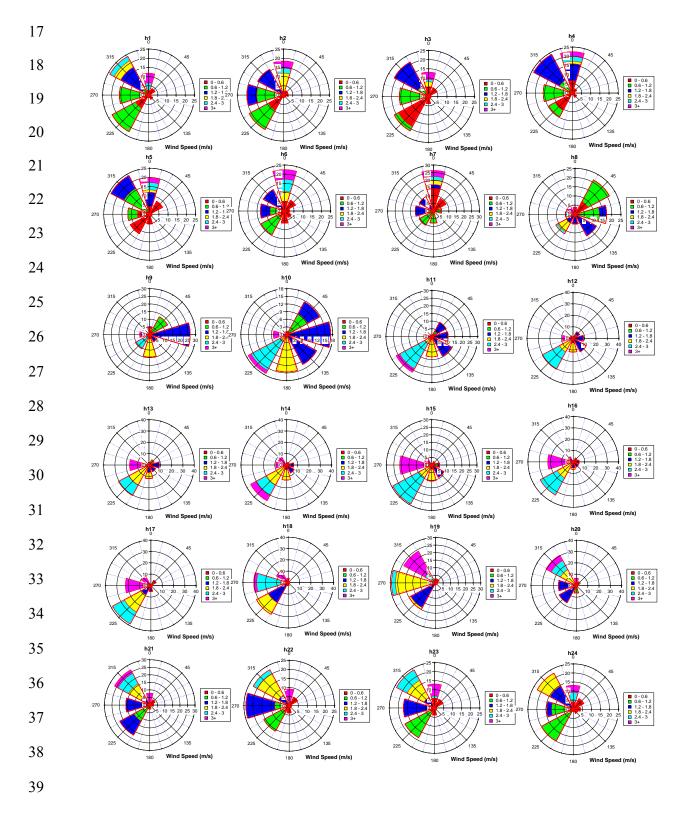
Species	Ratio	Aiken-Ambient	Canagaratna- Ambient
OA	O/C	0.38	0.49
	H/C	1.52	1.67
	OM/OC	1.66	1.82
HOA	O/C	0.11	0.15
	H/C	1.85	2.00
	OM/OC	1.31	1.37
COA	O/C	0.15	0.19
	H/C	1.68	1.83
	OM/OC	1.34	1.41
SVOOA	O/C	0.33	0.44
	H/C	1.55	1.73
	OM/OC	1.58	1.74
LVOOA	O/C	0.51	0.91
	H/C	1.41	1.46
	OM/OC	1.84	2.36

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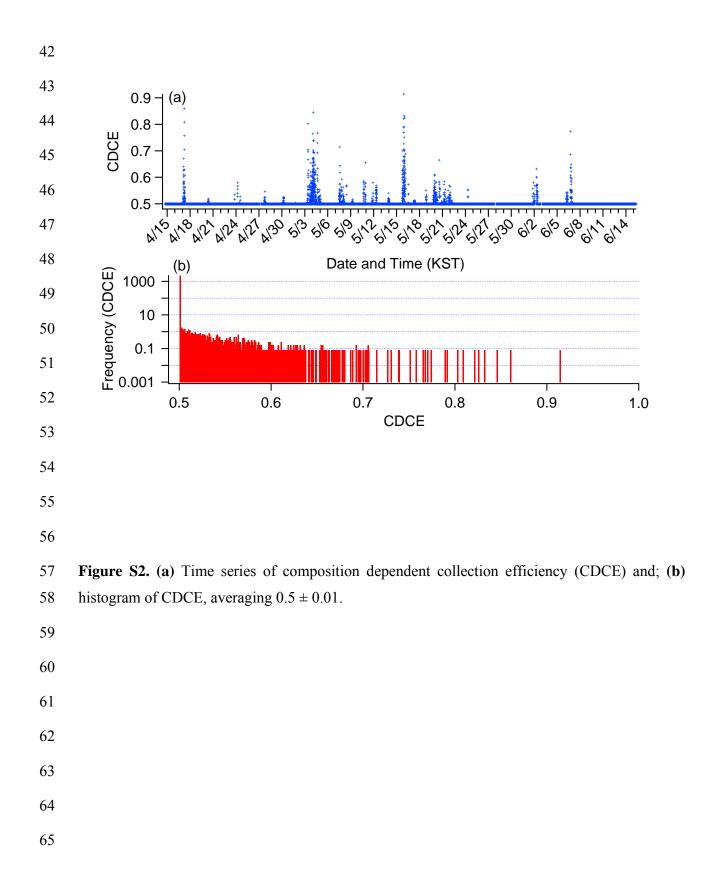
- 13 **Table S4.** Comparison of aerosol properties and meteorological parameters at different stages
- 14 during haze period.

	Overall	<b>S1</b>	<b>S2</b>	<b>S3</b>	S4
Average non-refractory submicrometer particulate matter (NR- PM <sub>1</sub> ) mass concentration (μg m <sup>-3</sup> ) (Average )	22.1	9.0	41.2	37.7	13.0
RH(%) / Temp(°C)	62/19	96/16	78/18	67/21	48/23
WS(m/s)	1.7	3.0	2.2	1.4	1.7
Trace gas conc.(CO (ppm) and /SO <sub>2</sub> / NO <sub>2</sub> O <sub>3</sub> /(ppb))	0.52/5.1/34/ 29	0.48/4.0/39/ 15	0.55/4.8/32/ 39	0.62/6.1/37/ 35	0.39/4.5/28.7 /33

15 PM, particulate matter; NR-PM<sub>1</sub>, non-refractory submicrometer particulate matter; RH, relative humidity



40 Figure S1. Wind Rose plots for every hour colored by wind speed. Radial scales correspond to the41 frequency.



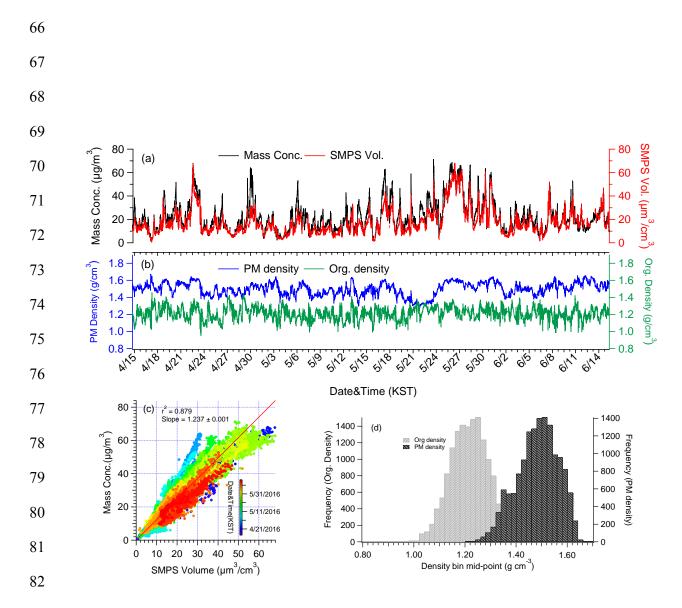
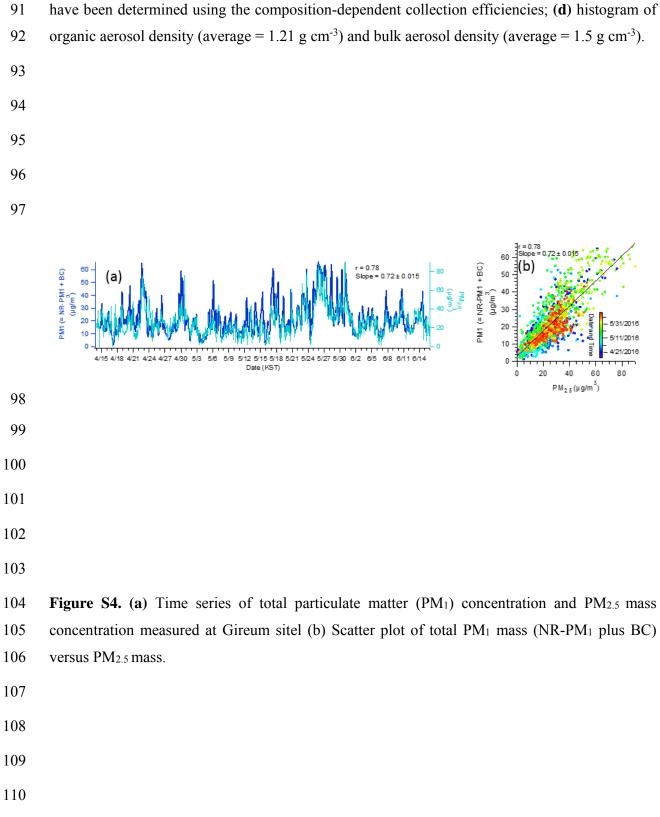


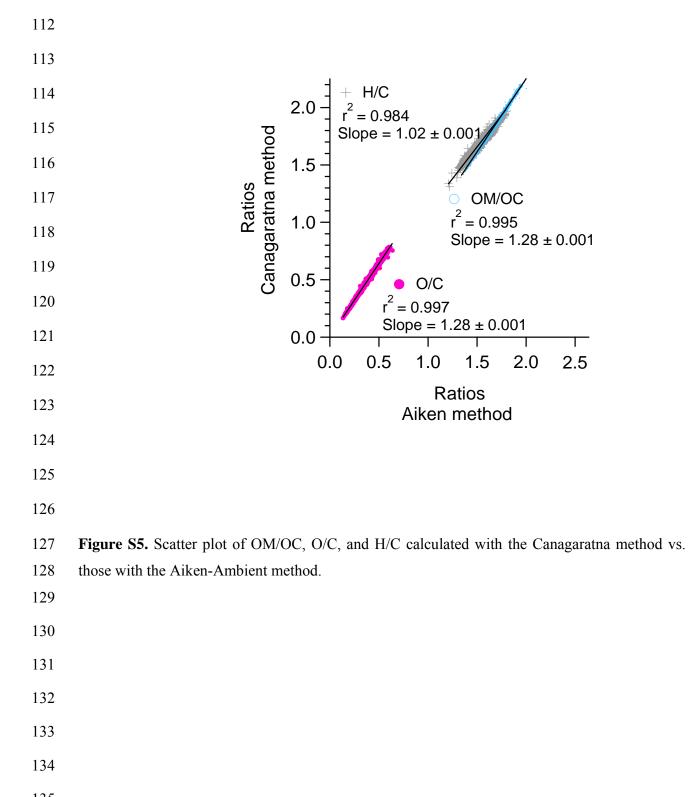
Figure S3. (a) Time series of total particulate matter (PM<sub>1</sub>) and scanning mobility particle sizer
(SMPS) volume concentrations. The correlation and ratio between PM<sub>1</sub> and PM<sub>2.5</sub> are shown ; (b)
Time series of the organic aerosol density estimated using the method reported in Kuwata et al.
(2012)

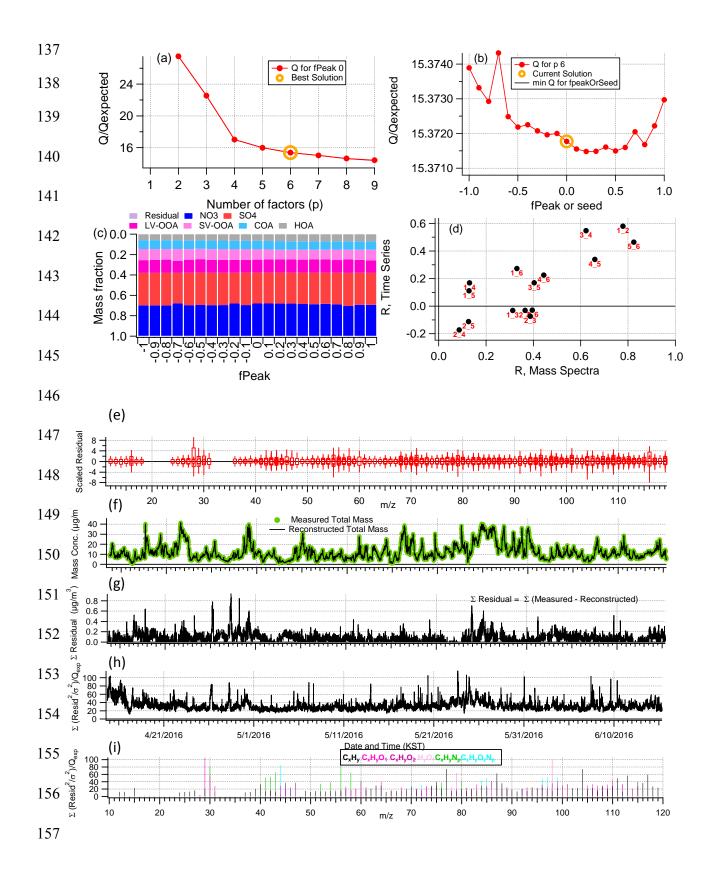
87  $\rho_{org} = [12 + 1 \cdot (H/C) + 16 \cdot (O/C)]/[7 + 5 \cdot (H/C) + 4.15 \cdot (O/C)]$ 

and bulk aerosol density estimated from the measured chemical composition, known inorganic
species density and the organic density estimated above (Zhang et al., 2005). (c) Scatter plot of the
total PM<sub>1</sub> mass (NR-PM<sub>1</sub> plus BC) versus SMPS volume, where the NR-PM<sub>1</sub> mass concentrations

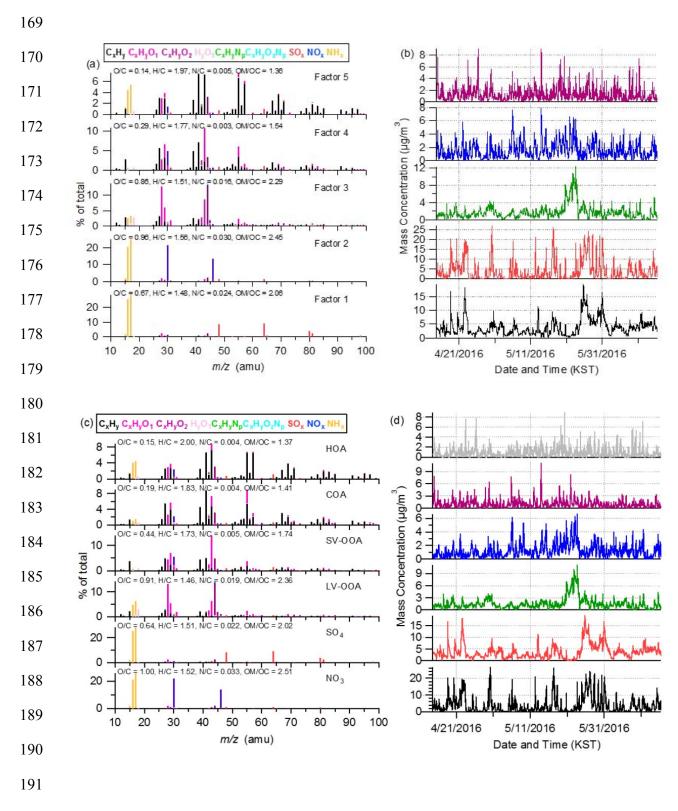


have been determined using the composition-dependent collection efficiencies; (d) histogram of

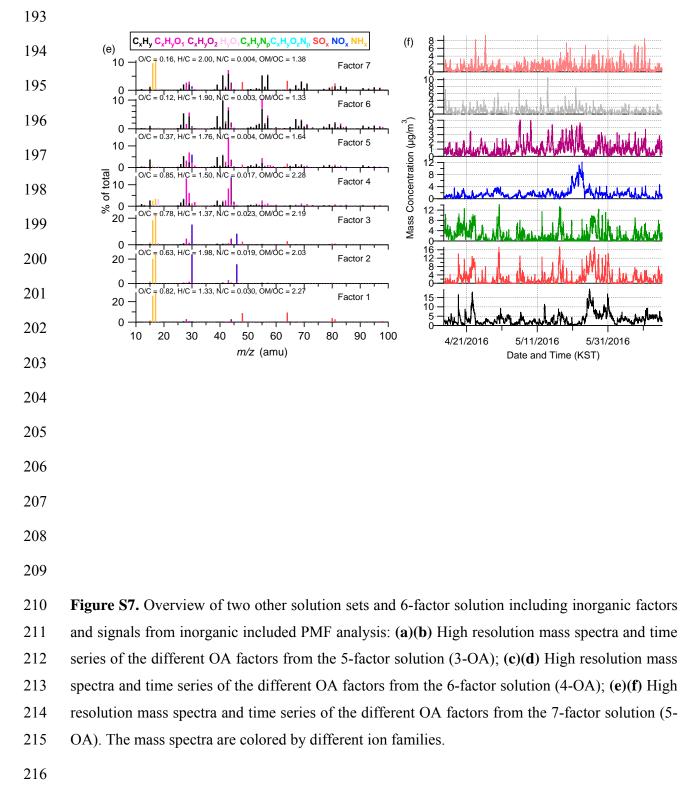




- 158 **Figure S6.** Summary of the key diagnostic plots of the chosen 6-factor solution (4 OA factor
- 159 solution) from PMF analysis of the organic aerosol fraction: (a)  $Q/Q_{exp}$  as a function of the number 160 of factors (*p*) explored in PMF analysis, with the best solution denoted by the open orange circle.
- Plots **b-i** are for the chosen solution set, containing 4 factors: (**b**)  $Q/Q_{exp}$  as a function of fPeak; (**c**)
- 162 mass fractional contribution to the total mass of each of the PMF factors, including the residual (in
- 163 purple), as a function of fPeak; (d) Pearson's r correlation coefficient values for correlations among
- 164 the time series and mass spectra of the PMF factors. Here,  $1 = NO_3$ ,  $2 = SO_4$ , 3 = LV-OOA, 4 =
- 165 SV-OOA, 5 = COA, 6 = HOA; (e) box and whiskers plot showing the distributions of scaled
- 166 residuals for each m/z; (f) time series of the measured mass and the reconstructed mass from the
- 167 sum of the 6 factors; (g) time series of the variations in the residual (= measured reconstructed)
- 168 of the fit; (h) the  $Q/Q_{exp}$  for each point in time; (i) the  $Q/Q_{exp}$  values for each fragment ion.







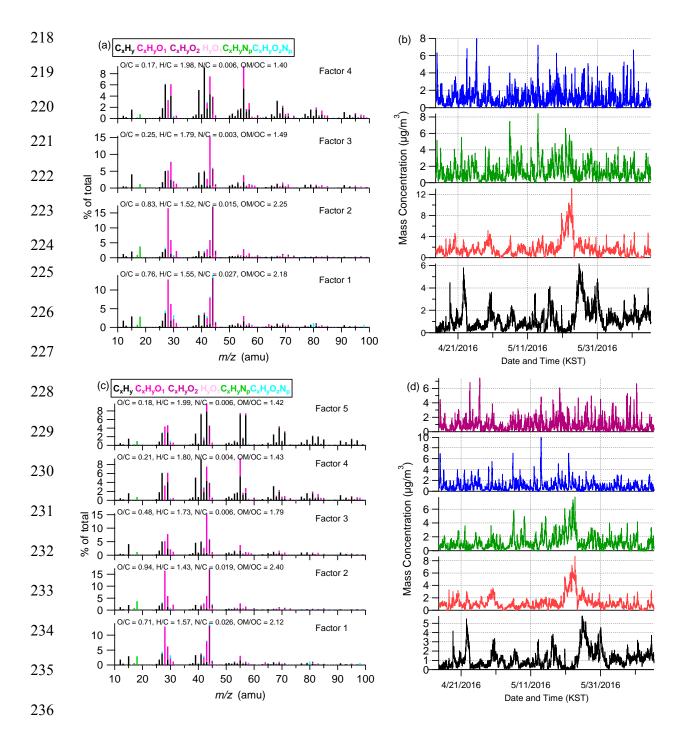


Figure S8. Overview of the 4-factor solution (a and b) and the 5-factor solution (c and d) from PMF analysis of the organic mass spectra only. The mass spectra are colored by different ion families.

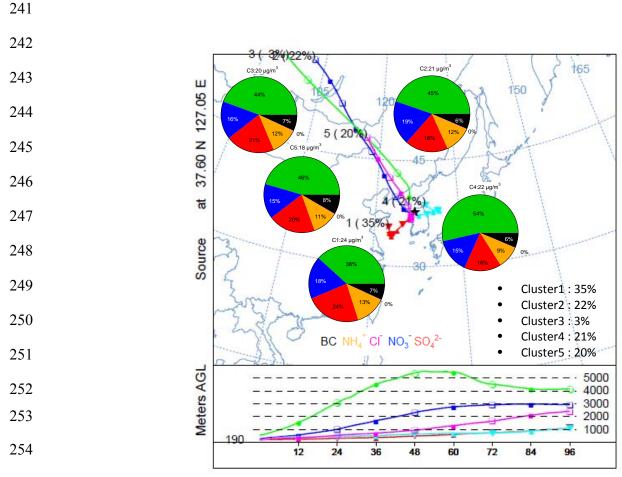
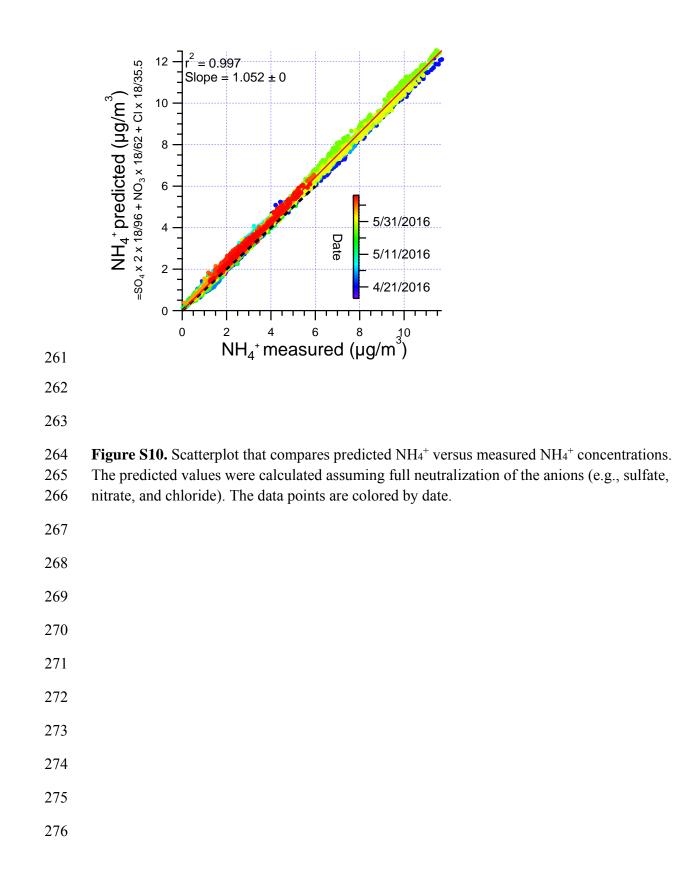


Figure S9. Averaged compositional pie chart of PM<sub>1</sub> species (non-refractory-PM<sub>1</sub> plus black carbon (BC)) in different clusters from the five cluster solution. The trajectories were released at

half of the mixing height at the KIST (latitude: 37.60N; longitude: 127.05E) and the average

arriving height for the back trajectories for this study was approximately 190 m.



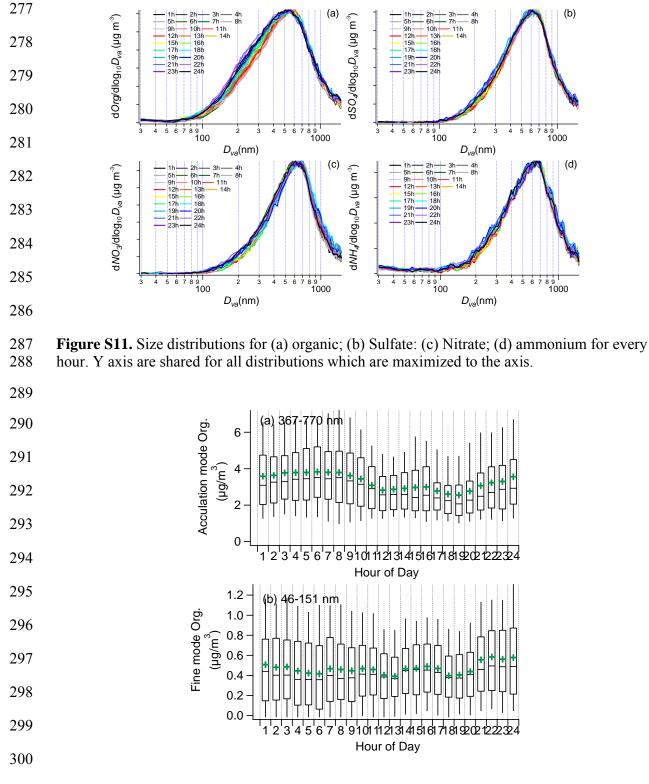
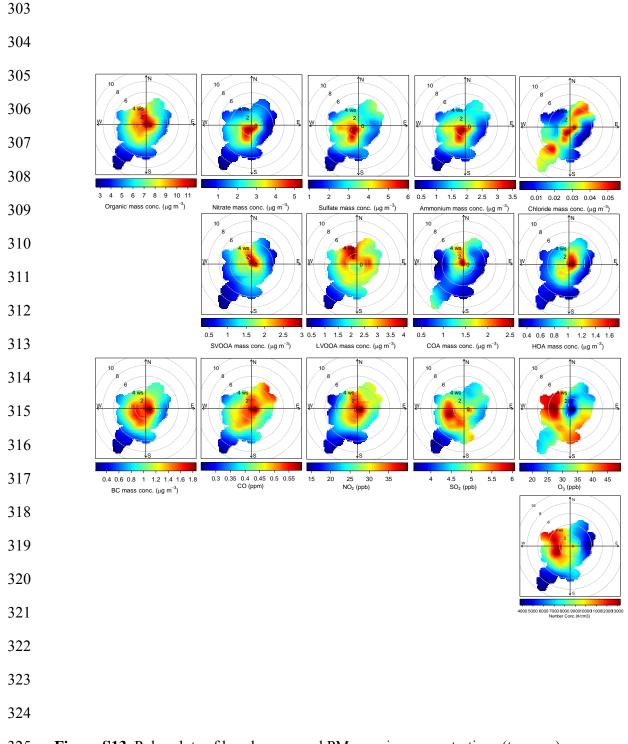


Figure S12. One-hour averaged diurnal profiles of organics in (a) accumulation mode and (b)
 fine mode.



**Figure S13.** Polar plots of hourly averaged PM<sub>1</sub> species concentrations (top row), mass

- 326 concentrations of the five OA factors identified from PMF analysis (middle row), the mixing327 ratios of various gas phase species (second row from the bottom), and the particle number
- 328 concentrations (bottom row) as a function of WS and direction.

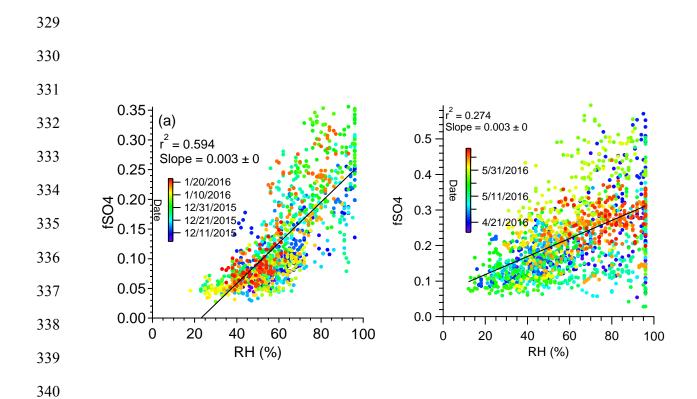


Figure S14. Scatterplot of the variations of fSO<sub>4</sub> ratios as a function of RH (a) during winter
(Kim et al., 2017); (b) during spring in this study.

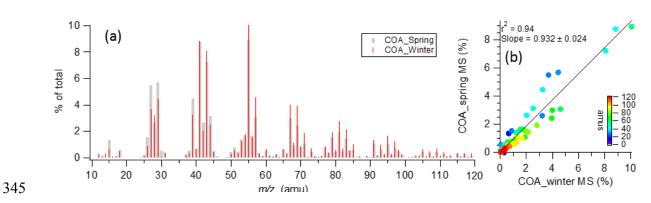


Figure S15. (a) Comparisons of mass spectra of the COAs from this study (spring) and the one
from winter, 2015-2016; (b) scatter plots of both COA mass spectra



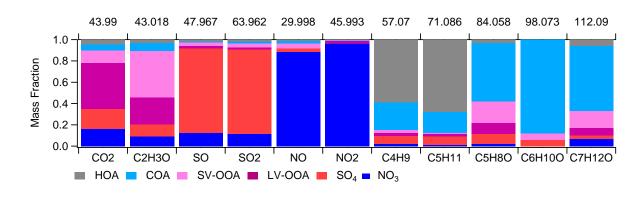




Figure S16. Mass fractional contribution of the factors (four OA factors + two inorganic factors)
 from PMF analysis to various ions that are relevant to each significant tracer.

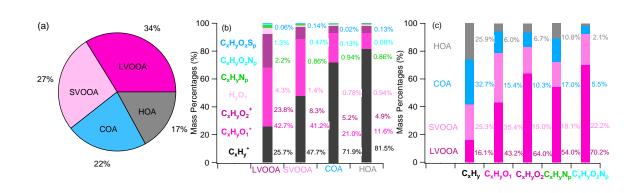
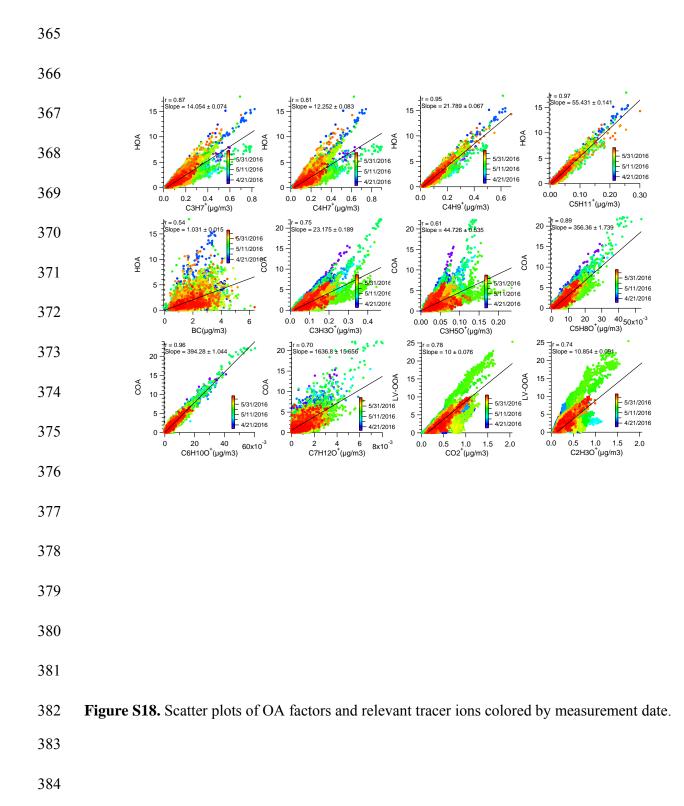
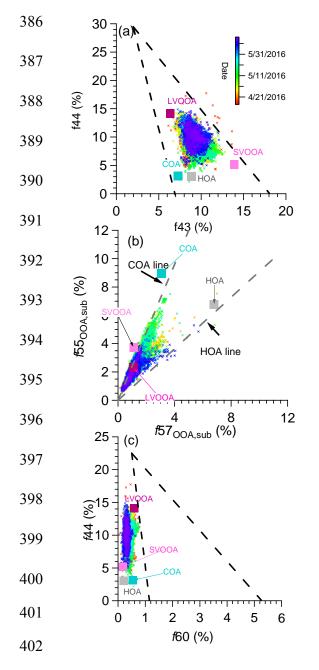
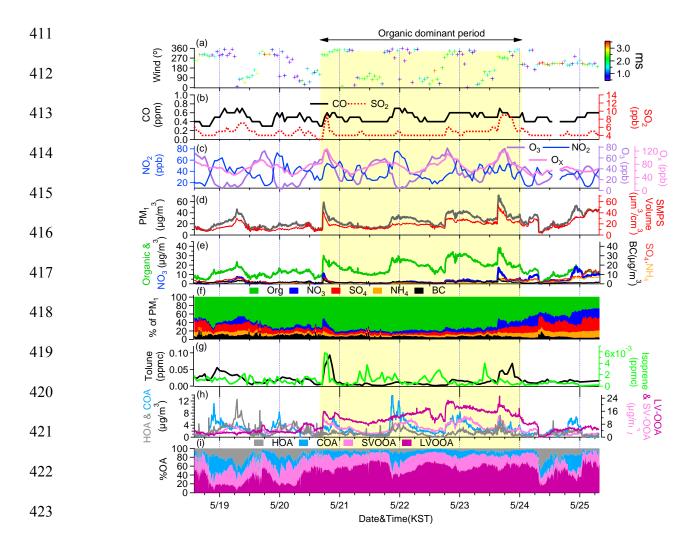


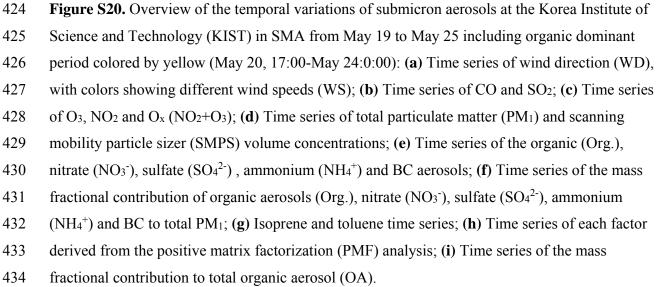
Figure S17. (a) Compositional pie chart of the average fractional contribution of each of the OA factors to the total OA over the campaign; (b) Average mass fractional contributions of seven ion families to each of the OA factors and; (c) Average mass fractional contributions of four OA factors to 4 each ion families

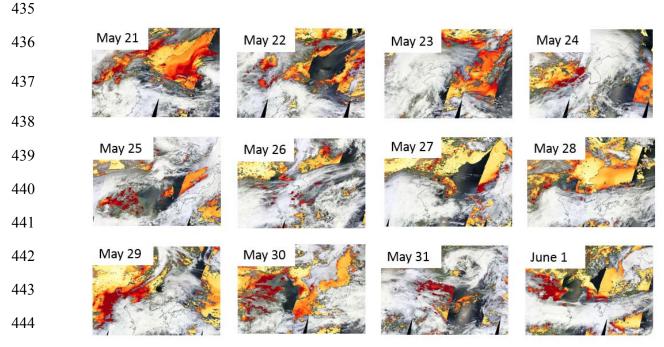




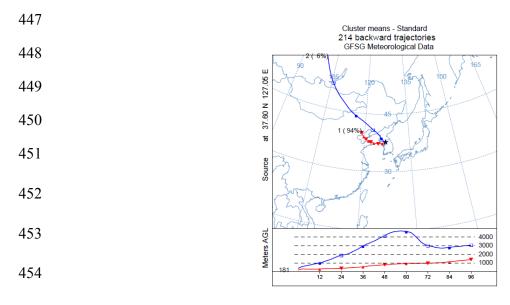
**Figure S19.** Triangular plots of (**a**) *-f*<sub>44</sub> versus *f*<sub>43</sub> and (**b**) *f*<sub>44</sub> versus *f*<sub>60</sub> (**c**) *f*<sub>55,00A sub</sub> versus *f*<sub>57</sub>, ooA sub for the five OA factors and all of the measured OA data (dots), colored by the time of the day. *f*<sub>43</sub>,*f*<sub>44</sub>, and *f*<sub>60</sub> are the ratios of the organic signal at m/z = 43, 44, and 60 to the total organic signal in the component mass spectrum, respectively. *f*<sub>55,00A sub</sub> and *f*<sub>55,00A sub</sub> are the ratios of the organic signal at m/z 55, 57 after subtracting the contributions from SV-OOA and LV-OOA (e.g., *f*<sub>55,00A sub</sub> = m/z 55- m/z55sv-ooA - m/z55LV-ooA; *f*<sub>57,00A sub</sub> = m/z 57- m/z 57sv-ooA - m/z 57LVooA)







445 Figure S21. Long range transportation of plums from China to Korea during Haze period. Plots446 are from MODIS, terra.

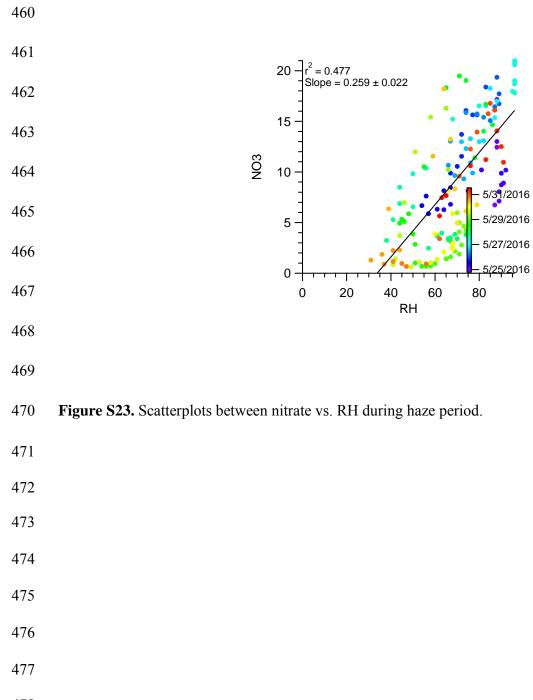


456 **Figure S22.** Two cluster solution of backtrajectroy analysis during haze period from (5/24 7:00-

457 6/2 24:00). The trajectories were released at half of the mixing height at the KIST (latitude:

458 37.60N; longitude: 127.05E) and the average arriving height for the back trajectories for this

459 study was approximately 181 m.



#### 479 1. Properties of Organic sources

#### 480 1.1 Hydrocarbon-like OA (HOA)

481 Alkyl fragments ( $C_nH_{2n+1}^+$  and  $C_nH_{2n-1}^+$ ) made a substantial contribution to the HOA 482 spectrum, with major peaks at m/z's 41, 43, 55, and 57 which were mostly composed of C<sub>3</sub>H<sub>5</sub><sup>+</sup>, 483  $C_{3}H_{7}^{+}$ ,  $C_{4}H_{7}^{+}$ , and  $C_{4}H_{9}^{+}$  ions, respectively (Fig. 8a). These major peaks and the overall picket 484 fence fragmentation pattern resulting from the  $C_nH_{2n+1}^+$  ions are typical features of the HOA 485 spectra reported in other studies and are due to the association of these aerosols with fossil fuel 486 combustion (e.g., Alfarra et al., 2007; Lanz et al., 2008; Sun et al., 2011b; Zhang et al., 487 2005a;Huang et al., 2010;Morgan et al., 2010;Ng et al., 2011;Collier, 2015;Ge et al., 2012a;Kim 488 et al., 2017; Young et al., 2016). In addition, strong correlations were observed between the time 489 series of HOA and the  $C_nH_{2n+1}^+$  and  $C_nH_{2n-1}^+$  ions, e.g.,  $C_3H_7^+$  (r = 0.87),  $C_4H_7^+$  (r = 0.81),  $C_4H_9^+$ (r = 0.95), and C<sub>5</sub>H<sub>11</sub><sup>+</sup> (r = 0.96) (Fig. S17 and Table 2). Due to the dominance of chemically 490 491 reduced hydrocarbon species, the O/C ratio of the HOA was low (0.15), whereas the H/C ratio 492 was high (2.00). On average, HOA contributed 17% of the total OA, which was similar to the 493 number observed in winter (16 %) (Kim et al., 2017).

#### 494 1.2 Cooking OA (COA)

495 COA has been widely reported in urban areas with high population densities (e.g., He et 496 al., 2010;Huang et al., 2010;Mohr et al., 2012;Sun et al., 2011b;Young et al., 2016;Ge et al., 497 2012a;Wang et al., 2016b;Xu et al., 2014;Hayes et al., 2013). In this study, COA was found to 498 account for 22% of the total OA mass, higher than HOA (Fig. S16). The diurnal pattern of COA 499 displayed a large enhancement at evening starting from ~ 19:00, i.e., dinner time, and a small lunch 500 time peak at ~ 12:00. Elevated COA concentration and larger fractional contribution to OA mass 501 were observed throughout the night (Figs. 3 and 8f).

Similar to HOA, the mass spectrum of COA contained many alkyl fragments, but to a lesser extent (71.9 % of the total signal in COA spectrum compared to 81.5 % of the total signal in HOA spectrum) (Fig. S16). COA also contained significantly larger amounts of oxygen containing ions than HOA (e.g.,  $C_xH_yO_1^+ = 21.0\%$  vs. 11.5% and  $C_xH_yO_2^+ = 5.2\%$  vs. 4.9%) (Fig. S16), and thus had a higher O/C ratio (0.19 vs 0.15) and a lower H/C ratio (1.83 vs. 2.00). The O/C ratio (0.19) of COA in spring was found to be higher than that in winter (0.14) but both values were within the range of the O/C ratios of COA observed in other studies (e.g., Barcelona (0.27) (Mohr et al., 509 2012), New York City (NYC) (0.23) (Sun et al., 2011a) and Fresno (0.14 in 2010 (Ge et al., 2012b) 510 and 0.19 in 2013 (Young et al., 2016)). Key tracers for identifying the presence of aerosols from 511 cooking related activities, such as C<sub>3</sub>H<sub>3</sub>O<sup>+</sup> (m/z 55), C<sub>3</sub>H<sub>5</sub>O<sup>+</sup> (m/z 57), C<sub>5</sub>H<sub>8</sub>O<sup>+</sup> (m/z 84) and 512  $C_{6}H_{10}O^{+}$  (*m/z* 98) (He et al., 2004; Adhikary et al., 2010; Mohr et al., 2009; Zhao et al., 2007; Ge et 513 al., 2012a;Sun et al., 2011b), all showed good correlation in time series with COA, e.g., C<sub>3</sub>H<sub>3</sub>O<sup>+</sup> 514 (r = 0.75),  $C_3H_5O^+$  (r = 0.61),  $C_5H_8O^+$  (r = 0.89),  $C_7H_{12}O^+$  (r = 0.70), and  $C_6H_{10}O^+$  (r = 0.99) (Fig. 515 S17 and Table 2) and COA was a major contributor to the signals of  $C_5H_8O^+$ ,  $C_6H_{10}O^+$ , and 516 C<sub>7</sub>H<sub>12</sub>O<sup>+</sup>, accounting for 62%, 94%, and 67%, respectively, of their signals (Fig. S15). Finally, the 517 ratios between f55 and f57 for OA in Seoul increased proportionally as the fractional contribution 518 of COA to total OA increased (Fig. S18b), with a "V" shape indicated by the two edges defined 519 by the COA and the HOA factors from several urban AMS data sets (Mohr et al., 2012). These 520 observations all confirm the identification of COA at SMA.

521 1.3. Semi-volatile and low volatile oxygenated OA (SV-OOA and LV-OOA)

In addition to the two POA factors, two OOA factors were identified and were found to account for an average of 61% of the OA mass (Fig. S16a) with LV-OOA and SV-OOA being 34 and 27%, respectively. OOA is ubiquitous in the atmosphere and usually a dominant component of submicrometer particles (Jimenez et al., 2009;Zhang et al., 2007). SOA formation was more important in spring that in winter: OOA contributed an average 61% of the OA mass in this study but only 41% during winter. In addition, the average OOA concentration in this study was higher than in winter (5.9 vs. 4.9  $\mu$ g m<sup>-3</sup>).

 $CO_2^+$  (m/z 44) and  $C_2H_3O^+$  (m/z 43) were major fragments of oxidized organics and 529 530 therefore frequently used as key tracers for identifying the presence of SOA in absence of biomass 531 burning influence. In this study, the time series of SV-OOA and LV-OOA correlated well with 532 these ions, e.g., LV-OOA vs  $CO_2^+$  (r = 0.78), SV-OOA vs  $C_2H_3O^+$  (r = 0.89) (Figs. S17 and Table 533 2). As shown in the triangle plots in Fig. S18, SV-OOA (O/C = 0.56; H/C = 1.90) resides within 534 the region representing fresher SOA, with a low f44. The mass spectrum of SV-OOA is 535 characterized by two prominent peaks, m/z 29 (mainly CHO<sup>+</sup>) and m/z 43 (mainly C<sub>2</sub>H<sub>3</sub>O<sup>+</sup>) (Fig. 536 8c). In addition, the SV-OOA show high fraction of  $C_xH_yO_1^+$  family (41 %), most likely from 537 carbonyl or alcohol functional groups, while much lower contribution of  $C_xH_yO_2^+$  family (8.3 %), 538 mainly from carboxylic functional groups (Fig. S16b). The diurnal variation of SV-OOA shows 539 higher concentrations at night and lower concentrations at daytime with a small afternoon peak 540 (Fig. 8g). The mass spectral features and diurnal variation of SV-OOA suggest that this factor was 541 driven by the partitioning of semivolatile organics between gas and particle phase and that SV-542 OOA formation was facilitated by the high humidity and low temperature during nighttime. In 543 addition, SV-OOA appeared to represent relatively fresh SOA formed from photochemical 544 reactions during daytime (Docherty et al., 2011). Indeed, a decrease of the PBL height, thus less 545 dilution, together with continued reactions of VOCs with the nighttime residual ozone (20–35 ppb 546 on average; Fig. 3) and other oxidants (e.g., nitrate radical) might have played a role in the 547 production of semivolatile secondary organic species at night. This kind of formation processes 548 also suggests that SV-OOA was mostly formed locally. Indeed, the polar plot of SV-OOA showed 549 that its high concentrations tended to be associated with lower wind speed.

550 On the other hand, the LV-OOA factor is characterized by high O/C ratio (=0.91) (Fig. 8d) 551 and high  $f_{44}$  (14 %) (Fig. S18a), which are in similar range as those previously reported for aged 552 and highly oxidized OA, including LV-OOAs, from various locations (e.g., Hayes et al., 553 2013;Mohr et al., 2012;Zhang et al., 2014;Ng et al., 2010). The diurnal profile of LV-OOA was 554 different than SV-OOA as well, showing a gradual increase from 8:00 till late afternoon despite a 555 rising PBL height (Fig. 8h). This diurnal pattern was very similar to  $O_x$  ( $O_x = O_3 + NO_2$ ) (Fig. 8). 556 In fact, LV-OOA correlated positively with  $O_x$  (r = 0.57; Fig. S19) and the correlation was higher 557 (r = 0.60; Fig. 9) between 10:00 and 16:00 when photochemical processing was intense. Note that 558 a high organic period (May 20-May 23) was excluded in this correlation because it appeared to be 559 driven by processes that were different than the other periods thus showed substantially higher 560 LV-OOA/O<sub>3</sub> ratios than the rest of the study (Fig. 9). Details on this episode is discussed in section 561 3.4. Unless otherwise indicated, the high organic period (May 20-May 23) was excluded from the 562 O<sub>x</sub> vs. OOA, LV-OOA and SV-OOA correlations. The relatively good correlation between LV-563 OOA and O<sub>x</sub> during daytime indicates that LV-OOA corresponded to photochemically produced 564 SOA (Fig. 8d). Furthermore, a gradual increase starting from ~08:00 till late afternoon indicated 565 that the photochemical production was coupled with transport of air masses from regions outside 566 of SMA. Indeed, the diurnal profile of ozone at the KIST site demonstrated a peak between 16:00– 567 17:00 (Figs. 3,8) associated with southwesterly winds that transported ozone produced outside of 568 SMA coupled with local daytime photochemical production. The high O/C (0.91) of the LV-OOA 569 as well as a dispersed feature of its polar plot (Fig. S12) all support the regional secondary 570 characteristics of LV-OOA.

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