

Rapid and reliable assessment of methane impacts on climate

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Abstract. It is clear that the most effective way to limit global temperature rise and associated impacts is to reduce human emissions of greenhouse gases, including methane. However, quantification of the climate benefits of mitigation options are complicated by the contrast in the timescales at which short-lived climate pollutants, such as methane, persist in the atmosphere compared to carbon dioxide. Whereas simple metrics fail to capture the differential impacts across all timescales, sophisticated climate models that can address these temporal dynamics are often inaccessible, timeintensive, require special infrastructure, and include high unforced interannual variability that makes it difficult to analyse small changes in forcings. On the other hand, reducedcomplexity climate models that use basic knowledge from observations and complex Earth system models offer an ideal compromise in that they provide quick, reliable insights into climate responses, with only limited computational infrastructure needed. They are particularly useful for simulating the response to forcings of small changes in different climate pollutants, due to the absence of internal variability. In this paper, we build on previous evaluations of the freely available and easy-to-run reduced-complexity climate model MAGICC by comparing temperature responses to historical methane emissions to those from a more complex coupled global chemistry-climate model, GFDL-CM3. While we find that the overall forcings and temperature responses are comparable between the two models, the prominent role of unforced variability in CM3 demonstrates how sophisticated models are potentially inappropriate tools for small forcing scenarios. On the other hand, we find that MAG-ICC can easily and rapidly provide robust data on climate responses to changes in methane emissions with clear signals unfettered by variability. We are therefore able to build confidence in using reduced-complexity climate models such as

MAGICC for purposes of understanding the climate implications of methane mitigation.

1 Introduction

Reduced-complexity climate models offer an ideal framework for evaluating greenhouse gas mitigation options if they can accessibly and rapidly reproduce the results of the more complex global chemistry–climate models (CCMs) that include more advanced and comprehensive treatments of chemistry and physics (Meinshausen et al., 2011a). However, there is a critical need to build confidence in the ability of reduced-complexity models to simulate temperature responses to individual greenhouse gases rather than just the suite of climate pollutants, because greenhouse gases have vastly different radiative properties and atmospheric lifetimes (Myhre et al., 2013; Fiore et al., 2015); it is important to confirm that individual species are represented appropriately if reduced-complexity models are to serve as an effective tool for assessing climate benefits of mitigation actions. This is especially central for the analysis of methane (CH₄) mitigation actions, on which the climate policy community has been increasingly focused (e.g. Shindell et al., 2012; Collins et al., 2018). Therefore, this paper builds on previous evaluations by comparing forcing and temperature responses to historical methane and carbon dioxide (CO₂) concentrations from a widely used reduced-complexity climate model with those from a state-of-the-art coupled global chemistry-climate model. While it is difficult to compare climate responses of simple models with those of complex ones because of the presence of unforced variability in the latter, we ultimately seek to determine if general temporal patterns and magnitudes are consistent enough to justify the use of a reduced-complexity climate model as a reliable tool for rapid assessment of methane mitigation measures.

Climate change impacts have now been observed on every continent and in every ocean (Stocker et al., 2013). If we want to reduce short- and long-term anthropogenically caused warming, then we need to reduce emissions of several climate pollutants. While limiting long-term climate warming requires drastically reducing CO2 emissions, reducing emissions of short-lived climate pollutants (SLCPs) - specifically, methane and black carbon (BC) - has been identified as one of the most impactful ways to reduce short-term warming (e.g. Ramanathan and Xu, 2010; Shindell et al., 2012; Rogelj et al., 2013; Shoemaker et al., 2013). Methane emissions in particular account for a quarter of the excess energy trapped by human emissions, and today's global anthropogenic methane emissions will have a larger impact on short-term warming than today's global fossil fuel CO2 emissions (based on forcing data provided in Myhre et al., 2013, and references therein; methane emissions provided in EPA, 2012; CO₂ emissions provided in IEA, 2015; and radiative efficiency estimates of methane provided in Etminan et al., 2016). Sustained methane emissions will also impact longterm warming (Allen et al., 2016). Furthermore, reducing methane emissions has air quality, health, and food security co-benefits (Shindell et al., 2012; West et al., 2013; Zhang et al., 2016; Melvin et al., 2016).

Most methane mitigation measures are assessed as a comparison to carbon dioxide warming impacts; almost all policy analyses rely on the simple metric global warming potential (GWP) because of its simplicity and ease of use (Fuglestvedt et al., 2010; Ocko et al., 2017). However, GWP is limited in its ability to quantify climate effects because it relies on the integrated impact of a pulse of emissions over a specified time horizon. Because methane and CO_2 have vastly different atmospheric lifetimes, their respective climate impacts occur over different timescales. Due to the inherent selection of a single time horizon, GWP is incapable of capturing these important temporal distinctions (e.g. Solomon et al., 2010; Alvarez et al., 2012) unless two time horizons that represent short- and long-term impacts are reported simultaneously (Ocko et al., 2017).

Assessment of SLCP climate impacts over different timescales can be performed using comprehensive global CCMs; however, a full assessment of various SLCP scenarios using sophisticated CCMs is computationally intensive and time-consuming, and forcing perturbations from slight changes in individual species are often too small for the response signal to be detected among the high unforced internal climate variability present in CCMs (e.g. Ocko et al., 2014). Determining robust climate responses to smallforcing scenarios using CCMs therefore requires a large number of ensemble simulations (Deser et al., 2012). Given that many institutions do not have access to CCMs nor the technical capacity or expertise to run these models, they must rely on partnerships with modelling centres that are often fo-

cused on model development. These characteristics of CCMs reinforce the use of the simple GWP metric for assessments of climate pollutant mitigation measures.

While detailed assessment of regional climate responses can only be provided by complex CCMs, reducedcomplexity climate models offer a useful alternative for global changes in major climate characteristics that is far more advanced than GWP but avoids the need for the tremendous amount of computational resources required to perform CCM simulations (and especially with enough ensemble members to average out unforced variability). These simpler models can rapidly analyse global average climate responses because they are easily accessible and quick to run, thereby providing immediate scientific guidance for mitigation assessments. Further, because they do not include unforced internal variability, they provide clear responses to small forcing scenarios without any noise.

There are several models that have been developed that fall within this intermediate complexity class - more advanced than simple metrics but far less sophisticated than CCMs. They range from simplified expressions (e.g. Shine et al., 2005) to more complex chemistry and physics with computations of only a few climate indicators averaged over large spatial domains (e.g. Meinshausen et al., 2011a; Hartin et al., 2015). One of the latter is the freely available Model for the Assessment of Greenhouse-gas Induced Climate Change (MAGICC), initially developed in the late 1980s (Wigley and Raper, 1987, 1992) and routinely updated since then (e.g. Meinshausen et al., 2011a). While not meant to replace atmosphere-ocean global climate models (AOGCMs) and carbon cycle models, MAGICC is a complementary, computationally inexpensive tool that is capable of efficiently analysing basic climate responses (such as radiative forcing, surface air temperature, and ocean heat uptake) to a suite of emission scenarios. Confidence in MAGICC results comes from a comprehensive effort to match several AOGCMs and carbon cycle models (Meinshausen et al., 2008, 2011a). Evaluations show that MAGICC closely reproduces temperature responses to aggregated forcings from the sophisticated Coupled Model Intercomparison Project CMIP3 atmosphere-ocean and C4MIP carbon cycle models (Meinshausen et al., 2011c).

While not the only model of its class, the reducedcomplexity model MAGICC is an especially great resource for mitigation analysis because of its widespread use in international climate reports and the ability it gives the user to modify future emissions of every radiatively active species. Therefore, when numerous scenarios exist and need to be evaluated for decision-making, a tool like MAGICC can provide rapid insight into the climate impacts of various options. However, to build confidence in MAGICC's evaluation of greenhouse gas mitigation strategies, we need to adequately assess its ability to reproduce climate responses to individual greenhouse gases beyond the aggregated forcings. Here, we analyse the capability of MAGICC in simulating climate responses to historical increases (1860–2014) in methane and CO₂ by comparing the results with that from a stateof-the-art coupled chemistry–climate model, the National Oceanic and Atmospheric Administration (NOAA) Geophysical Fluid Dynamics Laboratory (GFDL) CM3 model, which has been shown to adequately reproduce historical temperature trends (Golaz et al., 2013; Griffies et al., 2011; Donner et al., 2011; Winton, 2011; John et al., 2012; Levy et al., 2013). While it is difficult to compare simpler models with sophisticated ones for scenarios with small forcings – due to high interannual variability built into the latter – it is nevertheless important to do so because of the more advanced and comprehensive chemistry and physics in the more complex models.

We compare the response of the two models to assess similarities and differences, seeking to determine (i) whether the response of the forcings and temperature is comparable, (ii) whether the complexity of the CCM provides any benefits over the simple model, and (iii) whether the lack of variability in the simple model provide any advantages over the CCM when looking at small forcing amounts. Our goal is to build confidence in the simulation of the climate response to methane in order to justify future use of reduced-complexity climate models, such as MAGICC, to assess the climate impact of methane emissions mitigation scenarios. In this analysis, we add to previous evaluations by showing a high correlation between CM3's and MAGICC's radiative forcing and surface air temperature responses to changes in either CO_2 or methane in isolation, despite large unforced variability in CM3, thereby strengthening confidence in MAGICC's simulation of climate responses to individual greenhouse gases with vastly different radiative properties and lifetimes.

2 Models and simulations

2.1 MAGICC model description

We use MAGICC version 6, developed in 2011 (http://www. magicc.org/download, last access: 1 October 2018). MAG-ICC represents the complex coupled carbon-cycle climate system as a hemispherically averaged upwelling-diffusion ocean model coupled to an atmosphere layer and a globally averaged carbon cycle model. The atmosphere has four boxes (one over land and one over the ocean for each hemisphere) and is coupled to the mixed layer of the ocean hemispheres. The default number of ocean layers in each hemisphere is 50 including the mixed layer (though users can select the number of levels), and heat exchange is driven by vertical diffusion and advection. The terrestrial carbon cycle model is a globally integrated box model with one living plant box and two dead biomass boxes (one for detritus and one for organic matter in soils). The terrestrial carbon cycle does not feedback into carbon dioxide concentrations in the atmosphere. The sea-to-air carbon flux is determined by the partial pressure differential for carbon dioxide between the atmosphere and surface layer of the ocean.

From 1765 to 2005, the MAGICC v.6 radiative forcing is driven by global-mean concentrations of greenhouse gases (carbon dioxide, methane, nitrous oxide, ozone-depleting substances and their replacements); prescribed regional direct aerosol radiative forcings (sulfate, black and organic carbon, sea salt, mineral dust); land-use, volcanic, and solar radiative forcings; prescribed black carbon on snow radiative forcings; emissions of tropospheric ozone precursors (carbon monoxide, nitrogen oxides, non-methane volatile organic carbon); and indirect (first and second) aerosol forcings calculated from prescribed regional aerosol optical depths (parameterizations described in detail in Meinshausen et al., 2011a). For 2006 to 2014, the model is driven by emissions of gases and aerosols taken from the Representative Concentration Pathway (RCP8.5) scenario to capture a business-asusual trajectory. Climate responses (such as surface air temperature) are provided as global annual averages and also across four spatial boxes (over land and ocean and by hemisphere).

Historical greenhouse gas concentrations are from Meinshausen et al. (2011b); prescribed aerosol forcings and landuse historical forcings are from the National Aeronautics and Space Administration (NASA) GISS model (http://data.giss. nasa.gov/, last access: 1 October 2018), solar irradiance is provided by Lean (2010), and historical emissions of ozone precursors are from Lamarque et al. (2010). Present-day and future (2005–2100) forcings are driven by emissions of gases and aerosols and are taken from the RCP8.5 scenario to capture a business-as-usual trajectory, though we restrict our analysis here to climate responses from 1860 to 2014. Carbon dioxide radiative forcings are calculated using a standard simplified expression (Shine et al., 1990, with updated scaling parameter from Myhre et al., 1998). Methane radiative forcings are calculated using a radiative efficiency parameter in conjunction with standard simplified expressions from Myhre et al. (1998) and account for overlap between methane and nitrous oxide absorption bands.

For the most recent version of MAGICC, seven key climate parameters were calibrated to match 19 AOGCMs used in the Intergovernmental Panel on Climate Change (IPCC) Fourth Assessment Report, AR4 (see Meinshausen et al., 2011a). The parameters include equilibrium climate sensitivity, land-ocean warming ratio at equilibrium, vertical diffusivity in the ocean, sensitivity of feedback factors to radiative forcing change, sensitivity of vertical diffusivity at mixed layer boundary to global-mean surface temperatures (i.e. thermal stratification), land-ocean heat exchange coefficient, and an amplification factor for the ocean-to-land heat exchange. The MAGICC parameter set that best reproduces surface air temperatures and heat uptake of each AOGCM is determined via an optimization routine with 1000 iterations to find the combination that minimizes the squared differences between low-pass-filtered time series. The effective climate sensitivities in MAGICC v.6 vary over time due to spatially non-homogeneous varying feedbacks, until they reach the equilibrium climate sensitivity. The equilibrium climate sensitivity input into MAGICC depends on which AOGCM calibration is used; they range from 1.9 to $5.73 \,^{\circ}$ C across all 19 models, with a mean (median) of $2.88 \,^{\circ}$ C ($2.59 \,^{\circ}$ C). Multi-model ensembles are generated by running each simulation for all 19 AOGCM calibrations, which we refer to as "physics-driven ensemble members." The user of the downloaded MAGICC model can select which parameters to use for each simulation.

While the MAGICC model is particularly well-calibrated to more sophisticated models, the realism of MAGICC results relies on the realism of GCMs, which have their own sets of limits and uncertainties. Further limitations of MAG-ICC include incomplete knowledge of forcing patterns, unknown responses outside of the calibrated range, limited set of climate responses evaluated (such as temperature and heat uptake but not precipitation), reliance on a high level of parametrization (such as cloud feedbacks tuned to match those of more sophisticated GCMs), and possible errors in the data used for calibration. In addition, although the model is freely available, it is not open-source. However, despite these limitations, MAGICC has been shown to reasonably reproduce climate responses to all-forcing scenarios (Meinshausen et al., 2011a) and is one of the most prominent reduced-complexity climate models in use.

2.2 CM3 model description

We employ the GFDL global coupled atmosphere-oceanchemistry model (GFDL-CM3; Donner et al., 2011; Griffies et al., 2011) to assess the climate response to historical changes in methane and CO2. CM3 uses a finite-volume dynamical core on a cubed-sphere horizontal grid composed of six faces; each face includes 48×48 grid cells. The size of the grid cells range from 163 km at the corners to 231 km near the face centres. In the vertical, the model domain extends from the surface up to 0.01 hPa (86 km), with 48 vertical hybrid sigma pressure levels. The model simulates tropospheric and stratospheric chemistry interactively over the full vertical domain, with simulated ozone and aerosols influencing radiation calculations (Naik et al., 2013; Austin et al., 2013). Ensemble members for CM3 are generated by employing different sets of stochastically selected initial conditions (discussed in more detail in Sect. 2.3.), to which we refer herein as "initial condition-driven ensemble members." The equilibrium climate sensitivity of CM3 is 4.8 °C (Paynter et al., 2018), which is in the range of the MAGICC calibration models but higher than the median and mean.

Global mean concentrations of well-mixed greenhouse gases (WMGHGs), including carbon dioxide, nitrous oxide, methane, and ozone-depleting substances (ODSs), are specified for radiation calculations for the historical period (1860– 2005) from Meinshausen et al. (2011b) and for the period 2006 to 2014 following the RCP8.5 scenario. Within the chemistry module, global mean concentrations of methane are prescribed at the surface as the lower boundary condition and are allowed to undergo chemistry everywhere else in the model domain. Radiation calculations do not see the full three-dimensional methane field (simulated in the chemistry module) and only employ the global-mean concentrations; however, changes in ozone and water vapour are seen by the radiation. Further, CM3 CO_2 concentrations do not get altered by reactions that occur in the model.

CM3 is forced with emissions of short-lived species (including ozone precursors) as well as aerosols and their precursors, volcanic aerosols, solar radiation, and land-use change, as described in detail by Donner et al. (2011) and Naik et al. (2013). Anthropogenic emissions, including from biomass burning and ships, for the time period 1860-2005 are from the dataset of Lamarque et al. (2010) and were developed in support of the Couple Model Intercomparison Project Phase 5 (CMIP5). For the years 2006–2014, anthropogenic emissions follow the RCP8.5 scenario. Natural emissions of all precursor species, except isoprene, are included as described by Naik et al. (2013). Biogenic isoprene emissions are calculated interactively, as described by Lin et al. (2012), based on the Model of Emissions of Gases and Aerosols in Nature (MEGAN) (Guenther et al., 2012). "Explosive" volcanic eruptions are imposed via a time series of volcanic optical properties rather than from direct injection of sulfur into the stratosphere (Stenchikov et al., 2006; Donner et al., 2011).

Shortwave and longwave radiation algorithms in CM3 are described in Freidenreich and Ramaswamy (1999) and Schwarzkopf and Ramaswamy (1999), respectively, with some modification to enhance computational efficiency (GAMDT, 2004). The shortwave algorithm includes 18 bands in the solar spectrum, and the longwave algorithm includes 8 bands. Shortwave radiation parameterizations account for absorption by water vapour, carbon dioxide, ozone, molecular oxygen, molecular scattering, and absorption and scattering by aerosols and clouds. The longwave radiation parameterizations account for absorption and emission by water vapour, carbon dioxide, ozone, nitrous oxide, methane, halocarbons (CFC-11, CFC-12, CFC-13 and HCFC-22), aerosols, and clouds. Aerosols included are sulfate, carbonaceous (black and organic carbon), dust, and sea salt.

CM3 includes explicit representation of both the direct and indirect aerosol effects on radiation. For the calculation of the direct effect of aerosols on radiation, physical and optical properties of sulfate, black carbon, organic carbon, sea salt, and dust are considered in the model (Donner et al., 2011). Sulfate and black carbon are assumed to be internally mixed while all other aerosols are assumed to be externally mixed for radiation calculations. To account for the indirect effect of aerosols via aerosol–water cloud interactions, the model treats water-soluble aerosols, including sea salt, and organic aerosols as cloud-condensation nuclei (CCN), allow-

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Table 1. Direct experiments performed by MAGICC and GFDL-CM3 models, as well as derived simulations. All-forcing simulations include time-varying natural and anthropogenic forcings, but land use is held constant. Each experiment is run for 19 physics-driven ensemble members for GFDL-CM3 over the period 1860–2014.

	Experiments	Abbreviation	MAGICC v6	GFDL-CM3
	All-forcing	AllForc	Х	Х
	CO ₂ concentrations held constant at 1860 levels	CO ₂ 1860	Х	Х
	Methane concentrations held constant at 1860 levels	CH ₄ 1860	Х	Х
Direct simulations	All-forcing with methane chemistry turned off	CH ₄ nochem	Х	
	Methane concentrations held at 1860 levels with methane chemistry turned off	CH ₄ 1860nochem	Х	
	Methane concentrations held at 1860 for radiation	CH ₄ 1860chem		Х
	Methane concentrations held at 1860 for chemistry	CH ₄ 1860chem		Х
	CO ₂ -only		AllForc – $CO_2 1860$	
Derived	CH ₄ -only AllForc – CH ₄ 1860		41860	
simulations	CH ₄ -direct		CH_4 nochem – CH_4 1860nochem	AllForc - CH ₄ 1860rad
	CH ₄ -indirect		CH ₄ -only – CH ₄ -direct	AllForc – CH ₄ 1860chem

ing a physically based parameterization of CCN activation (Ming et al., 2006). The model does not consider the reduction in surface albedo caused by the deposition of black carbon on snow-covered surfaces.

2.3 Simulations

Three historical simulations are run for both MAGICC and GFDL-CM3 to derive climate responses to isolated CO_2 and methane concentrations, respectively. MAGICC was run from 1750 to 2100 by default, and CM3 was run from 1860 to 2014. As shown in Table 1, the direct runs for both models include an all-forcing simulation with all forcings varying with time except for land use, a simulation with CO_2 concentrations held at 1860 levels, and a simulation with methane concentrations held at 1860 levels. Subtracting temperature responses of the two latter runs from the former yields CO_2 -only and methane-only climate responses, respectively (see Eqs. 1 and 2). The same equations hold for the forcings as well.

$$\Delta T_{\rm CO_2} = T_{\rm AllForc} - T_{\rm CO_21860} \tag{1}$$

$$\Delta T_{\rm CH_4} = T_{\rm AllForc} - T_{\rm CH_41860} \tag{2}$$

For MAGICC, each simulation is run for all 19 AOGCMcalibrated configurations; each 350-year integration took approximately 1 s to run on a modern PC with a 3 GHz CPU processing speed. We use default MAGICC gas and aerosol properties, but update tropospheric ozone radiative efficiency and methane atmospheric lifetimes to IPCC Fifth Assessment Report (AR5) values (Myhre et al., 2013; Stevenson et al., 2013) to reflect the latest science. (Note that the updated atmospheric lifetime only impacts the model from 2006 to 2014 as it is driven by emissions and not concentrations during this period.) However, we do not include newer estimates of methane radiative effects that account for shortwave absorption in addition to longwave absorption (Etminan et al., 2016) to be consistent with the CM3 model that only includes longwave effects. Including the shortwave component increases methane's radiative efficiency by over 20 %. Further, we specifically do not tune MAGICC model climate and forcing properties to match that of CM3 because we are assessing how a "standard version" of the reducedcomplexity climate model compares with CM3; the goal is not to match MAGICC to CM3 but to assess whether a downloaded version of MAGICC broadly behaves similarly to CM3. However, two of MAGICC's physics-driven ensemble members are derived from two predecessors of CM3: CM2.0 and CM2.1 (Delworth et al., 2006).

The set-up of GFDL-CM3 simulations conducted here was similar to that adopted for simulations performed in support of the CMIP5, except we obtained initial conditions from a longer pre-industrial control (3000 years). Three-member initial condition-driven ensembles of transient CM3 simulations were performed with each ensemble member initialized stochastically at different points in the pre-industrial control simulation. Each 155-year integration of CM3 took about 15 days to complete on the NOAA's Remotely Deployed High Performance Computing System (RDHPCS) machine, known as "Gaea", running on 464 processors. While three ensemble members is relatively few, we are limited by com-



Figure 1. Atmospheric concentrations of carbon dioxide in parts per million (orange) and methane in parts per billion (blue) used in this study (Meinshausen et al., 2011b). Note that concentrations are prescribed for CM3 throughout this time period, but only prescribed for MAGICC through 2005, of which methane emissions inputs drive the model from 2006 to 2014. The resulting concentrations are plotted here.

putational resources, and studies have shown that forced changes in air temperature, as opposed to changes in atmospheric circulation and precipitation, can be detected with fewer ensemble members (Deser et al., 2012).

To compute CM3 radiative forcings for CO₂ and methane (direct and indirect) that are closest to the definition used by MAGICC (the forcing at the tropopause after stratospheric temperature adjustment), we performed simulations with the atmosphere-only version of CM3-AM3. The model configuration of AM3 was exactly the same as CM3 except AM3 model integrations over the period 1870 to 2014 were performed with observed sea-surface temperature and sea-ice cover (Rayner et al., 2003) and therefore do not include an ensemble driven by different initial conditions. Through the additional AM3 simulations, we were able to diagnose transient effective radiative forcing (ERF), i.e. the change in net radiation balance at the top of the atmosphere (TOA) following a perturbation to the climate system taking into account any rapid adjustments (Shine et al., 2003; Myhre et al., 2013) due to CO₂ and methane. Transient ERF calculated in this way follows the proposed protocol for the AerChem-MIP (Collins et al., 2017). While RF does not capture the full alterations in the energy balance, ERF is more uncertain than RF because it involves multiple climate interactions (Forster et al., 2016). However, several studies have found that ERF and RF are nearly equal for many situations, and especially for increased concentrations in CO₂ and methane (Myhre et al., 2013).

To separate the effect of methane due to its influence on ozone and water vapour (indirect effects) from its effect on radiation (direct effect), we ran two more simulations for MAGICC with methane chemistry turned off (an all-forcing run and methane held at 1860 levels run with methane chemistry turned off for both). Equation (3) shows how the direct methane forcings were calculated for MAGICC; subtraction between methane-only forcings and the direct forcings yielded the indirect responses to methane. We also ran two more simulations for AM3–CM3 with methane radiation calculations or chemistry held constant beyond 1860, respectively. Equations (4) and (5), respectively, show how the direct and indirect methane forcings were calculated in AM3. While we only show forcing calculations here via AM3 simulations with fixed sea surface temperatures, we also ran the simulations for the fully coupled CM3 model.

$\Delta I CH_4$, direct (MAGICC) – I CH ₄ nochem – I CH ₄ 1860nochem (-	3)	3)
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 $\Delta F_{\text{CH}_4, \text{ direct (AM3)}} = F_{\text{AllForc}} - F_{\text{CH}_4 \text{rad} 1860} \tag{4}$

 $\Delta F_{\text{CH}_4, \text{ indirect (AM3)}} = F_{\text{AllForc}} - F_{\text{CH}_4\text{chem}1860}$ (5)

The global mean historical concentrations of CO_2 and methane used by the models to calculate radiative forcings and therefore temperature changes are shown in Fig. 1. (Note that concentrations are prescribed for MAGICC only through 2005, and then emissions inputs drive the model thereafter; however, the resulting concentrations from these emissions are consistent with that input into CM3.) Results for both models are presented as an average of the individual ensemble members (initial condition-driven ensemble members for CM3 and physics-driven ensemble members for MAGICC). Surface air temperatures are taken to be 2 m above the surface. For both models, we calculate temperature changes as the difference between temperatures in year *t* compared to that in 1860.

A key difference between AM3-CM3 and MAGICC is that the full GCM has internally generated unforced variability. This occurs both when the model is coupled (CM3) and when it is run with prescribed sea surface temperatures and sea ice (AM3). The variability can be dampened by applying a smoothing to the annual time series. However, too long of a smoothing period removes much of the decadal level forcing that we hope to uncover in this study. Therefore, we employ a 5-year smoothing average to AM3-CM3 results to filter out some of the internal variability. Additionally, to better quantify and isolate the role of unforced variability in the AM3-CM3, we run control experiments of each with fixed forcing. For CM3 we ran a 500-year control simulation with all radiative forcing held constant at 1860 levels. For AM3 we ran a shorter 200-year control run, with all radiative forcing held fixed at 1860 levels, with annually repeating monthly averaged sea surface temperatures and sea ice characteristics taken from 30 years of the CM3 control run.

3 Results and discussion

Here we analyse AM3–CM3's and MAGICC's radiative forcing and surface air temperature responses to changes in either CO_2 or methane in isolation.

Given that an important difference between AM3–CM3 and MAGICC results is the role of unforced variability



Figure 2. Anomalies in (a) all-sky top-of-the-atmosphere (TOA) net radiation (W m⁻²) with respect to the 200-year mean of the control simulation of AM3 and (b) surface air temperature (°C) with respect to the 500-year mean of the control simulation of CM3. For both panels, pre-industrial (1860) forcings are held constant post-spin-up. Fluctuations indicate unforced variability. Results are shown for annual averages (grey line) and 5-year running means (black line).

in AM3–CM3, we first analyse the magnitudes of unforced variability in both AM3 and CM3. Although initial condition-driven ensemble member means and/or running averages are employed to dampen out some of the variability in AM3–CM3, it still plays a large role in forcing and temperature responses. CM3, in particular, has been shown to produce magnitudes of variability on the upper end of CMIP5 models (Brown et al., 2015).

The results of the control simulations with constant preindustrial (1860) external forcings are shown in Fig. 2. In the case of AM3 unforced changes in net radiation at the top of the atmosphere for all-sky conditions range from -0.18 to 0.21 W m^{-2} with a standard deviation of 0.07 W m^{-2} for a 5-year running mean. We find the maximum swing between two consecutive 5-year means to be 0.35 W m^{-2} . Sources of unforced variability in AM3 include a mixture of land snow and ice cover variability, clouds, and just year-to-year variability in the meteorology; soil moisture may also play a role. For CM3, unforced internal dynamics yield temperature responses ranging from -0.27 to $0.24 \,^{\circ}$ C for 5-year running



Figure 3. Radiative forcings (W m⁻²) after stratospheric adjustment due to all-forcing (grey), CO₂-only (orange), methane-only (blue), and for both AM3 (dashed) and MAGICC (solid) model simulations. Methane forcing includes its direct as well as indirect effect from influences on chemistry. AM3 radiative forcings are "effective" radiative forcings (ERFs), include tropospheric adjustments as well, and are calculated at the top of the atmosphere (TOA). MAGICC radiative forcings are calculated at the tropopause. AM3 data are 5-year running means. Correlation coefficients between MAGICC and AM3 forcings are shown inset.

means with a standard deviation of $0.1 \,^{\circ}$ C. We find the maximum swing between two consecutive 5-year means to be $0.2 \,^{\circ}$ C. The variability is driven by interactions among the ocean–atmosphere–land systems. While unforced variability is a key component to modelling the climate system, it can mask or amplify responses to external forcings over short timescales (e.g. Brown et al., 2017). This makes it difficult to clearly assess responses to small external forcings, and provides further motivation for using simpler models like MAG-ICC for analysis of small forcing scenarios.

3.1 Radiative forcing

Figure 3 shows the global-mean radiative forcings (RFs) in response to the all-forcing scenario as well as forcings attributed to isolated CO₂ and methane concentrations, respectively. Note that AM3 forcings are taken as top of the atmosphere and include rapid adjustments in the troposphere in addition to the stratosphere, and therefore they are considered an effective radiative forcing, while MAGICC-derived RF is calculated at the tropopause and only considers stratospheric temperature adjustment. While these are the standard forcing calculation methods for both types of models, we emphasize that comparing values of AM3 ERF to MAGICC RF can only allow for comparisons in broad patterns and relative magnitudes, especially because of large variability in ERF values when averaged over 1- to 5-year timescales (Forster et al., 2016) as discussed above.

MAGICC methane and CO_2 isolated forcings are much smoother than that of AM3 because of the lack of unforced variability in MAGICC. Some unexpected features in AM3 forcings (such as negative forcings in the earlier years despite increasing atmospheric concentrations) are likely due to unforced variability. Using the MAGICC forcings as a benchmark for a signal due to forced changes only, we find that nearly all of the deviations of AM3 fall within the range of internal variability as derived from the control simulation: 0.35 W m^{-2} . However, despite the slightly different forcing definitions and the unforced variability in AM3, all results are strongly correlated between AM3 and MAGICC (allforcing r = 0.81, CO₂-only r = 0.96, CH₄-only r = 0.93).

In the present day (model year 2014), AM3 and MAGICC yield an all-forcing ERF and RF of 2.0 and 2.5 W m⁻², respectively; note that land use is held constant in this analysis. This is consistent with the IPCC (2013) values that show an all-forcing ERF of 2.3 W m⁻² in 2011 (Myhre et al., 2013). The magnitudes for the AM3 and MAGICC all-forcing radiative forcings are offset after 1960 (-1 W m⁻² in 1960). This is due to AM3's strong aerosol indirect forcing (Golaz et al., 2011) beginning around this time when aerosol emissions in the mid-latitudes increased rapidly (Lamarque et al., 2010).

Isolating CO₂ and methane's contribution to overall forcings (Fig. 3), MAGICC RF is reasonably consistent with the AM3 ERF evolutions throughout the 20th century. Preindustrial to present-day forcings for CO₂ and methane simulated by AM3 and MAGICC are similar to those given by IPCC (Myhre et al., 2013) (IPCC values: 1.68 from CO₂ emissions and $0.97 \text{ W} \text{ m}^{-2}$ from methane emissions in 2011 relative to 1750 levels), although there are important differences, including baseline years (1750 for IPCC and 1870 for AM3 and MAGICC in this study to match that from AM3) and time series of atmospheric concentrations (Myhre et al., 2013). While the same radiation expressions are used for IPCC and MAGICC for CO2 and methane atmospheric concentrations, the representation of tropospheric ozone chemistry and its radiation effects in MAGICC is extremely simplified due to hemispheric averages in a four-box atmosphere. For a short-lived climate forcer that is highly spatially variable, this is a vastly different treatment than that by the IPCC, which employs multi-model assessments for tropospheric ozone forcings. We find that our direct methane forcing in MAGICC in model year 2011 is 0.45 W m⁻², extremely close to the IPCC's forcing of $0.48 \,\mathrm{W}\,\mathrm{m}^{-2}$ from changes in methane concentrations alone (recall, however, different baselines) (Myhre et al., 2013). However, when methane interactions with other chemical species are accounted for, MAGICC estimates a forcing of 0.7 W m^{-2} attributed to changes in methane compared to the IPCC's value of 0.97 W m⁻² in 2011 (Myhre et al., 2013).

In MAGICC, methane's RF is consistently around half the value of that by CO_2 . In AM3, methane's ERF is much closer to that of CO_2 until the year 2000 and beyond, where they diverge. While this divergence is consistent with global atmospheric methane concentrations levelling off for about a decade in the mid-1990s to mid-2000s, before rapidly increasing from 2007 onwards (Fig. 1), further simulations are



Figure 4. Same as Fig. 3, but showing direct (purple) and indirect (from methane's influence on ozone and water vapour, green) forcings (W m⁻²).

required (such as more ensemble members or adjustments to input conditions) in order to determine if the close methane and CO_2 ERFs before 2000 are an artefact of unforced variability or a substantiated feature. Based on our analysis of unforced variability in AM3, it is quite possible that they are features of internal variability.

Methane's role in radiative forcing can be divided into direct contributions via warming by methane as a greenhouse gas, and indirect contributions via production of other greenhouse gases (mainly tropospheric ozone) as it oxidizes to CO_2 in the atmosphere. Figure 4 compares the direct and indirect methane forcings from MAGICC and AM3, calculated via Eqs. (3) and (4), respectively. The results from AM3 further highlight the role of unforced variability in complicating perceived forcings from small concentration changes; the seemingly large swings in AM3 forcings deviate from that of MAGICC by around 0.25 W m^{-2} at most, which is within the realm of unforced variability (see Fig. 2). While correlation coefficients show consistency between MAGICC and AM3 (direct r = 0.87; indirect r = 0.78), the strong variability in AM3 makes comparisons of magnitude difficult. MAGICC attributes around 35 % of methane's present-day total radiative forcing to indirect effects, similar to the IPCC's attribution of 34 % (Myhre et al., 2013). AM3 shows magnitudes of indirect forcings in the present day that are around 30-50%of the total methane forcing, depending on the year; this variation is due to unforced variability.

3.2 Global surface air temperature change

To build confidence in the simulation of surface air temperature by both MAGICC and CM3, we compare the model results with 20th century reconstructions of surface air temperature, of which several datasets are available. Figure 5 shows the historical global-mean surface air temperature responses to changes in all forcings in MAGICC and CM3 compared with NOAA and National Aeronautics and



Figure 5. All-forcing global-mean surface air temperature responses in degrees Celsius (°C) for CM3 (solid red line) and MAGICC (solid grey line) model simulations compared to observations by NOAA (+) (https://data.giss.nasa.gov/gistemp/, last access: 1 October 2018) and NASA (X) (https://www.ncdc.noaa.gov/cag/time-series/global, last access: 1 October 2018). All annual temperature anomalies shown as change from 20th century average for each dataset. Individual initial condition-driven ensemble members for CM3 runs shown in thin dashed red lines. Physics-driven ensemble-member range for MAGICC shown as shaded grey. CM3 data are 5-year running means.

Space Administration (NASA) time series of global surface temperature anomalies, freely available online (https:// www.ncdc.noaa.gov/cag/time-series/global and https://data. giss.nasa.gov/gistemp/, last access: 1 October 2018). Following NOAA's methodology (NOAA, 2017), we compute the 20th century average temperatures from MAGICC, CM3, and NASA and calculate the annual temperature departures from this baseline.

The two observational datasets are perfectly correlated (r = 1.00). MAGICC and CM3 both have high correlations with NOAA and NASA data, although MAGICC's are higher (MAGICC r = 0.92 (NOAA) and 0.93 (NASA); CM3 r =0.76 (NOAA) and 0.75 (NASA)). Consistent with Fig. 3, CM3 shows lower temperature responses post-1960 due to the strong effect of aerosols (Golaz et al., 2013). We note, however, that the "lingering" temperature response in CM3 to major volcanic eruptions is an artefact of the 5-year running mean smoothing process; this is why CM3 temperature responses to volcanic eruptions persist longer than they are seen to persist in the observational records and by MAGICC. This is not found, however, to considerably impact the correlation coefficients between the CM3 data and the NOAA and NASA data. Overall, the general temperature anomaly temporal patterns reveal that both models adequately reproduce surface air temperature, providing confidence in both climate models of differing complexity levels.

The global mean surface air temperature responses attributed to CO_2 and methane forcings are shown in Fig. 6, calculated via Eqs. (1) and (2), respectively. The correlations of the ensemble means (19 physics-driven ensemble



Figure 6. Global mean surface air temperature responses in degrees Celsius (°C) for CM3 (dashed line) and MAGICC (solid line) model-derived simulations – CO_2 -only (orange) and methane-only (blue). Individual initial condition-driven ensemble members for CM3 runs shown in thin dashed lines. Range for MAGICC physicsdriven ensemble members shown in shaded colours. CM3 data are 5-year running means. Correlation coefficient between MAGICC and CM3 temperature responses are also shown.

members for MAGICC and 3 initial condition-driven ensemble members for CM3) are extremely high (CO₂ r = 0.98; methane r = 0.92). Figure 6 also shows individual CM3 initial condition-driven ensemble members and the range of MAGICC responses from all 19 AOGCM calibrations; however, we do not include MAGICC's highest climate sensitivity physics-driven ensemble member as the responses were a clear outlier to the rest of the members. Further, recall that the equilibrium climate sensitivity in CM3 is larger than the mean and median in MAGICC, and therefore we expect differences in the ensemble member-averaged responses from this characteristic alone.

We find that both CM3 and MAGICC attribute a nearly 1 °C rise in temperature from 1860 to 2014 from rising CO₂ concentrations (CM3: 0.9 °C; MAGICC: 0.9 °C). For methane, CM3 suggests a rise of 0.5 °C and MAGICC suggests a rise of 0.4 °C, consistent with the larger methane forcing in CM3 (Fig. 3). It is important to note that cooling from aerosols mask some of the warming that we otherwise would be experiencing from CO₂ and methane, which is why the combined warming from CO₂ and methane is larger than to-day's observed warming.

Two major features of the temperature response to methane in CM3, that are not present in MAGICC, further highlight the difficulty of extracting a small signal (and with a small ensemble) given the size of the unforced variability (Fig. 6); methane's forcing is considerably smaller than that of CO₂, making it difficult to extract a temperature response from the variability. The first feature is a global mean cooling response to methane forcings around 1900 to 1915, which is strongly apparent in two of the three initial condition-driven ensemble members. This cooling is likely a lagged response to negative methane ERF (at most -0.15 W m^{-2}) from 1895



Figure 7. Regional surface air temperature responses in degrees Celsius (°C) for CM3 (dashed line) and MAGICC (solid line) model indirect simulations – CO₂-only (orange) and methane-only (blue). Individual initial condition-driven ensemble members for CM3 runs shown in thin dashed lines. Range for MAGICC physics-driven ensemble members shown in shaded colours. CM3 data are 5-year running means. Correlation coefficient between MAGICC and CM3 temperature responses are also shown.

to 1900, seen both in the direct and indirect methane forcings (Figs. 3 and 4). The second feature is a strong warming signal in response to methane from 1980 to 1995, followed by cooling through 2000; while this is consistent with AM3 RFs (Fig. 3), the feature is more pronounced in the temperature response. Both of these features fall within the range of annual temperature swings due to unforced variability in CM3 (at most around 0.2 °C for a 5-year running mean). Therefore, we cannot conclude that they are robust responses to methane, but rather serve as a further example of why CCMs are difficult to employ for small individual forcings and the need for large ensembles.

To dig into these features further, we analyse regional surface air temperature responses to CO_2 and methane isolated forcings (Fig. 7). Methane-induced cooling between 1900 and 1915 is strongest in the Southern Hemisphere and especially over Southern Hemisphere oceans. On the other hand, the large methane warming in CM3 around 1990 is most prominent in the Northern Hemisphere, over both land and ocean.

When the global mean responses are parsed out by region (Fig. 7), the highest surface air temperature responses to methane and CO_2 are found over land in the Northern Hemisphere, with temperatures from CO_2 rising by well over 1 °C from 1860 to 2014 in both models. There is high correlation between MAGICC and CM3 for all regions. We expect and find methane correlations between the two models to be slightly lower than CO_2 because methane has more complex chemical interactions in the atmosphere than CO_2 that introduce more degrees of freedom than CO_2 , and are also potentially more simplified in MAGICC. We also find that correlations in the Southern Hemisphere are lower than in the Northern Hemisphere, especially for methane.

As seen and discussed earlier in Fig. 3 forcings, there are several time periods when the methane temperature responses are comparable in magnitude to that by CO_2 in CM3 global mean and regional responses (Figs. 6 and 7). We see this for all initial condition-driven ensemble members, and it is consistent with AM3 RFs (Fig. 3). In the ensemble mean, the comparable warming magnitudes between 1940 and 1950 are consistent with the rate of growth of CO_2 concentrations slowing down while methane concentrations consistently increase (Fig. 1).

Also discussed earlier and in contrast to the CO_2 and methane concentration trends from 1940 to 1950, the methane concentration growth rate slows down in the 1990s while the CO_2 concentrations consistently increase (Fig. 1). This is reflected in the CM3 temperature trends in addition to forcings (Fig. 3) as a divergence in the magnitude of temperature responses between methane and CO_2 to where they stand in the present day, with CO_2 yielding twice as much warming in 2014 as methane (Fig. 6).

4 Conclusions

The purpose of this study is to enhance confidence in reduced-complexity climate models, in the context of simulating temperature responses to methane and CO₂ atmospheric concentrations. We use the freely available (but closed source code) and computationally efficient model, MAGICC, for our analysis, motivated by the need to determine a quick and accessible, yet reliable, method for analysing impacts of future changes in methane emissions on climate warming. Given that sophisticated coupled climatechemistry models are generally inaccessible, time-intensive, and often employ high internal variability, they can be unsuitable for analysis of methane mitigation strategies when emissions changes are small, when many mitigation scenarios are considered, and when a large extent of parameter sensitivities are to be investigated. Employing a model like MAGICC, rather than resorting to simple GWP metrics, would significantly enhance the accuracy of mitigation assessments while still using basic infrastructure and providing immediate guidance for decision making.

To determine MAGICC's reliability for methane analysis, we performed several sets of experiments using MAG-ICC and CM3 – all-forcing with both time-varying natural and anthropogenic forcings but with land use held constant,

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simulations where CO_2 and methane concentrations are held constant at 1860 levels, and a simulation to isolate methane indirect effects resulting from its influence on ozone and water vapour (for MAGICC, we turned off methane chemistry; for CM3, we held methane radiative effects at 1860 levels). We also ran simulations using the atmosphere-only version of CM3, AM3, to calculate radiative forcings in response to the four sets of experiments. Finally, we ran control simulations for AM3 and CM3 to determine the role of unforced variability in influencing climate responses.

Both CM3 and MAGICC models adequately reconstruct surface air temperature records from NOAA and NASA from 1860 through 2014, especially for 1950 onwards. For isolated forcings, overall temporal patterns were consistent between MAGICC and CM3 temperature responses to methane and CO₂, including for indirect effects via methane chemical reactions. Correlation coefficients were very high at 0.98 and 0.92 in the global mean for CO₂ and methane, respectively, with overall magnitudes consistent. We therefore conclude that MAGICC is able to reproduce the general isolated greenhouse gas forcing results (temporal patterns and magnitudes) of a more sophisticated coupled global climate model, providing confidence in the use of MAGICC for understanding the climate implications of methane mitigation analyses.

Further, we find that methane accounts for a considerable fraction of 20th century and early 21st century warming – roughly half that of CO₂'s warming response. However, there are some features present in CM3 results without parallels in MAGICC. The features are, however, consistent in magnitude with forcing and temperature fluctuations due to unforced variability, and therefore are unable to be classified as robust responses. This highlights how unforced variability present in sophisticated models can make it difficult to ascertain robust responses to small changes in multiple forcings individually, further justifying the use of a model such as MAGICC beyond pure accessibility. To overcome this challenge, a larger number of ensembles could be employed or simulations could be run with a quasi-chemistry-transport model (Deckert et al., 2011).

Overall, we find that reduced-complexity climate models, with the MAGICC model as an example, are able to satisfactorily match the global mean temperature response to increases in isolated greenhouses gases as simulated by the GFDL-CM3, a complex chemistry–climate model. Furthermore, we find that the prominent role of unforced variability in AM3 and CM3 makes it difficult to clearly assess climate responses to small forcing changes, ultimately supporting further use of models like MAGICC, that have little to no unforced variability, for analysing climate responses to future changes in methane emissions.

Code availability. The MAGICC v6 model is available for download at http://www.magicc.org/download (last access: 1 October 2018) upon registration. Only the executable file is

available, and not the source code. The user manual can be accessed at http://wiki.magicc.org/index.php?title=Manual_ MAGICC6_Executable (last access: 1 October 2018). Full model details along with 19 sets of AOGCM-calibrated parameters used here for the physics-driven ensemble members are found in Meinshausen et al. (2011a). We update the default values of methane and tropospheric ozone radiative efficiency and methane atmospheric lifetime to values in Myhre et al. (2013).

The atmospheric model component (AM3) source code for GFDL-CM3 is available here: https://www.gfdl.noaa.gov/ am3-model/ (last access: 1 October 2018). The ocean model component (MOPM5) source code for GFDL-CM3 is available here: https://www.gfdl.noaa.gov/mom-ocean-model/ (last access: 1 October 2018).

Data availability. Results from CM3–AM3 simulations and from the MAGICC model are available from Vaishali Naik (vaishali.naik@noaa.gov) and Ilissa Ocko (icko@edf.org), respectively, upon request.

Author contributions. IBO designed and performed the analysis, ran the MAGICC simulations, and composed the paper. VN and DP ran CM3 and AM3 simulations, and provided paper feedback and additional analysis.

Competing interests. The authors declare that they have no conflict of interest.

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