



# Supplement of

# Long-path measurements of pollutants and micrometeorology over Highway 401 in Toronto

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### S1. Description of spectrum fitting method

As mentioned in the main text, high resolution reference spectra were taken from the HITRAN database at 296 K and 101325 Pa when available. For species not available in the HITRAN database, the reference spectra were taken from the PNNL database at 25 °C. Before measurements, the wavenumber calibration was performed by fitting a function for the instrumental line shape (ILS) (Harig et al., 2005). The ILS function also takes the reference spectra and calculates reference spectra into spectra with the same spectral resolution, 0.5 cm<sup>-1</sup> in our case. A known gas is selected, water H<sub>2</sub>O in our case, and a spectral range is fitted by adjusting ILS parameters.

As described in the main text, "at the beginning of the measurement period, a stray light spectrum was recorded by pointing the spectrometer away from the retroreflector. This stray light spectrum accounts for radiation back to the detector from reflections by internal parts inside the spectrometer, i. e. not from the retroreflector array, and was subtracted from all the measurement spectra before performing further analysis." The software which collects spectra and does spectral fitting analysis is called OPUS\_RS. OPUS RS calculates the spectral fitting by a non-linear curve fitting algorithm. Spectral fitting analysis was mainly done in the process of making reference files for each pollutant. Spectral windows/regions and interference gases for each gases were determined by examining the reference spectrum and picking the absorption features. Spectral windows in previous studies were also examined and adjusted to get a good fit. For CO, CO<sub>2</sub>, N<sub>2</sub>O HCN, HCHO, and NH<sub>3</sub>, we used spectral regions reported in previous studies (showing in Table 1 in the main text), since the fittings were good. It was also critical to include interference gases relevant in the spectral window under analysis. After many rounds of adjusting windows and interfering gases, an "optimized fit" was determined by examining the measurement spectrum, fitted spectrum, correlation coefficients and residuals. To further improve fittings, baseline calculations were performed under given spectral windows and interfering gases. Three methods are included in the baseline algorithm of the OPUS\_RS software: offset, linear baseline correction, and standard baseline algorithm. "Offset" and "linear baseline correction" provide a linear baseline calculation, and they were used when the base of the spectral region was roughly straight. The "standard baseline correction" algorithm uses Gaussian functions to simulate the baseline. It has two parameters: distance and width (both in cm<sup>-1</sup>). "Distance" controls how far apart the centers of multiple Gaussian simulations are and how many Gaussians are within the range of spectral window. "Width" controls the width of each Gaussian function. "Distance" and "width" parameters were adjusted to optimize fitting. In this study, up to 3 Gaussian functions in the spectral window were used. For each fitting, the OPUS RS returns a coefficient of correlation. Users have an option of setting a threshold of this coefficient. As noted in Table 1, OPUS\_RS reports mixing ratios when the correlation coefficient of fitting is above this threshold. When the correlation is below this threshold, the pollutant is not "identified" and the mixing ratio is reported as zero.

The temperature dependence of the reference spectra was studied for CO, NH<sub>3</sub>, CO<sub>2</sub> and CH<sub>4</sub>. HITRAN spectra at 278 K, 298 K and 323 K were used in the OPUS\_RS. One HITRAN spectrum at one of the three temperatures was loaded into the reference file which contained all the fitting information to produce a reference file for that temperature. The measurement conditions were set to 296 K and 101325 Pa in OPUS\_RS, and measurement spectra were loaded by using the reference file at that temperature. The analysis was repeated by using reference files at three temperatures. After fitting the measurement spectra to reference spectra at the three temperatures, the difference in the extracted mixing ratio between

323 K and 278 K divided by the mixing ratio at 278 K was calculated. The maximum difference in retrieved mixing ratio for the 45 °C range was 8.9 % for  $NH_3$ , 4.2 % for  $CH_4$ , 8.3 % for CO, and 4.1 % for  $CO_2$ . This is a conservative estimate for the temperature related uncertainty, since the observed temperature range was much smaller than 45 °C.

Examples of measured spectra, fittings, and residuals are shown in Fig. S1.

#### S2. Scintillometer theory

The structure function of the optical refractive index  $C_n^2$  is determined from the fluctuations of the lightbeam intensities received by the scintillometer receiver. The structure parameter of temperature  $C_T^2$  can be derived from  $C_n^2$  given the ambient temperature, humidity, pressure, and wavelength (Thiermann and Grassl, 1992). The determination of the sensible heat flux H based on  $C_T^2$  needs additional assumptions based on Monin-Obukhov Similarity Theory (MOST).

H is defined as

 $H = -\rho C_p \theta_* u_*$  (S1) where  $\rho$  is the air density,  $C_p$  is the specific heat capacity of air at constant pressure,  $\theta_*$  is the temperature scale, and  $u_*$  is the friction velocity.  $\theta_*$  and  $u_*$  are determined by the MOST functions (Wood et al., 2013):

$$\frac{C_T^2 z^{2/3}}{\theta_*^2} = \Psi_H \left(\frac{z}{L}\right) \tag{S2}$$

$$\frac{\kappa z U(z)}{\ln(\frac{z}{z_0})u_*} = \Psi_{\rm M}\left(\frac{z}{L}\right) \tag{S3}$$

where the Obukhov length L is defined as

$$L = \frac{T}{\kappa g} \frac{u_*^2}{\theta_*} \tag{S4}$$

and z is the height above the surface,  $z_0$  is the surface roughness length, the von Kármán constant  $\kappa$ =0.4, g is the gravitational constant, and U(z) is the mean horizontal wind speed. The scaling functions  $\Psi_H$  and  $\Psi_M$  can be calculated by Eqs. (S5) and (S6) (Thiermann and Grassl, 1992) and Eqs. (S7) (Paulson, 1970) and (S8) (Businger et al., 1971):

$$\Psi_H = 4\beta_1 \left[1 - 7\frac{z}{L} + 75(\frac{z}{L})^2\right]^{-1/3} \qquad for \ \frac{z}{L} < 0 \tag{S5}$$

$$\Psi_{H} = 4\beta_{1} [1 - 7\frac{z}{L} + 20(\frac{z}{L})^{2}]^{1/3} \qquad for \ \frac{z}{L} > 0 \qquad (S6)$$

with  $\beta_1=0.86$ 

$$\Psi_{\rm M} = -2\ln\left(\frac{1+\chi}{2}\right) - \ln\left(\frac{1+\chi^2}{2}\right) + 2\arctan(\chi) - \frac{\pi}{2} \qquad for \ \frac{z}{L} < 0 \qquad (S7)$$

Where 
$$\chi = (1 - 15\frac{z}{r})^{1/4}$$

and 
$$\Psi_{\rm M} = 4.7 \frac{z}{L}$$
 for  $\frac{z}{L} > 0$  (S8)

These calculations were performed with the software provided by Scintec (SRun, version 1.31; see http://www.scintec.com/english/web/scintec/Details/A012000.aspx). The procedure to calculate the sensible heat flux H from measurements is as follows:

An initial  $|L|=|L_{ini}| = 1000$  m is assumed, where the signs of L and H are determined by the simultaneous measurement of the vertical temperature gradient  $\Delta T/\Delta z$ . Then  $\Psi_H$  is calculated from Eqs. (S5) and (S6) using L and path height z.  $\theta_*$  can then be calculated with  $C_T^2$  using Eq. (S2). Next  $\Psi_M$  is calculated from Eqs. (S7) and (S8). Friction velocity  $u_*$  is then calculated by Eq. (S3) given the measured wind speed (from the NAPS station) and  $z_0$ . A new L can then be calculated from Eq. (S4) using the calculated  $\theta_*$  and  $u_*$ . This process is then repeated iteratively until the change in L is smaller than 0.1. Finally, H is calculated from Eq. (S1) using the last calculated values for  $\theta_*$  and  $u_*$ . In the end, H,  $u_*$  and L were calculated at a 1-minute resolution in this study.

## S3. Multi-year O<sub>3</sub> measurements from a regular NAPS station

It is also recognized that 2 weekends are not statistically enough to show the difference of  $O_3$  mixing ratios between weekdays and weekends. The NAPS station which provided measurement results for this manuscript was a recently installed station, so it was there only for that summer. However, there is another nearby NAPS station (Station # 60430) which has been in operation for several years. NAPS station #60430 is located 175 m south of the highway 401 and about 200 m southwest of the NAPS research trailer in Fig.1. Therefore, we have extracted the  $O_3$  mixing ratio from Station # 60430 for July and August from 2013 to 2015, to show some multi-year statistics. The diurnal variations of  $O_3$  in July to August from 2013 to 2015 are shown in Fig. S7. The median of the diurnal average  $[O_3]$  on weekends from 7 am to midnight are higher than median on weekdays, which is consistent with the results from the new research NAPS trailer shown in the main text and Fig.9. There is no resolvable difference in 2013. Therefore, the difference between weekdays and weekends we observed in our 16-day measurement agrees with longer observations in 2014 and 2015, which suggests it is representative.

It is also interesting to compare the 2015  $O_3$  mixing ratio from the NAPS research trailer at the edge of highway with this regular NAPS station. The maximum median  $[O_3]$  on weekends (54 ppb) from the

trailer at edge was higher than median  $[O_3]$  at the regular station (50 ppb). Both stations show differences of 20 - 24 ppb in the median  $[O_3]$  between weekdays and weekends.

### References

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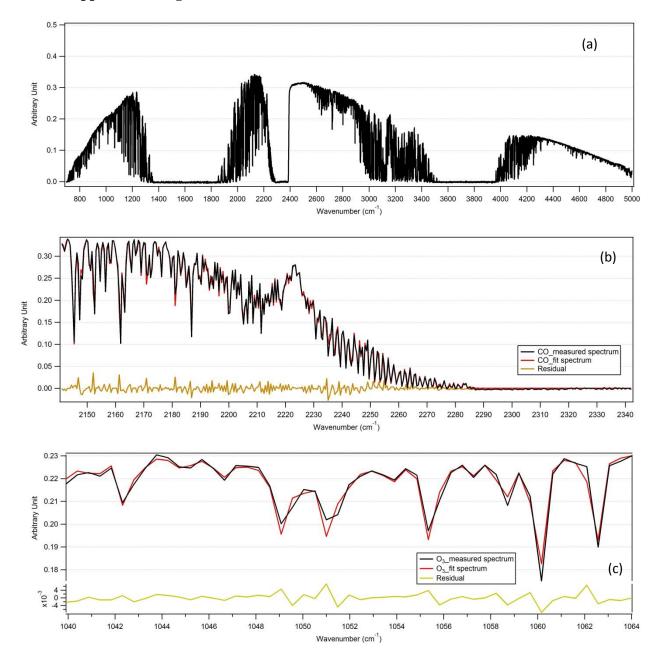
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Thiermann, V. and Grassl, H.: The measurement of turbulent surface-layer fluxes by use of bichromatic scintillation, Bound-Lay. Meteorol., 58, 367-389, doi: 10.1007/BF00120238, 1992.

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# **Supplemental Figures**



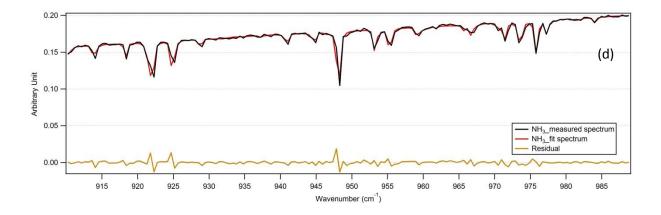


Figure S1: Examples of measured spectra at 16:01:45 on July 29, and fit modelled spectra made using the optimum spectral windows and OPUS\_RS non-linear fitting method. The black curve is the measured spectrum; the red curve is the fit spectrum; and the brown curve is the residual. (a) full measured spectrum; (b) Spectral range for analysis of CO; (c) Spectral range for analysis of O<sub>3</sub>; and (d) Spectral range for analysis of NH<sub>3</sub>.

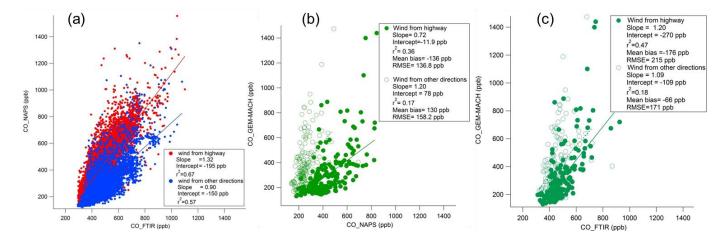


Figure S2: Comparison of CO mixing ratios: (a) NAPS and FTIR; (b) GEM-MACH and NAPS (c) GEM-MACH and FTIR. Red dots in (a) and solid circles in (b) and (c) show results when the wind was from the highway towards the NAPS trailer; Blue dots in (a) and open circles in (b) (c) show results when the wind was from other directions; the solid lines in (b) and (c) are the linear regression excluding 3 points when GEM-MACH results were much greater than measurement values; the dashed lines in (b) and (c) are the linear regression for these excluding the 4 most extreme outliers. Mean bias = [GEM\_MACH]-[measurements]. RMSE = root mean square error.

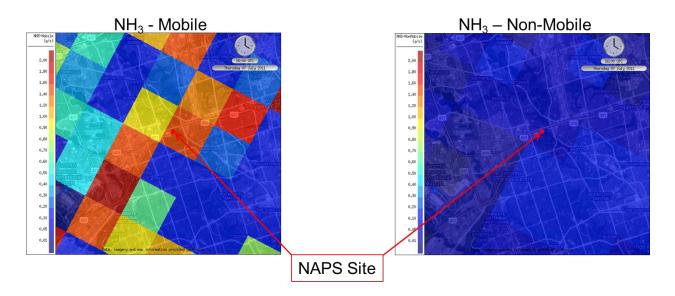


Figure S3: Gridded  $NH_3$  emissions field used by GEM-MACH. Each grid cell is 2.5 km by 2.5 km in size, and the red dot marks the location of the NAPS trailer. The color scale indicates the magnitude of the emissions of  $NH_3$  (g s<sup>-1</sup>) in July 2010 from on-road mobile sources (left) and non-mobile sources (right).

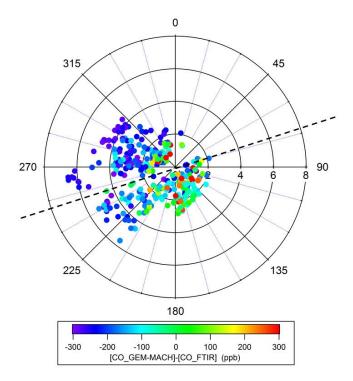


Figure S4: Polar plots of the difference of CO mixing-ratios between GEM-MACH output and FTIR measurements. Azimuth angle represents wind direction (meteorological convention), and radius indicates wind speed (m s<sup>-1</sup>). The color indicates the CO mixing ratio difference. The center corresponds to the location of the NAPS trailer. The black dashed line shows the orientation of the highway: above this line, the wind came across the highway to the trailer.

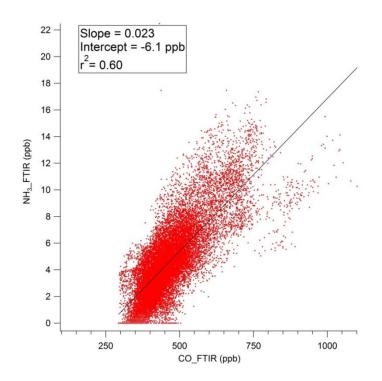


Figure S5: Linear relationship between NH<sub>3</sub> and CO mixing ratio from FTIR over the whole measurement period.

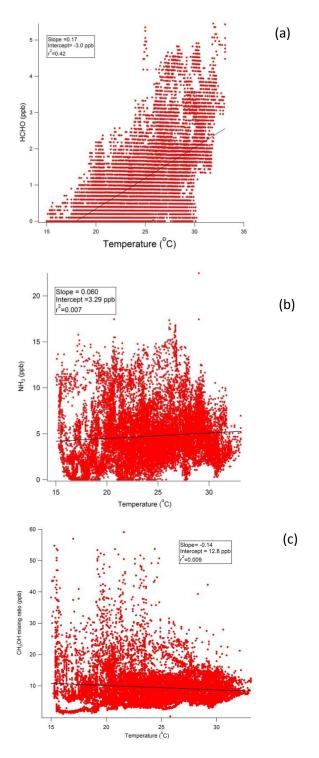


Figure S6: Correlation between mixing ratio of pollutants and ambient temperature for (a) HCHO; (b) NH<sub>3</sub>; (c) CH<sub>3</sub>OH

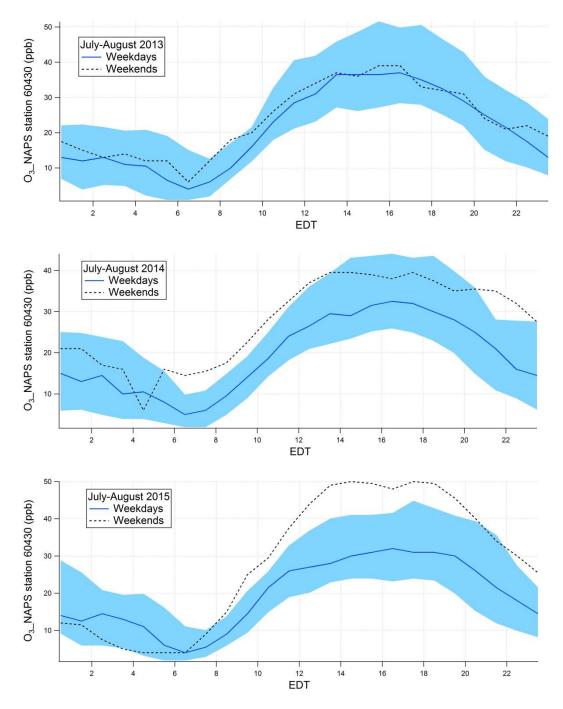


Figure S7: Average weekday and weekend diurnal cycles of mixing ratios of  $O_3$  from NAPS station# 60430 in July-August 2013 (top), July-August 2014 (middle), and July-August 2015 (bottom). Solid blue lines are medians and the shaded areas are the interquartile ranges on weekdays; dashed black lines are medians on weekends.