

Supplement of Atmos. Chem. Phys., 17, 13891–13901, 2017
<https://doi.org/10.5194/acp-17-13891-2017-supplement>
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Supplement of

Insight into the in-cloud formation of oxalate based on in situ measurement by single particle mass spectrometry

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19 **Instrumentation**

20 **SPAMS**

21 Individual particles are introduced into SPAMS through a critical orifice. They are
22 focused and accelerated to specific velocities, which are determined by two continuous
23 diode Nd:YAG laser beams (532 nm). Based on the measured velocities, a pulsed laser
24 (266 nm) downstream is trigger to desorp/ionize the particles. The produced positive and
25 negative molecular fragments are recorded. In summary, a velocity, a detection moment,
26 and an ion mass spectrum are recorded for each ionized particle, while there is no mass
27 spectrum for not ionized particles. The velocity could be converted to d_{va} based on a
28 calibration using polystyrene latex spheres (PSL, Duke Scientific Corp., Palo Alto) with
29 predefined sizes. The identified ion peaks have peak areas larger than 5 (arbitrary unit),
30 whereas the noise level is lower than 1.

31

32 **An discussion on the preferential association of oxalate within Fe-rich and Amine** 33 **particles**

34 As shown in Fig. 4, ~10% of oxalate was associated with Fe-rich particles, second
35 only to the K-rich particles. Regarding that the Fe-rich particles only accounted for $2.5 \pm$
36 0.4% of all the detected particles (Lin et al., 2017), it might reflect that the Fe facilitated
37 the formation of oxalate. Fenton reactions involving iron can produced more oxidants (e.g.,
38 $\bullet\text{OH}$) (Nguyen et al., 2013; Herrmann et al., 2015), which is an important factor for the
39 formation of oxalate (Ervens et al., 2014). While Sorooshian et al. (2013), Zhou et al.
40 (2015), and Cheng et al. (2017) have suggested that oxalate can be significantly lost

41 through the photolysis of iron-oxalato complexes. The difference between these
42 observations and this study might be attributed to the different radiation. Our observation
43 was conducted at a mountain site in winter, mostly covered with orographic cloud, resulted
44 in very low visibility (< 500 m), and thus low radiation was expected during sampling.
45 With sampling conducted on an aircraft, cloud water collected by Sorooshian et al. (2013)
46 included the below and top of cloud water samples, and thus photolysis is expected. On the
47 other hand, the highest fraction ($> 30\%$) of oxalate was found to be internally mixed with
48 metal-containing (e.g., iron, zinc, copper) particles in the Pearl River Delta region (Cheng
49 et al., 2017). The internally mixed oxalate and iron could account for $\sim 50\%$ of iron particles
50 at nighttime (Zhou et al., 2015). Additionally, oxalate was also found to be slightly
51 enriched in amine-containing particles, which is most probably attributed to the enhanced
52 partition of amine to wet aerosols (Rehbein et al., 2011; Zhang et al., 2012).

53 Table S1. Correlation analysis between the hourly detected number for species in
 54 cloud-free particles (N = 109) and *cloud RES* particles (N = 123). Most of the analysis
 55 shows significant correlation ($p < 0.001$) between the species, with the R^2 shown as follows.
 56 Results without significant correlation are marked with superscripts a and b.

57

	m/z -45	m/z -59	m/z -71	m/z -73	m/z -89	K-rich
m/z -45	1					
m/z -59	0.92/0.93	1				
m/z -71	0.77/0.33	0.92/0.35	1			
m/z -73	0.94/0.81	0.92/0.86	0.80/0.20	1		
m/z -89	0.22/0.32	0.38/0.45	0.46/0.12	0.33/0.64	1	
K-rich	0.52/0.58	0.33/0.59	0.21/0 ^a	0.57/0.72	0.05 ^b /0.41	1

58

59 ^a $p = 0.37$; ^b $p = 0.009$.

60 Table S2. Number fraction (%) of ion peaks for organic acids associated with all the
61 detected particles and K-rich particles, respectively.

62

63

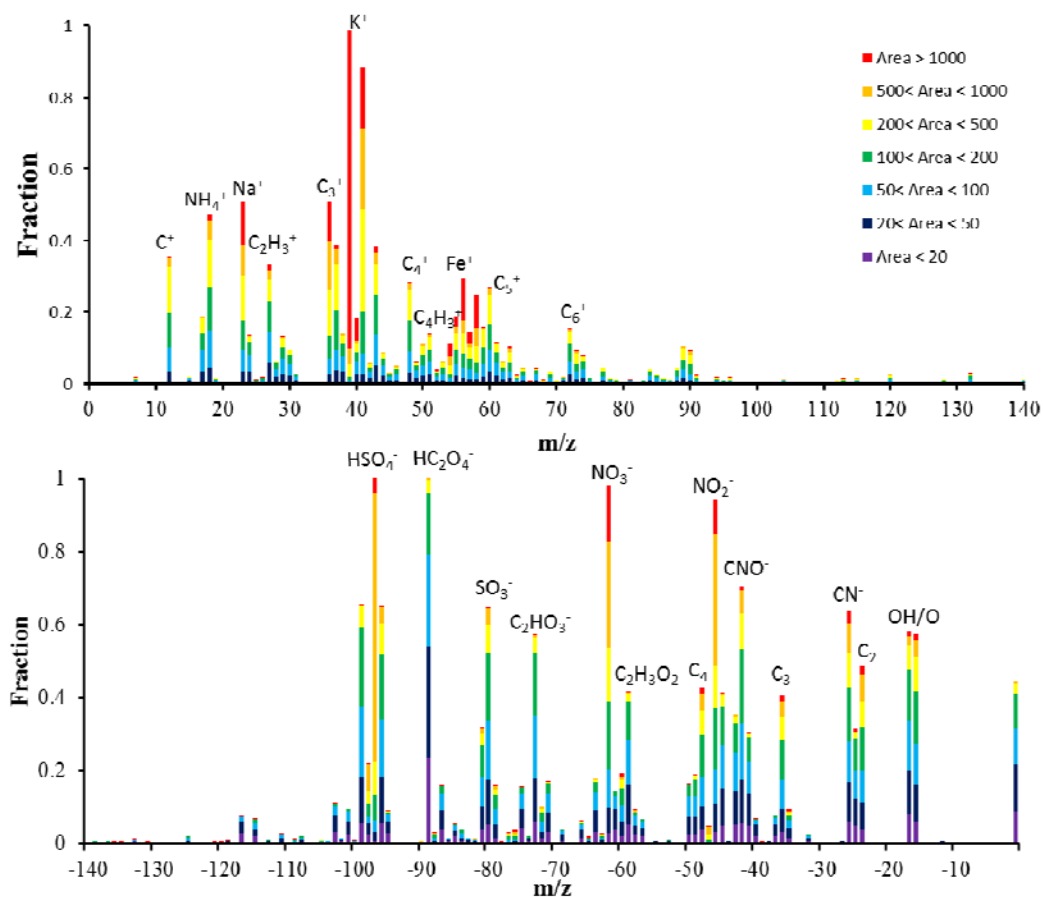
Ion peaks	All the detected particles (%)	K-rich particles (%)
m/z -45	12.4 ± 0.1	21.5 ± 0.3
m/z -59	8.6 ± 0.1	16.5 ± 0.3
m/z -71	2.8 ± 0.1	5.6 ± 0.1
m/z -73	12.6 ± 0.1	22.5 ± 0.3

65 Table S3. Number fraction (%) of the major OAs relative to all the detected particles, and
 66 visibility during each cloud event. Visibility was used here to indicate the cloud water
 67 content, since visibility is mainly controlled by the droplet number in cloud.

68

Ion peaks	Cloud I	Cloud II	Cloud III
m/z -45	16.5 ± 11.1	4.8 ± 1.2	8.6 ± 4.7
m/z -59	16.0 ± 9.6	3.9 ± 1.2	8.6 ± 5.5
m/z -71	8.7 ± 7.3	0.6 ± 0.4	4.0 ± 4.1
Visibility (km)	0.05 ± 0.03	0.31 ± 0.69	0.11 ± 0.17

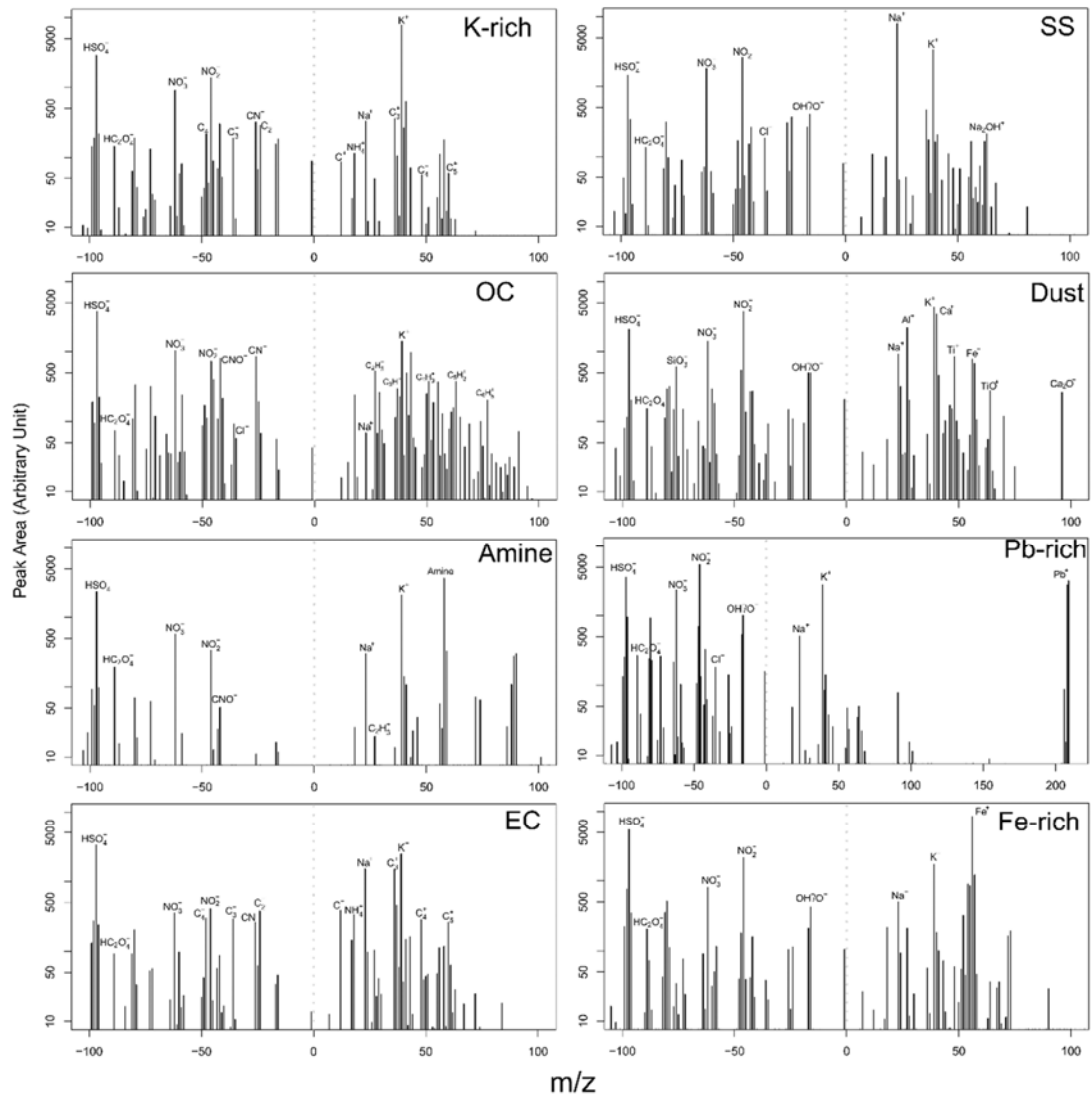
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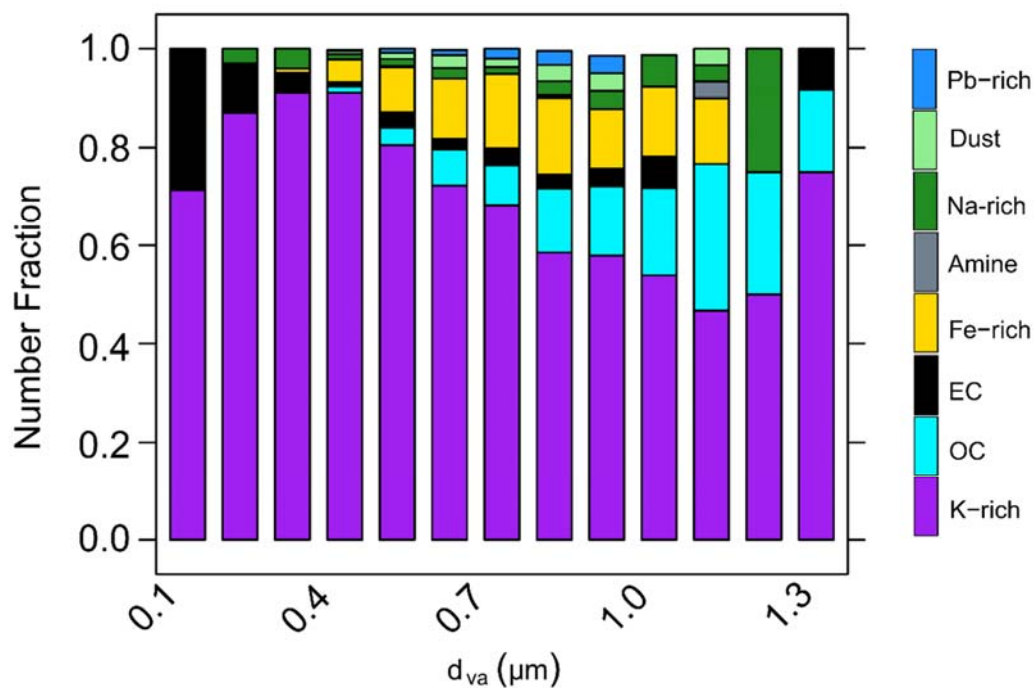
71 **Figure S1.** The number-based digitized mass spectrum of all the detected oxalate-
 72 containing particles. Compared with the number fraction of ammonium in Fig. 3, the
 73 result shows higher Nfs in oxalate-containing particles than ones in all particles, except
 74 m/z 18 (ammonium). As can be seen in Fig. 4, oxalate was dominantly distributed in K-
 75 rich particle type, which contained lower fraction of ammonium ($\sim 40\%$). However, as the
 76 dominant type in all the detected particles, EC type contained higher fraction ($\sim 80\%$) of
 77 ammonium. Therefore, the alkali nature (larger abundance of potassium, sodium) of the
 78 K-rich might explain the lower fraction of ammonium associated with the oxalate-
 79 containing particles.

80 (a)



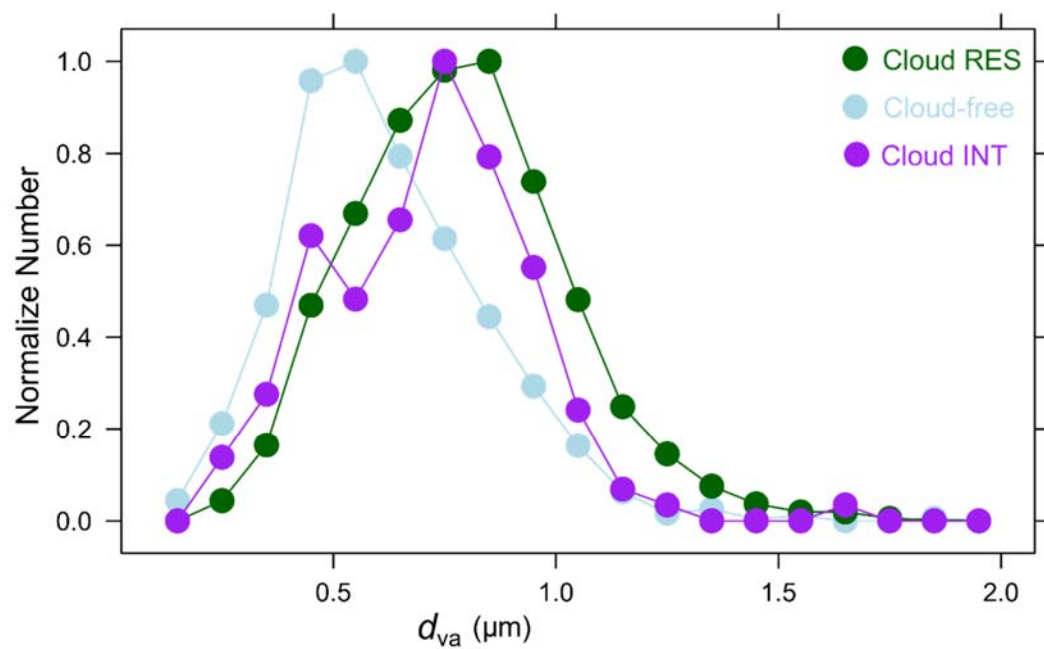
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82 (b)



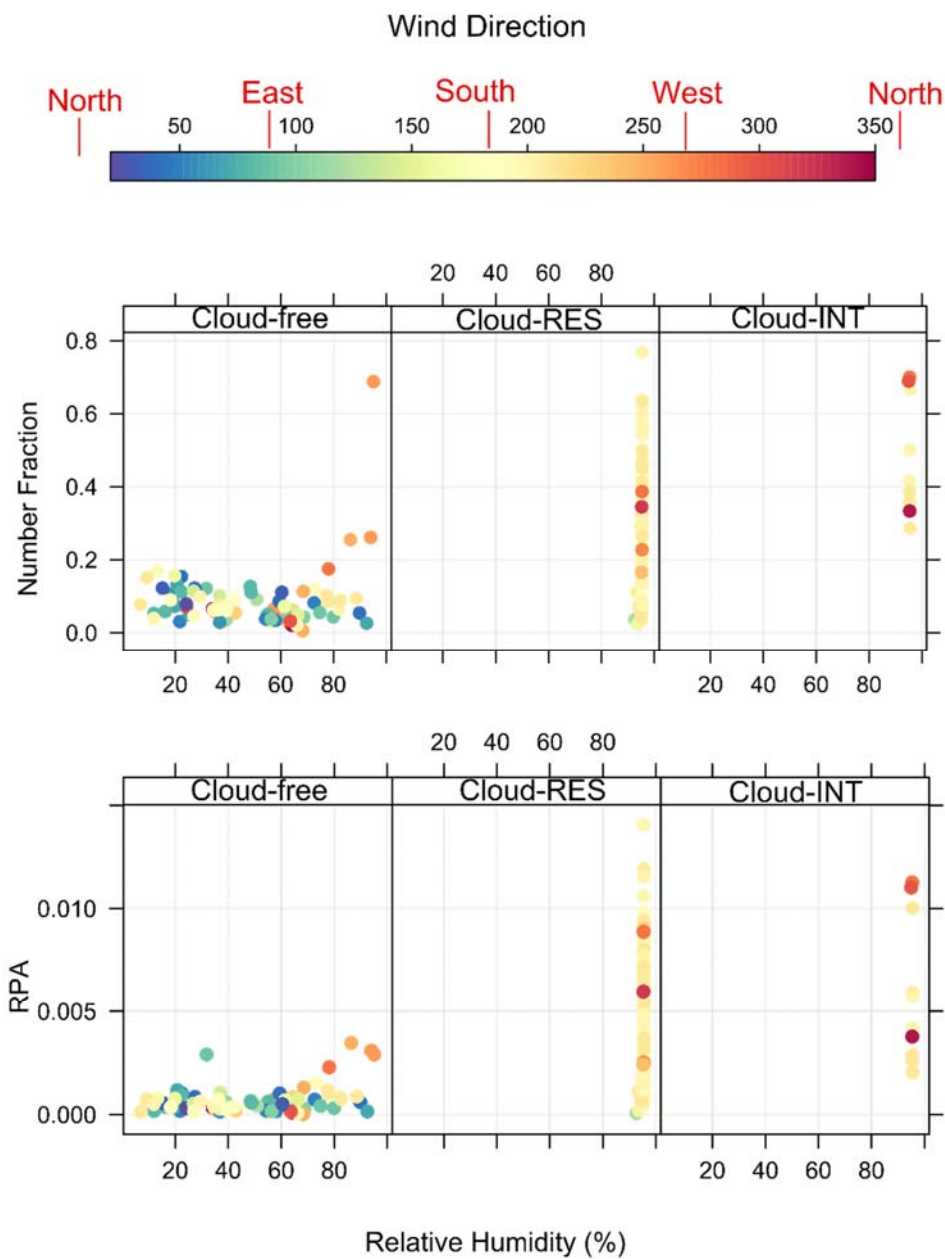
83

84 Figure S2. (a) Average mass spectra and (b) the size-resolved number fraction for each
85 particle type of oxalate-containing particles. Representative ions peaks were labeled for
86 each particle types. One may expect that oxalate at the largest d_{va} (1.3-1.4 μm) is
87 associated with aged sea salt and/or mineral dust particles. However, our result shows
88 that the aged biomass burning particles could contribute to the largest d_{va} (1.3-1.4 μm)
89 mode of oxalate. However, it shouldn't be conclusive since only 12 particles were found
90 at this size range.



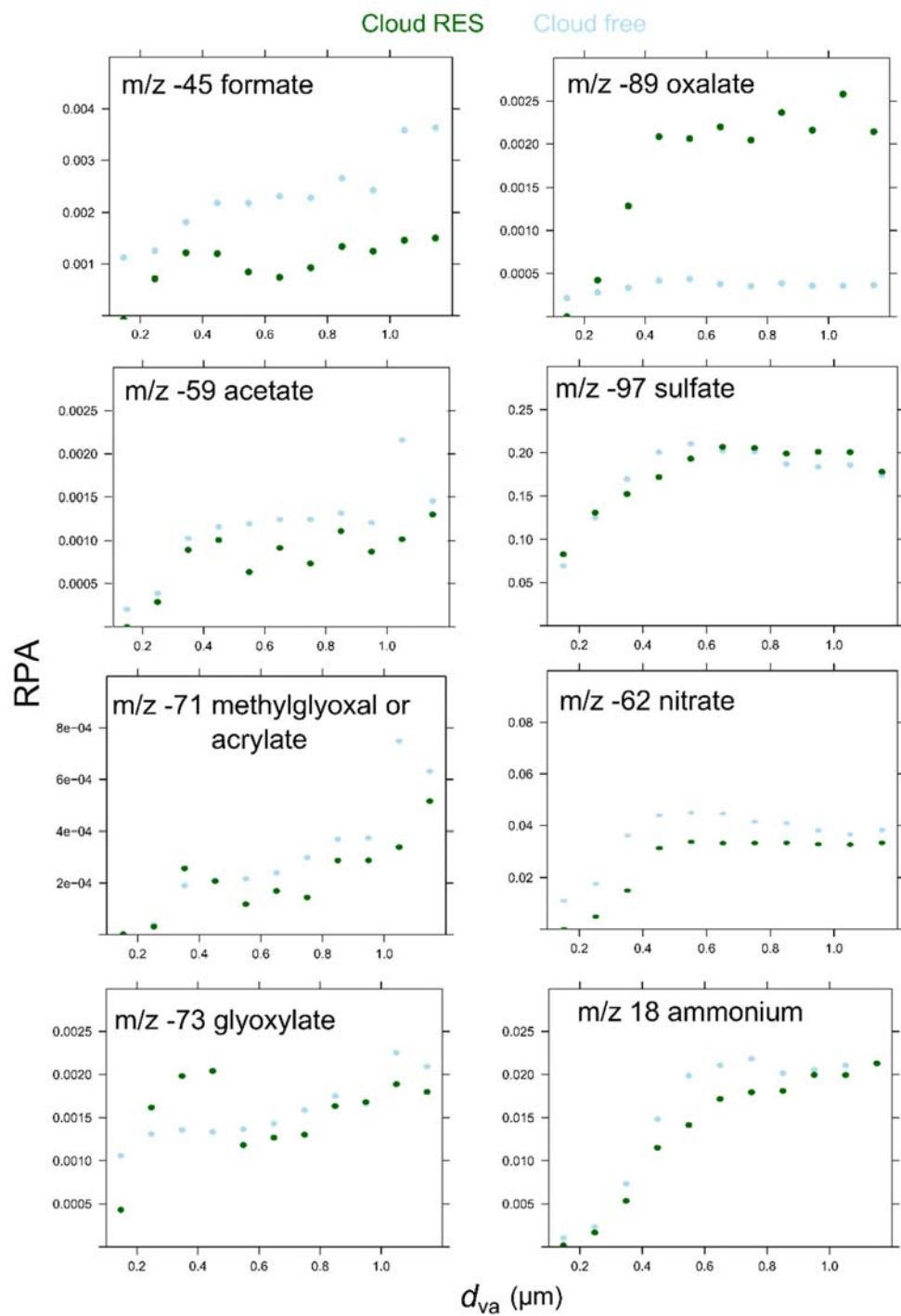
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92 Figure S3. The normalized unscaled number size distribution of oxalate-containing
 93 particles in cloud-free, RES, and INT particles, respectively.



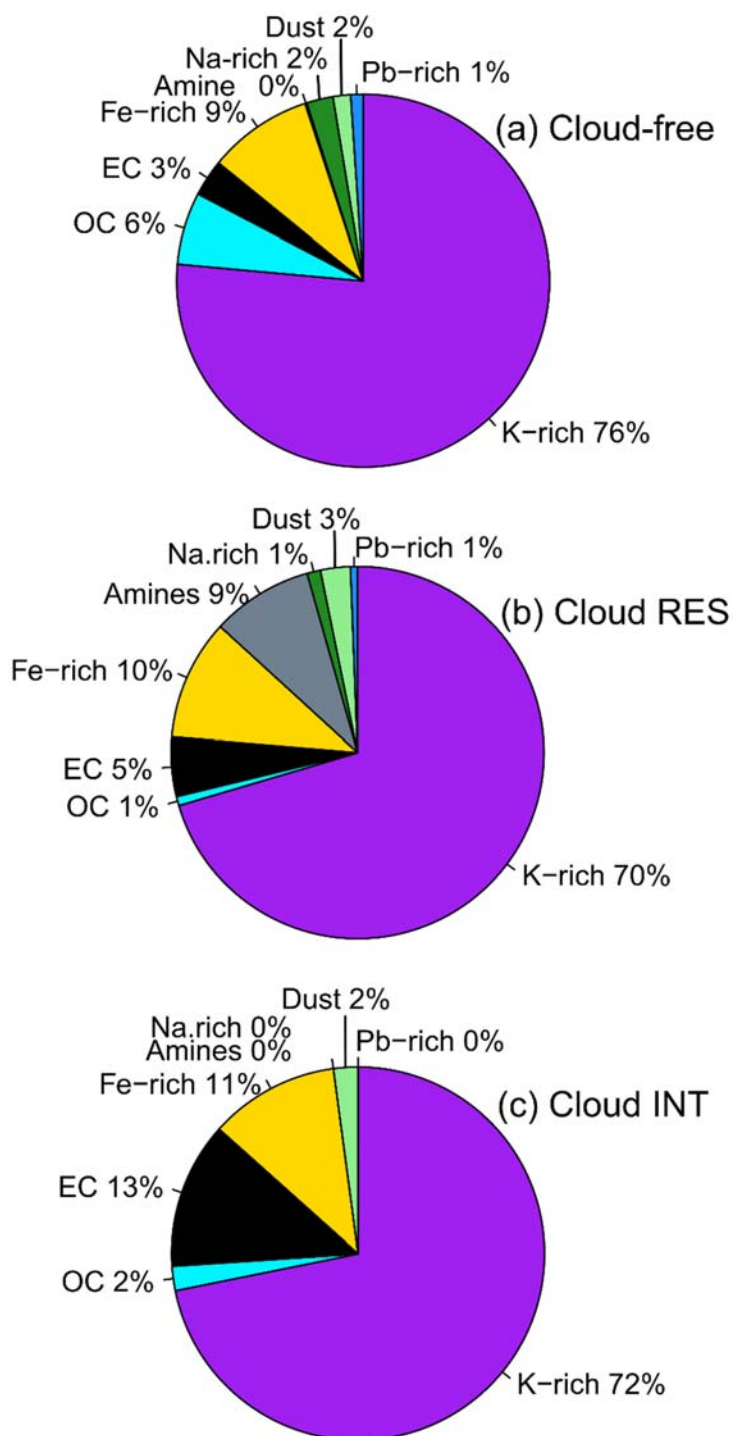
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95 Figure S4. Scattering plots of (upper) the number fraction and (lower) the RPA of the
 96 oxalate-containing particles versus relative humidity, separated for the cloud-free, cloud
 97 RES, and cloud INT particles . The coloration indicates the wind direction.



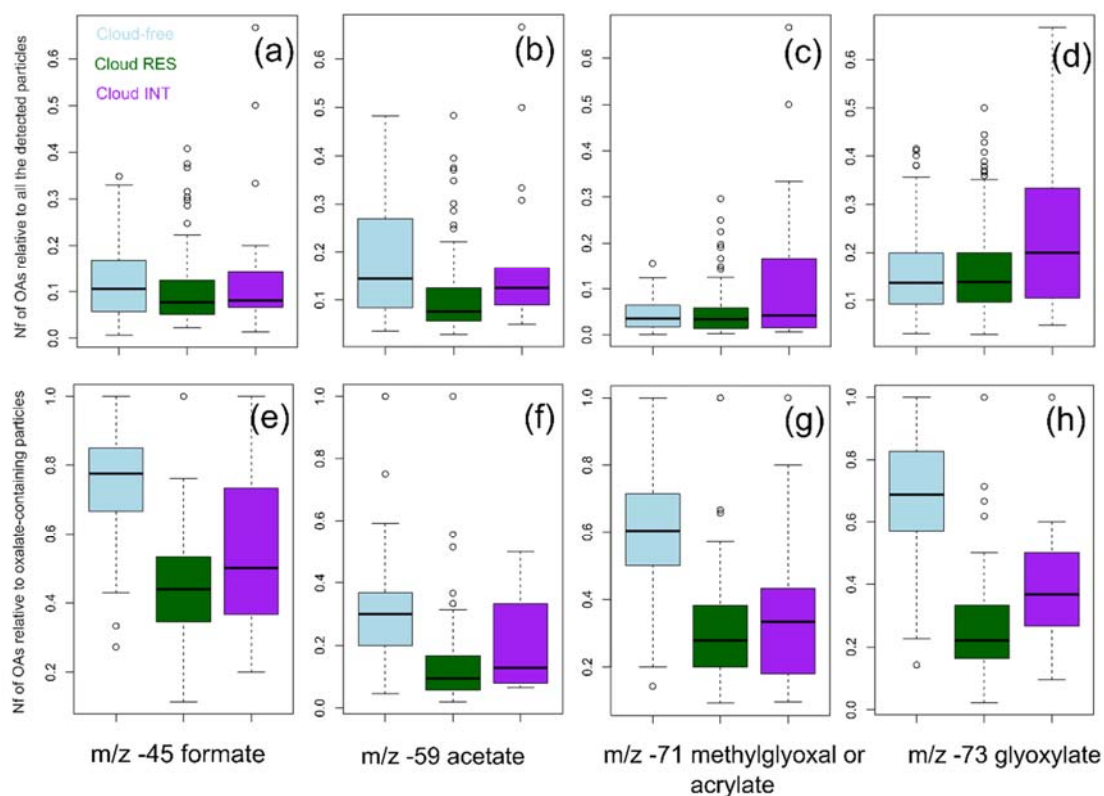
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99 Figure S5. Size-resolved distribution of RPAs for each species in the cloud-free and RES
 100 particles.



101

102 Figure S6. Number fraction of each oxalate-containing particle type for the (a) cloud-free,
 103 (b) cloud RES, and (c) cloud INT particles, respectively.



104
105

106 Figure S7. Box and whisker plot of the variations of number fractions for four OAs in (a-
107 d) all the detected particles, and (e-h) oxalate-containing particles, separated for cloud-free,
108 RES, and INT particles, respectively. In a box and whisker plot, the lower, median and
109 upper line of the box denote the 25, 50, and 75 percentiles, respectively; the lower and
110 upper edges of the whisker denote the 10 and 90 percentiles, respectively. Open circles
111 shows the data not included between the whiskers, which is larger than 90 percentiles or
112 lower than 10 percentiles of the data set.

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