

Aerosol optical properties and radiative forcing in the high Himalaya based on measurements at the Nepal Climate Observatory-Pyramid site (5079 m a.s.l.)

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Abstract. Intense anthropogenic emissions over the Indian sub-continent lead to the formation of layers of particulate pollution that can be transported to the high altitude regions of the Himalaya-Hindu-Kush (HKH). Aerosol particles contain a substantial fraction of strongly absorbing material, including black carbon (BC), organic compounds (OC), and dust all of which can contribute to atmospheric warming, in addition to greenhouse gases. Using a 3-year record of continuous measurements of aerosol optical properties, we present a time series of key climate relevant aerosol properties including the aerosol absorption (σ_{ap}) and scattering (σ_{sp}) coefficients as well as the single-scattering albedo (w_0). Results of this investigation show substantial seasonal variability of these properties, with long range transport during the pre- and post-monsoon seasons and efficient precipitation scavenging of aerosol particles during the monsoon season. The monthly averaged scattering coefficients range from 0.1 Mm^{-1} (monsoon) to 20 Mm^{-1} while the average absorption coefficients range from 0.5 Mm^{-1} to 3.5 Mm^{-1} . Both have their maximum values during the pre-monsoon period (April) and reach a minimum during Monsoon (July–August). This leads to dry w_0 values from 0.86 (pre-monsoon) to 0.79 (monsoon) seasons. Significant diurnal variability due to valley wind circulation is also reported. Using aerosol optical depth (AOD) measurements,

we calculated the resulting direct local radiative forcing due to aerosols for selected air mass cases. We found that the presence of absorbing particulate material can locally induce an additional top of the atmosphere (TOA) forcing of 10 to 20 W m^{-2} for the first atmospheric layer (500 m above surface). The TOA positive forcing depends on the presence of snow at the surface, and takes place preferentially during episodes of regional pollution occurring on a very regular basis in the Himalayan valleys. Warming of the first atmospheric layer is paralleled by a substantial decrease of the amount of radiation reaching the surface. The surface forcing is estimated to range from -4 to -20 W m^{-2} for small-scale regional pollution events and large-scale pollution events, respectively. The calculated surface forcing is also very dependent on surface albedo, with maximum values occurring over a snow-covered surface. Overall, this work presents the first estimates of aerosol direct radiative forcing over the high Himalaya based on in-situ aerosol measurements, and results suggest a TOA forcing significantly greater than the IPCC reported values for green house gases.

1 Introduction

Ambient aerosol particles play an important role in the overall energy balance of the atmosphere by scattering and absorbing incoming and outgoing solar and terrestrial radiation (the “direct effect”) and by modifying microphysical



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properties of clouds (the “indirect effects”) through their role as cloud condensation nuclei (CCN) and/or ice nuclei (IN). Current global mean estimates of direct anthropogenic aerosol radiative forcing (RF) at top of atmosphere (TOA) range from -0.63 to $+0.04 \text{ W m}^{-2}$ (IPCC, 2007). The mean value is significantly smaller than total greenhouse gas forcing of $+2.9 \text{ W m}^{-2}$ but the comparison of global average values does not take into account immense regional variability.

This is particularly true over Asia where anthropogenic emissions from industry, transport and incomplete fossil fuel combustion are rapidly rising (Richter et al., 2008), in particular in China and India. Intense emission of particulate matter at the surface increased by formation of secondary particles formed by condensation of gases leads to the formation of layers of particulate pollution as recently discussed by Ramanathan and Carmichael (2008). These layers of particulate pollution that can be observed also from satellites are often referred to the brown cloud (Lelieveld et al., 2001; Nakajima et al., 2007; Ramanathan et al., 2007a, b). The brown cloud refers to the ability of aerosols, such as black carbon from combustion processes, to not only scatter but also absorb solar radiation. Absorbing aerosols such as dust and black carbon (BC) and, to a lesser extent, brown carbon, which originate from incomplete combustion of fossil fuel or bio-fuel, can effectively absorb solar radiation and enhance atmospheric solar heating (Bond and Sun, 2005), and may contribute to atmospheric warming, in particular over world regions with intense anthropogenic emissions such as India. The Intergovernmental Panel on Climate Change (2007) recognizes BC as responsible for a global heating of approximately $+0.5 \text{ W m}^{-2}$, thus comparable to that of methane, which is the second most important next to CO_2 . Aerosol optical depth (AOD) in visible wavelengths measured over the most polluted regions in India ranges between 0.4 and 0.7 as compared to much lower AODs of 0.05 found in unpolluted air over the Indian Ocean (Welton et al., 2002). According to Ramanathan et al. (2005), absorbing aerosols over India have masked as much as 50% of the surface warming due to the global increase in greenhouse gases (GHG). Atmospheric forcing caused by the aerosol layer ranging from $+10$ to $+20 \text{ W m}^{-2}$ are calculated for the Indian subcontinent, in particular during the dry winter period (January–April).

Work performed within the INDIAN Ocean Experiment (INDOEX) also has revealed that this haze layer can be efficiently transported far beyond the source region, particularly during December to April and detected all over the Indian sub-continent (Clarke et al., 2002). Due to general circulation patterns, the Himalayan area is a strong receptor of the Indian/Pakistan source area. Atmospheric aerosols composed mainly of dust and black carbon from regional emissions, builds up over the Indo-Gangetic Plain, and is lifted up by the Himalaya orography. Lau and Kim (2006) suggest that absorbing aerosols in the elevated regions of Hindi-Kush-Himalaya HKH heat the mid-troposphere by absorbing solar radiation. Heating produces an atmospheric dynamical

feedback – the so-called elevated-heat-pump (EHP) effect resulting in increased precipitation over much of India.

Transport of optically active material to the very sensitive regions of the Himalayas is therefore a key issue to improve the description of the BC effect on Indian Summer Monsoon and, in particular, its impact on precipitation in the HKH regions and therefore on frozen water storage. In fact, BC not only affects the energy budget of the atmosphere, but deposits to snow surfaces, absorbs light, thus decreasing the albedo of the snow and modifying the energy budget of snow surfaces (Flanner et al., 2009 and 2007). Flanner et al. (2007) determined that the addition of 500 ppb (by weight) of black carbon to snow decreased its visible albedo from 0.98 to 0.88, and calculated that the instantaneous forcing over the Tibetan Plateau, due to the presence of BC in snow, exceeds 20 W m^{-2} in some places confirming that snow darkening is an important component of carbon aerosol climate forcing. This in turn affects snow melt, the duration of the snow cover and the seasonal availability of water. Predicting the effect of atmospheric aerosol loading on snow optical properties and energy budget is therefore essential to relate atmospheric aerosol loading and the seasonality of water flow.

The issue of the impact of optically active particulate matter on the local energy budget in the Himalaya has been put forward by recent observations showing that pollution aerosol from India and Pakistan can be transported by mountain breezes up to the high altitude (Carrico et al., 2003; Ramanathan et al., 2007; Bonasoni et al., 2008; Venzac et al., 2008; Komppula et al., 2009). Measurements of aerosol chemical composition and aerosol optical depth in the Nepal Himalaya have clearly shown the build up of aerosols in the pre-monsoon season during the winter and early spring, with relatively high values of light absorbing particulate matter including dust and black carbon (Carrico et al., 2003). Very recently, new insight into the mechanisms of aerosol transport from dust and pollutions sources in central and south-eastern Asia to the Tibetan and Himalayan regions was also provided by Huang et al., 2007 and Ramanathan et al., 2007. These studies indicate that aerosol in-situ and columnar concentrations over the south slope of Himalayas are strongly affected by the development of the boundary layer, responsible for transporting pollution aerosols upward, contributing to changes in solar irradiance (Dumka et al., 2006; Dumka et al., 2008; Ramachandran, 2008). Assessing the impact of aerosol particle on the local energy balance in the high altitude regions of Himalayas is therefore of importance for the whole area of HKH regions.

Despite efforts organized in the framework of the Atmospheric Brown Cloud Asia project and the INDIAN Ocean Experiment (Ramanathan et al., 2005; Lelieveld et al., 2001), there still exists a general lack of information on the spatial and temporal variability of aerosol optical properties in the Himalaya to adequately estimate the impact of aerosols on the atmospheric energy budget in the region. Results from measurement campaigns have been reported in a few studies

(Carrico et al., 2003; Pant et al., 2006; Dumka et al., 2006) but are often related to more populated areas. The only long-term study (1.5 y) of aerosol composition and optical properties in the Higher Himalayas was that of Carrico et al. (2003), conducted at Langtang (3920 m a.s.l.) at an altitude still below the ice front of glaciers. This study showed elevated concentrations of PM_{2.5} especially during the February–May season ($59 \pm 61 \mu\text{g}/\text{m}^3$) paralleled with substantial extinction (aerosol optical depth $\delta 500 \text{ nm} = 0.37 \pm 0.25$) with a dominance of both organic and elemental carbon from combustion sources as well as elements associated with dust emission. Additional measurements recently reported by Kompula et al., 2009, Hyvarinen et al., 2009, Kivekäs et al., 2009 for research stations located in India (Mukteshwar, 2180 m) and China (Mt. Waliguan, 3816 m) confirmed presence of elevated concentration of particles lifted up to high altitude by both local wind dynamics and regional/long range transport. Hyvarinen et al. (2009) focussed more specifically on the optical properties of particles at Mukteshwar station showing strong seasonal dependence of aerosol concentration and average value of the measured single scattering albedo of 0.81 at 525 nm, indicative of presence of substantial amount of absorbing material.

Recent work by Bonasoni et al. (2008) and Venzac et al. (2008) confirmed that pollution episodes are detected even further up in the glacierized area of Nepal, using measurements performed at the newly built Nepal Climate Observatory-Pyramid station (5100 m a.s.l.). In this paper we present observations of aerosol optical properties from NCO-P (Nepal Climate Observatory-Pyramid) with a 3-year observational record of aerosol light absorption and scattering coefficients, as well as aerosol optical depth (AOD). The objective is to determine the variability of aerosol optical properties in relation to air mass origins, and to derive statistically relevant information to describe the most common conditions observed at NCO-P. In addition, observed values of the scattering and absorption coefficient combined with the Aerosol Optical Depth (AOD) are used to estimate the direct aerosol radiative forcing for the region.

2 Site and instruments description

The Nepal Climate Observatory-Pyramid (NCO-P) station has been installed in March 2006 as part of Ev-K2-CNR “SHARE-Asia” (Stations at High Altitude for Research on the Environment) and UNEP “ABC” (Atmospheric Brown Clouds) projects. The station is located in the southern Himalayan region at the confluence of the secondary valley of Lobuche (oriented NNW-SSE) and the main Khumbu valley, on the top of a hill close (200 m distance and 100 m above) the ABC-Pyramid Observatory (Nepal, 27.95 N, 86.82 E, 5079 m a.s.l.). The nearest village from the station is Lobuche, a small village (100 inhabitants, 4850 m) located 5 km away. The largest village in the Valley is Namche-

Bazar (1000 inhabitants, 3440 m) located 25 km away. The Pyramid station is occupied year-round by station personnel but emissions are extremely reduced as power is mostly produced by photovoltaic cells. We can therefore exclude any artefact from local contamination in the measurement record. A complete description of the station can be found in Bonasoni et al. (2008) and Bonasoni et al. (2010).

The station was equipped to perform continuous measurements of chemical (organic and inorganic, soluble and insoluble), physical (mass and number size distribution) and optical (absorption and scattering coefficients) properties of aerosol. Surface O₃ and climate altering halocarbon concentrations are also measured at ABC-Pyramid. Aerosol sun photometry studies are carried out as well within the AERONET (AErosol RObotic NETwork) program (Holben et al., 1998). In the present paper, we will focus on optical aerosol properties derived from integrating nephelometer and absorption photometers with a data record extending from March 2006 up to March 2009.

Aerosol total and back scattering coefficients at three wavelengths (450, 550 and 700 nm) are derived by an integrating nephelometer (model TSI 3563), installed in March 2006. A PM_{2.5} cyclone before entering the nephelometer limits sampling to aerosol particles with aerodynamic diameter less than 2.5 μm . We have followed operating procedures described in Anderson and Ogren (1998) in particular for calibration procedures and corrections for truncation errors. One limitation, however, was that CO₂ calibration procedures could only be performed once/year during the yearly mission of expert personnel from France and Italy. Relative humidity in the Nephelometer never exceeded 30% even during Monsoon periods thus not affecting the scattering coefficient (Fierz-Schmidhauser et al., 2010). Measurements are performed at 301 min^{-1} with a time integration of 5 min, averaged every hour. Over the course of the 3-year period, a number of problems limited the measurement record to the following periods: from 1 November 2006 to 10 February 2007, from 10 December 2007 to 20 February 2008 and from 15 November 2008 up to the end of the measurement record used in this study (March, 2009). In fact, instrument failures occurred during the autumn and could not be fixed until late winter of the following year.

The aerosol light absorption at 670 nm was measured with a Multi-Angle Absorption Photometer (MAAP 5012, Thermo Electron Corporation). Measurements are performed below a PM₁₀ inlet with a time-integration of 5 min and are averaged every hour. Black carbon concentrations are calculated using a mass absorption coefficient of 6.6 gm^{-2} as recommended by Petzold and Schönlinner (2004). Additional information on MAAP measurements, calibration procedures and comparisons with Elemental Carbon EC measurements are provided by Marinoni et al. (2010).

Overall, concurrent measurements of absorption and scattering coefficients (σ_{ap} and σ_{sp} , respectively) are available for slightly less than 20 000 h of measurements. Unfortunately,

seasons are not evenly sampled due to nephelometer failures with a majority of concurrent absorption/scattering coefficient samples during the post-Monsoon season (October/November) and the pre-Monsoon and Monsoon seasons (April/May and June/September) while limited information is available for winter season (December/February).

Overall, the 3-year record consists of slightly less than 20 000 concurrent hours of measurements for σ_{ap} and σ_{sp} representing more approximately 75% of the maximum sampling time available. Due to nephelometer failures encountered predominantly during the winter period, sampling has been performed mostly during Pre-, post-, and Monsoon periods. In fact, the number of points available during the winter season has been considered too low to be included in data reduction of the present paper. The entire available record has been used to calculate diurnal and seasonal variations. The number of observations used to derive average information for specific conditions of air masses encountered at NCO-P is a sub-sample of the entire record.

In addition, we have used in this paper, measurements from a Cimel CE 318 sun photometer. It is an automatic sun-tracking and sky radiometer for measuring the aerosol optical depth at 8 wavelengths between 340 and 1020 nm. This instrument is part of the AERONET project, registered as number 367. Both sun photometer data and specifications can be downloaded from the AERONET web site (<http://aeronet.gsfc.nasa.gov>, station EVK2-CNR). See Gobbi et al. (2010) for further information. Determination of aerosol properties is completed by a Scanning Mobility Particle Sizer (SMPS) measuring aerosol number concentration and size from 15 to 500 nm in diameter. The description of the SMPS system can be found in Villani et al. (2007)

Finally, all meteorological information are derived from a Vaisala WXT510 meteorological unit, measuring temperature, pressure, relative humidity, rain, wind intensity and direction. All measurements presented in this paper are available at the following url (<http://www.rrcap.unep.org/abc/data/abc/ncop.html>). It should be noted that data used in this paper have been manually checked and in the case of a doubt concerning data quality removed through an analysis of consistency with the other data.

3 Synoptic and local meteorology and air mass classification

Synoptic meteorology of the Khumbu Valley area is discussed by Bonasoni et al. (2010). For the purpose of the present paper, we can summarize synoptic meteorology at NCO-P as follows: (1) Four seasons are identified thereafter defined as pre-Monsoon season (March to May), Monsoon season (June to September), post-Monsoon season (October to November) and winter season (December to February), (2) Origin of air masses during all seasons is predominantly from the West, over Middle East/central Asia, except during Mon-

soon season for which a transport from the South is dominant. Bonasoni et al. (2010) discussed the air mass classification at NCO-P using a hierarchical cluster analysis to group the trajectories into clusters. They identified 3 main clusters corresponding to (1) SW south-westerly (SW-AP: Arabian Peninsula, SW-AS: Arabian Sea, SW-BG: Bengal Gulf); (2) W westerly (W-NA: North Africa, W-EU: Europe, W-ME: Middle East) and REG regional.

In addition to synoptic scale meteorology, NCO-P is under the influence of orographic induced local circulation. The most common pattern is characterized by upslope wind bringing the air masses from the low-altitude Nepal plains during the afternoon. Opposite during night-time and morning hours, with down-slope wind bringing the air aloft from the free troposphere to the station. The local wind-slope system weakens during the monsoon season. Most atmospheric signals measured at NCO-P are strongly influenced by slope winds. In particular, the diurnal variation of anthropogenic species such as BC regularly exhibits very low night-time concentrations and daytime values peaking in the middle of the afternoon. Consequences of local circulation on aerosol dynamics are discussed by Venzac et al. (2008) and Panday et al. (2009). Local circulation is not considered in the cluster analysis of Bonasoni et al. (2010).

As will be discussed later, both large scale transport and local meteorology influence the atmospheric signals at NCO-P. However, the variability of aerosol optical properties does not appear to be linked to back-trajectories using cluster classification defined by Bonasoni et al. (2010). Therefore, in this paper, we have not based data reduction on air-mass backtrajectory analyses. Instead, it is observed that the variability of σ_{ap} and σ_{sp} is primarily driven by both synoptic scale meteorology (seasonal dependency) and local scale dynamics. The data reduction process has therefore been done with the following distinctions:

1. Background periods (BG) corresponding to air flowing from the Tibetan plateau and characterized by very low concentration values. BG samples are selected on the basis of the local dynamics when wind flows from $<90^\circ$ and $>270^\circ$ at NCO-P, and restricted to early morning periods (5 a.m.–9 a.m. (local time)). Night-time samples were not considered to avoid artefact by recirculation of pollutants emitted the night before, which spend the night in elevated layers over the valley. Over the duration of the entire record, data sampled during background periods represent about 2000 h of measurements (approximately 60% of the maximum sampling time available in the period). BG samples are predominantly available for Pre-Monsoon period (48% of observations) as respect to both Monsoon and post-Monsoon periods (33% and 18%, respectively).
2. Regional pollution (RP) periods correspond to air flowing from the Nepal plains (Panday et al., 2009). As

mentioned before, these conditions are driven by the timing of thermal-induced ventilation from the valley showing strong day-to-day similarity, with a strong westerly wind blowing through the valley from late morning until dusk. The timing of the up-slope flow initiation depends on the seasons but takes place at NCO-P between 9 am and 11 am. In order to avoid sampling BG air, RP samples are selected with wind flows between 90° and 270° and restricted to afternoon periods (3 p.m.–6 a.m. (local time)). Overall, RP samples represent approximately 1600 h of measurement and 65% of the maximum sampling time available. As for BG samples, RP samples are predominantly from Pre-Monsoon periods.

- In addition to BG and RP samples, we have defined samples called “Special Events (SE)”. Special events characterize situations of long-range transport of pollution not necessarily linked to mountain breeze and identified based on scattering and/or absorption coefficient afternoon values larger than twice the median coefficients for afternoon periods in the same season and corresponding to elevated concentrations of particles (from SMPS measurements), BC (from MAAP) and PM_{10} , $PM_{2.5}$ and PM_{10} (from GRIMM) as respect to typical average values for the season.

Over the 3-year record, we have identified 32 SE events representing almost 1600 h of measurements, mostly from pre-Monsoon period (11 events but 53% of the measurements). Overall, we believe both the analysis of seasonal and diurnal variations and the derivation of average values for representative conditions at NCO-P is based on a very representative number of samples. In addition, no sampling bias is affecting the data set besides the fact that the winter season is not considered in the study. In any case, it represents the longest record available of aerosol optical properties for this area of the globe.

4 Seasonal and diurnal variations of scattering, σ_{sp} , and absorption, σ_{ap} , coefficients and single-scattering albedo, w_0

4.1 Seasonal variations of measured properties

Monthly averaged scattering (σ_{sp}) and absorption (σ_{ap}) coefficients over the course of the record are – provided in Fig. 1a and b. They are provided at 700 nm and 670 nm, respectively with their mean and standard deviation. We chose to present σ_{sp} at 700 nm only because this wavelength was used to calculate the dry single scattering albedo (w_0). The good correlation between EC and BC shown in Marinoni et al. (2010) and Decesari et al. (2010) at the NCO-P site shows that light absorption is dominated by carbonaceous material which is predominantly in the sub-micron fraction. The use

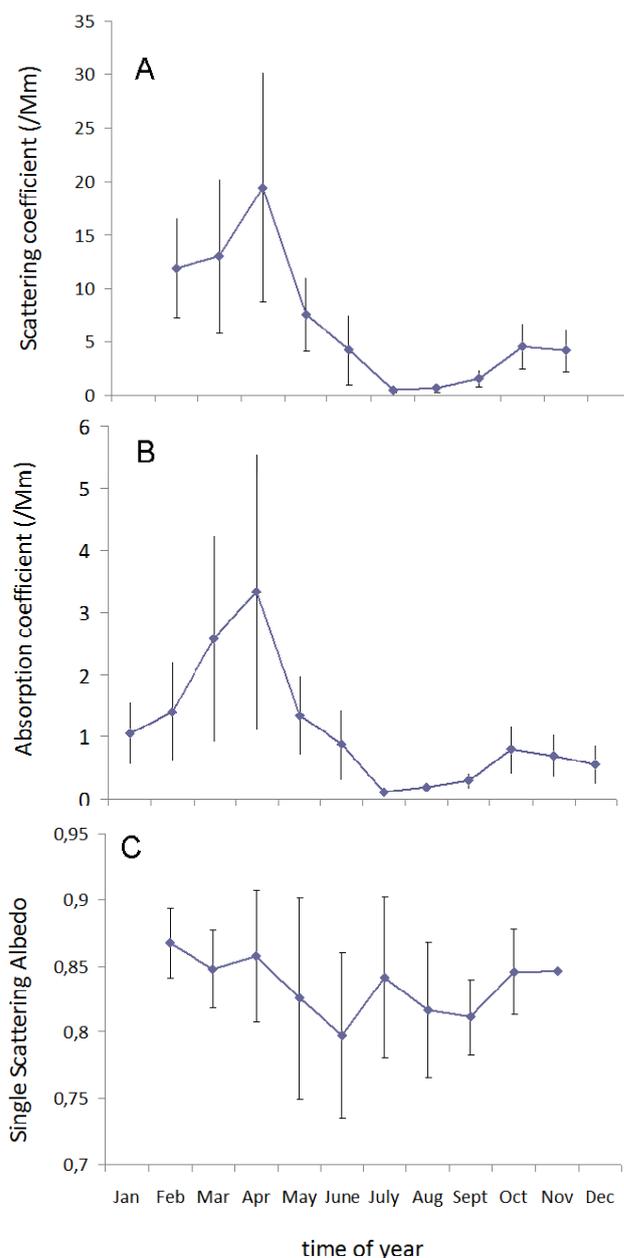


Fig. 1. Monthly averaged values of scattering coefficient (A), absorption coefficient (B) and single scattering albedo (C) measured at NCO-P. Error bars correspond to ± 1 standard deviation.

of different cut-size for Nephelometer and MAAP inlets will therefore not affect w_0 values. For reasons explained above, scattering coefficients for December and January are not provided.

The average scattering coefficient (not corrected for pressure) at the site is $6.3 \pm 10.7 \text{ Mm}^{-1}$, that includes relatively high σ_{sp} values during the pre-Monsoon periods ($19.4 \pm 21.2 \text{ Mm}^{-1}$ in March) as well as low σ_{sp} values for the Monsoon season in July/August ($0.4 \pm 0.5 \text{ Mm}^{-1}$ in July).

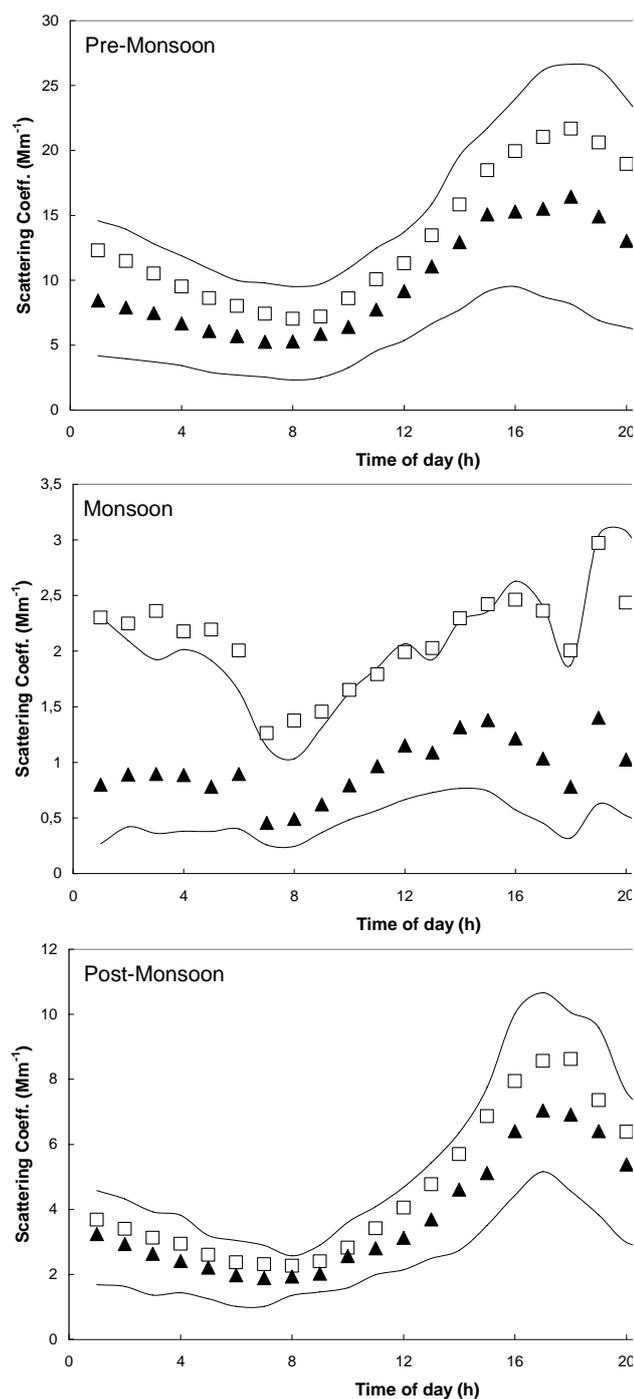


Fig. 2. Diurnal variability of hourly averaged (squares) and median (triangle) scattering coefficient values at NCO-P for pre-monsoon, monsoon and post-monsoon periods. Dashed lines represent 25 and 75-percentile, respectively.

These summer-Monsoon values are even lower than those observed at other remote sites such as Mauna Loa, Cape Grim, South Pole or Barrow where σ_{sp} ranges from 5 to 10 Mm^{-1} (these measurements are at 550 nm and for a size

cut of $D_p < 1 \text{ }\mu\text{m}$ at $\text{RH} < 40\%$) (Ogren, 1995; Carrico et al., 1998, 2000; Koloutsou-Vakakis et al., 2001; Delene and Ogren, 2002). The seasonal variations of σ_{sp} reflect a feature already observed for other aerosol variables at NCO-P and in the Himalaya with elevated concentrations during winter and pre-Monsoon seasons and much lower concentrations found during the Monsoon season (Carrico et al., 2003; Bonasoni et al., 2008, Venzac et al., 2009). This is to be compared to measurements performed at the altitude site of Mukteshwar station in India where much larger values of σ_{sp} are found closer to the source areas for both Monsoon and pre/post Monsoon seasons. Variations of σ_{ap} closely follow those of σ_{sp} with average annual value of $1.1 \pm 2.1 \text{ Mm}^{-1}$ and maxima during pre-Monsoon periods ($3.4 \pm 4.4 \text{ Mm}^{-1}$ in April) and minima during Monsoon season ($0.3 \pm 0.2 \text{ Mm}^{-1}$ in July). This obviously reflects wet scavenging of the aerosol component during the Monsoon season and stronger influence of air masses originating from Nepal and Indian plains during the pre-Monsoon seasons (Panday et al., 2009). As for σ_{sp} , data can be compared to measurements performed at Mukteshwar station showing highest σ_{ap} in April (23.2 Mm^{-1}) and the lowest one (4.5 Mm^{-1}) in August. Here again, distance from the strong Indian emission areas and high altitude explains differences between NCO-P and Mukteshwar stations.

Monthly averaged dry w_0 values are derived from σ_{ap} (670 nm) and σ_{sp} (700 nm) and range from 0.80 (in July) to 0.87 (in February) as shown in Fig. 1c. These variations reflect a higher proportion of absorbing material during the Monsoon that can either be related to preferential scavenging of more hygroscopic aerosol particles, such as sulphate and/or nitrate, or due to less-hygroscopic absorbing organic and BC material. It can also reflect changing emission source areas from long range transport of material from western regions of India and Pakistan in winter to more direct emissions from Indian and Nepal plains during Monsoon (summer) season (Bonasoni et al., 2010). Very low w_0 values (as low as ~ 0.75) have been measured by several authors within the boundary layer in India (Ramanathan et al., 2001) and also at the Mukteshwar station where w_0 at 525 nm range from 0.74 to 0.85. Our study therefore adds additional evidence that strongly absorbing particles are efficiently transported to the high remote Himalaya regions.

4.2 Diurnal variations of measured properties

The variability of the monthly-averaged optical properties is very strongly influenced not only by day-to-day changes but rather by significant diurnal variation. As discussed before, daily variations originate from thermal winds developing in the valley during the day. Both σ_{ap} and σ_{sp} are strongly dependent upon local wind circulation with very low values during the night and increase upon wind-shift in the late morning. Hourly averaged σ_{sp} for pre-Monsoon, Monsoon and post Monsoon seasons are presented in Fig. 2.

Diurnal variations show two distinct periods with low values in the early morning (1 a.m. to 11 a.m.) followed by strong increase associated with up-slope wind conditions from 11:00 on to evening hours). Diurnal variations are well pronounced during pre-Monsoon and post Monsoon seasons with up to a 3-fold increase in the mid-afternoon as respect to early morning values. On the contrary, the diurnal signal during the Monsoon season is rather weak, reflecting limited thermal wind circulation and/or scavenging process.

Scattering coefficient values in free tropospheric air sampled in the early morning at NCO-P typically range from 2 to 5 Mm^{-1} depending on the season. This is close to values measured in remote areas (Clarke, 1989). The free-tropospheric air flowing from the Tibetan plateau at night is extremely clean and does not show any influence of emission in Southern China. During afternoon hours, averaged values of σ_{sp} typically range between 10 and 20 Mm^{-1} during pre-Monsoon and post-Monsoon seasons, slightly less than 5 Mm^{-1} during Monsoon season. Upslope winds transport primary and secondary particles from the Nepal and Indian plains from more than 100 km away to the high altitude regions of Himalayas. Emission sources are likely linked to biofuel combustion sources as shown by the elevated levels of organic material measured in the particulate phase (Decesari et al., 2010).

Very similar diurnal variations are observed for the absorption coefficient, ranging from 0.4 Mm^{-1} in the morning for Monsoon periods to 4 Mm^{-1} in the afternoon for pre-Monsoon periods – see Fig. 3. Seasonal changes of the diurnal variation also show the strong influence of wet scavenging with very limited increase during the afternoon. Both σ_{sp} and σ_{ap} diurnal variations are in strong agreement with those measured by independent particle instruments at NCO-P.

Instead, diurnal variations of the single scattering albedo are not so well pronounced (see Fig. 4). During pre-Monsoon periods, w_0 shows a maximum (0.87) in the morning hours in correspondence with the nucleation peak seen by Venzac et al. (2008). Although scattering intensity by small particle is limited, presence of elevated levels of non-absorbing 20–40 nm particles might explain the w_0 increase. The w_0 then steadily decreases throughout the day reaching minimum values below 0.84 in the evening. Presence of a higher fraction of absorbing material in aerosol is likely due combustion aerosol in regionally-transported air masses. Residual layers from regional pollution seem to stay up until the evening although wind direction already shifted northwards. Daily variability during other seasons is less clear. During the Monsoon season, high humidity during the afternoon may be responsible for higher scattering (thus higher single scattering albedos). Relative humidity in the nephelometer is not controlled and scattering may be increased by hygroscopic growth of soluble particles. Lowest w_0 values (0.82) are found during the evening period as well, possibly explained by preferential scavenging of soluble material before reaching the high altitude and consequent enrichment in wa-

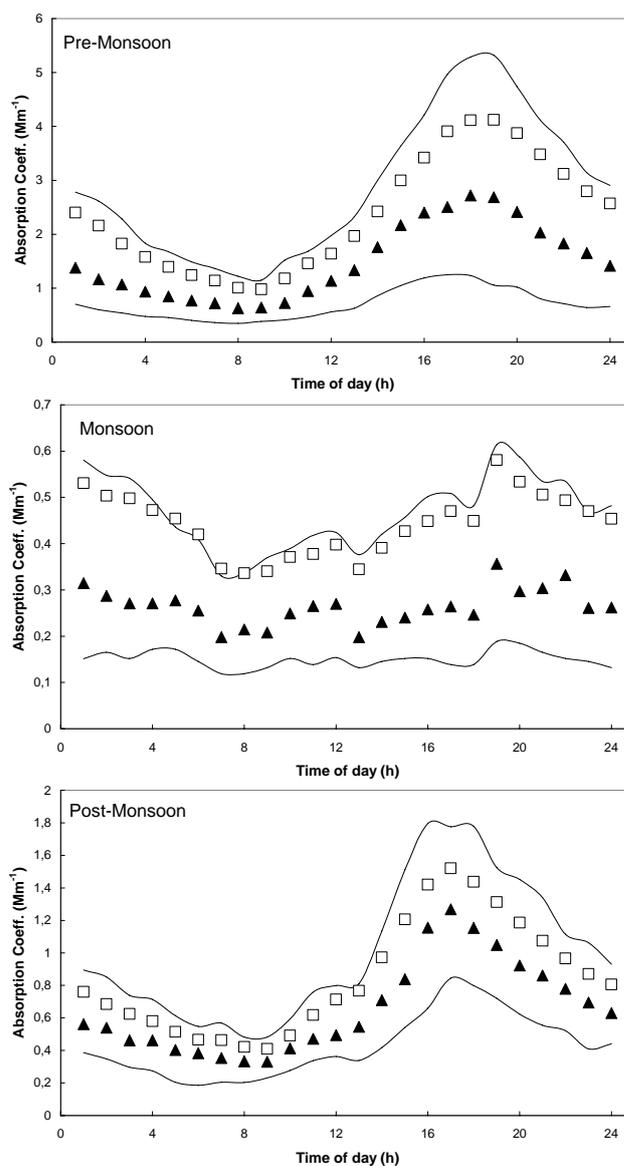


Fig. 3. Diurnal variability of hourly averaged (square) and median (triangle) absorption coefficient values at NCO-P for pre-monsoon, monsoon and post-monsoon periods. Dashed lines represent 25 and 75-percentile, respectively.

ter insoluble BC fraction. During post-Monsoon season, w_0 varies from 0.84 to 0.86, with maximum values in the morning hours. The daily variability is however very limited.

An interesting feature of the w_0 record at NCO-P is that the minimal values are consistently found at night, regardless of the season. 25th percentile of w_0 is close to 0.6 during Monsoon, slightly higher during the pre-Monsoon and post-Monsoon seasons (0.8), showing a relative enrichment of absorbing material in the night-time aerosol. This enrichment is either resulting from residual layers enriched in BC or to the presence of dust emitted from the arid plateau of

Table 1. Averaged aerosol optical properties found in the 3 types of events for each considered season excluding winter season.

Season	Event type	σ_{scatt} (Mm^{-1})	σ_{abs} (Mm^{-1})	AOD 550 nm	particles number ($D > 10$ nm) (cm^{-3})	w_0	asymmetry factor g
<i>f</i> Monsoon	BG	0.7±0.8	0.2±0.2	0.031±0.000	384±189	0.76	0.68
	RP	1.3±2.2	0.3±0.4	0.043±0.010	546±389	0.81	0.68
	SE	8.3±9.3	1.2±1.1	0.077±0.015	839±191	0.88	0.65
Post-Monsoon	BG	2.3±1.9	0.4±0.3	0.008±0.000	728±272	0.85	0.71
	RP	7.0±5.6	1.2±0.9	0.023±0.015	1127±624	0.85	0.74
	SE	14.4±6.1	2.8±1.0	0.098±0.027	1479±436	0.84	0.63
Pre-Monsoon	BG	5.9±4.6	0.8±0.6	0.010±0.000	656±269	0.89	0.78
	RP	12.9±9.3	2.3±2.3	0.025±0.015	1194±672	0.84	0.75
	SE	36.2±27.9	7.8±6.0	0.093±0.035	2810±1252	0.82	0.70

Tibet. Chemical analysis and analysis of BC/PM₁ variability seem to support the later hypothesis (Decesari et al., 2010; Marinoni et al., 2010).

The difference between median and mean w_0 is substantial during the Monsoon season while very similar (albeit slightly negative – mean $w_0 <$ median w_0). Both preferential scavenging of non-absorbing aerosols and the presence of relatively frequent episodes of long-range pollution may explain these findings.

4.3 Aerosol properties according to air masses

As mentioned previously, both seasonal and diurnal variations of aerosol optical properties appear to be strongly linked to: (1) wet removal during the Monsoon season lowering particle concentration and, thus, scattering and absorption coefficients, and (2) thermal wind development during the day explaining much of the daily variability of almost all aerosol parameters. Instead, origin of 5-day air mass back-trajectories cannot be used to explain the observed variability. In the following, we have, therefore, derived averaged values of optical properties using the classification proposed in Sect. 3, i.e. splitting between Background (BG), Regional Pollution (RP) and Special Events (SE). As mentioned earlier, BG and RP samples are restricted to very specific time periods during a typical day at NCO-P (5 a.m. to 9 a.m. for BG and 3 p.m.–6 p.m. for RP). The distinction between BG and RP on one hand, and special events on the other is instead made on a case-by-case basis after reviewing the data. The strong influence of SEs can be seen in both Figs. 2 and 3 by the substantial difference between median and averaged σ_{sp} and σ_{ap} in the record. SE events occur regularly, limit diurnal variability and are characterized by significant increases in BC, O₃ and particle Number concentrations for several days in a row.

Aerosol optical depth derived in Table 1 only used the fine fraction component of AOD to limit artefacts due to elevated clouds and high humidity, as recommended by Gobbi

et al. (2010). There are substantial differences between the 3 air mass types in addition to the seasonal dependence within each air mass type. Clearly, SE show very different values as respect to both BG and RP air masses. This is seen for all other parameters related to optical and chemical properties. Special events are characterized by large scale changes as discussed by Bonasoni et al. (2010). However, it should be noted that day/night variability still exist during SEs but, contrary to RPs, nighttime is not characterized by low PM concentrations. Absorption and scattering coefficients, however, can reach much higher values with respect to both BG and RP, such as other aerosol variables. There are no clear trends in w_0 related to SE. While SE during Monsoon shows relatively high w_0 (0.88), absorption is relatively higher during pre- and post- Monsoon SEs (0.82 and 0.84, respectively). The very high values of σ_{sp} ($36 \pm 26 \text{ Mm}^{-1}$) are comparable to, for example, those observed (at 550 nm) in North American or European sub-urban areas like Bondville, IL, southern Great Plains, OK, Sable Island, NS, and Sagres, Portugal where values of 30–50 Mm^{-1} are reported (Delenne and Ogren, 2002; Carrico et al., 2000); which also corresponds to measurements by Hyvärinen et al. (2009) at the Mukteshwar station much closer to emission sources.

Single scattering albedo w_0 during pre and post-Monsoon seasons vary from 0.84 to 0.89 for RP events and from 0.89 to 0.83 for BG events. This opposite behaviour is difficult to explain. Similar changes in w_0 are measured at Mukteshwar, although highest values are found during the rainy season, opposite to what is seen at NCO-P. The variability at Mukteshwar is possibly explained by changing agricultural practices from winter to summer which may also be the case at NCO-P. Comparisons with Elemental Carbon (EC) fraction measured on bulk filters (Decesari et al., 2010) confirm that the EC fraction of the Total Carbon content is relatively higher during the pre-Monsoon season as respect to other seasons. This which would explain the relatively lower w_0 measured during RPs which are over-represented in the

average presented by Decesari et al. (2010). The very low w_0 measured during the Monsoon season BGs is however difficult to compare with filters due to the very limited number of filters collected under these conditions.

5 Direct aerosol radiative forcing estimation

5.1 Model description

Data from Table 1 are used to derive the local aerosol radiative forcing for different air mass origins. It should be noted that we applied correction procedures recommended by Nessler (2005a and b) to scattering coefficients to account for high relative humidity (RH) at the station especially during the Monsoon season (mean RH above 90%). No correction is applied to absorption coefficient, thus considering BC mostly hydrophobic. To derive the local direct aerosol forcing, we have used the GAME code (Global Atmospheric Model code; Dubuisson et al., 1996) set in the configuration described in Roger et al. (2006). The absorption is based on the results of a line by line code (Scott, 1974) and includes the gas-phase absorbers in the short-wave region. The multiple scattering effects are treated using the Discrete Ordinates Method (DOM) (Stamnes et al., 1988) following the plane-parallel approximation. The DOM employs a Legendre polynomial decomposition for the phase function and the radiance. Interactions between multiple scattering and gaseous absorption are accurately treated using the correlated-k method (Lacis and Oinas, 1991). The GAME code takes into account the surface albedo as well as the asymmetry factor (given in Table 1) calculated using the method described in Marshall et al. (1995).

For computations, the GAME code divides the atmosphere into several layers. For this study, we have made the following hypotheses: (1) Above 9000 m a.s.l. the atmosphere is divided into several layers, which optical properties remain constant. We have used the global standard properties for this altitude (Hess et al., 1988). (2) the atmosphere between 5000 m a.s.l. and 9000 m a.s.l. is divided into eight uniform 500 m-thick layers. Aerosol intensive properties within each layer are originally derived from surface measurements with different degree of mixing with typical free tropospheric air. Layer 1 (5000 m a.s.l. to 5500 m a.s.l.) is the polluted layer for which optical aerosol properties are given by the direct measurements at NCO-P. The thickness of the layer is derived from modelling studies of thermal wind developments in the Alps, showing a thermal wind depth of a few hundred meters (Henne et al., 2005). Layers 2 to 5 (5500 m a.s.l. to 7500 m a.s.l.) are resulting from the mixing between layer 1 and layers above. We considered that the mass of particles in all four layers as equal to 10% of the mass of the particles from layer 1 (polluted layer) plus 90% of the mass of the particles from layer 6 to 8 (free troposphere). Layers 6 to 8 (7500 m a.s.l. to 9000 m a.s.l.) are the free tropospheric

layers and are considered to have uniform optical properties. These properties are given by the direct measurements at the station during night time when the downslope wind brings clean air masses from the free troposphere to the station.

For each layer, the GAME code inputs consist of (1) relative humidity, pressure and temperature derived from climate data for high altitudes atmospheric layers but adapted to fit the measures from the station, (2) wavelength-dependence aerosol optical properties (in practice 10 values in the solar range are entered and others are deduced by interpolation): extinction coefficient σ_{ext} (scattering coefficient + absorption coefficient), asymmetry factor and single scattering albedo w_0 . The coefficients are calculated to fit the layering of the atmosphere that we supposed and the AOD is divided into the layers so that the contribution of each layer to the AOD is proportional to the concentration of aerosols in it; and (3) Solar angle and surface wavelength-dependence ground albedo.

The outputs of the GAME code are the upward and downward net radiative fluxes between the different atmospheric layers computed for all solar spectrum (intervals of 100 cm^{-1} from $0.3 \mu\text{m}$ to $3.0 \mu\text{m}$) and their radiative heating rate. Computations are done for two cases: with and without aerosol. The direct radiative forcing is then derived by subtracting these two values. The conventions of this forcing at the ground (Bottom Of Atmosphere – ΔF_{BOA}), in the atmosphere (ΔF_{ATM}) and at the Top Of the Atmosphere (ΔF_{TOA}) are the ones commonly used, i.e. imply an aerosol cooling effect when ΔF is negative. It should be kept in mind that calculated forcing are maximum theoretical values because it is assumed that all energy is converted to heating. Clouds are not considered in the study, assuming only clear-sky condition. Snow albedo is that of pure fresh snow and therefore, synergistic impact of BC deposition to snow and albedo modification is not considered in the study.

5.2 Direct radiative forcing by aerosols

Several runs have been performed for the different conditions listed in Table 2. For the pre-Monsoon season, we considered two cases; with snow surface and with rock surface conditions corresponding respectively to 58% and 42% of frequency. Results are summarized in Table 2 where we report forcing in W m^{-2} for the first layer of the atmosphere (F_{ATM}), at the surface (F_{BOA}) and at top of atmosphere (F_{TOA}) due to the presence of the aerosol layer. We point out that F_{TOA} is calculated only considering the atmospheric column above the sampling station and is therefore of limited relevance. Only F_{ATM} is discussed in the context of the present paper. We also report in Table 2 the corresponding heating of the first atmospheric layer above surface for snow-covered and bare-rock conditions. The maximum heating rate calculated from the 3 h afternoon period is also indicated.

Table 2. Daily aerosol direct radiative forcing ($\text{W m}^{-2}/\text{day}$) for the different seasons and different air masses (BG for Background; RP for regional pollution; SE for special event) at , within all the Atmosphere (F_{ATM}), at the surface (F_{BOA}) and at top of atmosphere (F_{TOA}). We also report the daily heating (in K) just above the surface; the mean value of the 3 h afternoon period around the maximum is reported in parenthesis.

Season	Air mass	F_{ATM}		F_{BOA}		F_{TOA}		heating above surface (max value)	
		snow	rocks	snow	rocks	snow	rocks	snow	rocks
Pre-Monsoon	BG	3.7	1.9	-1.6	-1.7	2.1	0.2	0.1 (0.2)	0.0 (0.1)
	RP	10.2	5.2	-4.4	-4.4	5.9	0.7	0.6 (3.0)	0.4 (2.0)
	SE	42.1	22.3	-19.5	-19.0	22.6	3.3	1.1 (3.5)	0.7 (2.1)
Monsoon	BG		7.1		-6.2		0.9		0.1 (0.3)
	RP		10.1		-8.8		1.3		0.5 (1.8)
	SE		23.2		-19.4		3.8		0.7 (2.0)
Post-Monsoon	BG		1.0		-1.0		0.0		0.0 (0.0)
	RP		3.5		-3.4		0.1		0.3 (1.4)
	SE		18.5		-17.2		1.3		0.5 (1.5)

For all background (BG) conditions, the daily radiative impact due to the presence of aerosol particles is almost negligible with none significant heating at surface, which is expected since BG air masses correspond to the very clean conditions of the free troposphere.

Radiative forcing during RP and SE conditions is significantly higher than for BG conditions. In Fig. 5, daily variations of the aerosol direct radiative forcing are shown using conditions corresponding to the pre-Monsoon conditions. In the absence of snow (bare rocks conditions), aerosol forcing on the first 500 m atmospheric layer (F_{ATM}) and the associated forcing is limited for RP events, ranging from 3.5 to 10.1 $\text{W m}^{-2}/\text{day}$ according the season. Almost all effective forcing is concentrated over the 11 a.m.–3 p.m. period corresponding to the peak of particle concentration due to regional transport. In that case, instantaneous radiative impact above a rocky surface ranges from up to 20 W m^{-2} corresponding to 1.4 K to 2.0 K warming above surface. In all cases, the ATM radiative impact is more important during the pre-Monsoon season due to the presence of snow.

The aerosol radiative forcing is higher in the presence of snow. This is due to higher surface albedo as respect to bare rocks favouring direct and multiple reflections between the surface and the atmosphere. During pre-Monsoon period, we have considered the presence of snow for 58% of the time, based on meteorological observations. This leads to instantaneous ATM forcing of up to 40 W m^{-2} corresponding to a maximum of 3.0 K warming in the atmosphere above surface for a given day.

Radiative impact is even higher during SEs. This is not only due to very elevated concentration of particles, as seen in previous sections, but also because radiative impact and heating is more persistent throughout the day as respect to RP conditions. The average daily radiative impact for pre-Monsoon seasons ranges from 18 to 23 $\text{W m}^{-2}/\text{day}$ over rocks

depending on the seasons. Over snow, the instantaneous forcing is even higher, reaching more than 50 W m^{-2} in the morning and 140 W m^{-2} in the afternoon. This corresponds to maximum heating rate of 3.5 K/day above surface.

Overall, when we consider that SE event account for a maximum of 5% of the time and that daylight period is, on average, split between BG in the morning (30% of the time) and RP in the late-morning/afternoon (events 70% of the time), we calculate an average radiative forcing of approximately 10 W m^{-2} above bare rocks and 20 W m^{-2} above snow surface. These are obviously considering only clear-sky conditions. Radiative forcing and heating rates calculated for the conditions of NCO-P are slightly higher but on the same order, than those estimated by Ramana et al. (2004) for the Himalaya area, using a different method. There, daily heating rate are on average 1.2 K per day with in the first kilometer corresponding to 5 W m^{-2} . Since our estimate is on one side a maximum value as explained in section 5.1 and on the other side, only referring to a very local layer of a first 500 m, we can consider that the two estimates show similar results.

At the surface, the computed forcing is negative, ranging from -1.5 W m^{-2} for BG periods up to -20 W m^{-2} for SEs (see Table 2). Presence of snow is, in that case, not influencing BOA forcing. This forcing estimate is similar to that derived by Ramana et al. (2004) but for location much closer to emission sources. The amount of radiation received at the surface is clearly modified by the presence of aerosol particles even at very high elevation.

6 Conclusions

Our results confirm that heating of the lower atmosphere due to absorbing aerosols transported from regional to long-range distances is significant in the high altitude areas of

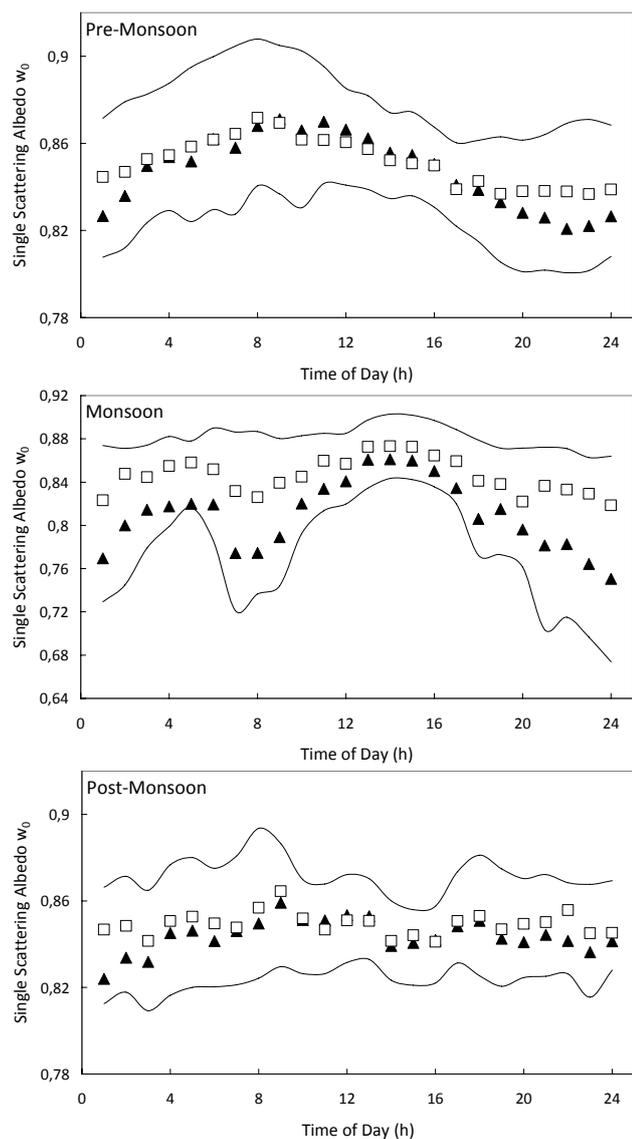


Fig. 4. Diurnal variability of the hourly averaged (square) and median (triangle) single scattering albedo values at NCO-P for pre-monsoon, monsoon and post-monsoon periods. Dashed lines represent 25 and 75-percentile, respectively.

Himalayas, and locally higher than IPCC estimated GHG associated warming. Our estimates range from 10 to 20 W m^{-2} for the first atmospheric layer (500 m above surface). Warming this layer is extremely dependent upon the presence of snow at the surface.

Warming takes place preferentially during episodes of regional pollution occurring on a very regular basis in the Himalayan valleys. A very strong component of the forcing is linked to the chemical characteristics of aerosol particles containing a strong absorption component. Single scattering albedos of 0.85, and lower, are regularly measured at NCO-P, in particular during episodes of regional pollution dur-

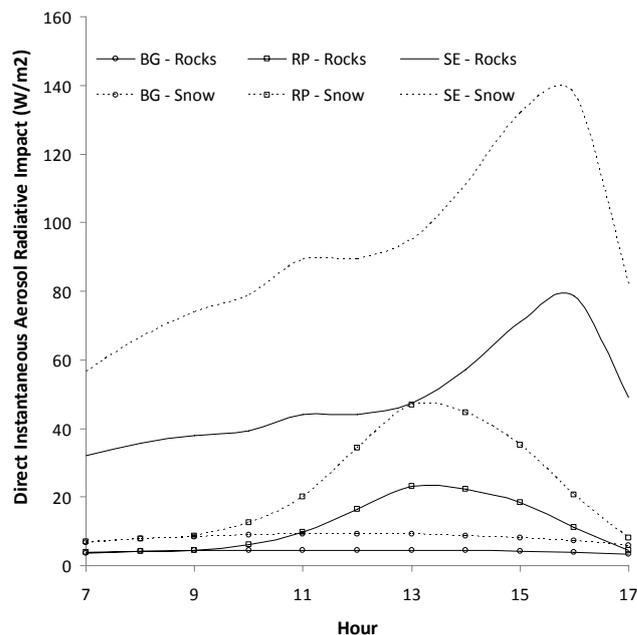


Fig. 5. Direct instantaneous radiative forcing in the first atmospheric layer (F_S) during insulation hours in the pre-monsoon season for the 3 defined air mass types (BackGround, Regional Pollution and Special Events) and for snow-covered and bare rock surfaces.

ing the pre-Monsoon seasons. On relatively polluted days, instantaneous radiative forcing can reach values as high as 140 W m^{-2} corresponding to more than a 3 K warming of the lower atmosphere. Warming of the lower layer is accompanied by a substantial decrease of the amount of radiation reaching the surface. This is estimated on the order of -4 to -20 W m^{-2} for RP and SE events, respectively.

Implications of these results are important for the local atmospheric radiative balance in the high Himalaya, in particular with regards to the seasonal snow cover dynamics. The dynamics of snowpacks is, in fact, a complex function of the local meteorological conditions throughout the year (air temperature, irradiance, humidity and precipitation) and variables related to the snowpack conditions (i.e. snow albedo) influencing the local energy balance. On one hand, surface air temperature is increased by presence of absorbing particles, possibly leading to increased melting rates. On the other hand, direct amount of radiation received by the snowpack is substantially reduced, implying a decrease in surface temperature. This effect will, in turn, limit snow-pack melting. An important issue in this case is linked to the presence of absorbing material in snow which can compensate for the reduction of incoming solar radiation by decreasing snow albedo. Calculation of such effects is beyond the scope of this paper but recent study of Yasunari et al. (2010) for the area of NCO-P showed significant melting potential of BC in snow. In the future, a more detailed study, coupling the

radiative balance of atmosphere with that of snow is required to provide a more complete estimate of the impact of aerosol particles in the HKH area.

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