

1 Airborne measurements of the spatial distribution
2 of aerosol chemical composition across Europe
3 and evolution of the organic fraction:
4 Supplementary material

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11 **1 Scope**

12 The supplementary material outlined in this document is provided in order to present
13 the meteorological context of the flight operations and support the analysis techniques
14 and data quantification steps outlined in the main paper. The meteorological fields
15 corresponding to each flying period are presented and further information regarding
16 the photochemical context of the operations is presented. Further details regarding
17 the volume closure between the Aerosol Mass Spectrometer (AMS) and the Passive
18 Cavity Aerosol Spectrometer Probe (PCASP) are discussed. Comparison of the esti-
19 mated HOA with primary combustion tracers is included. The relationship between
20 the fractional contribution of Low-Volatility Oxygenated Organic Aerosol (LV-OOA)
21 to the organic mass and the normalised organic signal at m/z 44 is also shown. Further
22 information is provided regarding the Positive Matrix Factorisation (PMF) analysis ex-
23 amples from the main text, as well as a summary of some PMF diagnostics for the
24 whole dataset. The PMF analysis was performed using the tools presented by Ulbrich
25 et al. (2009).

26 **2 Meteorological summary**

27 Figs. S1 and S2 display the typical meteorological conditions prevalent during each
28 period considered by the analysis. The periods are relatively consistent in terms of
29 their transport patterns, with the air masses transporting pollution from continental
30 Europe downwind to either the UK region or into the Eastern Atlantic Ocean. Thus the
31 flights are predominantly focused upon either sampling such pollution over continental
32 Europe itself or at a range of scales downwind.

33 The evolution of the aerosol chemical composition during the LONGREX-2 period
34 was examined based upon the relatively consistent transport patterns prevalent during
35 the period. Fig. S3a displays the back trajectories for each flight during this period
36 based upon Straight and Level Runs (SLRs) during each flight. The trajectories display
37 highly consistent behaviour during the period, which is unsurprising given the relative
38 stability of the high pressure system located over Northern Europe during this period.
39 Fig. S3b highlights the back trajectory from the 14 May 2008, which was initialised
40 from a SLR during B374 in the Eastern Atlantic Ocean. B374 represented the end-
41 point in our operations during this period both in terms of the geographical location
42 of the missions and also the distance from continental European sources i.e. the most
43 aged polluted air mass. The back trajectory indicates that the spatial coverage of the
44 flight operations closely matches the air mass transport during the period leading up to
45 the 14 May 2008. Specifically, flights B370-B374 took place across Northern Europe
46 during this period covering close to 5 days of air mass transport.

47 **3 AMS versus PCASP comparison**

48 Validation of the collection efficiency treatment applied to the dataset following the
49 principles developed by Matthew et al. (2008) is accomplished by comparing the AMS
50 data with the volume estimated concentrations from the PCASP instrument. The AMS
51 total mass concentrations were converted to total volume concentrations using the den-
52 sities reported by Cross et al. (2007), which correspond to 1.27 g cm^{-3} for organics
53 and 1.77 g cm^{-3} for inorganics. A comparison of the estimated volume from the AMS
54 and PCASP is shown for SLRs below 3000 m in Fig. S4. Over all of the considered
55 flights, the estimated AMS volume concentrations were 26% higher than the estimated
56 PCASP volumes. This average agreement is predominantly determined by the LON-
57 GREX flights, which were quite consistent in terms of the agreement from flight-to-
58 flight. The ADIENT flying periods sit on either side of the overall regression slope,
59 with the ADIENT-2 flying indicating that the PCASP volume was 48% of the AMS
60 volume. These discrepancies are considered tolerable given the large uncertainties pre-
61 viously reported in the literature for PCASP volume estimates (e.g. Moore et al., 2004;
62 Hallar et al., 2006) and the uncertainties in the AMS volume estimates.

63 For B357, the PCASP volume estimate was more than two times greater than the
64 AMS volume estimate, which is outside of the bounds of uncertainty for the two instru-
65 ments. The reason for this discrepancy is unknown but could reflect an artifact in either
66 instrument or the presence of material that is not detected by the AMS. The discrep-
67 ancy between the two instruments is also reflected in the calculated volume-scattering

68 relationship when comparing the measurements with a nephelometer system. The main
69 difference between B357 and the other flights in the dataset is the sampling altitude of
70 the aircraft, where in B357, the aircraft operated at a constant altitude of 200 m for
71 the majority of the flight. The other flights in the dataset operated at altitudes higher
72 in the boundary layer. Potentially, the nephelometer and PCASP measurements could
73 be perturbed by the constant low-level flying in a humid environment as the aerosol
74 sampled may not be sufficiently dried in the inlet lines and by the heater respectively.
75 The AMS volume estimates do not include water, so this could potentially cause the
76 discrepancy. Additionally, the PCASP and nephelometer may be measuring refractory
77 material or particles above the cut off of the AMS aerodynamic lens. This would also
78 lead to the AMS underestimating the volume relative to the PCASP.

79 **4 Photochemical context**

80 The relationship between O_3 and CO with the $O_3:NO_x$ ratio discussed in the main pa-
81 per is presented in Fig. S5. The results indicate that O_3 increases and CO decreases
82 steadily in the 1-100 $O_3:NO_x$ range, which is a reflection of photochemistry and di-
83 lution respectively. Beyond an $O_3:NO_x$ ratio of 100, the concentrations decrease with
84 CO returning to background levels and O_3 remaining relatively constant in the 40-60
85 ppb range.

86 **5 Positive Matrix Factorisation**

87 Potentially, the most challenging and subjective aspect of PMF analysis is the selec-
88 tion of the appropriate number of factors. For AMS datasets, this is usually accom-
89 plished using internal PMF diagnostics, similarity to reference mass spectra and exter-
90 nal measurement parameters. An example of an internal diagnostic is the parameter
91 Q/Q_{expected} , which is defined as the total sum of the scaled residuals, divided by its
92 expected value. This expected value is derived based upon the error estimates for the
93 data matrix (Ulbrich et al., 2009). A value of unity for the Q/Q_{expected} parameter in-
94 dicates that the expected variance associated with random errors can be explained by
95 the solution set. Values greater than unity indicate that there is additional variance not
96 accounted for by the solution set. The suite of aerosol and gas phase instrumentation
97 available on the aircraft provides several necessary external parameters to facilitate
98 validation of the solution. Reference spectra utilised in this study are taken from the
99 AMS spectral database (<http://cires.colorado.edu/jimenez-group/AMSsd/>)
100 described in Ulbrich et al. (2009). The factor solutions are interpreted based upon com-
101 parisons with external parameters and reference mass spectra. Comparisons between
102 mass spectra are accomplished using the Uncentered Correlation (UC, Ulbrich et al.,
103 2009) coefficient, while the comparisons with external parameters use the Pearson's R
104 coefficient.

105 Here we include additional information regarding the PMF solutions used in the
106 main manuscript. This includes example PMF solutions from a range of flights in
107 different conditions, a summary of the results for the whole dataset and further details

108 regarding the estimation of HOA from our data.

109 **5.1 Example PMF solutions**

110 Several PMF solutions will be illustrated in the following section based upon some
111 example flights from the dataset. The flights chosen are B357 (16 April 2008), B362
112 (6 May 2008), B369 (10 May 2008) and B406 (25 September 2008) as these broadly
113 represent the range of operations conducted in terms of their proximity to pollution
114 sources. B357 was conducted primarily downwind of the major conurbations of Manch-
115 ester and Liverpool, in the North-West of England. B362 was conducted on a South-
116 North transect originating from Oberpfaffenhofen in southern Germany, culminating
117 in a sequence of Straight-and-Level Runs (SLRs) in the North Sea in the outflow from
118 continental Europe. B406 was conducted in the outflow from continental Europe along
119 the southern coast of the UK. B369 was conducted in the Baltic Sea region and reflects
120 background conditions for comparison with the more polluted examples.

121 The results for B357 are shown in Fig. S6, for B362 in Fig. S7, for B369 in Fig. S8
122 and B406 in Fig. S9. All the examples show excellent agreement between the measured
123 and reconstructed organic mass concentrations.

124 **5.1.1 Close to source case study – B357**

125 For B357, a 2-factor solution was deemed most appropriate as increasing the number
126 of factors led to a phenomenon known as “splitting”. This leads to multiple factors
127 being assigned with numerically similar factor profiles and time series. In this case,
128 the additional factors merely represented variability within the factors identified in the
129 2-factor solution. Consequently, the 2-factor solution was retained for analysis. The
130 Q/Q_{expected} was equal to 1.38. Reconstructed OM concentrations made up 98% of
131 the measured OM concentrations. Factor 1 (OOA-1) exhibits a correlation coefficient
132 of 0.89 and 0.98 when compared to reference spectra for an ambient rural case (from
133 Canada) and laboratory generated fulvic acid respectively, which has chemical func-
134 tionality that are representative of aged OM (e.g. McFiggans et al., 2005) and hence
135 is indicative of an LV-OOA type factor. Factor 2 (OOA-2) is interpreted as HOA as
136 it exhibits a high correlation (0.91) with the derived HOA mass spectrum from Pitts-
137 burgh (Zhang et al., 2005a). Additionally, it has a correlation coefficient of 0.87 when
138 compared to diesel exhaust from a chase study in New York (Canagaratna et al., 2004).
139 The factor 2 time series correlates with NO_x ($r=0.65$), CO ($r=0.73$) and BC ($r=0.79$),
140 with the strong gradients in the time series coincident with large increases in NO_x , CO
141 and BC. Cororally, factor 1 exhibits low correlation ($r<0.35$) with these combustion
142 tracers. Comparison of the emission ratio of POA to NO_x from B357, which is equal
143 to $29.9 \mu\text{g sm}^{-3} \text{ppm}^{-1}$, with the previous studies summarised in Table S1 yields good
144 agreement. Furthermore, the emission ratio relative to CO of $8.0 \mu\text{g sm}^{-3} \text{ppm}^{-1}$ falls
145 within the literature values summarised in Table S1. Therefore, it appears that factor 2
146 in this case is likely attributable to primary sources of OM in the form of HOA.

147 5.1.2 Continental European scale case studies – B362 and B406

148 While B357 presented a relatively straight forward 2-factor solution, B362 and B406
149 present highly complex examples of the factor analysis. When more than 3 factors were
150 examined, the “splitting” phenomenon was observed. The 3-factor solutions contained
151 a factor profile strongly resembling LV-OOA, with m/z 44 dominating the mass spec-
152 trum and a high correlation with the reference spectra for fulvic acid and the ambient
153 rural case. The second factor is characterised by a mass spectrum with similar intensi-
154 ties at m/z 43 and 44 and an enhanced base peak at m/z 55. Such a spectrum is fairly
155 typical of a SV-OOA type component (Ulbrich et al., 2009; Jimenez et al., 2009). The
156 remaining factor appeared to represent a second SV-OOA mass spectrum but with en-
157 hanced mass peaks associated with commonly resolved hydrocarbon peaks (m/z 's 27,
158 29, 41, 43, 55, 57, 69, 71, ...), indicating a contribution of HOA. This third factor was
159 correlated with NO_x , CO and BC but for B362 the slope was found to be approximately
160 a factor of 2 greater than the range of literature values cited in Table S1. A regression
161 for the same factor from B406 with NO_x and BC also yielded a similar discrepancy
162 although a CO measurement was not available during the flight so a comparison with
163 CO was not possible. Enhanced signal is not identified at m/z 60 or 73 during either
164 flight. These are typical mass spectral markers for wood burning emissions (Alfarra
165 et al., 2007), therefore solid fuel burning is not considered to be a potential source of
166 the enhanced mass. Consequently, it appears that the 3-factor solution is “blending”
167 more than one distinct organic component into a single factor. This is likely a conse-
168 quence of there being more chemical variability in the SV-OOA component (arising
169 from evaporation/condensation, aging etc.) than in the HOA component. Thus the
170 PMF solution identifies a factor which represents a ‘mathematical mixing’ of the more
171 recently formed OOA with a HOA component. This arises due to the commonality
172 between some of the major peaks in their respective mass spectra e.g. m/z 29 and 43.

173 The number of factors was increased under the supposition that an increase beyond
174 3-factors would reveal a more realistic contribution of HOA to the OM by separating
175 it from the more volatile, fresher OOA fraction. While OOA factor profiles gener-
176 ally display significant variability as a result of differing processes such as aging and
177 partitioning, HOA factor profiles are far more chemically distinct. Additionally, the
178 largely linear association between HOA concentrations and urban primary emission
179 markers makes source identification more straightforward than for OOA. Based upon
180 these criteria, the number of factors was increased and inspected at each step until a
181 single factor was present that most resembled HOA. For B362, this occurred for a 6-
182 factor solution while for B406, a 7-factor solution was chosen. These factors were
183 chosen due to their strong resemblance to HOA based upon their mass spectra and
184 comparisons with combustion tracers. The comparison with the combustion tracers
185 for B362 is shown in Fig. S14b, which indicates that increasing the number of factors
186 does appear to bring the HOA-type component closer to the literature emission ratio.
187 The main deviation from the relationship is from a low-level SLR in a highly moist
188 layer ($\text{RH} > 95\%$) where total mass concentrations exceeded $50 \mu\text{gsm}^{-3}$ and ammo-
189 nium nitrate was the dominant chemical component. The layer is likely characteristic
190 of freshly formed secondary material and it appears that some of the freshest-OOA
191 mass has been apportioned to the HOA profile. The number of factors was increased

192 to 15 but this did not significantly alter the mass apportionment during this event.

193 **5.2 Bootstrapping analysis**

194 The numerical stability of the factor solutions was quantitatively evaluated using a
195 bootstrapping analysis (Ulbrich et al., 2009, and references therein), where random
196 resampling of the data matrix is performed in the time dimension. This analysis was
197 performed using 20 iterations, with the results being grouped according to the UC
198 coefficient between mass spectral profiles. The results of this analysis are summarised
199 in Fig. S14a for B362 by comparing the contribution to total mass versus m/z 44 for
200 each factor as the solutions are stepped through an increasing number of factors. The
201 derived mean and standard deviations from the bootstrapping analysis are compared
202 with the base solutions. The analysis indicates that by increasing the number of factors,
203 the solutions become increasingly unstable in a numerical sense. This is a consequence
204 of the aforementioned chemical variability inherent in the air masses sampled during
205 the flight operations, which results in large scope for different factor solutions for larger
206 numbers of factors. The bootstrapping analysis suggests that the 2-factor solution is the
207 most appropriate solution, especially as the OOA-1 component is highly robust with a
208 close match between the base solution and the bootstrapping solution. The enhanced
209 standard deviation in the m/z 44 for the OOA-2 component is likely a consequence
210 of the variability in the chemical nature of the OM. Very similar results were derived
211 for B406, with the 2-factor solution being more numerically robust than subsequent
212 solutions.

213 Further results from the bootstrapping analysis are shown for B357 in Fig. S13, for
214 B362 in Fig. S14, for B369 in Fig. S15 and for B406 in Fig. S16. The bootstrapping
215 results for all the flights are summarised in Table S2. The variance of the solutions in
216 both the time series and factor profile dimensions is evaluated using suitable metrics.
217 The time series diagnostic is the mean of the standard deviation for each factor, reported
218 as a percentage of the overall mean mass concentration. For the mass spectra, the
219 greatest standard deviation for each factor profile from the bootstrapping analysis is
220 reported. A mean is not deemed appropriate to evaluate the stability of the mass spectra
221 as the chemical assignment of factors is performed based on a limited number of peaks
222 (i.e. less than 10).

223 The OOA-1 (LV-OOA) factor profiles are highly robust with little deviation be-
224 tween the average mass spectrum from the bootstrapping analysis and the base solu-
225 tion. Furthermore, the standard deviations are typically low. This is a consistent result
226 throughout the dataset, which is shown by the low scores for the diagnostics in Table
227 S2. The OOA-2 factor profiles are more variable for B362, B369 and B406, with less
228 stability in the signal intensity at m/z 44. This reflects the continuum nature of the OM
229 discussed in the main text, whereby there is significant variability on a given flight in
230 the level of oxidation. Thus by randomly resampling the dataset in the time dimension
231 using the bootstrapping procedure, conditions with either enhanced or diminished m/z
232 44 in the OOA-2 may be encountered and this is reflected by enhanced standard devi-
233 ations in both the time series and mass spectra. The normalised standard deviation for
234 the time series of OOA-2 for B369 is much greater than the other flights in the dataset.
235 This is predominantly a result of the low concentrations encountered during the flight,

236 coupled with large standard deviations which are associated with large changes in the
237 composition of the OOA-2 component in this instance. For much of the flight, signal
238 at m/z 57 is close to zero indicating that HOA makes a minimal contribution to the
239 OOA-2 component. However, during the peak OM event at 15:47 UTC, signal at m/z
240 57 is enhanced and thus the contribution of HOA is likely elevated. This event co-
241 incides with the maximum in the NO_x concentration and is associated with potential
242 sampling of low-level urban outflow from Stockholm into the Baltic Sea. During this
243 period, the standard deviations for the OOA-2 factor increase. Thus such changes in
244 the OOA-2 composition from SV-OOA dominated to HOA dominated are reflected by
245 large increases in the standard deviation values from the bootstrapping analysis. The
246 much lower normalised standard deviation values associated with the time series of
247 the OOA-2 components for the rest of the dataset suggest this feature to be atypical.
248 This is a reflection of the regional nature of the measurements, with few instances of
249 prolonged exposure to intense urban signatures.

250 The results presented here demonstrate the robustness of the chosen 2-factor solu-
251 tions in terms of both the mass spectra and time series of the components. The OOA-1
252 (LV-OOA) components are highly numerically stable, while the decreased numerical
253 stability of the OOA-2 (SV-OOA and HOA) components is entirely consistent with the
254 continuum of oxidation/aging discussed in the main text.

255 **5.3 Application to the entire dataset**

256 The remaining flights in the dataset were analysed in an identical manner to the frame-
257 work established in the previous section. This resulted in broadly similar behaviour in
258 terms of the inability to accurately and quantitatively resolve HOA. A consistent theme
259 was that increasing the number of factors in order to attempt to separate the HOA con-
260 tribution led to a numerically unstable solution. Thus we chose to use the 2-factor
261 solutions as these consistently represented a more quantitative solution set.

262 A summary of the Q/Q_{expected} parameter is shown in Fig. S12, which indicates that
263 the parameter is generally less than 2. Four flights had Q/Q_{expected} values greater than
264 2, with the largest value being 5.01 (B374). Such values are greater than is generally
265 considered optimal if attempting to produce a perfect characterisation of the dataset but
266 given the difficulty in deriving robust results when more than 2 factors are chosen, it
267 is not possible to reduce Q further. Consequently, the additional Q contribution prevalent
268 in the dataset is considered as ‘model error’.

269 The available solutions include some rotational ambiguity, which is explored by
270 varying a parameter known as “fPeak” (Ulbrich et al., 2009, and references therein).
271 An fPeak range from from -2.5 to 2.5 is investigated in order to explore the numerical
272 variability in factor profiles and time series for small changes in Q/Q_{expected} . Invest-
273 igation of the rotational freedom in the solutions using fPeak was accomplished by
274 inspecting the mass spectra and time series in relation to external tracers for a subset of
275 fPeak values from -2.5 to 2.5 . The most appropriate value was then chosen, which for
276 this dataset was determined to be zero in all cases. A test of the numerical uniqueness
277 of the solution sets is the dependence upon the initiation seed, which is described by
278 Ulbrich et al. (2009). Each of the 2-factor solutions was examined using this technique
279 and little variation was exhibited for a range of different seeds.

280 **5.3.1 Estimation of HOA and comparison with OOA components**

281 Included in Fig. S11a are correlations of the estimated HOA concentration with Black
282 Carbon (BC), NO_x and CO. These indicate that for 8 of the flights, the correlations
283 of the estimated HOA with these primary emission tracers are greater than 0.5. Cor-
284 relations lower than 0.5 are generally encountered on flights where these tracers and
285 the estimated HOA are very low, thus the correlations break down at values when the
286 relationships exhibit enhanced noise due to low signal. This is demonstrated in Fig.
287 S11b and c, where at low concentrations the relationships are relatively flat but at en-
288 hanced concentrations, the correlation is greater. Given the simple nature of the HOA
289 estimate, these correlations and relationships do provide some confidence that the es-
290 timated HOA provides a qualitative indicator of the contribution of HOA to the OM
291 burden. The HOA estimate using this approach is likely an upper limit as the contribu-
292 tion of any oxidised fragments at m/z 57 has not been removed.

293 Also shown in Fig. S11a are the correlations between the Low-Volatility Oxidised
294 Organic Aerosol (LV-OOA) organic mass fraction and the m/z 44:OM ratio described
295 in the main paper.

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Table S1: Summary of POA:NO_x and POA:CO emission ratios used in this study. Emission ratios are given in $\mu\text{g sm}^{-3} \text{ppm}^{-1}$.

Study	Location	POA:NO _x	POA:CO
Allan et al. (2010)	London, UK	31.6	N/A
Allan et al. (2010)	Manchester, UK	N/A	20.5
de Gouw et al. (2005)	Northeastern USA	N/A	9.4
Kirchstetter et al. (1999)	Tunnel study, USA	11.0	N/A
Lanz et al. (2007)	Zurich, Switzerland	15.9	20.4
Zhang et al. (2005b)	Pittsburgh, USA	N/A	4.3

Table S2: Summary of the diagnostics relating to the bootstrapping analysis from the dataset for each flight.

Flight	SD _{ts} /TS (%)		Max (SD _{ms}) (%)	
	OOA-1	OOA-2	OOA-1	OOA-2
B357	2.0	2.5	0.77	0.85
B362	5.6	9.7	1.00	2.44
B365	2.8	3.0	0.62	1.31
B366	13.4	11.8	1.66	1.46
B369	19.1	45.9	2.23	1.66
B370	2.9	3.6	0.52	0.46
B371	4.8	8.0	0.41	1.82
B373	2.3	3.9	1.06	0.86
B374	1.1	1.7	0.44	0.68
B379	6.6	9.3	0.61	1.63
B380	1.4	2.4	0.36	0.37
B401	11.1	13.3	0.24	0.32
B402	5.5	14.9	0.76	1.23
B406	1.4	1.8	1.06	0.79

Table S3: Summary statistics regarding the AMS chemical composition for each zone referred to the main paper. Concentrations are reported in $\mu\text{g sm}^{-3}\text{t}$ at the 25th, 50th and 75th percentiles.

Species	Statistic	Zones						
		1	2	3	4	5	6	7
Organics	25 th	3.69	0.57	3.21	3.49	2.65	2.84	0.66
	50 th	4.13	0.91	3.69	4.13	3.63	3.68	1.40
	75 th	4.55	1.29	4.40	4.83	4.96	4.39	3.40
Nitrate	25 th	3.04	0.58	1.15	2.10	0.24	0.33	0.05
	50 th	3.51	2.11	2.92	4.84	1.24	1.00	0.14
	75 th	4.40	3.02	4.98	6.85	2.72	3.16	1.45
Sulphate	25 th	2.78	0.83	2.88	1.99	0.98	1.13	0.49
	50 th	3.25	1.20	3.82	3.02	1.93	1.44	0.92
	75 th	3.65	1.78	4.89	3.65	4.12	1.93	2.12
Ammonium	25 th	2.60	0.55	1.48	1.67	0.51	0.61	0.17
	50 th	2.87	1.14	2.06	2.25	1.11	1.23	0.45
	75 th	3.06	1.47	3.06	3.36	2.03	1.77	1.46

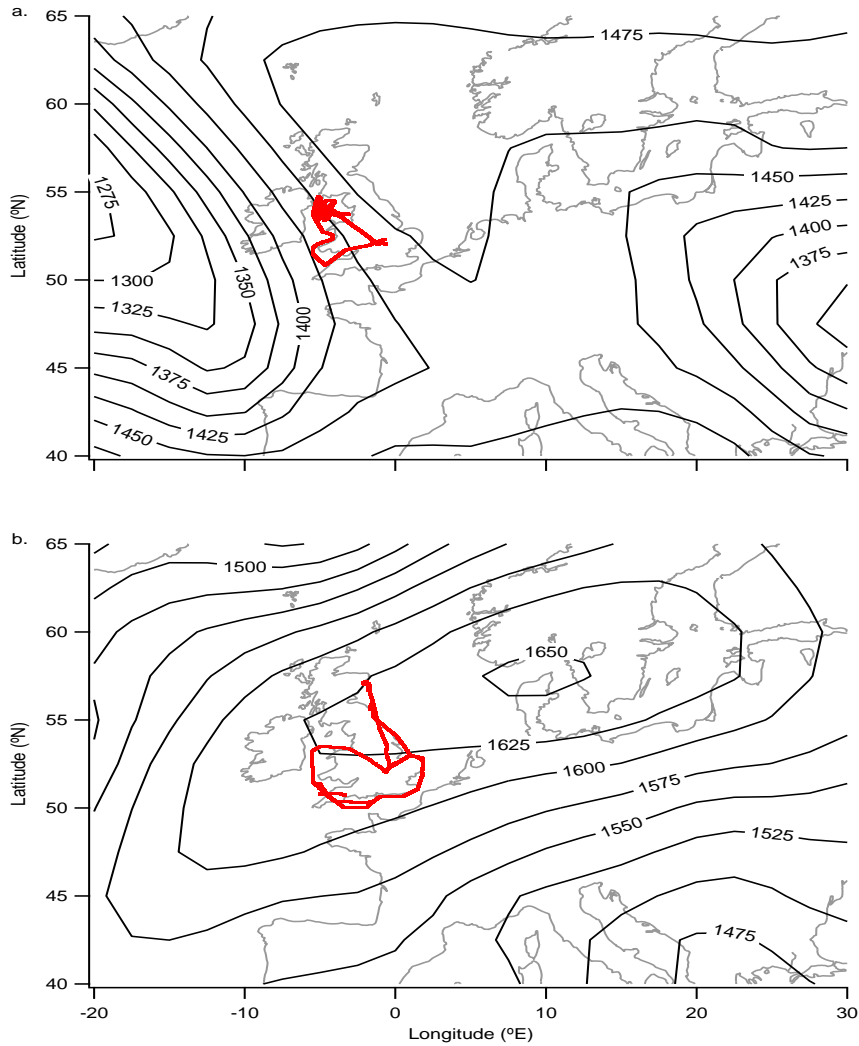


Figure S1: Flight tracks of the BAe-146 considered by this analysis for the ADIENT periods. Also shown are ECMWF 850 hPa geopotential height fields for each period considered by the analysis, where the field is either pertinent to a particular flight or is representative of the overall meteorological regime of the period. All geopotential height fields are for 12UTC. (a) summarises the flights for the UK-based ADIENT flying in April 2008 and the geopotential height field is from 16 April 2008 (B357). (b) summarises the flights for the UK-based ADIENT flying in September 2008 and the geopotential height field is from 25 September 2008 (B406).

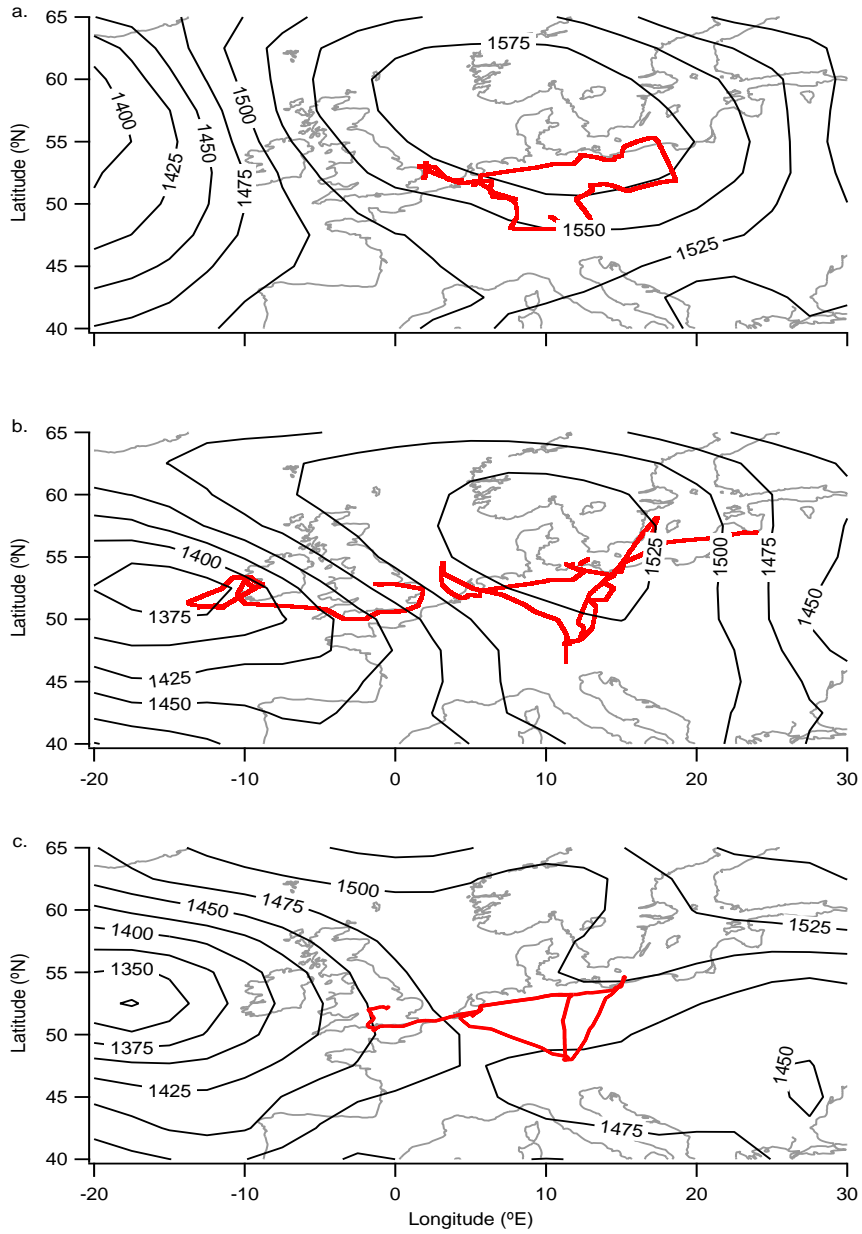


Figure S2: Same plots as Fig. S1 but now for the LONGREX flying period. (a) summarises the flights for the LONGREX-1 period with a geopotential height field for the 06 May 2008. (b) summarises the flights for the LONGREX-2 period with a geopotential height field for the 14 May 2008. (c) summarises the flights for the LONGREX-3 period with a geopotential height field for the 22 May 2008.

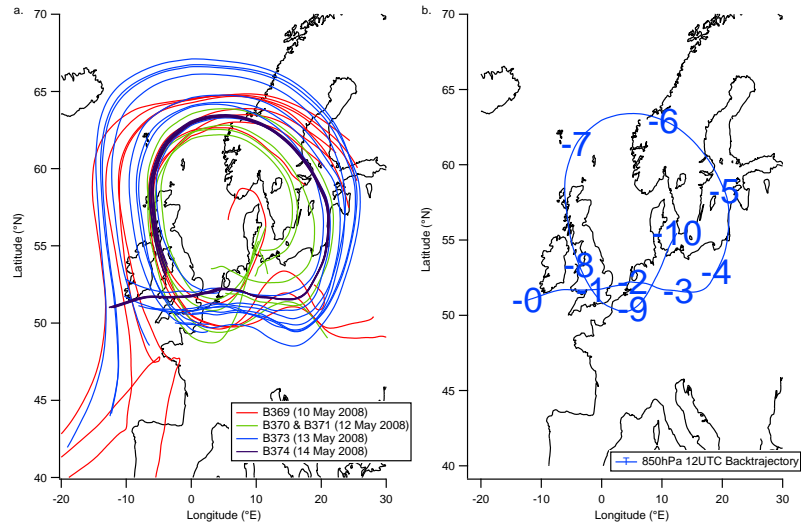


Figure S3: (a) Back trajectories at 850 hPa initialised from 1200 UTC on each flight day from the LONGREX-2 period. The initialisation points correspond to several SLRs during each flight. (b) Air mass back trajectory initialised from 1200 UTC on 14 May 2008 at 850 hPa. The numbered points relate to the number of days passed since the air mass was in that location.

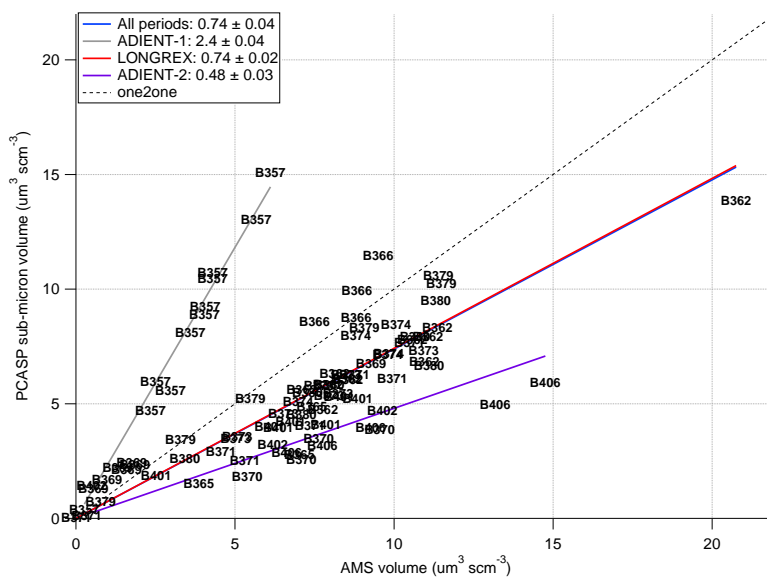


Figure S4: Comparison of inferred volume from the AMS with estimated sub-micron volume derived from the PCASP. The markers refer to SLRs below 3000 m i.e. within the boundary layer. The text markers refer to the flight which the point is from. Linear regression lines are shown for both individual flying periods and the study as a whole.

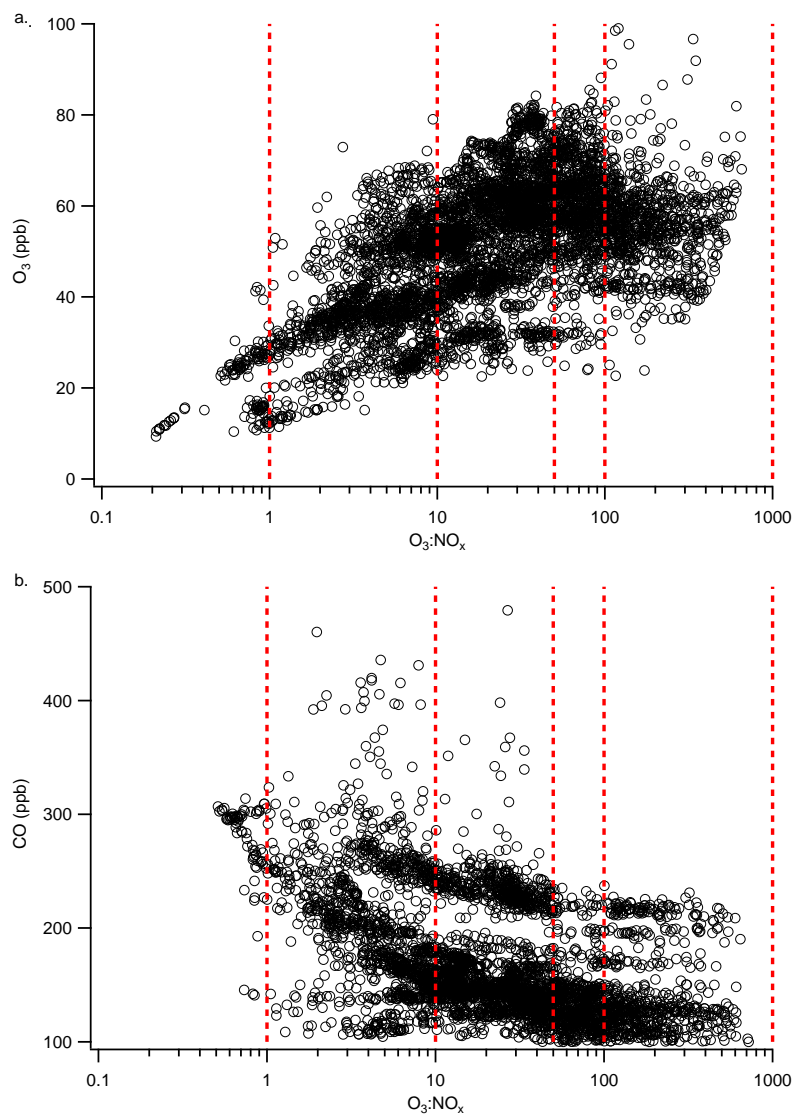


Figure S5: (a) Relationship between O_3 and $O_3:NO_x$ for all flights. (b) Relationship between CO and $O_3:NO_x$ for all flights except for ADIENT-2. The red dashed lines refer to the distance from source boundaries discussed in the main paper.

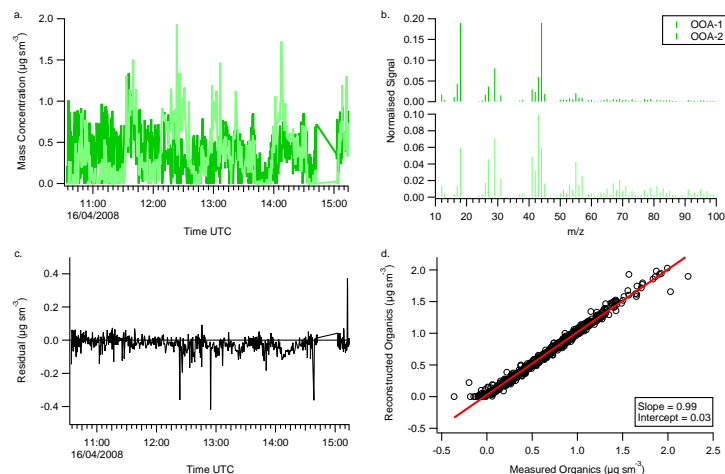


Figure S6: PMF solution for the two factor case from flight B357 including factor component time series (a) and mass spectra (b). The absolute residual is also shown in (c), whilst a comparison between the factor analysis reconstructed mass and measured organic signal is displayed in (d).

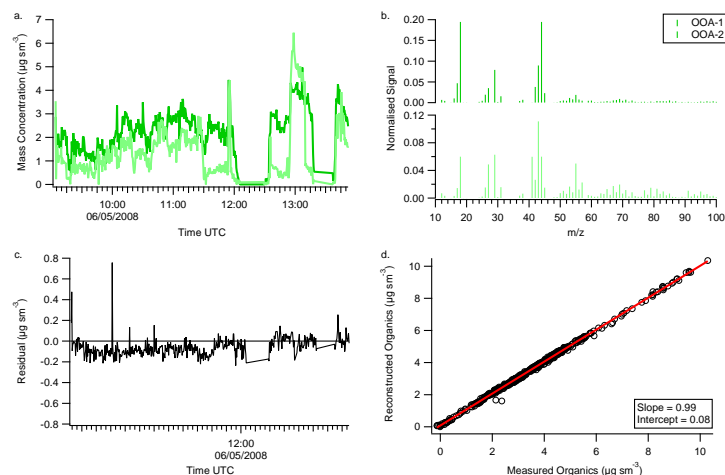


Figure S7: PMF solution for the two factor case from flight B362 including factor component time series (a) and mass spectra (b). The absolute residual is also shown in (c), whilst a comparison between the factor analysis reconstructed mass and measured organic signal is displayed in (d).

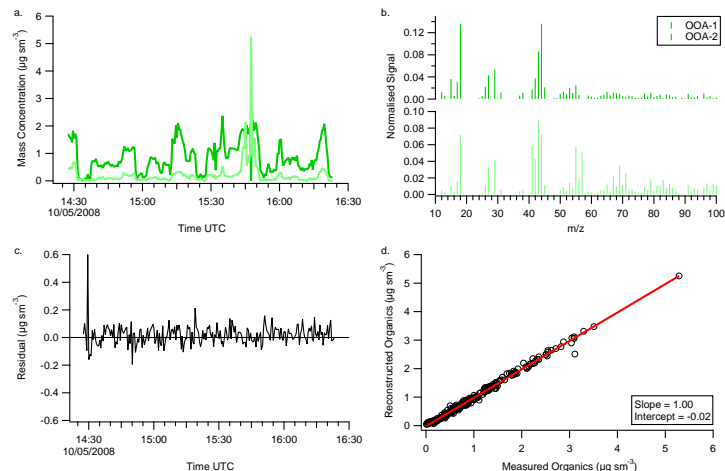


Figure S8: PMF solution for the two factor case from flight B369 including factor component time series (a) and mass spectra (b). The absolute residual is also shown in (c), whilst a comparison between the factor analysis reconstructed mass and measured organic signal is displayed in (d).

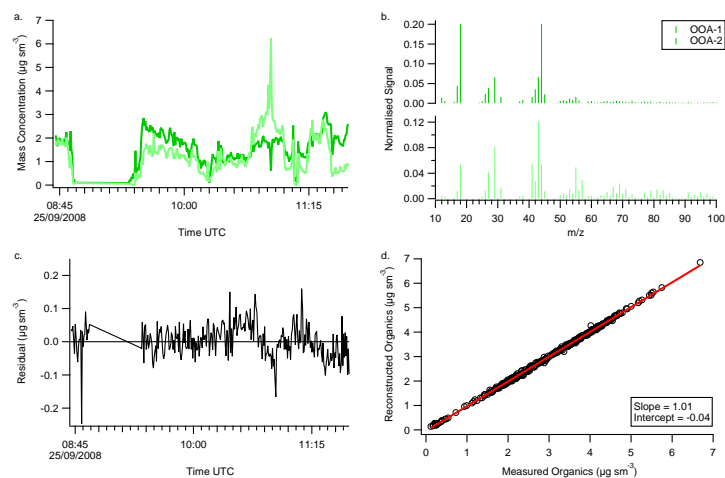


Figure S9: PMF solution for the two factor case from flight B406 including factor component time series (a) and mass spectra (b). The absolute residual is also shown in (c), whilst a comparison between the factor analysis reconstructed mass and measured organic signal is displayed in (d).

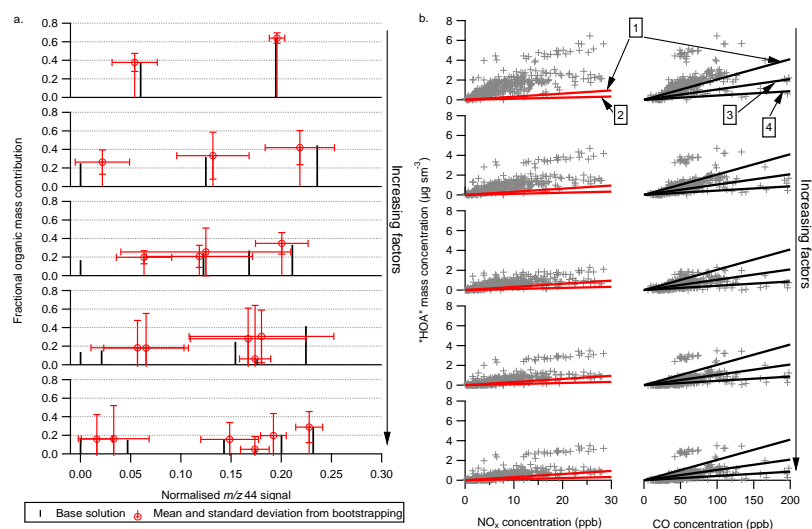


Figure S10: **(a)** Example from flight B362 of the relationship between the fractional mass contribution of a given factor to its normalised signal at m/z 44 for PMF solutions from two through seven factors. The black vertical bars refer to the base solution, while the red vertical and horizontal bars are the results from a resampling analysis known as bootstrapping. Increased standard deviations and mismatching between the base and bootstrapping solutions suggest a numerical unstable solution. **(b)** Relationship between the most HOA like factor profile with NO_x (red line) and CO (black line) for the factor solutions in (a). Solid red and black lines refer to literature emission ratios where (1) is from Allan et al. (2010), (2) is from Kirchstetter et al. (1999), (3) is from Lanz et al. (2007) and (4) is from Zhang et al. (2005b).

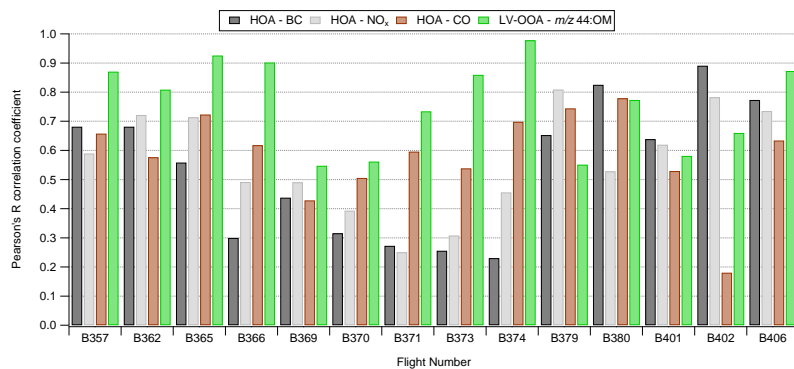


Figure S11: Summary of correlations for estimated HOA with Black Carbon (BC), NO_x and CO. Also shown are the correlations for LV-OOA organic mass fractions with the normalised organic signal at m/z 44.

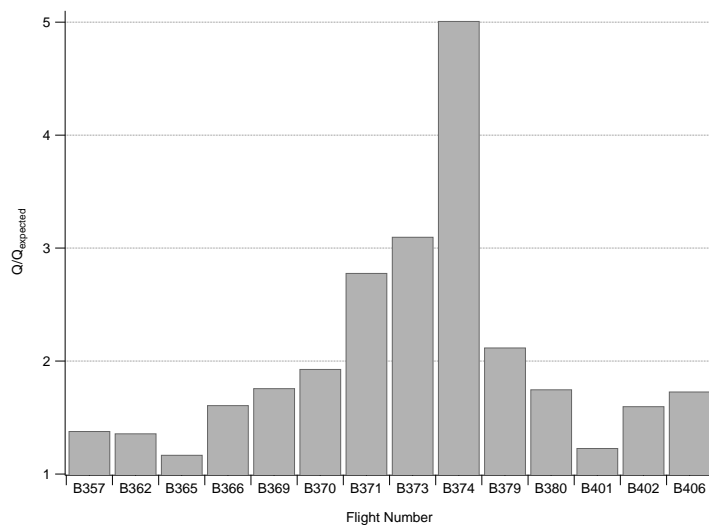


Figure S12: Summary of the $Q/Q_{expected}$ parameter for each flight in the dataset. All values are for the two-factor solutions with an fPeak of zero.

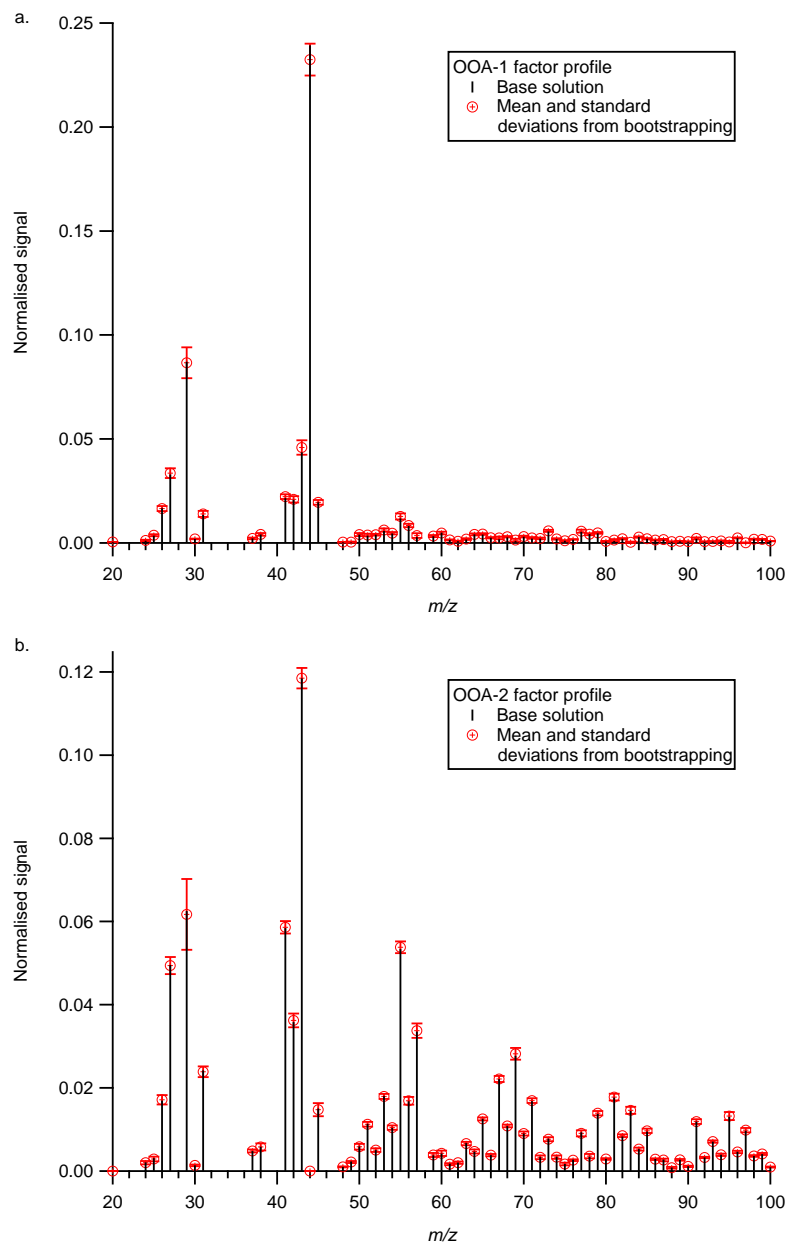


Figure S13: Results of the bootstrapping analysis for the two factor solution mass spectra for flight B357. (a) displays the mass spectrum for OOA-1, while (b) displays the mass spectrum for OOA-2.

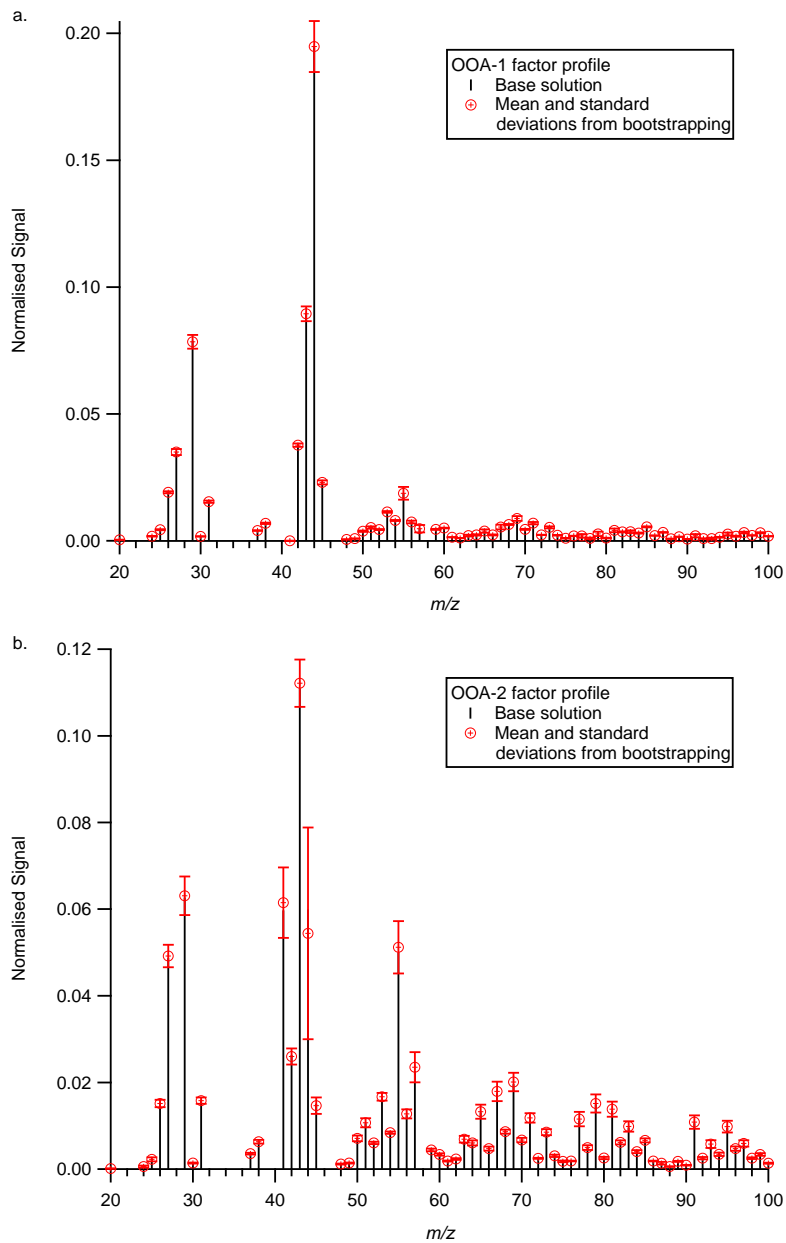


Figure S14: Results of the bootstrapping analysis for the two factor solution mass spectra for flight B362. (a) displays the mass spectrum for OOA-1, while (b) displays the mass spectrum for OOA-2.

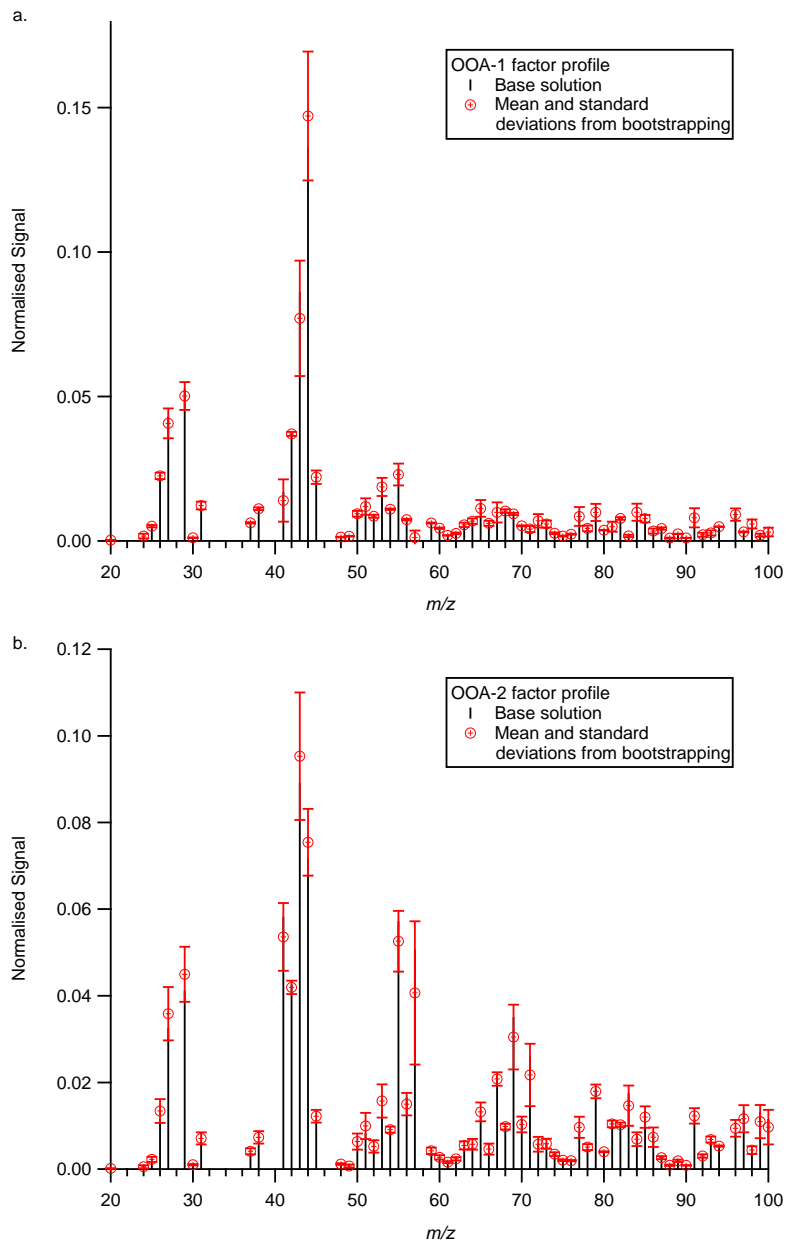


Figure S15: Results of the bootstrapping analysis for the two factor solution mass spectra for flight B369. (a) displays the mass spectrum for OOA-1, while (b) displays the mass spectrum for OOA-2.

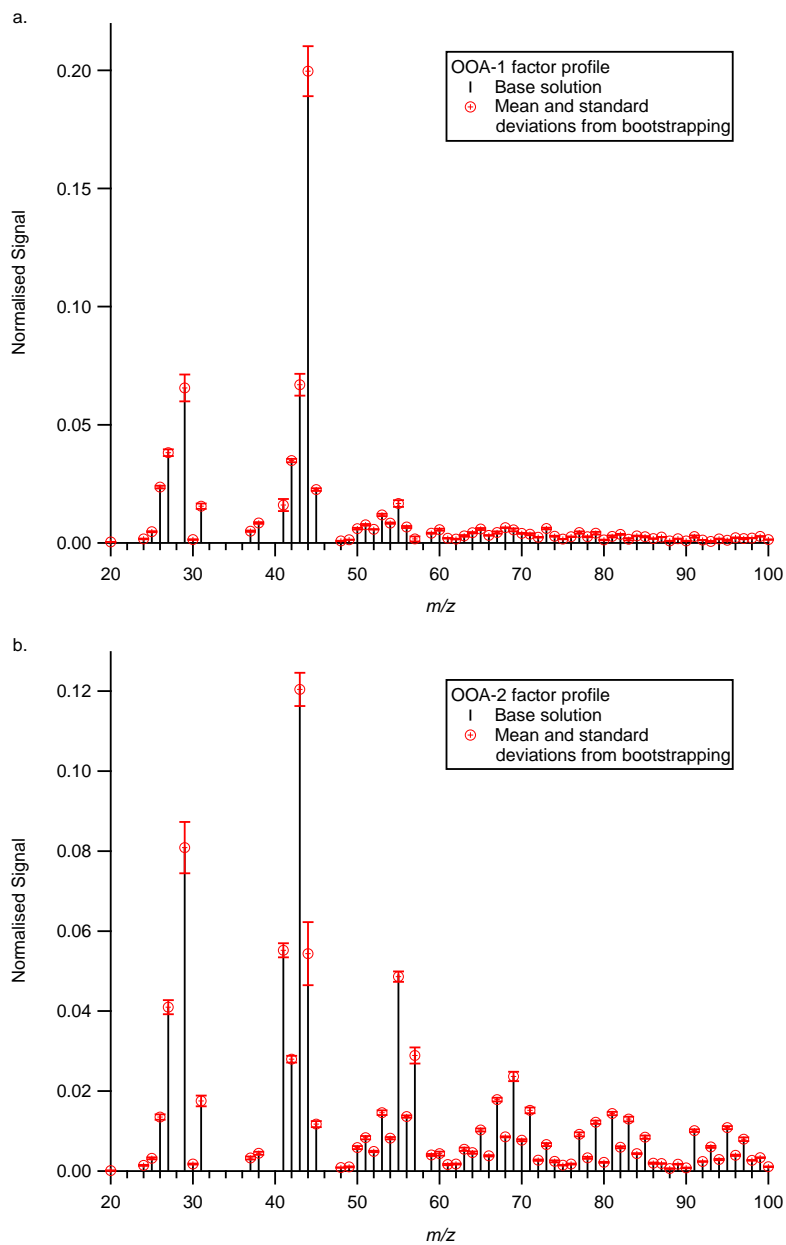


Figure S16: Results of the bootstrapping analysis for the two factor solution mass spectra for flight B406. (a) displays the mass spectrum for OOA-1, while (b) displays the mass spectrum for OOA-2.